

# Long-Lived Particles at Spallation Sources

In collaboration with C. Argüelles (Harvard) & S. Urrea (IFIC, Valencia)



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**FNAL theory seminar**

**Sep 5 2024**

$\mathcal{L}_{\text{SM}}$

**Incompleteness problems:**

Dark matter

Neutrino masses and mixing

**Predictivity problems:**

Topological angle  $\theta_{\text{QCD}} \rightarrow 0$   
(strong CP problem)

Fermion mass pattern

Higgs mass and naturalness

Matter-antimatter asymmetry

$\mathcal{L}_{\text{SM}}$

**DARK SECTOR (DS)**

$\mathcal{L}_{\text{DS}}$

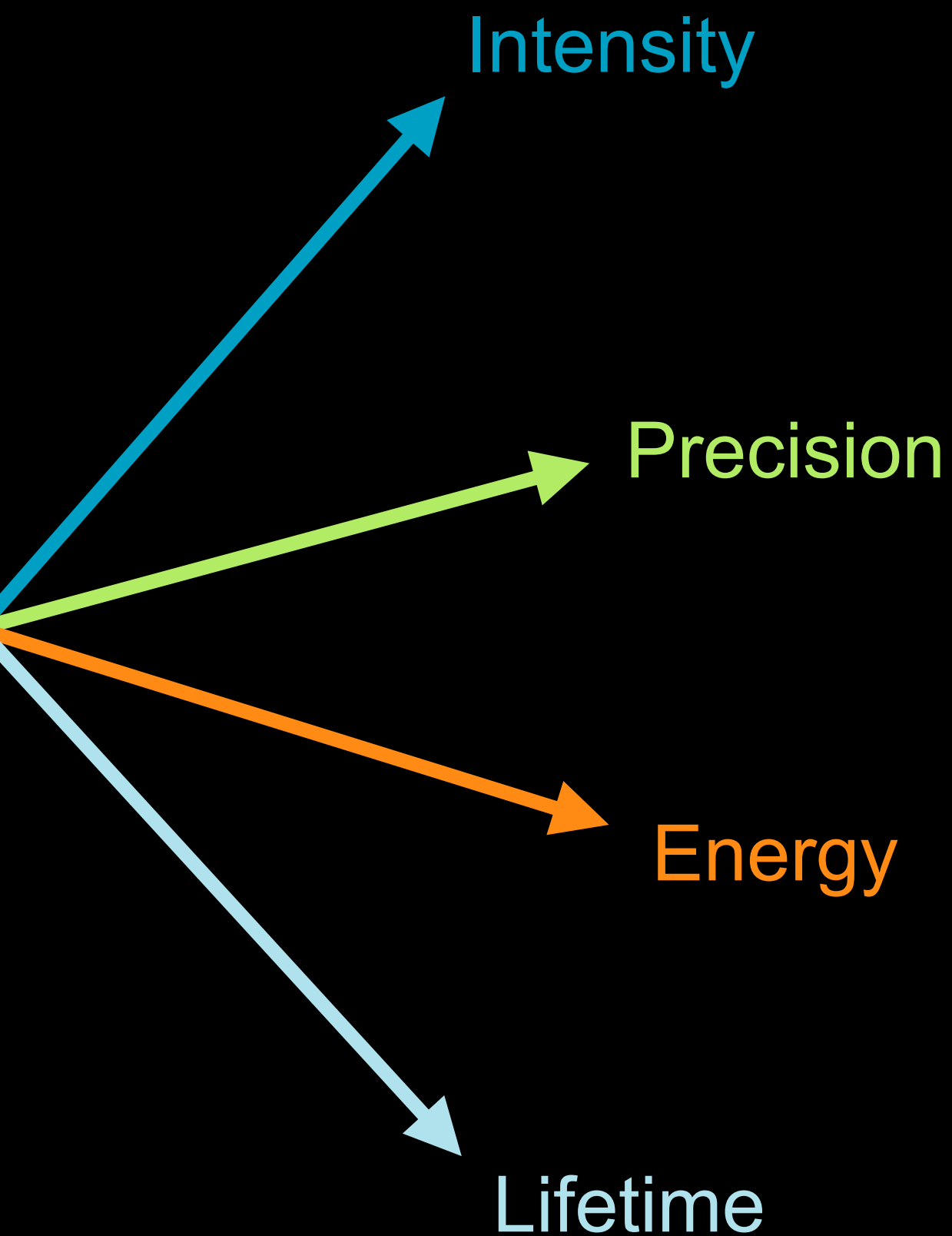
No SM charge

New fundamental mass scales

New particles can be light

**Testable!**

## Prongs of intensity frontier



### Long-lived particles (**Intensity** & **Lifetime**)

The existence of new light particles is not *mandatory* for solving the *predictivity and incompleteness* issues of the Standard Model.

But it sure would provide major hints of the direction forward.

Searching for new “dark sectors”  
is **cheap\*** but also **exciting!**

\*but not free, we do have to think about it.

# Outline:

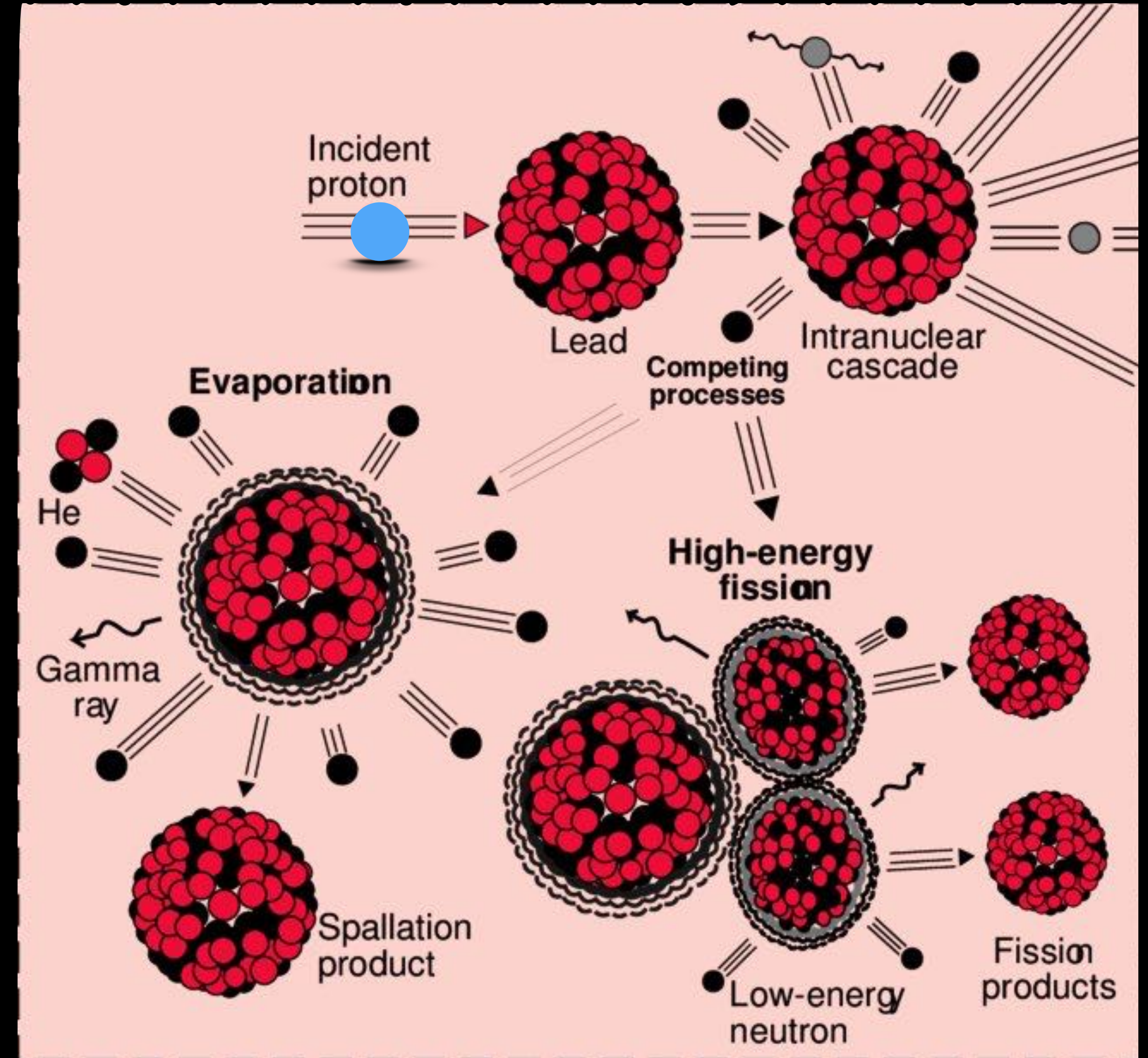
- 1) Spallation (neutron) sources: what and why?**
- 2) Opportunities with neutrino detectors
- 3) Long-lived particles below  $K$ ,  $\pi$ , and  $\mu$  masses

# What is a Spallation (Neutron) Source?

$\mathcal{O}(\text{GeV})$  proton beam on dense targets

→ bright neutron sources

→ efficient pion (+ kaon) production



# Why a Spallation (Neutron) Source?

Neutron science and applications...  
lots of it.

But why not, e.g., a reactor?

ILL reactor:

58 MW ( $1e15$  neutrons/cm<sup>2</sup>/s)

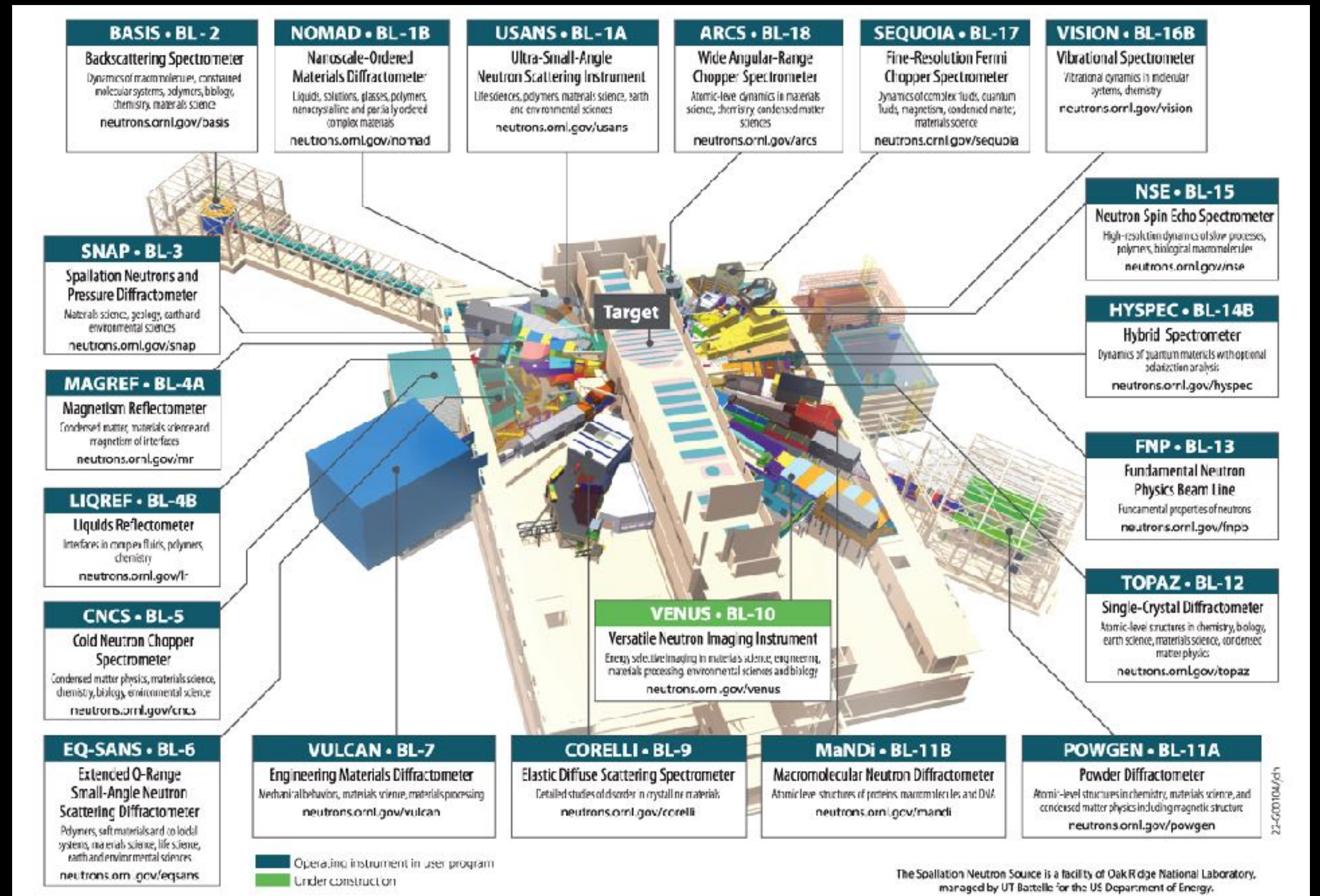
Oak Ridge Spallation:

1.4 MW ( $1e14$  neutrons/cm<sup>2</sup>/s — Average)

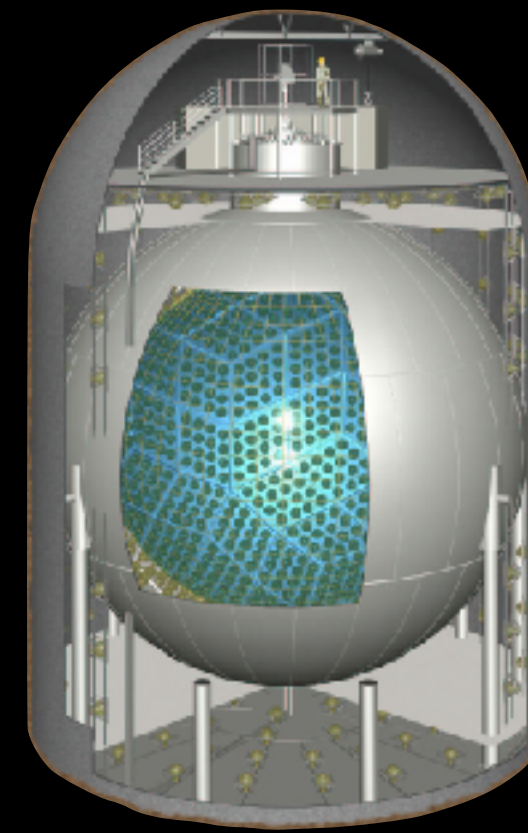
Atmosphere:

cosmic ray power ( $\sim 0.1$  neutrons/cm<sup>2</sup>/s)

# Oak Ridge Spallation Source instruments



# Why a Spallation (Neutron) Source?



KamLAND

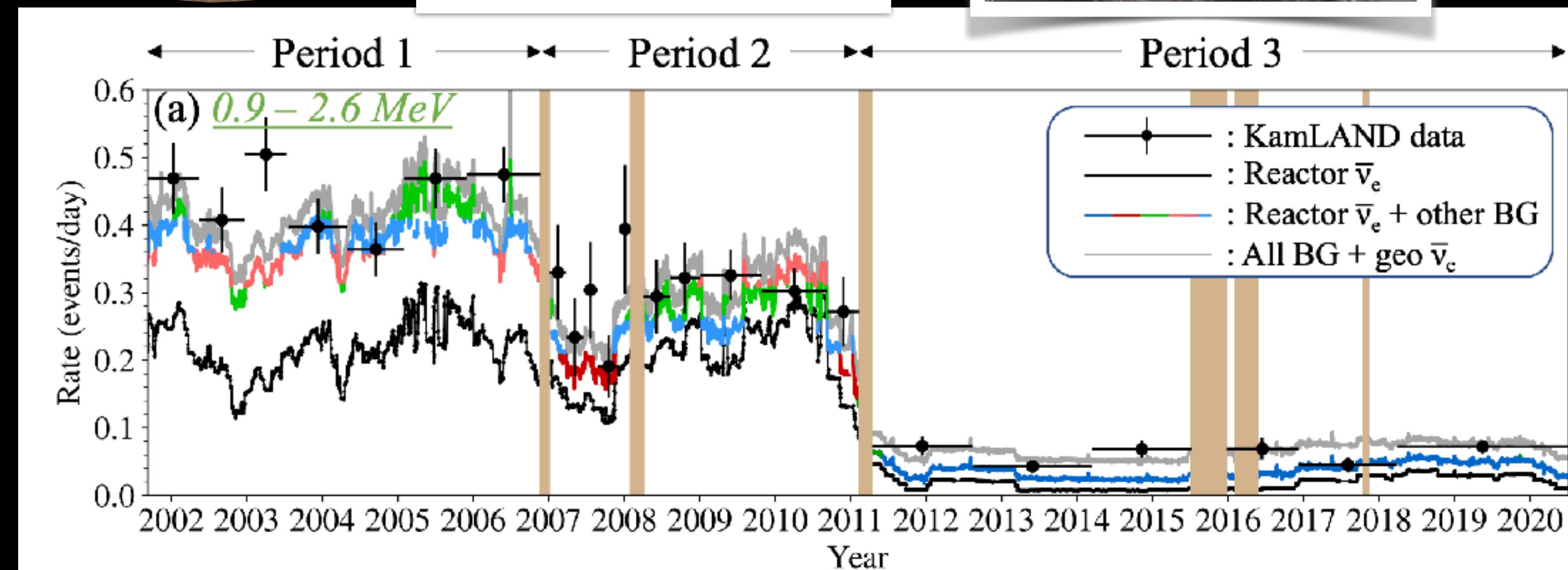
Tōhoku Tsunami  
Fukushima disaster



## Beams give you more control:

- 1) no secondary processes (no criticality needed — *endothermic*)

Purification periods



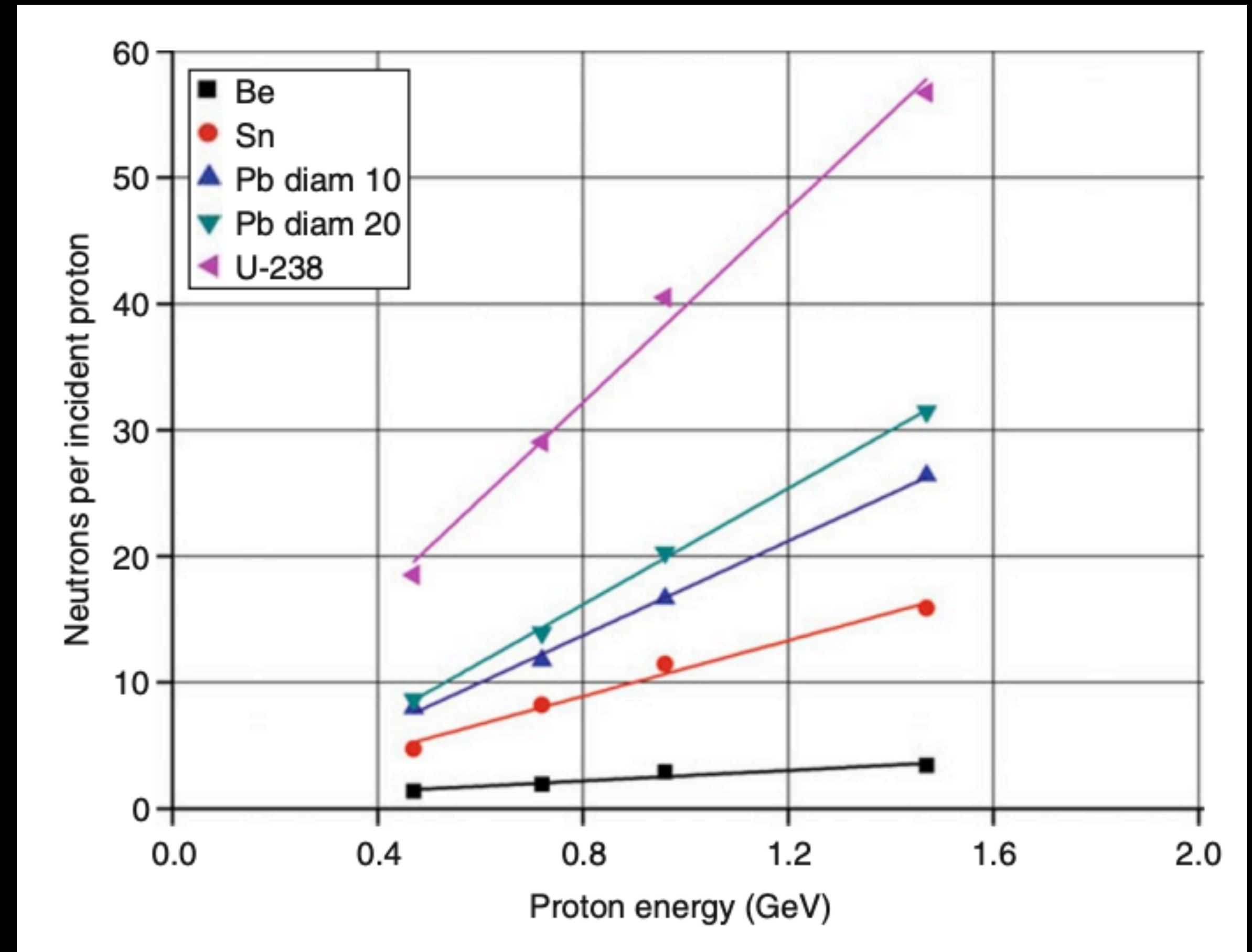
Impact of the reactor shutdown in Japan on the KamLAND program



## Why a Spallation (Neutron) Source?

### Beams give you more control:

- 1) no **secondary** processes  
(no criticality needed — *endothermic*)
- 2) **More neutrons per proton**



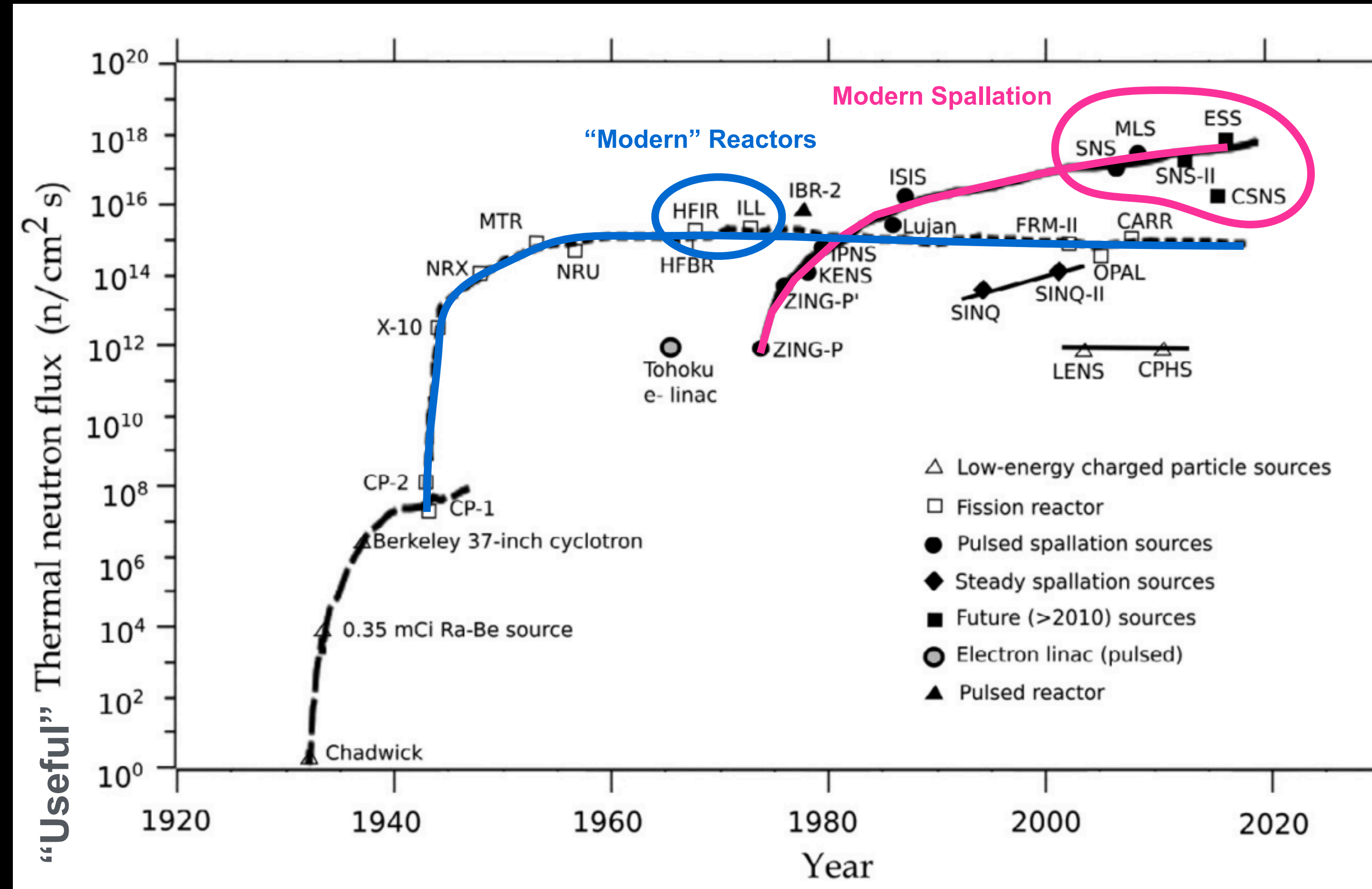
# Why a Spallation (Neutron) Source?

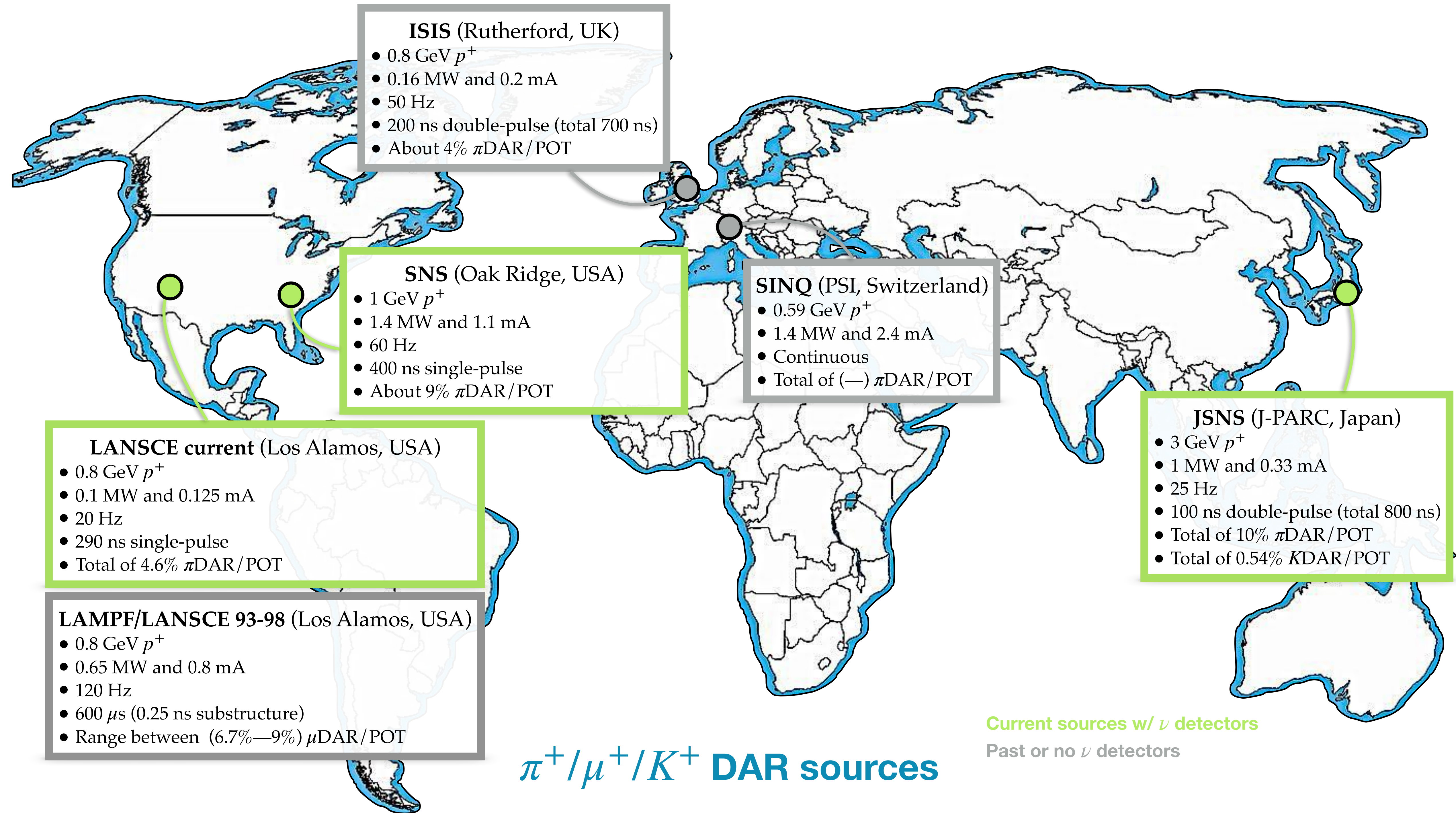
## Beams give you more control:

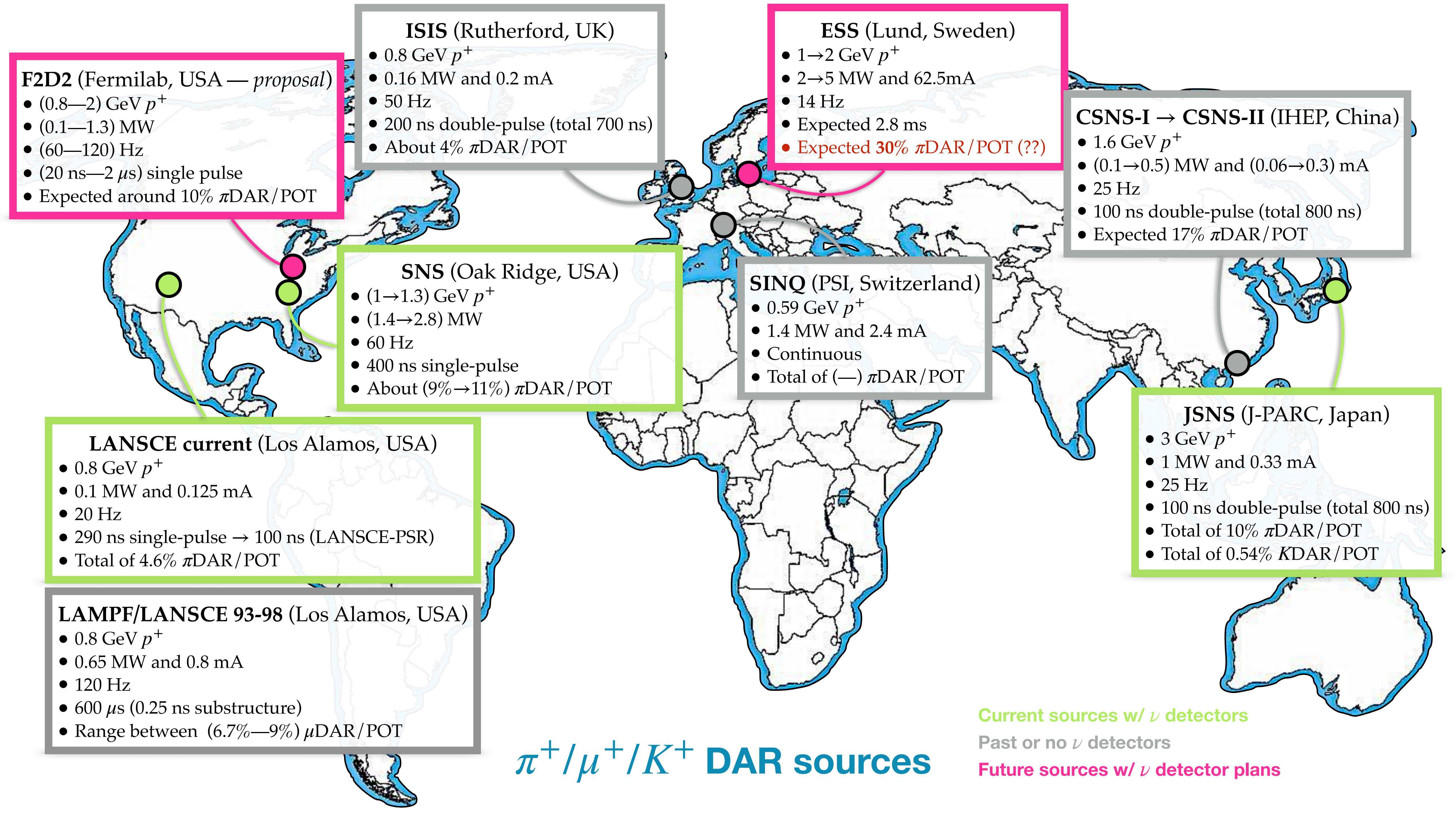
- 1) no secondary processes (no criticality needed — *endothermic*)
- 2) More neutrons per proton
- 3) Sources are pulsed

We will continue to see development of spallation sources around the world.

