



UNIVERSITY OF
OREGON

Dark Sectors at Muon Colliders

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INAUGURAL US MUON COLLIDER COMMUNITY MEETING

August 8, 2024

Why Muon Colliders for Dark Sectors?

- dark sectors are a sandbox for model building and systematic way of exploring model space
- can give rise to “weird” and unique signatures > help inform detector design
- connection to big questions like nature of dark matter
- muon collider provides synergy to upcoming era of cosmological probes

WIMP Dark Matter

“Minimal Dark Matter”: consider dark matter to be the lightest member of an electroweak multiplet

T. Han, Z. Liu, L.-T. Wang, X. Wang [arXiv:2009.11287]

Model (color, n , Y)		Therm. target	5σ discovery coverage (TeV)				
			mono- γ	mono- μ	di- μ 's	disp. tracks	
Higgsino-like Wino-like	(1,2,1/2)	Dirac	1.1 TeV	—	2.8	—	3.2 – 8.5
	(1,3,0)	Majorana	2.8 TeV	—	3.7	—	13 – 14
	(1,3, ϵ)	Dirac	2.0 TeV	0.9	4.6	—	13 – 14
	(1,5,0)	Majorana	14 TeV	3.1	7.0	3.1	10 – 14
	(1,5, ϵ)	Dirac	6.6 TeV	6.9	7.8	4.2	11 – 14
	(1,7,0)	Majorana	23 TeV	11	8.6	6.1	8.1 – 12
	(1,7, ϵ)	Dirac	16 TeV	13	9.2	7.4	8.6 – 13

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stronger
channel

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Higgsino-like
Wino-like

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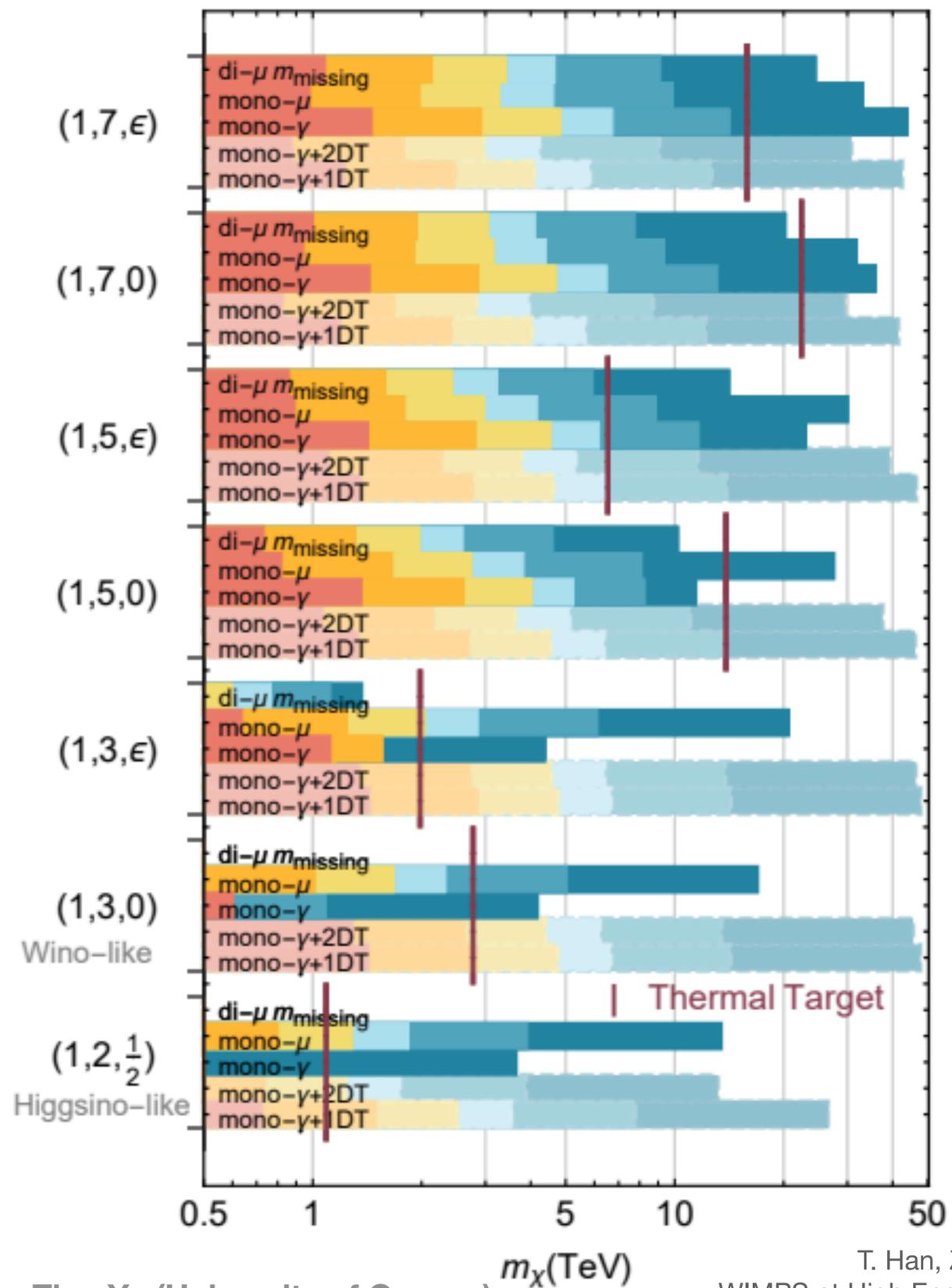
← even more promising!

↑
stronger
channel

further enhancements shown in
R. Capdevilla, F. Meloni, J. Zurita [arXiv:2405.08858]

Muon Collider 2σ Reach

$(\sqrt{s} = 3, 6, 10, 14, 30, 100 \text{ TeV})$

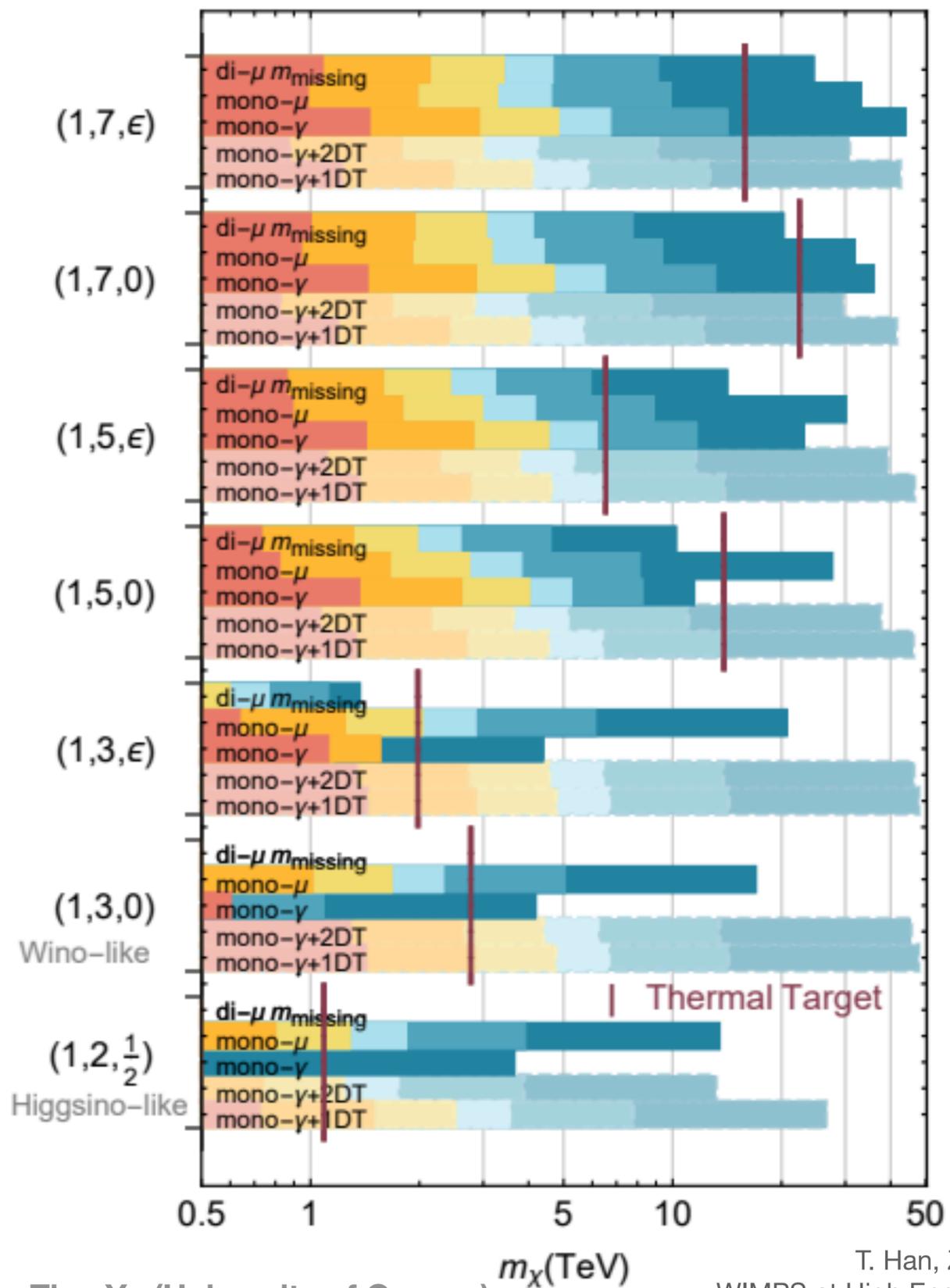


T. Han, Z. Liu, L.-T. Wang, X. Wang

WIMPS at High Energy Muon Colliders [arXiv:2009.11287]

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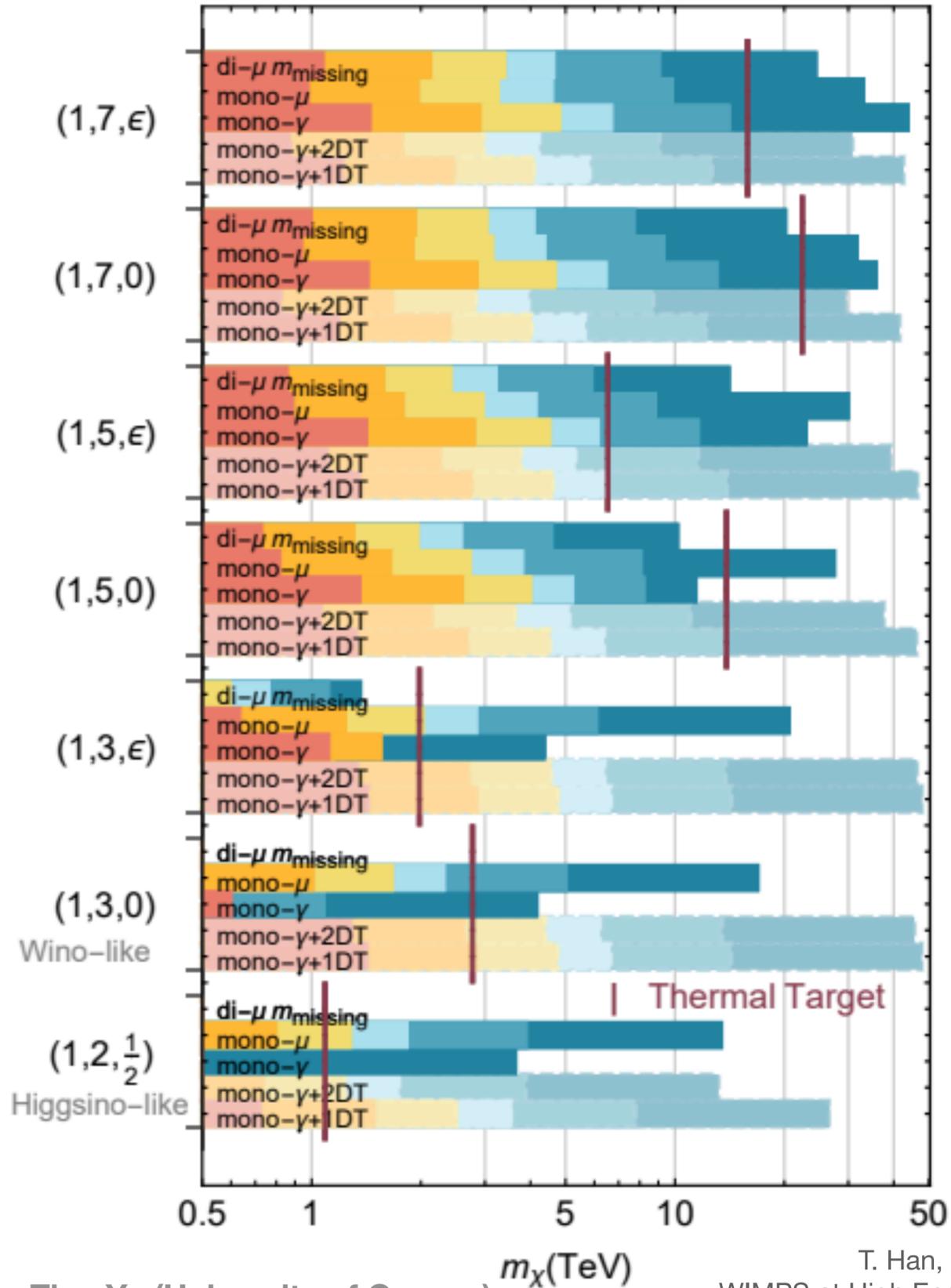
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Muon collider can definitively test Wino and Higgsino WIMPs at 10 TeV!

