# Constraining Systematics for Future Sterile Neutrino Analysis at NOvA Experiment

New Perspective Meeting

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# **Neutrino Oscillations**

#### **Neutrino Oscillations**



- Neutrinos produced in one flavor state change its flavor during its travel across the distance.
- $\nu_l$ , flavor eigenstate which is a superposition of  $\nu_i$ , mass eigenstates.

#### **Neutrino Oscillations**

- In most of the long baseline experiments, we use the  $\nu_{\mu}$  disappearance or  $\nu_e$  appearance channels to study the neutrino oscillation parameters.
- As an example, in two flavor approximation  $\nu_{\mu}$  disappearance probability is defined as:

$$P(\nu_{\mu} \rightarrow \nu_{\mu}) \approx 1 - \left(\sin^2 2\theta_{23}\right) \sin^2 \left(\frac{\Delta m_{23}^2}{E_{\nu}}\right)$$

- mixing angle determines the magnitude of oscillations.
- mass splitting determines the frequency of oscillations.



#### Is Three Flavor Picture Enough?

• Several anomalous results observed by various experiments could suggest a possible explanation beyond the active three-flavor oscillations.



• LSND observed a  $3\sigma$  excess above the expected beam background.

- One possible solution is adding an extra neutrino flavor.
- Any additional neutrino flavor should not interact with the known forces (not even through weak interactions).
- More than one sterile neutrino is possible, but the minimal solution uses the 3+1 model.
- This leads to adding an extra dimension to the PMNS mixing matrix, also leading to an additional oscillation frequency  $\Delta m^2_{41}$ .

# **NOvA Experiment**

#### **NOvA Experiment**

#### NuMI Beam



- 120 GeV protons from the Fermilab Main Injector strike the target to produce secondary particles.
- Two focussing horns focus the secondary particles that decay into the decay tunnel to produce the  $\nu(\bar{\nu})$  beam.



- NOvA is a long baseline experiment with two functionally identical liquid scintillator detectors.
- The Near Detector is placed 100 m underground at 1 km from the source, and the far detector at 810 km on the surface from the near detector.

# Sterile Neutrino

#### Sterile Neutrino at NOvA: Neutral Currents

• Neutral Current Disappearance gives a clean measurement of 3+1 oscillations because of their flavor independency.



- Oscillations begin to manifest at ND for  $\Delta m_{41}^2 > 0.5 \text{eV}^2$ .
- Highlighted text is the short baseline approximation.

$$P(\nu_{\mu} \rightarrow \nu_{5}) \approx \frac{1 - \cos^{4}\theta_{14}\cos^{2}\theta_{34}\sin^{2}2\theta_{24}\sin^{2}\Delta_{41}}{-\sin^{2}\theta_{34}}\sin^{2}2\theta_{23}\sin^{2}\Delta_{31}$$
$$+ \frac{1}{2}\sin\delta_{24}\sin\theta_{24}\sin2\theta_{23}\sin\Delta_{31}$$

• Sensitivity to  $\sin^2 \theta_{34}$  at FD NC can be measured independent of  $\sin^2 \theta_{24}$ .

#### Sterile Neutrino at NOvA: $u_{\mu}$ disappearance

• Any additional  $\nu_{\mu}$  disappearance above the expected 3-flavor oscillation can be manifested as sterile neutrino.



- Highlighted text is the FD oscillation intermixed with the 3-flavor oscillations.
- Charged Current  $\nu_{\mu}$  is sensitive to the  $\theta_{24}$  at both ND and FD.

# Motivation: Constraining Systematics

#### Sterile Neutrino at NOvA

- NOvA 2022 Sterile Neutrino mode results showing a leading limit on  $\sin^2 \theta_{24}$  at high  $\Delta m_{41}^2$ .
- On one hand, the low  $\Delta m_{41}^2$ region is driven by the FD data and is statistically limited.
- On the other hand, at high  $\Delta m_{41}^2$  region where sensitivity is driven by ND is systematically limited.



**Figure 2:** NOvA's 90 % confidence limits in (a)  $\sin^2 \theta_{24}$  vs  $\Delta m_{41}^2$  space with other allowed regions and exclusion contours.

#### Sterile Neutrino at NOvA



Figure 3: Sensitivity Contour (at 90% CL) for  $\sin^2 \theta_{24}$  vs  $\Delta m_{41}^2$ 

• We are taking more and more data, which improves the statistics, but with more statistics, we also need to deal with the systematics.

- The figure on the left shows the Sensitivity Contour (at 90% CL) for  $\sin^2 \theta_{24}$  for different systematic groups.
- We can see that the cross-section and flux systematics are the dominant ones, and the future analysis includes constraining the systematics.

We used a new approach to implement the ND NC sample, where instead of using the sample as a whole, we divided it into subsamples based on the number of prongs associated with the event.

