FERMILAB-SLIDES-24-0142-PPD

Scintillating Bubble Chambers for Direct Dark Matter Detection, and an Update on SBC-LAr10

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On behalf of the SBC Collaboration

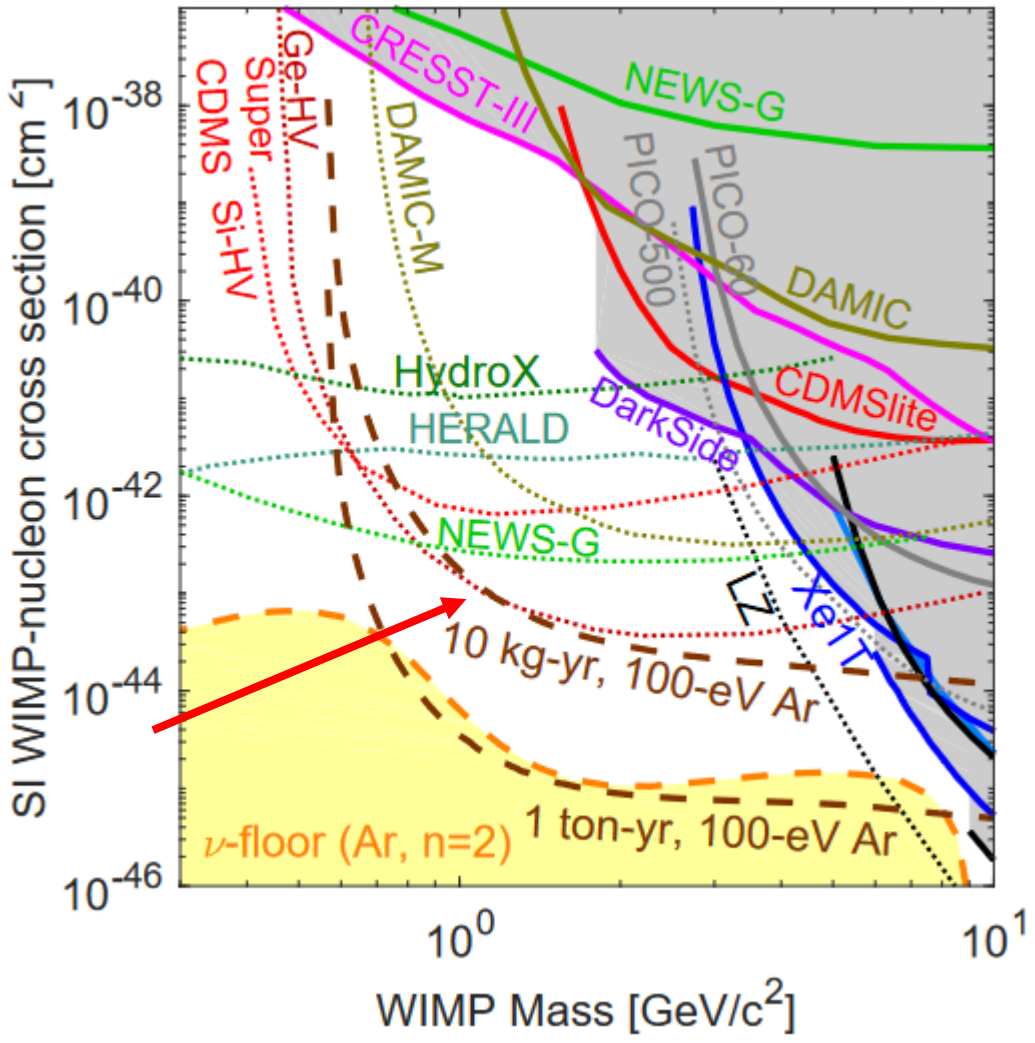
New Perspectives 2024

2024-07-09

Northwestern

Department of
Physics and Astronomy

Scintillating Bubble Chamber Technology



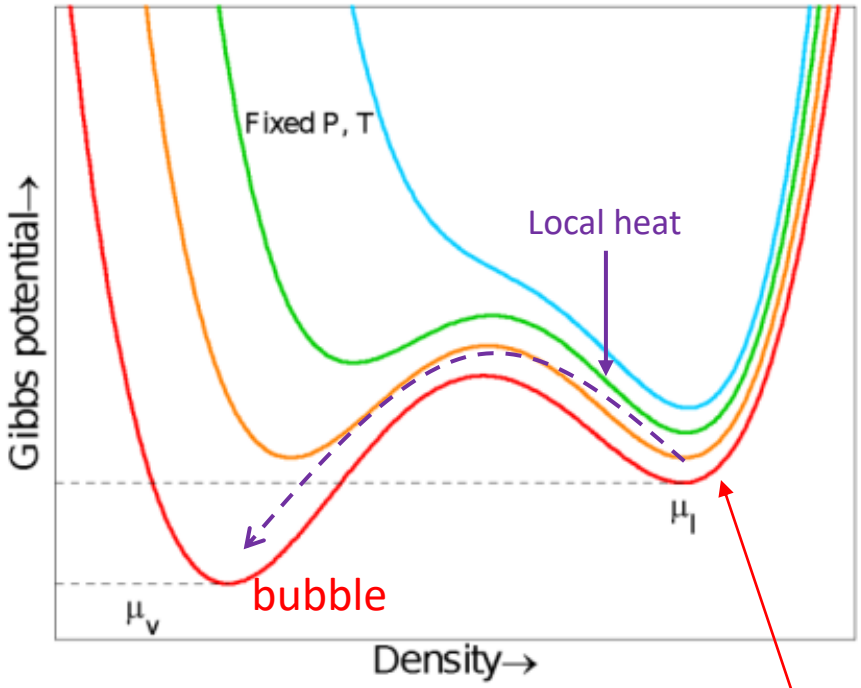
1. Scintillating Bubble Chamber (SBC) combines **powerful electron recoil (ER) discrimination** technique with **event-by-event energy resolution** of liquid-noble scintillators

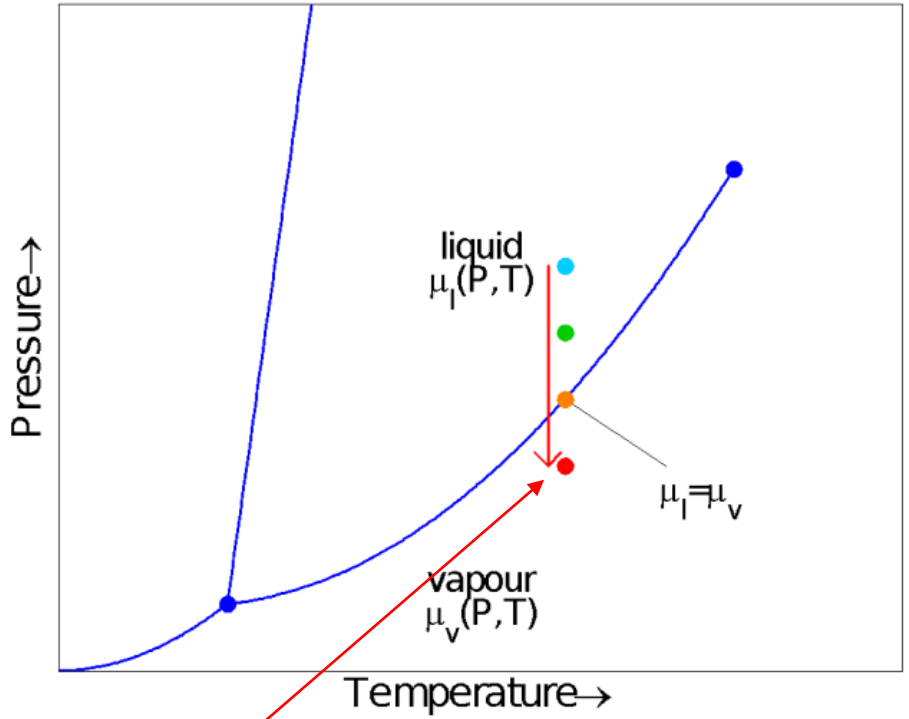
2. DM masses down to ~1GeV (~100eV nuclear recoil on argon nucleus)

3. SBC is the only technology that can achieve this has the **discrimination** and **scalability** needed for a ton-year exposure at 1 GeV WIMP mass

