Heavy Neutral Leptons searches on **SBND** µπ channel **New Perspectives 2024**

8-9th July 2024

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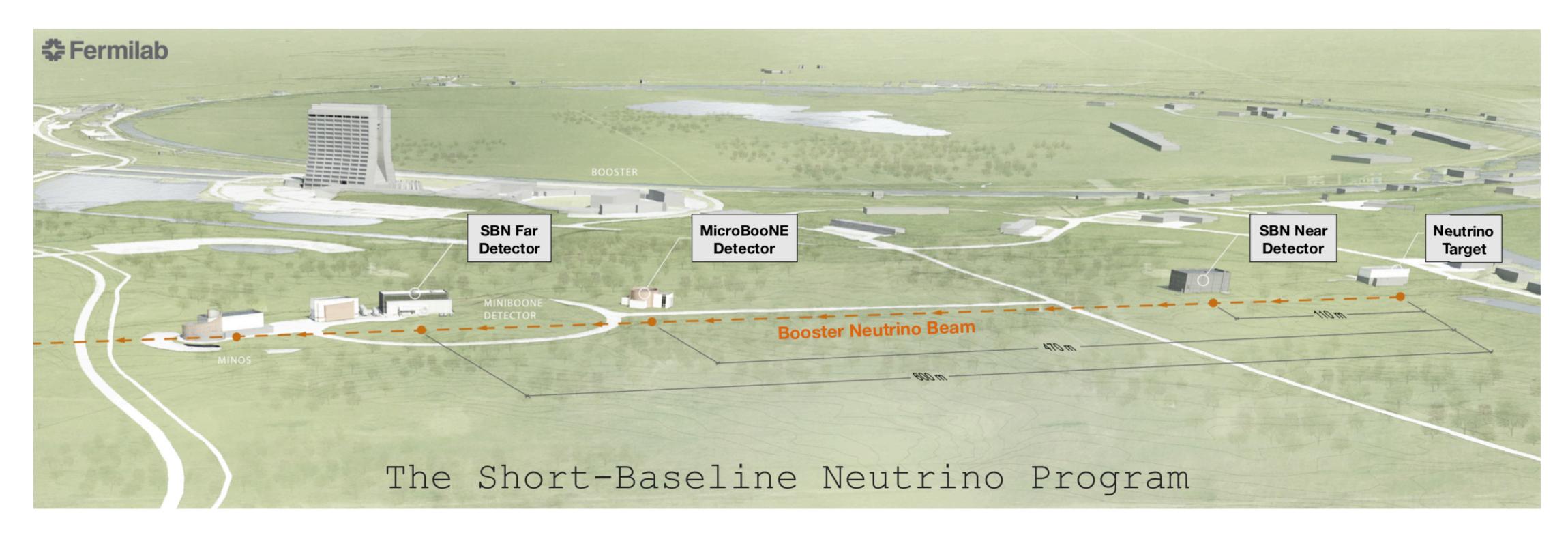
UNIVERSIDAD **DE GRANADA**



The Short-Baseline Neutrino Program

Purpose

Investigate LSND & MiniBoone anomalies in neutrino oscillations: Test the low energy excess under different models



