

DUNE Systematic Flux Uncertainties

Ian D. Kotler FRAS on behalf of the DUNE Collaboration

[APS DPH-PHENO 2024](#)

May 14th, 2024

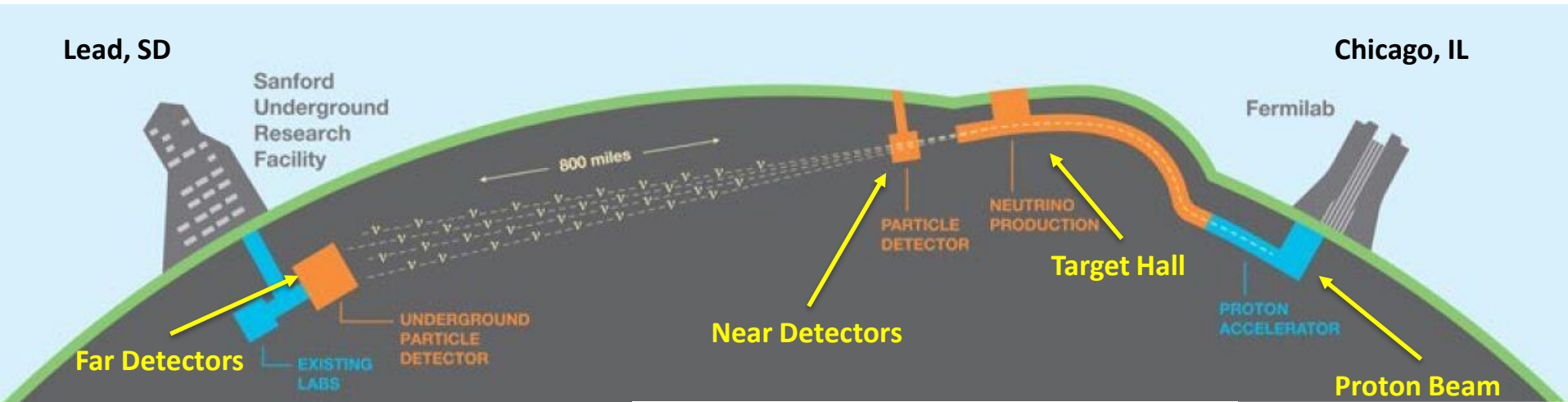
Content

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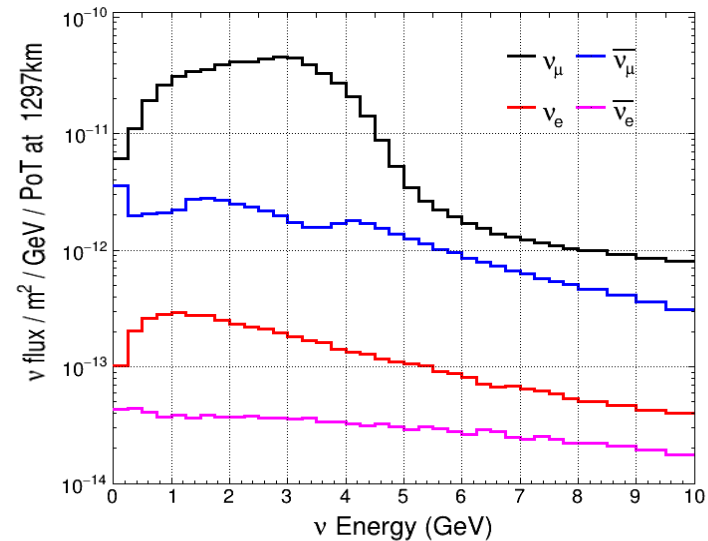
What is DUNE?

- The **D**eep **U**nderground **N**eutrino **E**xperiment **hosted at** Fermilab
- **Comprised of** 1400+ collaborators **across** 35+ countries.
- **Physics goals include** (but not limited to):
 - **Address Baryon Asymmetry of the Universe (BAU)**
 - ◆ Measure δ_{CP} in lepton sector.
 - **Determine the neutrino mass ordering**
 - ◆ Sign of $|\Delta m_{32}^2|$?
 - **Determine the octant of θ_{23} .**
 - ◆ Is θ_{23} greater or less than $\frac{\pi}{4}$?
 - **Near Detector Complex hosts a suite of rich physics programs .**
 - ◆ Suite of detectors {LAR, GAR, SAND, TMS, **PRISM** ...}
 - **And so much more!**
 - ◆ Interested in joining **DUNE**? Get started [here](#).

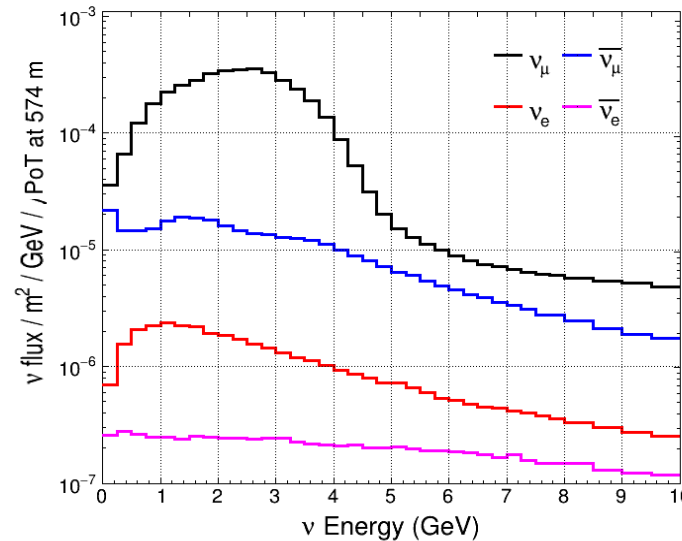
How does DUNE work?



ν Flux Prediction at Far Det.