SBC Snowmass White Paper, arXiv: 2207.12400v1

Superheated Fluid and Bubbles





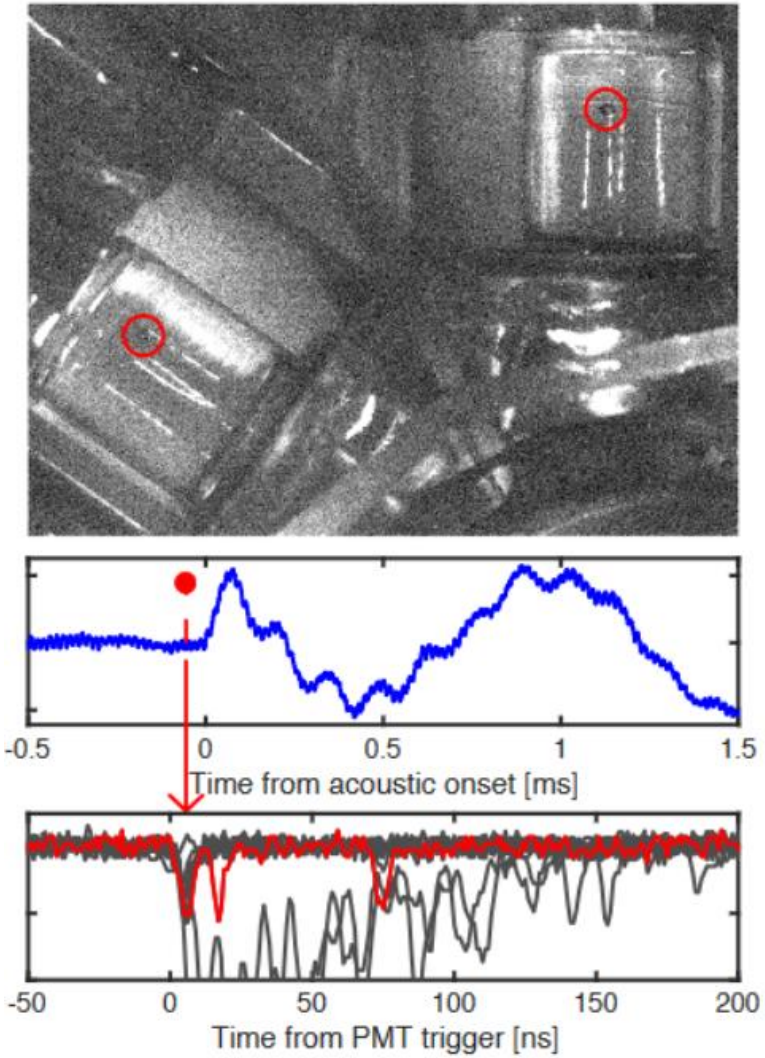
Superheated fluid

Fast Neutron / Dark Matter → Nuclear recoil → Large dE/dx → Heat deposited locally → Bubble

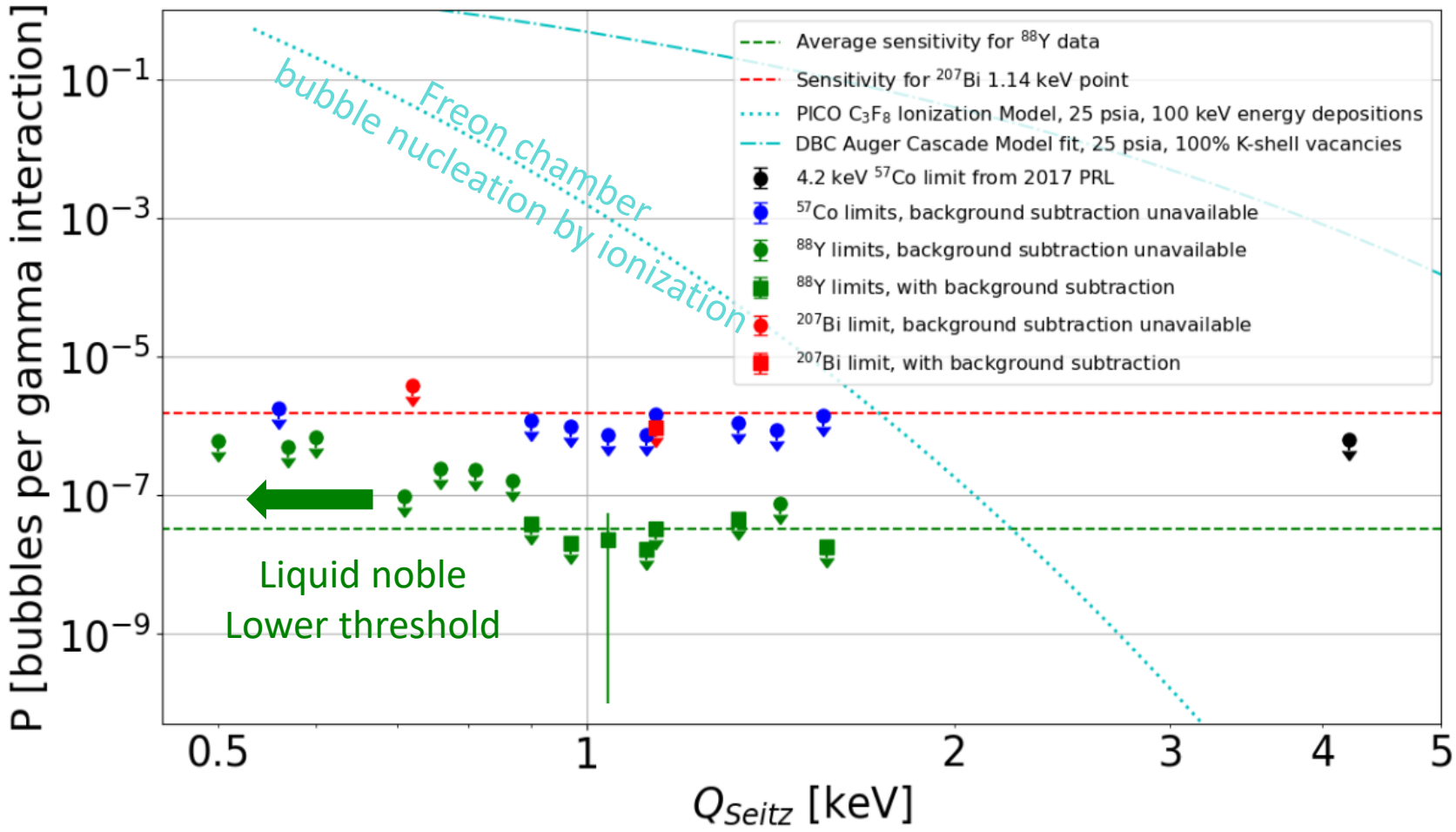
Beta / Gamma → Electron recoil → Small dE/dx → Heat deposited over long distance → No bubble