This is an opportunity for fundamental physics.

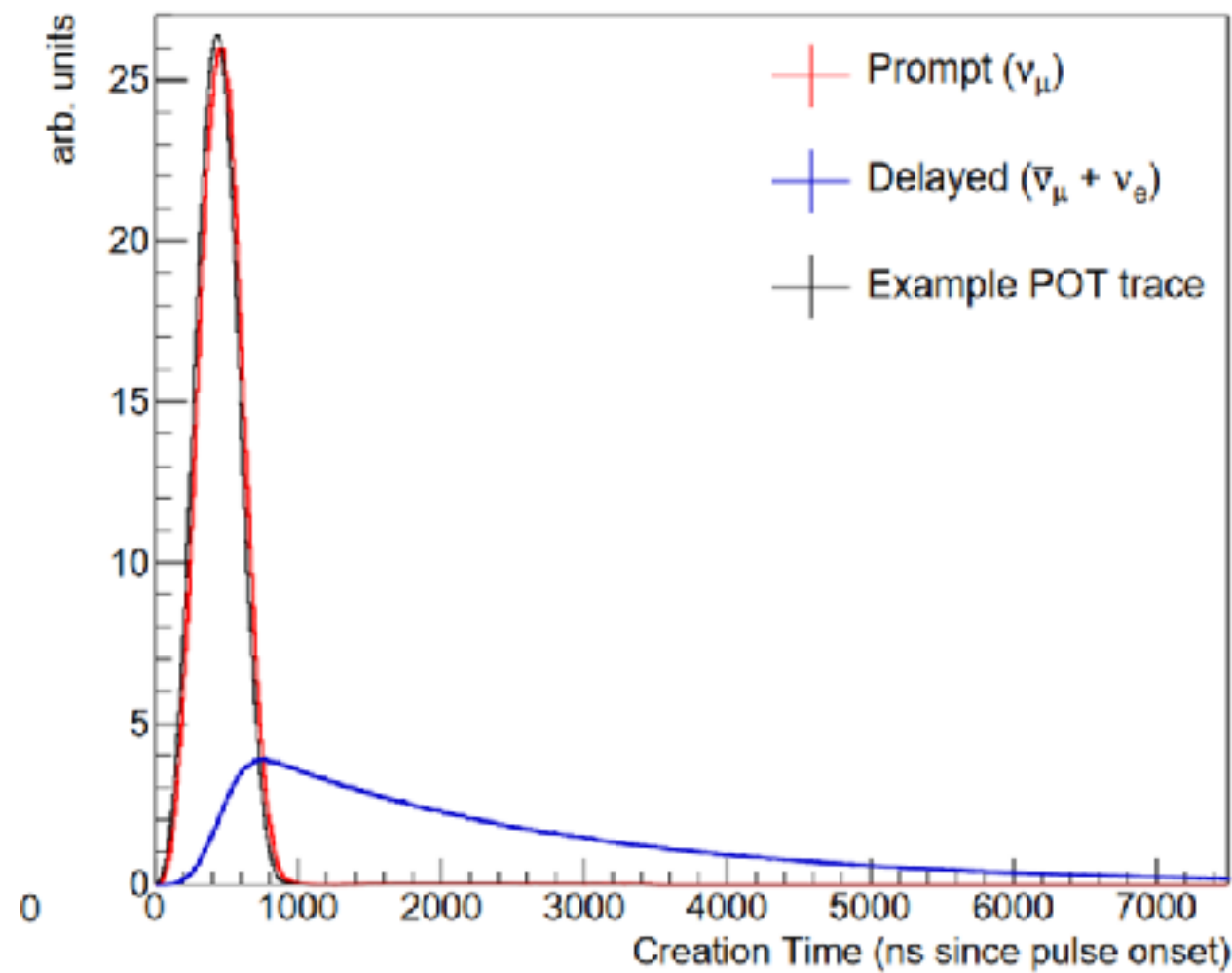






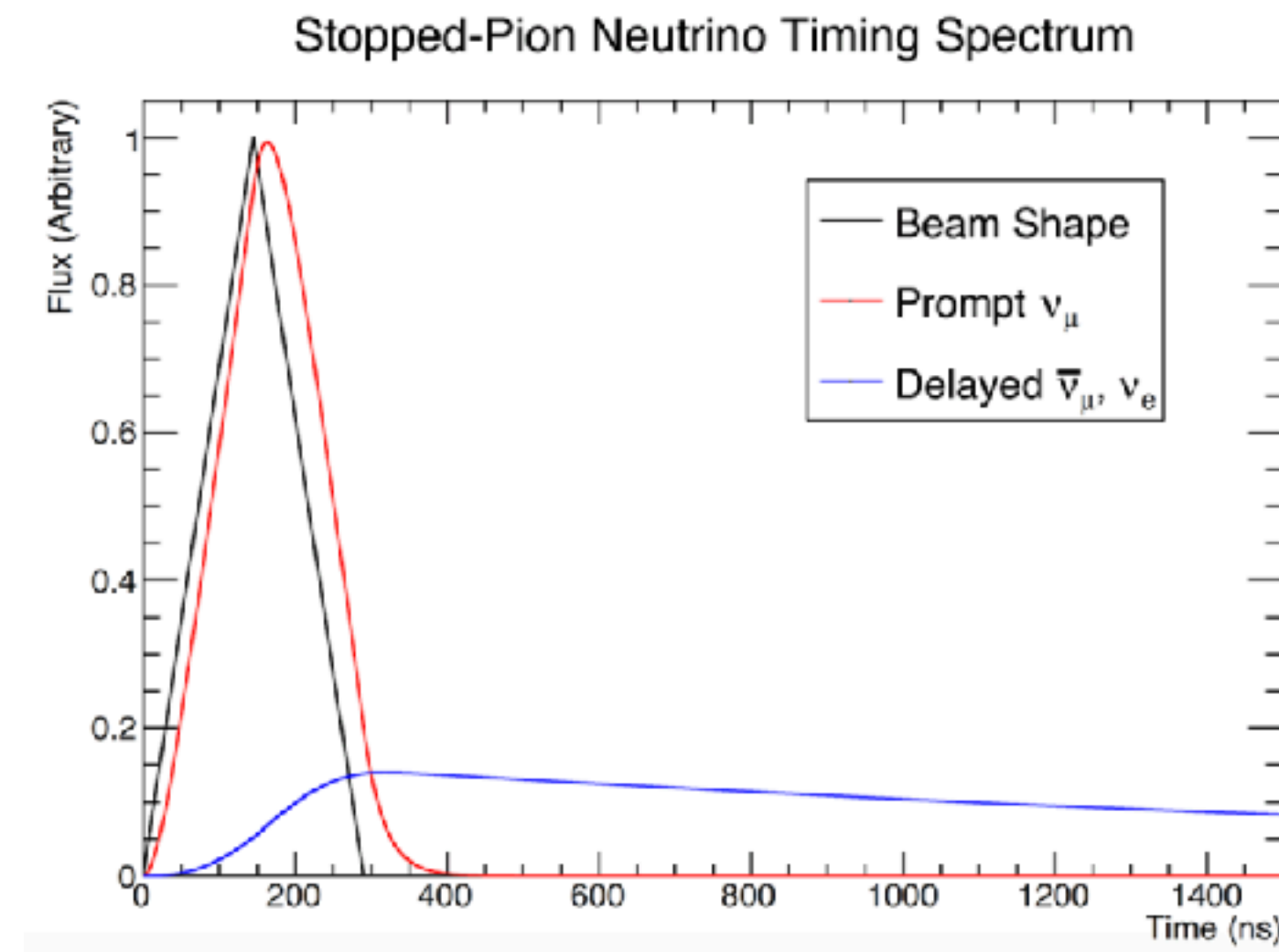
# Pulse shape

## Oak Ridge



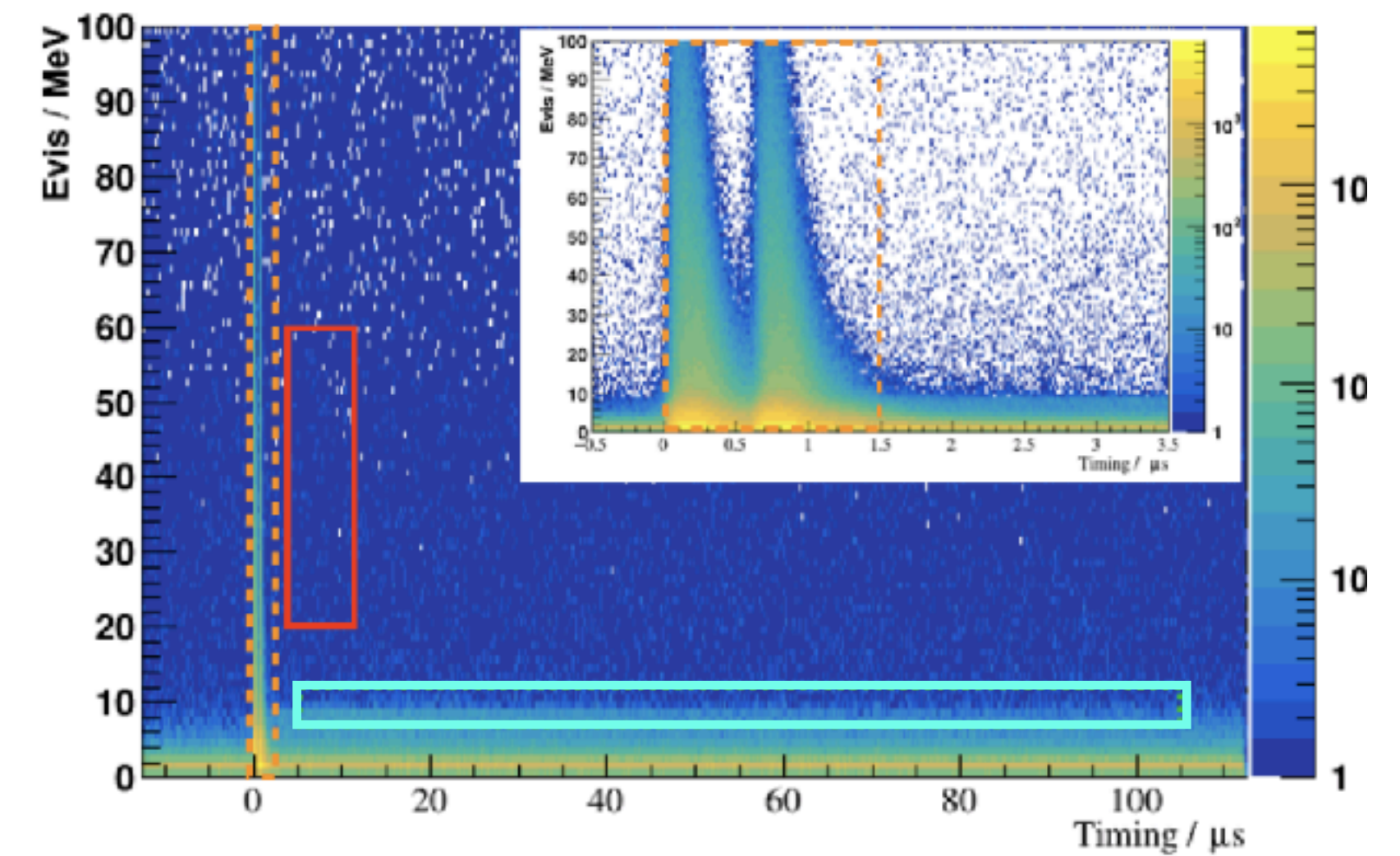
400 ns single-pulse

## Los Alamos



290 ns single-pulse  
Upgrade to 100 ns with LANSCE-PSR

## J-PARC



100 ns double-pulse (total 800 ns)

# Spallation sources as a neutrino source

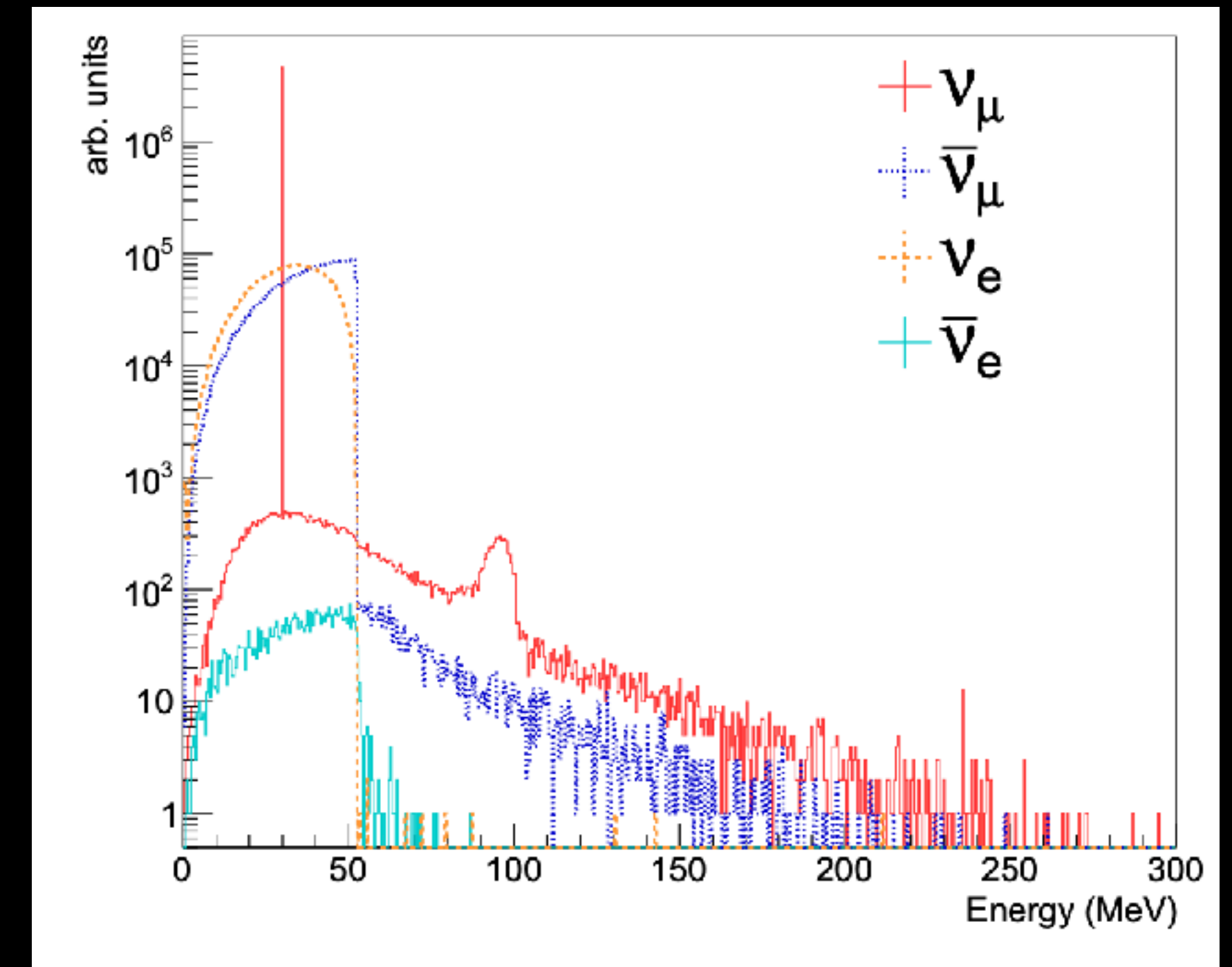
$10^{20} - 10^{22}$   $\pi^+$ ,  $K^+$ , and  $\mu^+$  decays at rest

Production data on p+Be target...

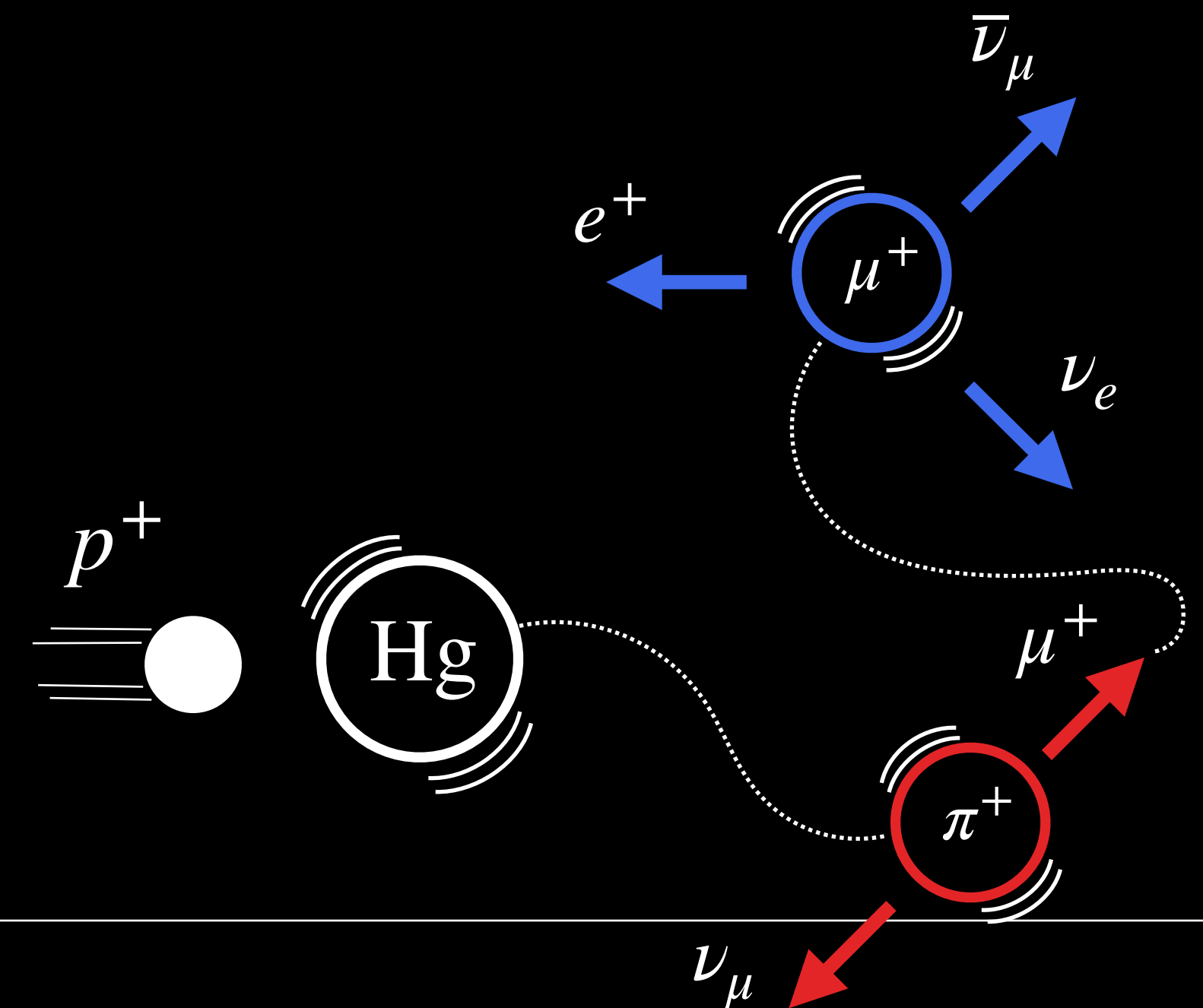
Produced Hadron	Exclusive Reaction	$M_X$ (GeV/c <sup>2</sup> )	$\sqrt{s_{thresh}}$ (GeV)	$E_{thresh}^{beam}$ GeV	KE of beam (MeV)
$\pi^+$	$pn\pi^+$	1.878	2.018	1.233	295
$\pi^-$	$pp\pi^+\pi^-$	2.016	2.156	1.54	602
$\pi^0$	$pp\pi^0$	1.876	2.011	1.218	280
$K^+$	$\Lambda^0 p K^+$	2.053	2.547	2.52	1582
$K^-$	$ppK^+K^-$	2.37	2.864	3.434	2496
$K^0$	$p\Sigma^+K^0$	2.13	2.628	2.743	1805

Very few locations!

J. Conrad

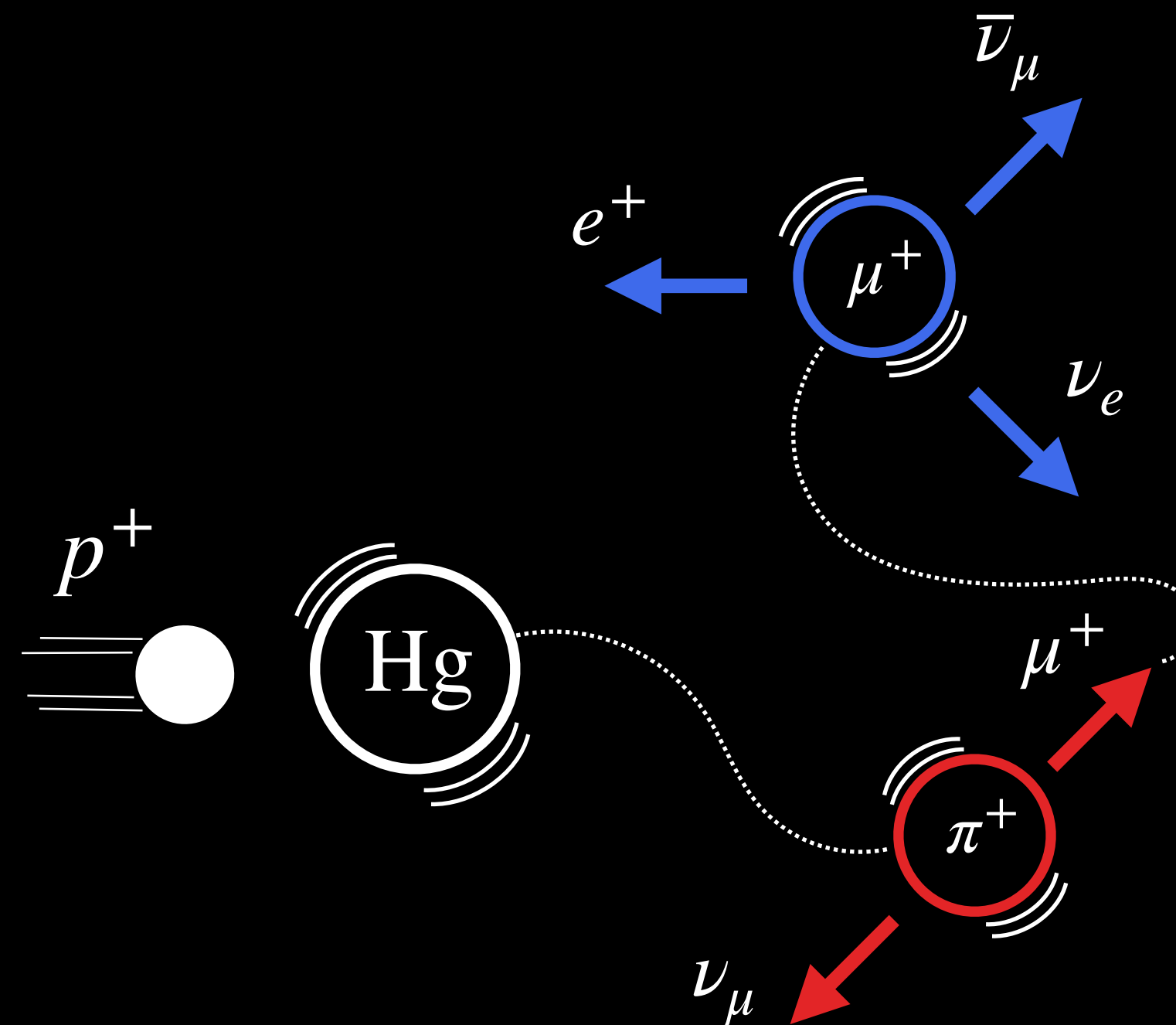
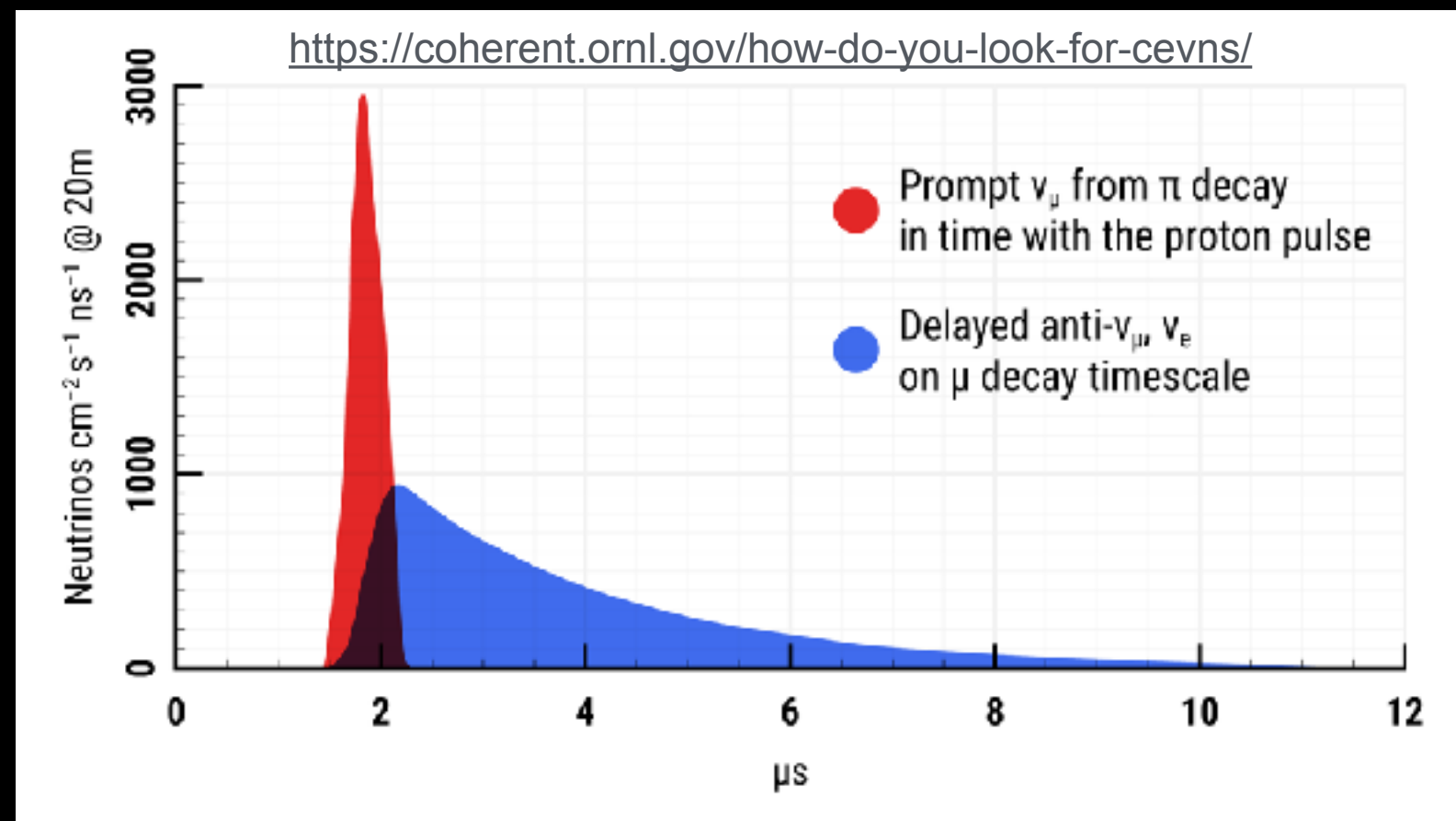


CEvNS and neutrino physics program



# Long-lived particles at Spallation sources

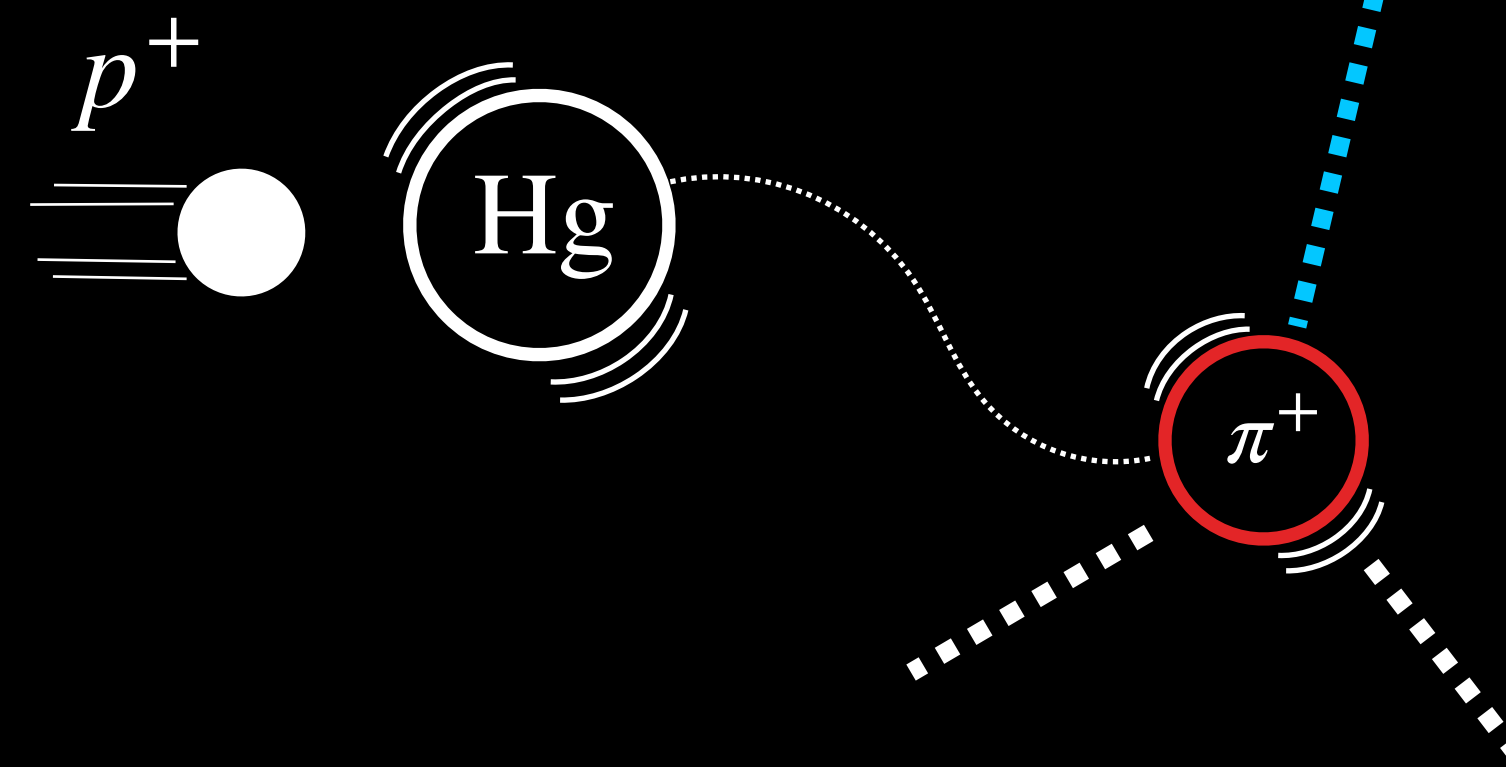
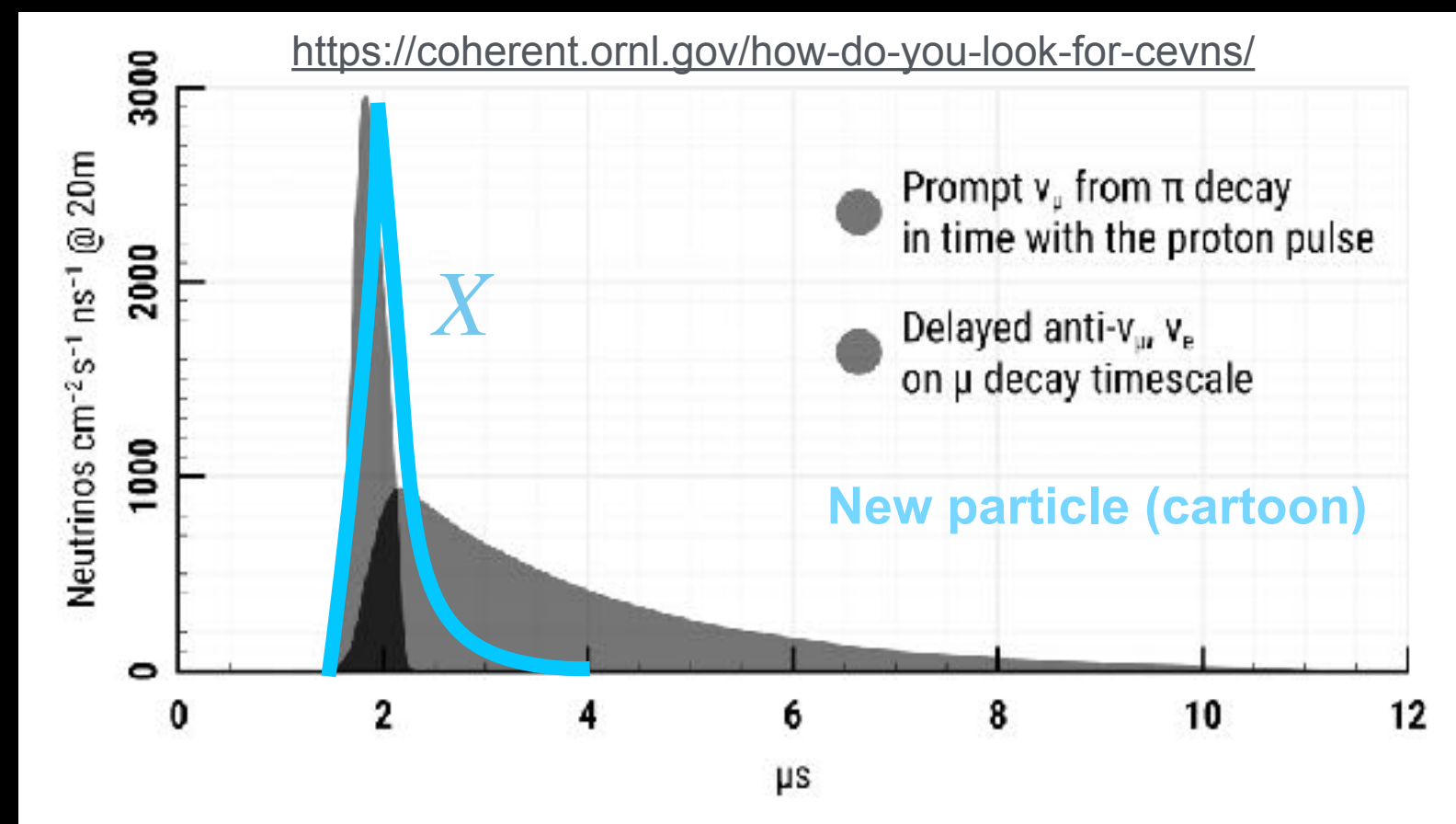
$10^{20} - 10^{22}$   $\pi^+$ ,  $K^+$ , and  $\mu^+$  decays at rest



# Long-lived particles at Spallation sources

$10^{20} - 10^{22}$   $\pi^+$ ,  $K^+$ , and  $\mu^+$  decays at rest

Can also search for a flux of new states.