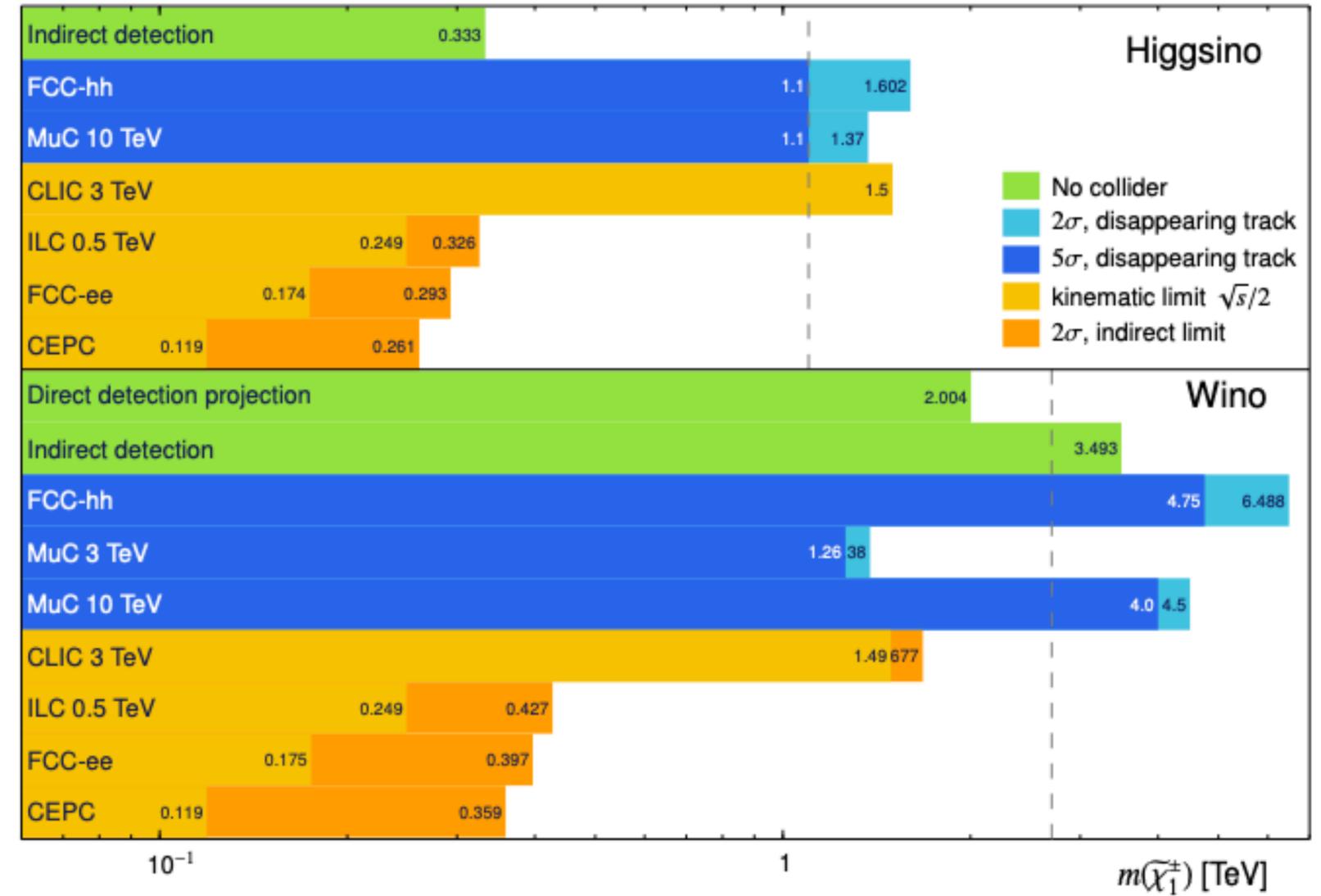
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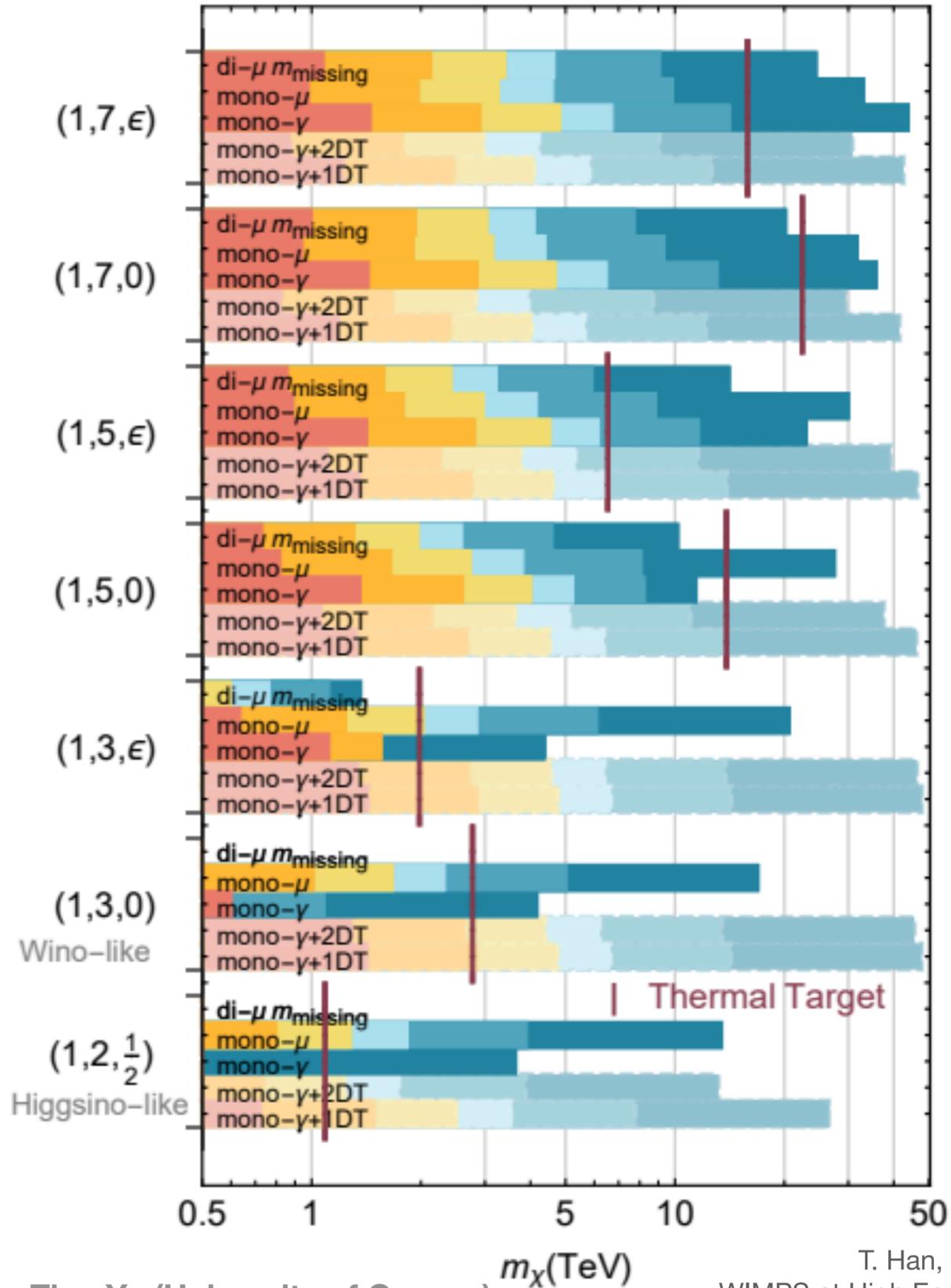
Comparison with other approaches:



R. Capdevilla, F. Meloni, R. Simoniello, and J. Zurita, JHEP 06 (2021) 133, [arXiv:2102.11292]

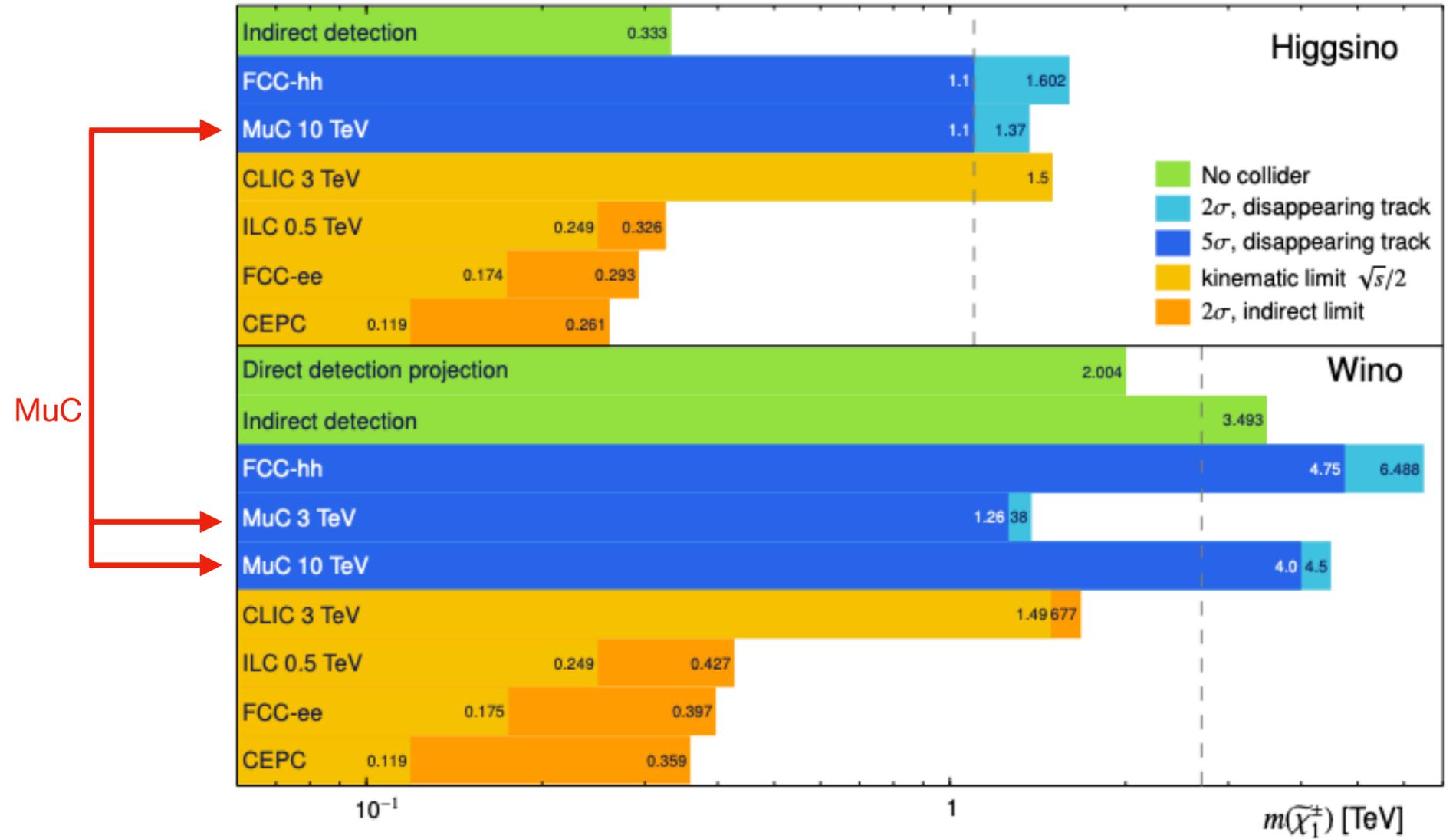
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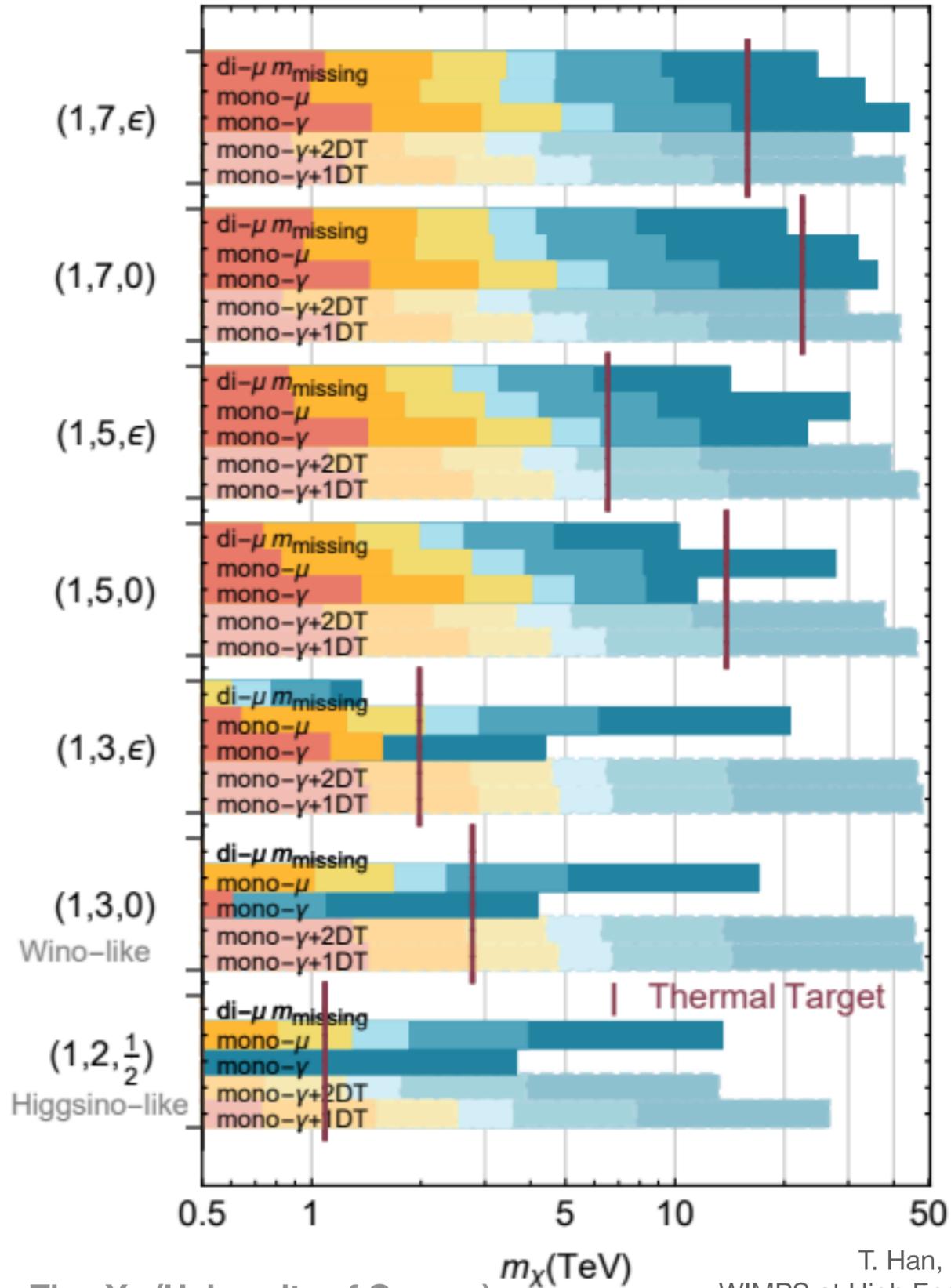
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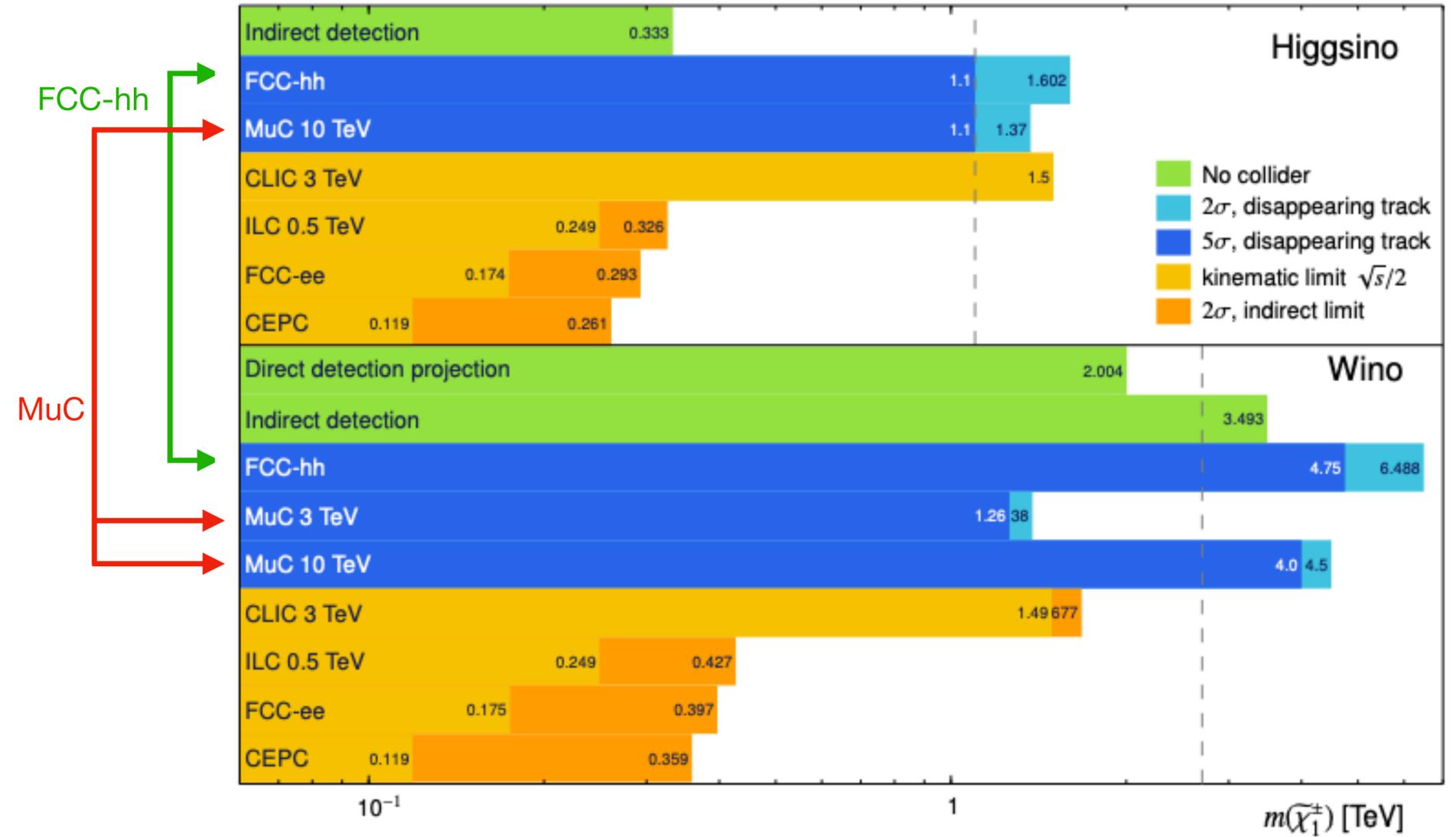
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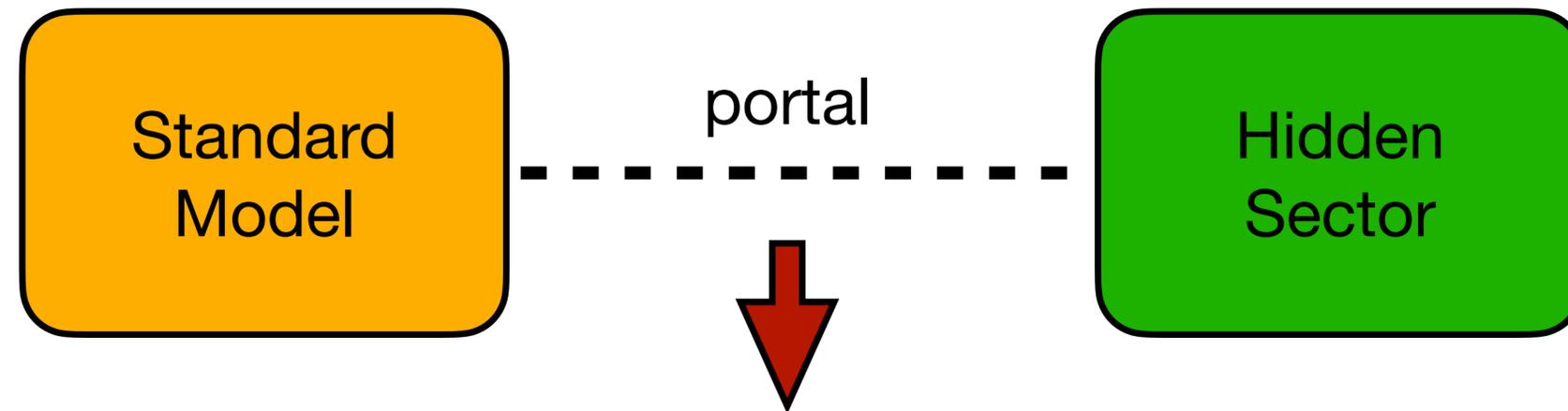
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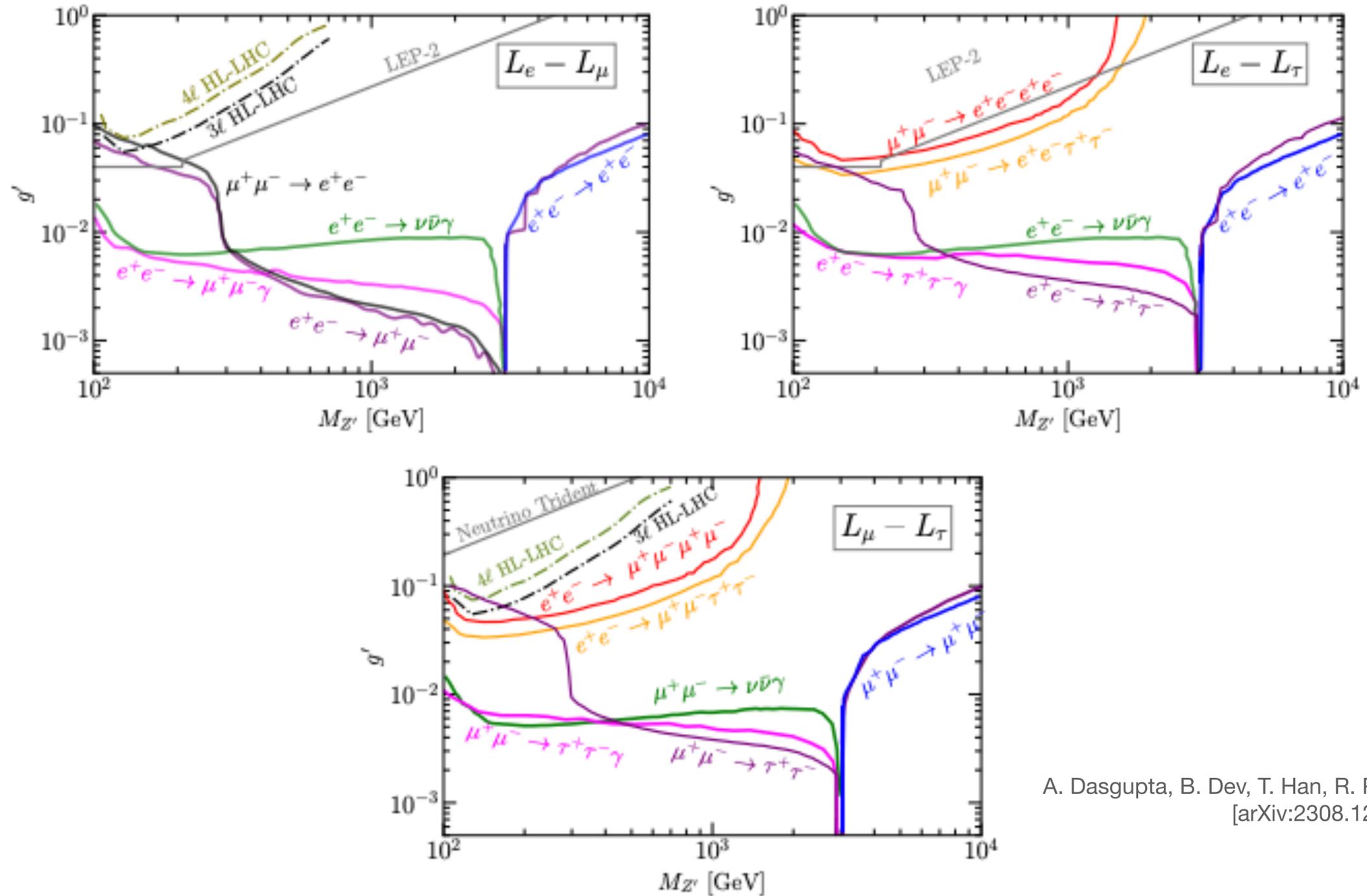
Dark Sector Portals