(In liquid noble) → No heat, only radiation → Less than no bubble

Electron Recoil Discrimination

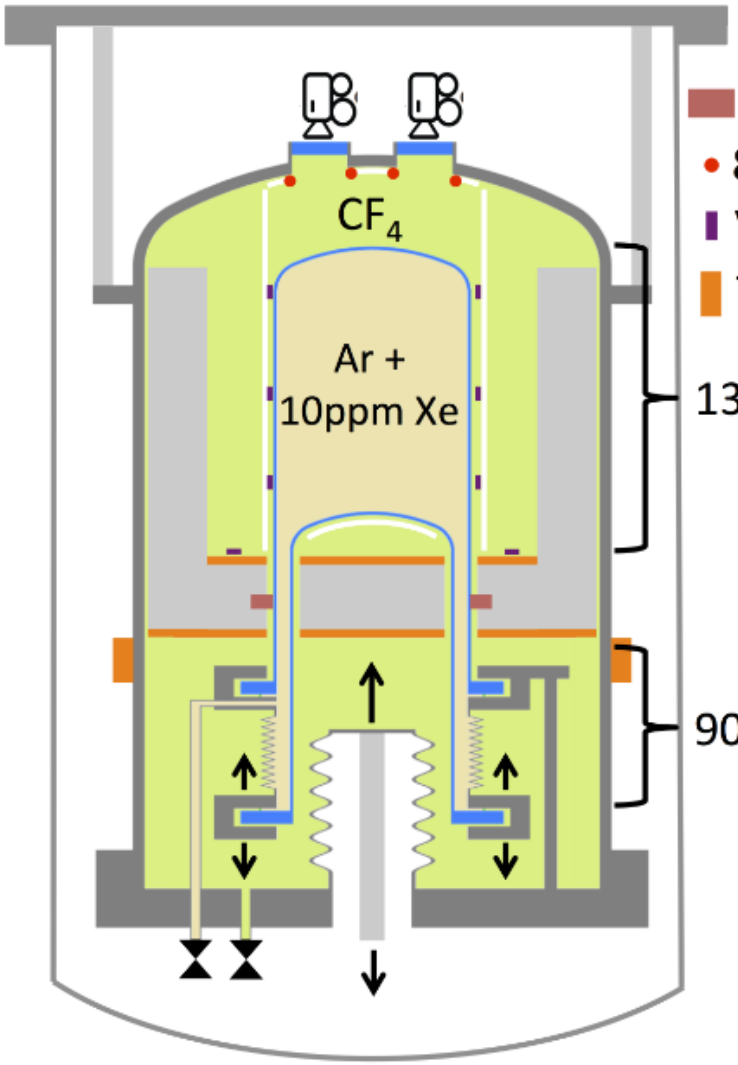


Prototype Xenon bubble chamber



SBC Snowmass White Paper, arXiv: 2207.12400v1

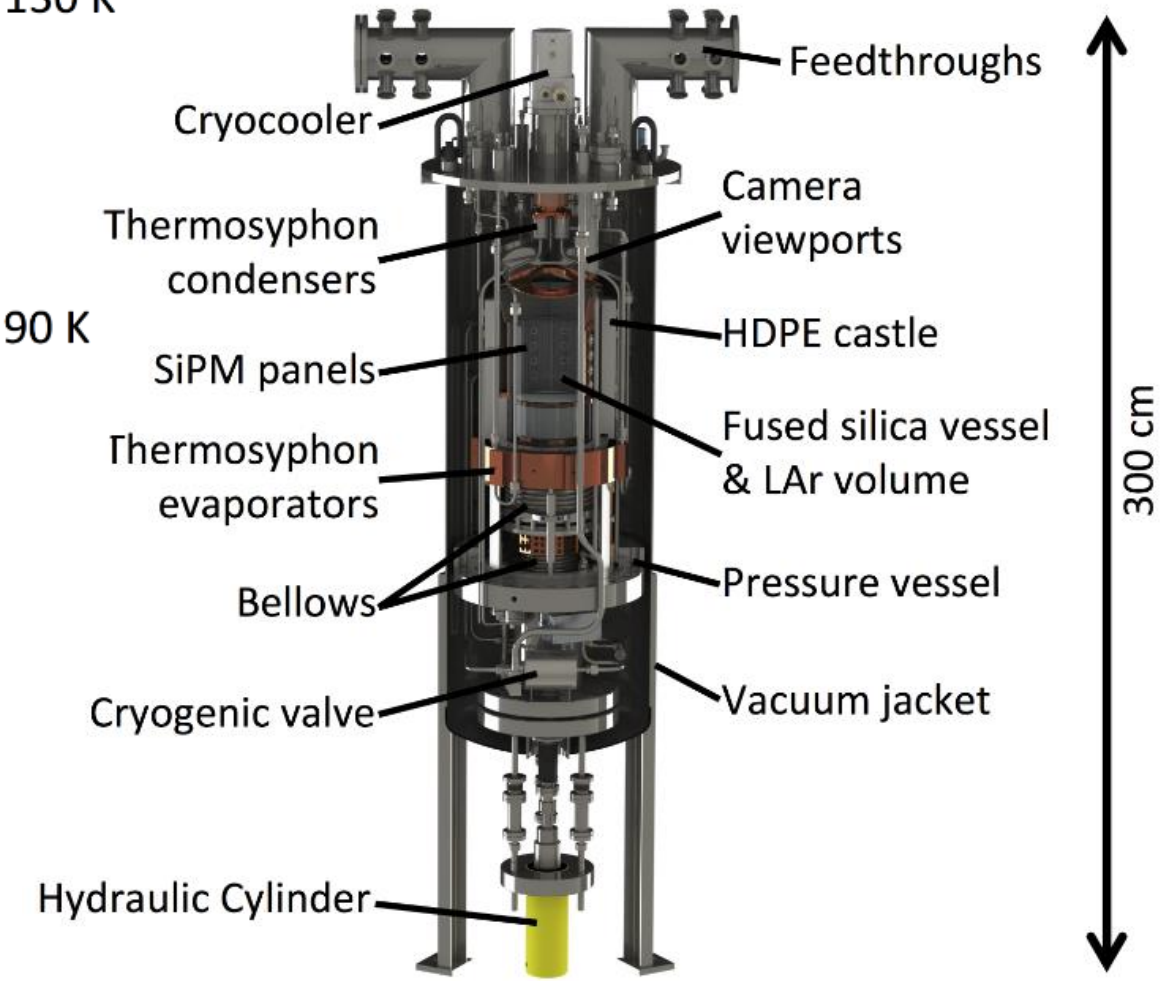
SBC-LAr10 Fermilab Design and Update



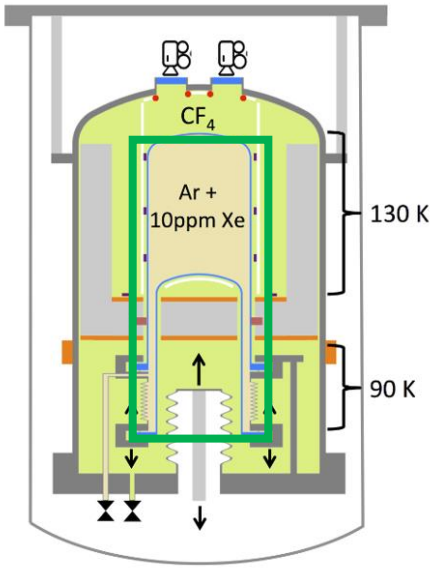
SBC-LAr10 Fermilab
10kg calibration chamber recently installed in MINOS

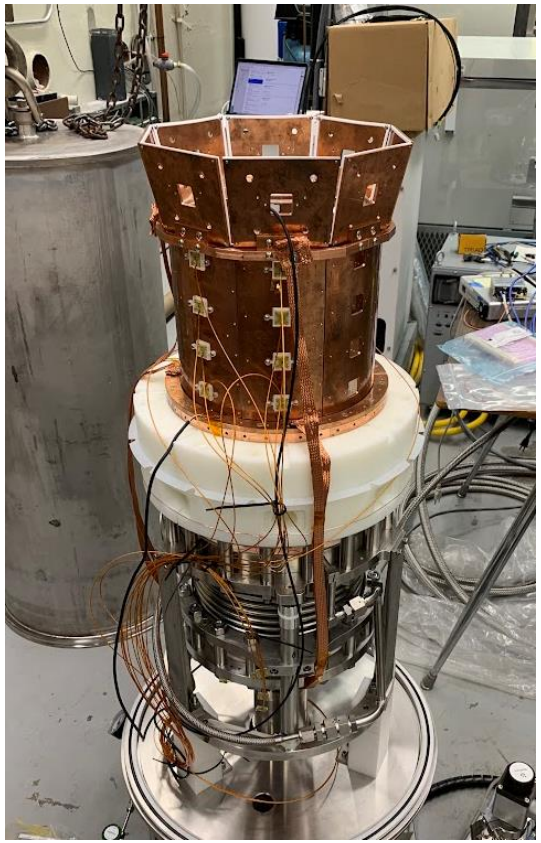
SBC-LAr10 SNOLAB
Functionally identical radiopure chamber
Dark matter search

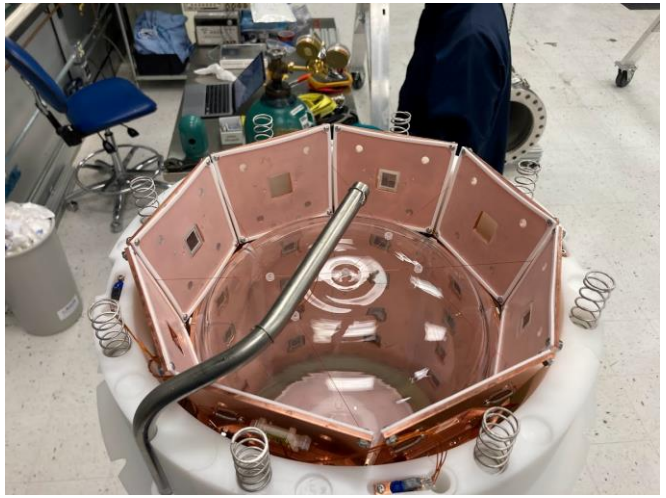
Ton-year chamber
Explores solar CEvNS at 1GeV