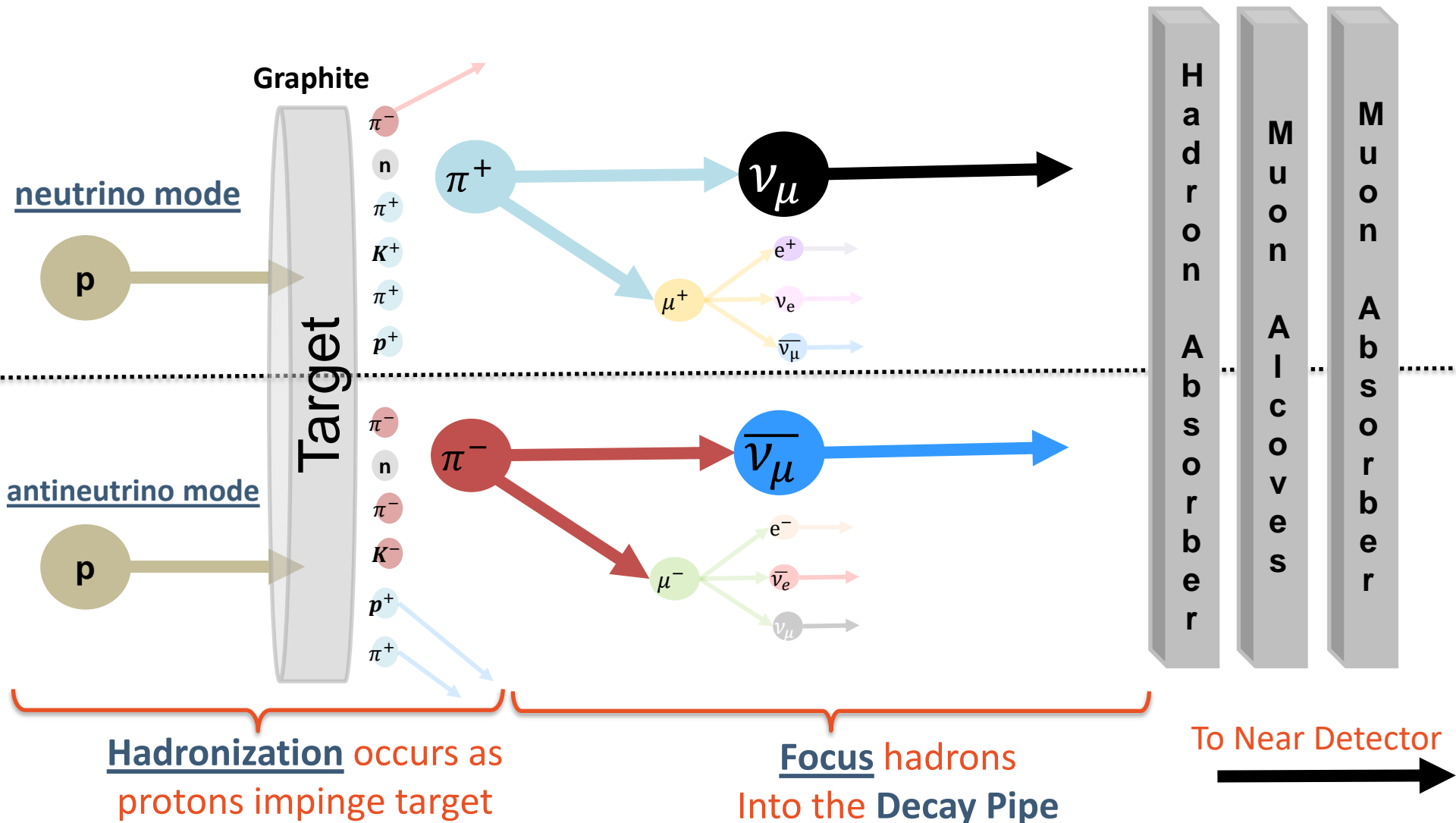


ν Flux Prediction at Near Det.



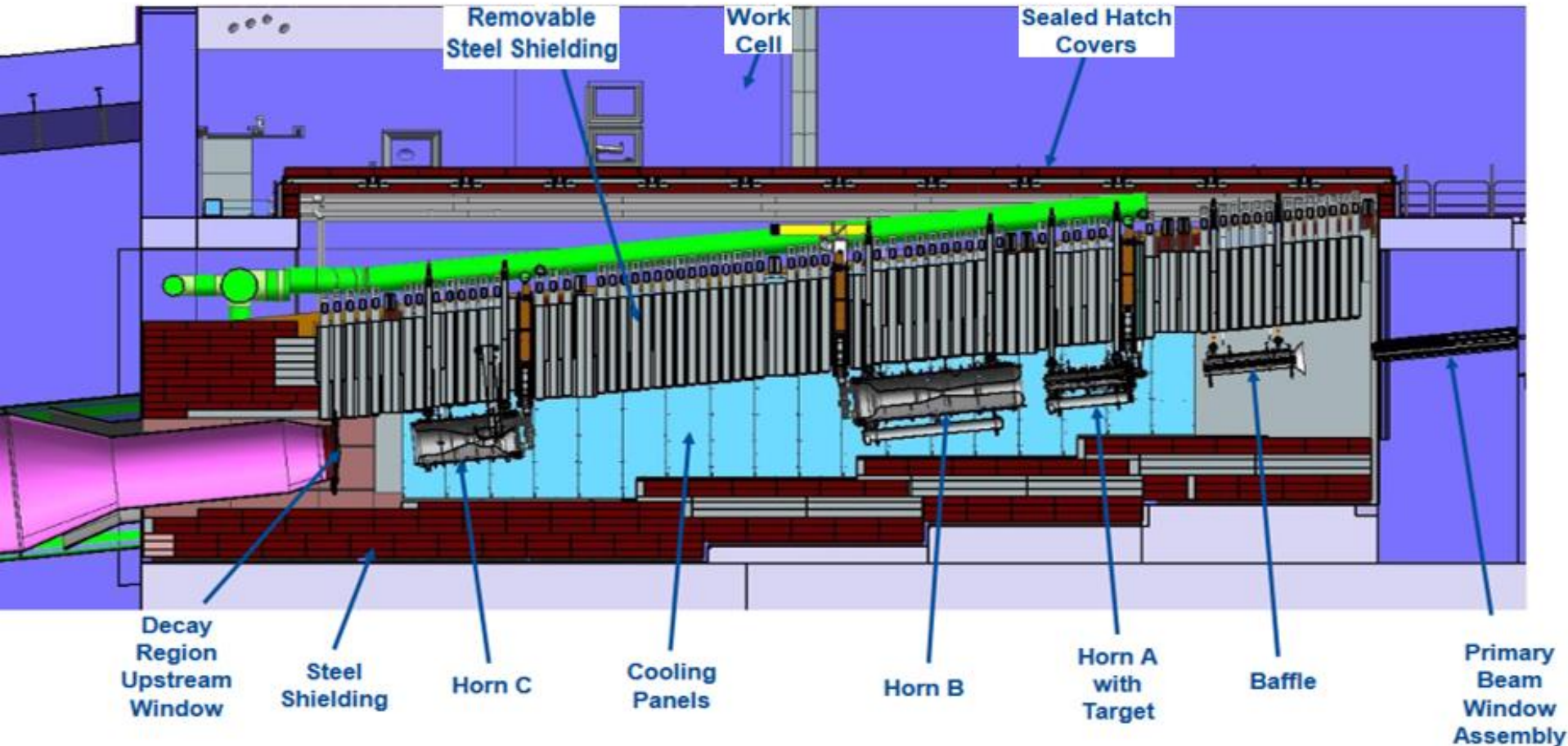
- Protons impinge target.
- Secondary beam of pions is created and focused.
- Pions decay to neutrinos and muons.
- Near Det. samples the neutrino beam.
- Neutrinos oscillate and experience matter effect.

Hadron Production for DUNE



Focusing Effects

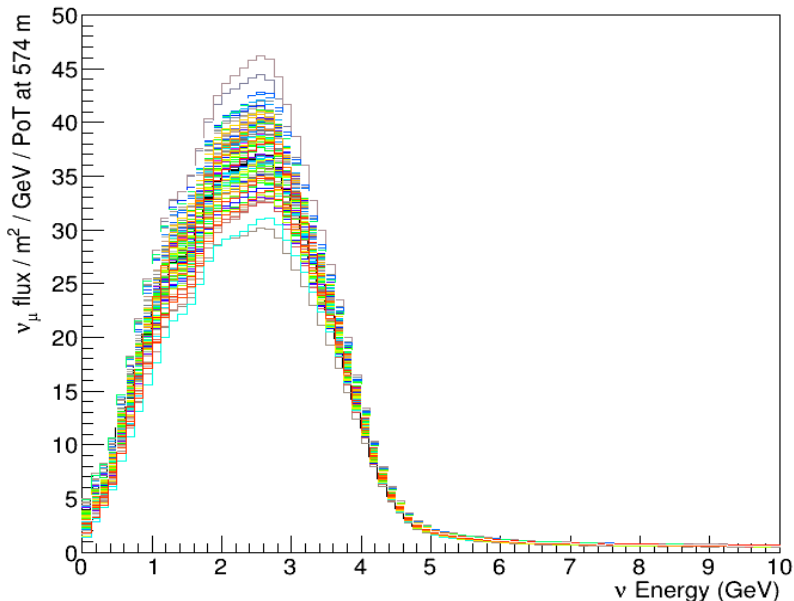
- 3 Horn Focusing System
- 2 modes of operation
 - ◆ Current: $\pm 300\text{kA}$
- 60+ sources of beam focusing uncertainties
- Target Hall components contribute to Hadronization