- Single prong Sample : the events with single prong are mostly enriched with the SIS, QE and dominated by Res.
- 2 and 3 Prong Sample : This sample is highly enriched in Res but has a contribution from SIS interaction as well.
- 4 Prong Sample : DIS starts appearing, but this region is dominated by SIS events.
- >4 Prong Sample : Once we have more than 4 prongs, the DIS interaction highly dominates the events.

### **Results and Conclusion**

#### **Results and Conclusion**



- The distribution in light blue shows the uncertainty at FD without any constraint from the ND.
- Dark Blue distribution represents the FD uncertainty knowing the information about the ND without splitting.
- Pink distribution represents the FD uncertainty with additional information with ND splitting.

- Conditional uncertainty distributions show better constraints on the cross-section uncertainties.
- This split sample approach will allow us to disentangle the signal and systematic effects and help improve the sensitivity at higher  $\Delta m^2_{41}$  region.
- More studies are underway, including zero horn current and  $\nu$ -on-e studies to improve the flux systematic uncertainties.

# **NOvA Collaboration**



### Thank You

# **Backup Slides**

### Fraction with each prong:: True NC only (ND NC CVN >0.1)

pngs	Coh	DIS	SIS	QE	Res
1	0.029( <b>0.071</b> )	0.055 ( <b>0.046</b> )	0.236 ( <b>0.254</b> )	0.147 ( <b>0.091</b> )	0.531 ( <b>0.535</b> )
2	0.016 (0.038)	0.051 ( <b>0.051</b> )	0.315( <b>0.326</b> )	0.039 ( <b>0.022</b> )	0.577 ( <b>0.561</b> )
3	0.005( <b>0.011</b> )	0.103( <b>0.115</b> )	0.384( <b>0.393</b> )	0.022( <b>0.011</b> )	0.484( <b>0.467</b> )
4	0.001(0.004)	0.220( <b>0.247</b> )	0.414(0.418)	0.013(0.006)	0.350( <b>0.322</b> )
5	0 (0.001)	0.382 (0.400)	0.379 ( <b>0.387</b> )	0.006 ( <b>0.004</b> )	0.230 ( <b>0.206</b> )
6	0 (0.001)	0.527 ( <b>0.534</b> )	0.322 ( <b>0.328</b> )	0.001 (0.001)	0.146 (0.133)
7	0 (0)	0.644 ( <b>0.641</b> )	0.261 ( <b>0.278</b> )	0 (0.001)	0.091 (0.077)
8	0	0.725 ( <b>0.727</b> )	0.198 ( <b>0.278</b> )	0 (0)	0.075 ( <b>0.066</b> )
9	0	0.751 ( <b>0.763</b> )	0.190 ( <b>0.210</b> )	0	0.058 ( <b>0.026</b> )

Table 1: Fraction of each interaction with # of prongs for CVN>0.1

- The table shows the different interaction fractions with loose CVN scores.
- Losening the CVN score reduces the fraction of QE events and increases the DIS and Res fractions.

pngs	Coh	DIS	SIS	QE	Res
1	0.056 ( <b>0.151</b> )	0.048 ( <b>0.054</b> )	0.149 ( <b>0.172</b> )	0.280(0.147)	0.465( <b>0.473</b> )
2	0.048 (0.113)	0.058 ( <b>0.039</b> )	0.206( <b>0.216</b> )	0.046 (0.024)	0.640( <b>0.605</b> )
3	0.018 (0.044)	0.150 ( <b>0.123</b> )	0.219 ( <b>0.256</b> )	0.030 ( <b>0.014</b> )	0.580 ( <b>0.559</b> )
4	0.008 (0.004)	0.403 ( <b>0.247</b> )	0.206 ( <b>0.418</b> )	0.019 ( <b>0.006</b> )	0.362 ( <b>0.322</b> )
5	0.003 (0.001)	0.655 ( <b>0.400</b> )	0.155 ( <b>0.987</b> )	0.010 (0.004)	0.175 ( <b>0.206</b> )
6	0.000 (0)	0.792 ( <b>0.534</b> )	0.122 ( <b>0.328</b> )	0.001 (0.001)	0.081 ( <b>0.133</b> )
7	0 (0)	0.830 ( <b>0.641</b> )	0.114 (0.278)	0 (0.001)	0.054 ( <b>0.077</b> )
8	0 (0)	0.841 ( <b>0.727</b> )	0.110 (0.205)	0 (0)	0.047 ( <b>0.066</b> )
9	0 (0)	0.879 ( <b>0.763</b> )	0.051 ( <b>0.210</b> )	0 (0)	0.069 ( <b>0.026</b> )

Table 2: Fraction of each interaction with # of prongs for CVN>0.98

- To see the different interactions, DIS has been split into two categories (DIS and SIS) based on the  $Q^2$  and W value.
- It appears that, beyond five prongs, most events fall into the DIS category.