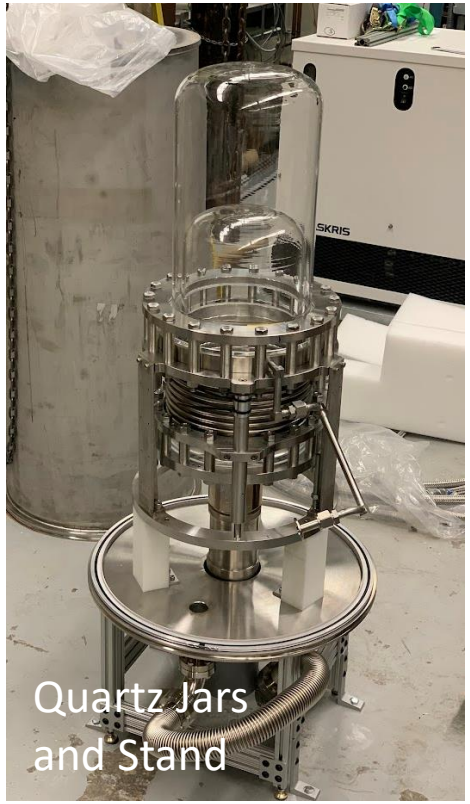


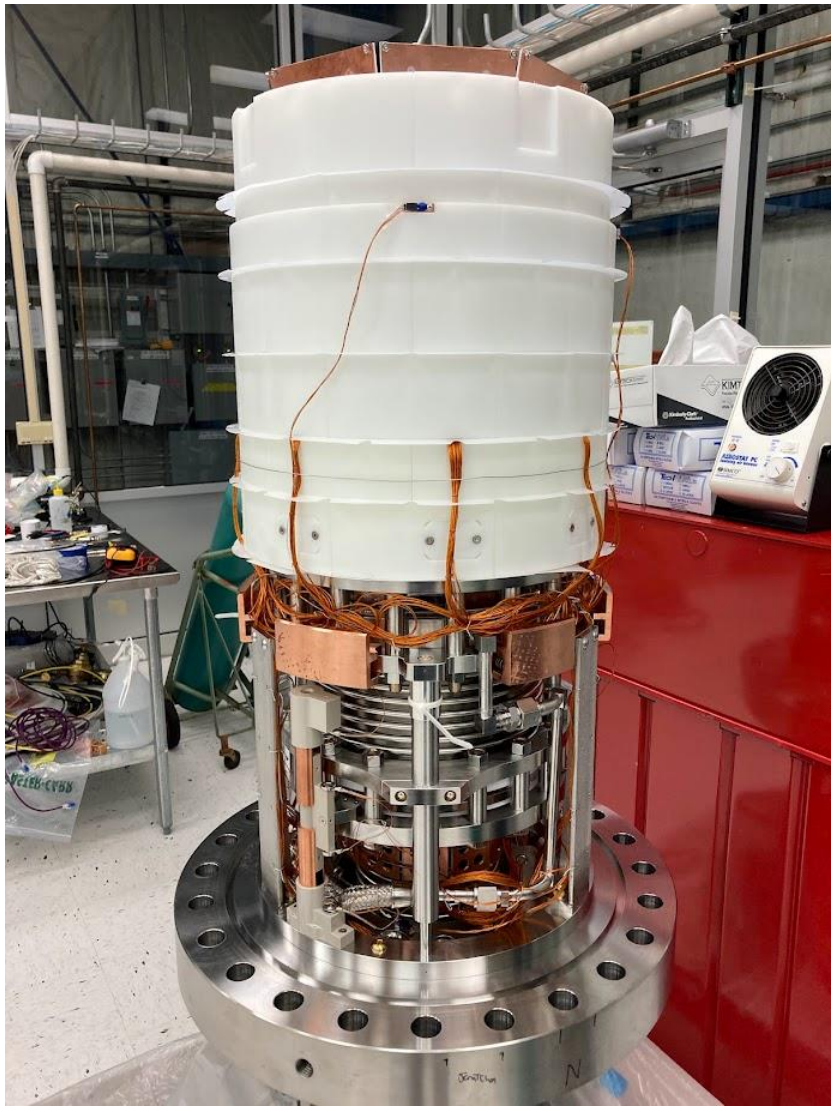
SBC Snowmass White Paper, arXiv: 2207.12400v1