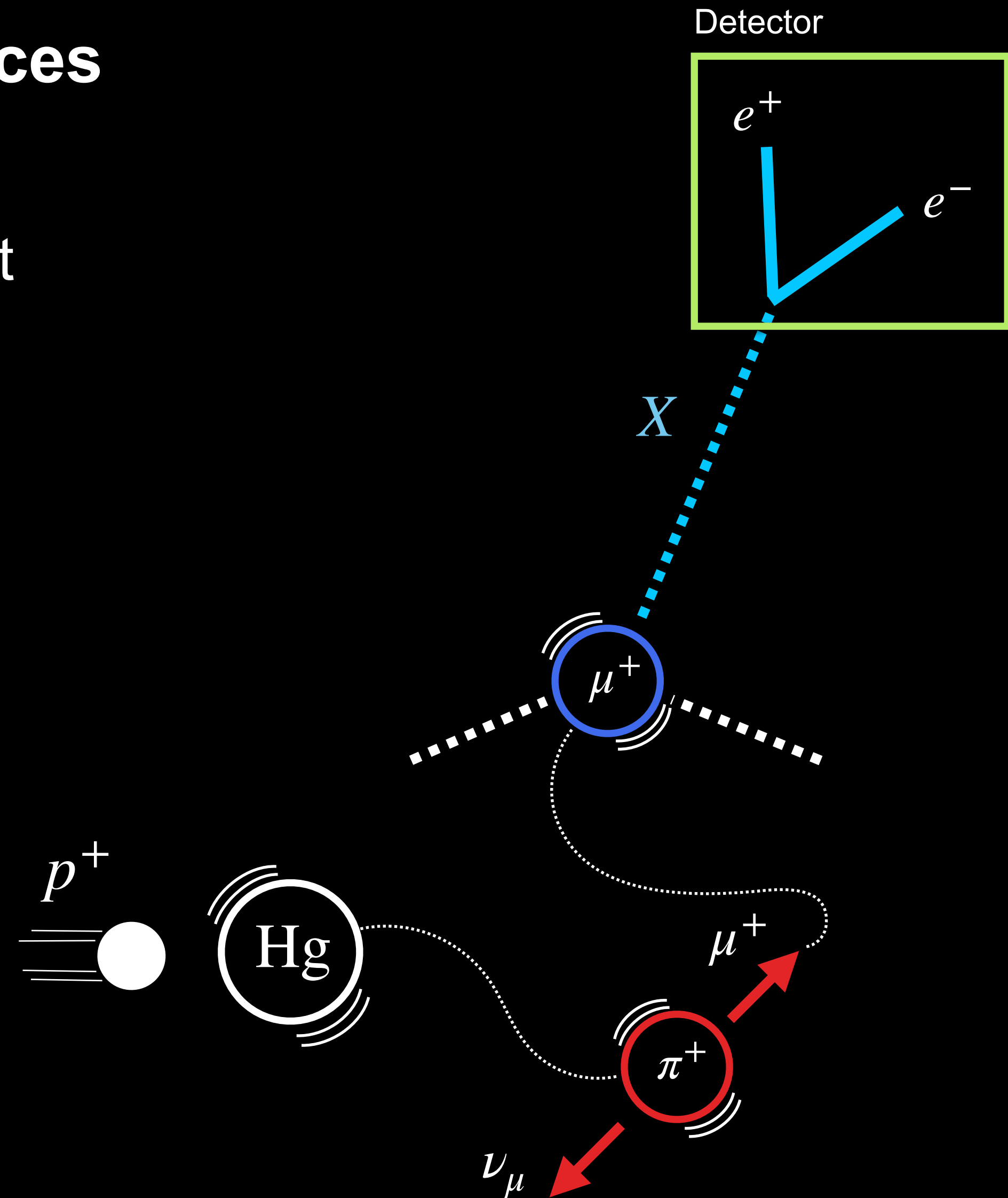
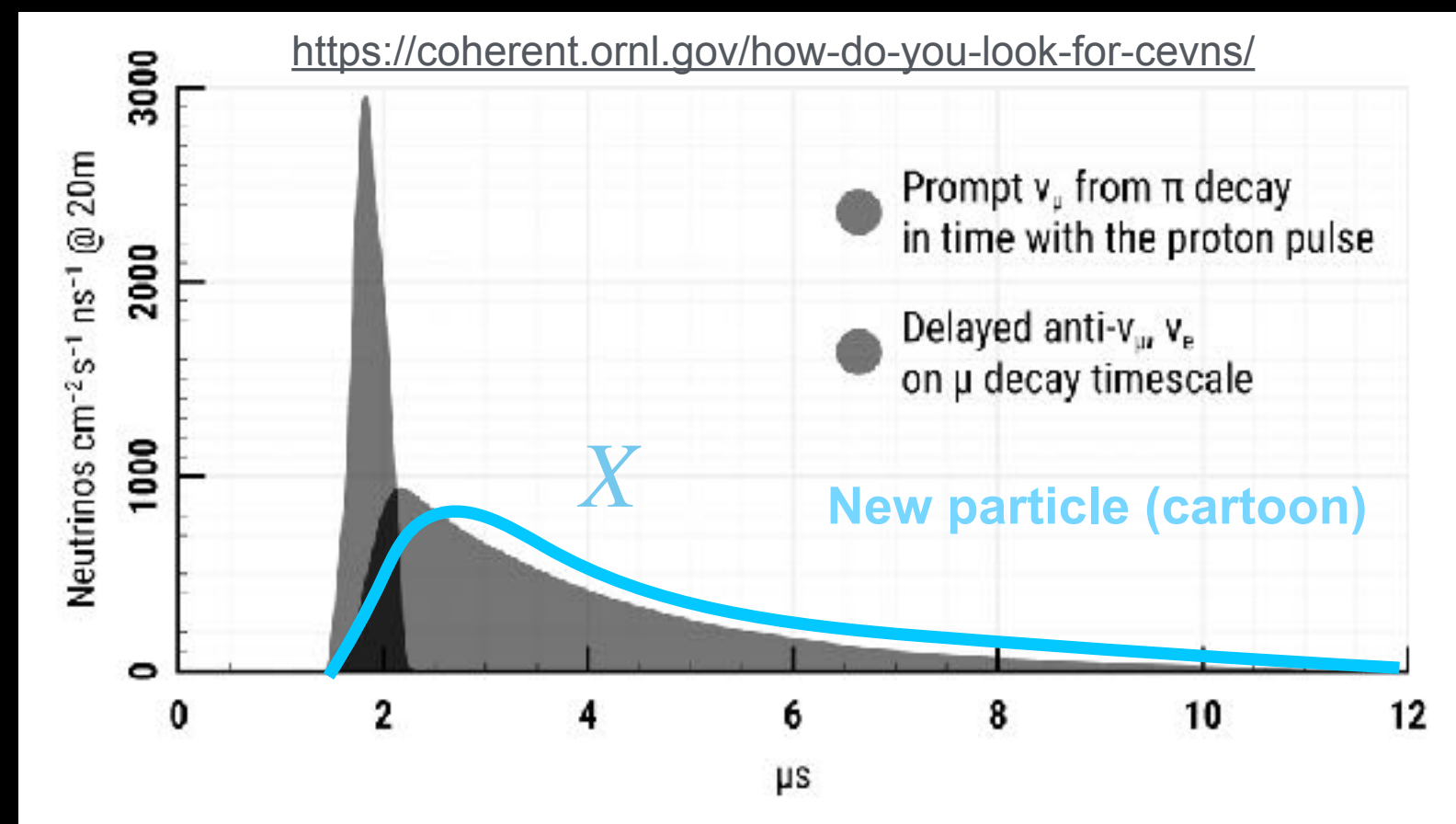




# Long-lived particles at Spallation sources

$10^{20} - 10^{22}$   $\pi^+$ ,  $K^+$ , and  $\mu^+$  decays at rest

Can also search for a flux of new states.



# Outline:

- 1) Spallation (neutron) sources: what and why?
- 2) Opportunities with neutrino detectors**
- 3) Long-lived particles below  $K$ ,  $\pi$ , and  $\mu$  masses

# What do we already have?

## The LSND $\nu - e$ scattering measurement

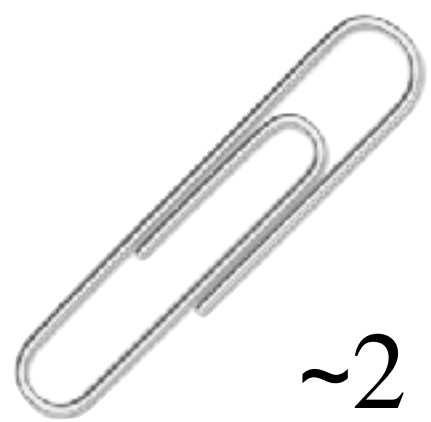
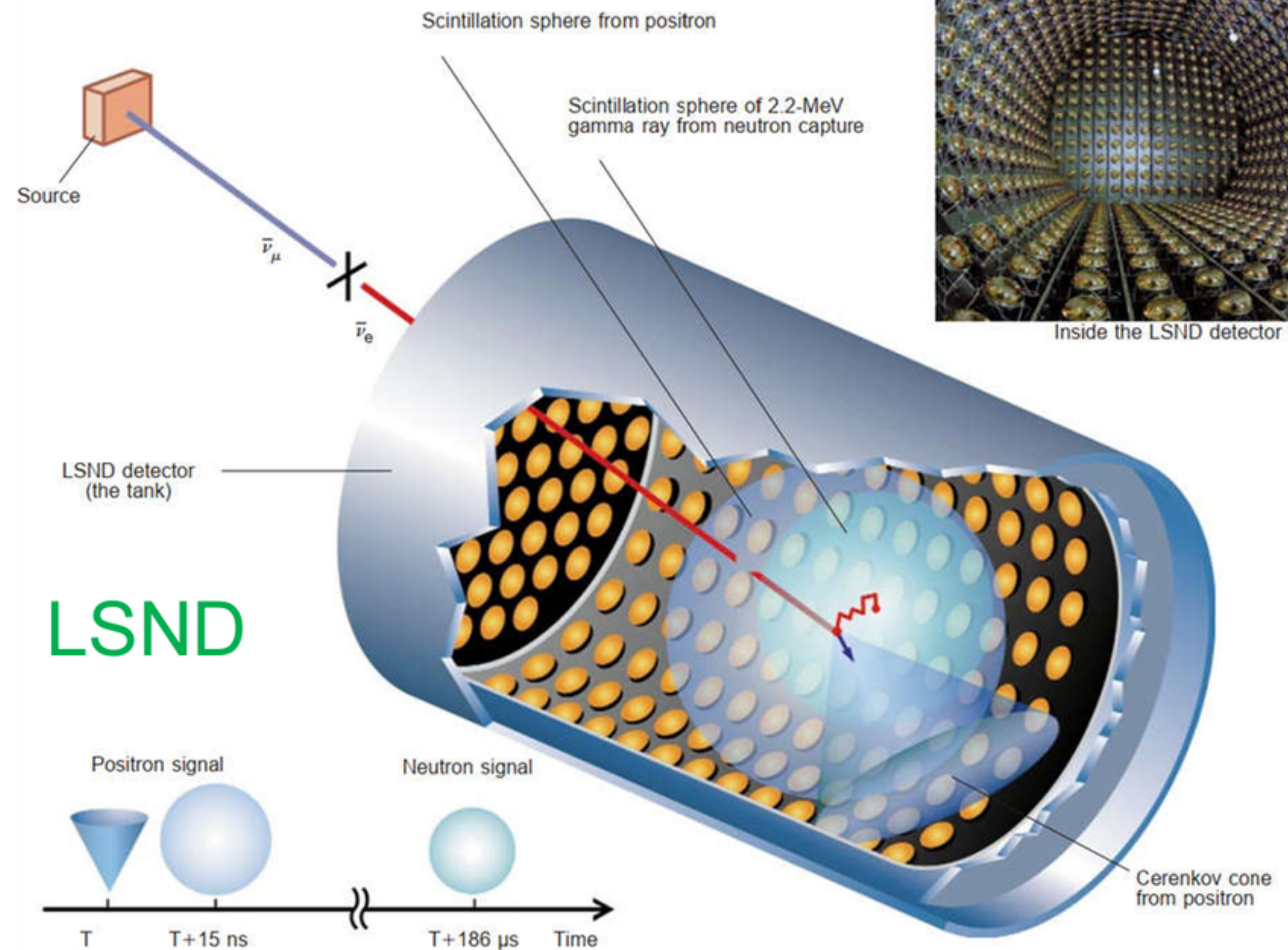
Extremely high intensity beam from ~93 - 98

~160 ton of liquid scintillator

$\sim 2 \times 10^{23}$  POT (about a paperclip's worth of protons)

A striking measurement of  $\nu + e \rightarrow \nu + e$  scattering with about 10% precision on the cross section.

(\*not the same channel as the LSND anomaly.)



# What do we already have?

## The LSND $\nu - e$ scattering measurement

**LSND:**  $\nu + e \rightarrow \nu + e$  scattering

(\*not the same channel as the LSND anomaly.)

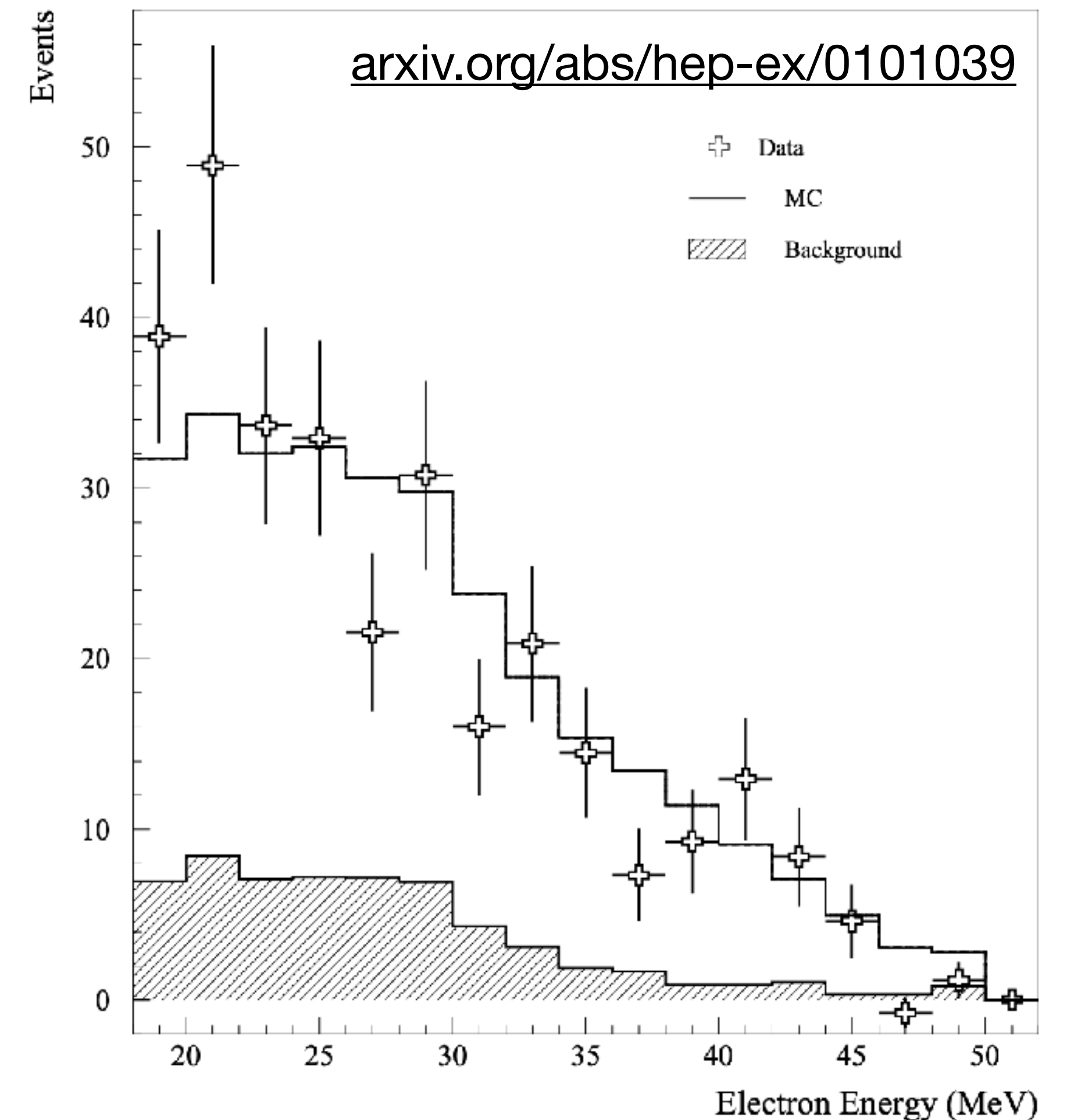
Very useful for BSM applications.

But it has its limitations:

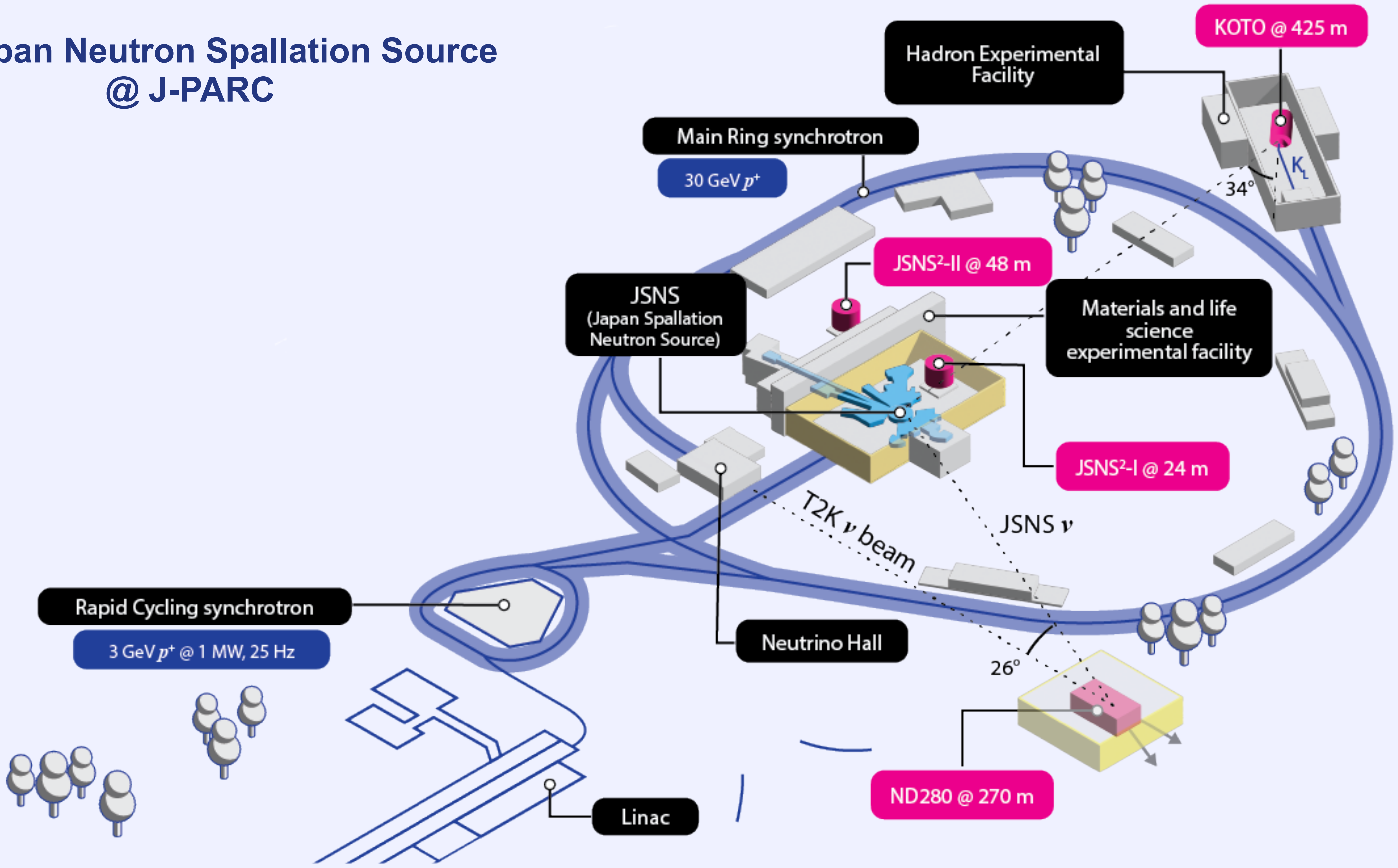
- 1) Very wide beam pulse (600  $\mu$ s) and low E ( $\pi$  and  $\mu$  only).
- 2) Only single EM showers (how to account for misID of  $e^+e^-$ ?)
- 3) Limited energy range:  $18 \text{ MeV} < E_{\text{vis}} < 50 \text{ MeV}$
- 4) Only the most forward electrons:  $\cos \theta_{\text{vis}} > 0.9$

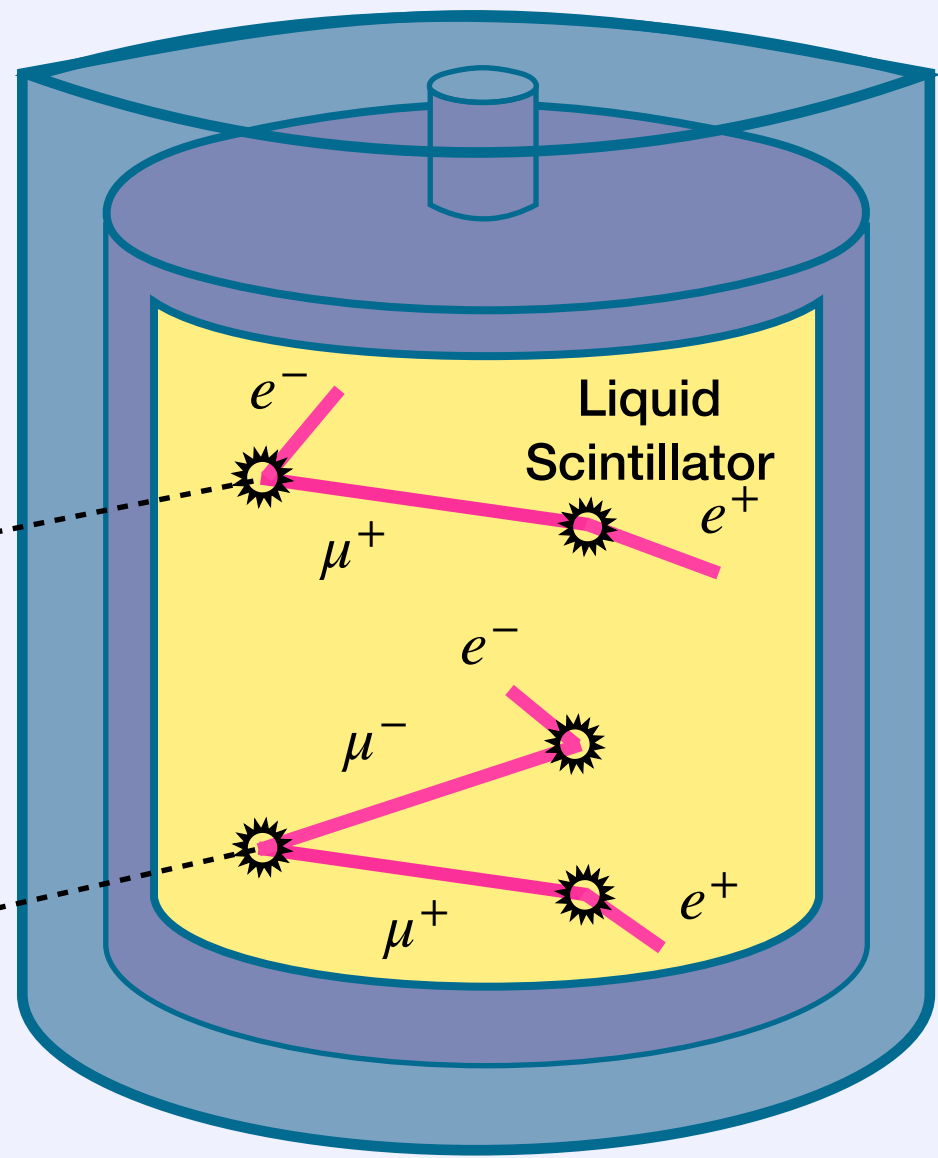
**No data release... how to model efficiencies?**

**All bounds come from theorists recast of this plot  $\rightarrow$**



# The Japan Neutron Spallation Source @ J-PARC

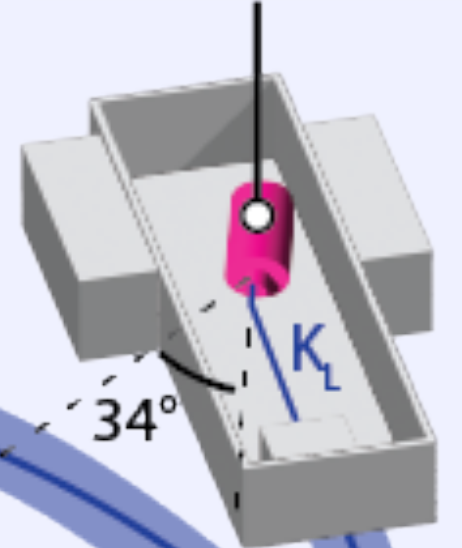




**JSNS<sup>2</sup> (I and II):**

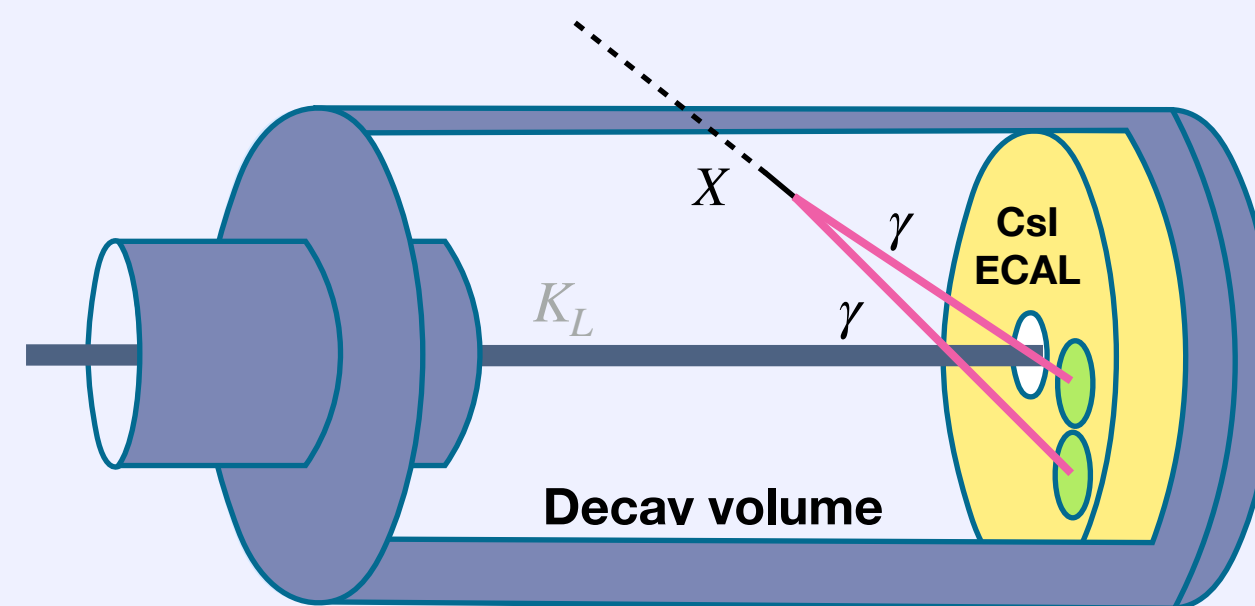
**Pros:** Closest to the source and largest vol  
**Cons:** larger backgrounds, single flash events (e.g.,  $e^+e^-$ ) very challenging for  $E_{\text{vis}} \lesssim 30$  MeV.  
**Best for:** double/triple flash ( $\mu\mu$ ,  $\mu\pi$ , or  $\nu\mu e$ ).