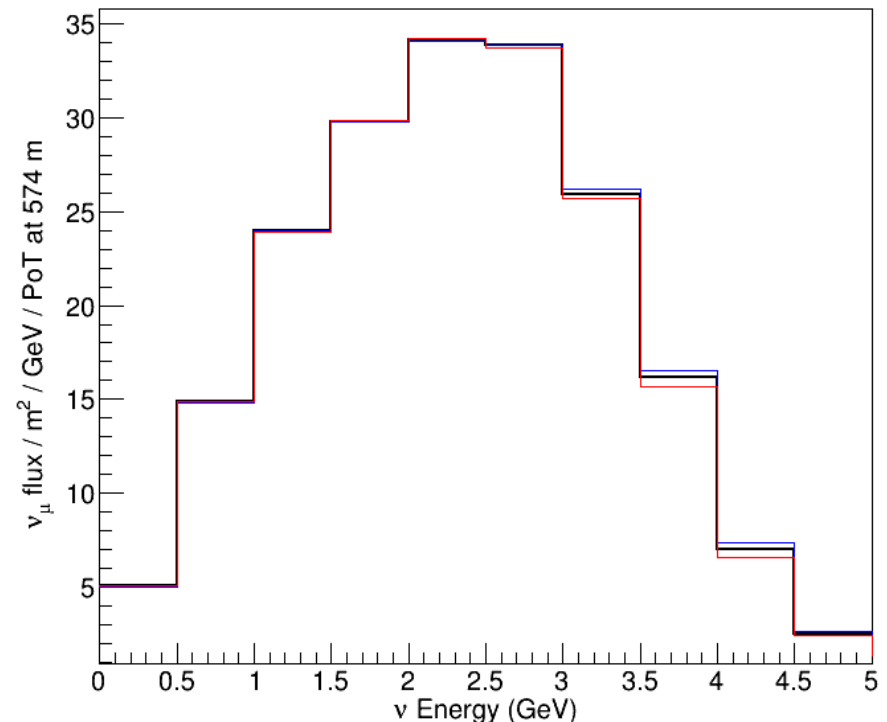
Modeling the DUNE Flux

- Nominal flux is input into PPFX.
- Varies the flux parameters across 100 universes.
- Specialized reweighters and external inputs account for Hadron Production processes.
- Nominal flux is generated in g4lbne
- Varies nominal by engineering tolerance.
 - Results in 2 universes, $\pm 1\sigma$.

PPFX Multi-Universe FHC ν_μ Flux



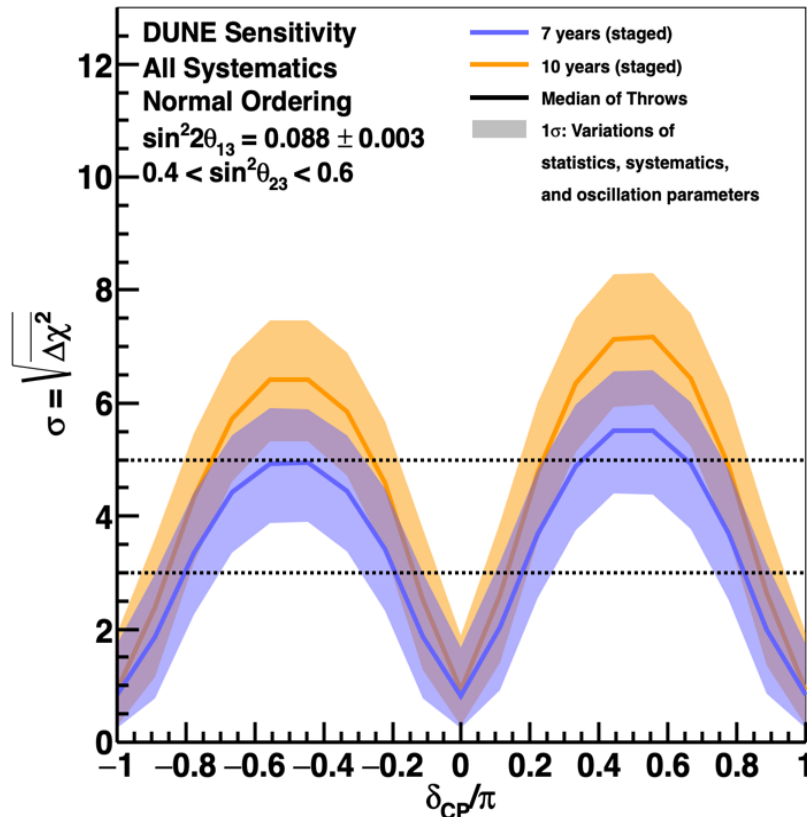
BFU Multi-Universe FHC ν_μ Flux



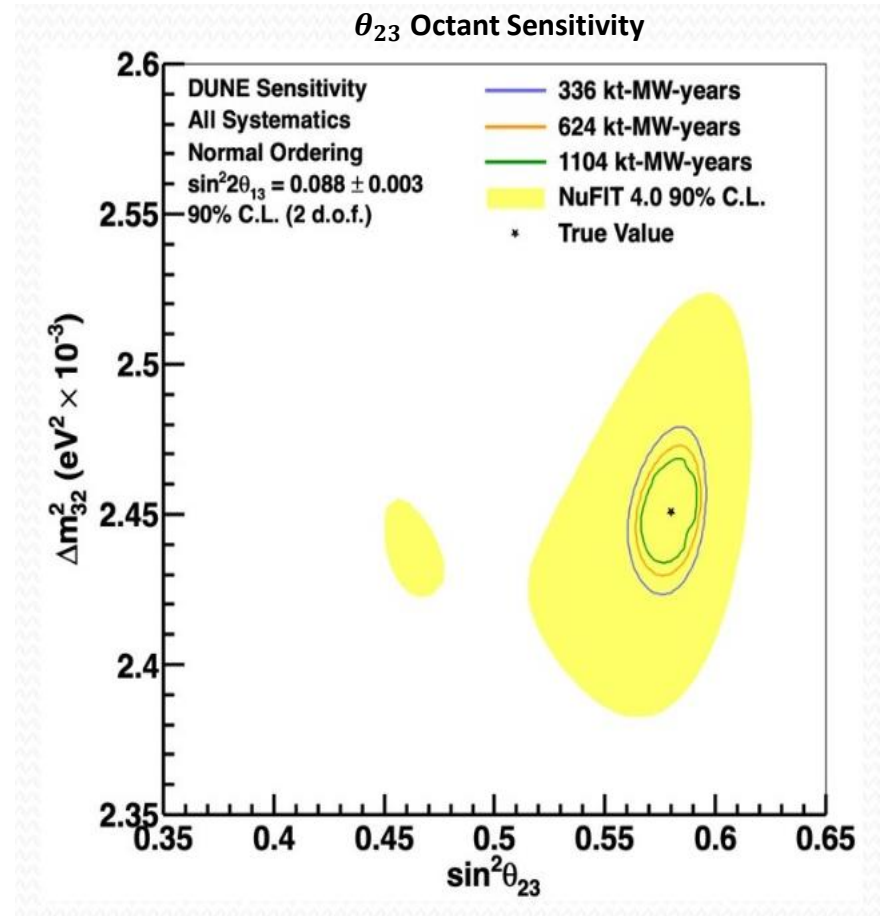
The Importance of Systematics

- See upcoming publication, “The DUNE Neutrino Flux Simulation” details on covariance.

