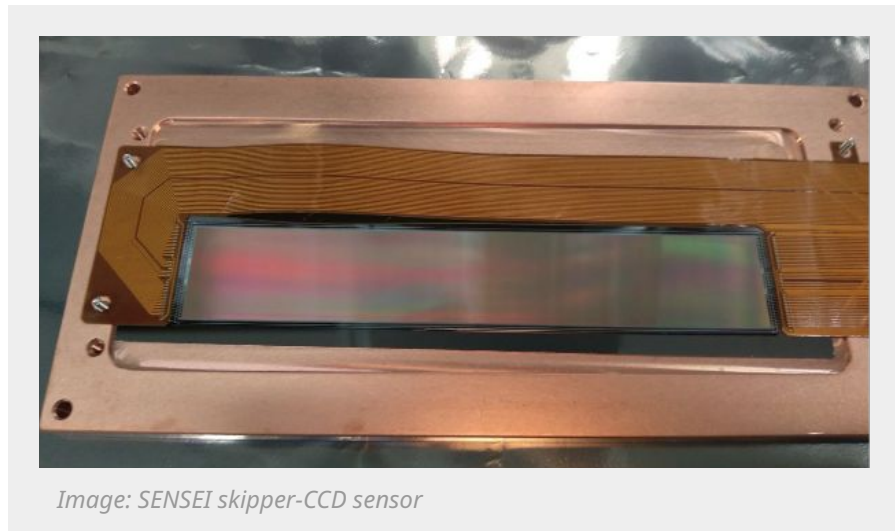


Measuring the universe one electron at a time with Skipper-CCDs

Ana Martina Botti*
Fermilab Users Meeting 2024
July 11, 2024

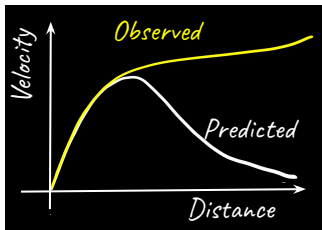


* Fermi National Accelerator Laboratory and Kavli Institute for Cosmological Physics, University of Chicago
abotti@fnal.gov

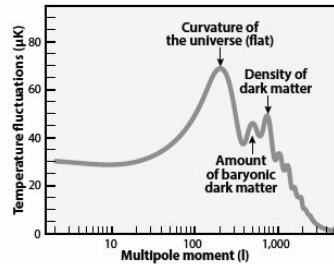
Dark-matter

Overwhelming evidence from cosmology and astrophysics for the existence of dark matter...

Galaxy gas rotation

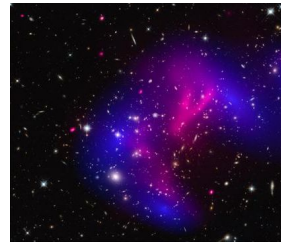


CMB



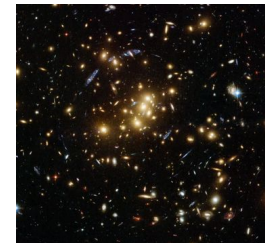
Astronomy: Roen Kelly, after Wayne Hu

Cluster collision



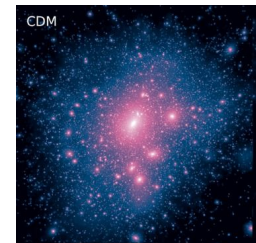
NASA, ESA, STScI, CXC. Copyright

Gravitational lensing



NASA, ESA, M.J. Jee, and H. Ford (Johns Hopkins University)