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Facilities

3 Liquid Argon Time-Projection Chambers (LArTPC) along the Booster Neutrino Beam

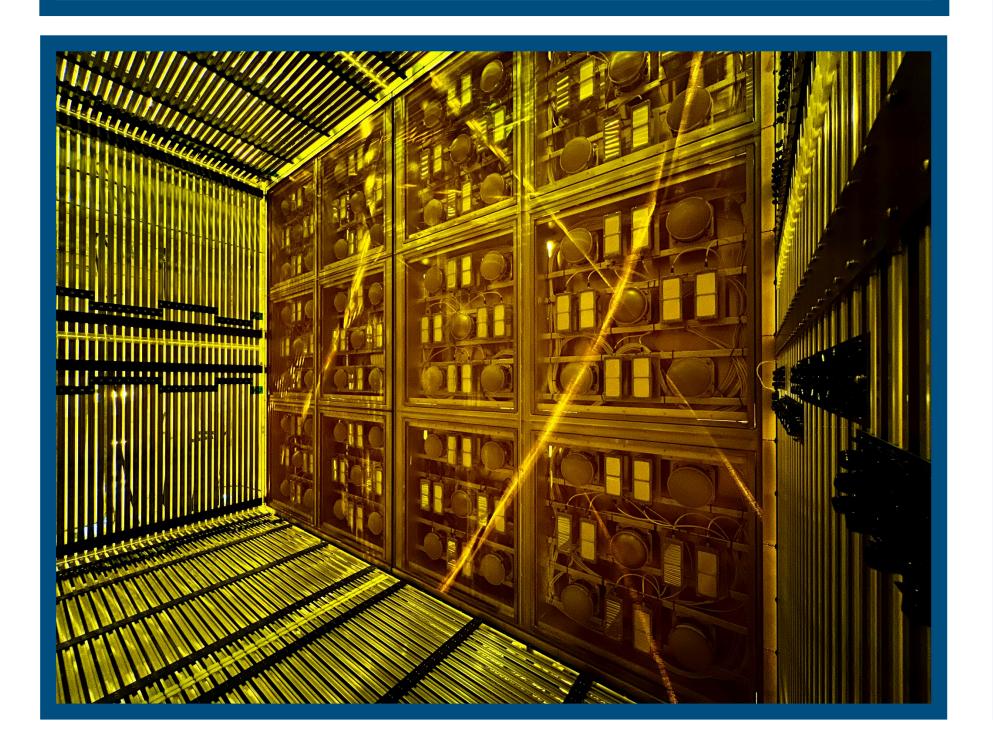


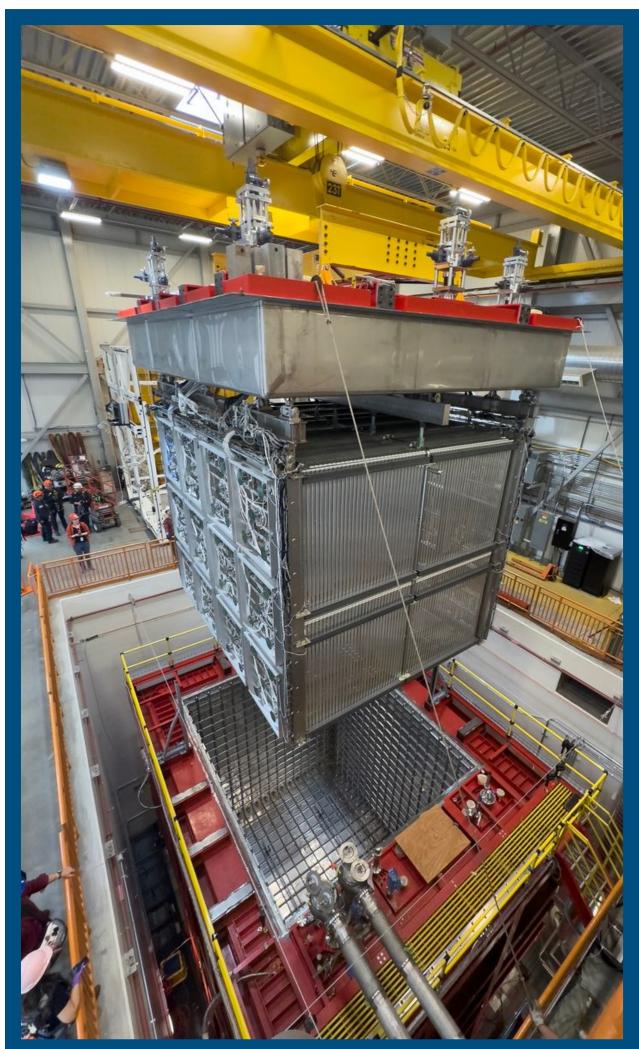




SBN Near Detector

The Short Baseline Near Detector (SBND), currently in commissioning, is a 112 ton active volume TPC located 110 m from the BNB target







Rich physics program due to proximity to the target:

- High precision measurement of the unoscillated BNB Flux
- Largest neutrino dataset of neutrino-argon interaction
- Exclusive and Inclusive cross-section measurements
- **Beyond Standard Model (BSM)** searches of particles produced in the neutrino beam