CP Violation Sensitivity



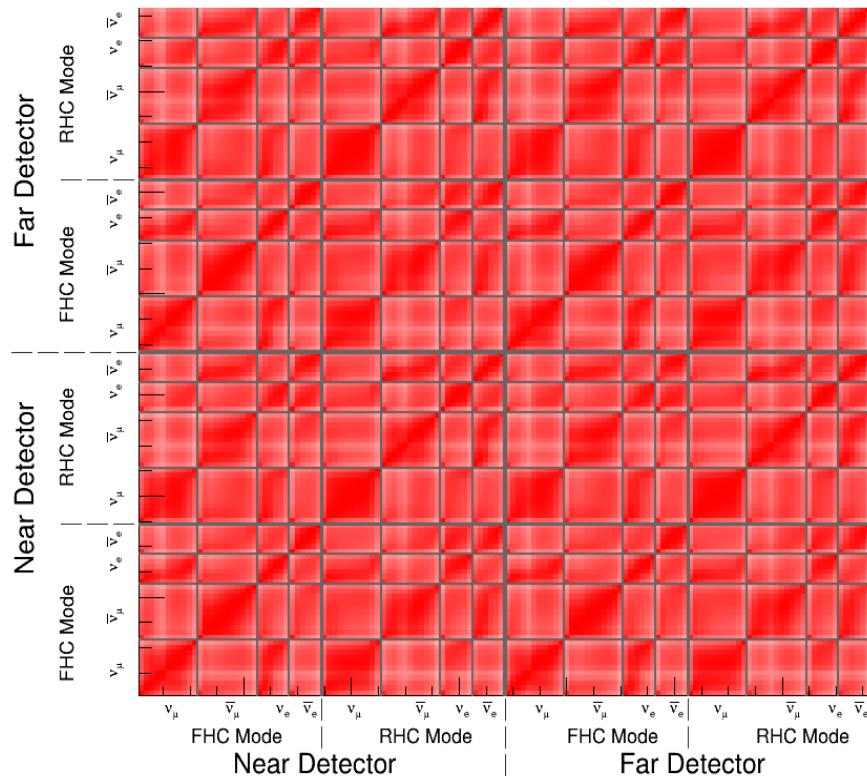
θ_{23} Octant Sensitivity



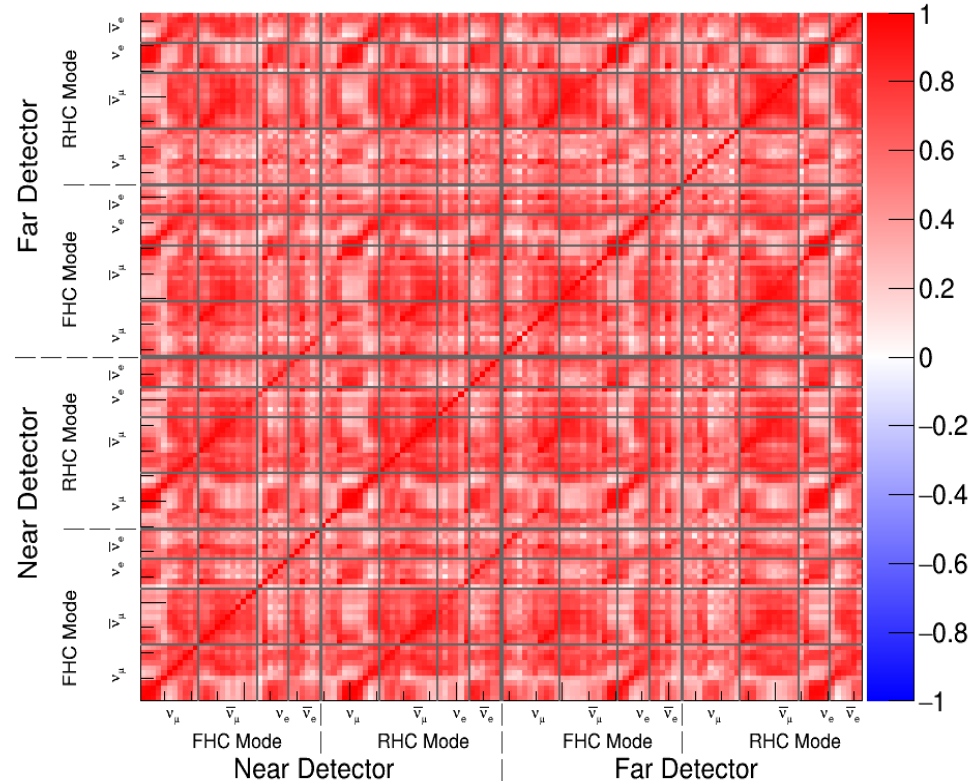
Determining the Correlations

- The **Correlation Matrices** reveal the magnitude of the relations amongst the various sources of uncertainty across all modes, detector locations and neutrino species.

Hadron Production Correlations

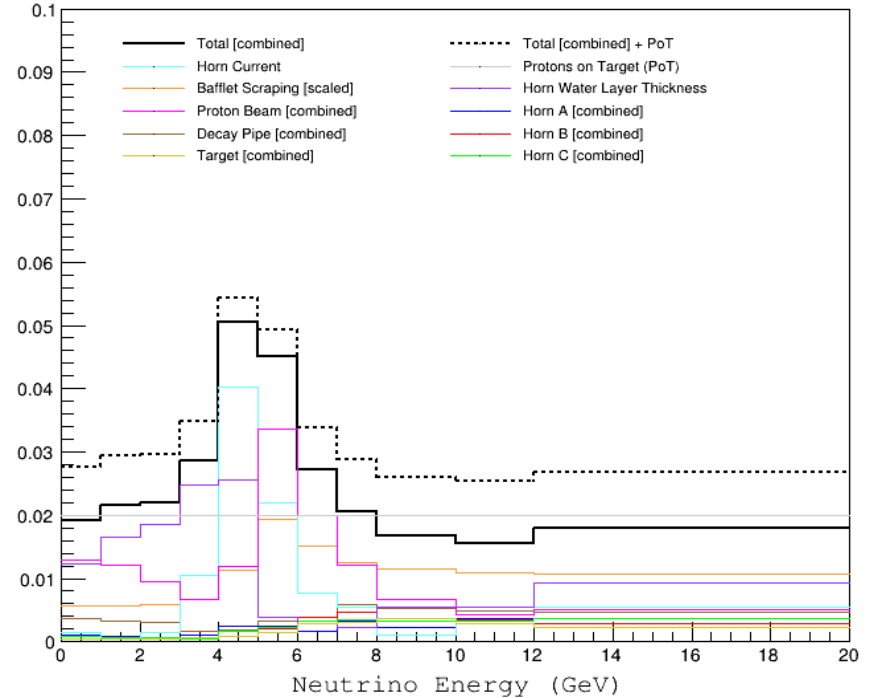
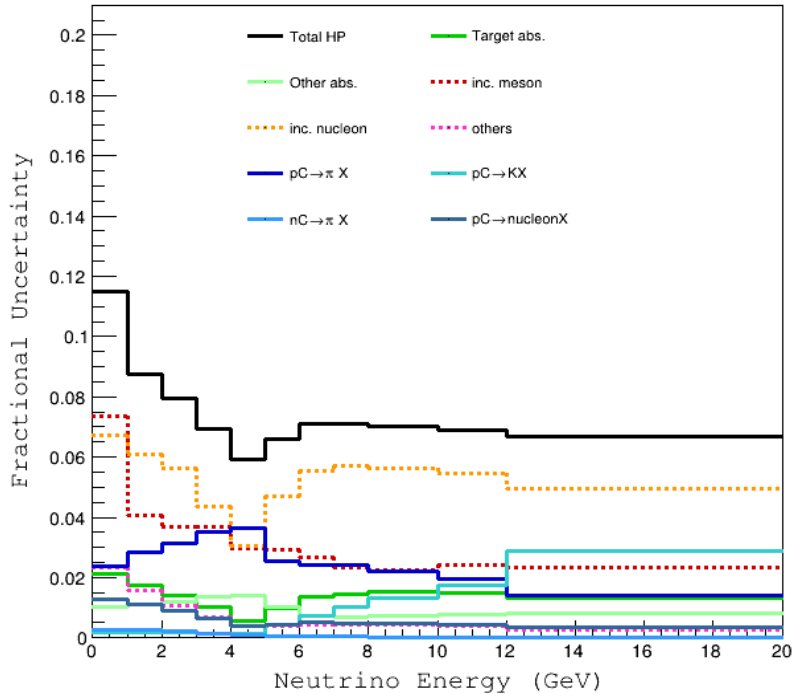