Structure formation

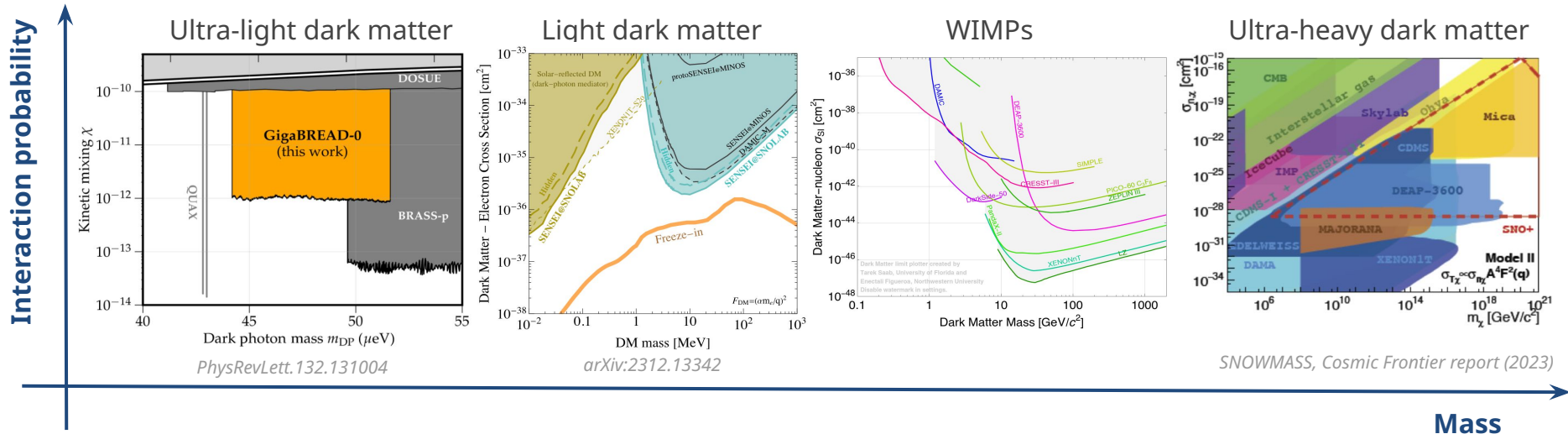


Mark Vogelsberger 2020

More than **84% of the matter in the universe is dark...** and we have no idea what dark is.

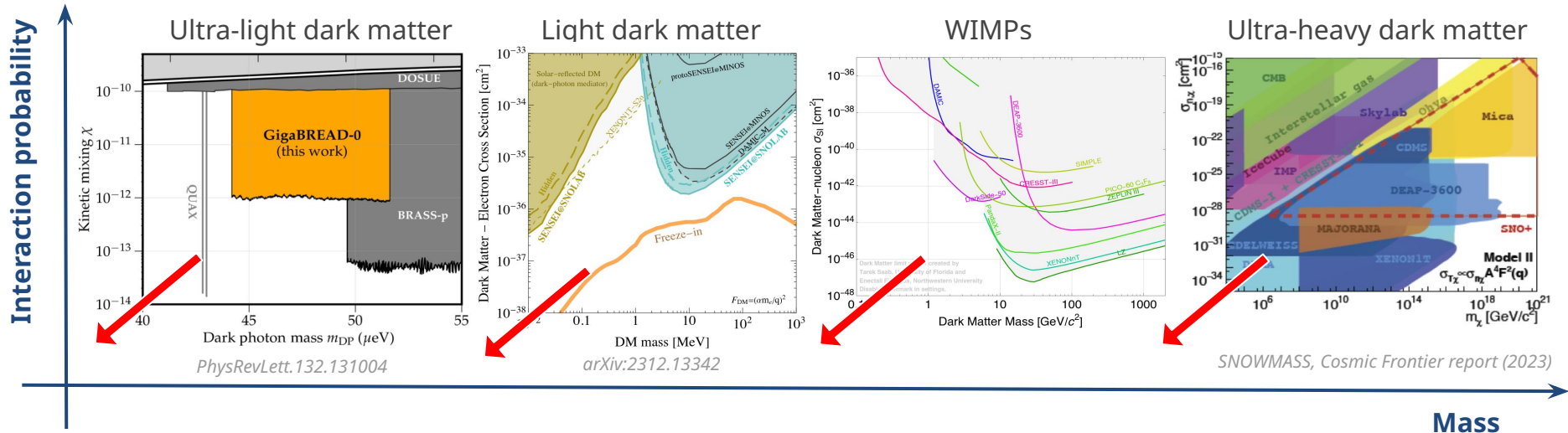
The **standard model of particles** cannot explain these phenomena... dark matter is promising gateway towards **physics beyond** the standard model

But... we have some ideas of what is not!