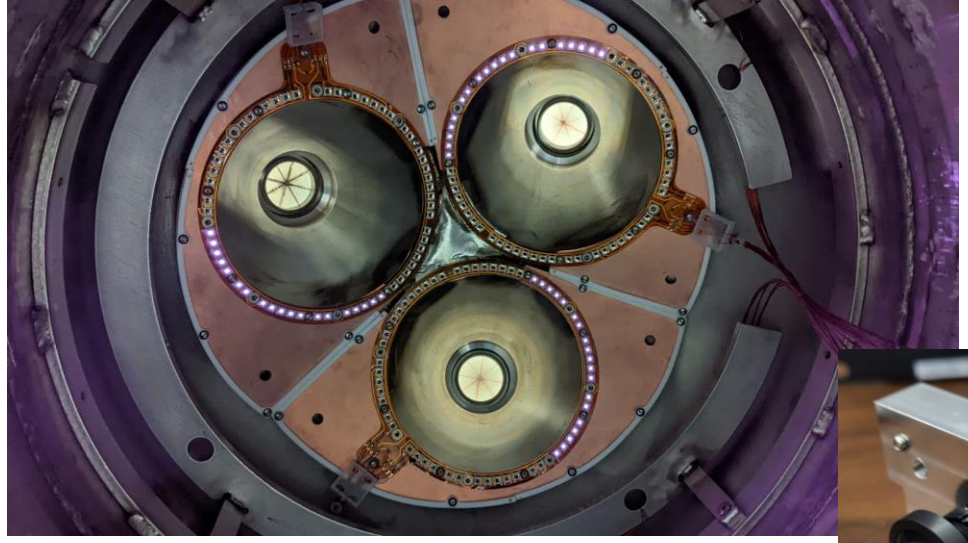




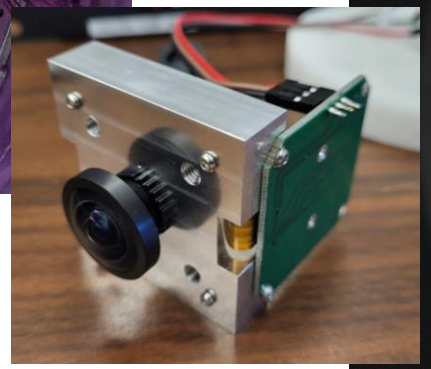




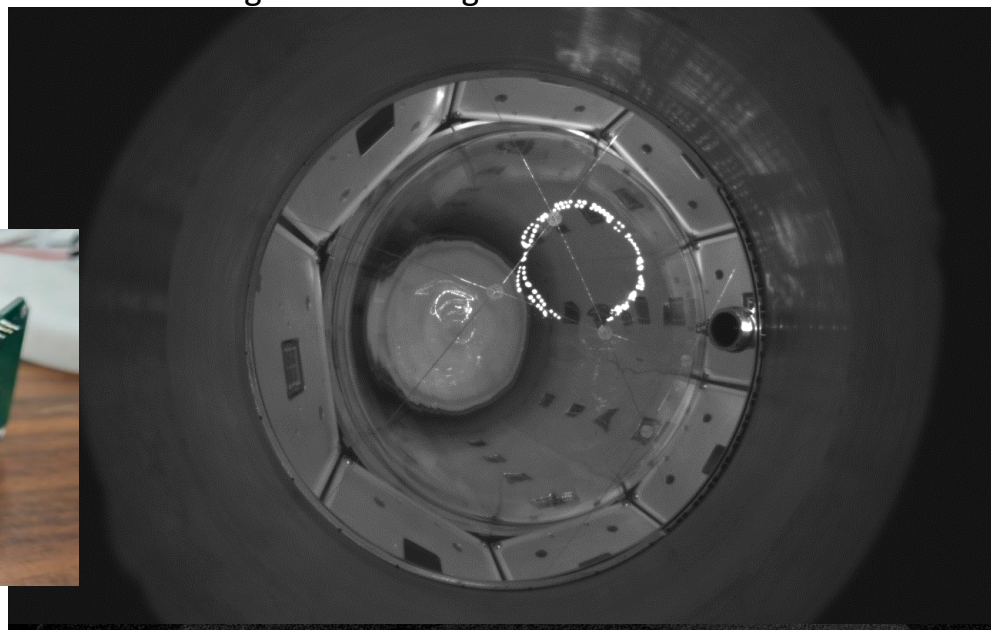
Inner Assembly

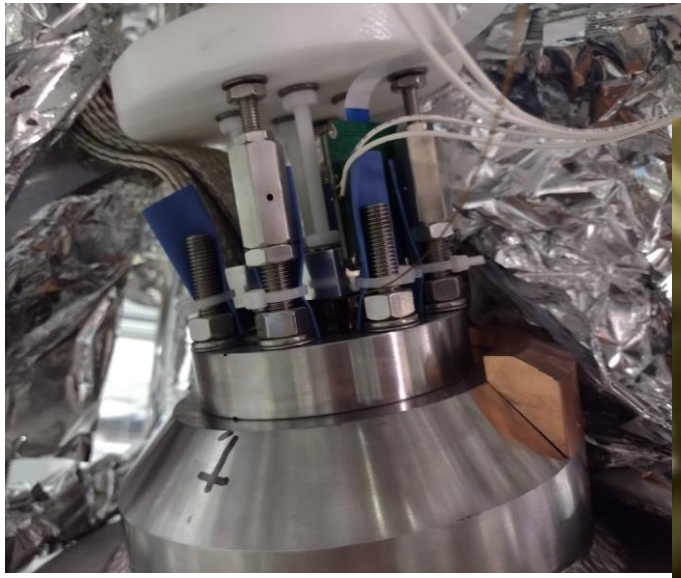


Three camera ports with LED rings

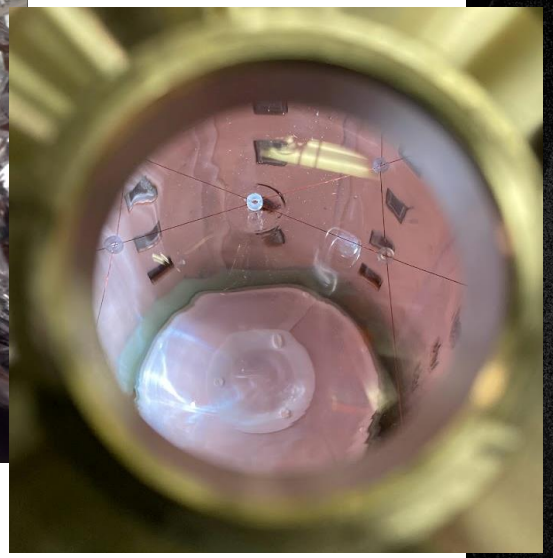


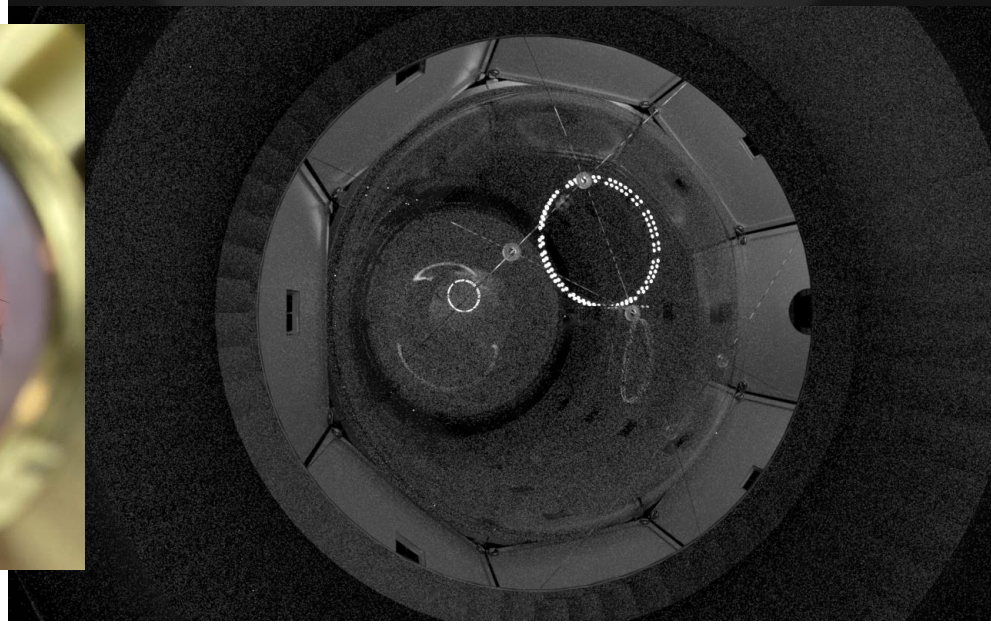
Real image taken during cold test run



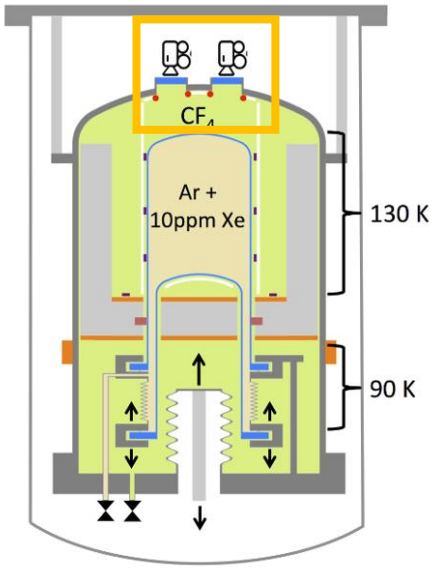


Camera mount on view port

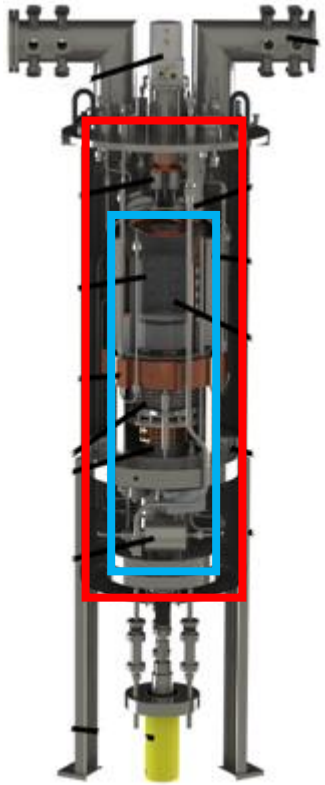


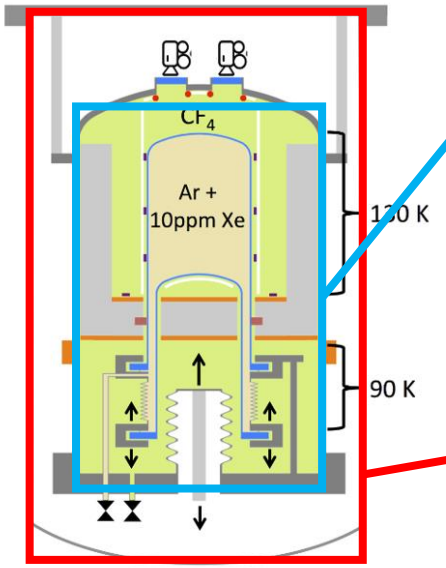


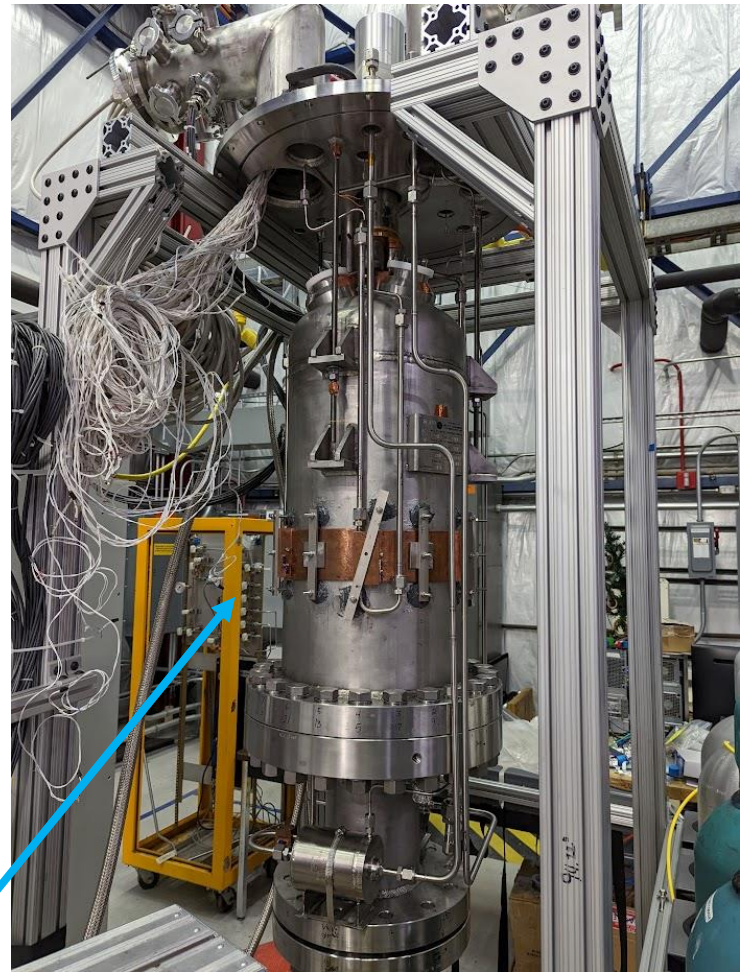
Rendering using Mitsuba 3



Bubble imaging (cameras)







