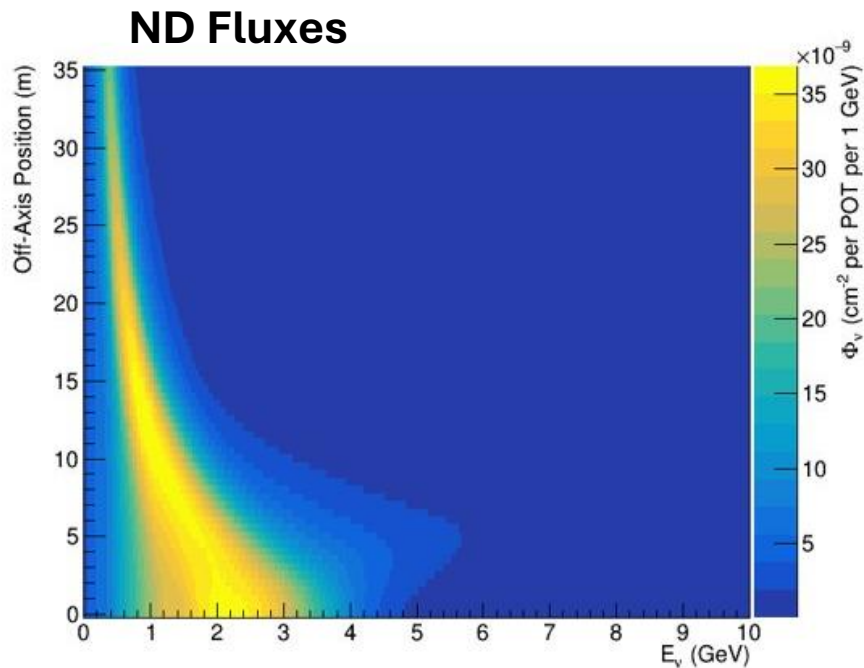


# DUNE-PRISM Gaussian Fluxes

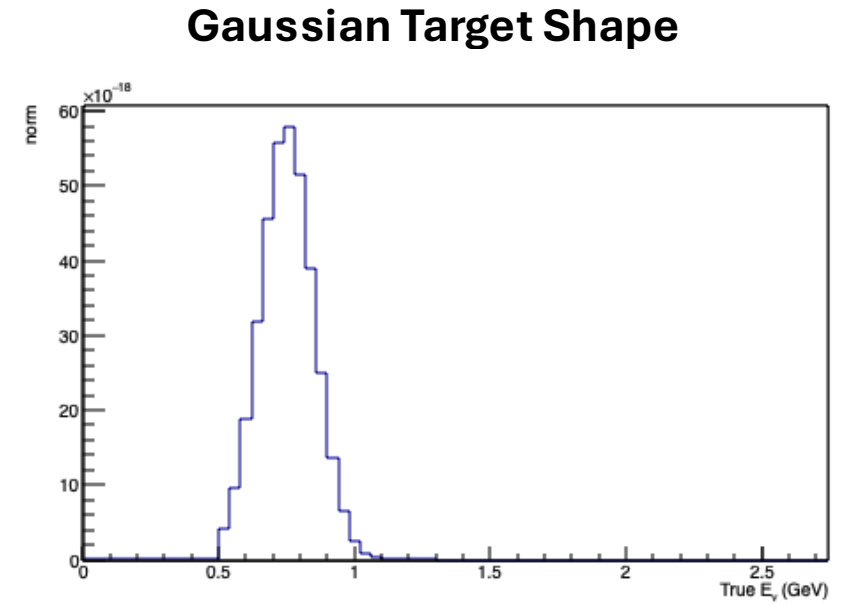
Ciaran Hasnip  
DUNE-PRISM Meeting  
08/08/2024

# Matching a Gaussian Target

- Match the ND  $\nu_\mu$  fluxes to a narrow gaussian shape
- Just solving a **linear algebra problem** with the flux
- Mathematically, this is  $\mathbf{Nc} = \mathbf{F}$  – we solve for  $\mathbf{c}$ !

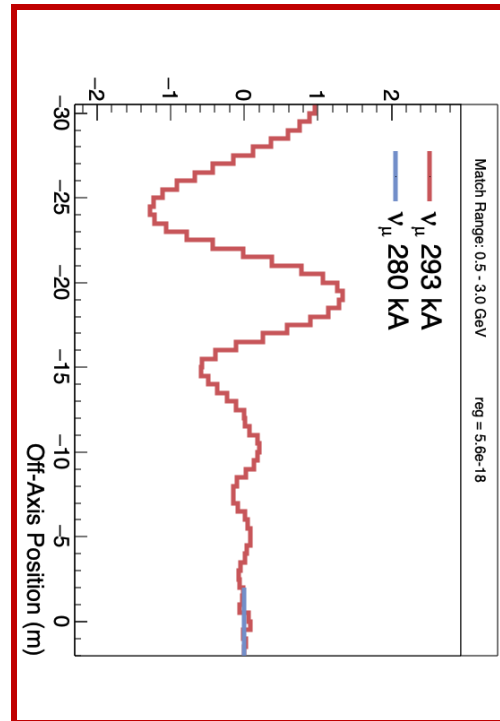
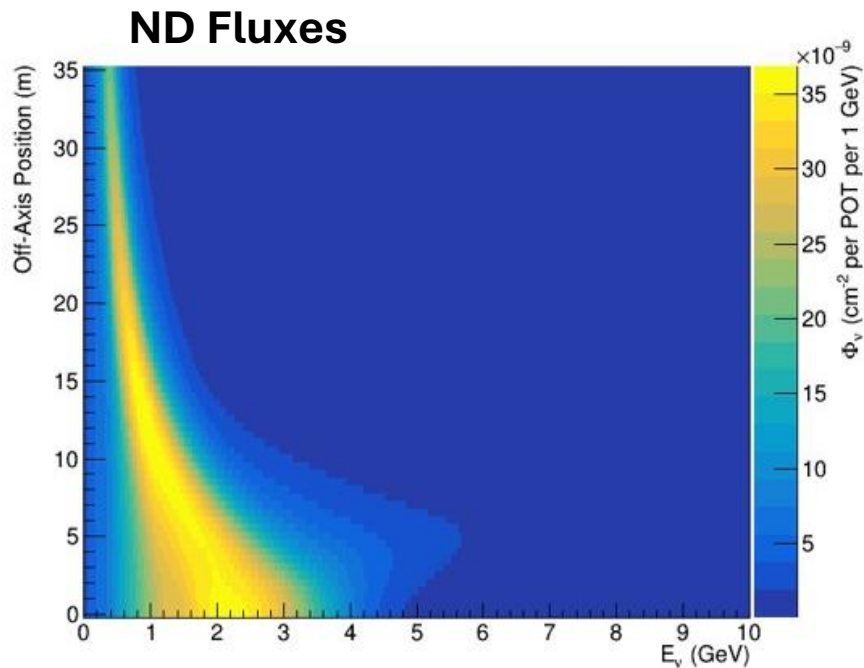