Beam Focusing Correlations



Individual Uncertainties

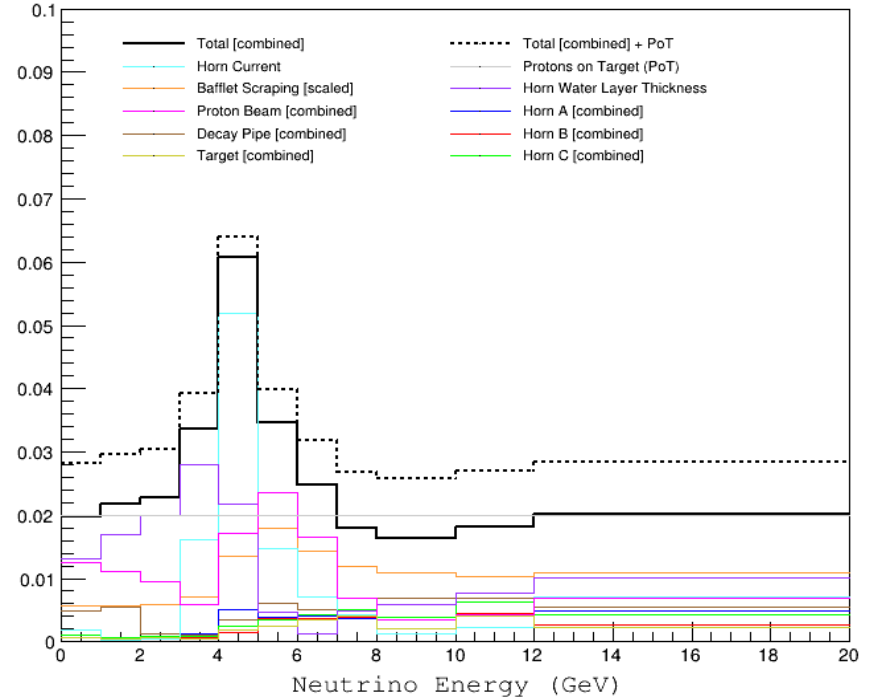
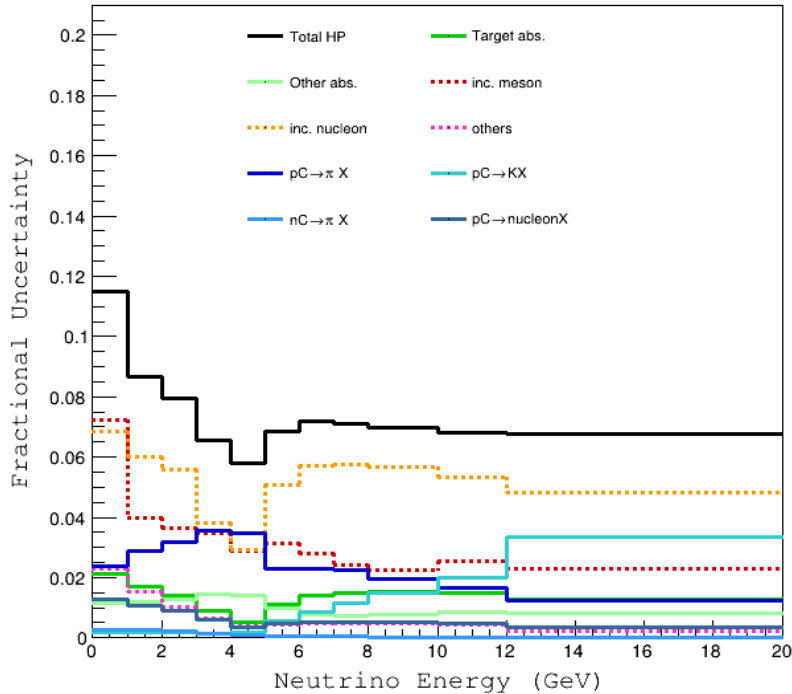
- Taking the square root of the diagonals of each matrix yields the individual uncertainties.



Mode	Location	Species	Mode	Location	Species
neutrino	Far Det.	ν_μ	neutrino	Far Det.	ν_μ
Hadron Production			Beam Focusing		

Individual Uncertainties

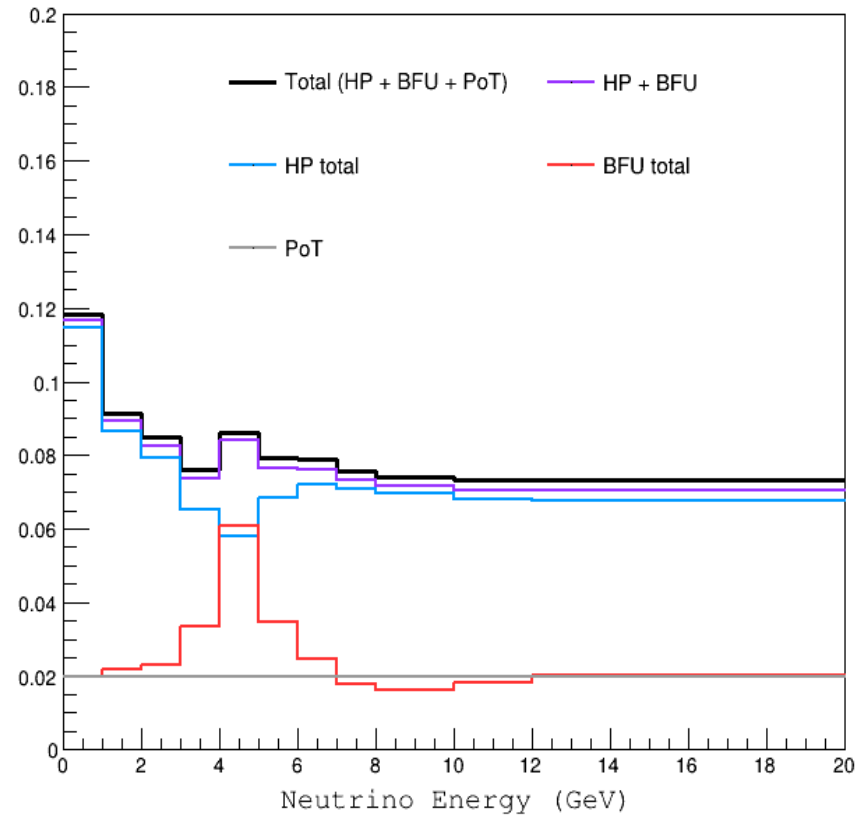
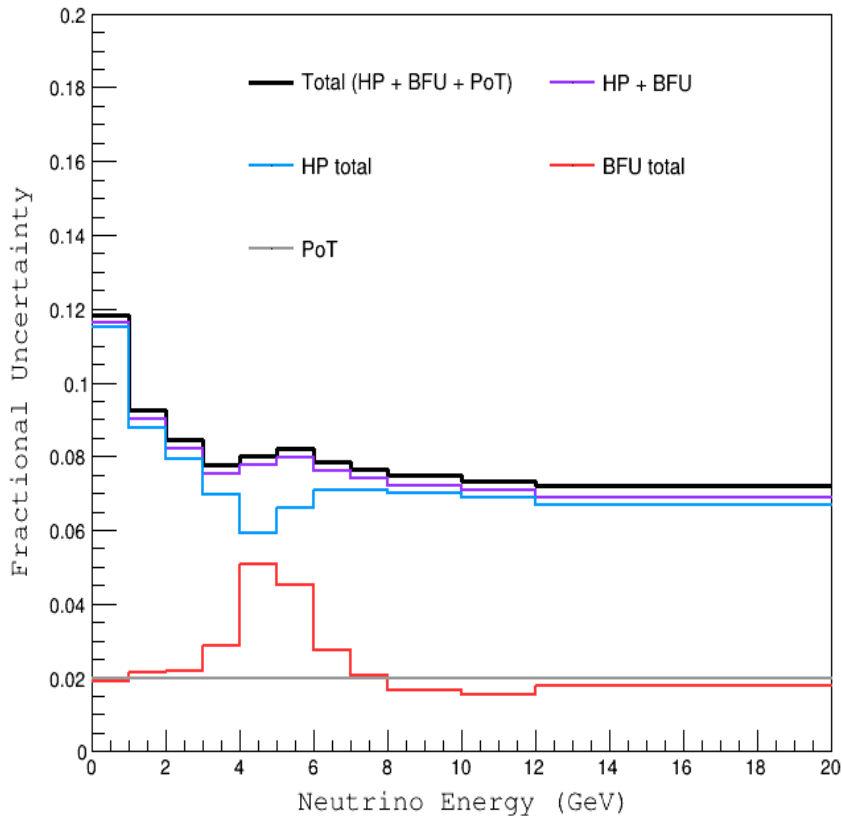
- Taking the square root of the diagonals of each matrix yields the individual uncertainties.