> 80 orders of magnitude in **mass** and **thousands** of **theoretical models** to explore
 Drawing a line in these plots takes years of work

But... we have some ideas of what is not!

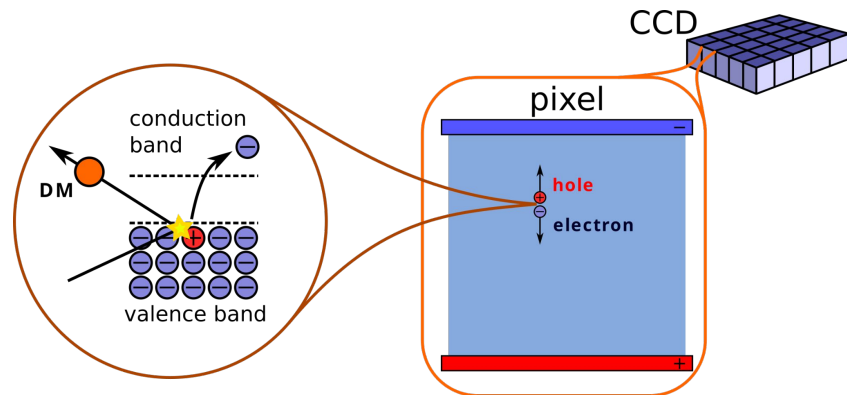


We need **bigger** and **more sensitive** detectors to push these limits

We all want to **find dark matter**. We are **running** a race! New detectors, new runs, new analysis every year.

Skipper-CCDs: a (rather) new approach

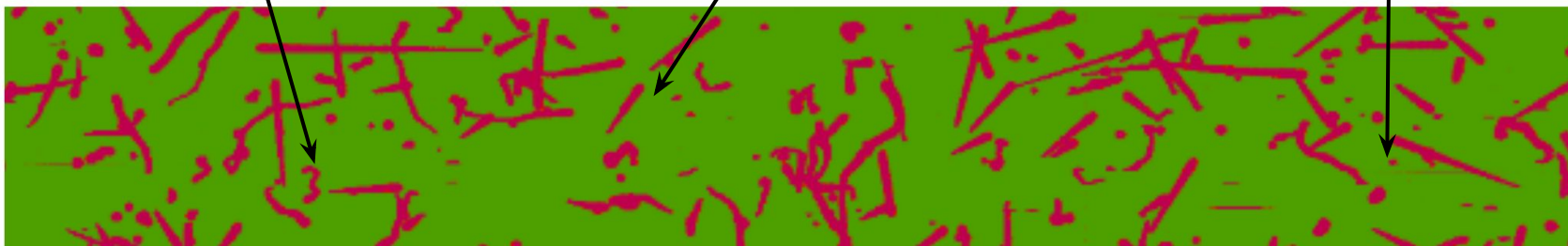
New generation Charge Couple Devices (CCD)
 Pixeled silicon detectors
 S. Holland (LBNL) Design
 First demonstrated in **Fermilab** (Tiffenberg et al.)



Electrons (curly tracks)