moving beyond the WIMP



$$\mathcal{L} \supset \begin{cases} (\mu S + \lambda S^2) H^\dagger H & \text{scalar} \\ \frac{a}{f} \tilde{F}_{\mu\nu} F^{\mu\nu} & \text{pseudoscalar} \\ -\frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} F^{\mu\nu} & \text{vector} \\ y_n L H N & \text{neutrino} \end{cases}$$

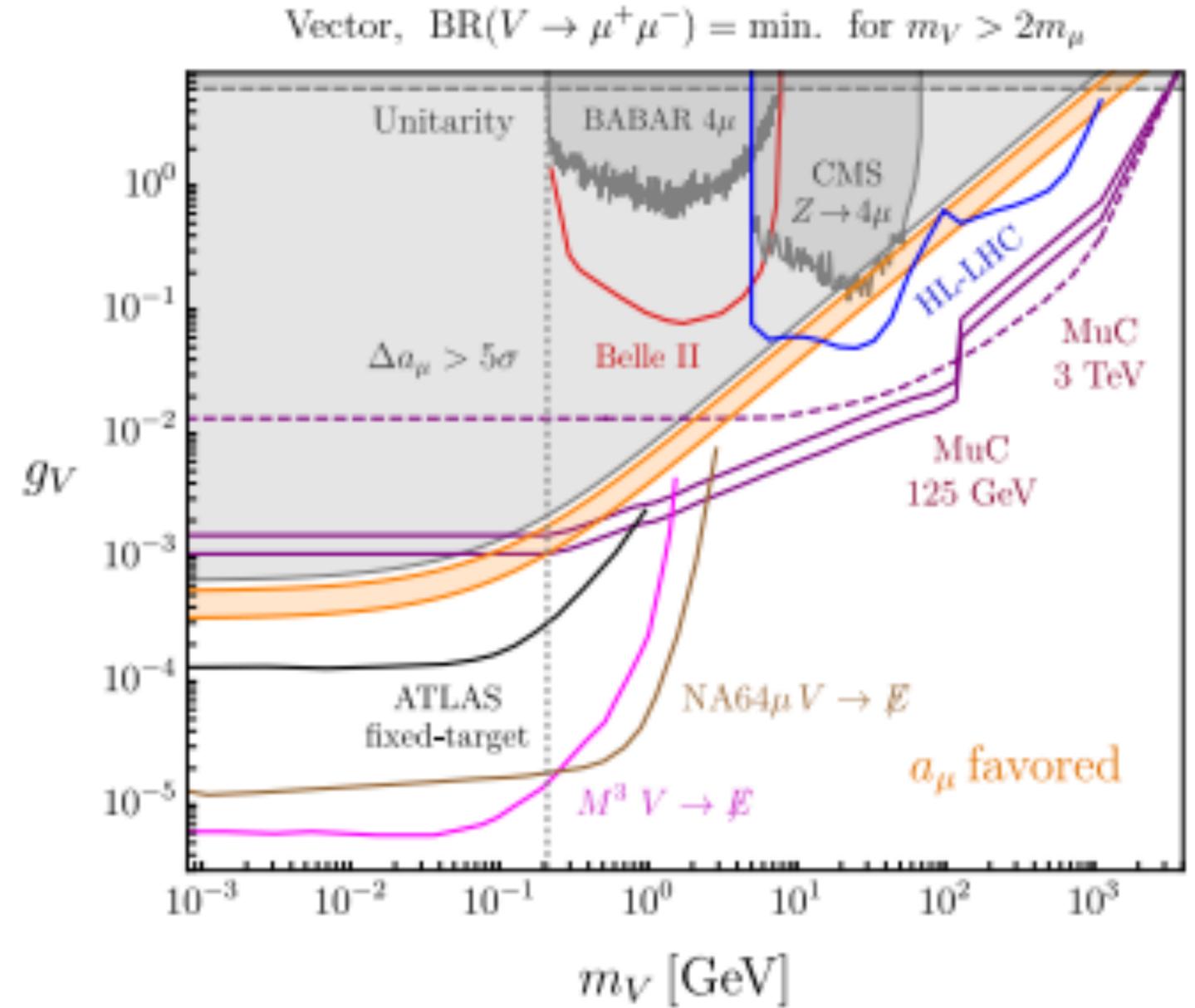
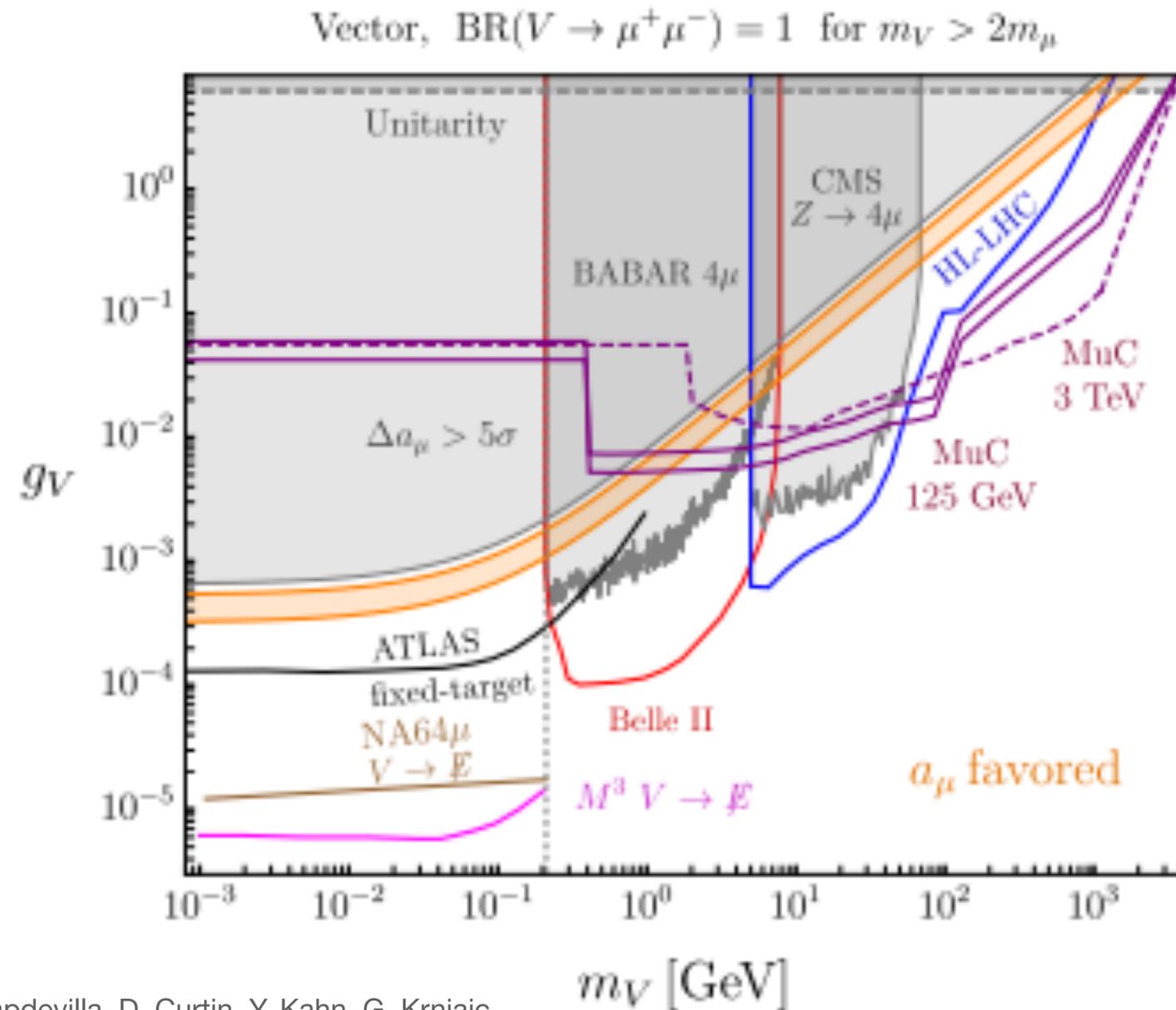
Vector Portal



$$\sqrt{s} = 3 \text{ TeV}$$

A. Dasgupta, B. Dev, T. Han, R. Padhan, S. Wang, K. Xie
[arXiv:2308.12804]

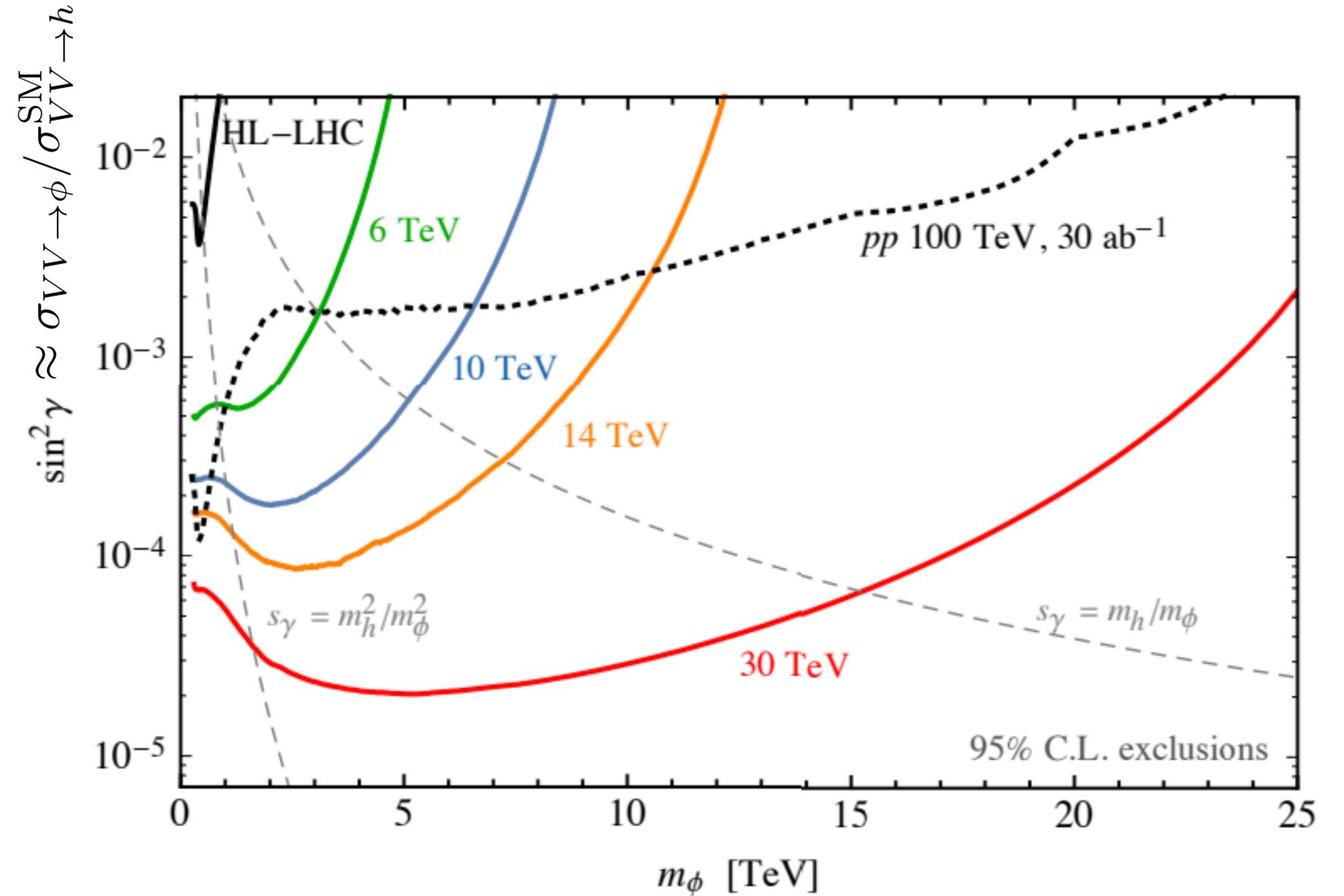
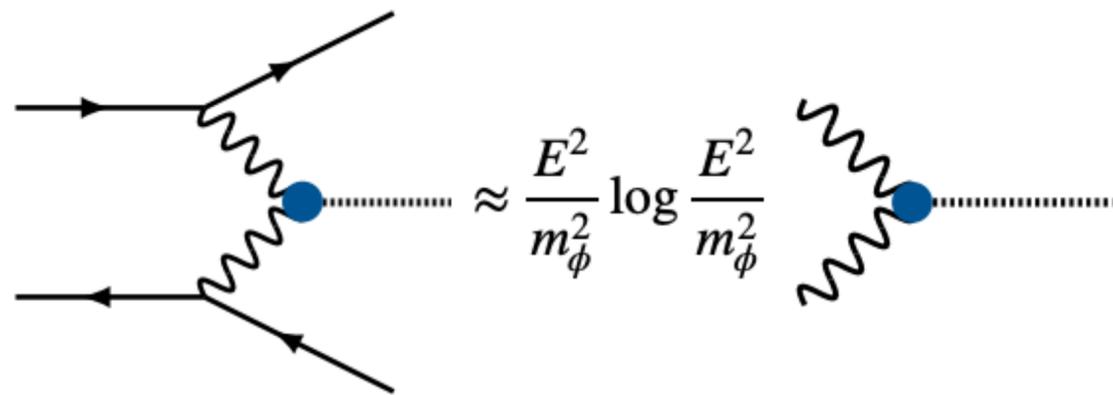
Vector Portal