**KOTO @ 425 m**

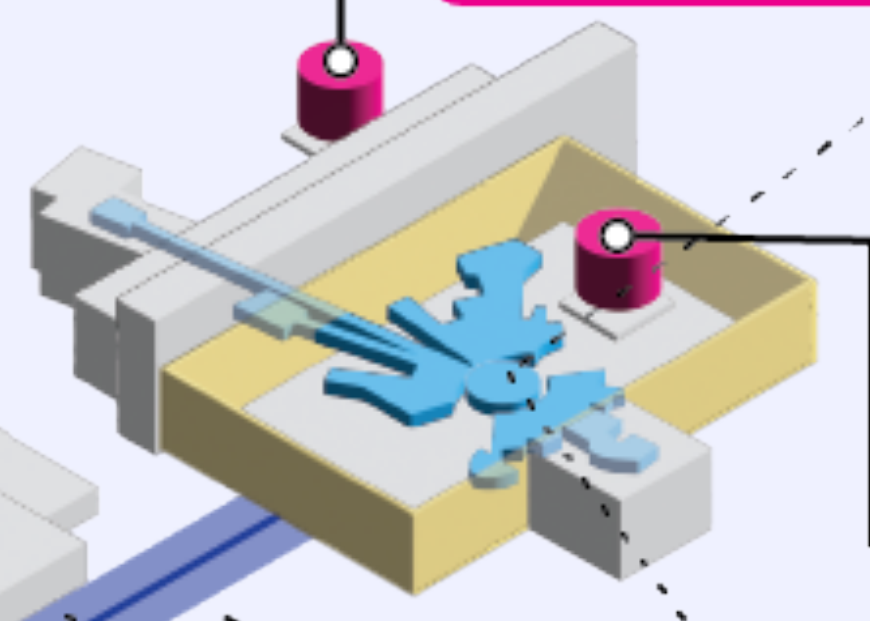


**KOTO:**

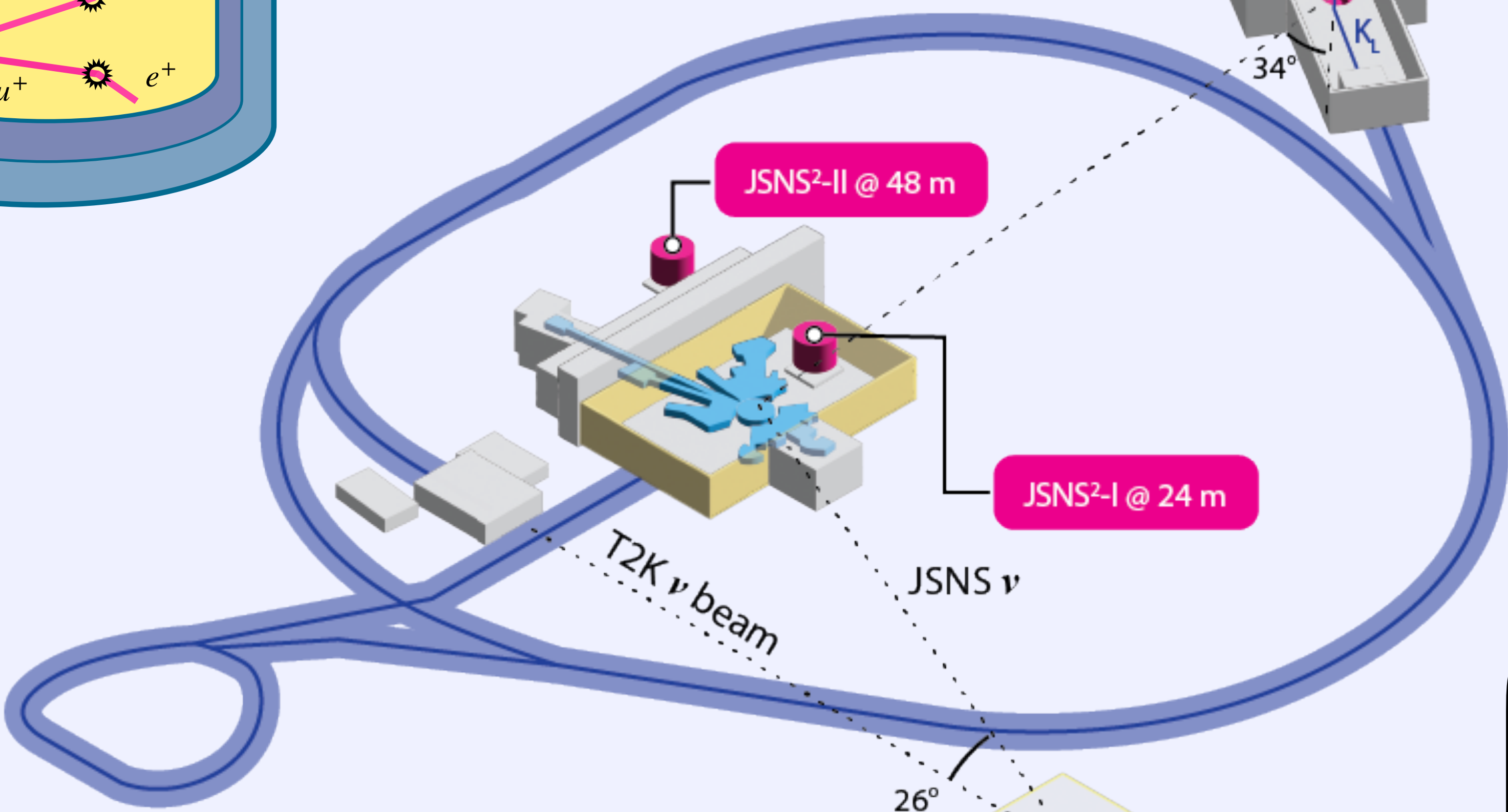
**Pros:** Low-density vol and low bkg  
**Cons:** Further away  
**Best for:**  $\pi^0$  and  $\gamma\gamma$



**JSNS<sup>2</sup>-II @ 48 m**



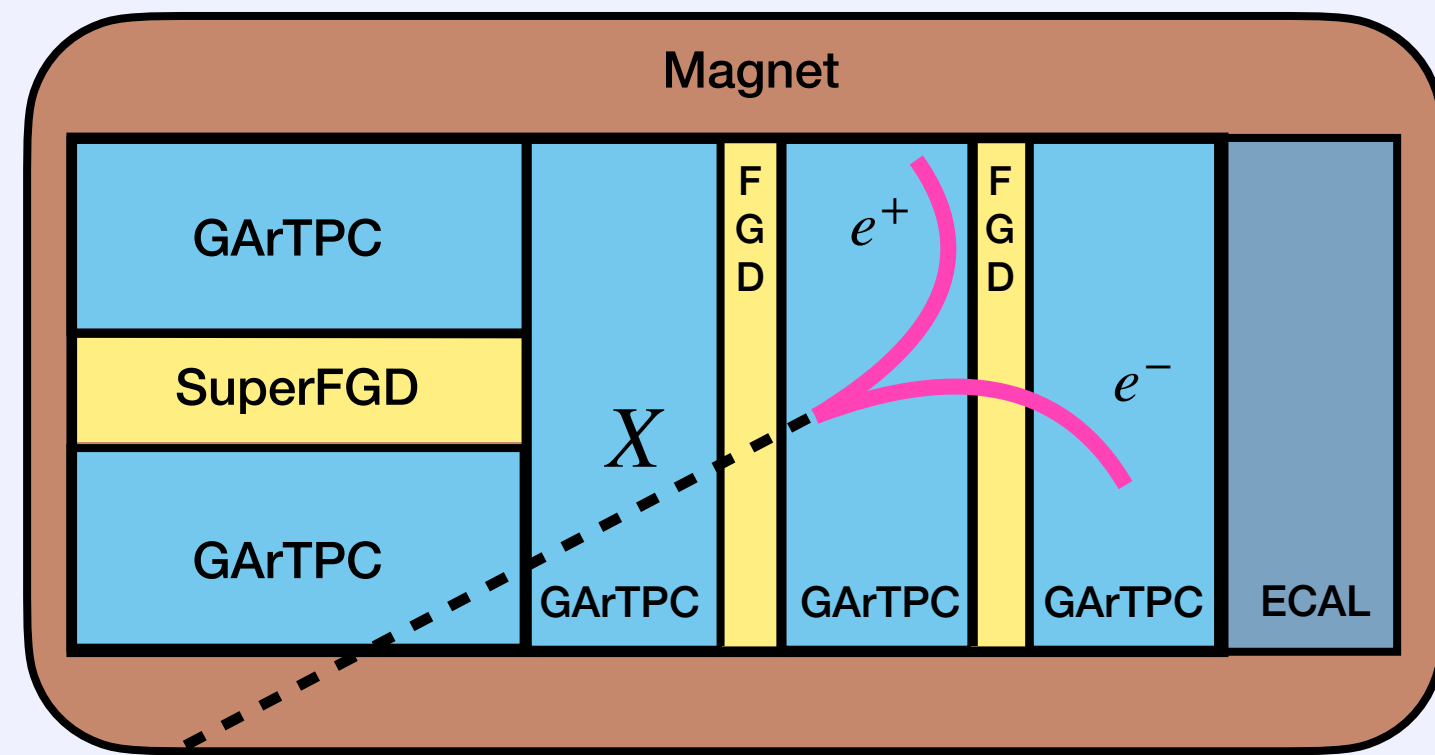
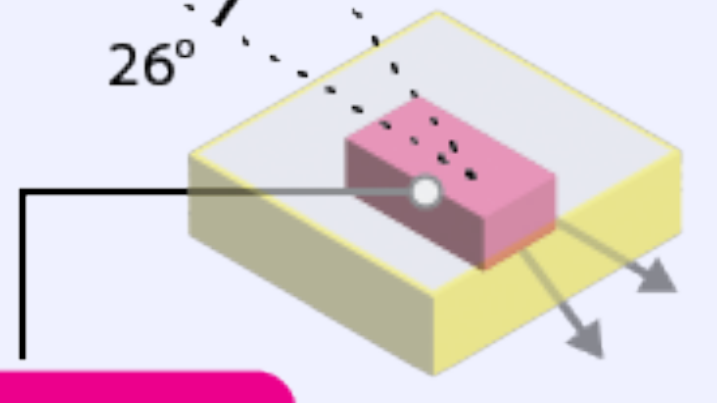
**JSNS<sup>2</sup>-I @ 24 m**

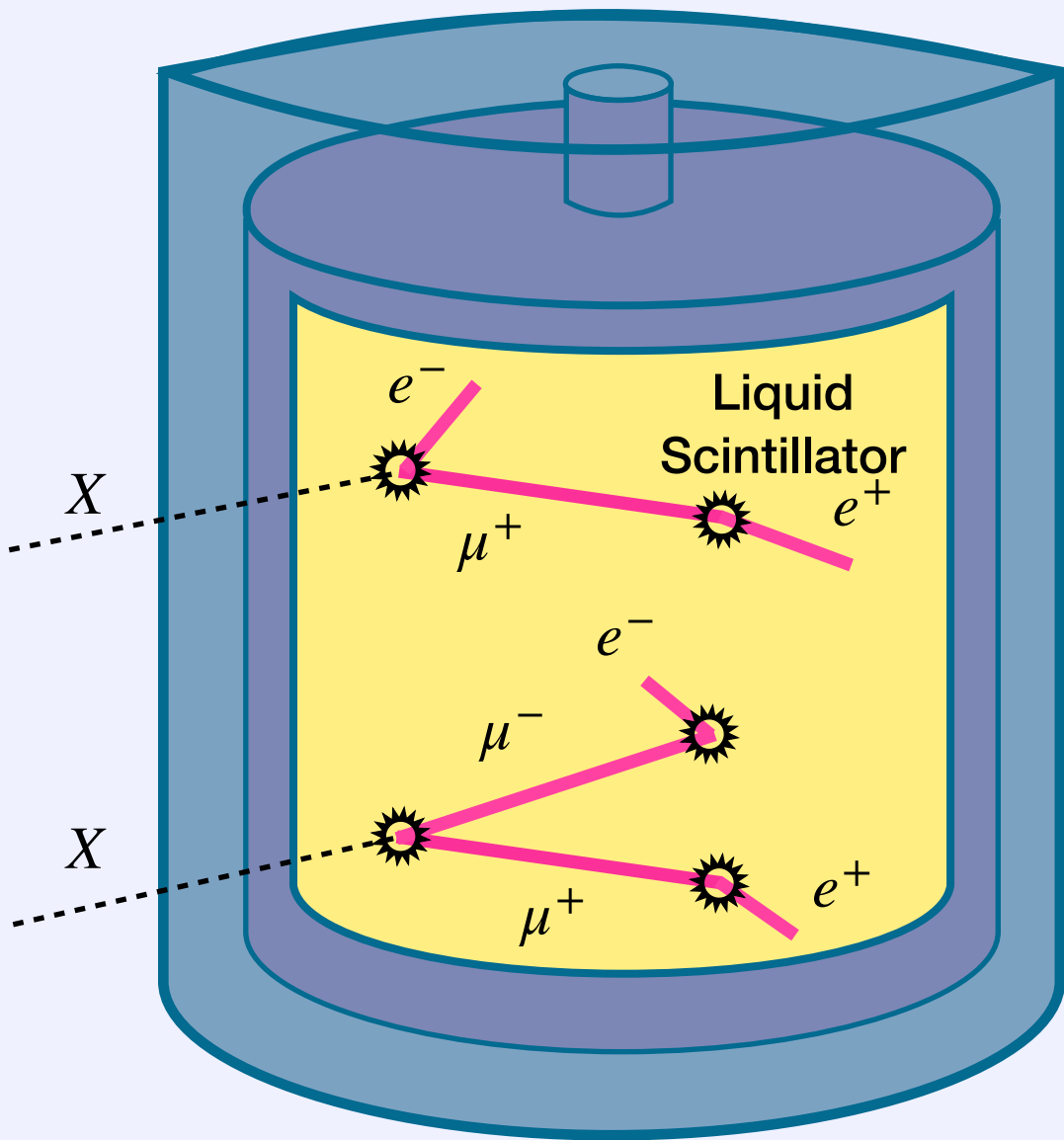


**ND280:**

**Pros:** Low-density and magnetized  
**Cons:** Further away  
**Best for:** any charged final state

**ND280 @ 270 m**





**JSNS<sup>2</sup> (I and II):**

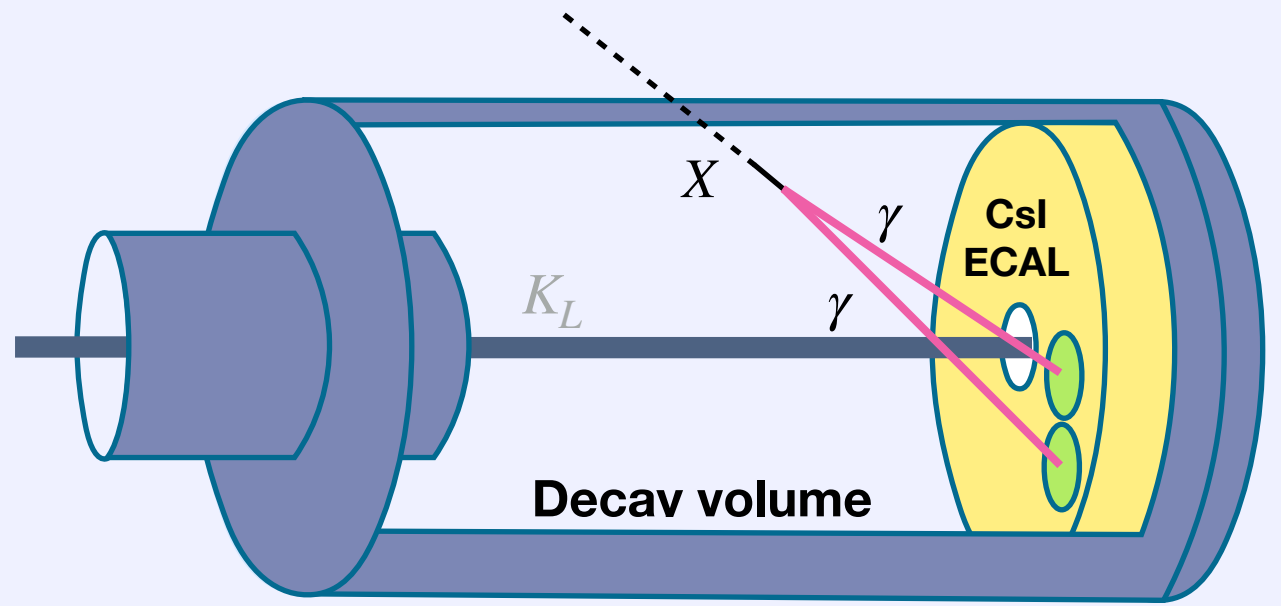
**Pros:** Closest to the source and largest vol  
**Cons:** larger backgrounds, single flash events  
 (e.g.,  $e^+e^-$ ) very challenging for  $E_{\text{vis}} \lesssim 30$  MeV.  
**Best for:** double

**KOTO @ 425 m**

Decay modes	
$X \rightarrow e^+e^-$	
$X \rightarrow e^+\mu^-$	<b>ND280</b>
$X \rightarrow \mu^+\mu^-$	<b>JSNS<sup>2</sup></b>
$X \rightarrow \gamma\gamma$	<b>KOTO</b>
$X \rightarrow e^-\pi^+$	
$X \rightarrow \mu^-\pi^+$	

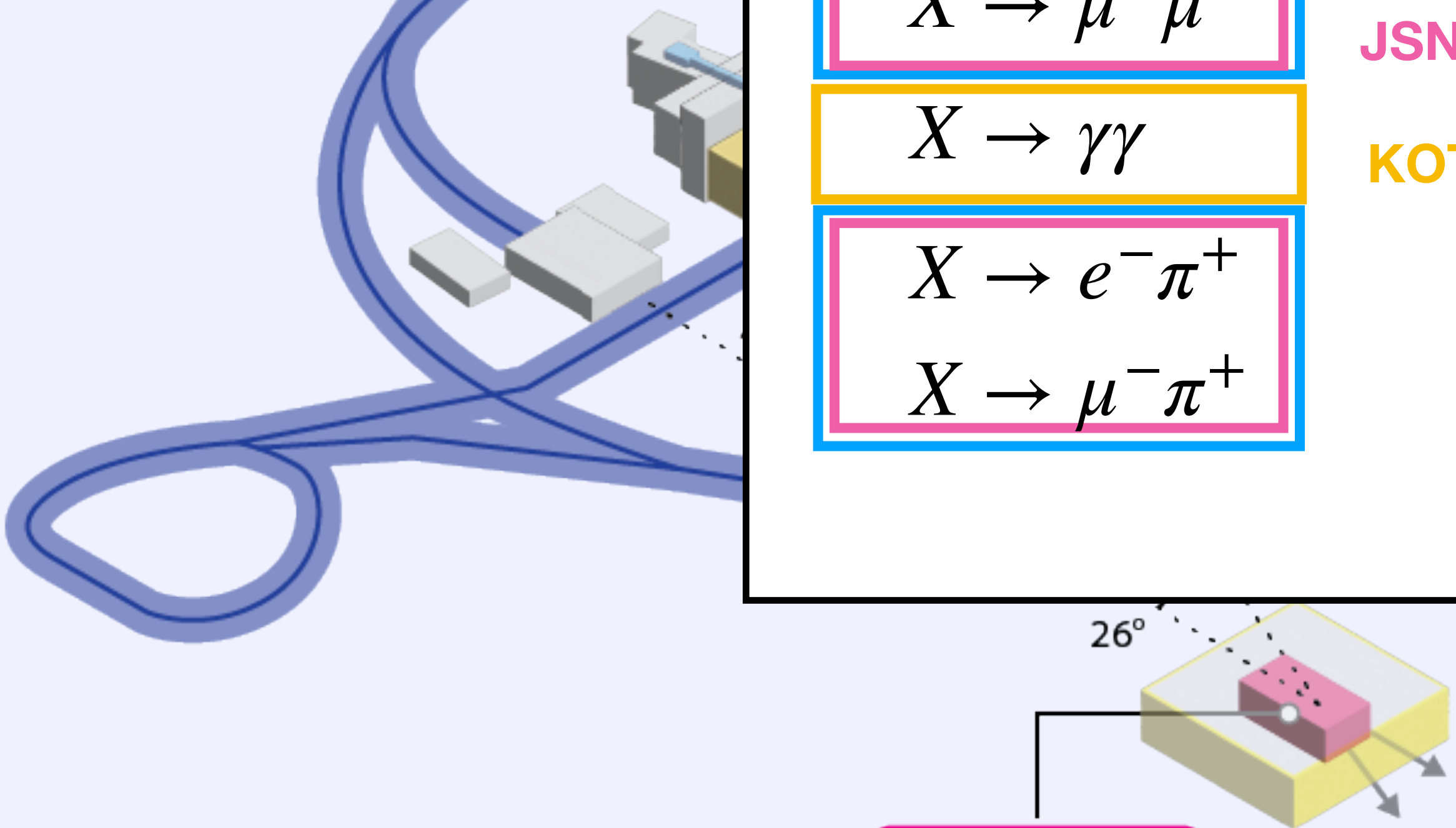
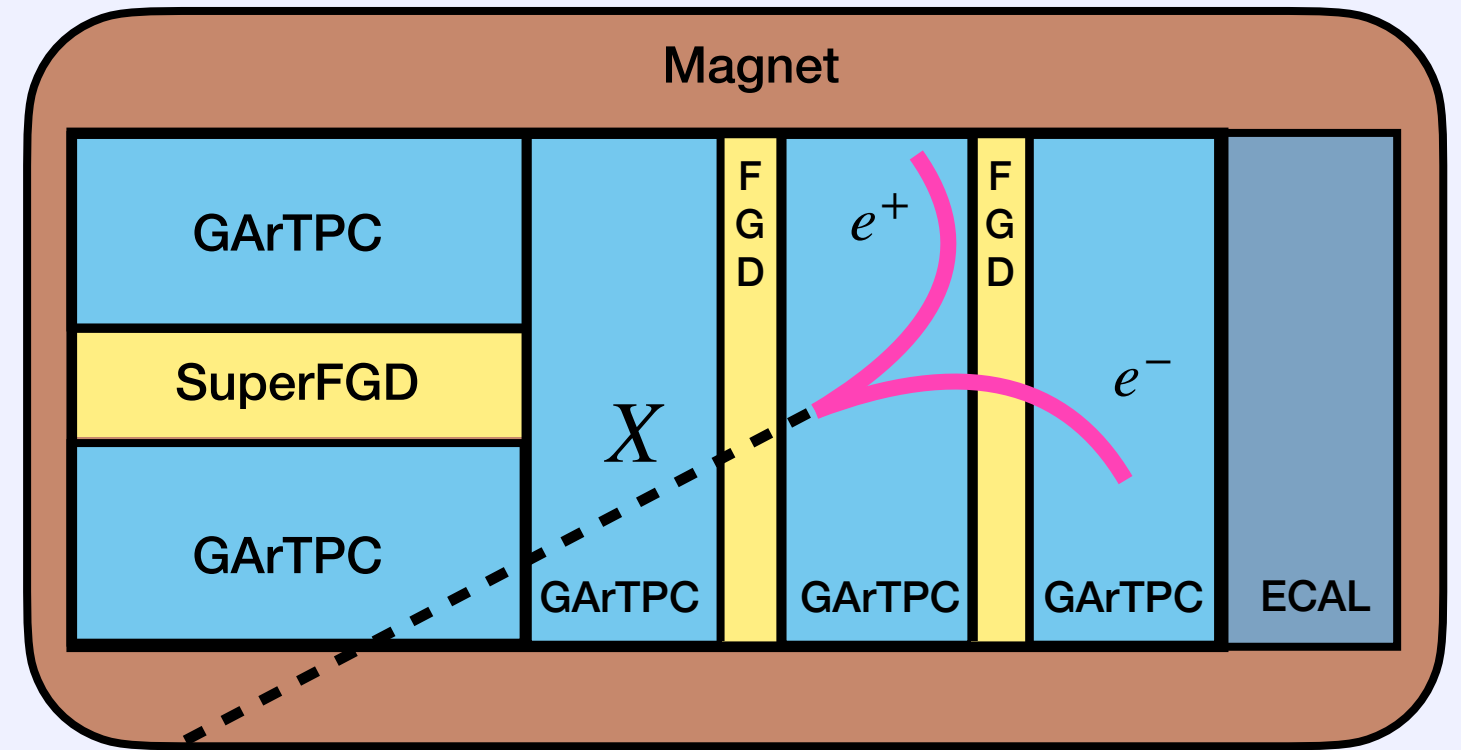
**KOTO:**

**Pros:** Low-density vol and low bkg  
**Cons:** Further away  
**Best for:**  $\pi^0$  and  $\gamma\gamma$



**ND280:**

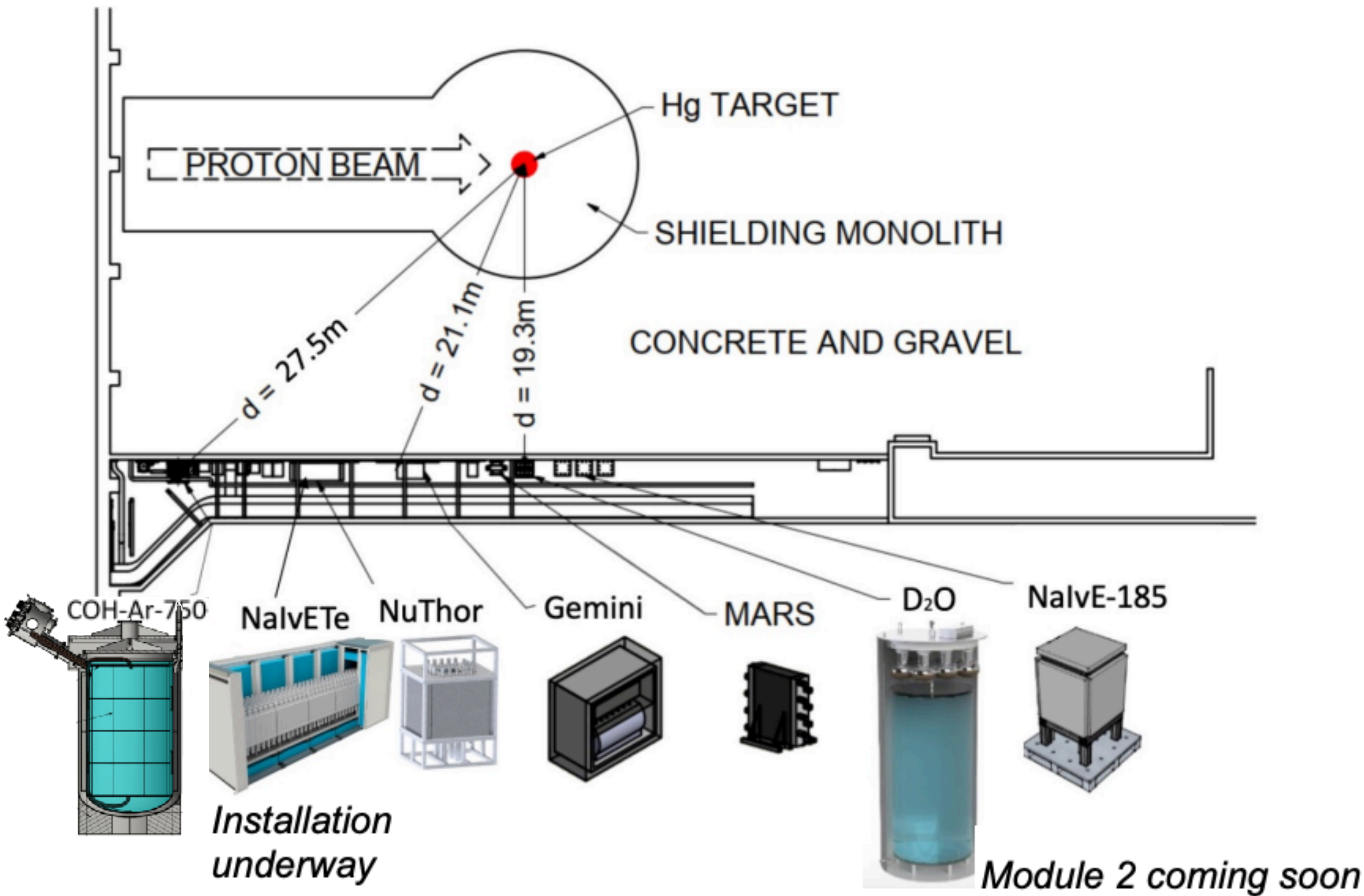
**Pros:** Low-density and magnetized  
**Cons:** Further away  
**Best for:** any charged final state