**some vector,  $\mathbf{c}$**  =

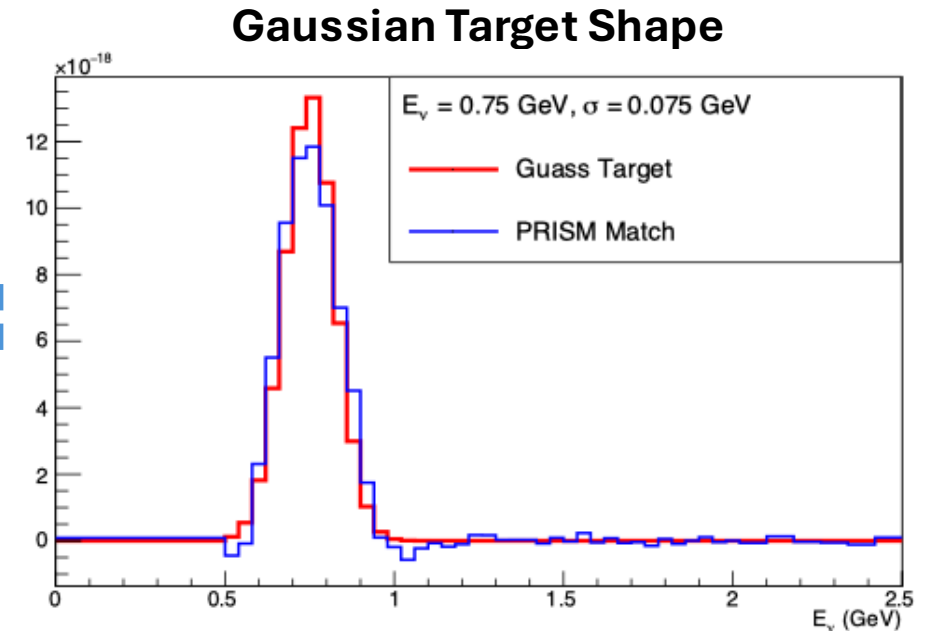


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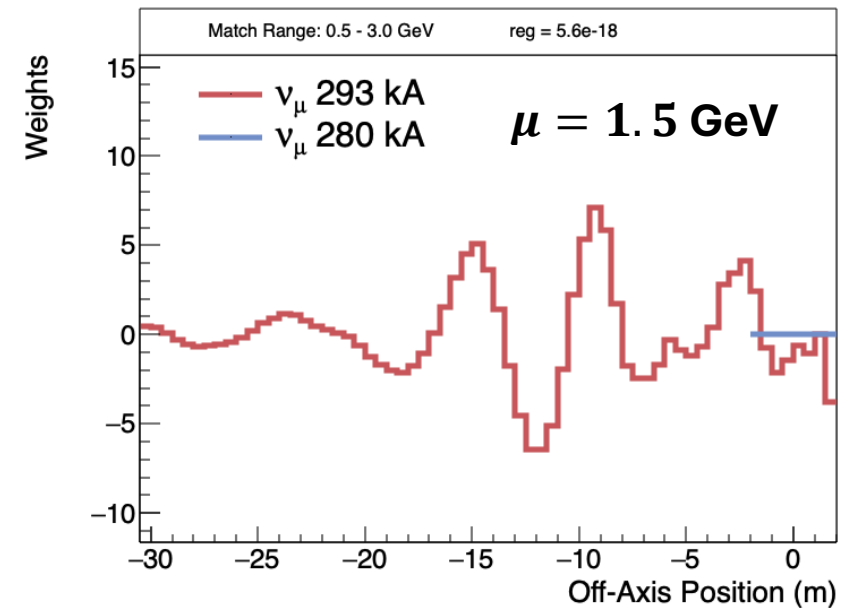