R. Capdevilla, D. Curtin, Y. Kahn, G. Krnjaic
[arXiv:2112.08377]

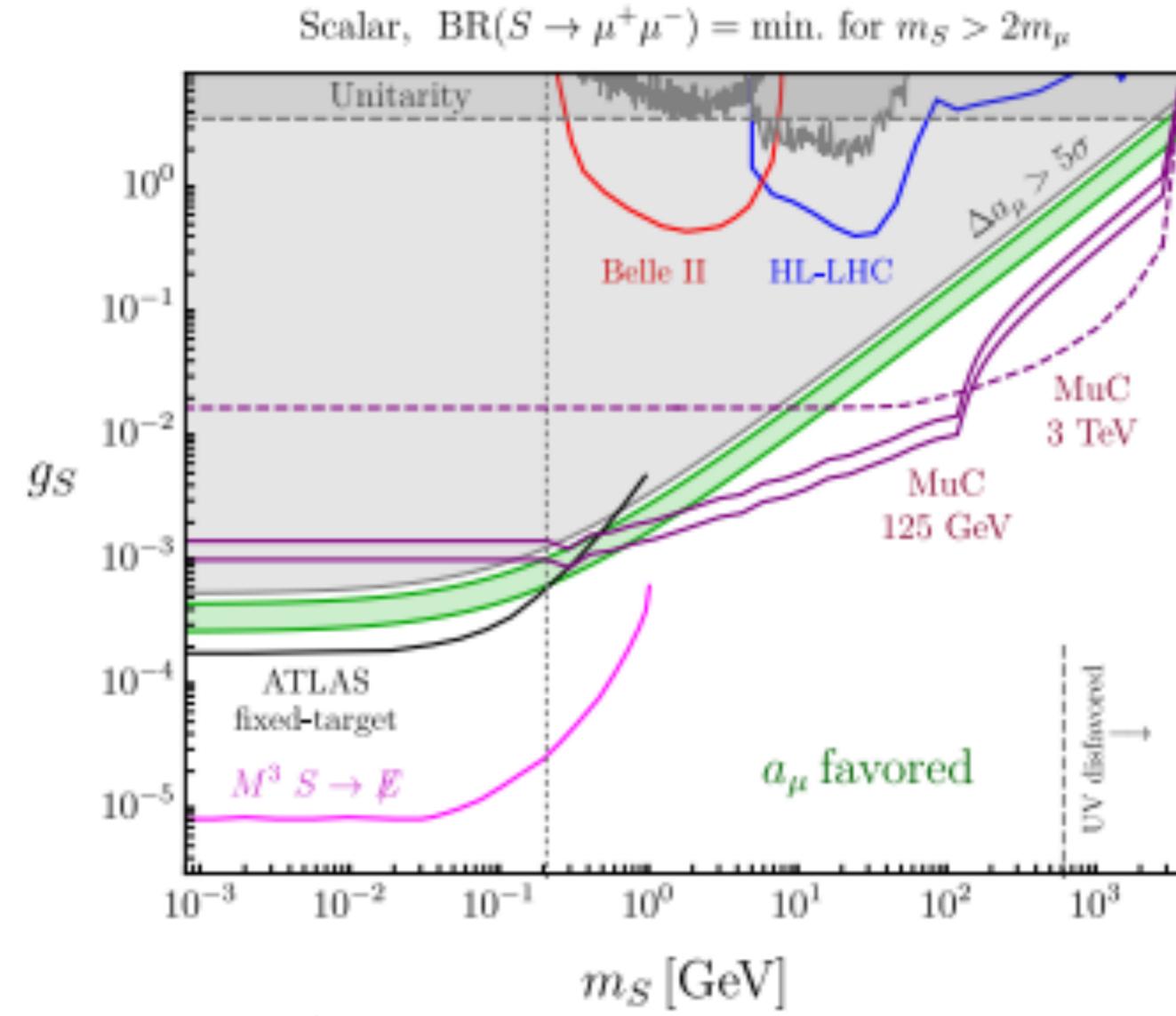
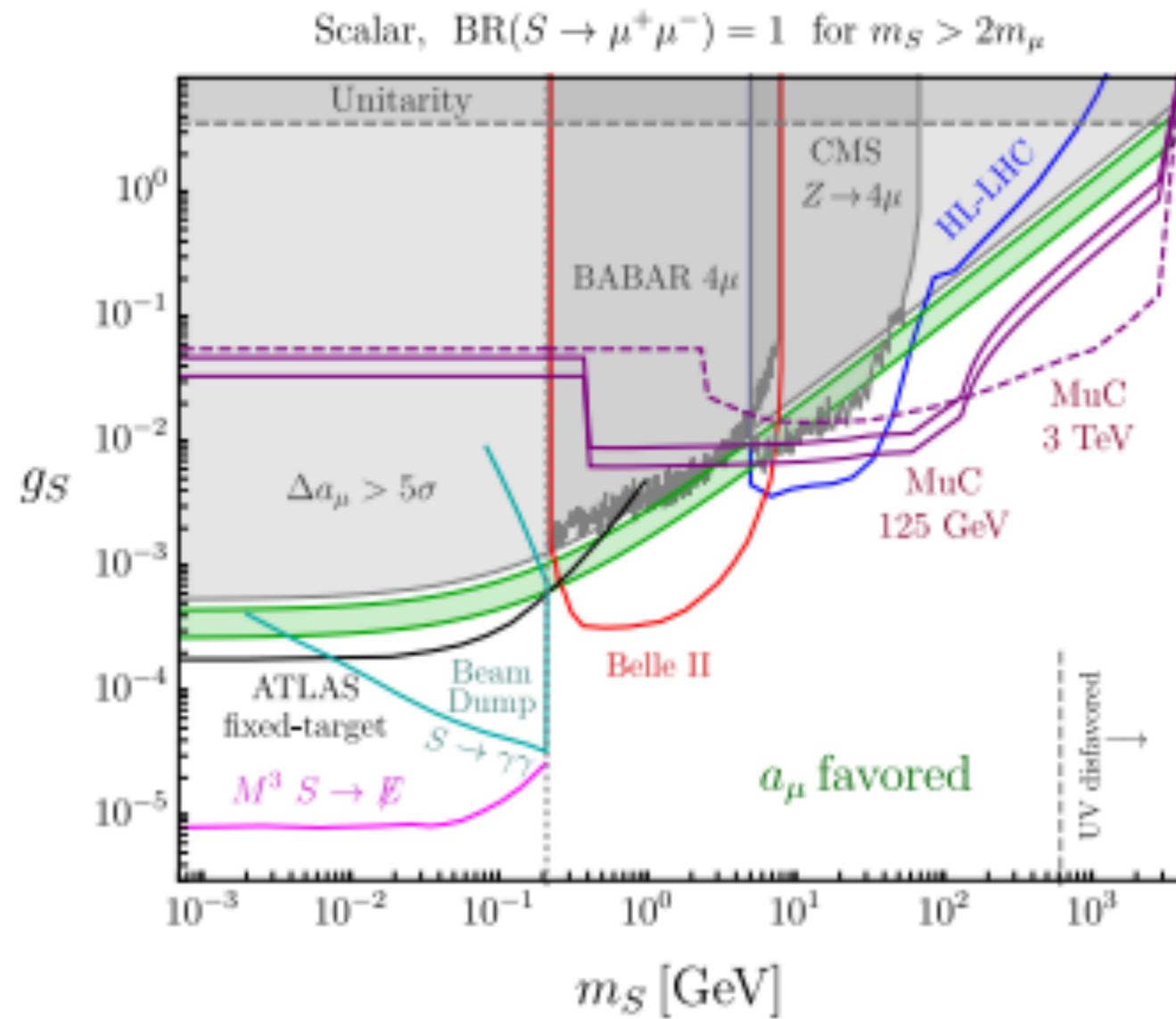
$$\mathcal{L} \supset g_V V_\gamma (\mu_L^\dagger \bar{\sigma}^\gamma \mu_L + \mu^{c\dagger} \bar{\sigma}^\gamma \mu^c)$$

High-Energy Lepton Collider as a “Vector Boson Collider” great for scalar and pseudoscalar portal!



D. Buttazzo, D. Redigolo, F. Sala, A. Tesi
[arXiv:1807.04743] + [arXiv:2103.14043]

Scalar Portal



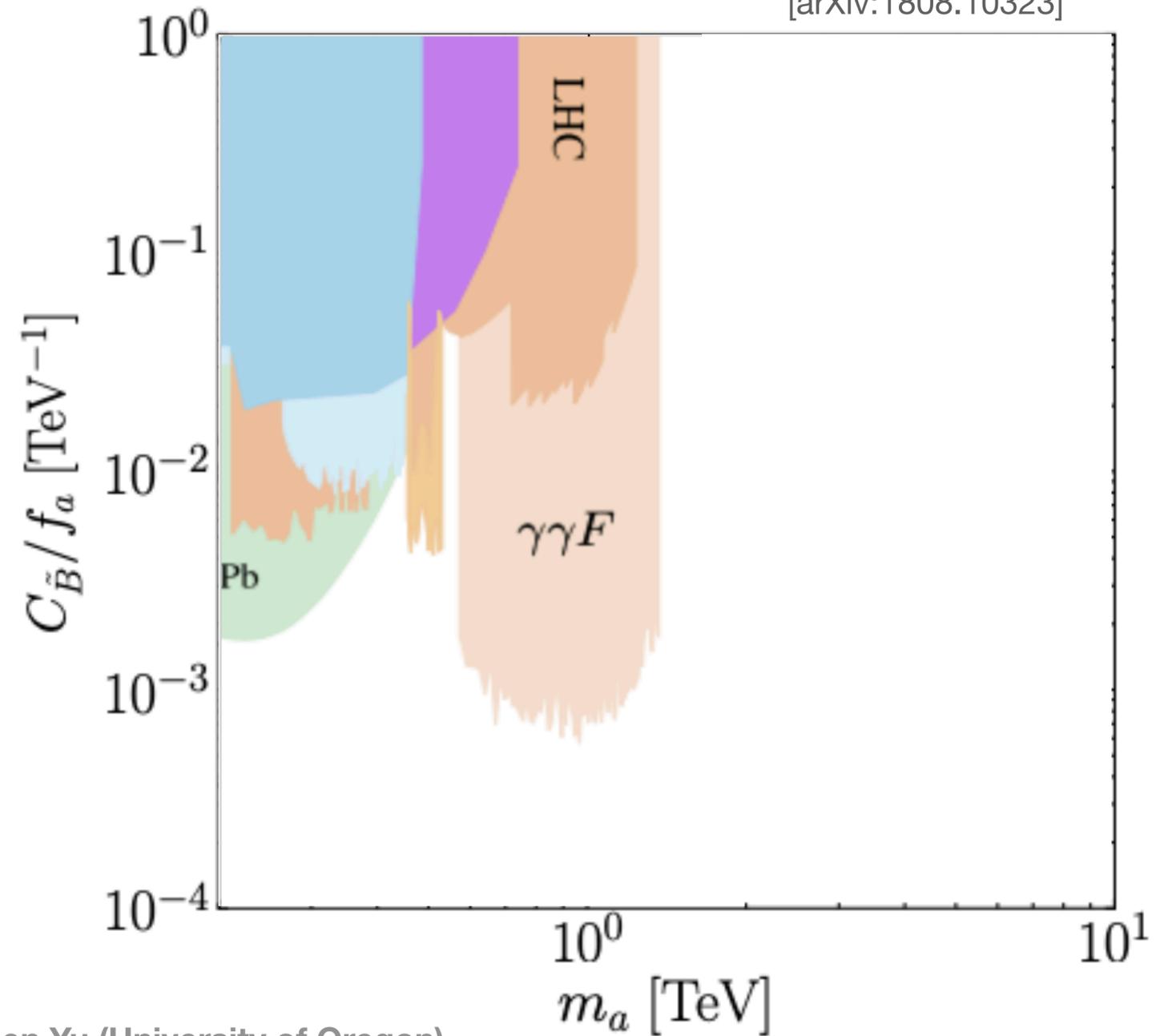
R. Capdevilla, D. Curtin, Y. Kahn, G. Krnjaic
[arXiv:2112.08377]

$$\mathcal{L} \supset g_S S (\mu_L \mu^c + \mu^{c\dagger} \mu_L^\dagger)$$

Pseudoscalar Portal

Axions and ALPs

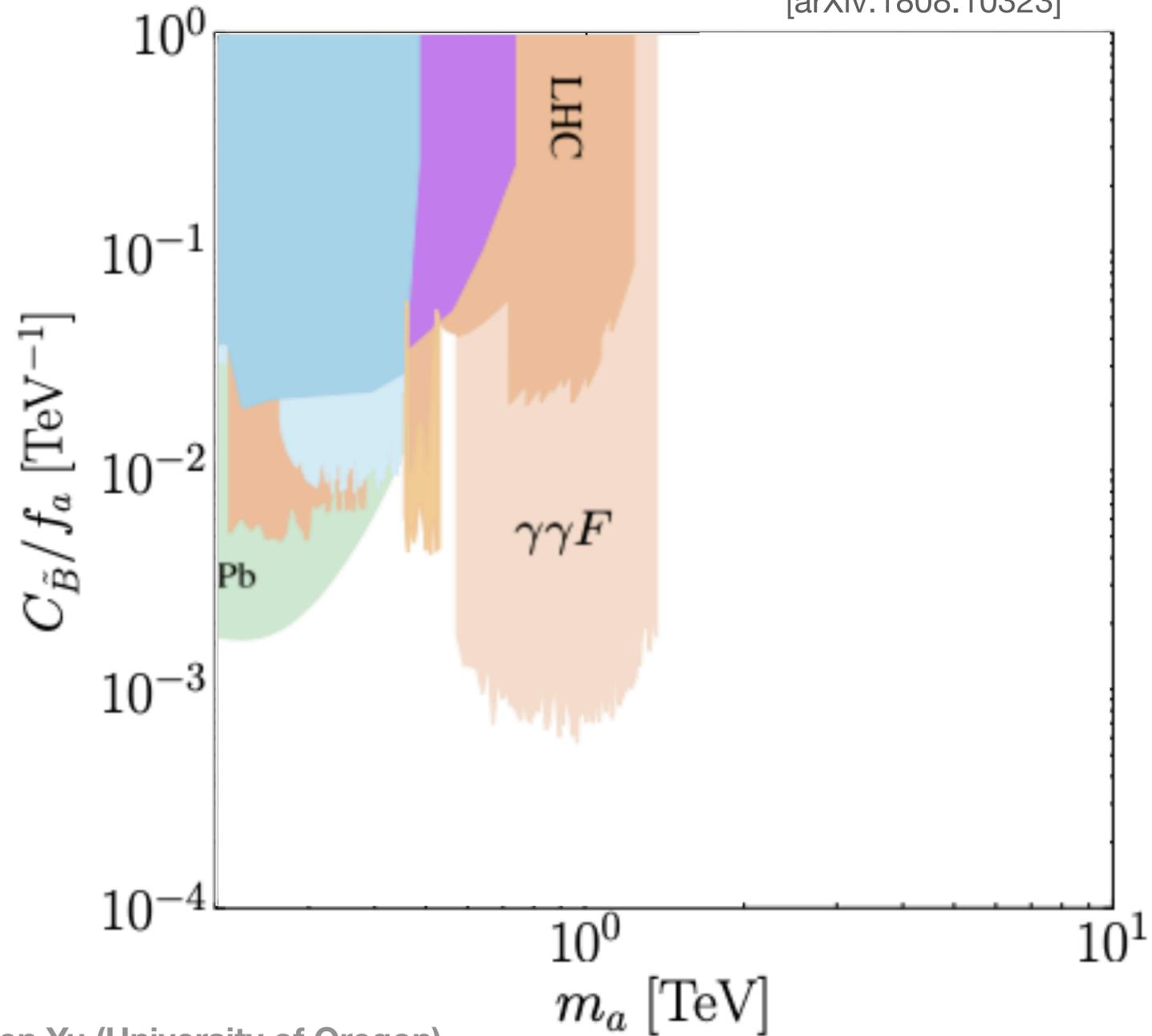
adapted from
M. Bauer, M. Heiles, M. Neubert, A. Thamm
[arXiv:1808.10323]