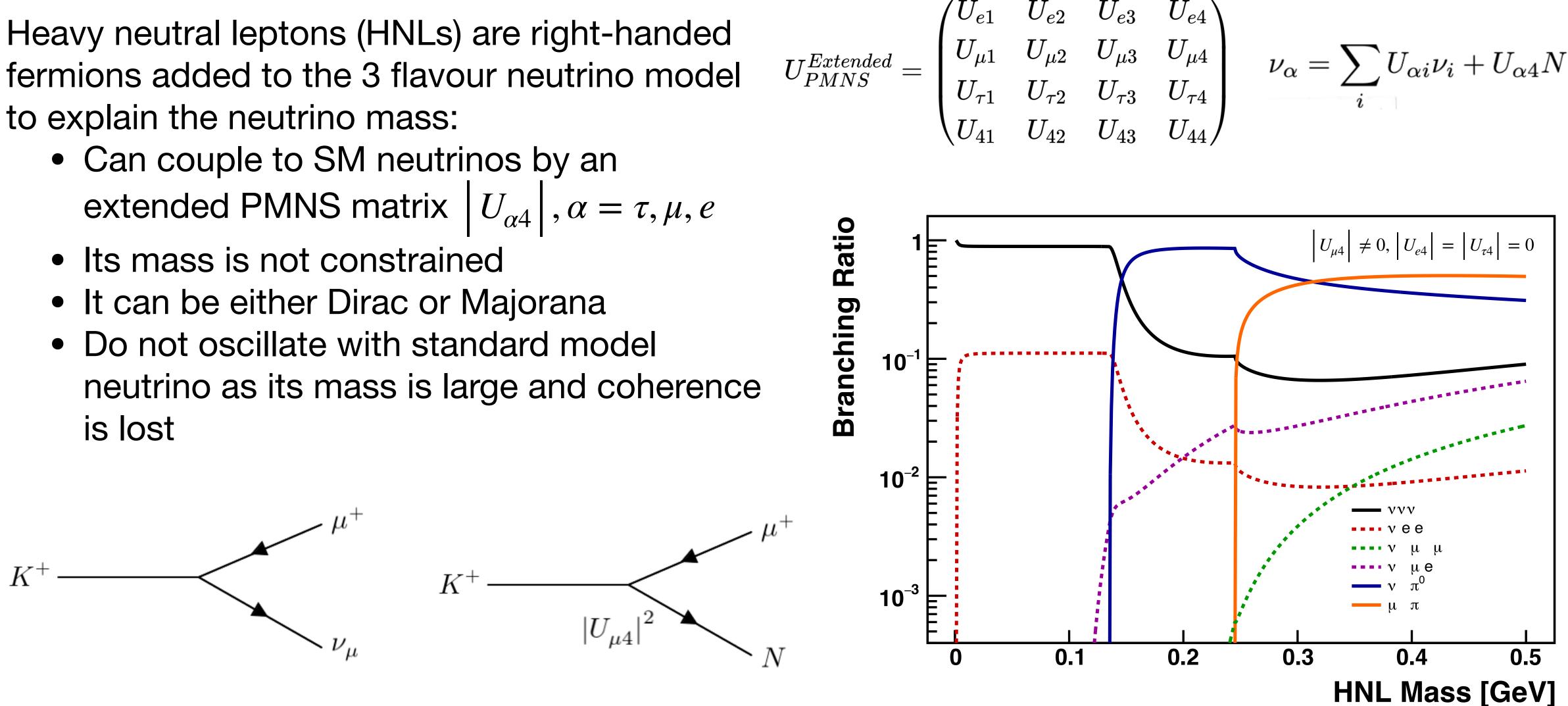






Heavy neutral leptons Definition

- is lost















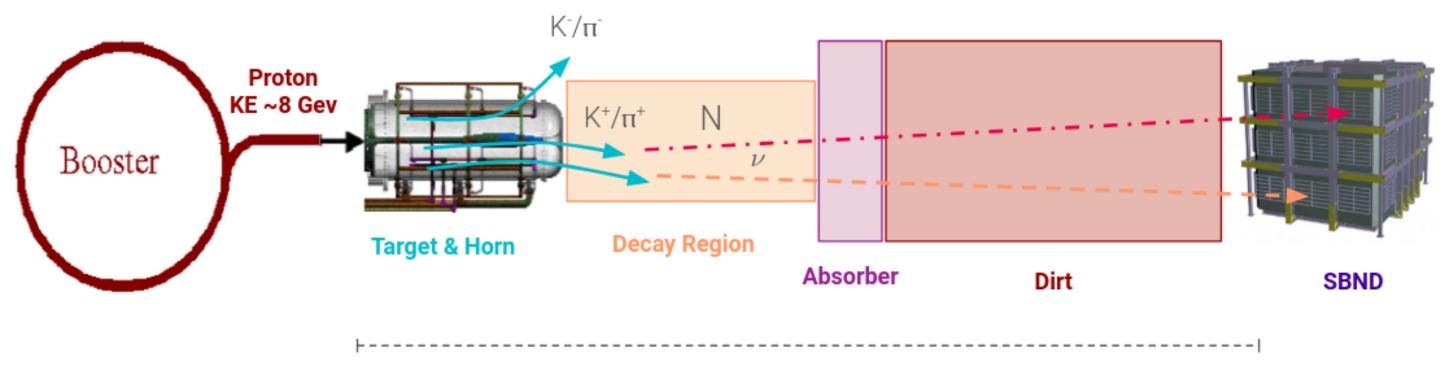
HNLs in SBND

Production and detection:

In the given theory HNLs are produced along with neutrinos in K^+ decays, constraining the explorable μ coupled HNLs mass to < 388 MeV (m_K-m_{μ})

HNLs decay inside the detector into standard model particles by three main channels:

- $\nu e^+ e^- \to m_{HNL} \in [1.02, 388] \text{ MeV}$
- $\nu \pi^0 \to m_{HNL} \in [135, 388] \text{ MeV}$
- $\mu\pi \rightarrow m_{HNL} \in [240, 388] \text{ MeV}$