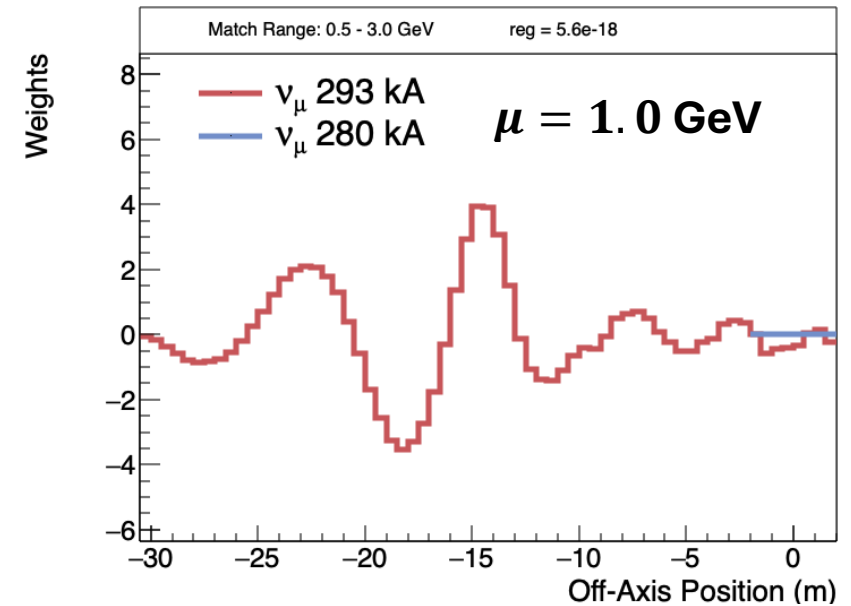
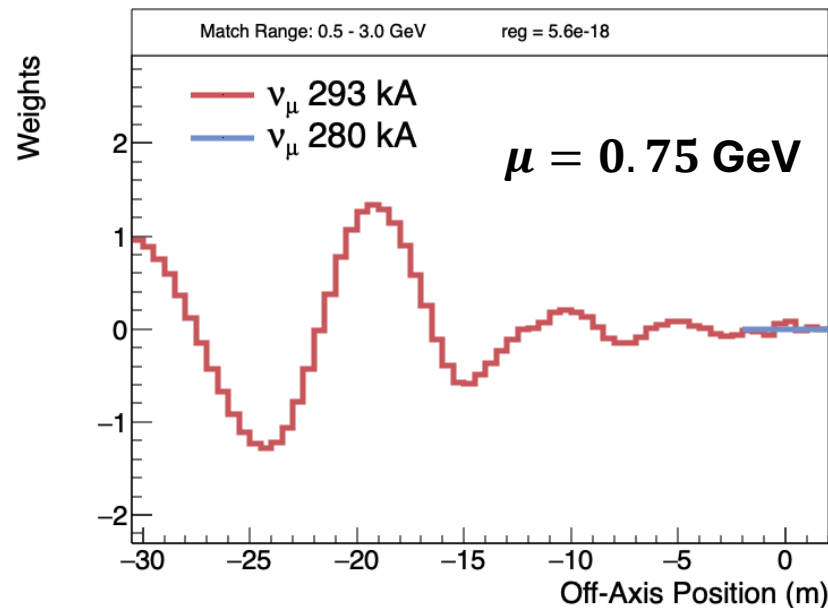


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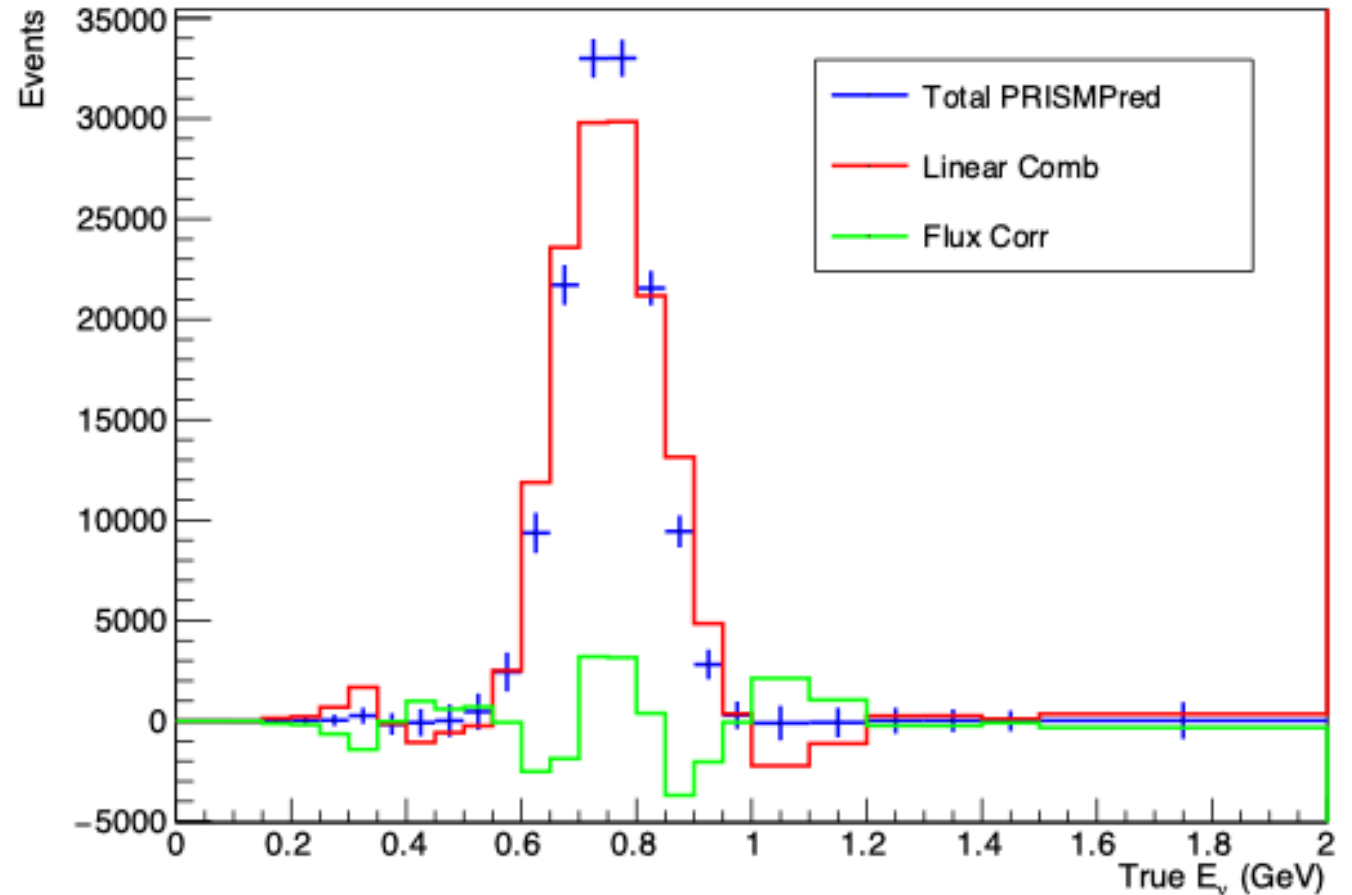
# Matching a Gaussian Target