Pseudoscalar Portal

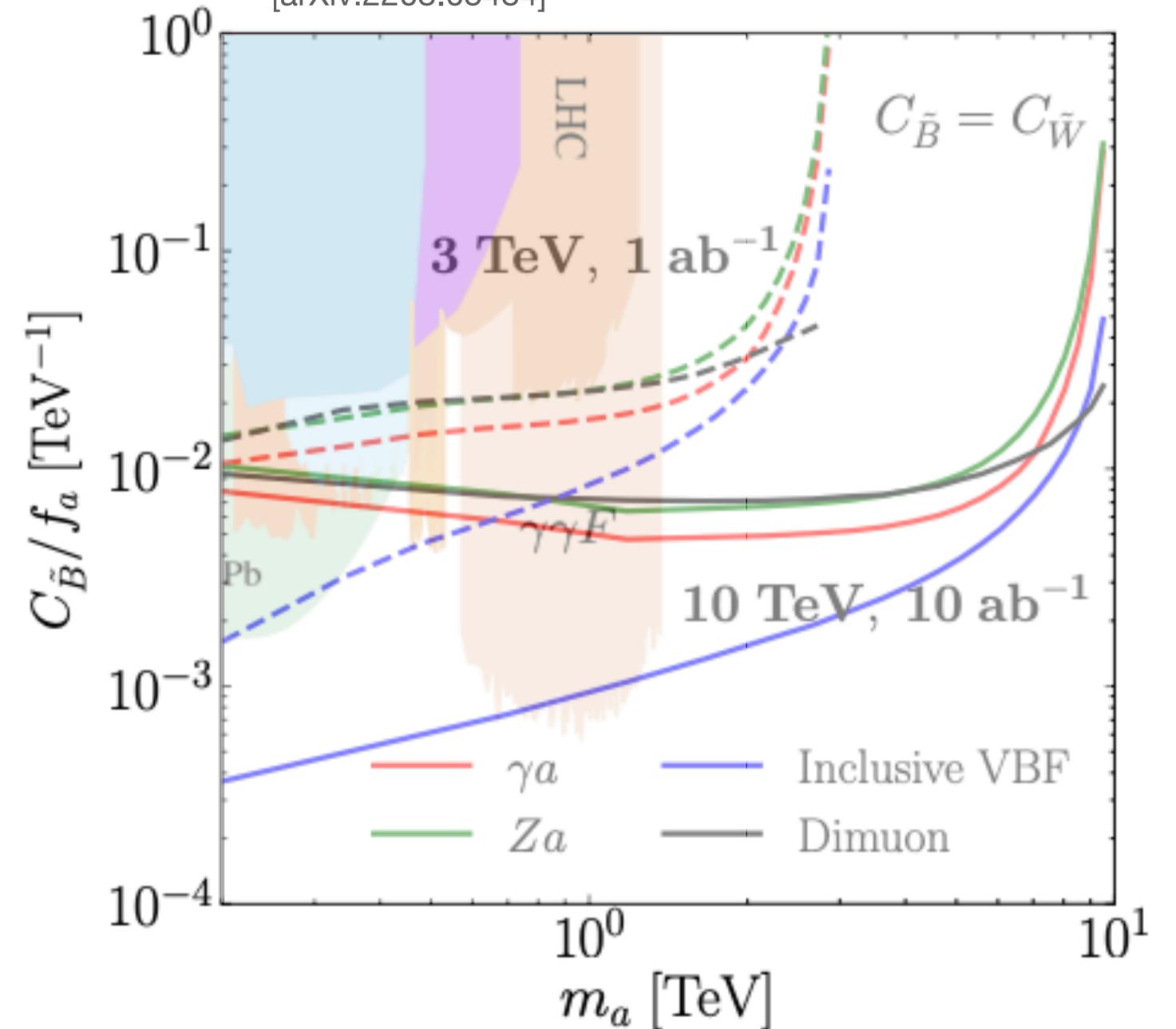
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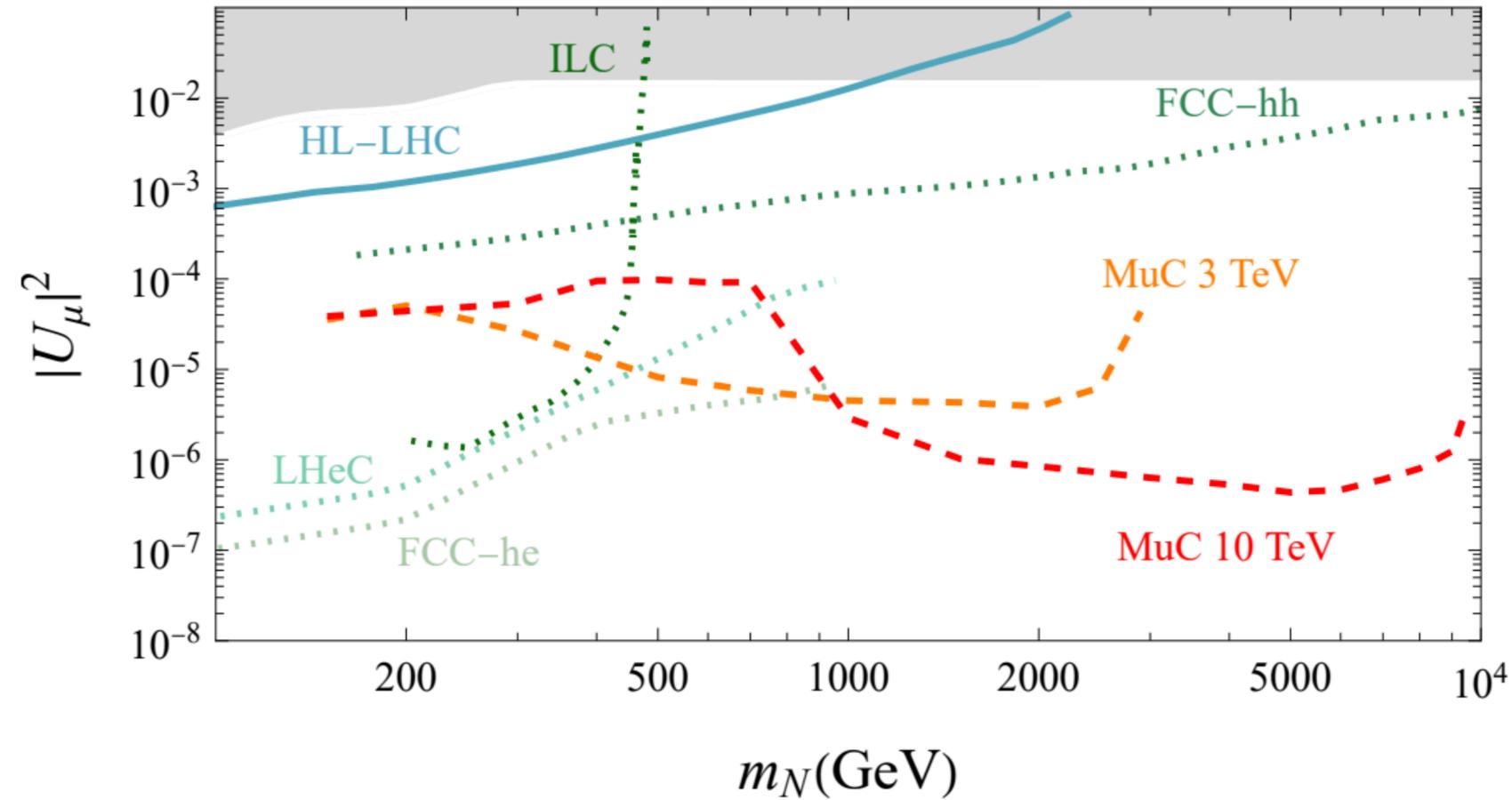
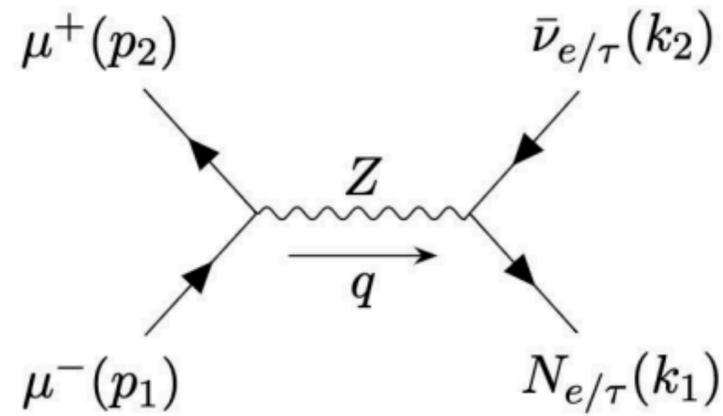
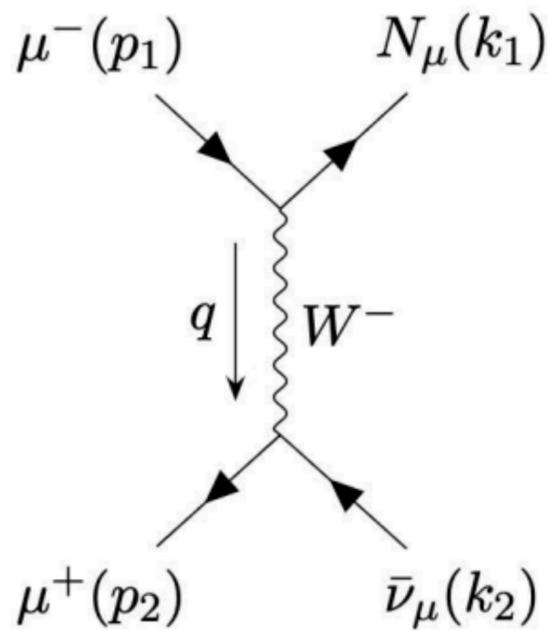
+ T. Han, T. Li, X. Wang
 [arXiv:2203.05484]

see also Y. Bao, J. Fan, L. Li
 [arxiv:2203.04328]



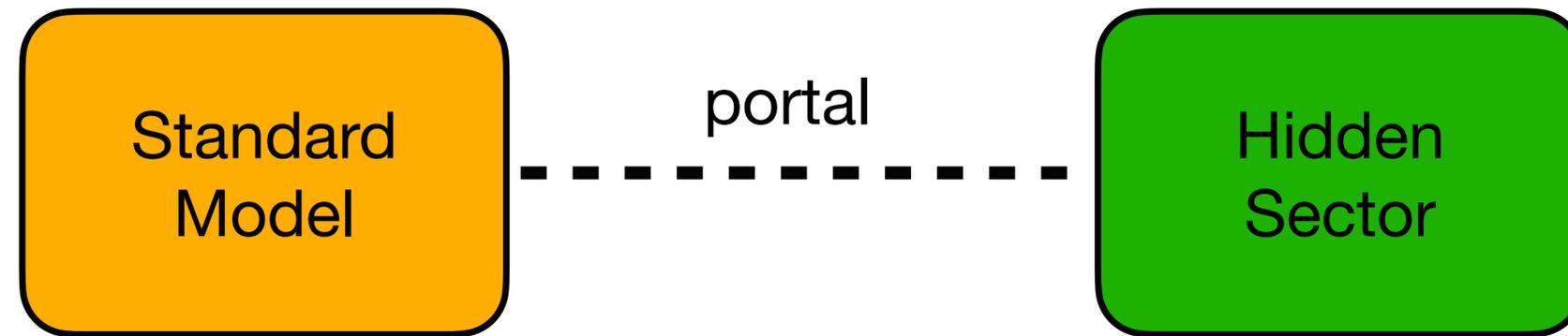
Neutrino Portal

$$\mathcal{L}_\nu \supset -\lambda_\nu \bar{L} \tilde{H} N + h.c.$$



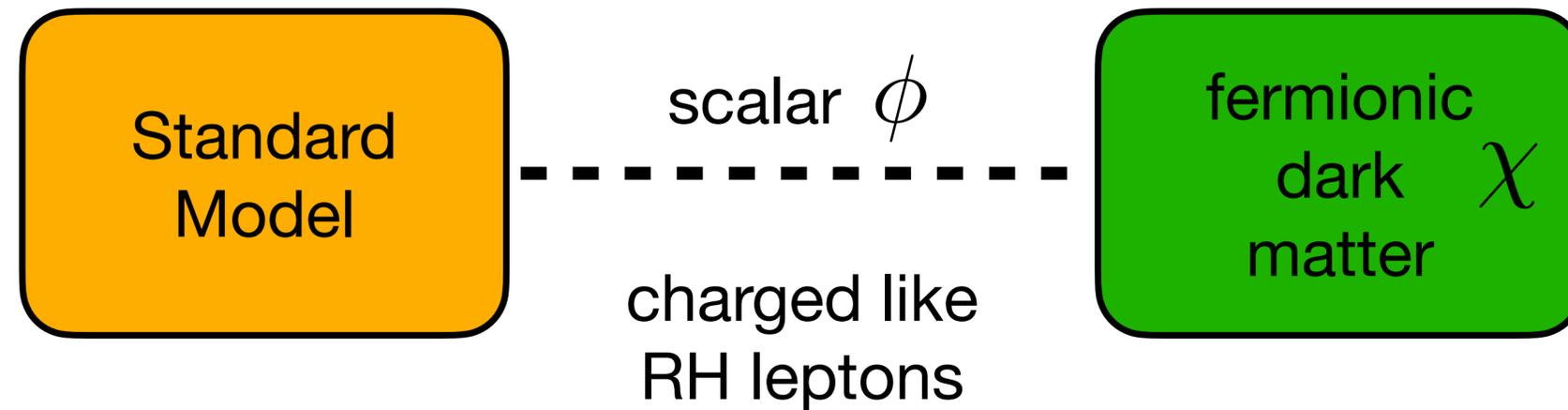
P. Li, Z. Liu, K-F. Lyu
[arXiv:2301.07117]