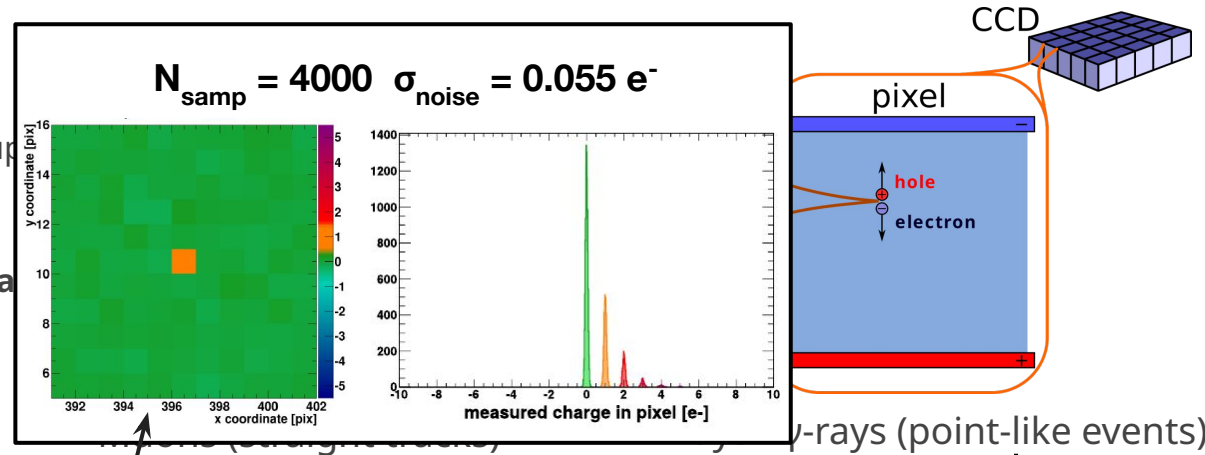
Muons (straight tracks)

X-rays / γ -rays (point-like events)



Skipper-CCDs: a (rather) new approach

New generation Charge Coupled
 Pixelated silicon detectors
 S. Holland (LBNL) Design
 First demonstrated in **Fermilab**



SENSEI to push sensitivities...

First experiment using **skippers** for dark matter
40 g of mass (19 CCDs) after 3 commissioning **campaigns**

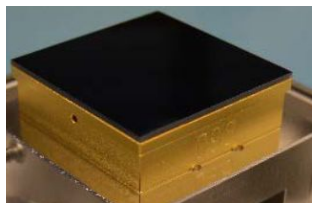


SENSEI to push sensitivities...

First experiment using **skippers** for dark matter
40 g of mass (19 CCDs) after 3 commissioning **campaigns**

Lowest dark current on
silicon detectors
(1 e-/pixel every 200
years)

Best constraints
in Light dark
matter on
2023-2024!!!!



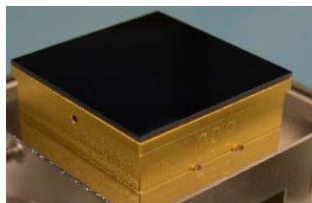
LSST focal
plane CCD
DC ~ 1000
e⁻/pix/day



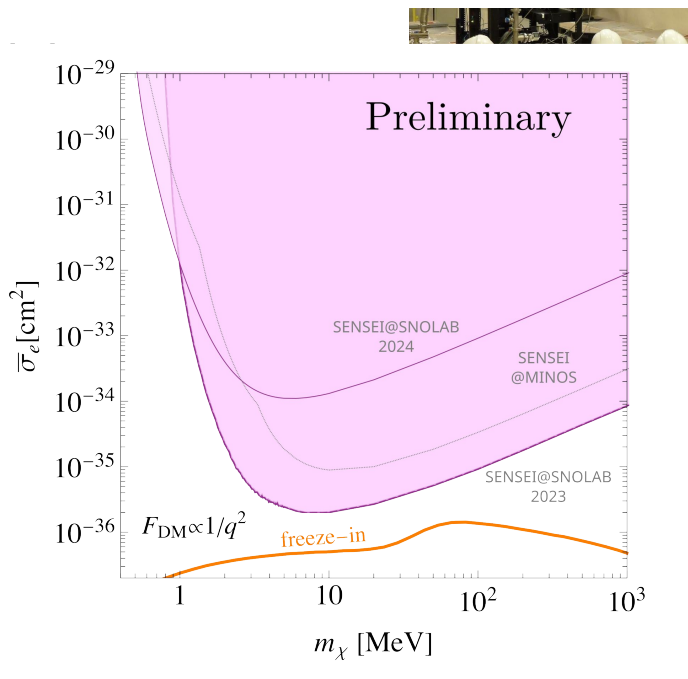
SENSEI to push sensitivities...