- Low energy mean – higher coefficients at far off axis position
- High energy mean – larger coefficients at more on-axis positions



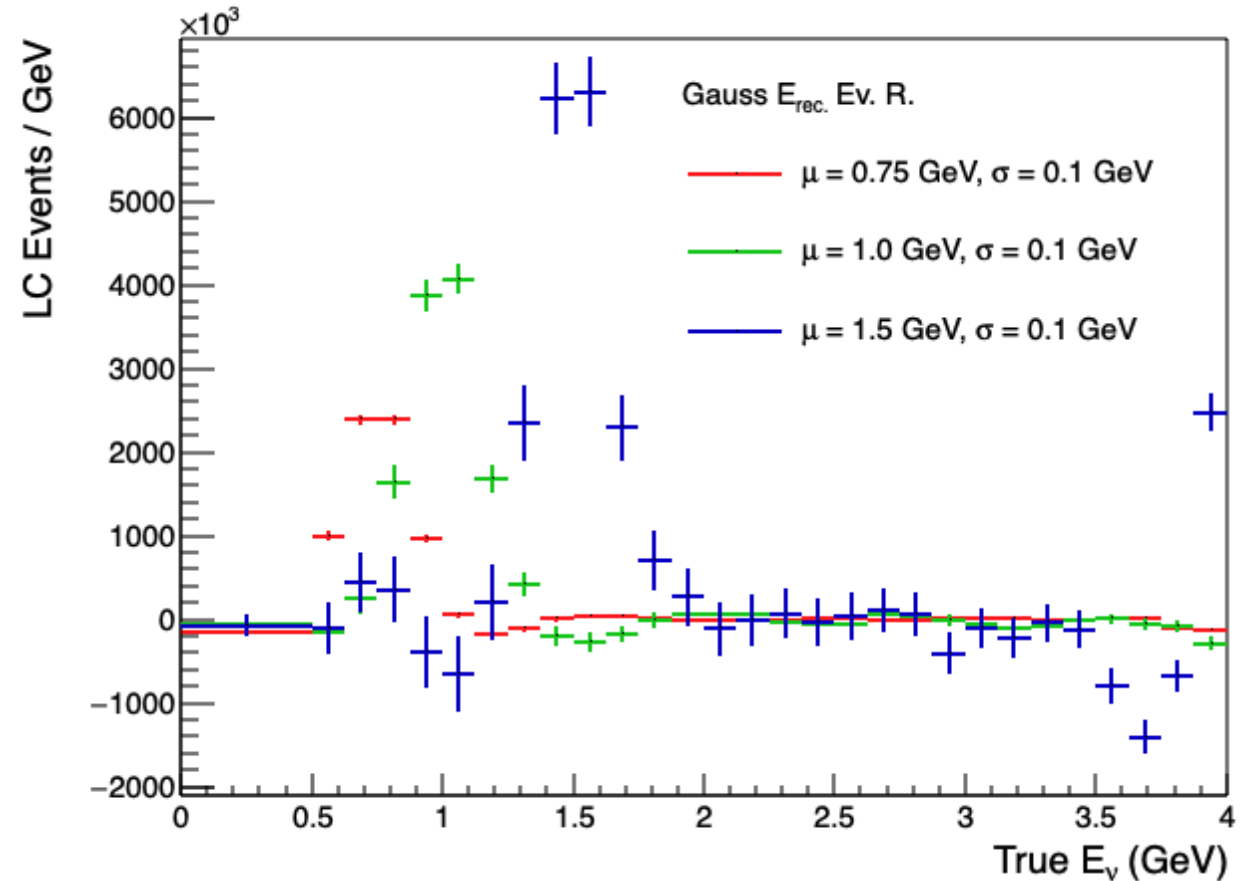
# Flux Mismatch Correction

- Like in the oscillation analysis – the coefficient calculation is **not perfect**
- Correct with a **flux mismatch correction**
- Calculated by **weighting on-axis ND MC** by flux-mismatch **residual**
- Seems to **work okay**