Pressure vessel

Super-insulation

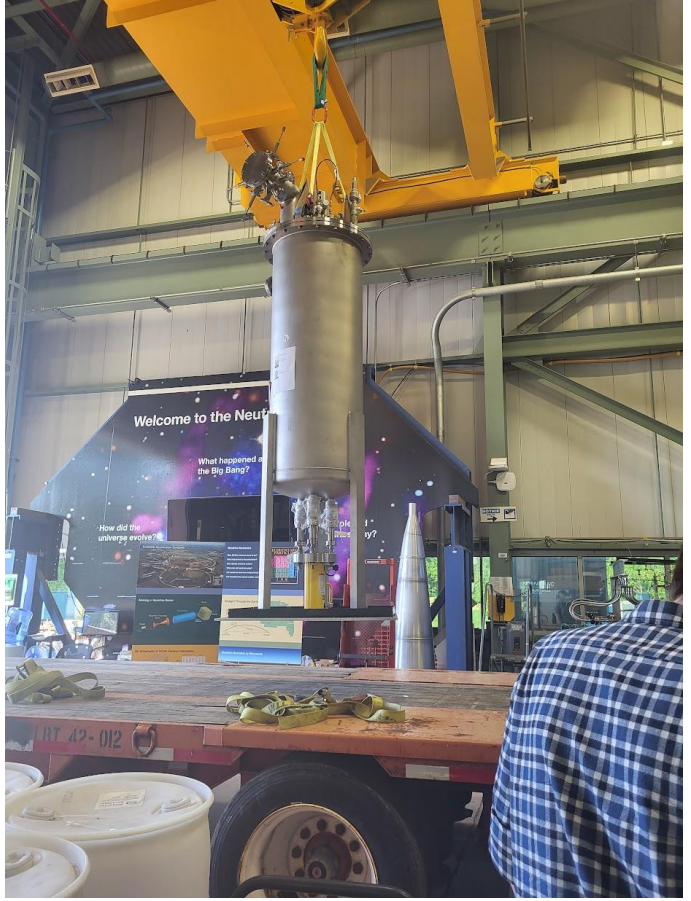
Vacuum jacket

Vacuum jacket and pressure vessel

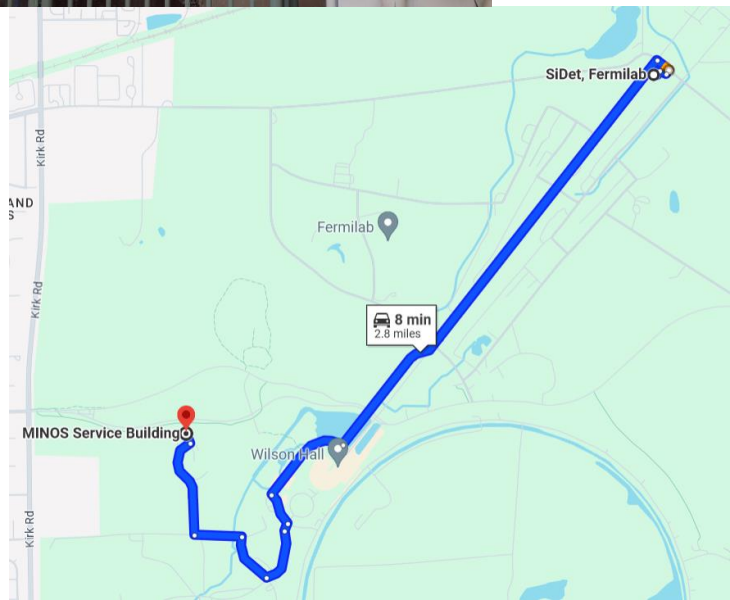
Moving to MINOS

May 15, 2024

Departing SiDet Lab B



Arriving at MINOS building



QUIET

DUNE 2x2

NEXUS

Elevator

Moving to MINOS

May 15, 2024

Arriving underground

SBC-LAr10@Fermilab Calibration Plan

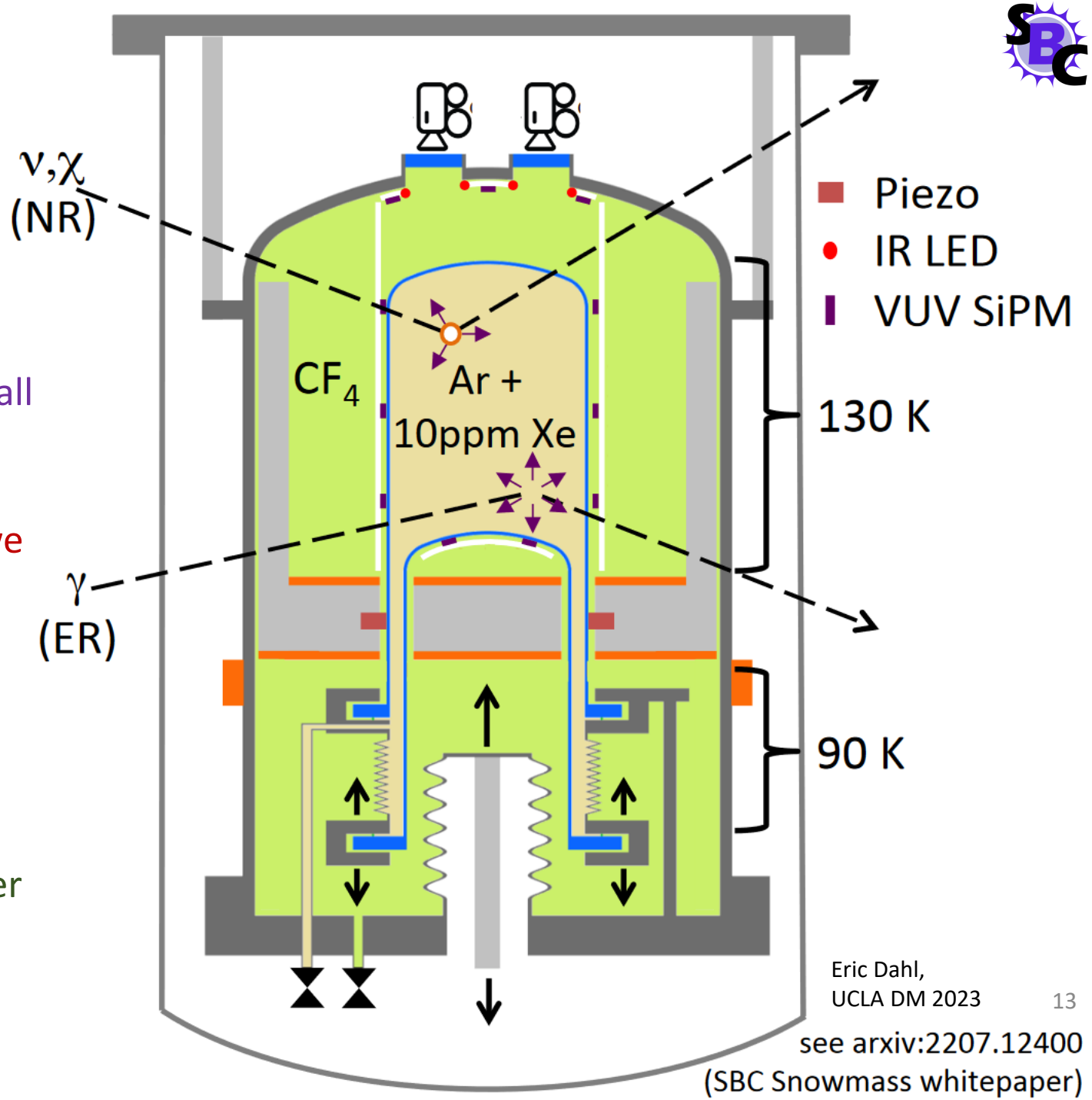
Questions to answer with the calibration chamber:

Can we operate the chamber as expected? Do all the data streams work together?

What's the lowest thermodynamic threshold we can operate the bubble chamber at while still being insensitive to ERs?

What's the lowest energy of NRs that can nucleate bubbles at that threshold?

Can we set a (even if not very good) dark matter limit with the background data collected?



see arxiv:2207.12400
(SBC Snowmass whitepaper)

Electron recoil (ER) calibration

- Operation of chamber at multiple thermodynamic thresholds
- Multiple gamma sources – explore limit of ER blindness
  - Consistent elevated bubble rate above background
- Used to determine ideal operating conditions
- Scintillation can also be used to study ER light yield in LAr and LCF4

How superheated can we get?