First experiment using **skippers** for
40 g of mass (19 CCDs) after **3** com

**Lowest dark current on
 silicon detectors
 (1 e-/pixel every 200
 years)**



LSST focal
 plane CC
 DC ~ 10C
 e-/pix/day

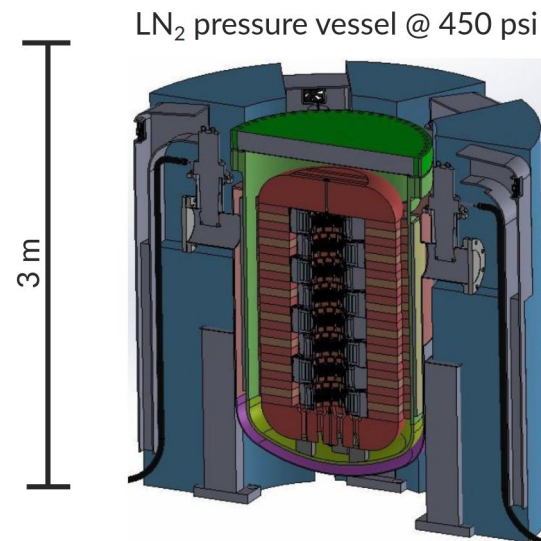
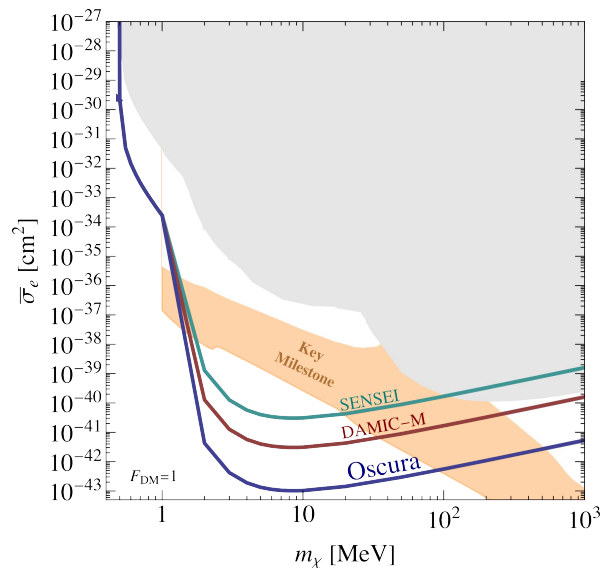


**+ 2/3 of URA
 Tollestrup
 award**



OSCURA to make it **BIG**

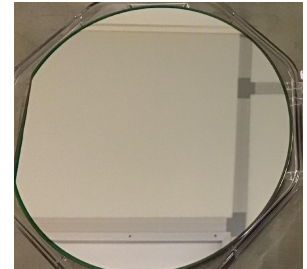
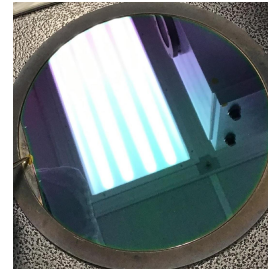
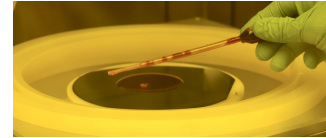
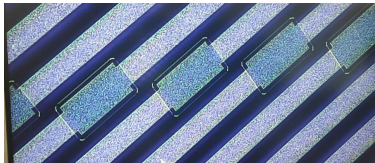
10 kg of skipper-CCDs