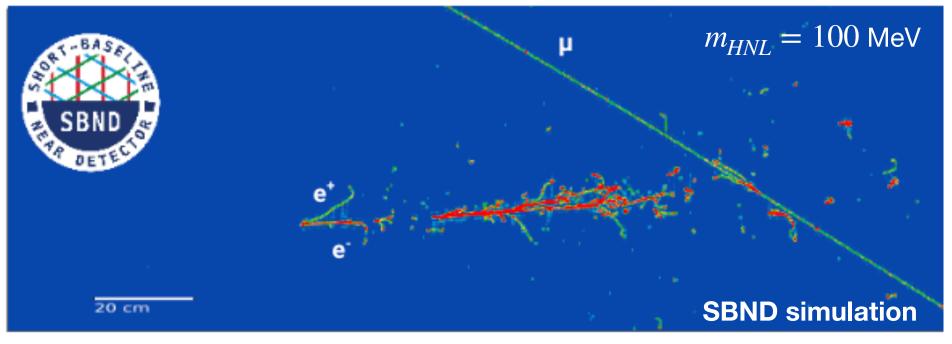


110 m

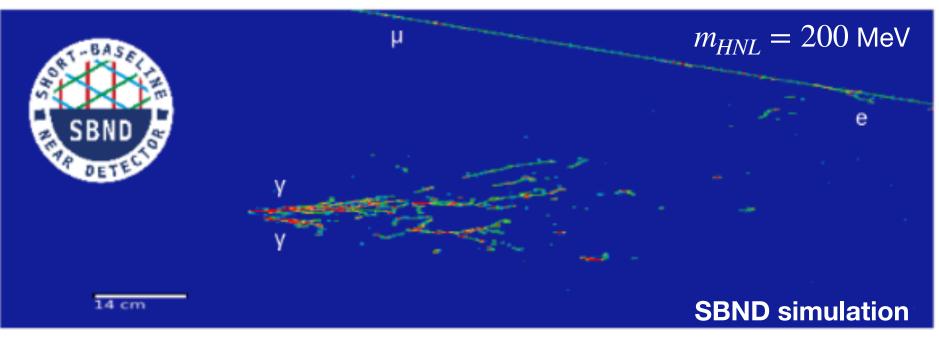
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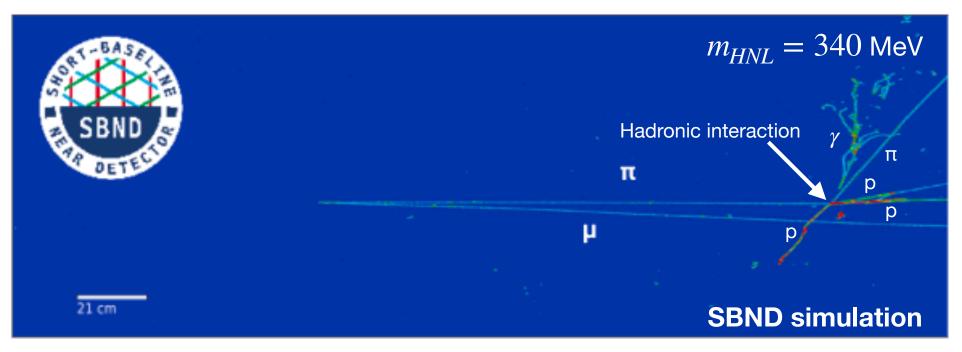
Event display for vee



Event display for vπ⁰



Event display for $\mu\pi$







μπ HNL decays on SBND

Most important channel at $m_{HNL} \in [245, 388]$ MeV

Signal definition

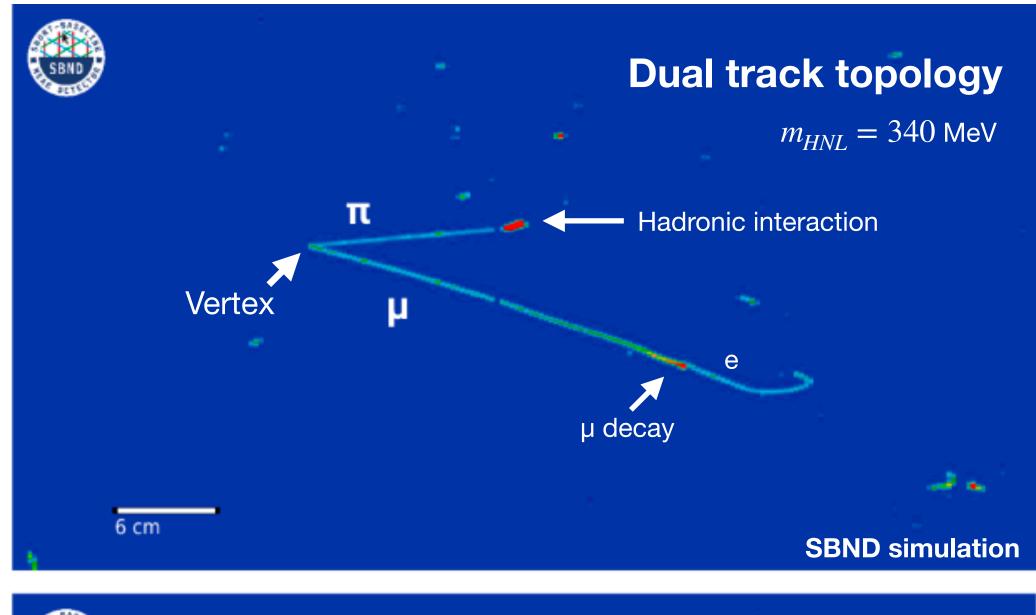
Two boosted muon-like tracks with a common vertex Delayed compared to a SM neutrino interaction