**ND280 @ 270 m**

# Opportunities at Oak Ridge

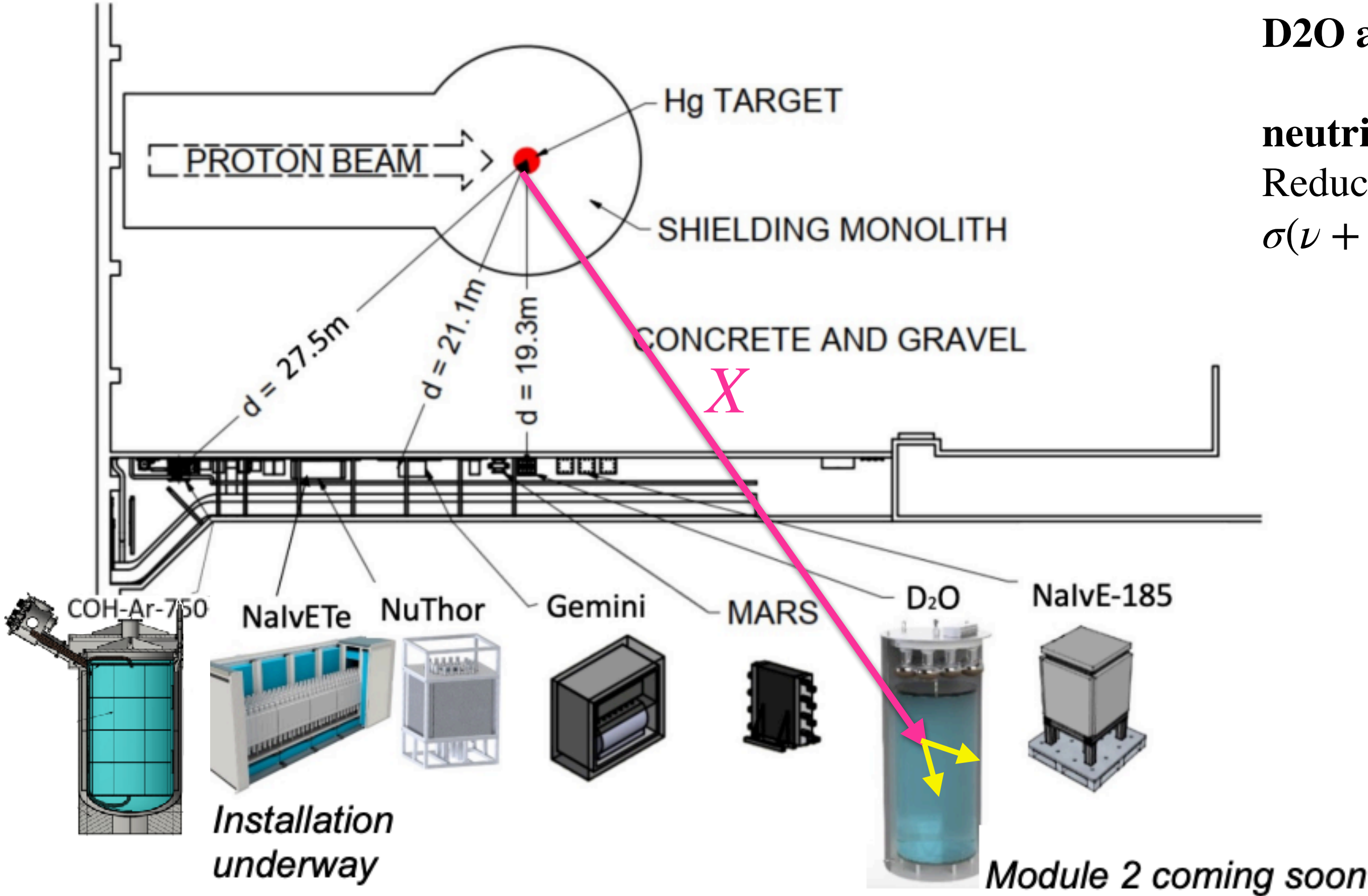
## SNS COHERENT detectors





# Opportunities at Oak Ridge

## SNS COHERENT detectors

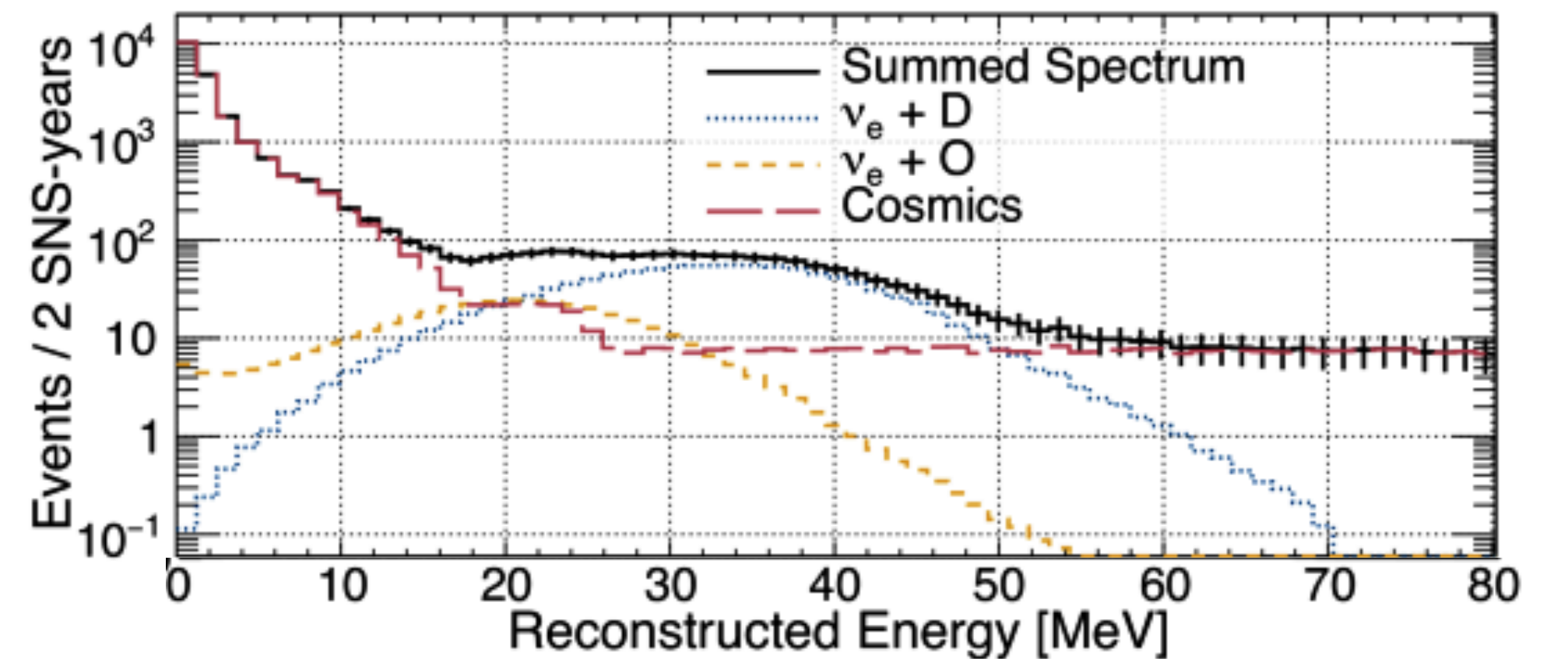
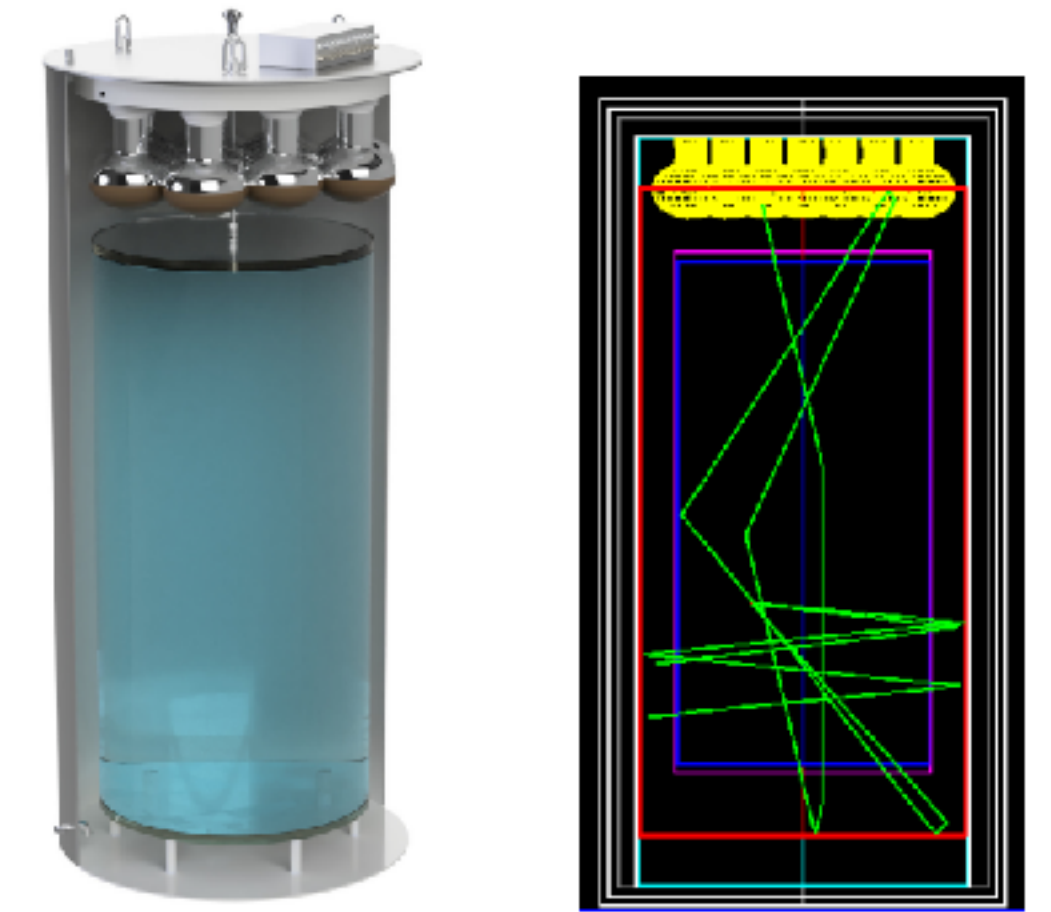


## D2O and H2O modules:

### neutrino flux uncertainty :

Reduction of 10%  $\rightarrow$  2-3% in 5 years  
 $\sigma(\nu + d \rightarrow e + p + p)$  known to 2%).

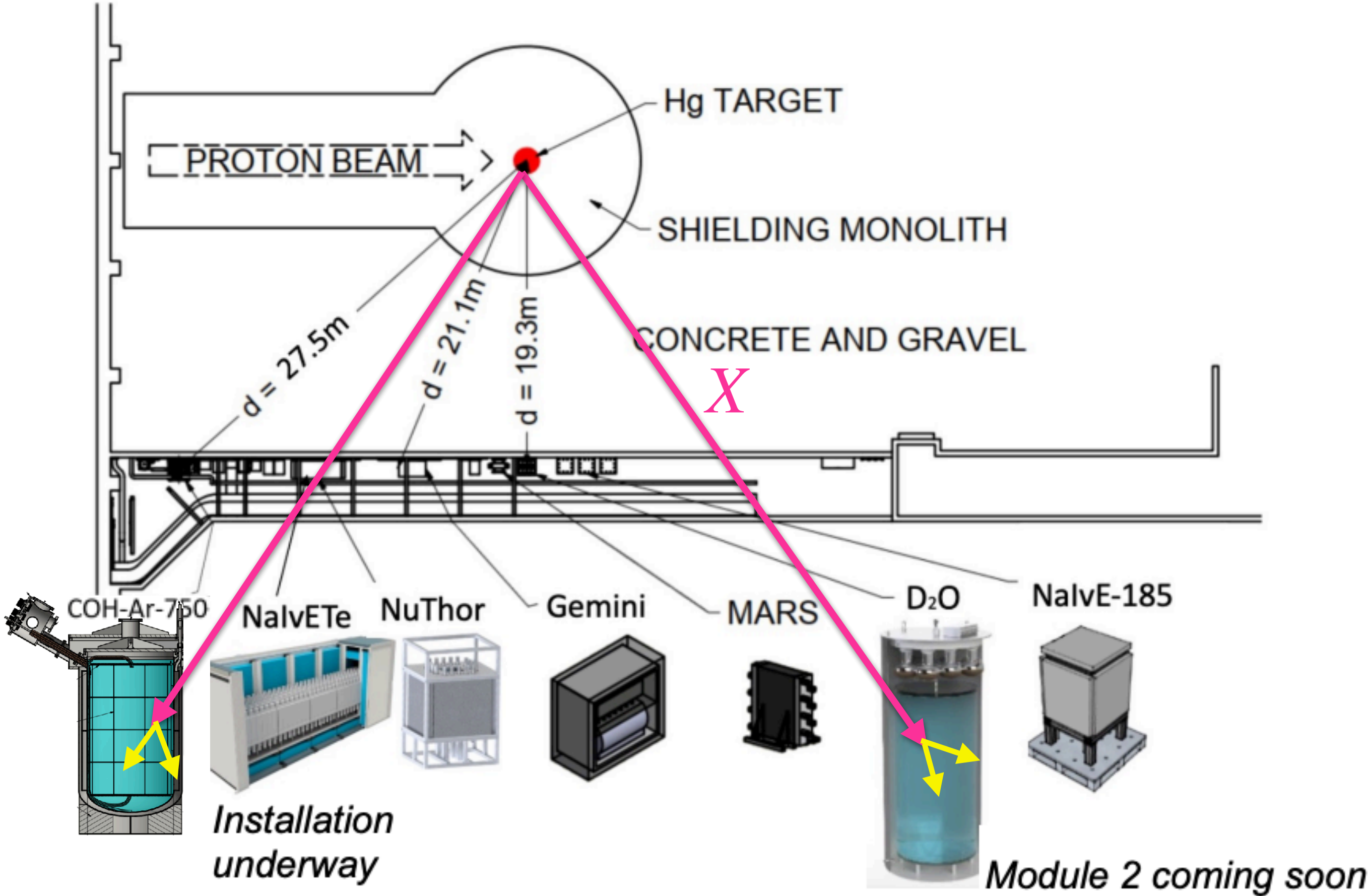
d = deuterium



\* D<sub>2</sub>O already has data and H<sub>2</sub>O is under construction.

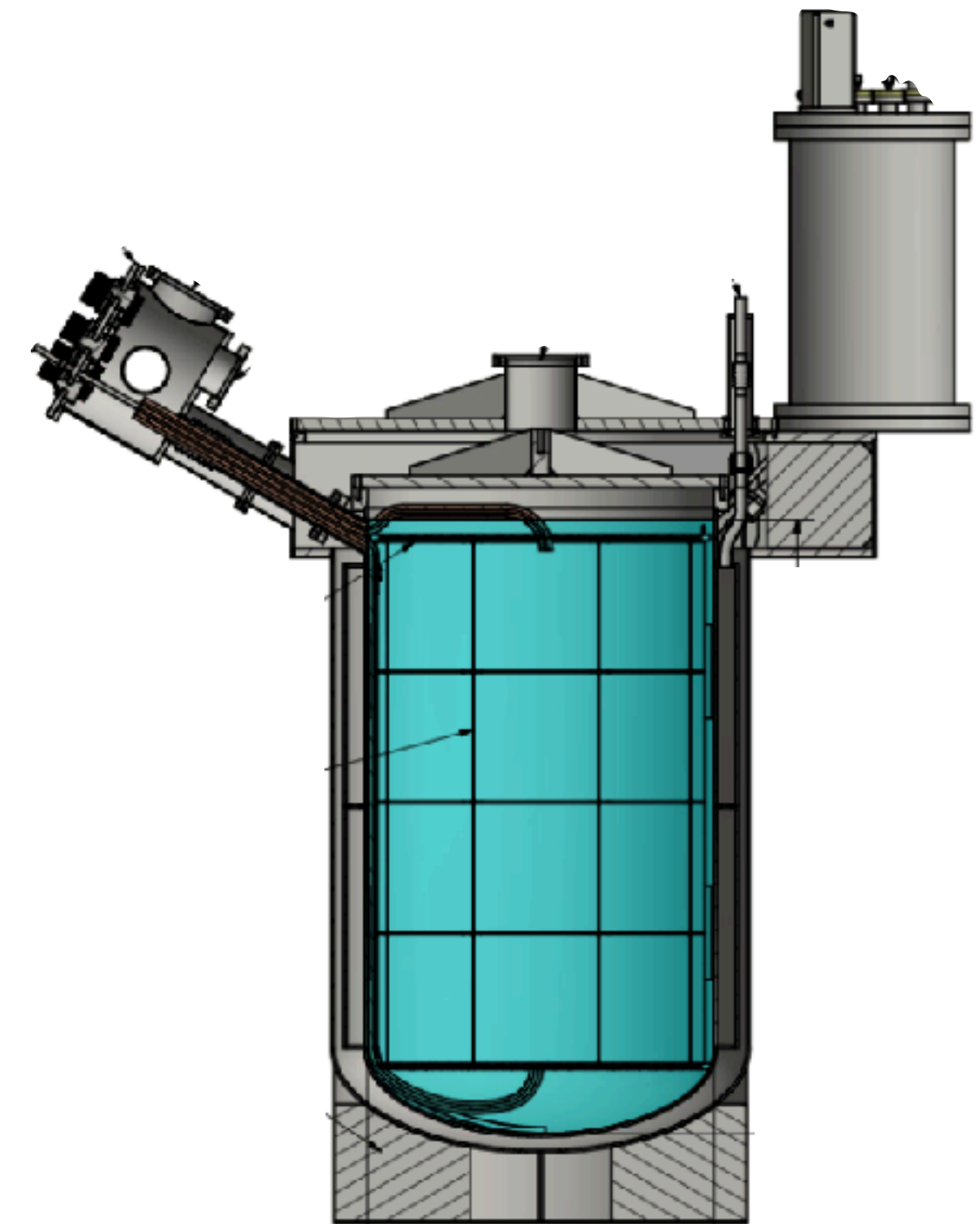
# Opportunities at Oak Ridge

## SNS COHERENT detectors



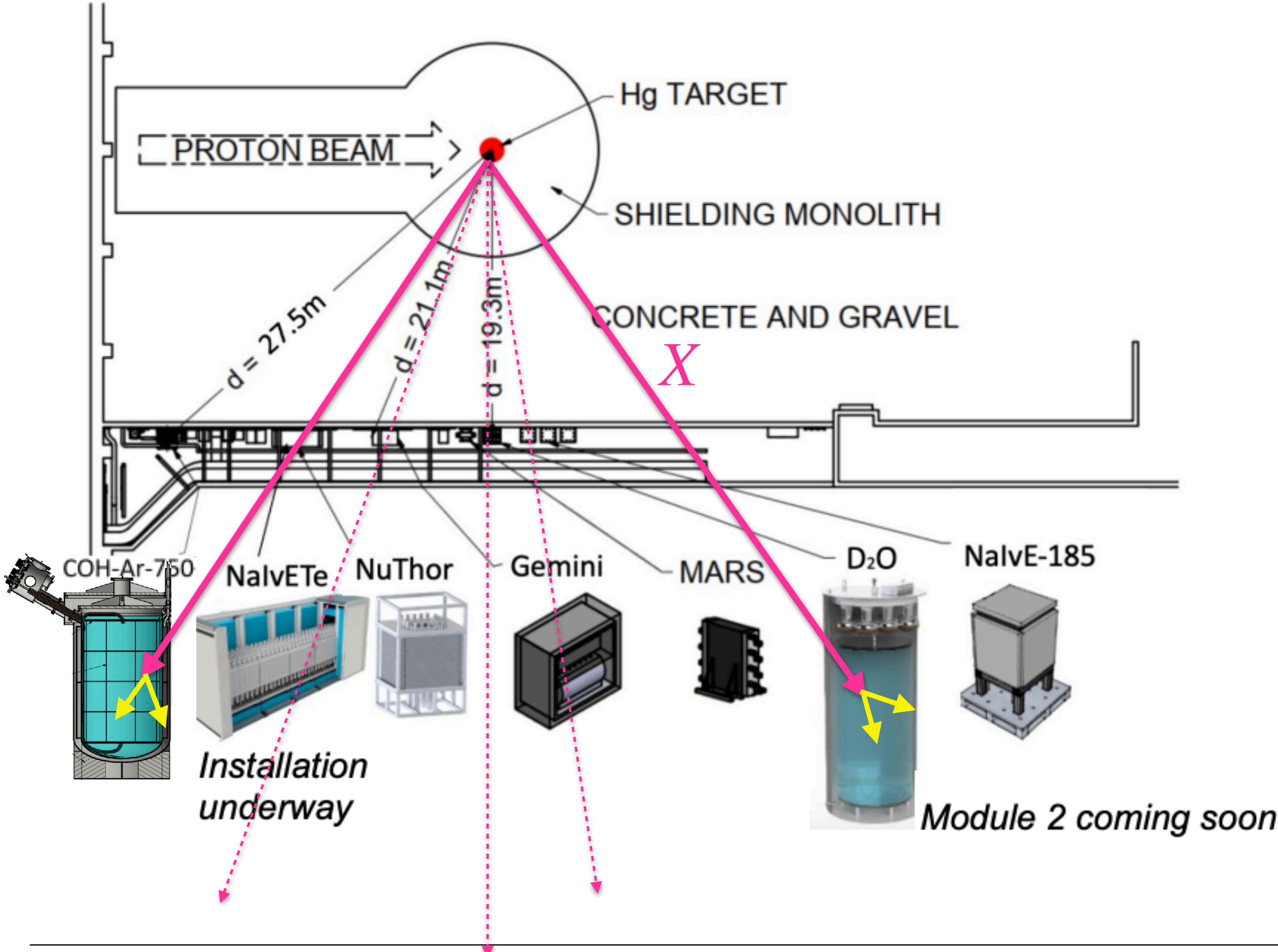
## COH-Ar-750

- Scintillation calorimeter
- 750 kg of purified liquid Ar
- Measure  $\nu_e$  Ar CEvNS



# Opportunities at Oak Ridge

## SNS COHERENT detectors



For this study, I will require:

**50 events/3 years of operation:**

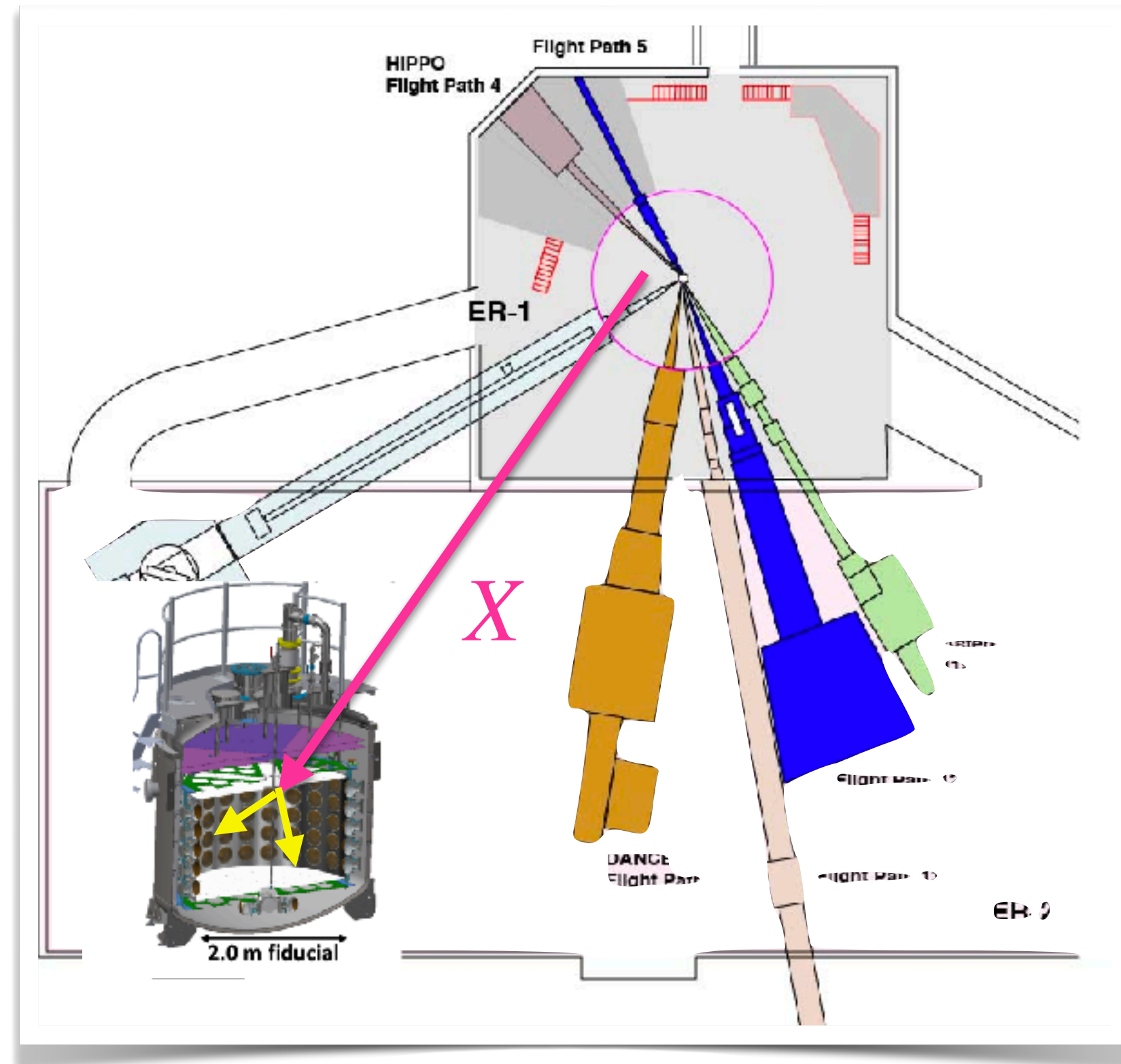
- 750 kg (LAr)
- 508 kg (H<sub>2</sub>O) + 549 kg (D<sub>2</sub>O)

Other detectors? For example:  
Surface deployment of 4-ton detector  
possible with  $L \sim 22\text{ m}$ .

Possible for PROSPECT-II?  
(under investigation with B. Littlejohn)

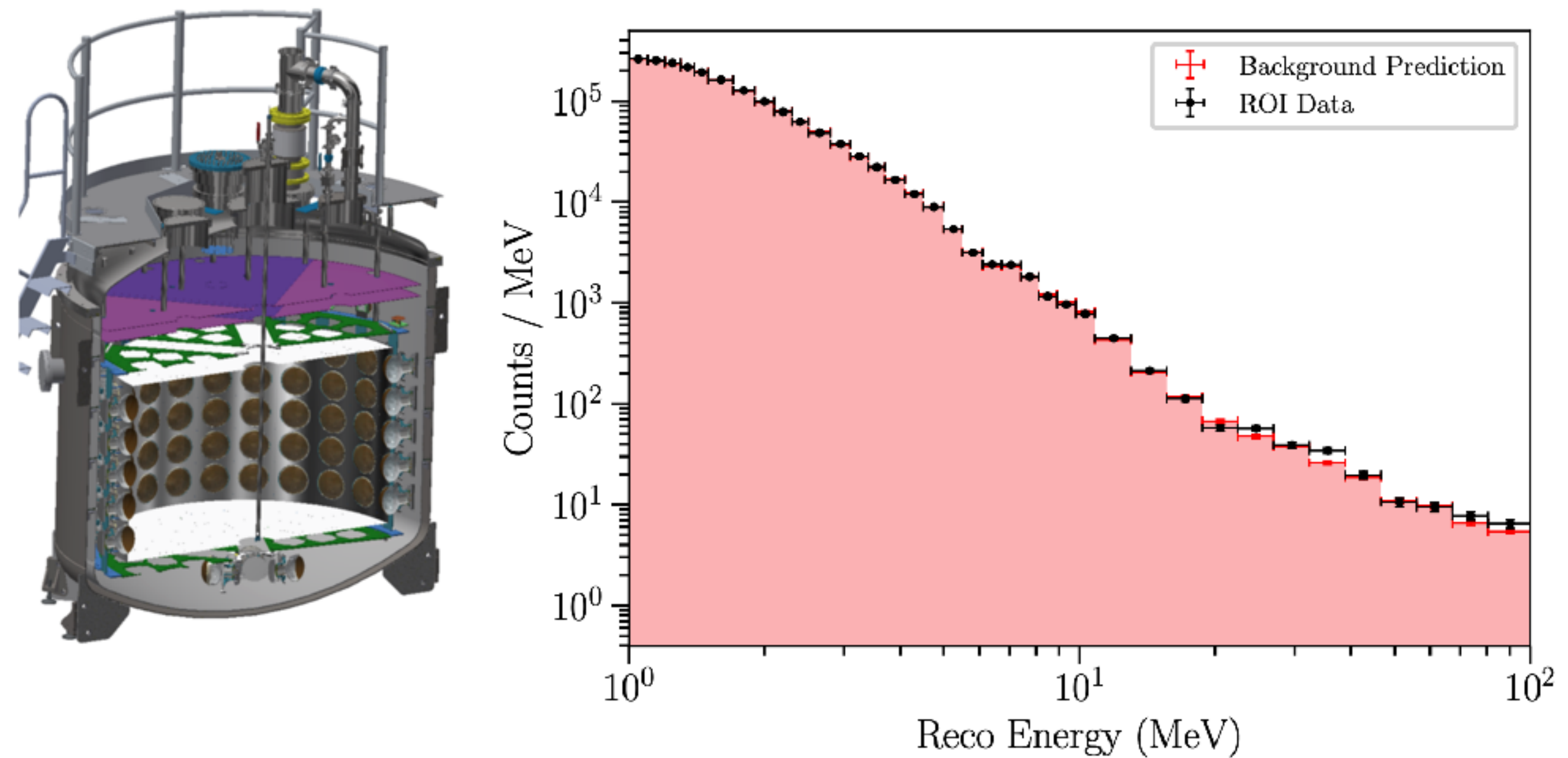
# Opportunities at Los Alamos

## LANSCCE Coherent-Captain-Mills (CCM)



**CCM:** 7 tons of LAr  
Detects scintillation of Ar

Rate of background events for CCM120 (120 inward pointing PMTs)



Projected 100 times smaller bkg rate for CCM200 (200 PMTs)

**In this study, we project sensitivity for  
40 events/3 years of operation**

# Outline:

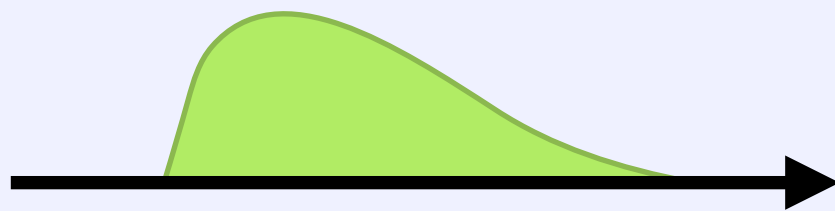
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# Timing profile of LLP signatures

Prompt  $\mathcal{O}(100)$  ns (on time from  $\pi^+/K^+$  decay)



$\mathcal{O}(2)$   $\mu$ s delayed ( $\mu^+$  time)




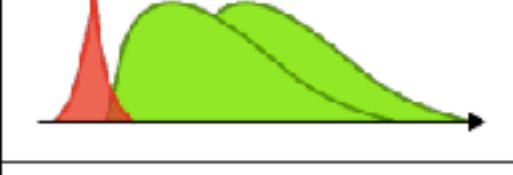
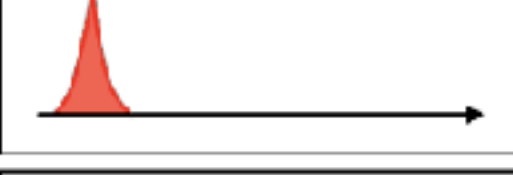

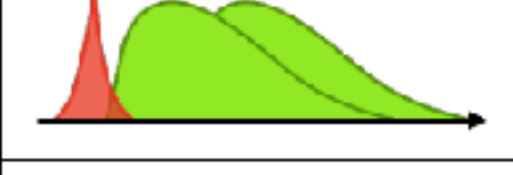


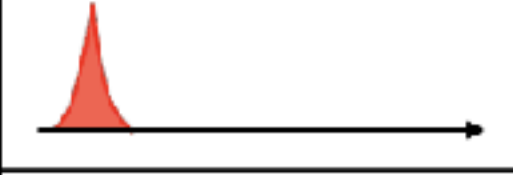

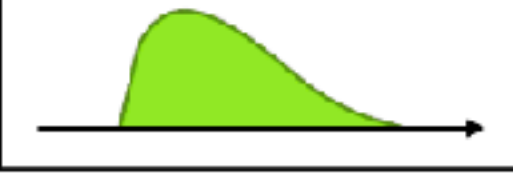



Model	Production	Decay	Timing signature	J-PARC Detector
Heavy Neutral Leptons	$\mu^+ \rightarrow e^+ \nu N$	$N \rightarrow \nu e^+ e^-$		ND280
	$\pi^+/K^+ \rightarrow \ell N$	$N \rightarrow \nu e^+ e^-$		ND280
		$N \rightarrow \nu \mu^+ e^- / \pi^+ e^-$		JSNS <sup>2</sup> and ND280
		$N \rightarrow \nu \mu^+ \mu^- / \pi^+ \mu^-$		JSNS <sup>2</sup> and ND280
		$N \rightarrow \nu \pi^0$		KOTO
Higgs Portal Scalar	$K^+ \rightarrow \pi^+ S$	$S \rightarrow e^+ e^-$		ND280
		$S \rightarrow \mu^+ \mu^- / \pi^+ \pi^-$		JSNS <sup>2</sup> and ND280
		$S \rightarrow \pi^0 \pi^0$		KOTO
Muon Portal Scalar	$\mu^+ \rightarrow e^+ \nu \nu S_M$	$S_M \rightarrow \gamma \gamma$		KOTO
ALP: Higgs Coupling	$K^+ \rightarrow \pi^+ a_\phi$	$a_\phi \rightarrow e^+ e^-$		ND280
		$a_\phi \rightarrow \mu^+ \mu^-$		JSNS <sup>2</sup> and ND280
ALP: Flavor Violating	$\mu^+ \rightarrow e^+ a_{FV}(\gamma)$	$a_{FV} \rightarrow e^+ e^-$		ND280
ALP: Weak Violating	$\pi^+ \rightarrow e^+ \nu_e a_{WV}$	$a_{WV} \rightarrow e^+ e^-$		ND280

# Timing profile of LLP signatures

A lot of these come from  $\mu^+$  and  $\pi^+$  decays.

$\sim 1$  GeV  $p^+$  beams are in the game.