Case study: flavored dark matter



Other Hidden Sector Models

case study: flavored dark matter

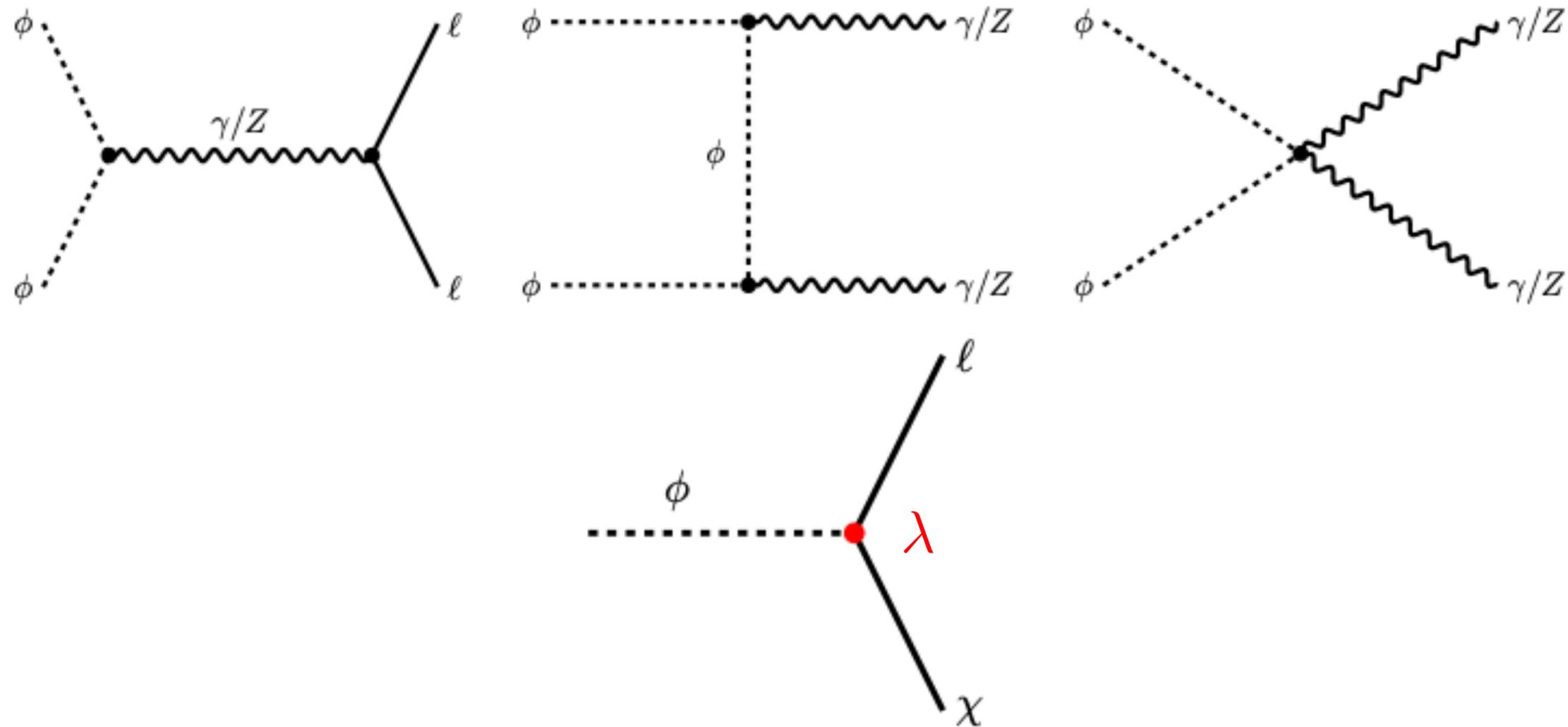


$$\mathcal{L} \supset -m_\chi \bar{\chi}_\alpha \chi^\alpha - m_\phi^2 |\phi|^2 - \lambda_{i,\alpha} \phi \bar{e}_i \bar{\chi}_\alpha$$

\downarrow
flavor indices

Lepton-Flavored Dark Matter

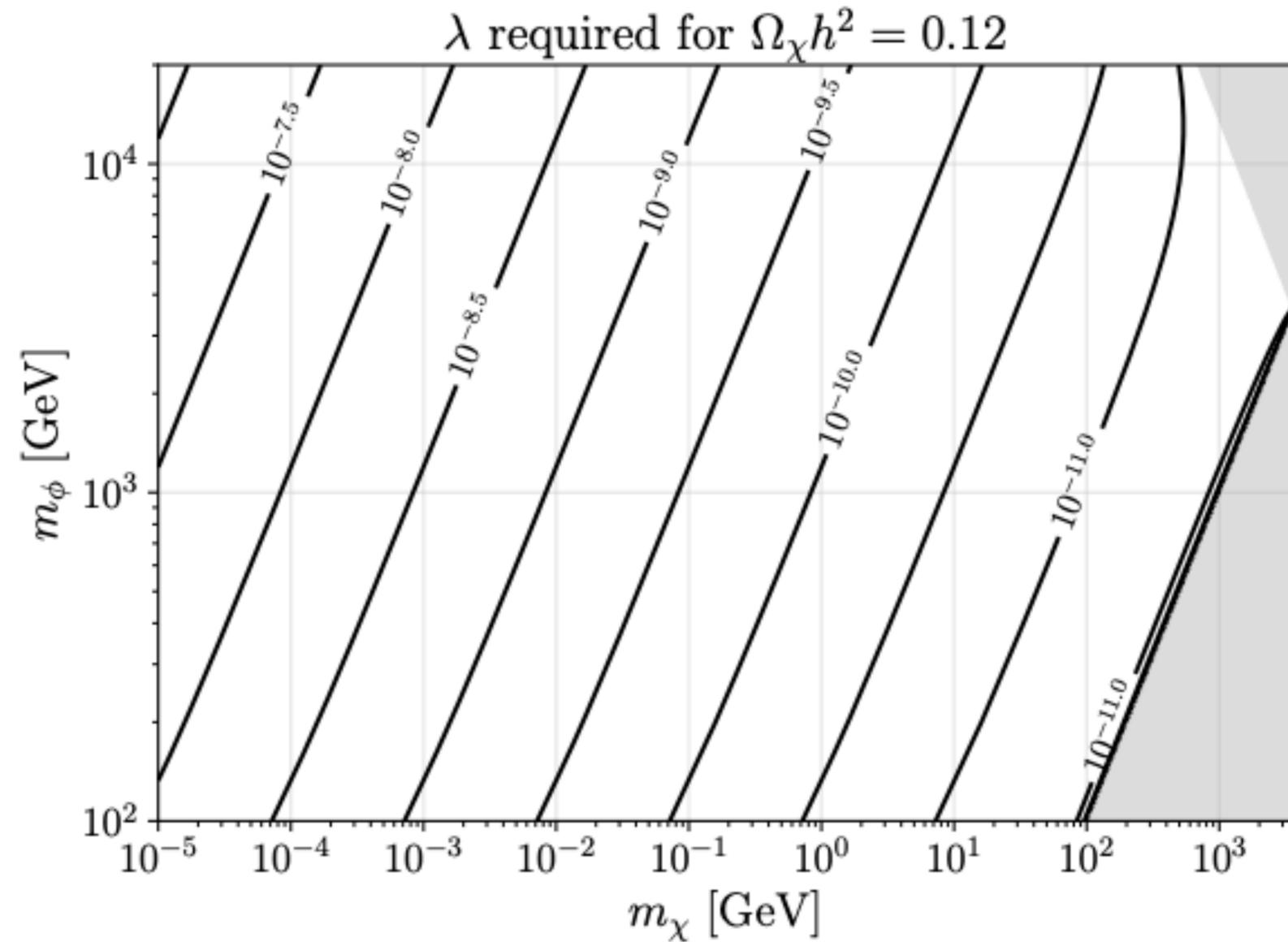
relic abundance



correct relic abundance fixes coupling

Lepton-Flavored Dark Matter

relic abundance

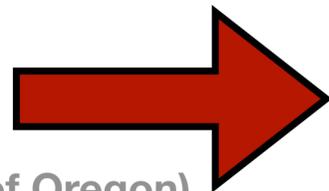


DM must populate
relic abundance

DM parameter space
is fully bounded!

DM must be stable

small λ

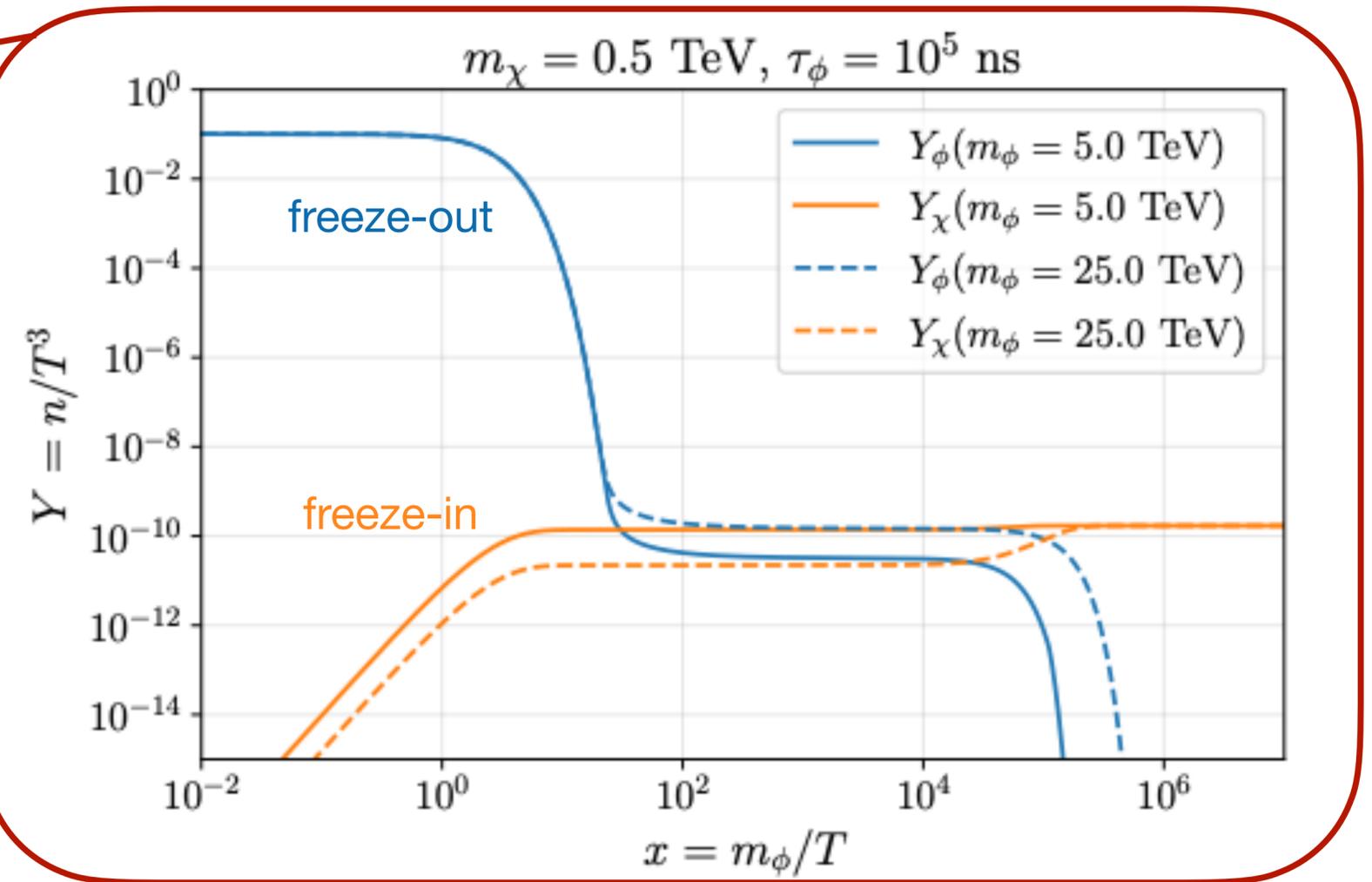
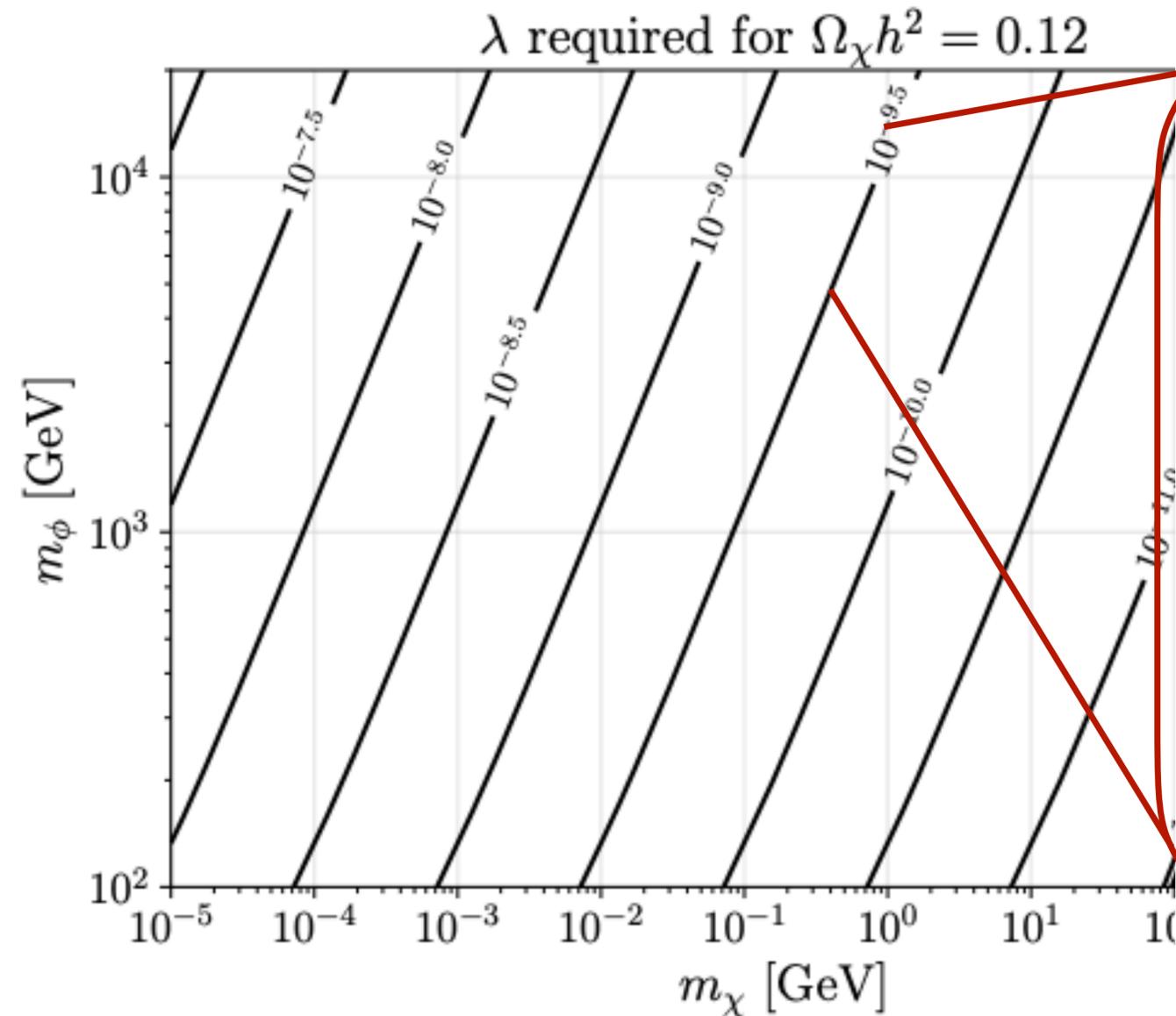


Dark Matter χ produced via freeze-in

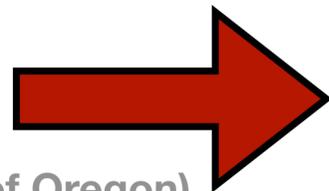
P. Asadi, A. Radick, TTY [arXiv:2312.03826]