OSCURA to make it **BIG**

10 kg of skipper-CCDs

New low-background compact packages

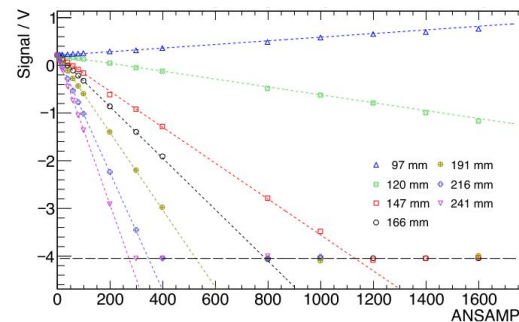
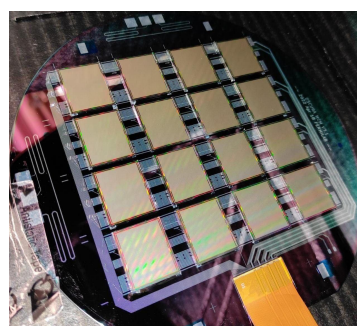
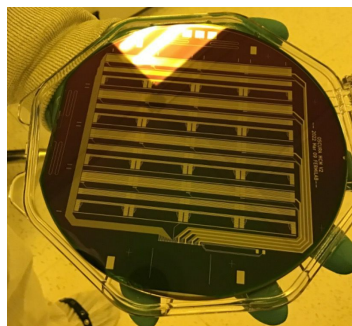


OSCURA to make it **BIG**

10 kg of skipper-CCDs

New low-background compact packages

Silicon multi-chip modules to scale 2 orders of magnitude in mass



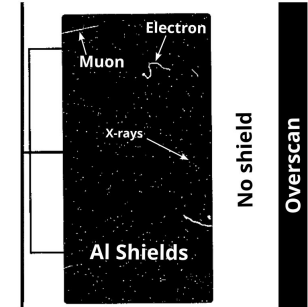
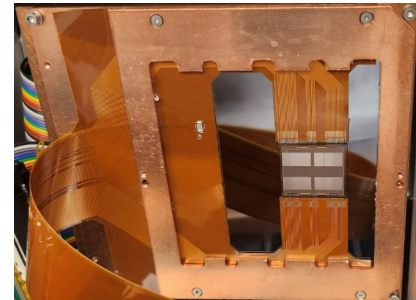
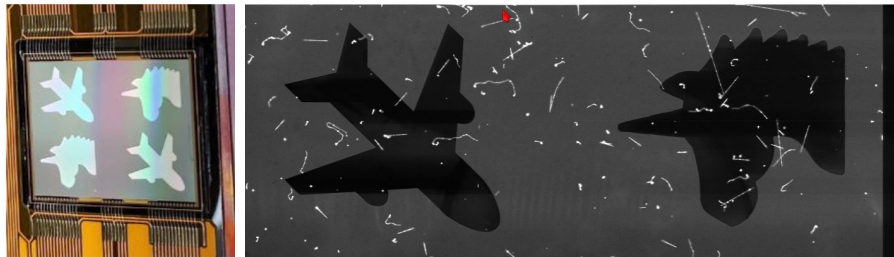
OSCURA to make it **BIG**

10 kg of skipper-CCDs

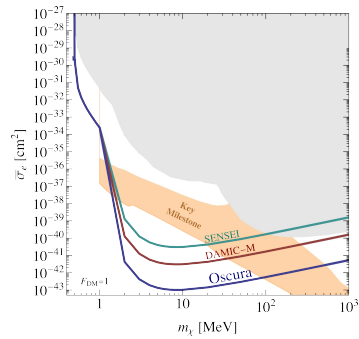
New low-background compact packages

Silicon multi-chip modules to scale 2 orders of magnitude in mass

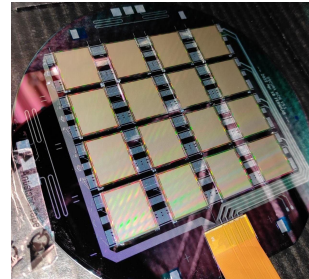
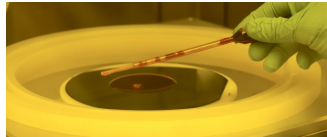
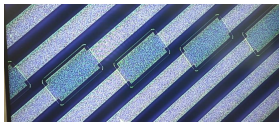
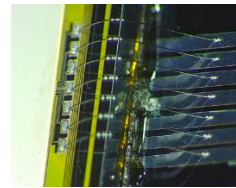
Aluminum shield to reduce the backgrounds orders of magnitude