Model	Production	Decay	Timing signature	J-PARC Detector
Heavy Neutral Leptons	$\mu^+ \rightarrow e^+ \nu N$	$N \rightarrow \nu e^+ e^-$		ND280
	$\pi^+ / K^+ \rightarrow \ell N$	$N \rightarrow \nu e^+ e^-$		ND280
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# Outline:

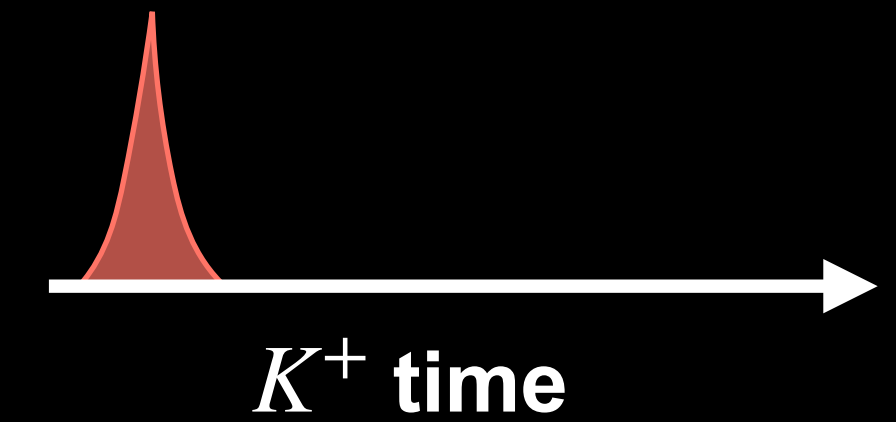
- 1) Spallation (neutron) sources: what and why?
- 2) Opportunities with neutrino detectors
- 3) Long-lived particles below  $K$ ,  $\pi$ , and  $\mu$  masses

Higgs portal scalar: the simplest case



# Long-lived particles at spallation sources

## Higgs portal scalar



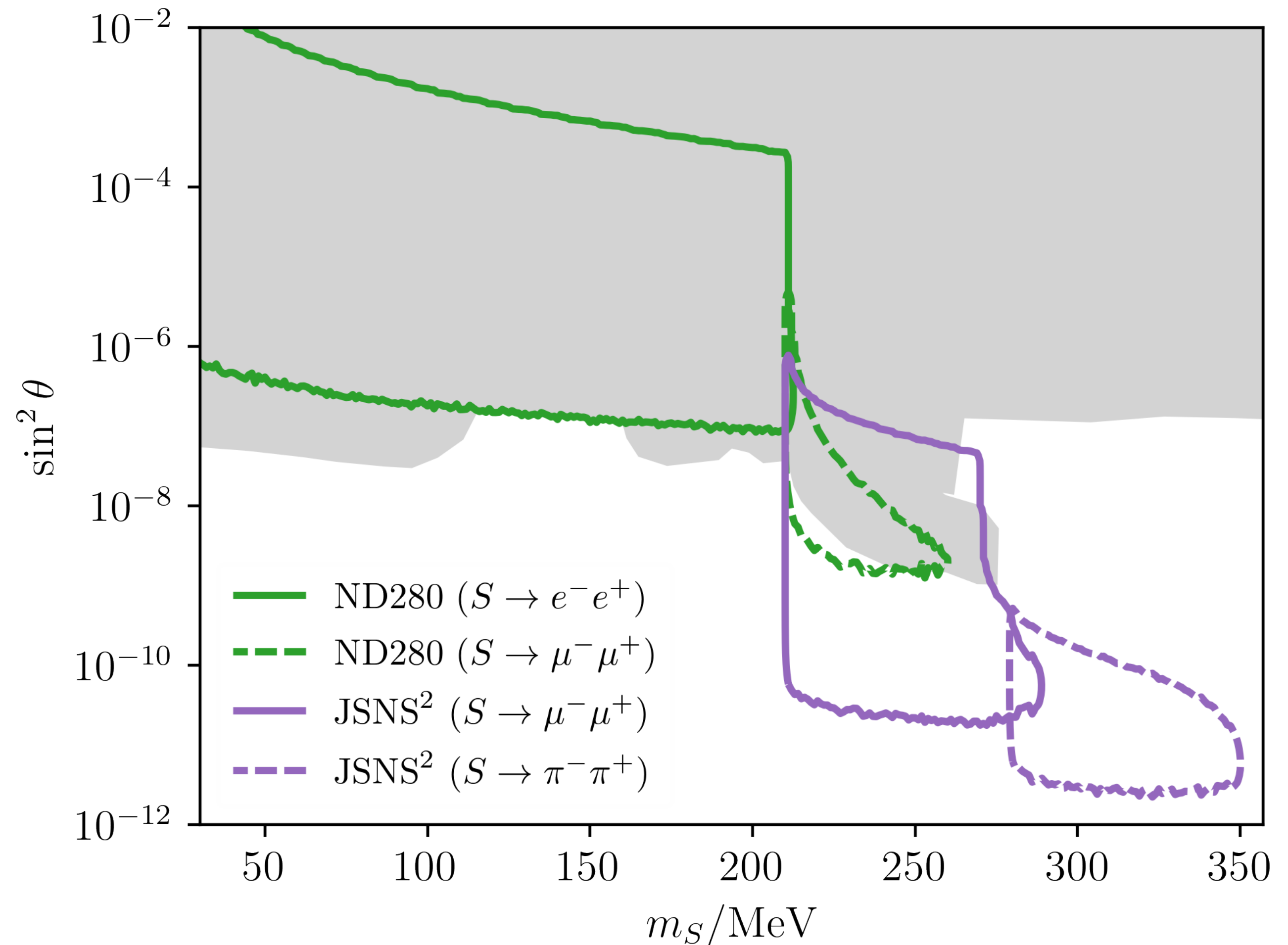
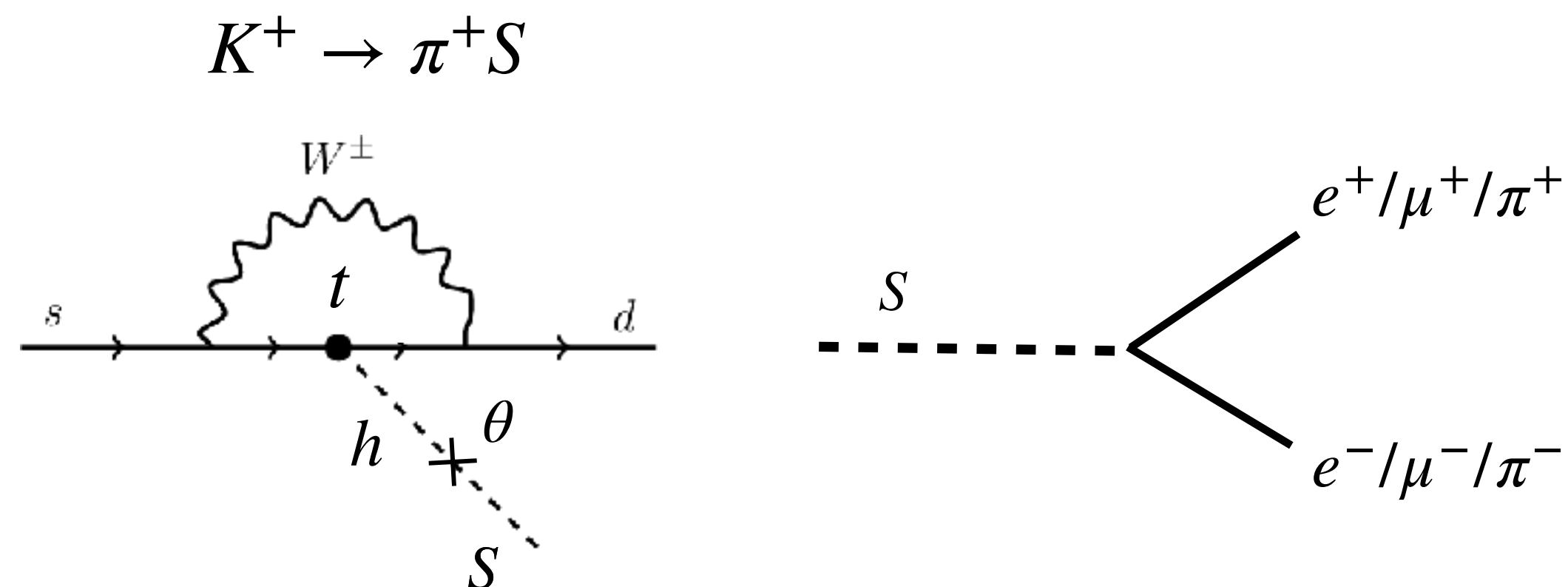
Arguably the simplest extension of the SM:

Singlet scalar particle  $S$  that mixes with the Higgs boson  
a.k.a. Higgs Portal Scalar (HPS).

Production (almost) exclusively through  $K^+$  decays.

J-PARC is most well suited for this.

(+ accelerators like T2K and FNAL's SBN program)



## Outline:

- 1) Spallation (neutron) sources: what and why?
- 2) Opportunities with neutrino detectors
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Axion-like particle: flavor-violating couplings

# Long-lived particles at spallation sources

## A lepton-flavor-violating axion-like particle (LFV ALP)

### Light goldstone boson to probe lepton flavor violation

Complementary to  $\mu \rightarrow e$  searches (Mu2e, Mu3e, MEG-II).

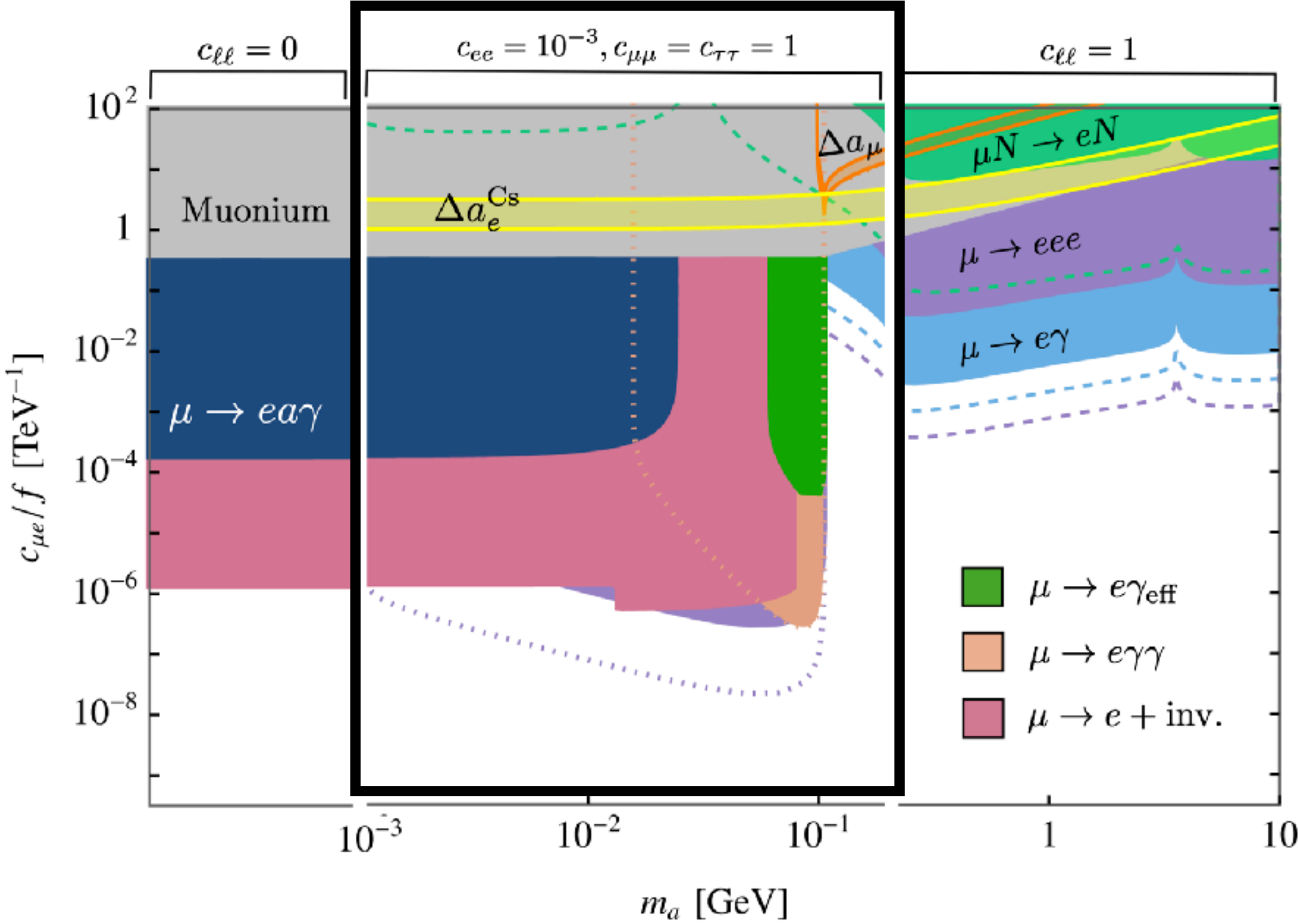
Indirect limits constraint products of couplings (not a killer).

Direct limit:  $\mathcal{B}(\mu^+ \rightarrow e^+ a_{\text{fV}}) \lesssim 10^{-5}$  if  $a$  is long-lived.

That would lead to about  $10^{14} - 10^{15} a_{\text{fV}}/\text{year}$  in typical spallation sources...



$$j_{\text{FV}}^\rho = y_{e\mu}^L \frac{\overline{L_e H} \gamma^\rho H L_\mu}{\Lambda^2} + c_{e\mu}^R \bar{e}_R \gamma^\rho \mu_R \xrightarrow{\text{EW}} \bar{e} \gamma^\rho (c_{e\mu}^L P_L + c_{e\mu}^R P_R) \mu$$



M. Bauer et al, 2110.10698 and L. Calibbi et al 2006.04795

# Long-lived particles at spallation sources

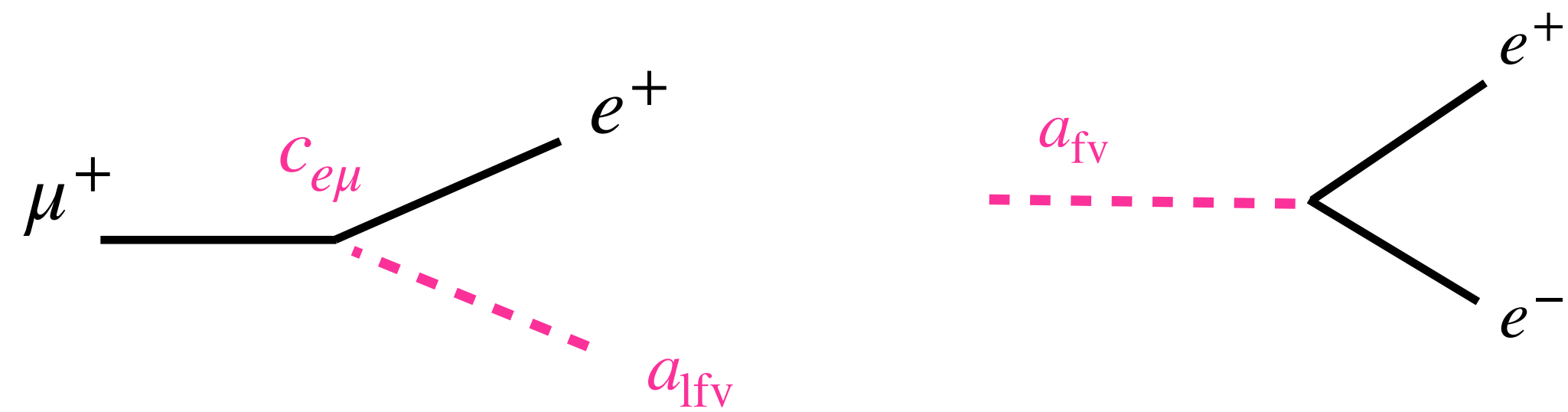
## A lepton-flavor-violating axion-like particle (LFV ALP)



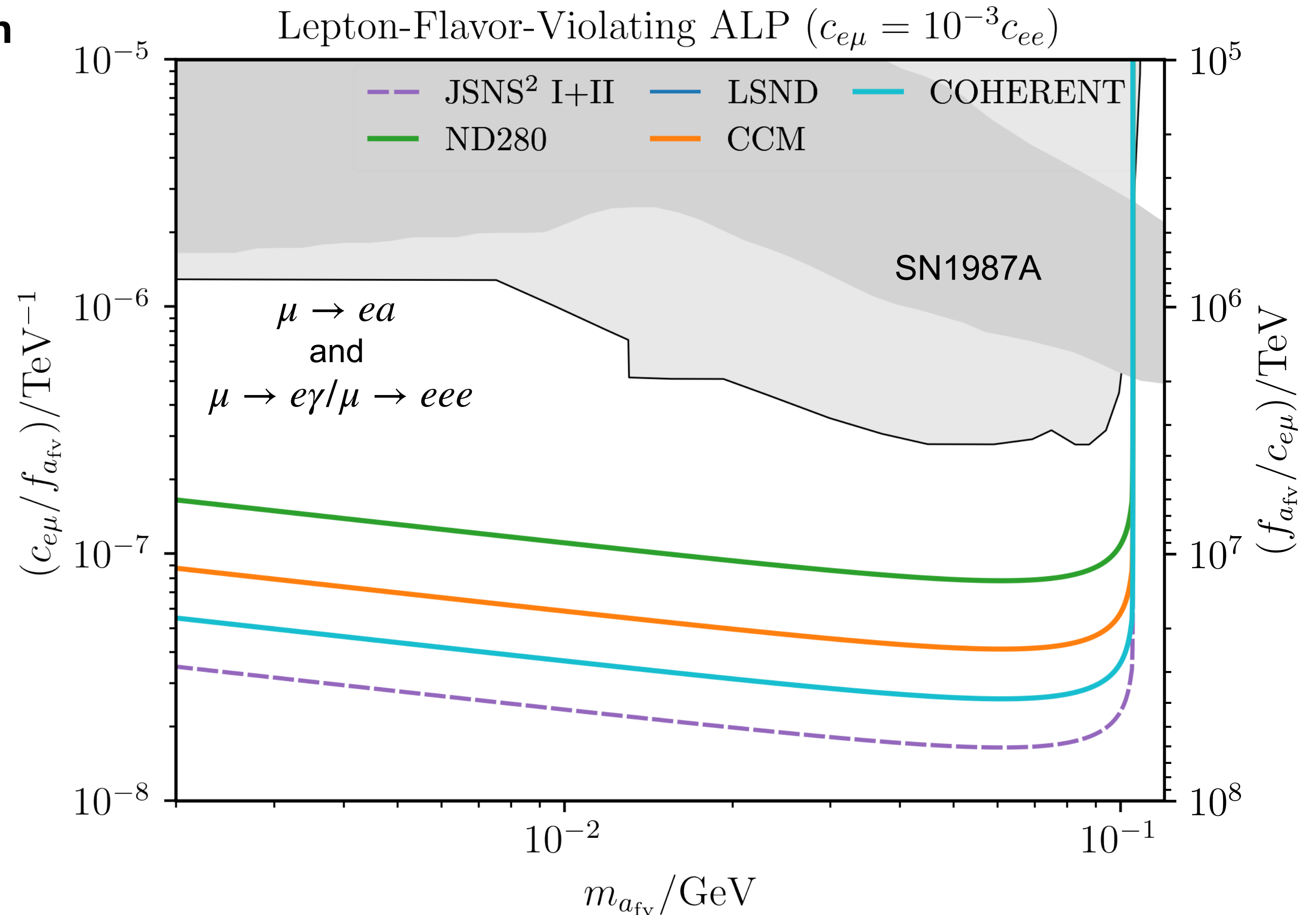
### Light goldstone boson to probe lepton flavor violation

Far more sensitive than indirect probes in this region.

This is a high-energy final state  $\simeq m_\mu/2$ , so LSND “threw it away”. Easy to see with new searches.



\* Supernovae will also be impacted due to large  $\mu$  content (afaik, this has not been worked out yet).



# Outline:

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- 3) Long-lived particles below  $K$ ,  $\pi$ , and  $\mu$  masses

Minimal muonphilic scalars

# Long-lived particles at spallation sources

## Muonphilic scalar

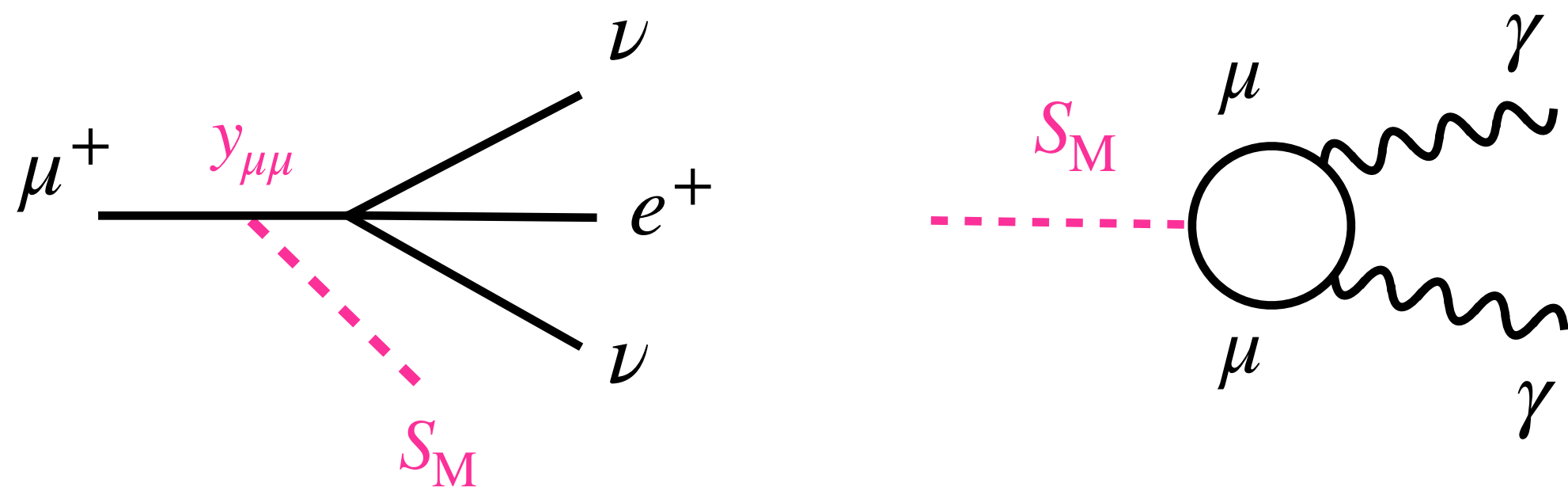
### Exotic force that couples only to muons

A popular effective model for discrepancy(?) in  $(g - 2)_\mu$

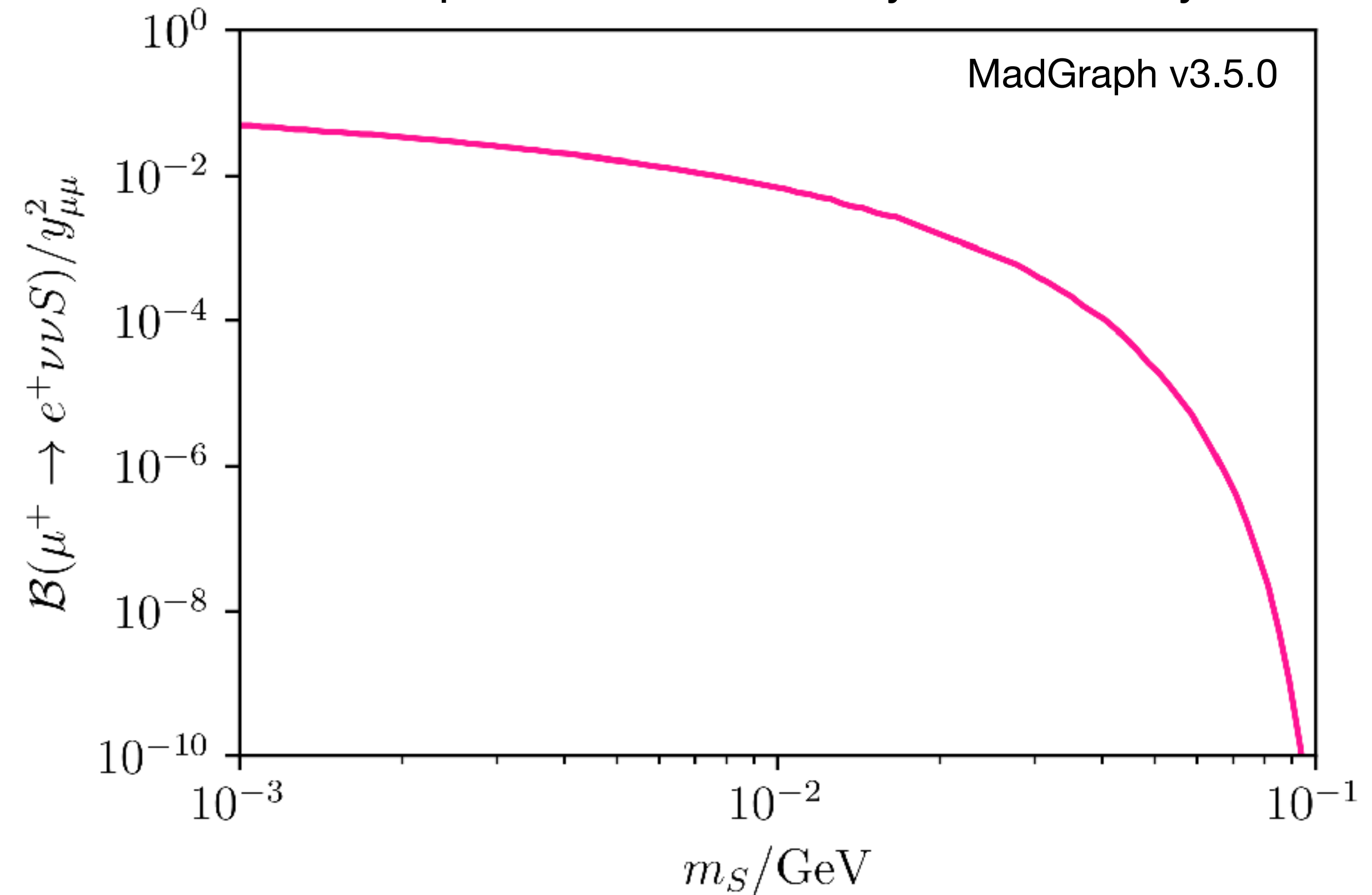
$$-\mathcal{L}_M \supset \frac{Y_{\mu\mu}}{\Lambda} S_M \bar{L}_\mu H \mu_R + \text{h.c.} \xrightarrow{\text{EW}} y_{\mu\mu} S_M \bar{\mu} \mu.$$

Very hard to constrain — no coupling to neutrinos.

Below dimuon threshold ( $m_S < 2m_\mu$ ), the scalar is long-lived:



Scalar production in 4-body muon decays:



# Long-lived particles at spallation sources

## Muonphilic scalar

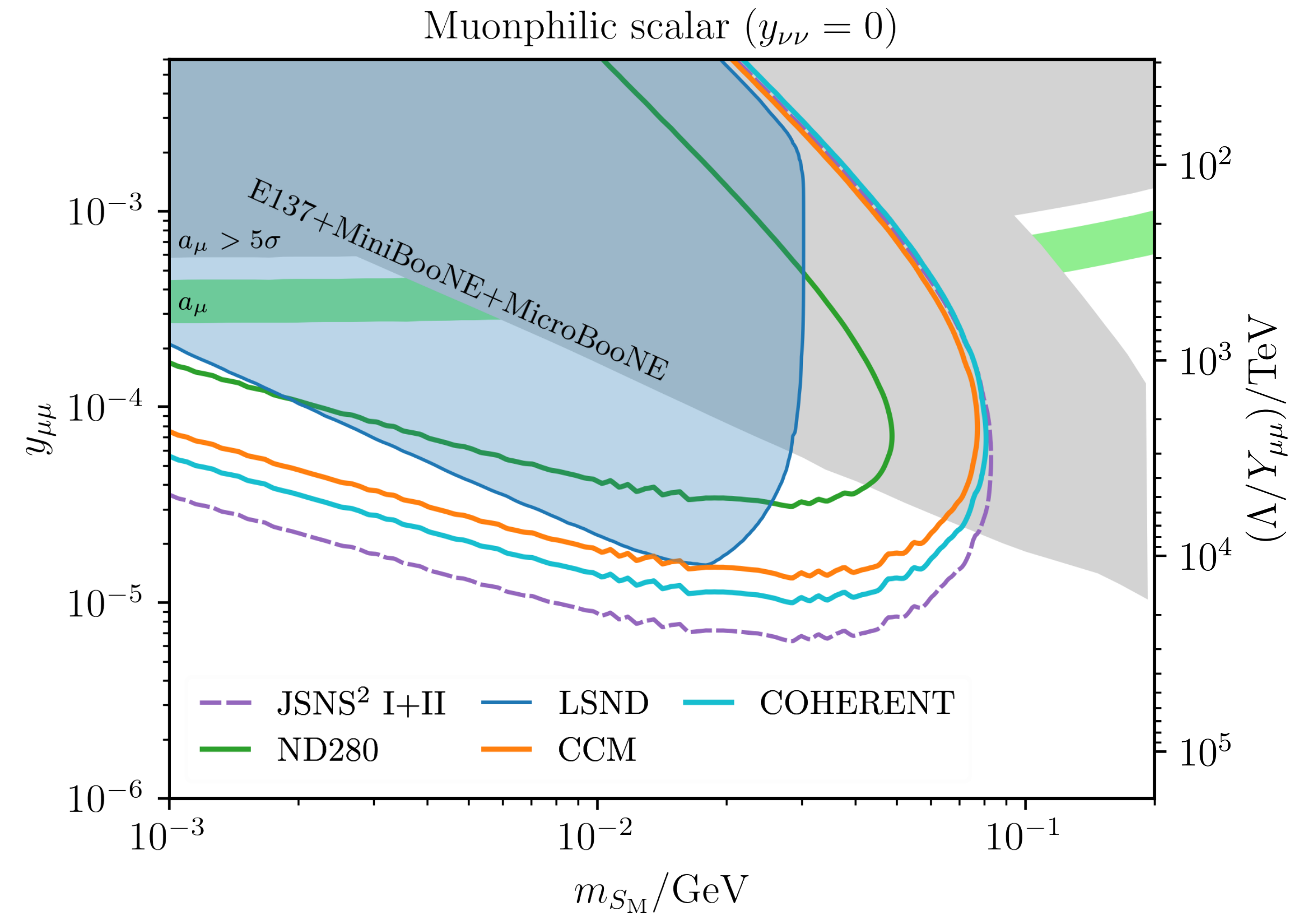
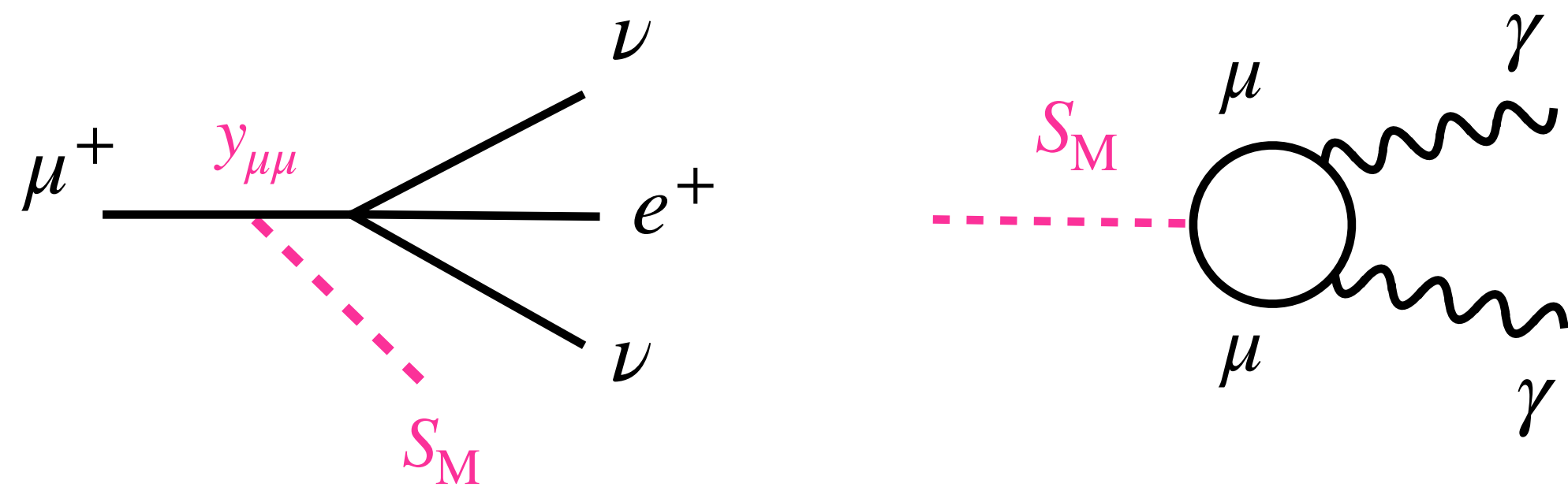


### Exotic force that couples only to muons

Setting new limits with LSND recast — no g-2 at low masses.