Lepton-Flavored Dark Matter

relic abundance



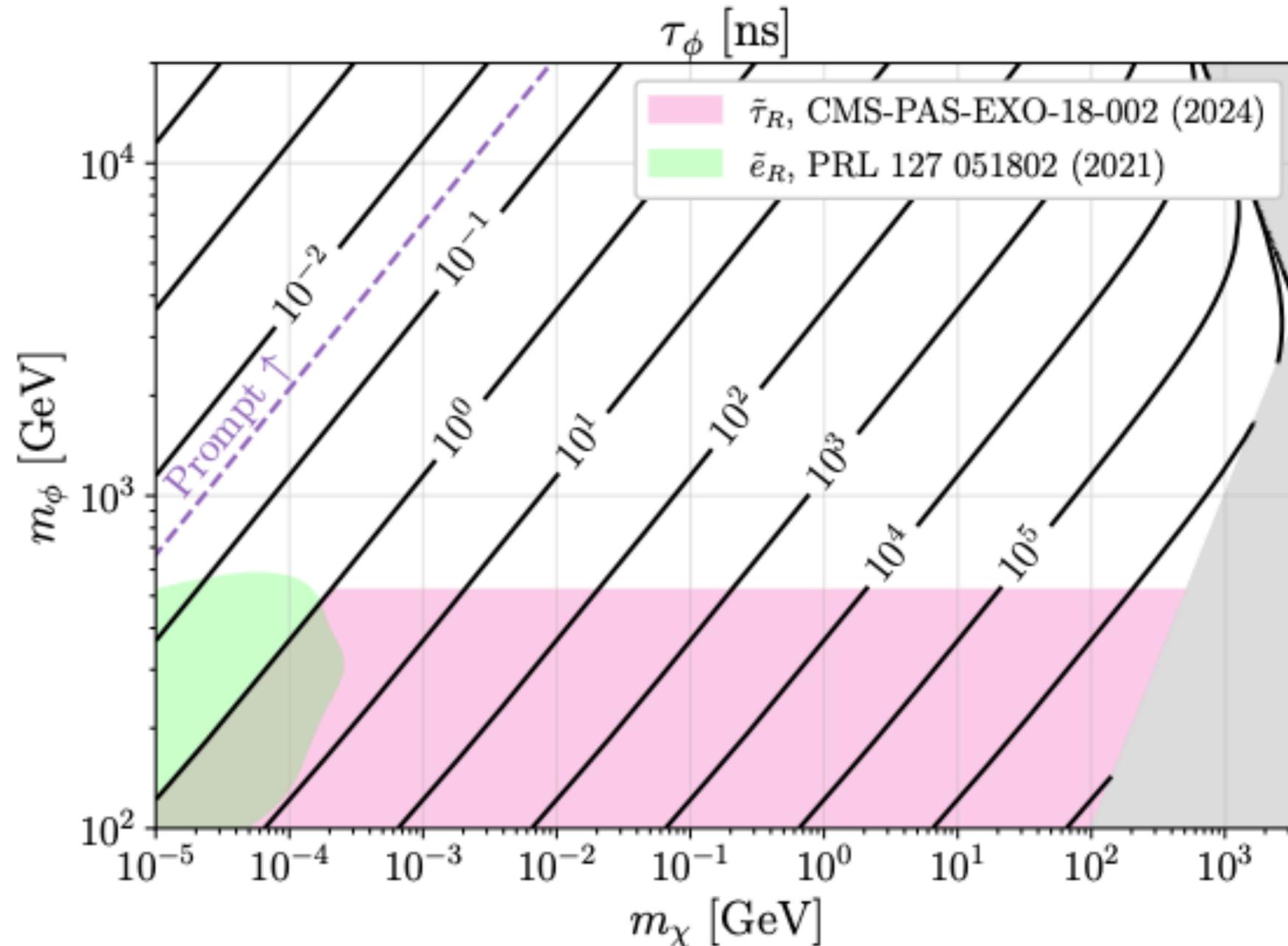
small λ



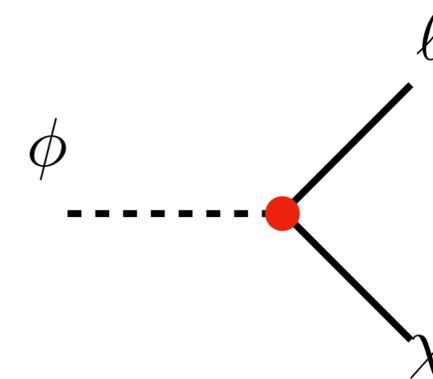
Dark Matter χ produced via freeze-in

Lepton-Flavored Dark Matter

scalar lifetime



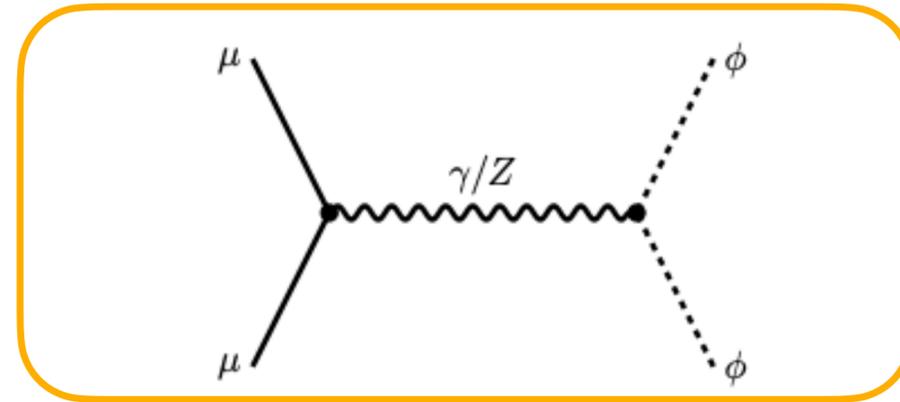
for much of the parameter space, the scalar mediator is long-lived!



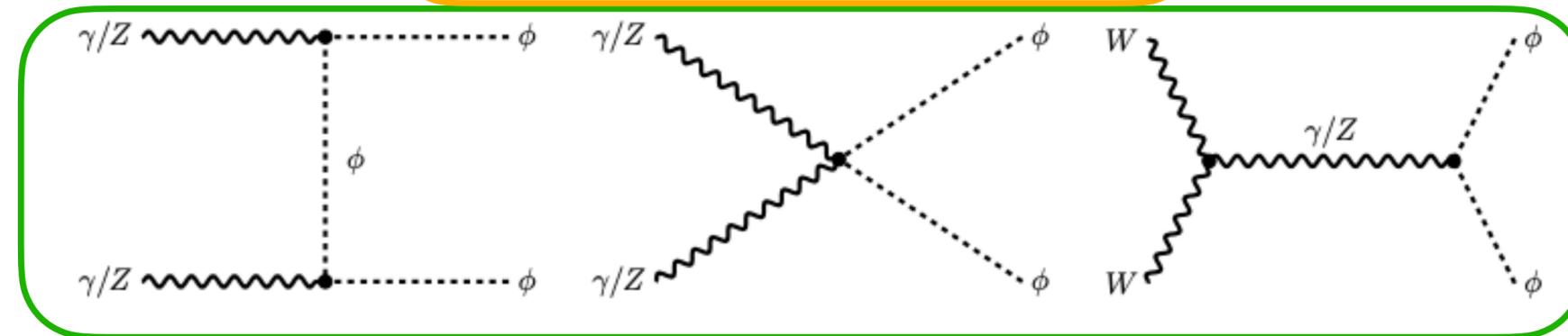
Lepton-Flavored Dark Matter

production channels at muon collider

Drell-Yan

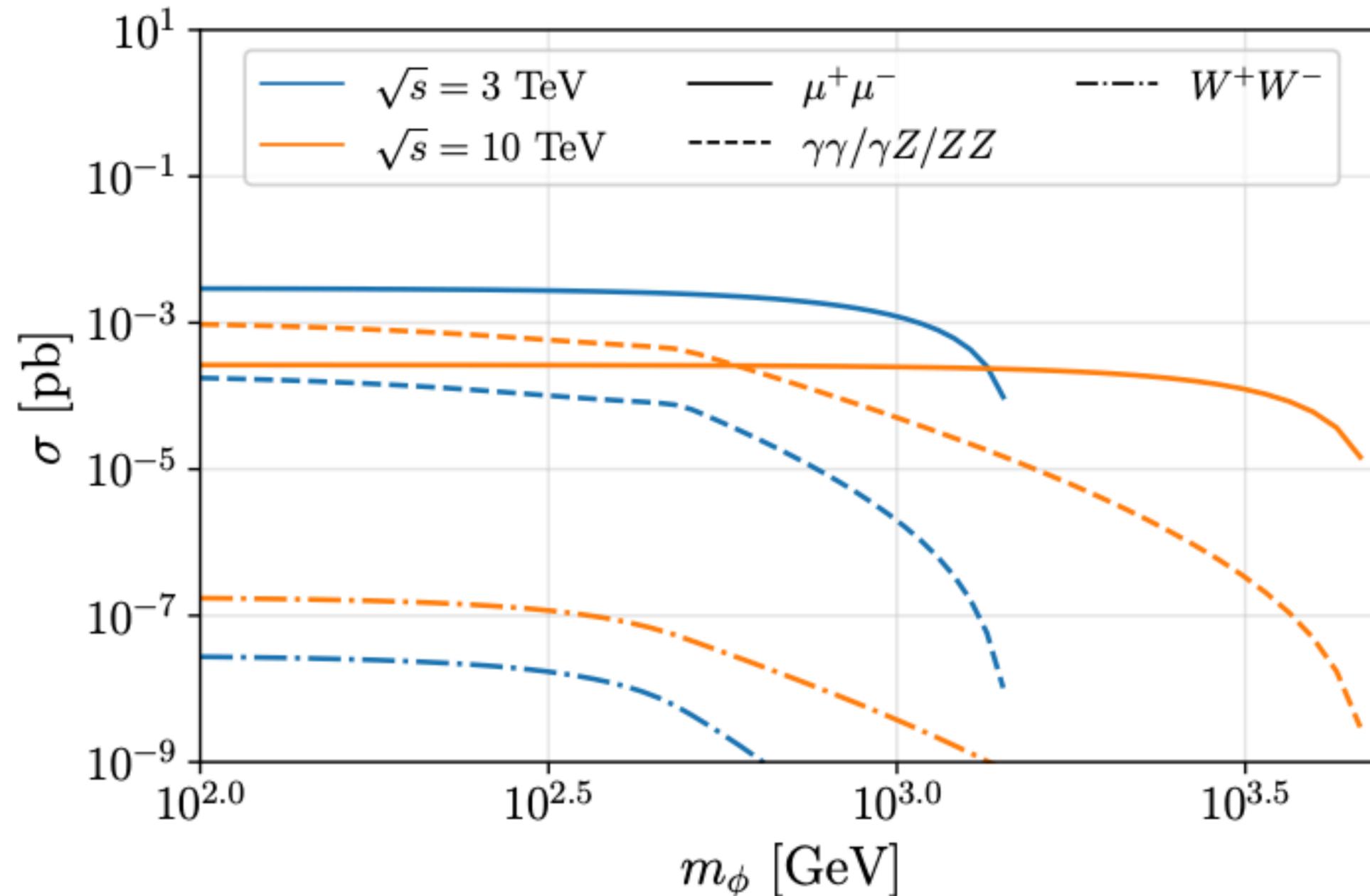


VBF



Lepton-Flavored Dark Matter

collider signatures



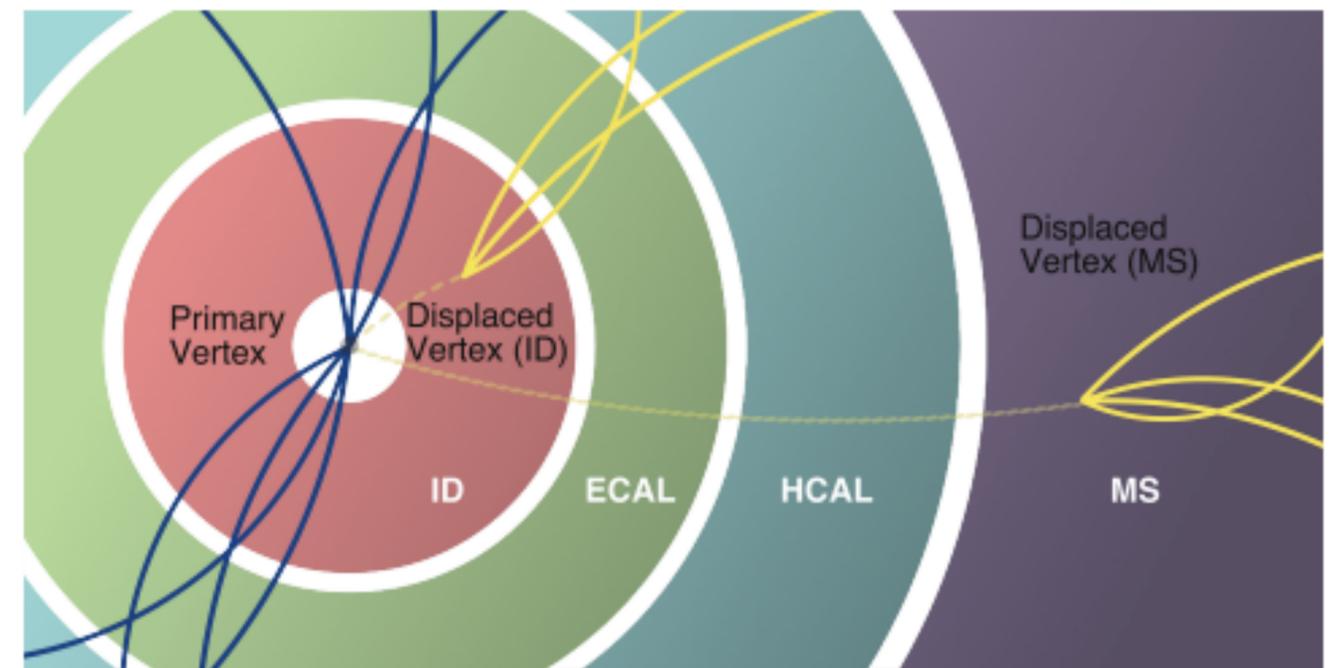
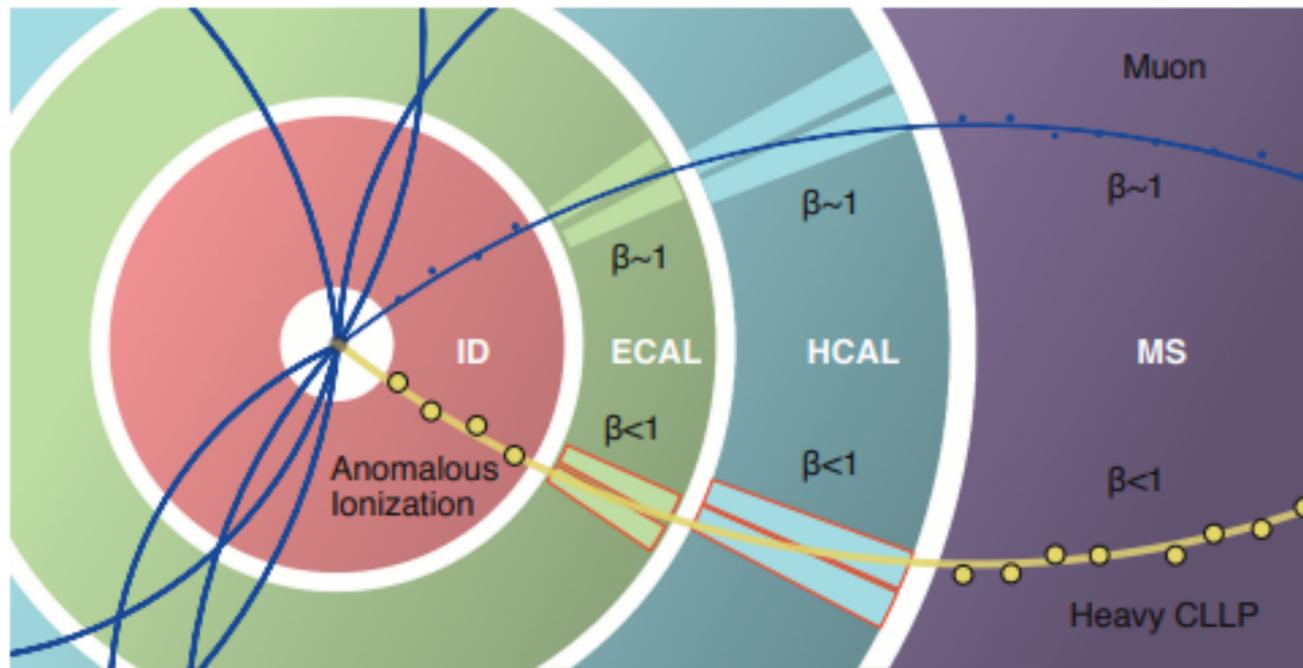
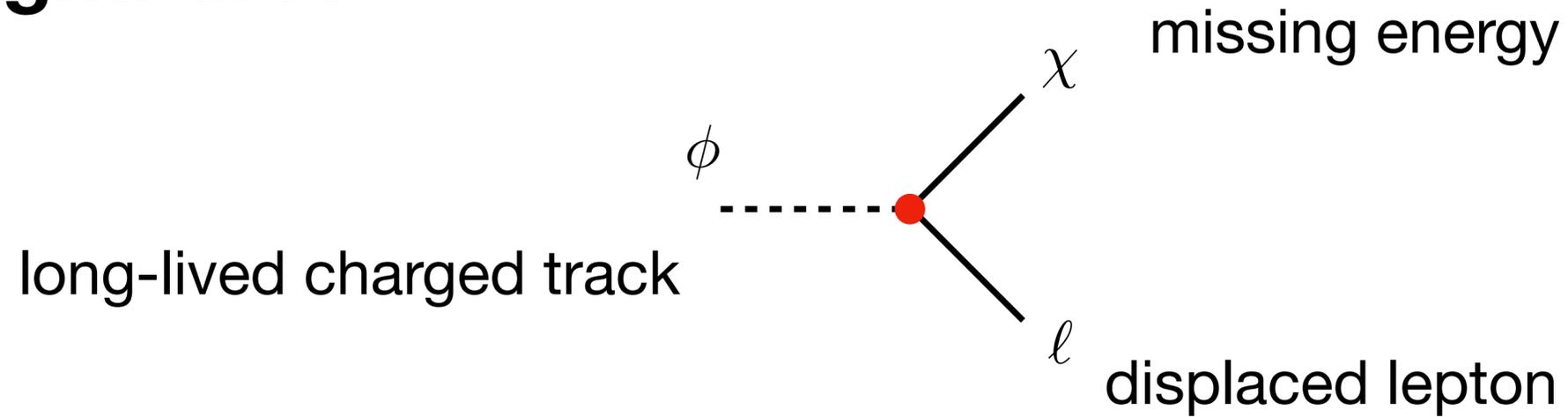
Drell-Yan dominates at higher mediator mass, VBF at lower masses