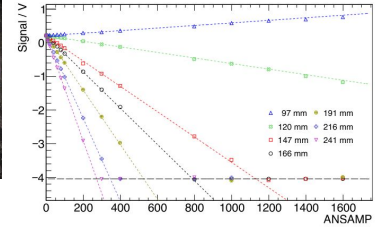
OSCURA to make it BIG



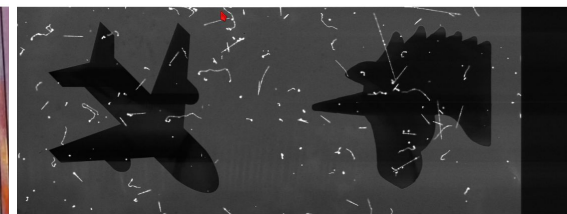
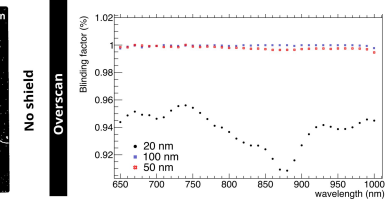
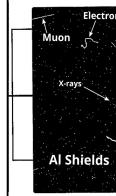
+ 1/3 of URA Tollestrup award



Silicon multi-chip interconnect



num light shield



And more... exploring new ideas:

The current leading **light dark-matter** direct detection **technology** comes from **astronomical instruments**.

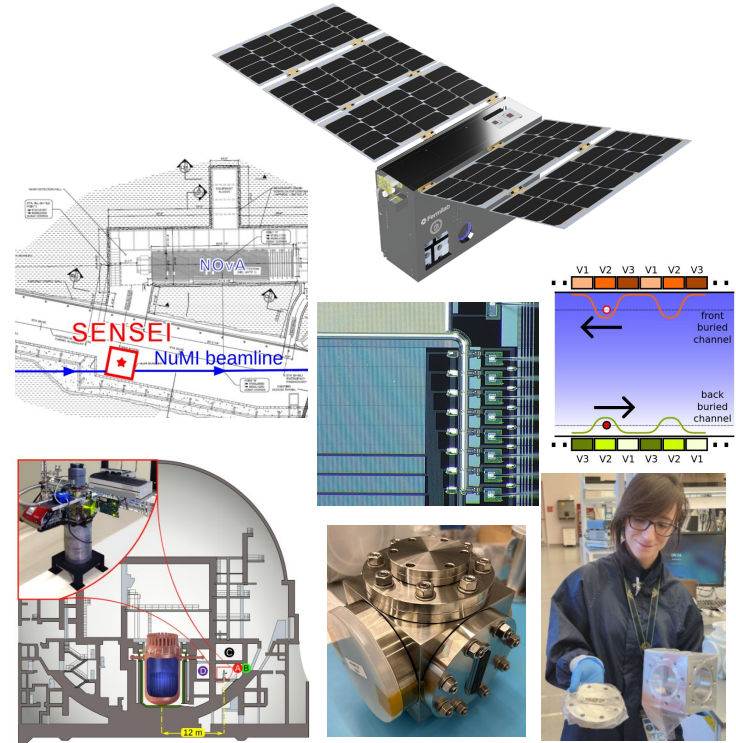
(Near)Far-future experiments will probably not come from HEP.

Maybe **CM**, **Quantum**, **weird thermodynamics**, etc.

This is the time to explore new ideas and foster cross-disciplinary collaborations:

- DarkNESS
- MiliCharge @ beams and reactors
- Dual-side CCD
- MAS-CCD
- Nuclear reactor neutrinos
- Quantum skipper-CCD camera

The next generation of **astronomical instruments** might come from the current leading **light dark-matter** direct detection **technology**



Many thanks to:

Fermilab, ANL and SNOLAB technical staff

CCD Group @ Fermilab:

- J. Estrada
- G. Fernandez Moroni
- J. Tiffenberg
- B. Cervantes
- N. Saffold
- S. Uemura
- C. Chavez

+ our many collaborators/visitors from UBA, UNS, UNC, etc.

SENSEI and OSCURA collaborations

URA Tollestrup selection committee and URA board and personnel

Family and friends (of course)