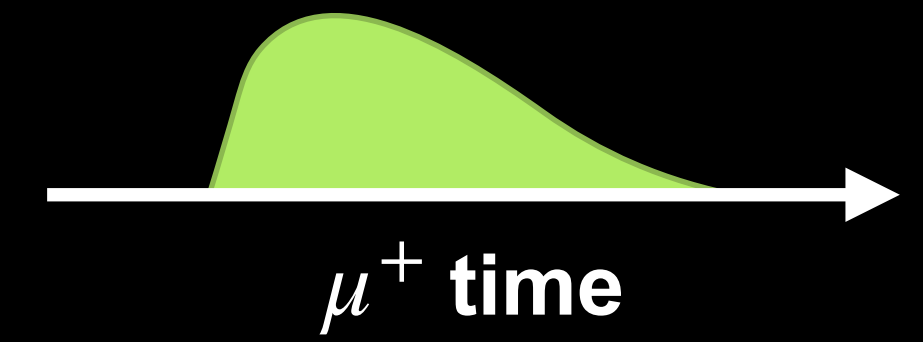
Improvements expected in all cases.

(Hard for JSNS<sup>2</sup> as it is only a single flash).



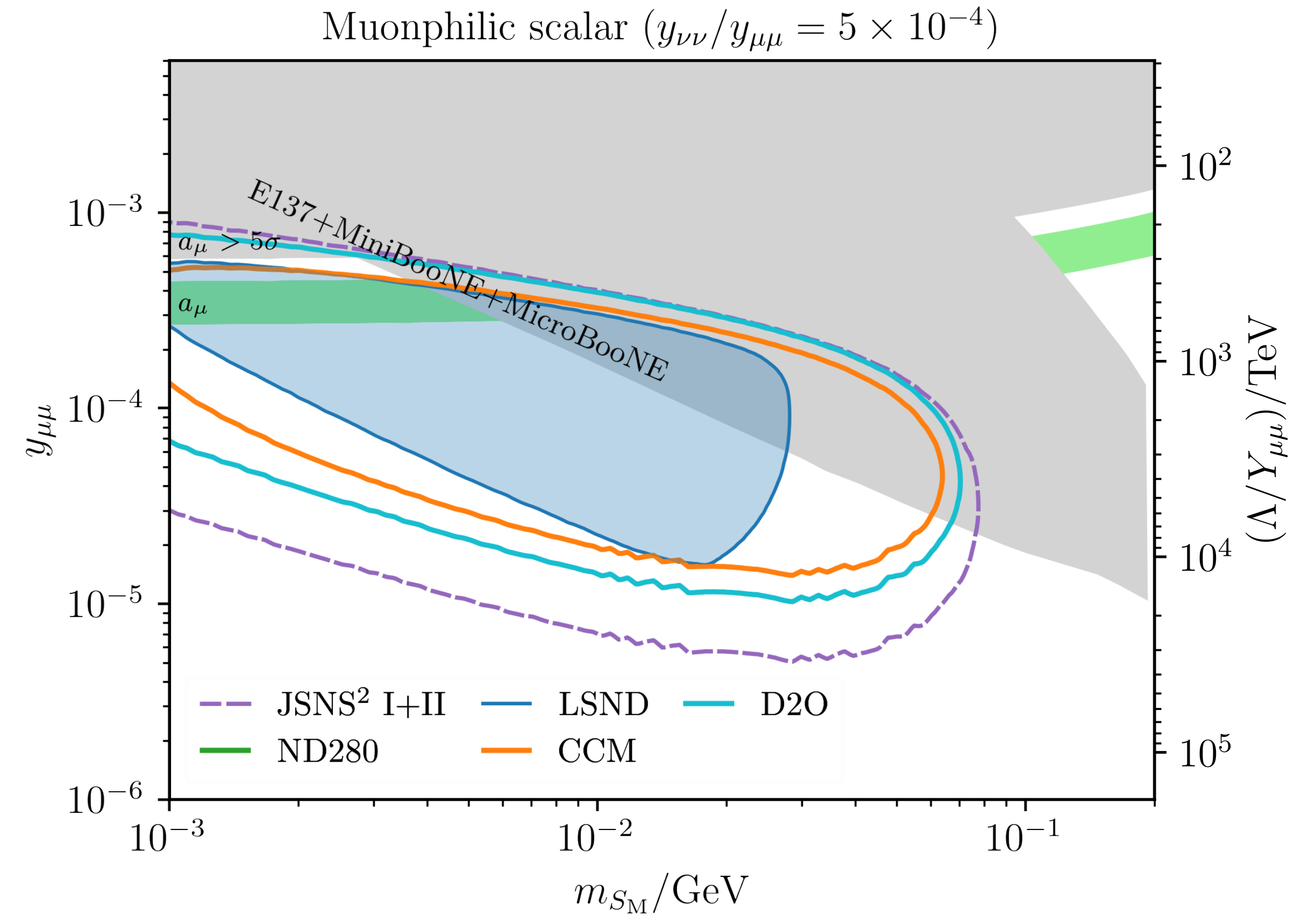
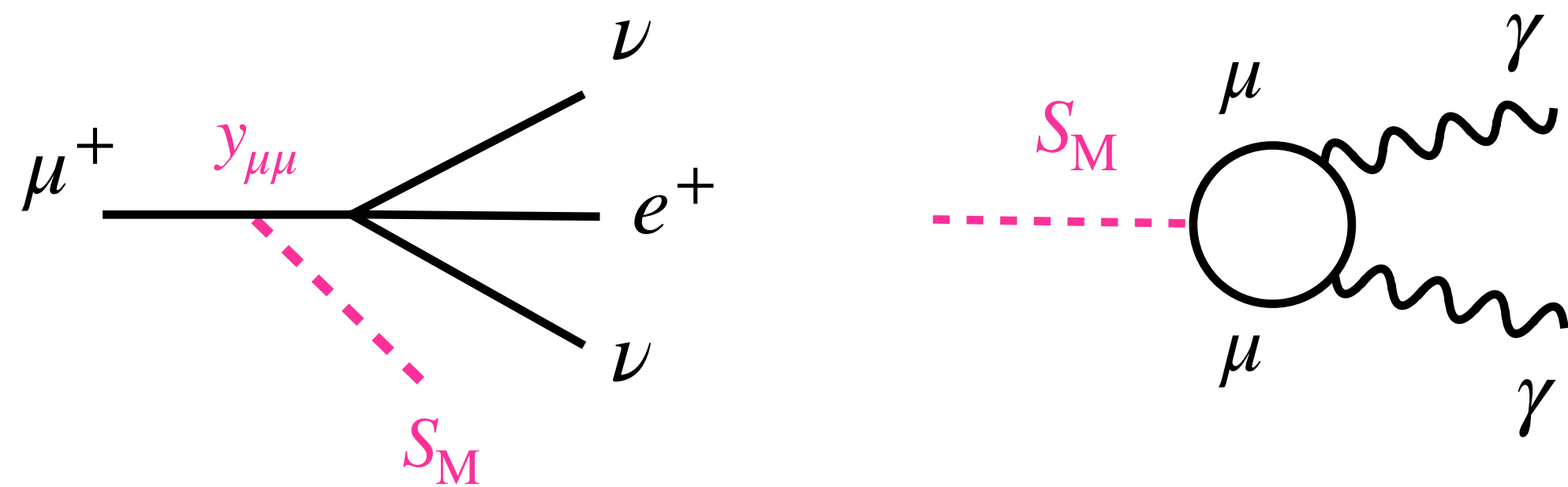
# Long-lived particles at spallation sources

## Muonphilic scalar



### Exotic force that couples only to muons

Turning on an invisible branching ratio for  $S_M$  may be possible, but requires a very specific hierarchy of couplings.





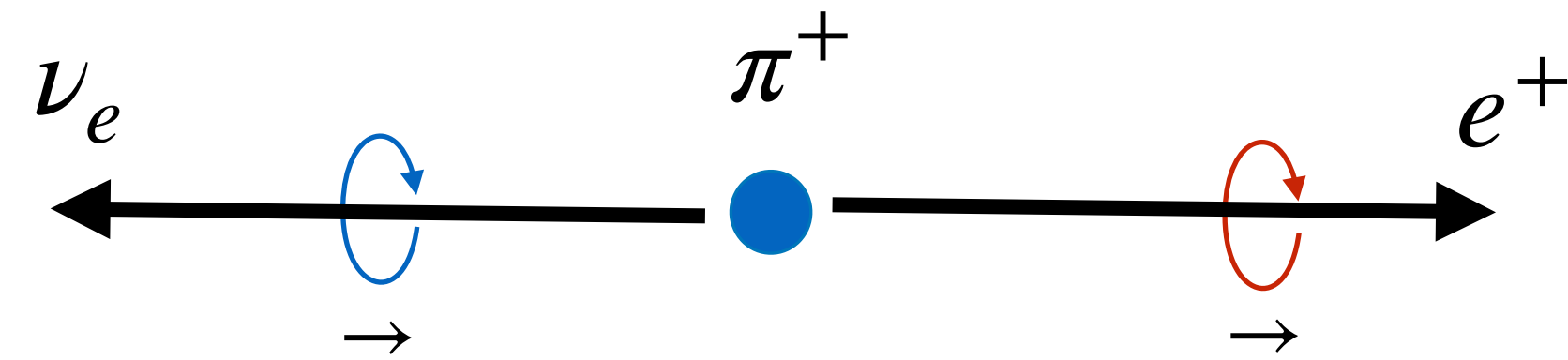
# Outline:

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How about dark particle production in charged pion decay?

# Radiative pion decay

## Pion decays — helicity flip



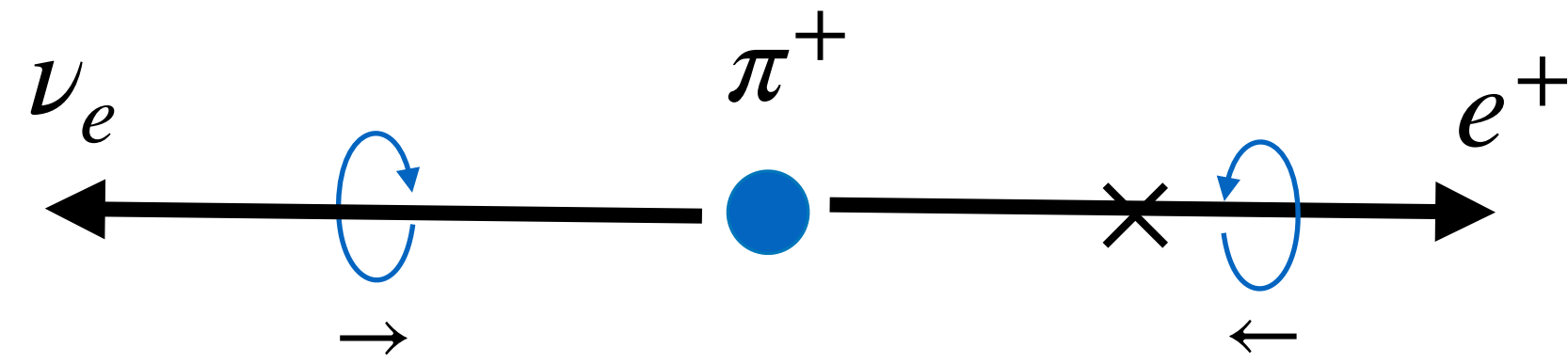
Weak interactions in the SM are left-handed.

Angular momentum needs to be conserved.

The pion is a spin-0 particle, so **neutrino and positron helicities have to be anti-aligned!**

# Radiative pion decay

## Pion decays — helicity flip



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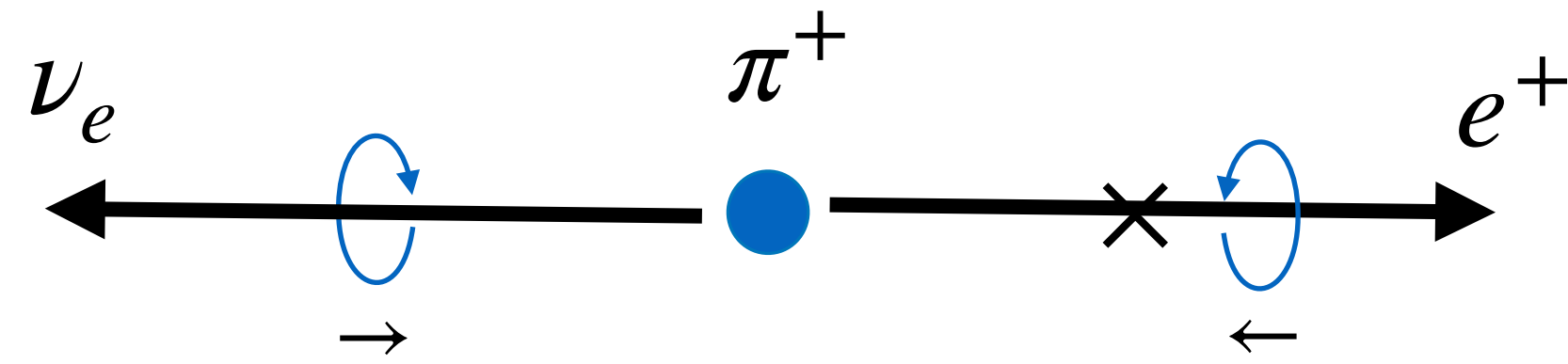
The pion is a spin-0 particle, so **neutrino and positron helicities have to be anti-aligned!**

**Helicity suppression:**

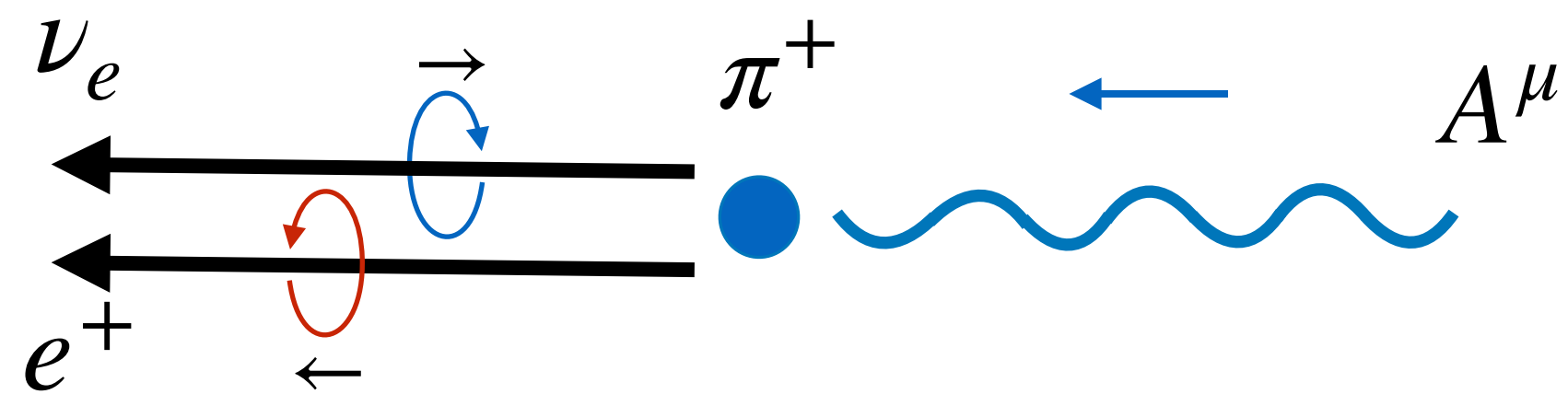
$$\Gamma \propto G_F^2 f_\pi^2 m_\pi^3 \times \left( \frac{m_e^2}{m_\pi^2} \right) \rightarrow \mathcal{B} \sim 10^{-4} \text{ branching ratio for } \pi^+.$$

# Radiative pion decay

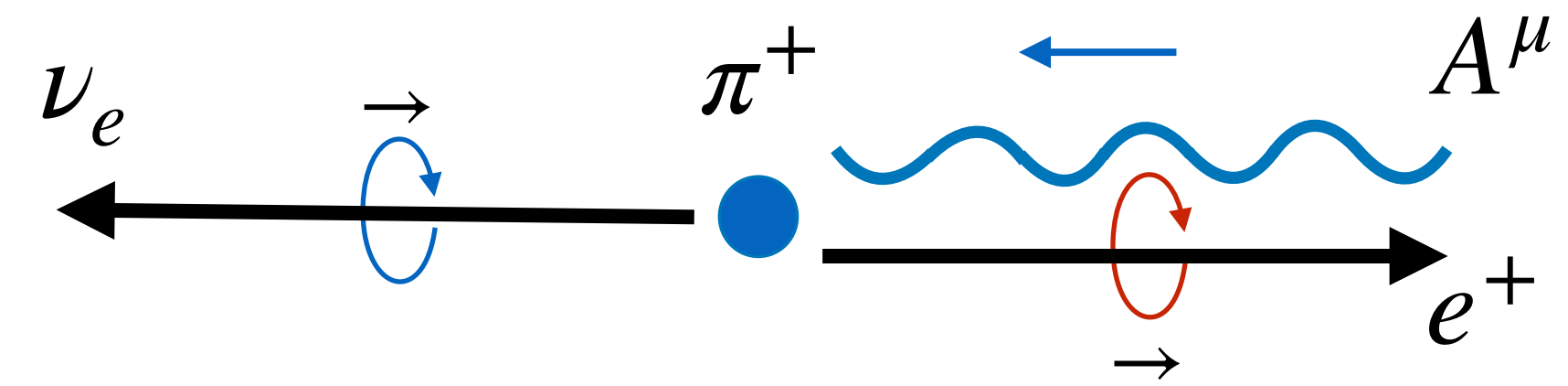
## Pion decays — helicity flip



Well, there's an easy way out: make it a three-body decay! **Radiative pion decays:**



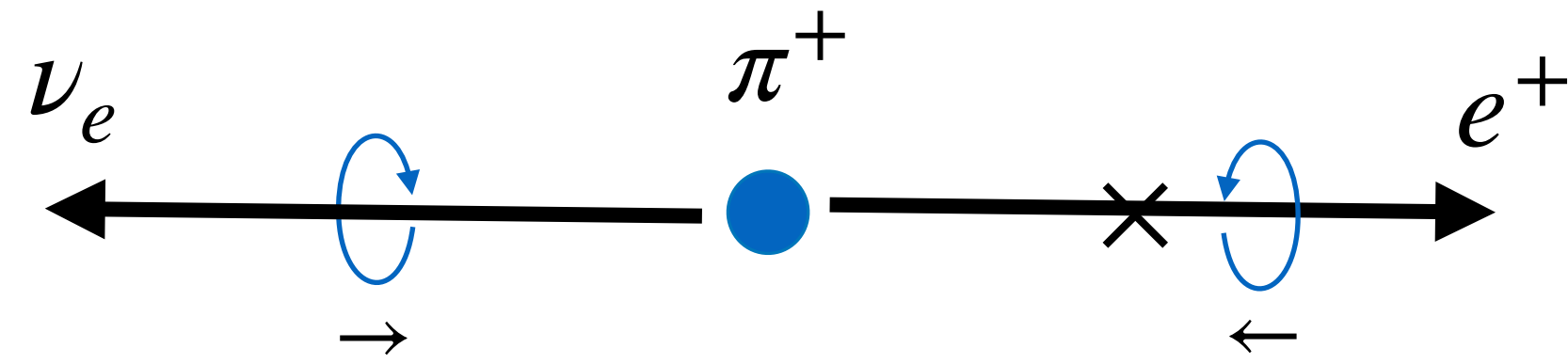
No helicity flip.



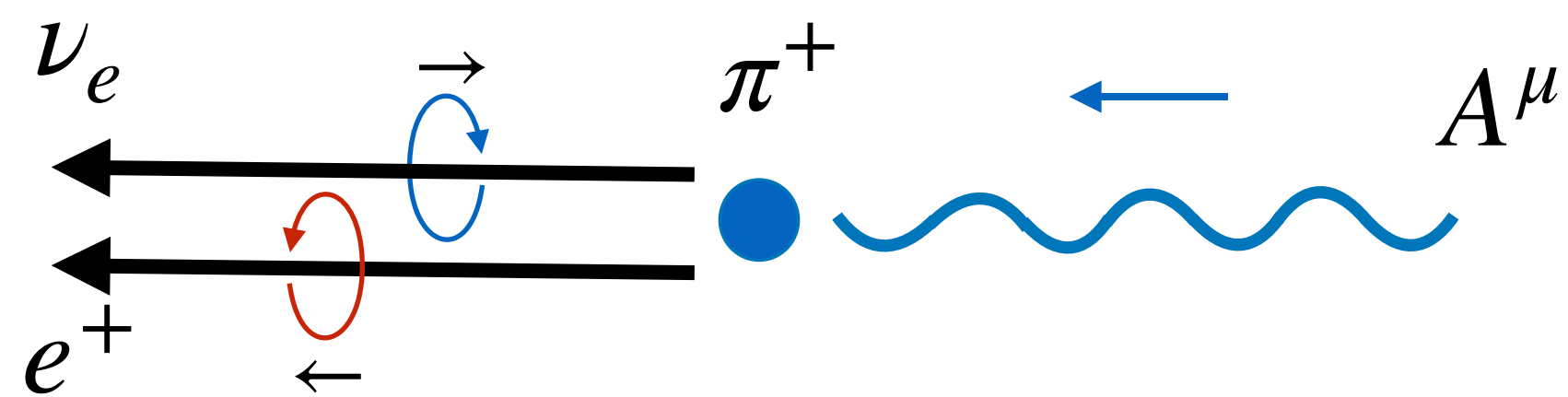
No helicity flip.

# Radiative pion decay

## Pion decays — helicity flip

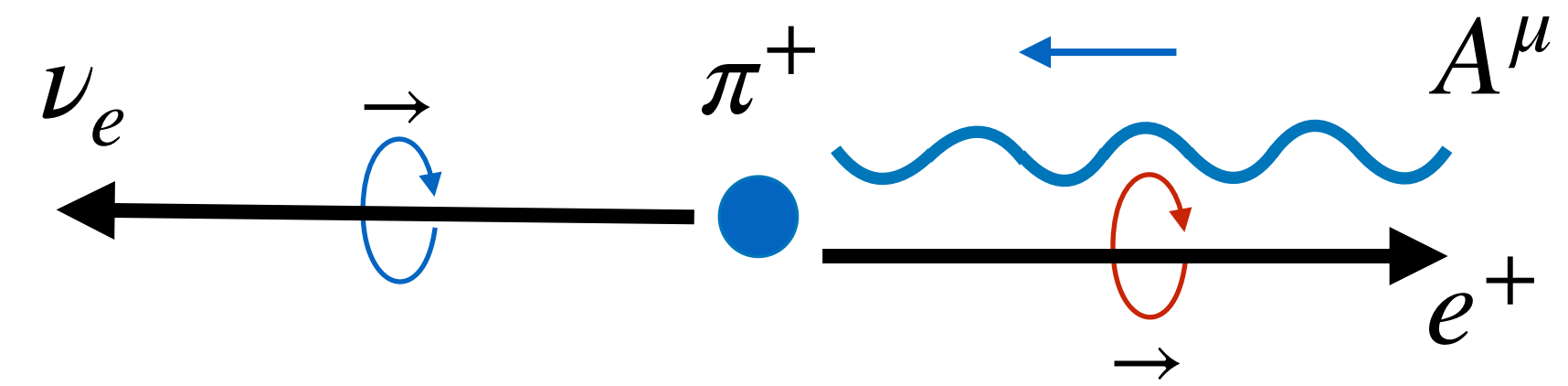


Well, there's an easy way out: make it a three-body decay! **Radiative pion decays:**



No helicity flip.

Feels like cheating...



No helicity flip.

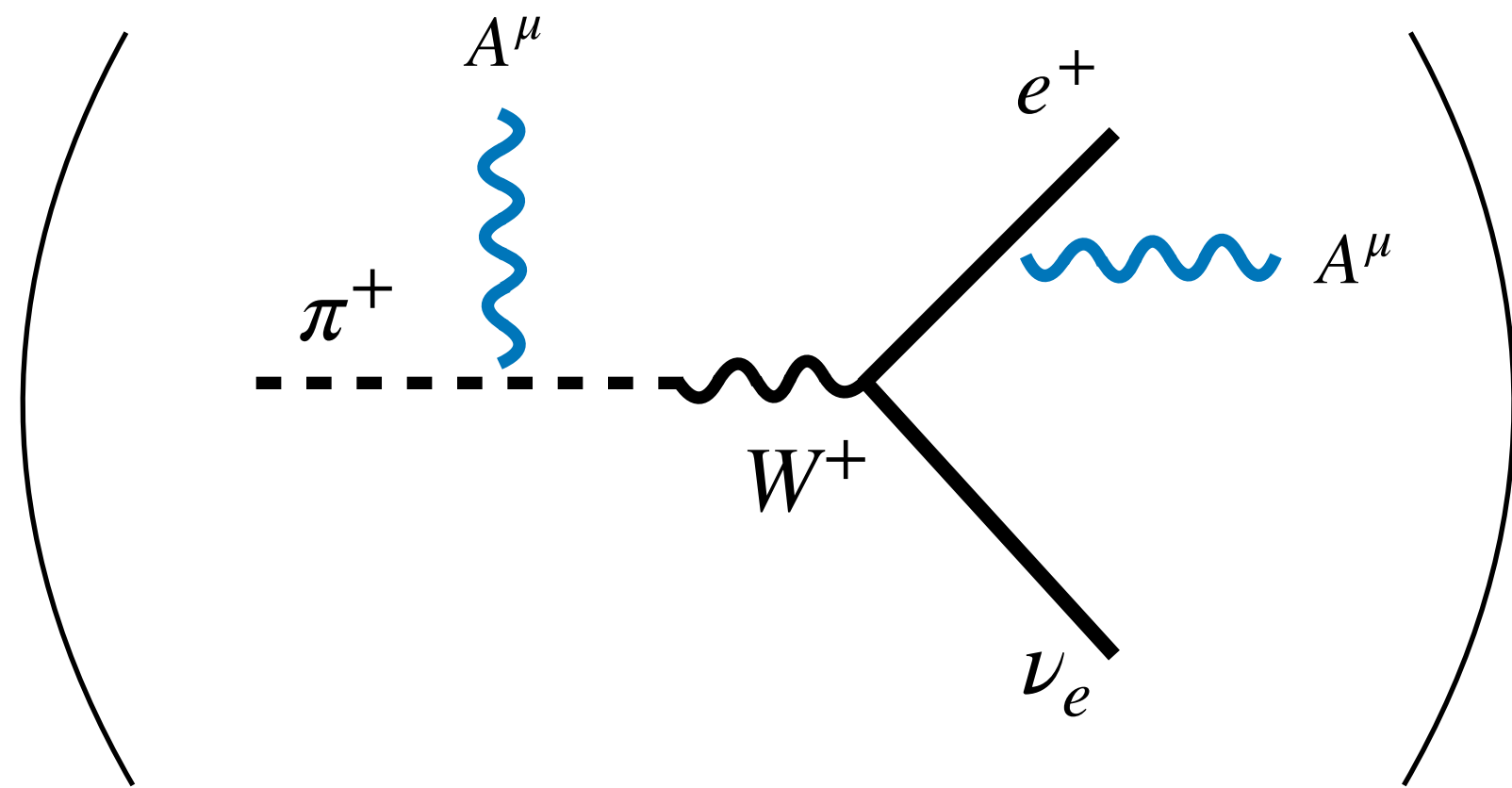
$$\frac{\Gamma_{e\gamma}}{\Gamma_e} \sim \left( \frac{m_\pi^2}{m_e^2} \right) \left( \frac{\alpha}{4\pi} \right) \sim 40$$

Can the radiative mode really be that much larger than the core leptonic process?

# Radiative pion decay

## Pion decays — helicity flip

Internal Bremsstrahlung 1 and 2



$\propto m_e$

Not helicity suppressed.