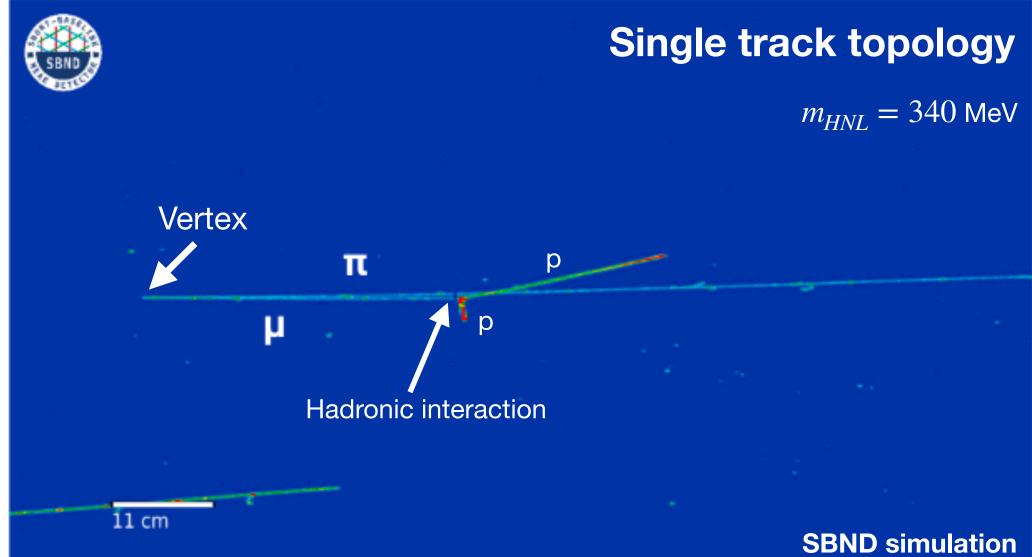
Current workflow working with full beam and detector Monte Carlo simulations. Due to detector resolution the reconstruction of this signal can be split in 2 cases:

- I track case: Some tracks appear merged near the vertex ($\theta_{\mu\pi} \sim 5^{\circ}$), when $\theta_{\mu\pi} \simeq 180^{\circ}$ some tracks are also merged. More relevant at lower mass ranges
- 2 track case: The tracks can be resolved by our reconstruction algorithms











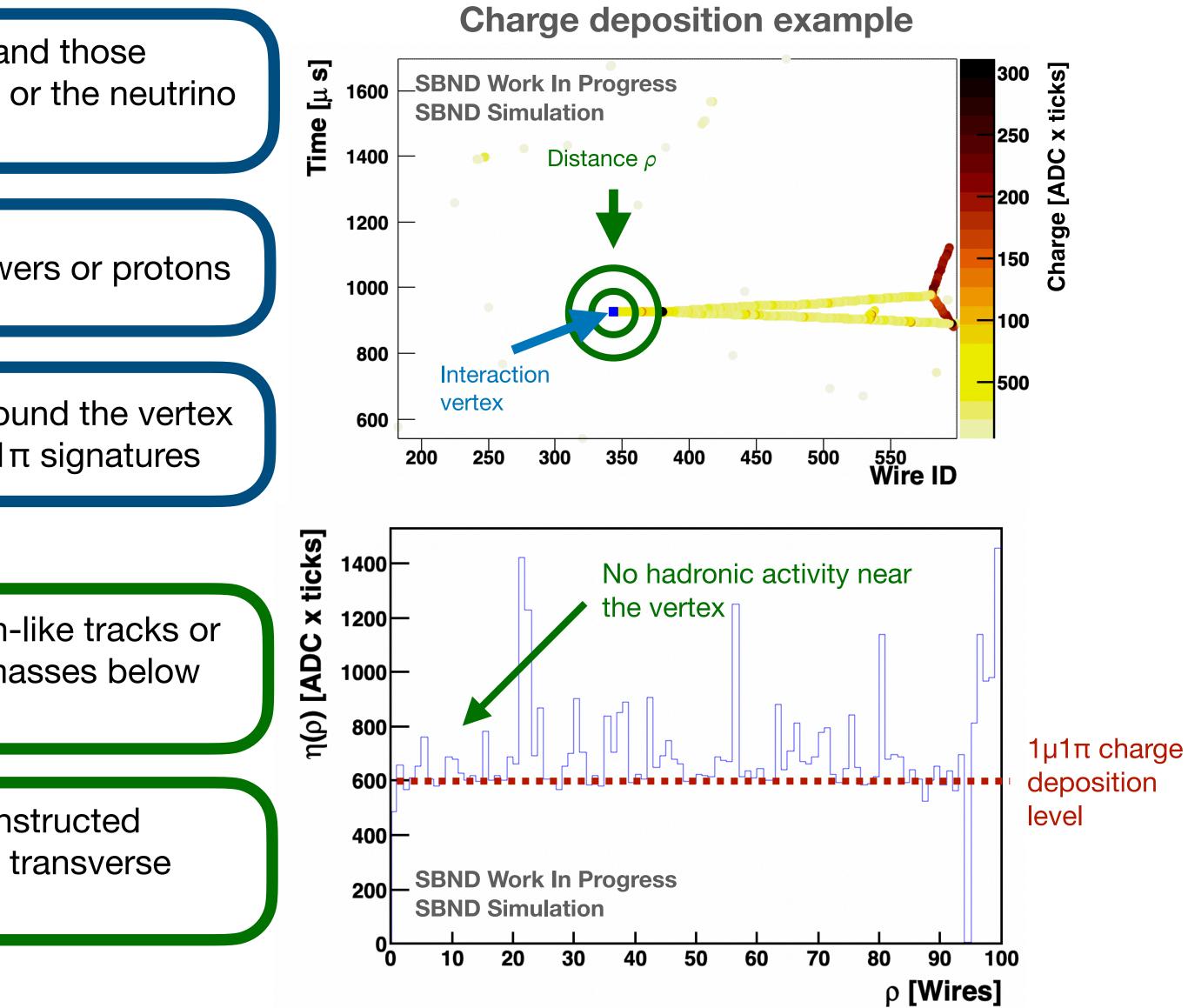


Selection strategy

1	Pre-selection & Cosmic rejection	Reject cosmic-like events an outside the fiducial volume of window
2	Particle Identification (PID) Cut	Reject any event with show
3	Charge deposition (dQdx) Cut	Use Charge deposition arouted 1µ11
Cuts independent on HNL Mass		
4	Muon candidate and topology cut	Select events with 2 muon- 1/2 muon-like tracks for ma 320 MeV
5	Kinematic Cuts	Cuts on higher level recons variables (Invariant mass, t momentum)