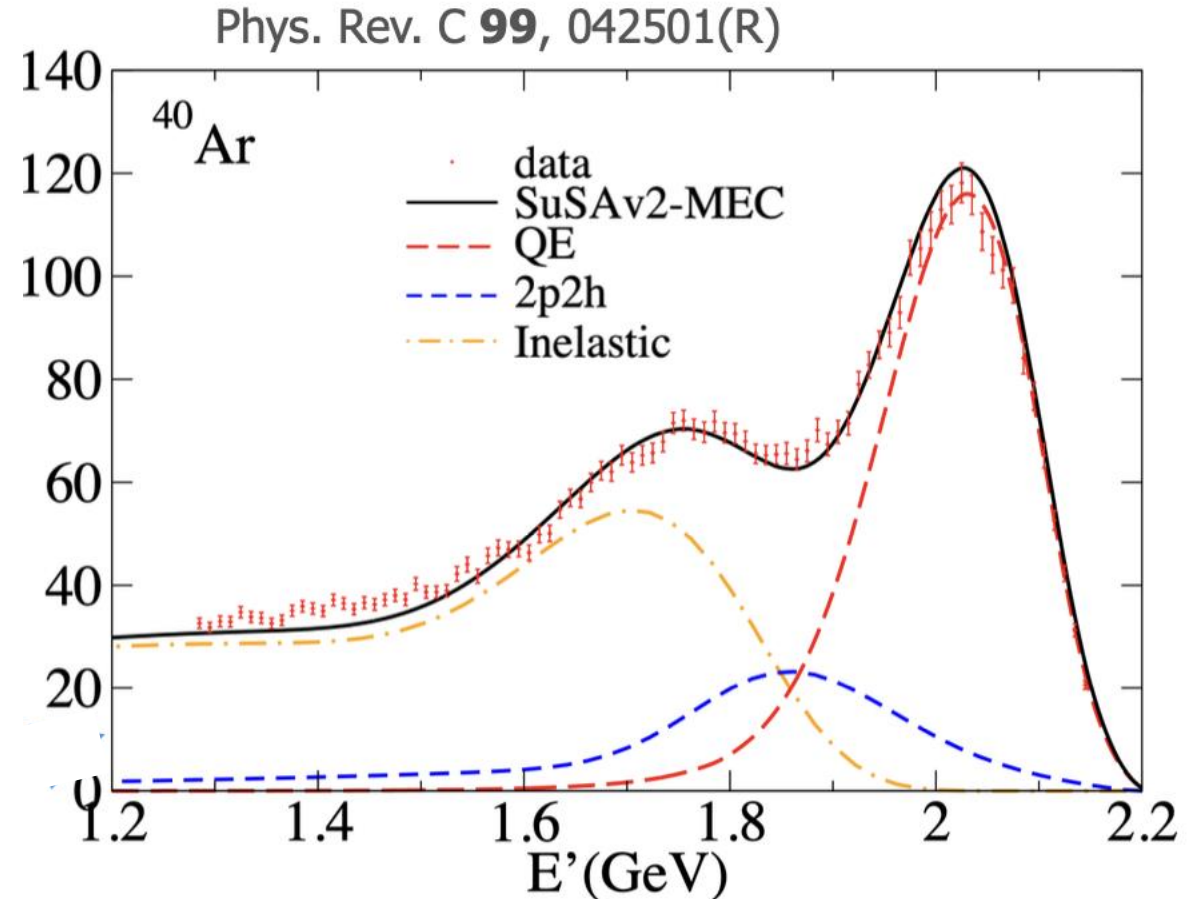
# Removing Detector Effects

- In a cross section analysis probably want to **remove detector effects**
- Can **unfold ND data** and correct for **efficiencies** before linear combination
- Here shows unfolded gaussian fluxes for  $E_{\nu}^{rec} \rightarrow E_{\nu}^{true}$