Combination can lead to interesting LLP signatures

use to help inform detector design

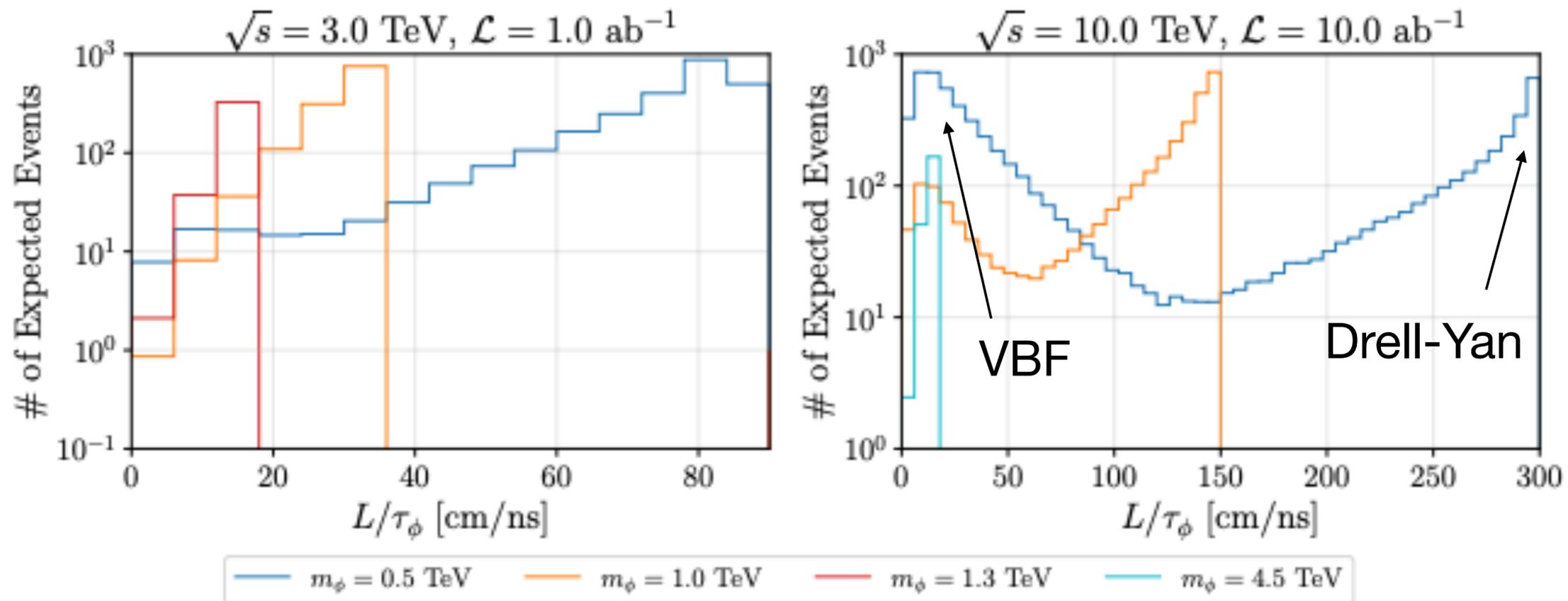
Lepton-Flavored Dark Matter

collider signatures



Lepton-Flavored Dark Matter

LLP collider signatures

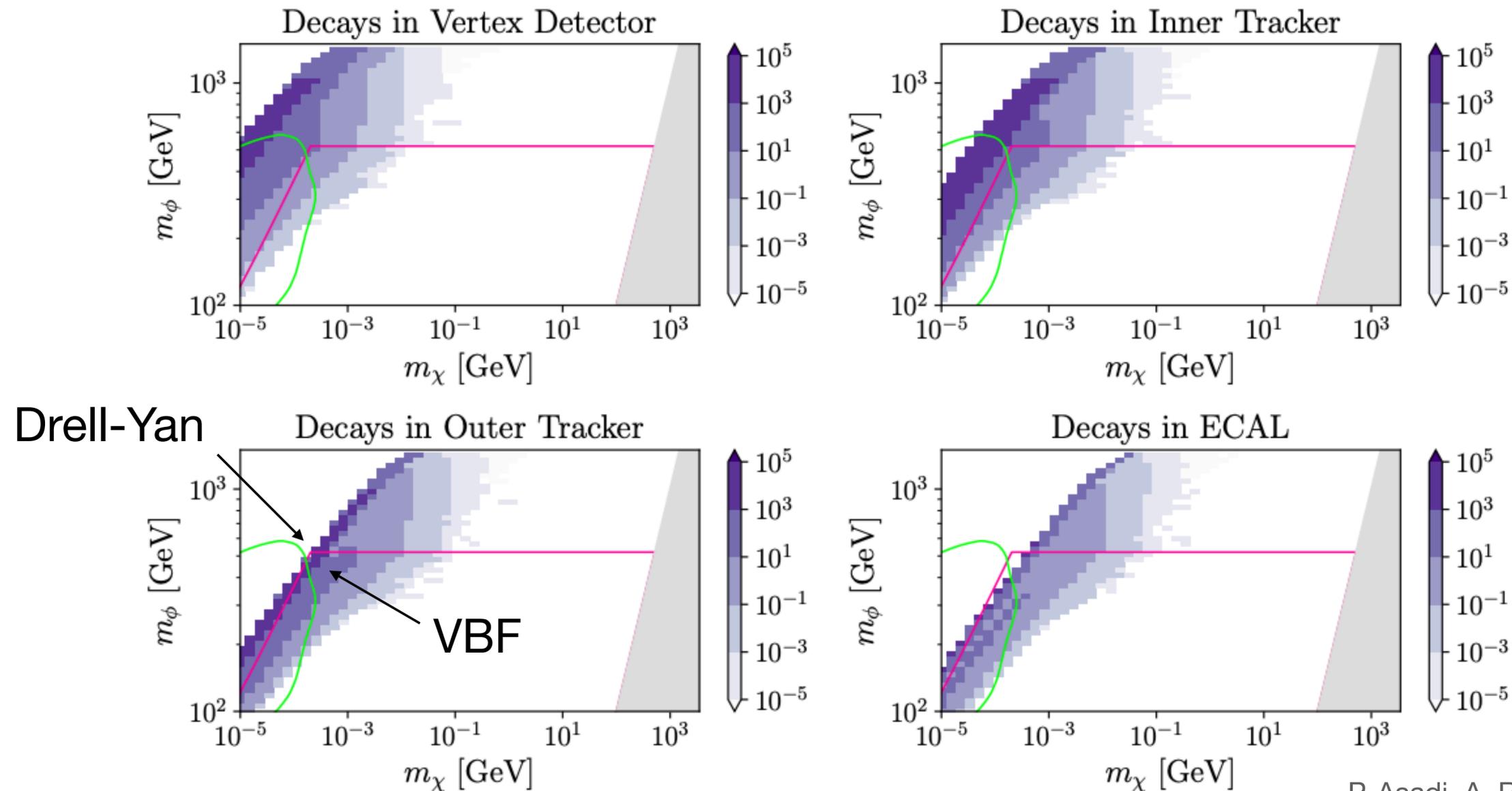


P. Asadi, A. Radick, TTY [arXiv:2312.03826]

Lepton-Flavored Dark Matter

LLP collider signatures

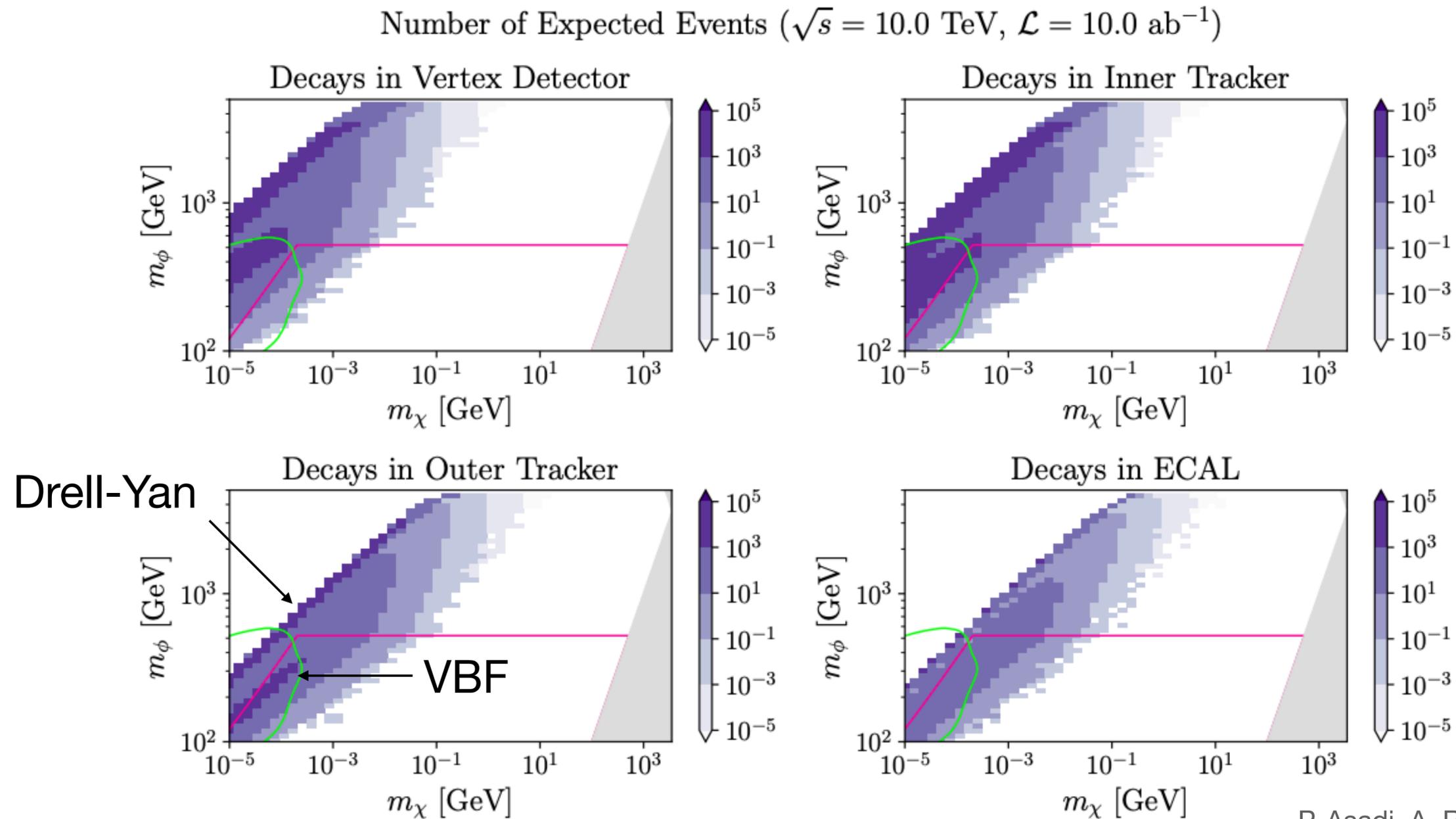
Number of Expected Events ($\sqrt{s} = 3.0 \text{ TeV}$, $\mathcal{L} = 1.0 \text{ ab}^{-1}$)



P. Asadi, A. Radick, TTY [arXiv:2312.03826]

Lepton-Flavored Dark Matter

LLP collider signatures



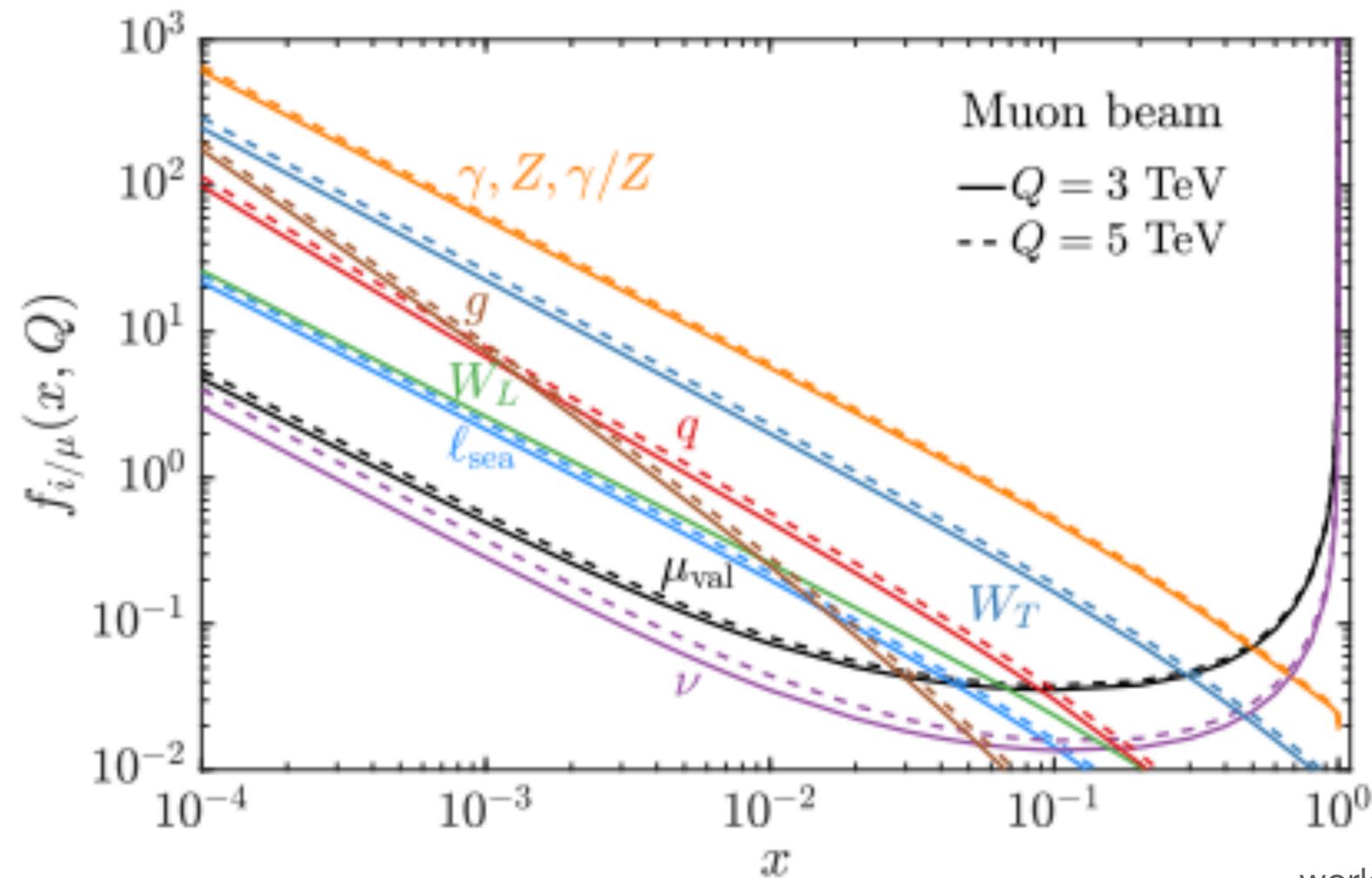
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Future directions

- implement muon PDFs into study

T. Han, Y. Ma, K. Xie, [arXiv:2103.09844]

see also LePDF: F. Garosi, D. Marzocca, S. Trifinopoulos [arXiv:2303.16964]



work in progress with P. Asadi, S. Homiller, A. Radick

Future directions

- implement muon PDFs into study
- expand to other models (quark flavored, LH lepton flavored, etc)

$$\mathcal{L}_{\text{int}}^{\text{RH quark}} \supset -\lambda_{i,\alpha} \phi \bar{u}_i \bar{\chi}_\alpha + h.c.$$

$$\mathcal{L}_{\text{int}}^{\text{LH lepton}} \supset -\lambda_{i,\alpha} \Phi^\dagger L_i \bar{\chi}_\alpha + h.c.$$

$$\mathcal{L}_{\text{int}}^{\text{LH quark}} \supset -\lambda_{i,\alpha} \Phi^\dagger Q_i \bar{\chi}_\alpha + h.c.$$

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- consider other mediator-DM combinations
- consider other “weird” signatures

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- consider other “weird” signatures
- but even before we get to the muon collider...
 - staging can be used as muon beam dump, e.g.

C. Cesarotti, S. Homiller, R. Mishra, M. Reece [arXiv:2202.12302]
expanded upon in C. Cesarotti, R. Gambhir [arXiv:2310.16610]