Cuts dependent on HNL Mass





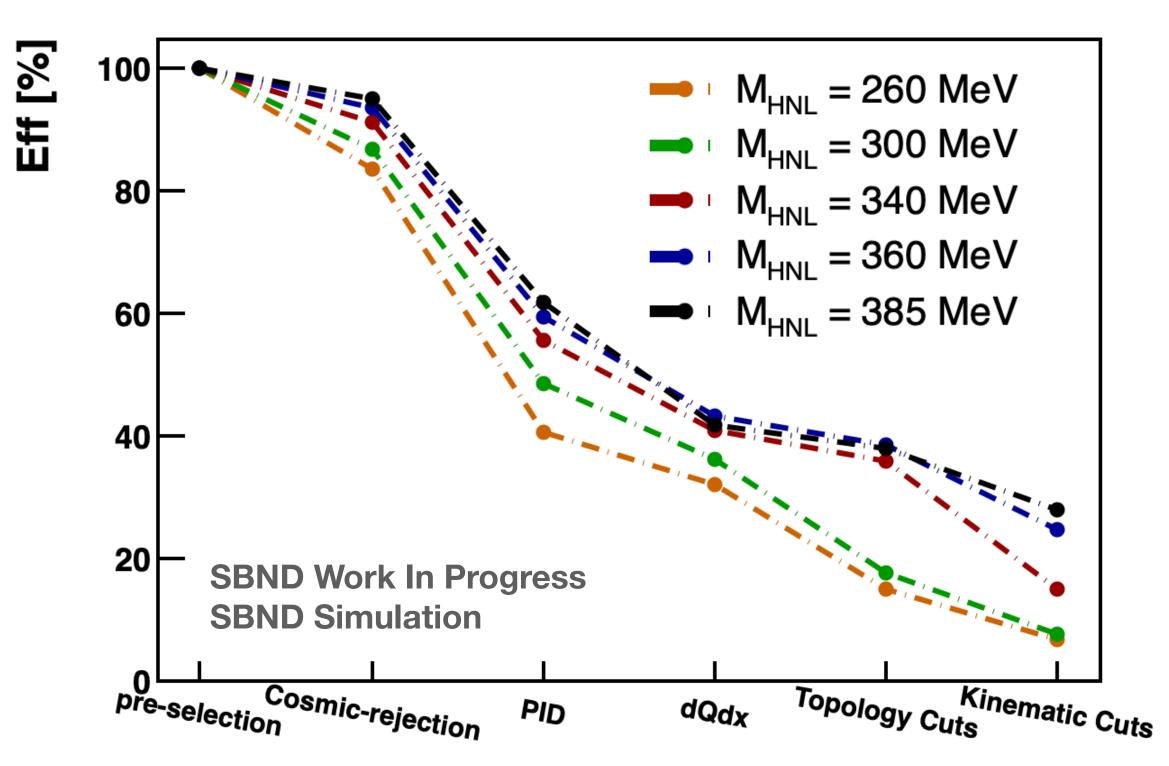






μπ selection Results

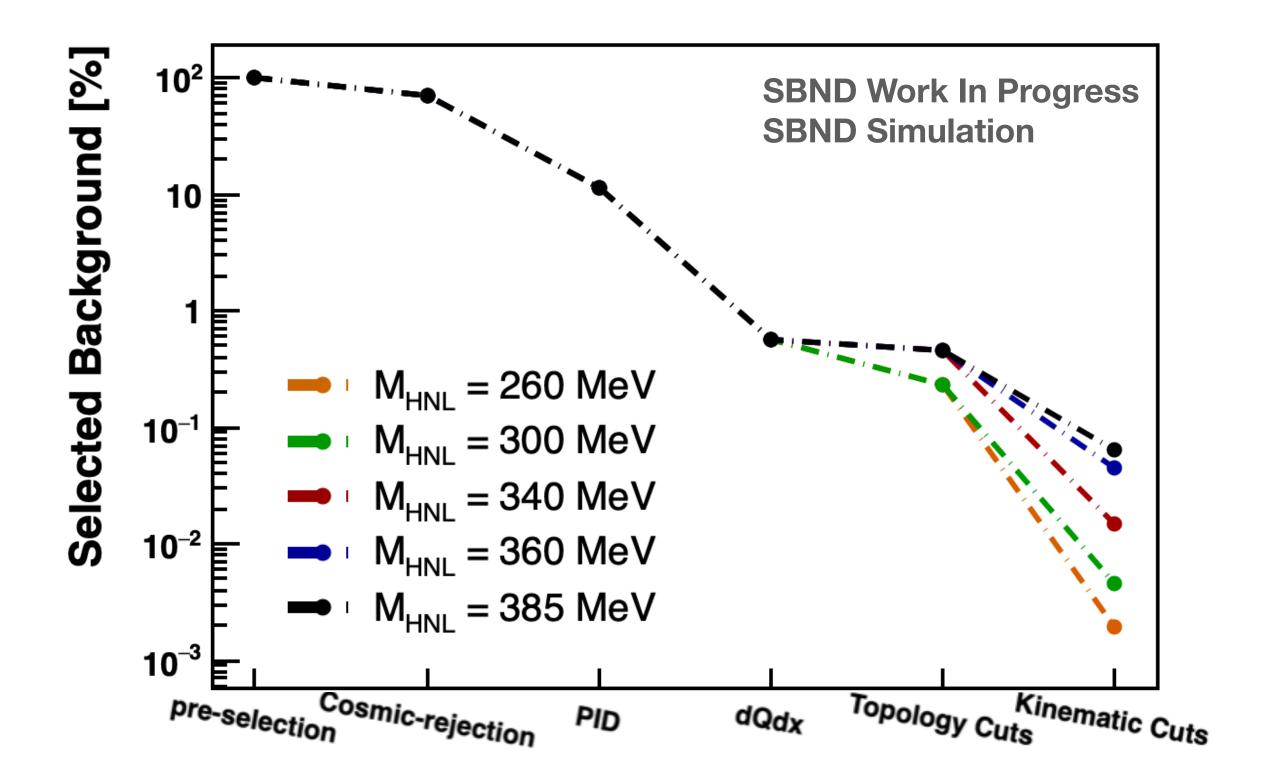
- Background reduction ranges from 3 to 5 orders of magnitude with no timing information
- Between 6.7% and 28% selection efficiency





Most effective Cuts

- dQdx Cut: Reduces background by 2 orders of magnitude
- Kinematic Cuts: Increases the background rejection at lower masses







Setting HNL sensitivity PYHF

PYHF is a python-interface of HistFactory, a statistical modelling tool built by LHC experiments

Using it we test two different hypothesis:

- Null hypothesis: signal + background