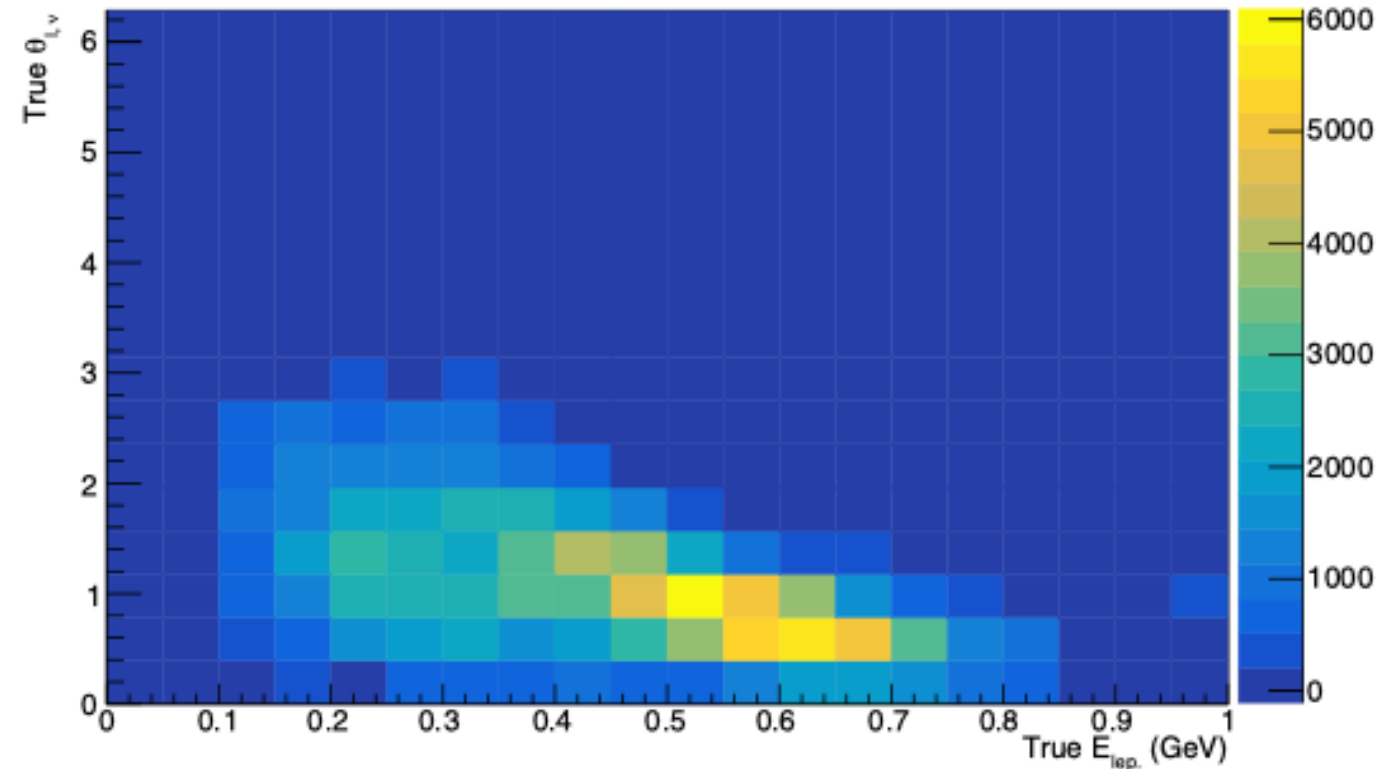
# Different Analysis Axes

- Inspiration from **electron scattering** and see [report](#) by Amir Gruber
- Known **incoming electron energy**
- Measure outgoing scattered electron energy at specific **scattering angle**
- Enables **separation** of nuclear effects
- Gaussian neutrino flux gives a **pseudo-known** incoming **neutrino energy**
- Can measure outgoing lepton energy and angle from neutrino interaction