Mode	Location	Species	Mode	Location	Species
neutrino	Near Det.	ν_μ	neutrino	Near Det.	ν_μ
Hadron Production			Beam Focusing		

Total Systematic Uncertainties

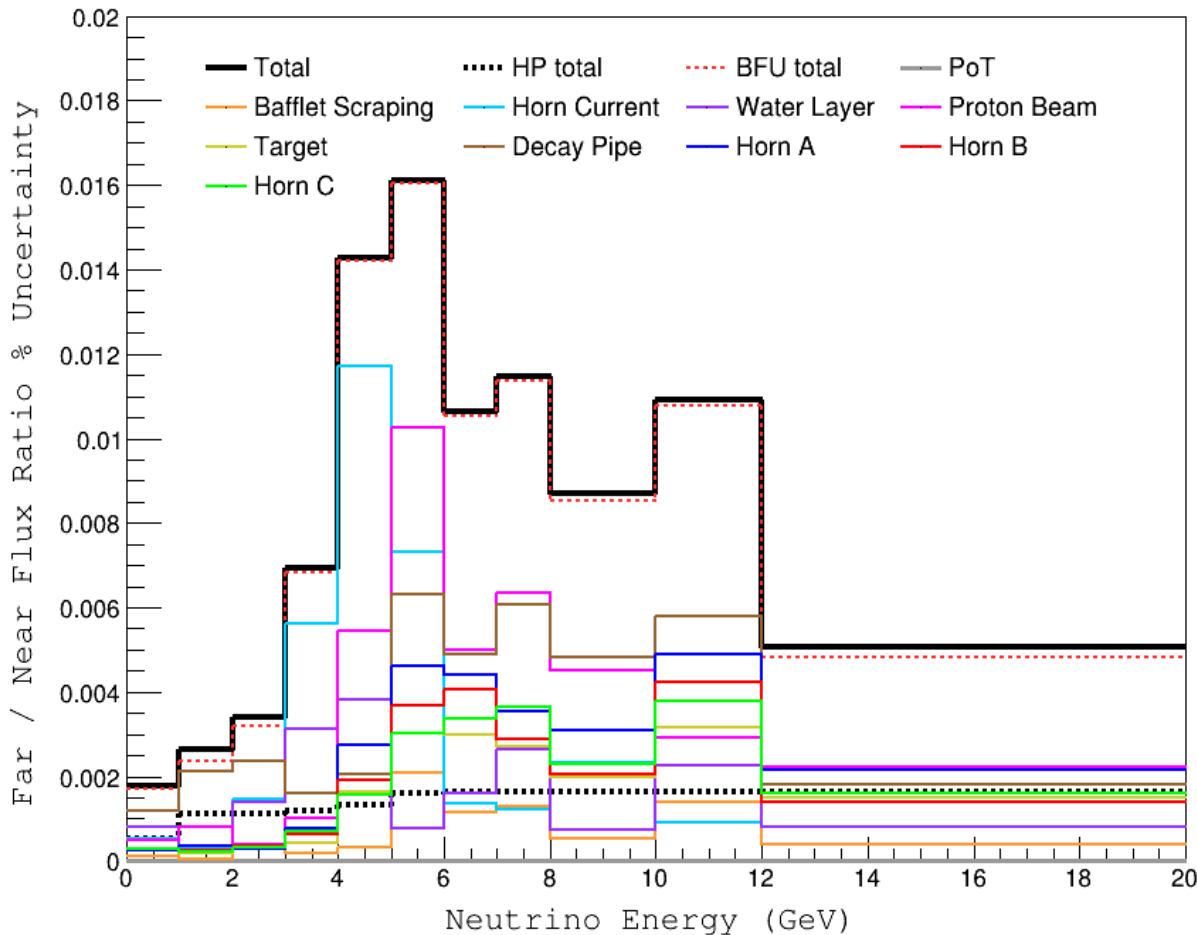
- Adding Hadron Production and Beam Focusing Covariances to obtain Total Beam Covariance.



Mode	Location	Species	Mode	Location	Species
neutrino	Far Det.	ν_μ	neutrino	Near Det.	ν_μ

The Far to Near Flux Ratio

FHC ν_μ Far / Near Ratio % Uncertainties



- **Beam Focusing dominates.**
- **Hadron Production now < 0.2%**
- **Previous max of 6% at 4.2 GeV ~ 1.5%**
- **Flux Peak Unc. at 2.5 GeV ~ 0.4%**

Conclusions

- **DUNE** is an accelerator-based neutrino experiment hosted at Fermilab
- Among **DUNE's** many goals includes determining:
 - ◆ δ_{CP}
 - ◆ neutrino mass hierarchy
 - ◆ octant of θ_{23}
- To achieve the high sensitivity required to measure parameters requires covariance matrices for all Systematic Uncertainties
- The covariance matrix encapsulates the all information regarding uncertainties and correlations.
- Hadron Production and Beam Focusing are the largest contributors to beam systematics uncertainties.

Back Up Slides

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(2:30 - 2:45) pm

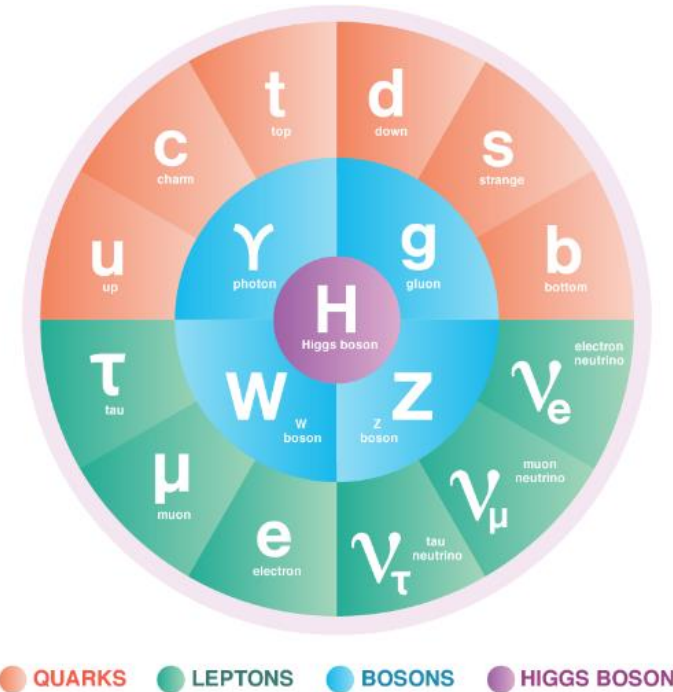
David Lawrence Hall, 107

University of Pittsburgh



What are neutrinos?