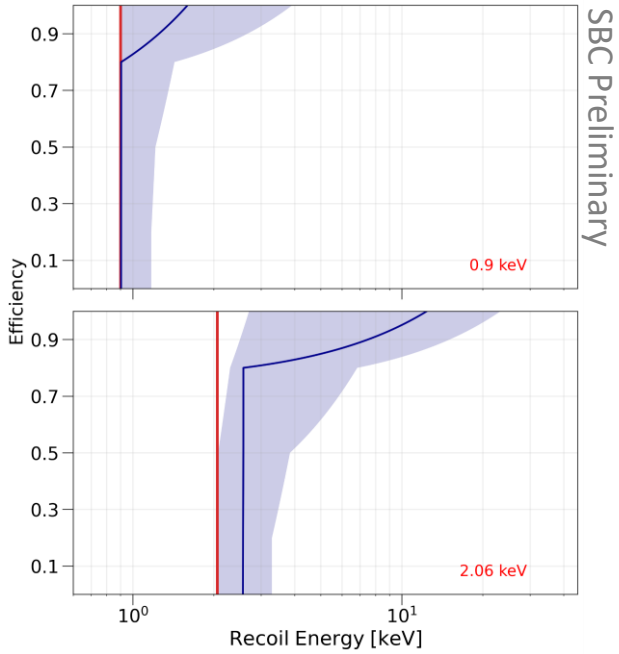
Liu 2023 – All metal argon bubble chamber

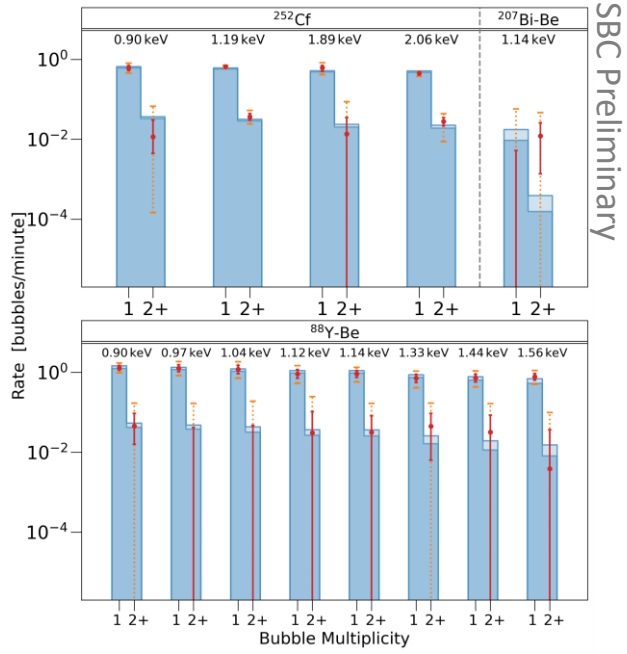
Nuclear recoil (NR) strategy

- Run the chamber at target (100eV) thermodynamic threshold
- Using multiple sources with argon nuclear recoil energies spanning 100eV to 3keV range
- See which of the NR sources nucleate bubbles

Why this is necessary:

- calculated threshold is only accurate to a factor of 2 at best (red and blue lines not overlapping)
- The threshold calculation has never been tested below 1keV





Red: calculated thermodynamic threshold

Blue: measured bubble nucleation efficiency from various sources (each at its NR energy)

XeBC nuclear efficiency plots (WIP)

NR calibration sources

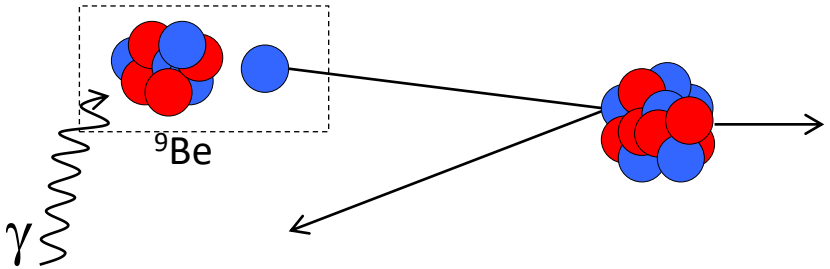
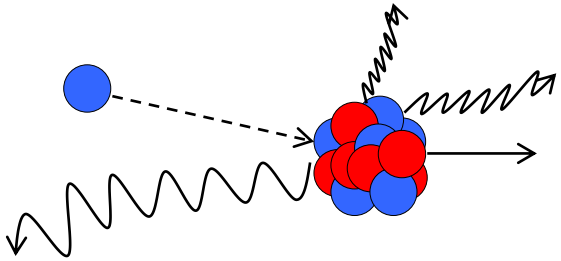


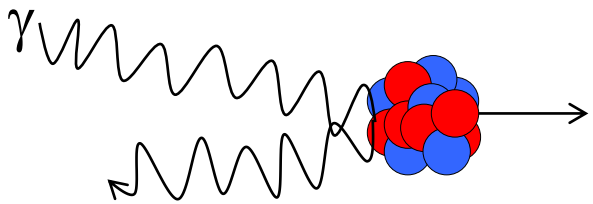
Photo-neutrons (> 500 eV E_r)

- Multi-bubble events
- Study bubble nucleation efficiency



Thermal neutrons (200 – 500 eV E_r)

- With scintillation

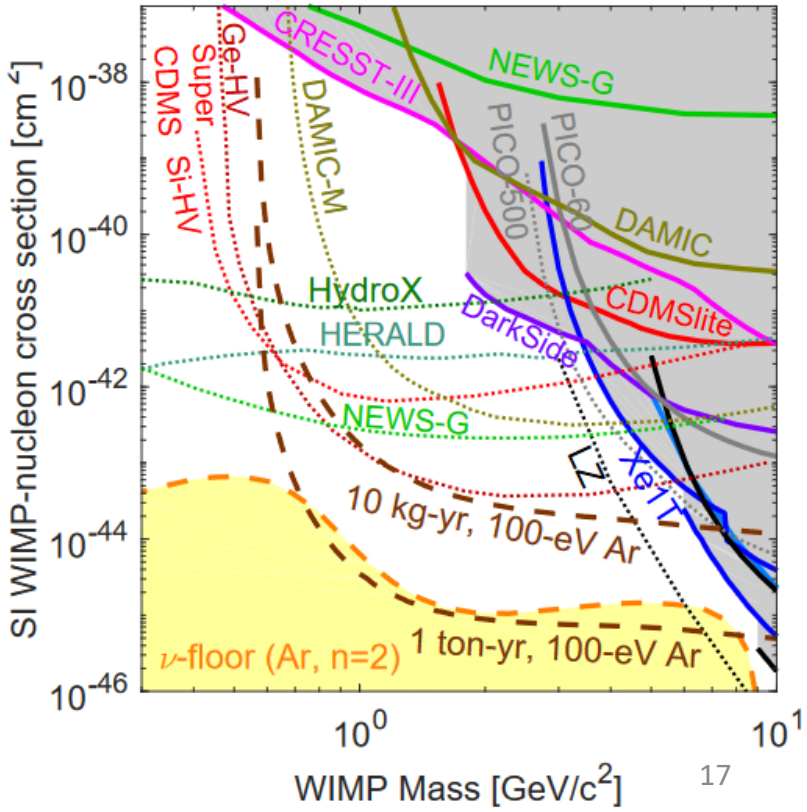


Gamma sources (< 300 eV E_r)

- Photo-nucleus Thomson scattering

Dark matter analysis

- Using background data during ER calibration
- Get upper limit using total number bubbles seen in data
- Compare to background model in MINOS
  - no background subtraction



Northwestern University

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- **Ken Clark**
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- Jonathan Corbett
- Austin De St Croix
- Koby Dering
- Hector Hawley
- Gary Sweeny
- Ezri Wyman

University of Alberta

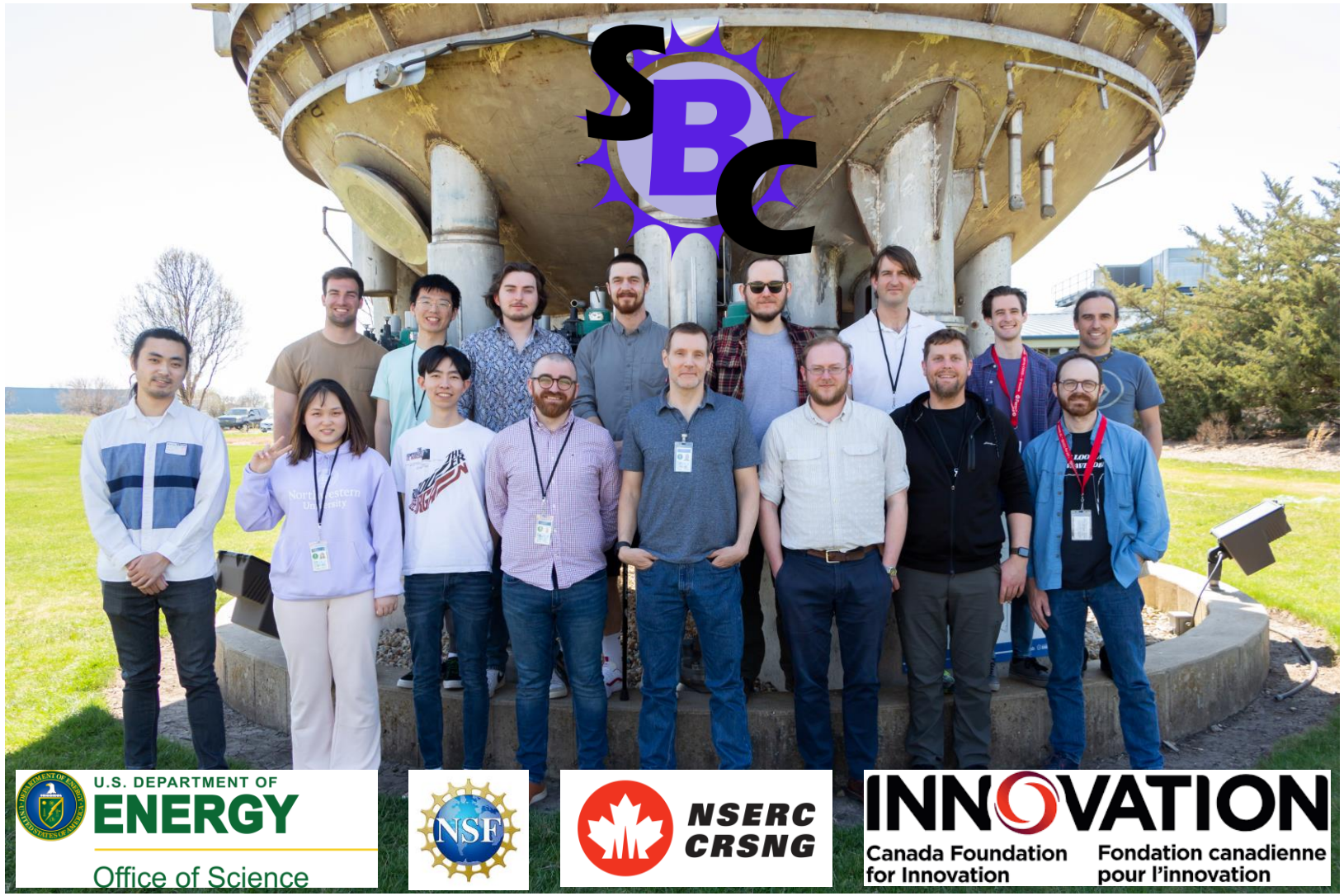
- Marie-Cécile Piro
- Carsten Krauss
- Mitchel Baker
- Daniel Durnford
- Youngtak Ko