- Test hypothesis: background only

Using a multibinned histogram a likelihood function is created and an exclusion limit at 90% Confidence Level is found by varying model parameters

$$L(\mu, \theta) = \prod_{i=1}^{N} \frac{(\mu s_i + b_i)^{n_i}}{n_i!} e^{-(\mu s_i + b_i)} \prod_{\theta \in \theta} c_\theta(a_\theta | \theta)$$

Product of Poisson Constrain on how much

probability bin by bin s = signal, b = background, μ = signal normalisation

signal or background can fluctuate i.e. statistical uncertainty, etc.

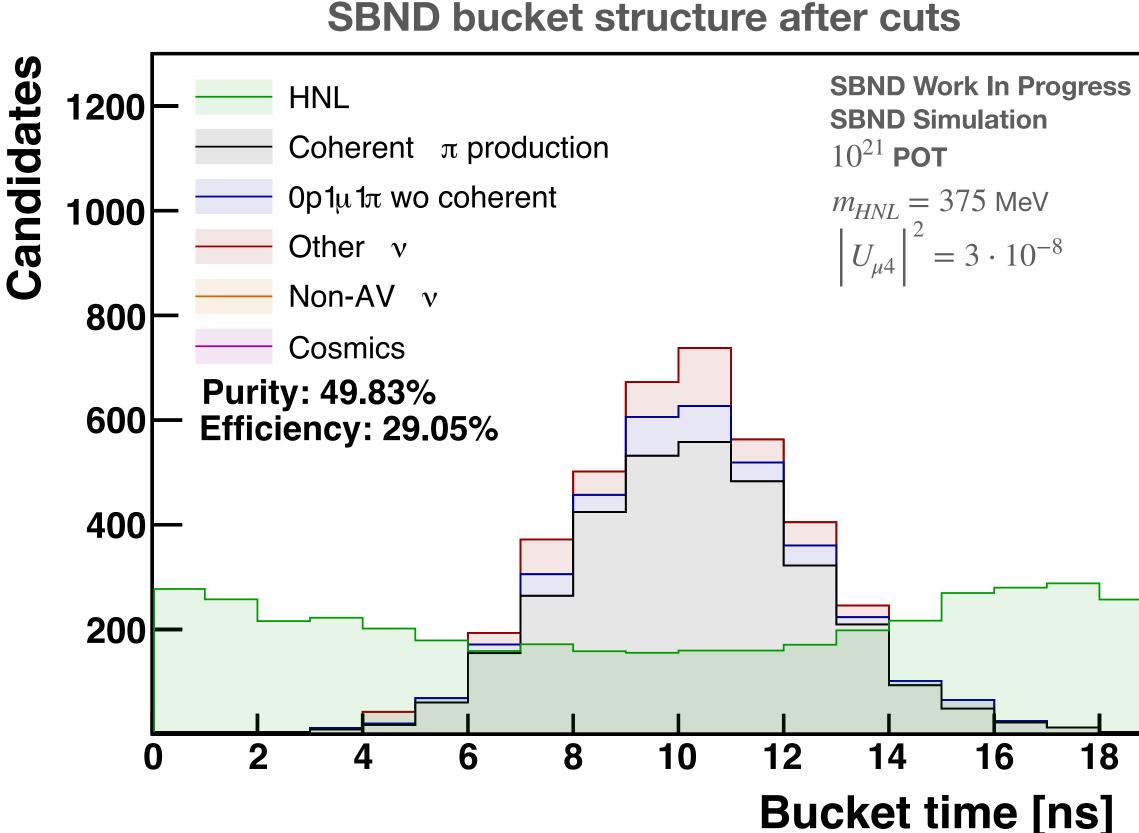
https://pyhf.readthedocs.io/en/v0.7.6/ https://arxiv.org/abs/1007.1727 https://inspirehep.net/literature/532313