# Different Analysis Axes

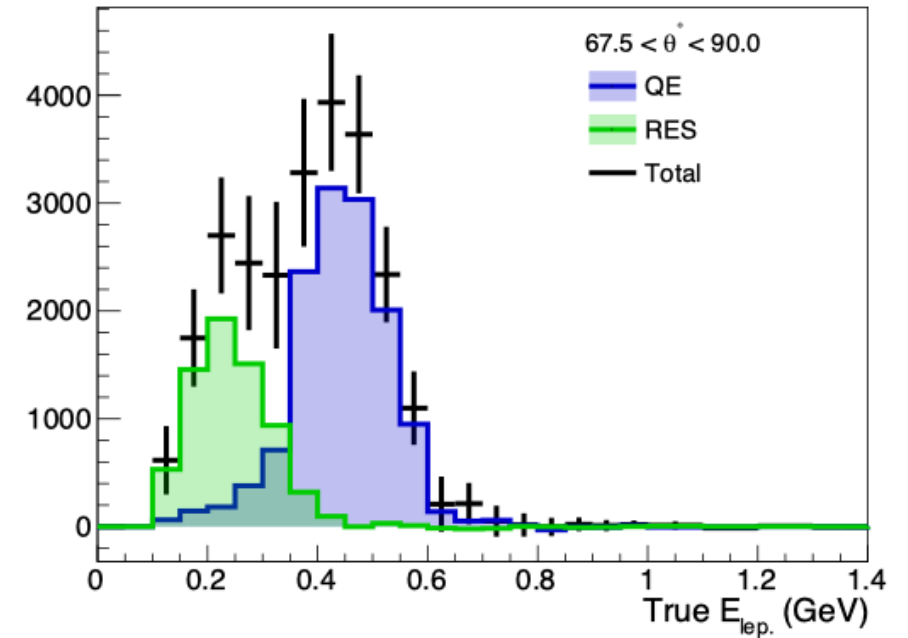
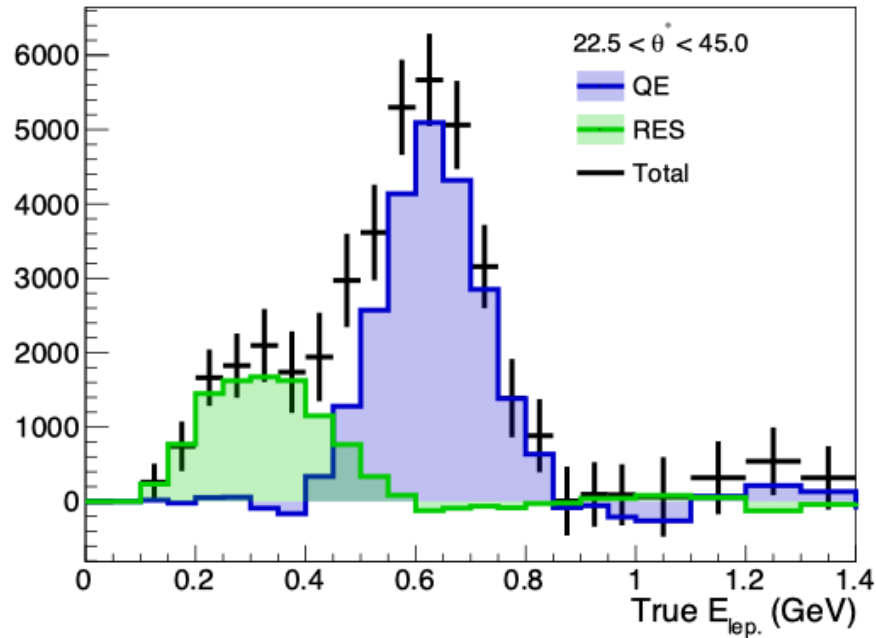
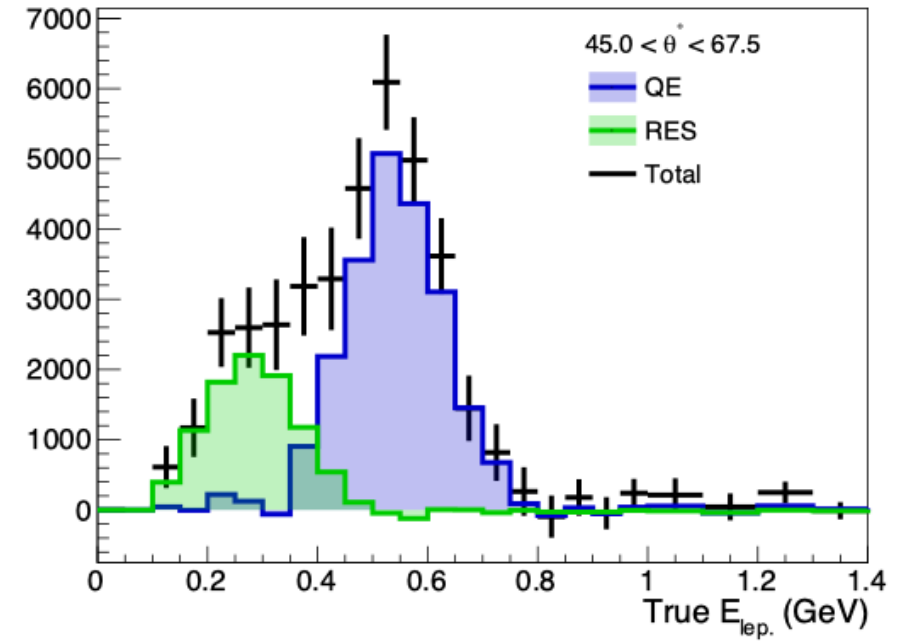
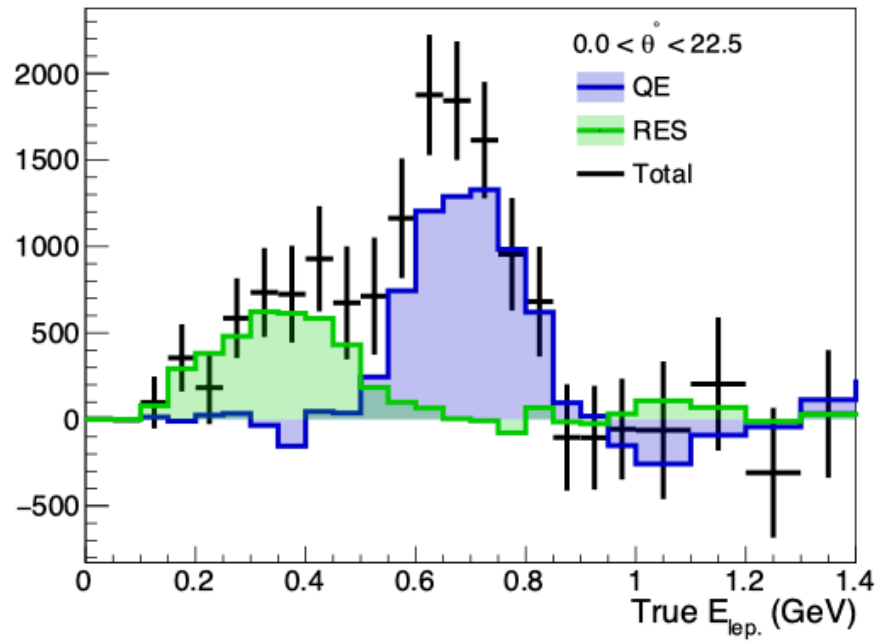
- Follow the **same procedure** for gaussian linear combination
- Analysis axis is now **Lep-  
NuTheta vs Elep**
- Look at **individual angular bins** – just like **electron scattering** experiments looking at a **single scattering angle**