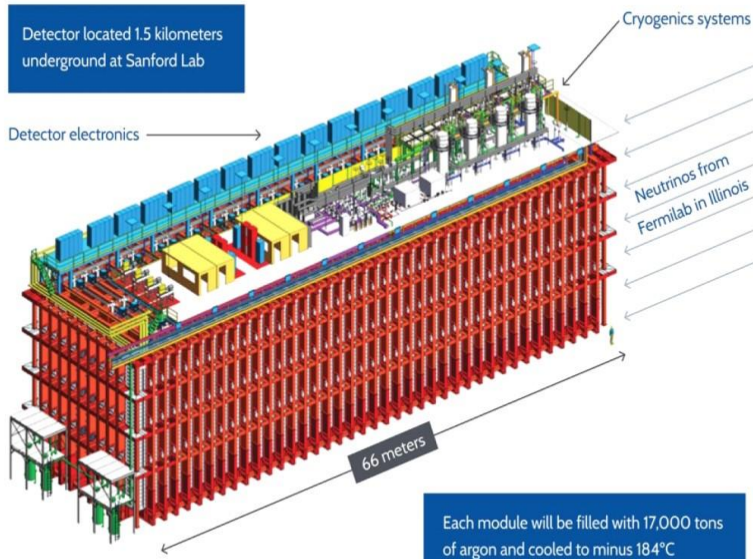
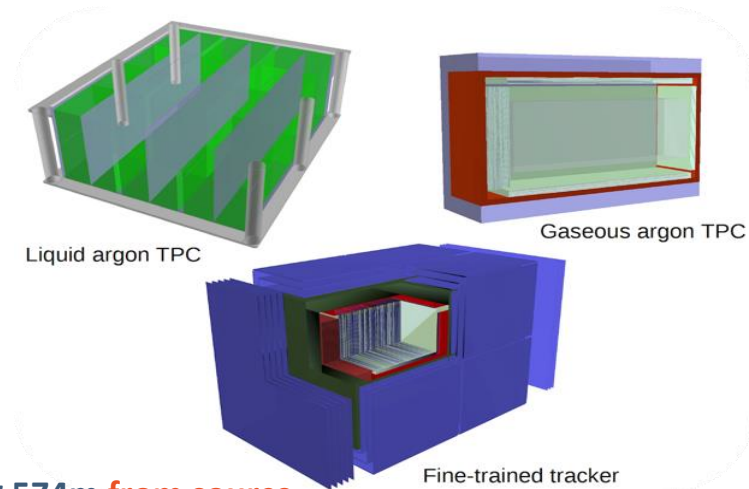
- Fundamental particles of the SM.
- Colorless, neutral leptons
- 3 distinct flavors: ν_e, ν_μ, ν_τ
- 3 distinct masses: ν_1, ν_2, ν_3
- Can oscillate between flavors, governed by the PMNS matrix.



$$|U| = \begin{bmatrix} |U_{e1}| & |U_{e2}| & |U_{e3}| \\ |U_{\mu1}| & |U_{\mu2}| & |U_{\mu3}| \\ |U_{\tau1}| & |U_{\tau2}| & |U_{\tau3}| \end{bmatrix} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{CP}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{CP}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{CP}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{CP}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{CP}} & c_{23}c_{13} \end{bmatrix}$$

DUNE Detection Systems

- 2 Detector Design (Near and Far Detectors)
- Far Det. ~ 1297km from source located at SURF
- Comprised of 4, 17kt LAr TPC's
- Cooled to ~ 90 Kelvin
- Fiducial Volume of 10kt per TPC
- Primary Goal: Measure Oscillated Flux



- Near Det. ~ 574m from source
- Comprised of 3 subdetectors:
 - LAr TPC (50-ton)
 - GAr TPC (1-ton)
 - SAND (8-ton plastic scintillator)
- LAr TPC measures unoscillated flux
- GAr TPC monitors the muon flux
- SAND (Fine-trained tracker) measures:
 - On-Axis beam flux, possible Neutron detection.
- Argon TPC's can move off axis (PRISM)
 - Deconvolves flux from ν cross-section on Argon.
- Primary Goals:
 - Measure Unoscillated Flux /
 - Characterize the Beam

What is Hadron Production?

- Largest source of systematic uncertainty for DUNE flux prediction.
- Sources of Hadron Production in DUNE include:
 - Protons impinging on Graphite target: $p + {}^{12}\text{C} \rightarrow \pi^{\pm} + X$
 $p + {}^{12}\text{C} \rightarrow K^{\pm} + X$
 $p + {}^{12}\text{C} \rightarrow p(n) + X$
 - Secondary Interactions of neutrons: $n + {}^{12}\text{C} \rightarrow \pi^{\pm} + X$
 - Hadron Absorption both inside and outside the target.
 - Secondary meson and nucleon interactions
 - And many others!
- Simulating these Hadron Production uncertainties requires:
 - Input data from dedicated experiments [NA49, SHINE, NA61*]
 - Package to Predict the Flux (PPFX), developed originally for Minerva by Leonidas Aliaga Soplín of U. Houston.

Beam Focusing Effects

- 2nd largest source of systematic uncertainty in DUNE flux prediction.
- Over 60 sources, all arising from engineering tolerances, such as:
 - Horn Current ($\pm 300\text{kA}$)
 - Thickness of Water Layer cooling Horns.
 - Scraping of proton beam against the Baffle.
 - Various characteristics of:
 - ◆ Proton Beam characteristics (Radius, Position, Angle, ...)
 - ◆ Target characteristics (Density, Position, Length, ...)
 - ◆ Horns A,B,C characteristics (Position, Ellipticity, Tilt, ...)
 - ◆ Decay Pipe characteristics (Radius, Position, Cross-Section, ...)

Calculating BFU Covariance

- Calculate individual covariances for each source of uncertainty (i) in both universes.

$$\text{Cov}_{\text{BFU},+}^{(i)}(x_j, x_k) = \frac{(x_j^{(i)} - \bar{x}_j)(x_k^{(i)} - \bar{x}_k)}{\bar{x}_j \bar{x}_k}$$

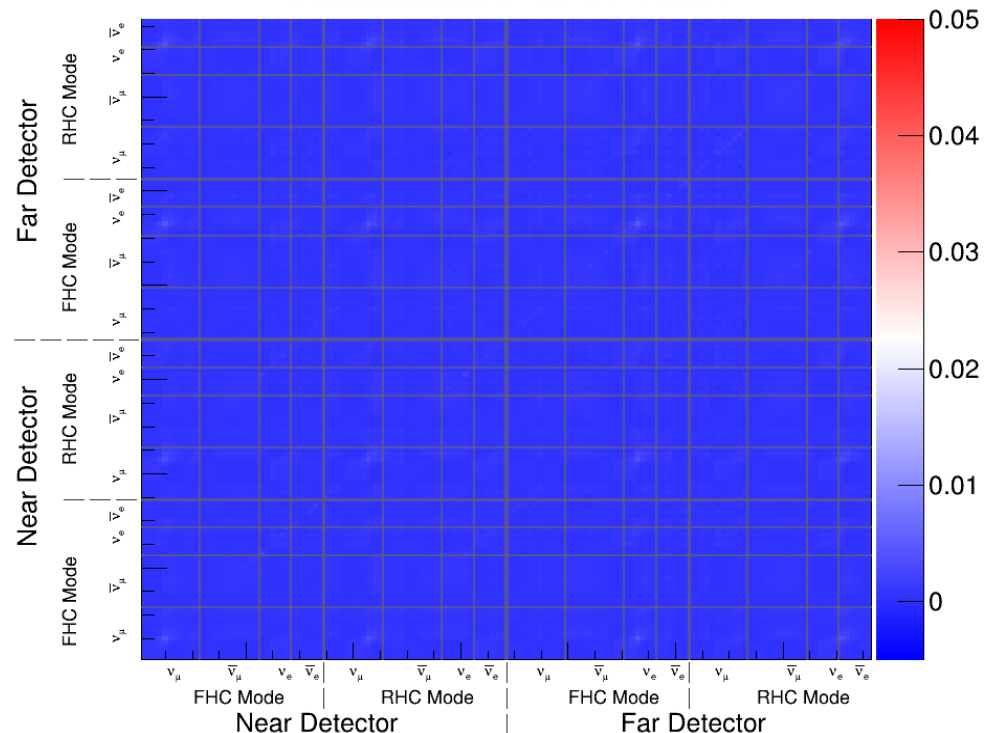
$$\text{Cov}_{\text{BFU},-}^{(i)}(y_j, y_k) = \frac{(y_j^{(i)} - \bar{y}_j)(y_k^{(i)} - \bar{y}_k)}{\bar{y}_j \bar{y}_k}$$

- Total BFU Covariance is average of universe covariances.

$$\left\langle \text{Cov}_{\text{BFU}}^{(i)}(x_j, x_k, y_j, y_k) \right\rangle = \frac{1}{2} \left[\text{Cov}_{\text{BFU},+}^{(i)}(x_j, x_k) + \text{Cov}_{\text{BFU},-}^{(i)}(y_j, y_k) \right]$$

- Here we see the BFU Covariance is quite small indicating the magnitudes of the focusing uncertainties are likewise, small.