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HNL timing as PYHF input

As HNL are slower than neutrinos the arrival time can be used as a good input for PYHF







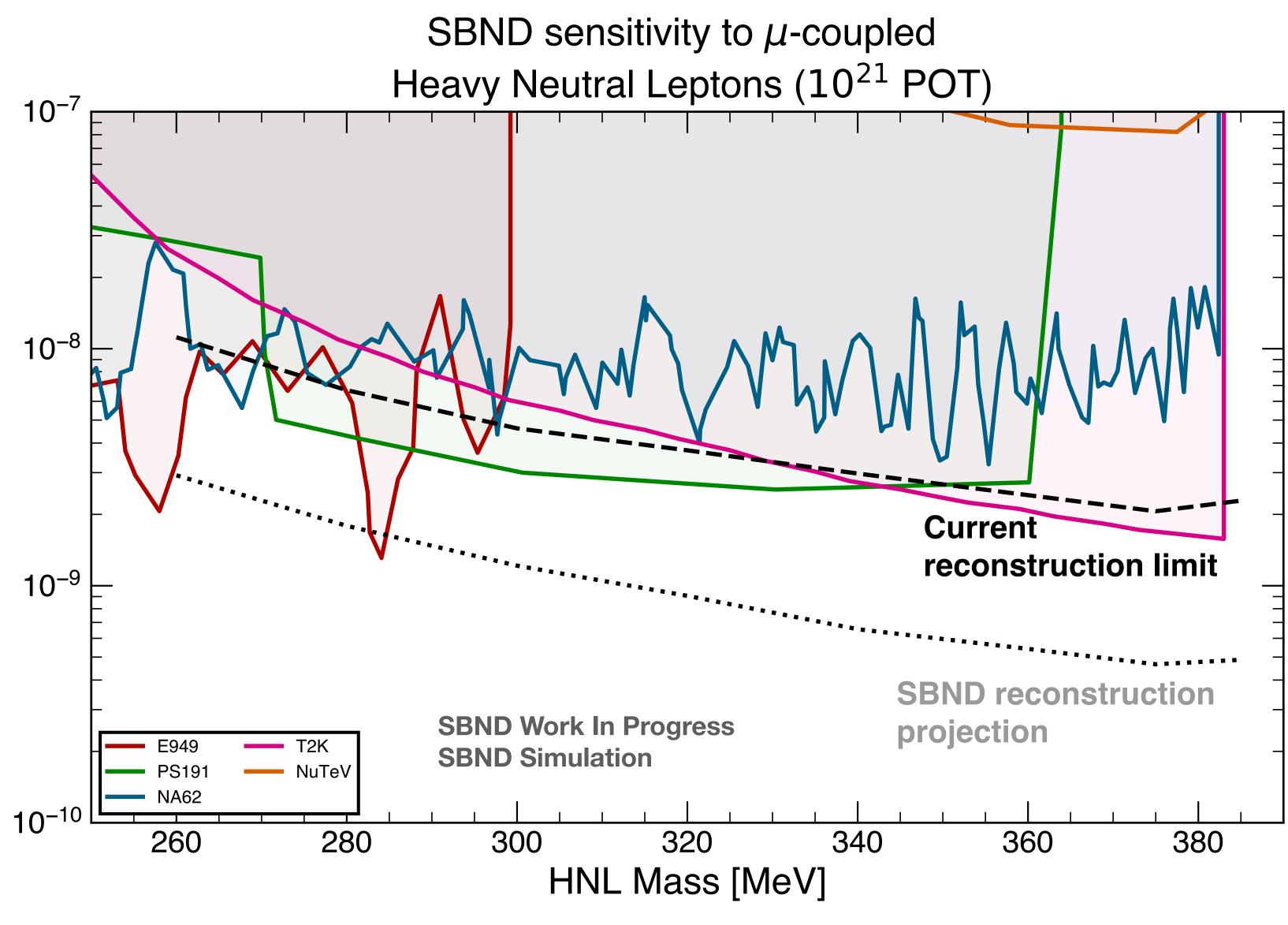


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Current and future sensitivity projection

Limit given for two cases:

- Current reconstruction limit: Ongoing development on improving SBND reconstruction tools
- SBND limit projection: Limit calculated by assuming the same efficiency and background rejection as current limit but better S/N ratio linked to an improved timing resolution of ~1.73 ns









¡¡Muchas Gracias!!!



SBND Collaboration Meeting, Fermilab, July 2024

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