SNOLAB

- Jeter Hall
- Alex Claveau

TRIUMF

- Pietro Giampa



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- Orin Harris

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Universidad Nacional Autónoma de México

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- Ernesto Alfonso-Pita

Drexel University

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- Julian Fritz-Littman
- Noah Lamb
- Daniel Pyda

Indiana University South Bend

- Ilan Levine
- Ed Behnke
- Cody Cripe

UC Santa Barbara

- Hugh Lippincott
- Logan Joseph
- TJ Whitis
- Runze Zhang

Fermilab

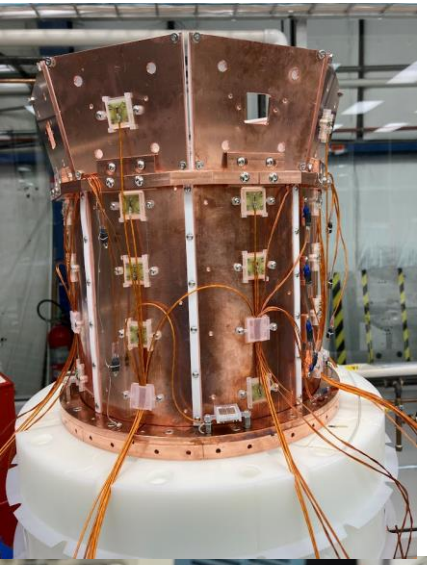
- Gray Putnam
- Vrushank Patel

PennState

- Shashank Priya

Back Up Slides

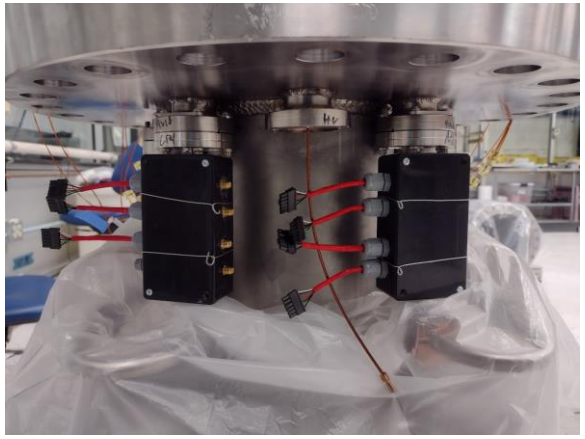
Data Acquisition System



SiPM panel

Piezo housing

Acoustic preamplifier



CAEN digitizer

SiPM amp

Back-Up: NR calibration sources

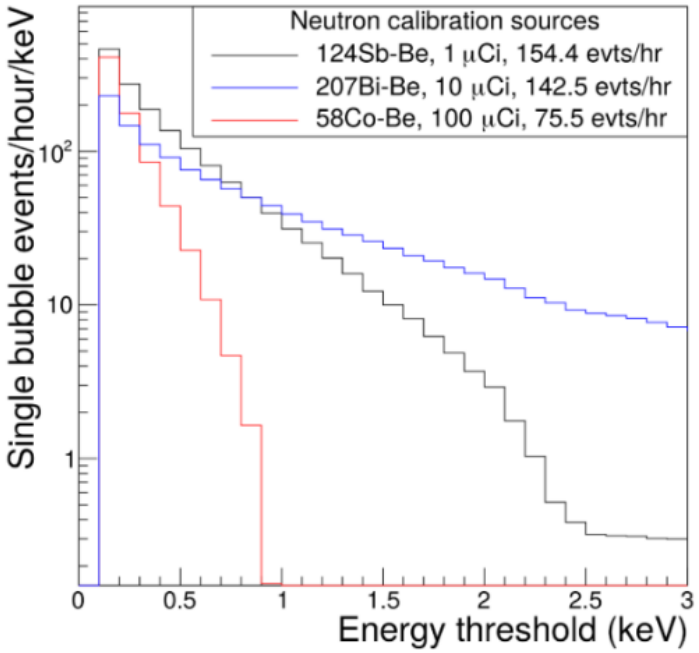
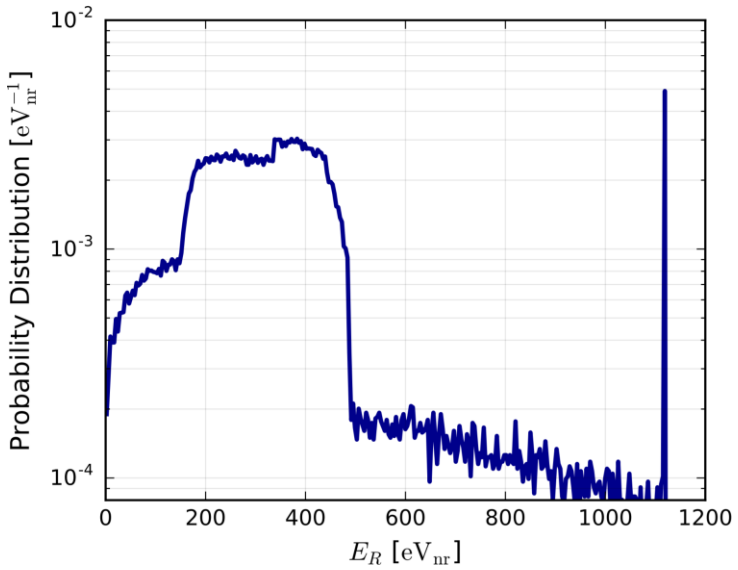
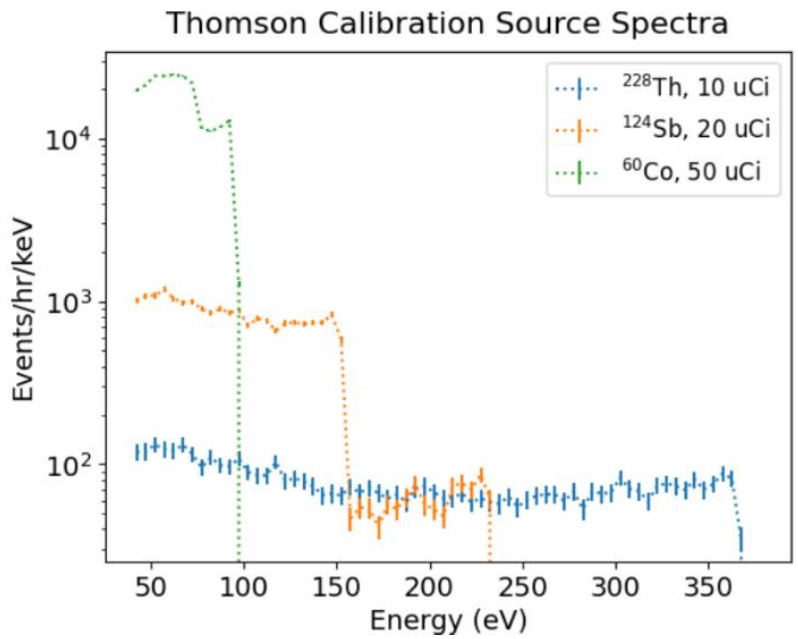


Photo-neutron scattering
(recoil > 500eV)



Neutron capture
(200eV < recoil < 500eV)



Photon-nucleus scattering
(recoil < 300eV)