Total Beam Focusing Covariance



Calculating HP Covariance

- Calculate individual covariances for each source of uncertainty (i) in both universes.

$$\left\langle \text{COV}_{\text{HP}}^{(i)}(z_j, z_k) \right\rangle = \frac{1}{\mathcal{N}} \sum_{u=1}^{\mathcal{N}} \text{COV}_{\text{HP}}^{(i,u)}(z_j, z_k) \quad \therefore \begin{cases} \mathcal{N} = 100 & 100 \text{ universes} \\ u = [1, \mathcal{N}] & \text{universe \#} \end{cases}$$

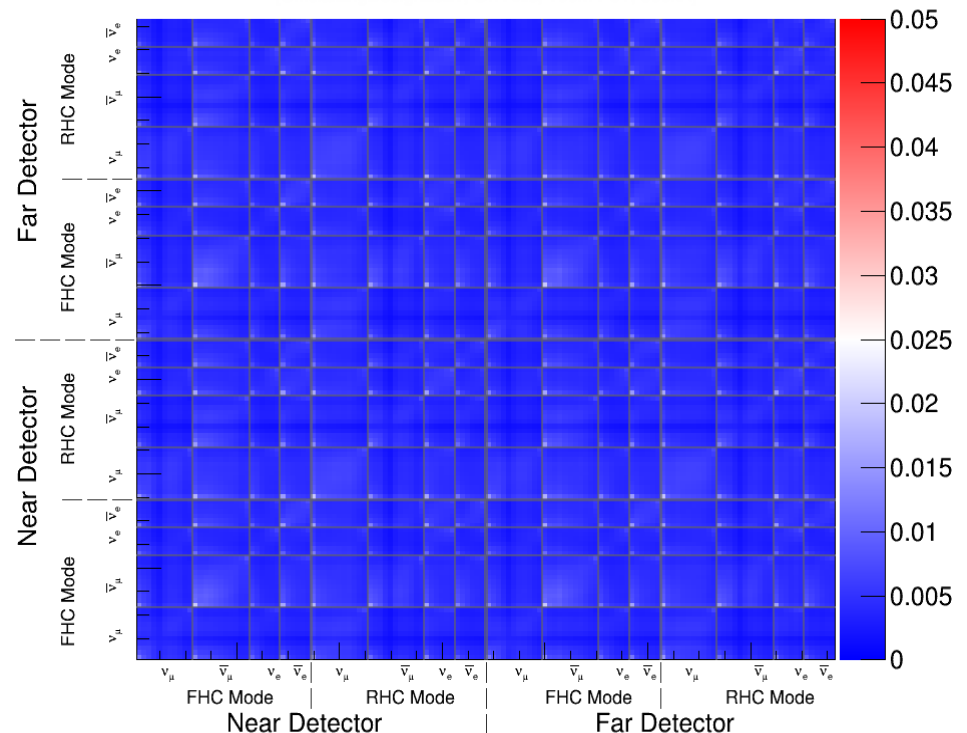
- Total BFU Covariance is average of universe covariances.

$$\text{COV}_{\text{HP}}^{(\text{total})}(z_j, z_k) = \sum_{i=0}^{\mathcal{N}} \left\langle \text{COV}_{\text{HP}}^{(i)}(z_j, z_k) \right\rangle$$

$$\therefore \begin{cases} \mathcal{N} = 9 \end{cases}$$

- Here we see the HP Covariance is likewise small indicating the magnitudes of the Hadron Production uncertainties are also, small.

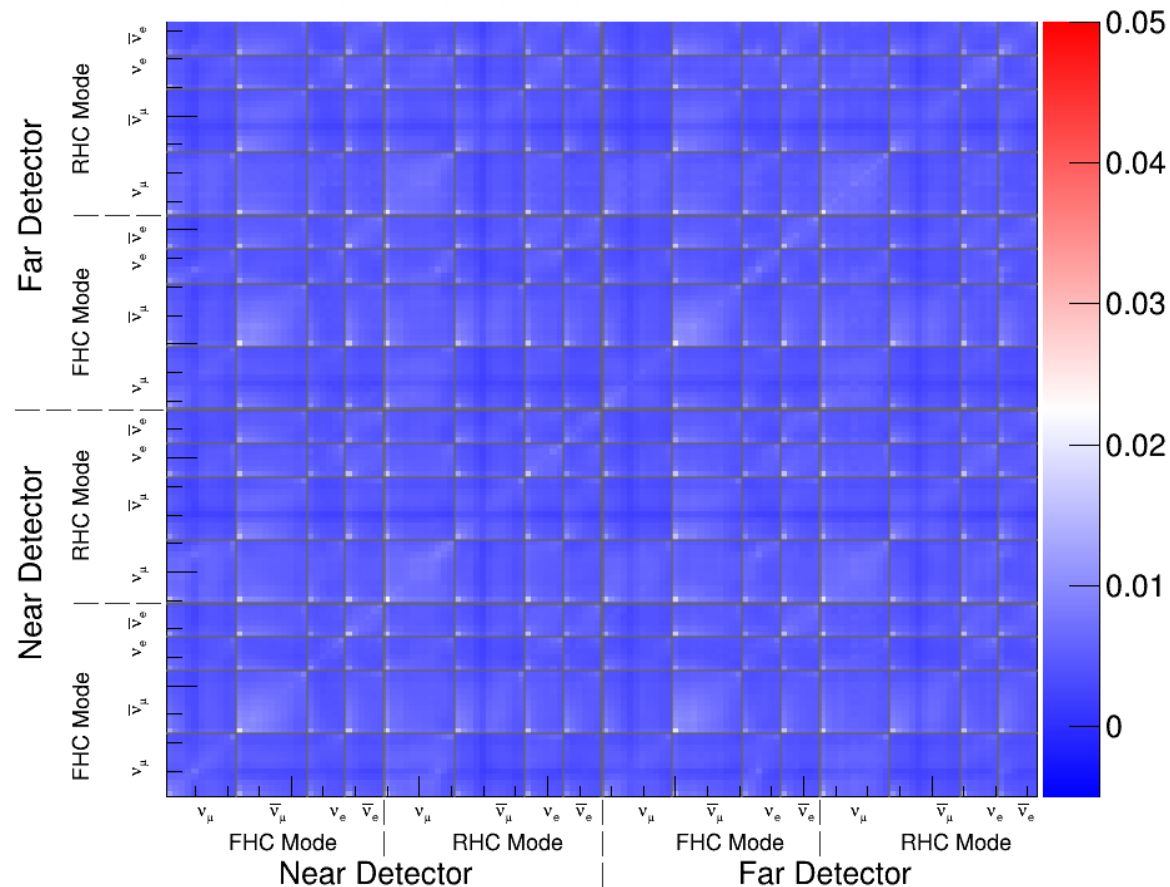
HP Covariance Matrix



The Total Covariance

- Sum of the Hadron Production and Beam Focusing Covariance matrices.

Total DUNE Flux Covariance



Total Flux Correlation

- The Total Correlation Matrix reveals the magnitude of the relations amongst the various sources of uncertainty across all modes, detector locations and neutrino species.

Total DUNE Flux Prediction Correlations