**Gauss:**  
 $\mu = 0.75 \text{ GeV}$   
 $\sigma = 0.075 \text{ GeV}$

**2 Years of ND  
data**



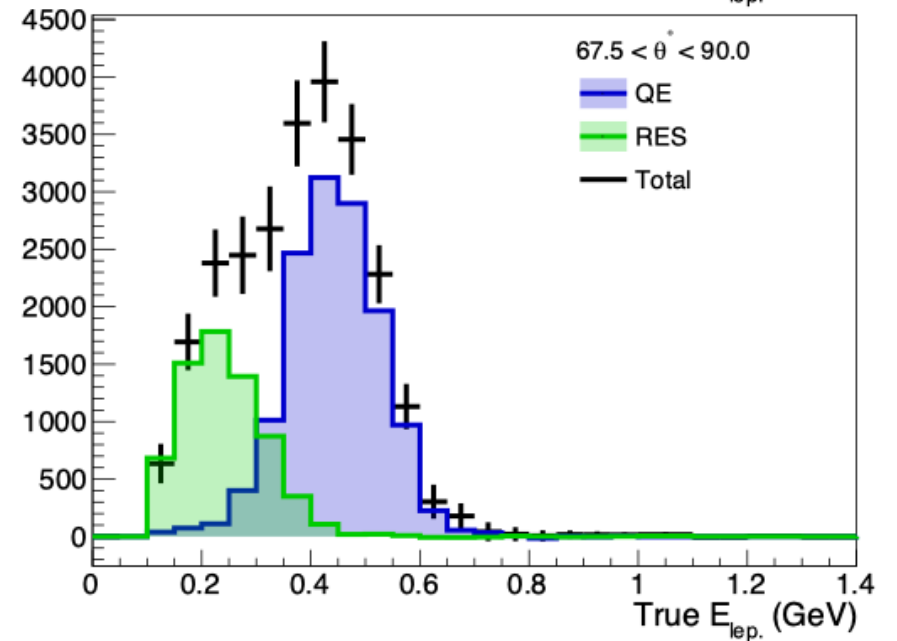
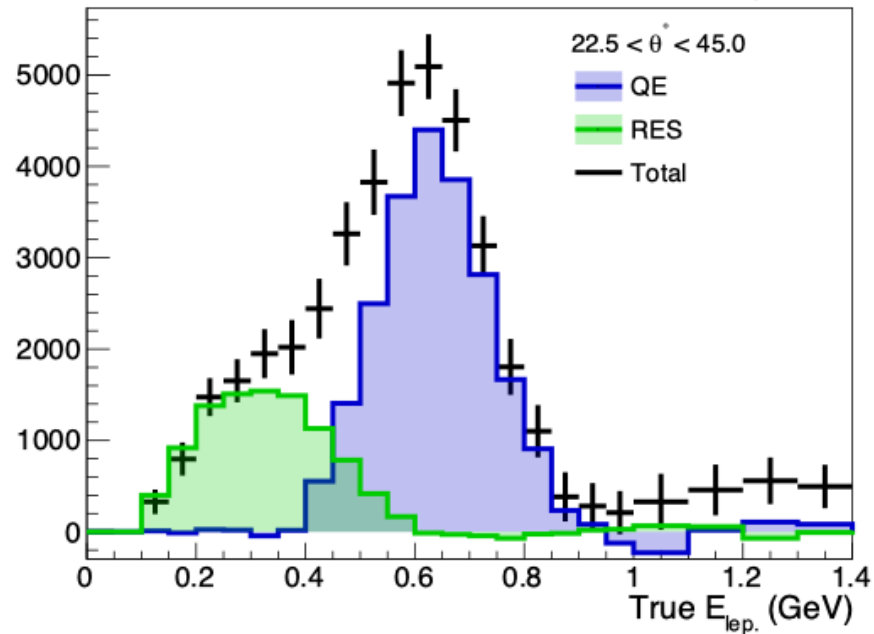
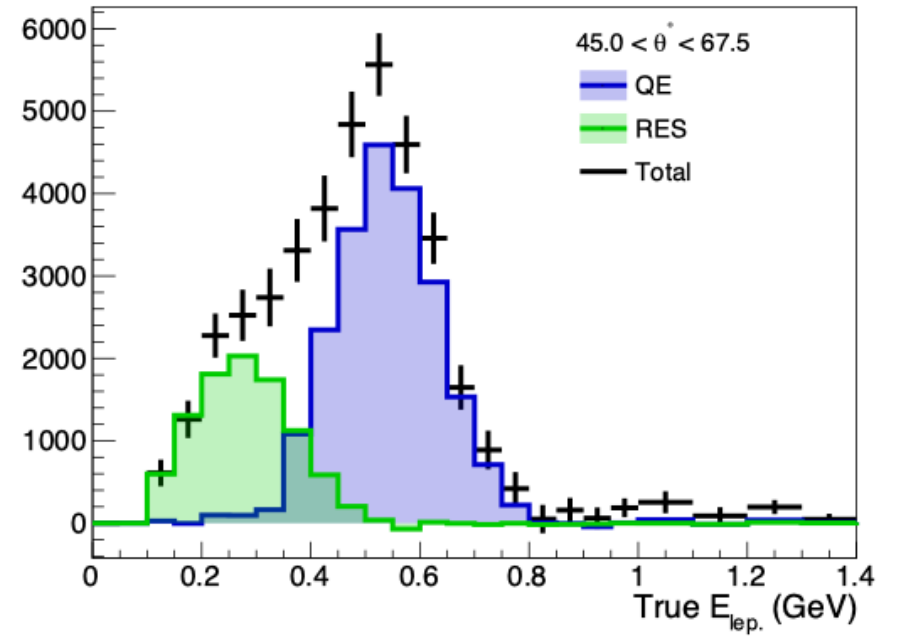
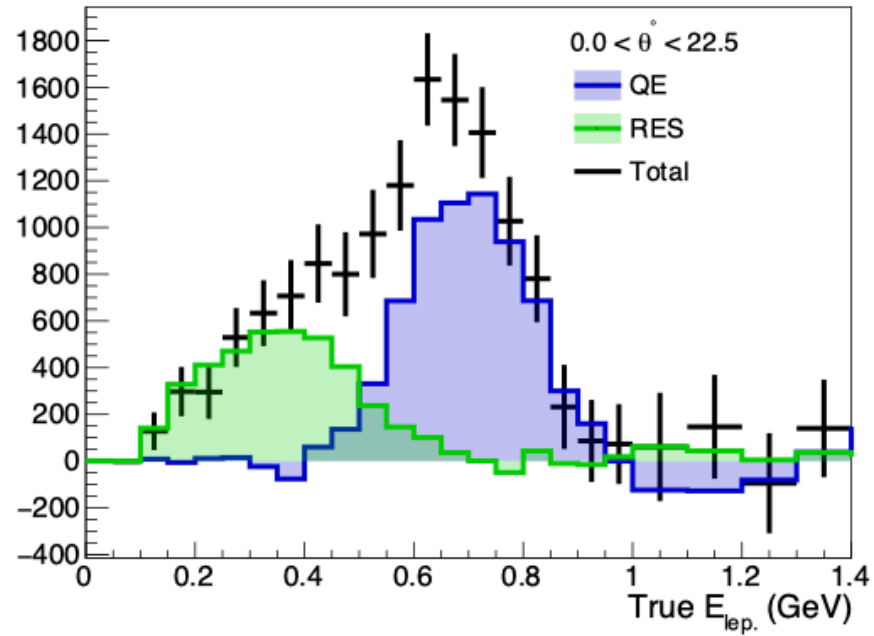
**Gauss:**

$\mu = 0.75 \text{ GeV}$

$\sigma = 0.1 \text{ GeV}$

Slightly wider gaussian:

- Better control of **statistical errors**
- Weaker **separation** of QE and RES



# Next Steps

- Main components are **working in CAFAna**
- Controlling **statistical uncertainties** is challenging but (I hope) manageable – more work needed
- Implementation of **all systematics** we should get for **free**
- Unfolding **detector effects** fine for 1D – need to get this working for **2D as well**
- Need to have a think of how to **demonstrate this is useful** in a simulated **cross section analysis**

**Thanks for listening!**