Primary vertex for pion decay:

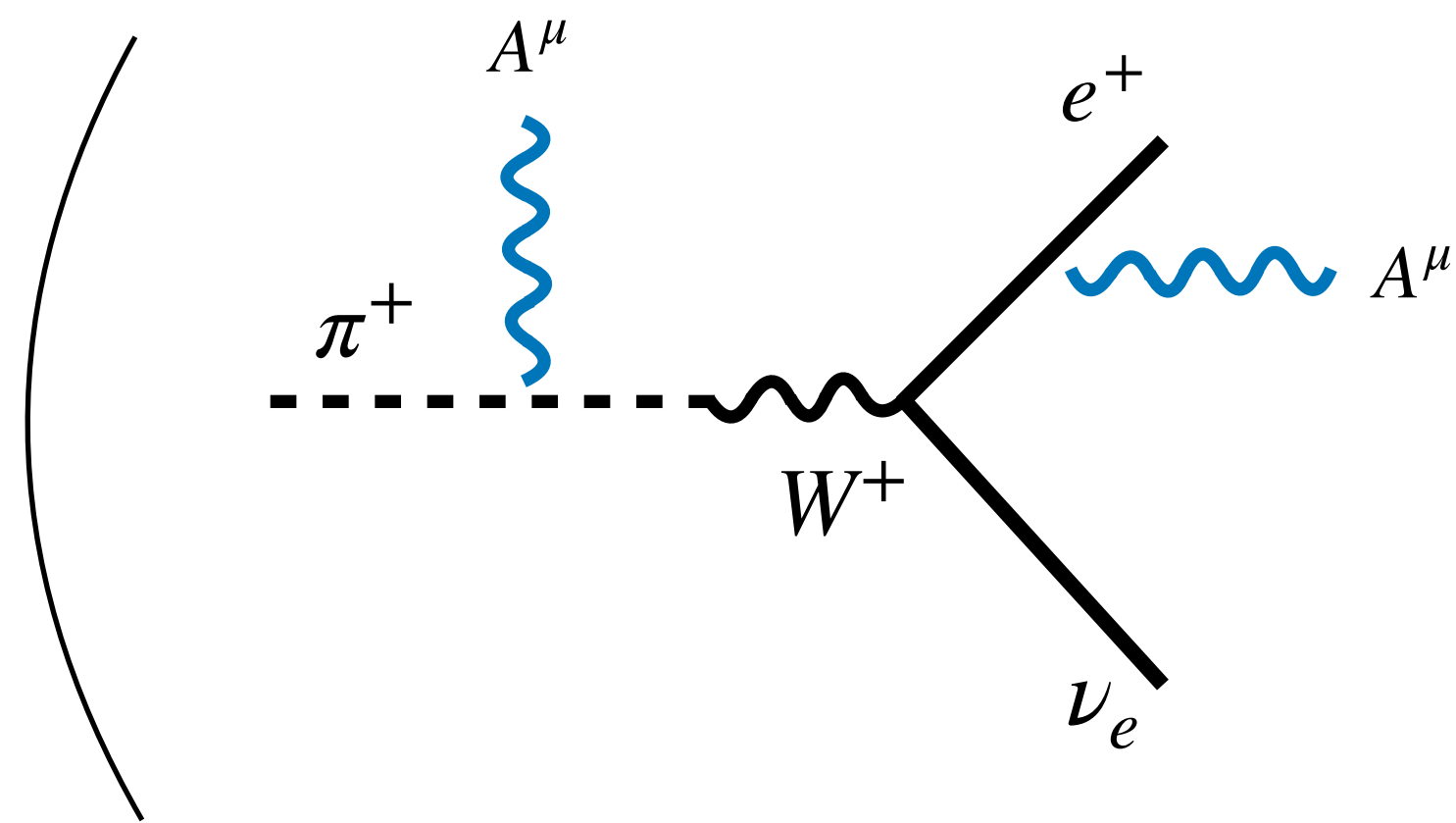
$$\mathcal{L} \supset G_F f_\pi \times \partial_\mu \pi (\bar{\ell} \gamma^\mu \nu_\ell)$$

Point-like pion & leptonic bremsstrahlung

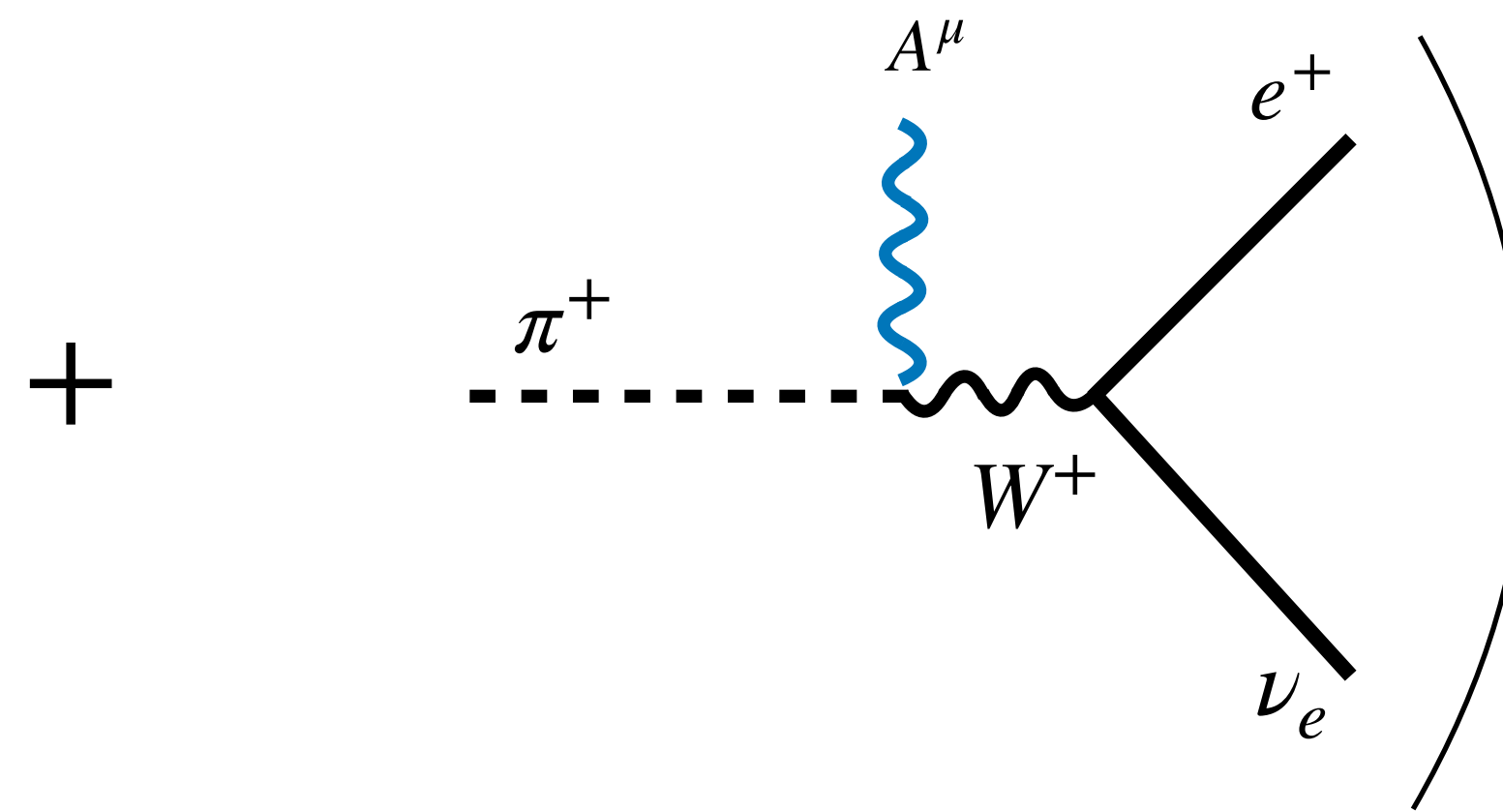
# Radiative pion decay

## Pion decays — helicity flip

Internal Bremsstrahlung 1 and 2



Internal Bremsstrahlung 3



$$\propto m_e$$

**Still helicity suppressed!**

Primary vertex for pion decay:

$$\mathcal{L} \supset G_F f_\pi \times \partial_\mu \pi (\bar{\ell} \gamma^\mu \nu_\ell)$$

Point-like pion & leptonic bremsstrahlung

**Not gauge invariant.** Need a replacement:

$$\partial_\mu \pi (\bar{\ell} \gamma^\mu \nu_\ell) \rightarrow (\partial_\mu - ieA_\mu) \pi (\bar{\ell} \gamma^\mu \nu_\ell)$$

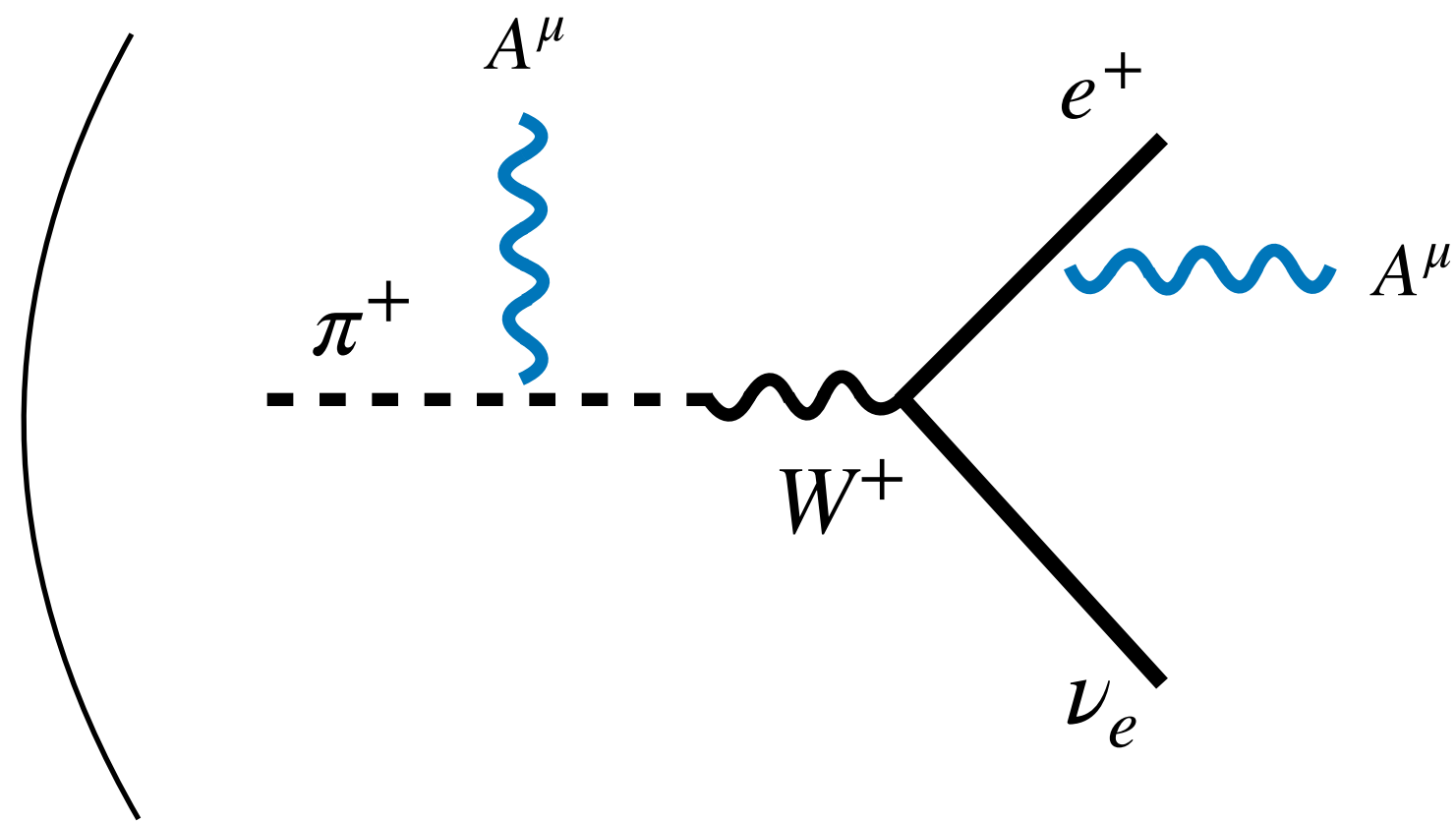
**Seagull** diagram unavoidable in gauge invariant theory.



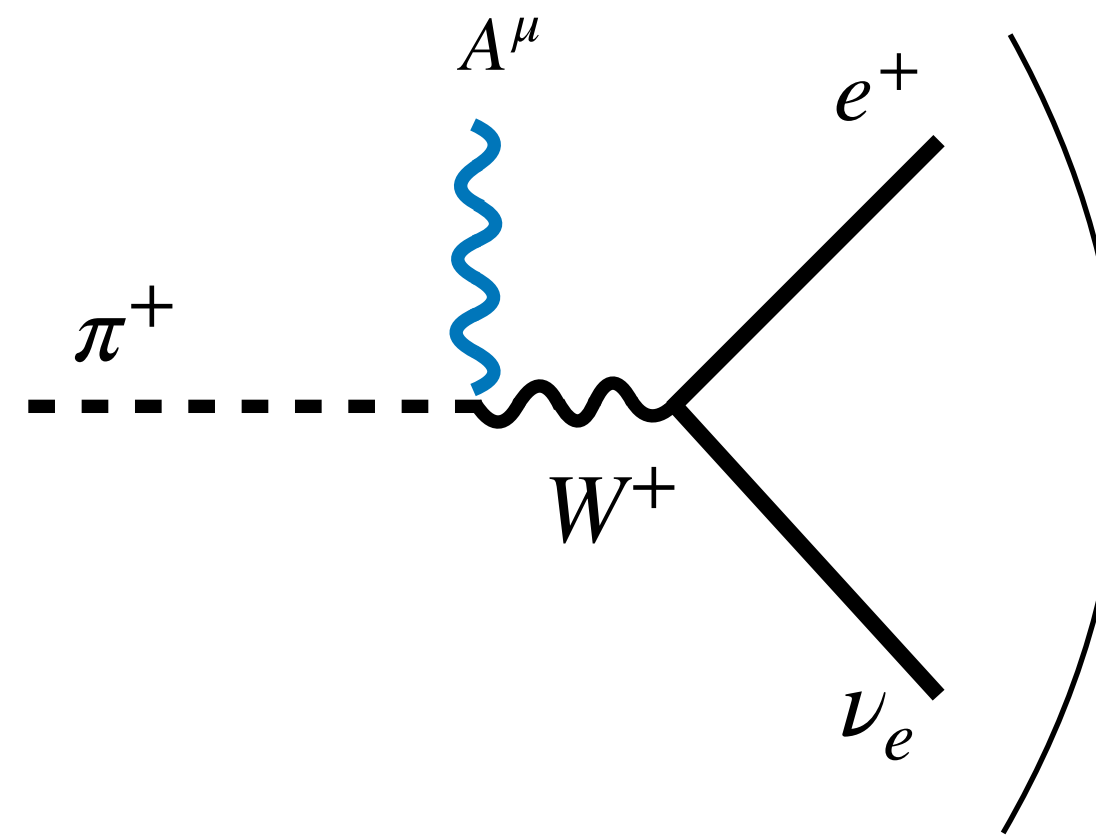
# Radiative pion decay

## Pion decays — helicity flip

Internal Bremsstrahlung 1 and 2

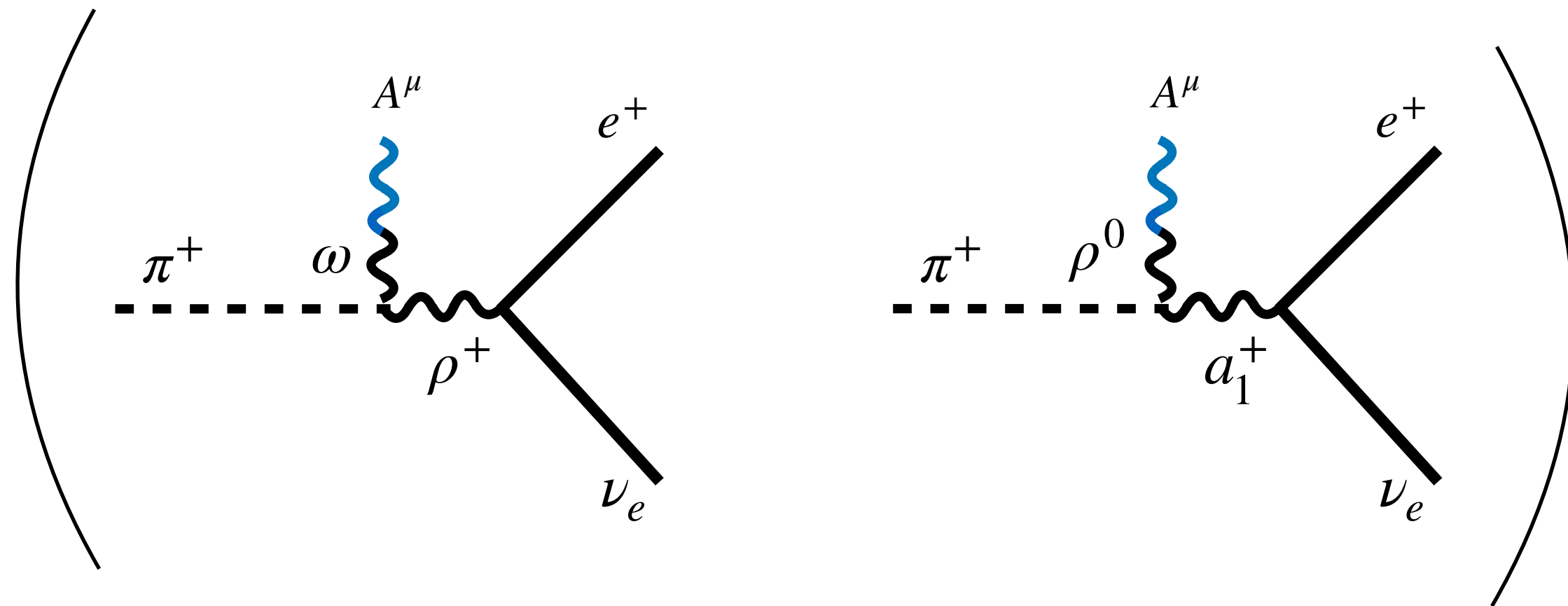


Internal Bremsstrahlung 3



$$\propto m_e$$

**Still helicity suppressed!**



$$\not\propto m_e$$

**Structure Dependent V and A**

$\pi$  no longer “point-like”

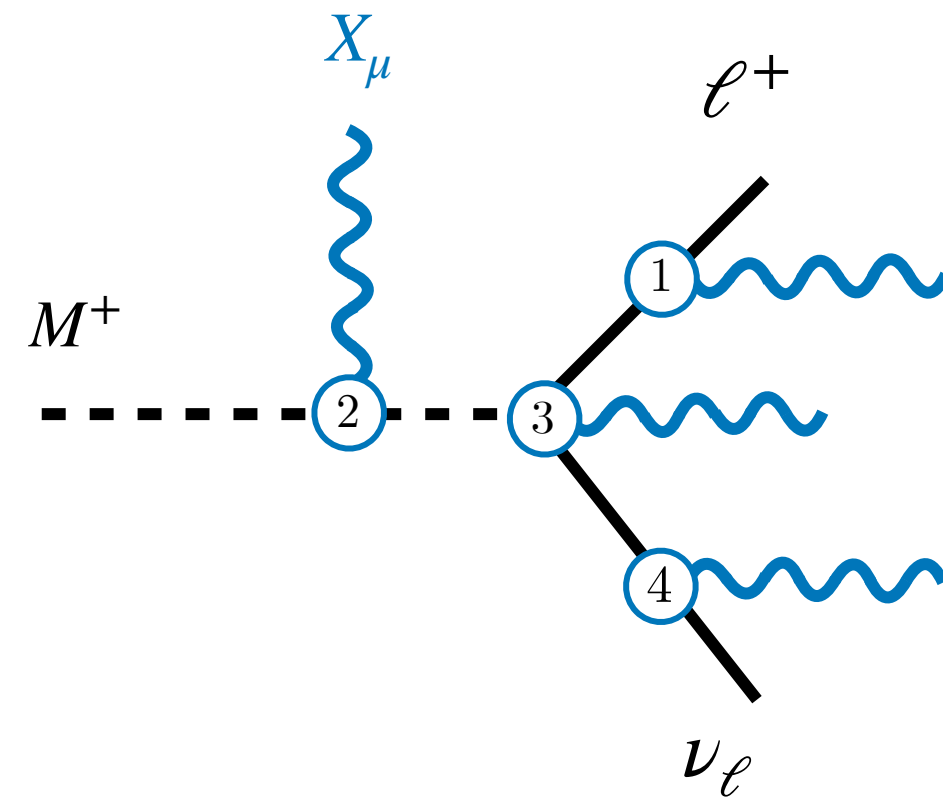
These are suppressed by  $M_\rho^{-1}$  instead so can be safely neglected!



# New vector bosons

## Radiative pion decays

MH, M. Pospelov, PRD 108, 055011

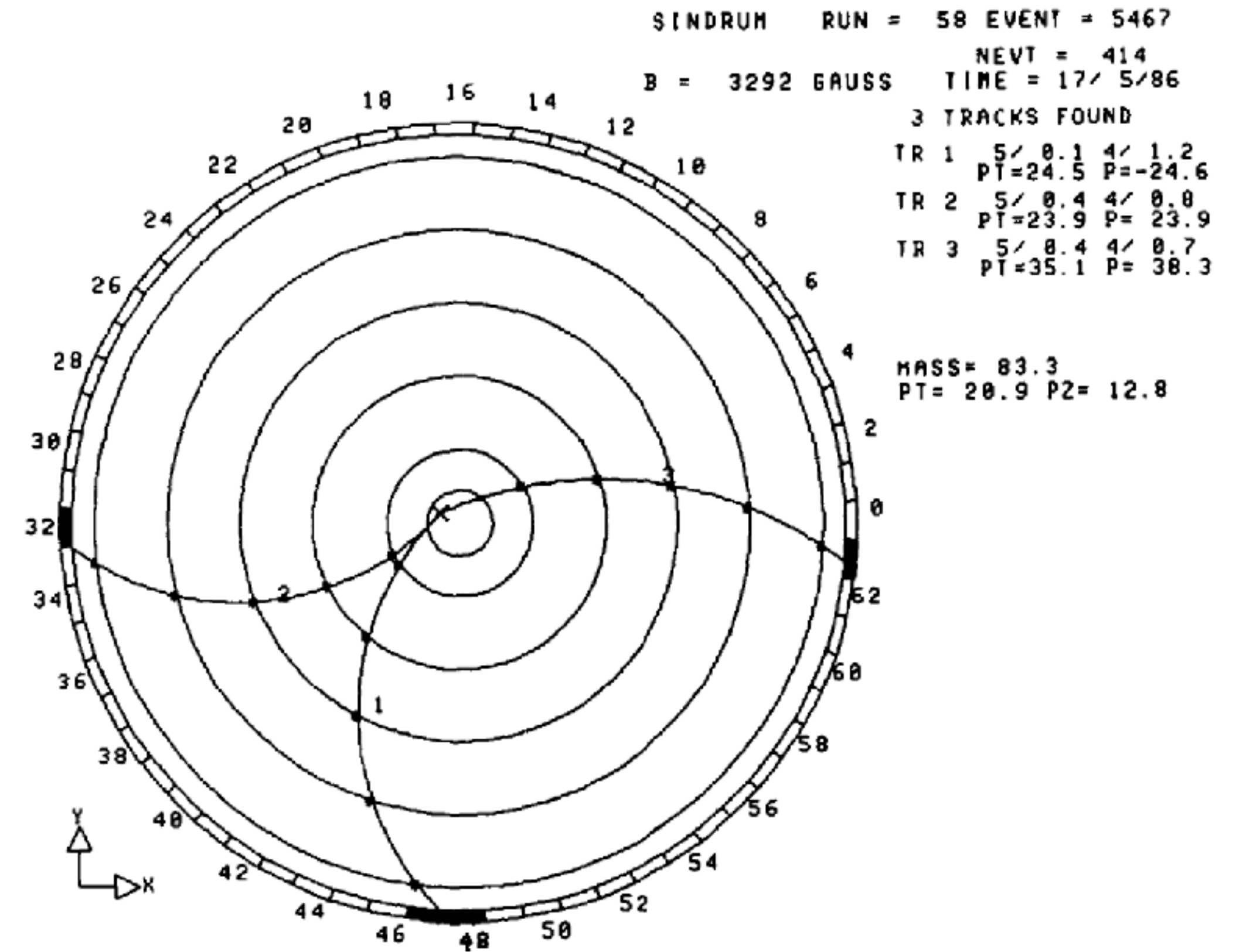


$$k^\mu \mathcal{M}_\mu \propto (Q_{\pi^+} - (Q_\nu - Q_e)) \bar{\nu}_e (k^\mu \gamma_\mu P_L - m_e P_R) e$$

Enhancement by:

- 1) helicity flip,
- 2) longitudinal mode emission.

$$\Gamma_{\text{protophobic}} \sim \Gamma_{\text{SM}} \times \left( \frac{m_\pi^4}{m_e^2 m_X^2} \right)$$

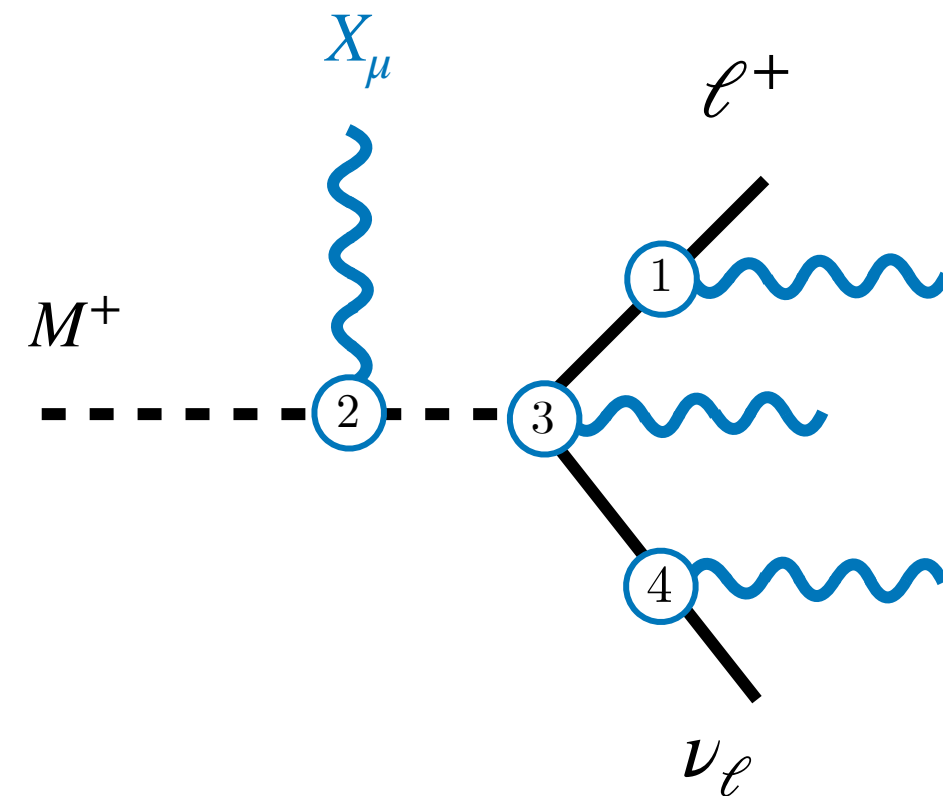


SINDRUM-I search for bumps in  $\pi^+ \rightarrow e^+ \nu_e e^+ e^-$  decays.

# New vector bosons

## Radiative pion decays

MH, M. Pospelov, PRD 108, 055011

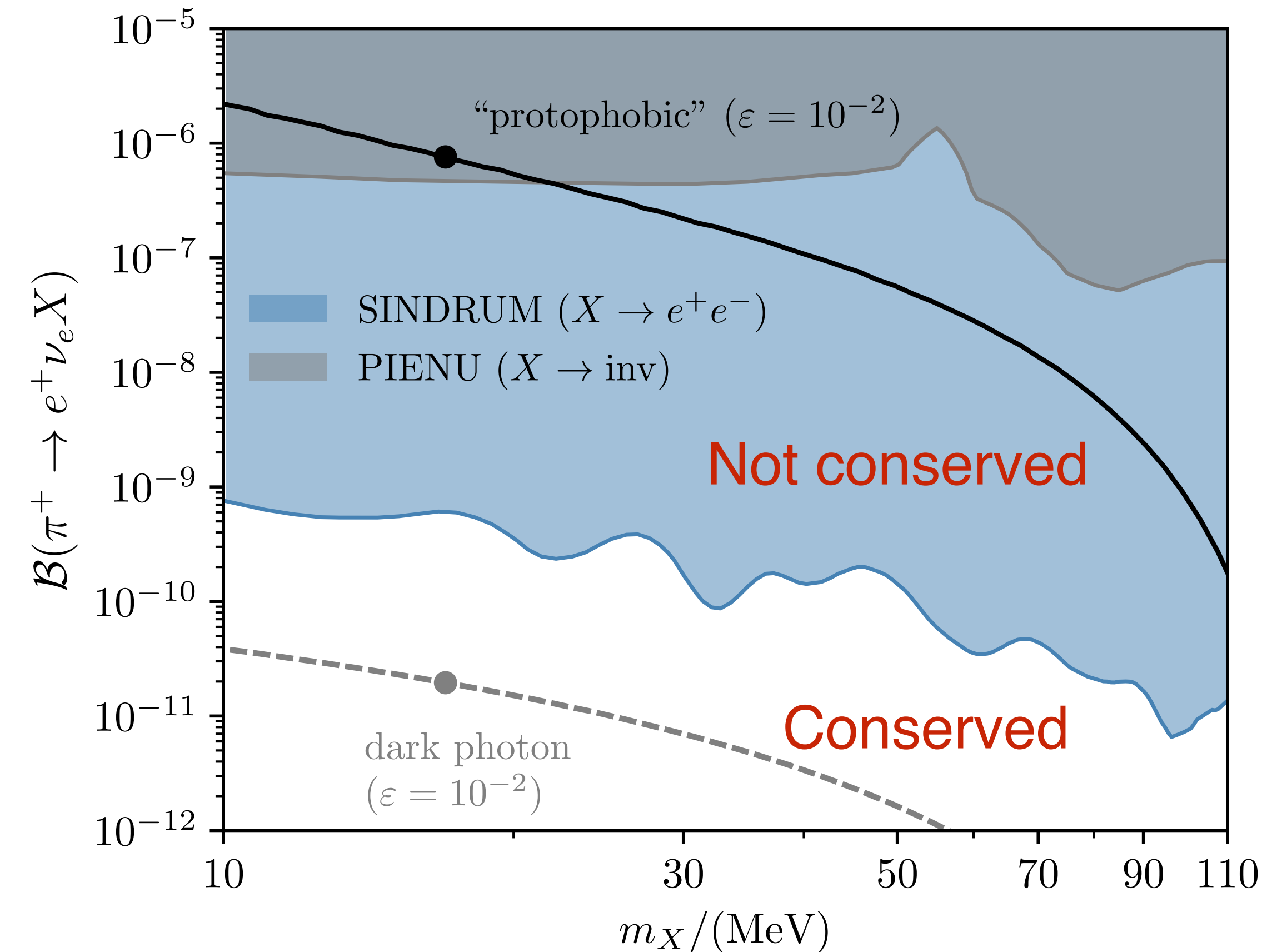


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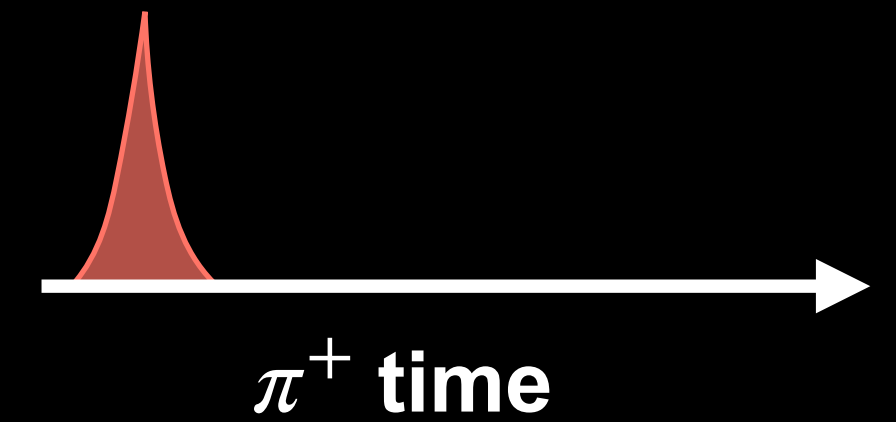


BTW: This basically excludes all exotic 17 MeV vector Boson explanations of ATOMKI anomaly



# Long-lived particles at spallation sources

## “Weak-violating” axion-like-particle (WV ALP)



Light goldstone boson that probes exotic electron couplings

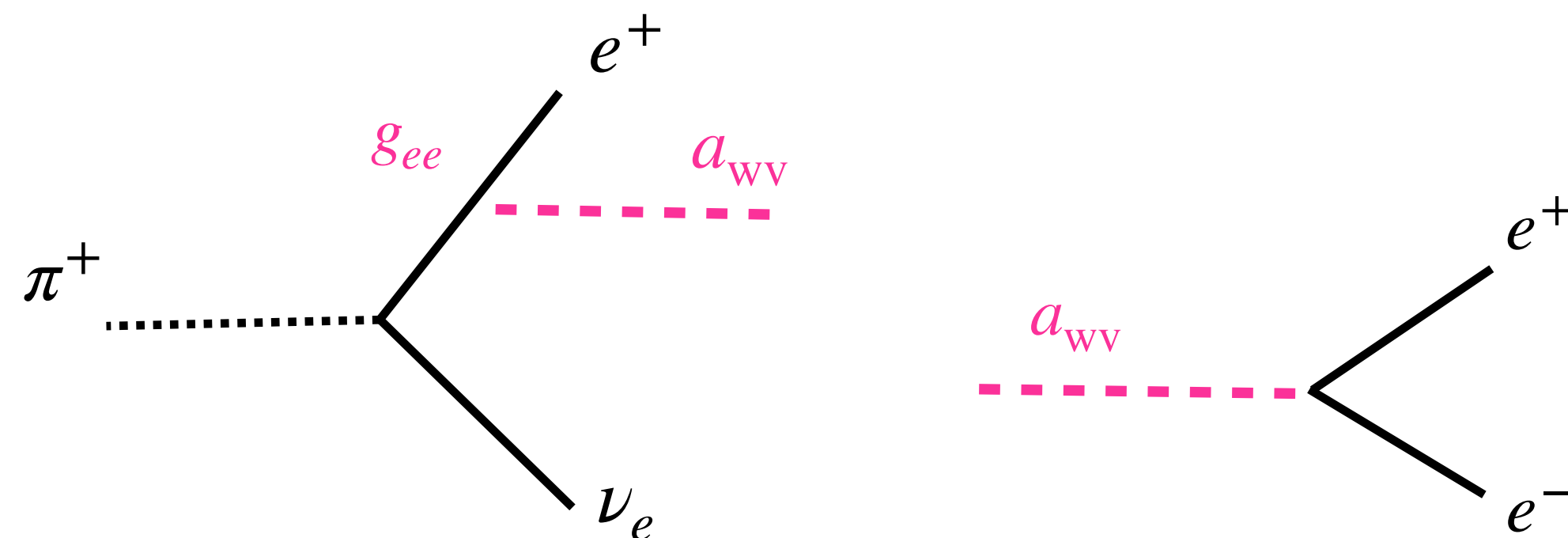
$$j_{PQ}^\mu = \frac{\bar{g}_{\ell\ell}}{2m_\ell} \bar{\ell} \gamma^\mu \ell + \frac{g_{\ell\ell}}{2m_\ell} \bar{\ell} \gamma^\mu \gamma_5 \ell + \frac{g_{\nu\ell}}{2m_\ell} \bar{\nu}_\ell \gamma^\mu P_L \nu_\ell.$$

Lifting helicity suppression in 3-body  $\pi^+$  decay is not easy, but can be done in a class of “weak-violating” ALP models.

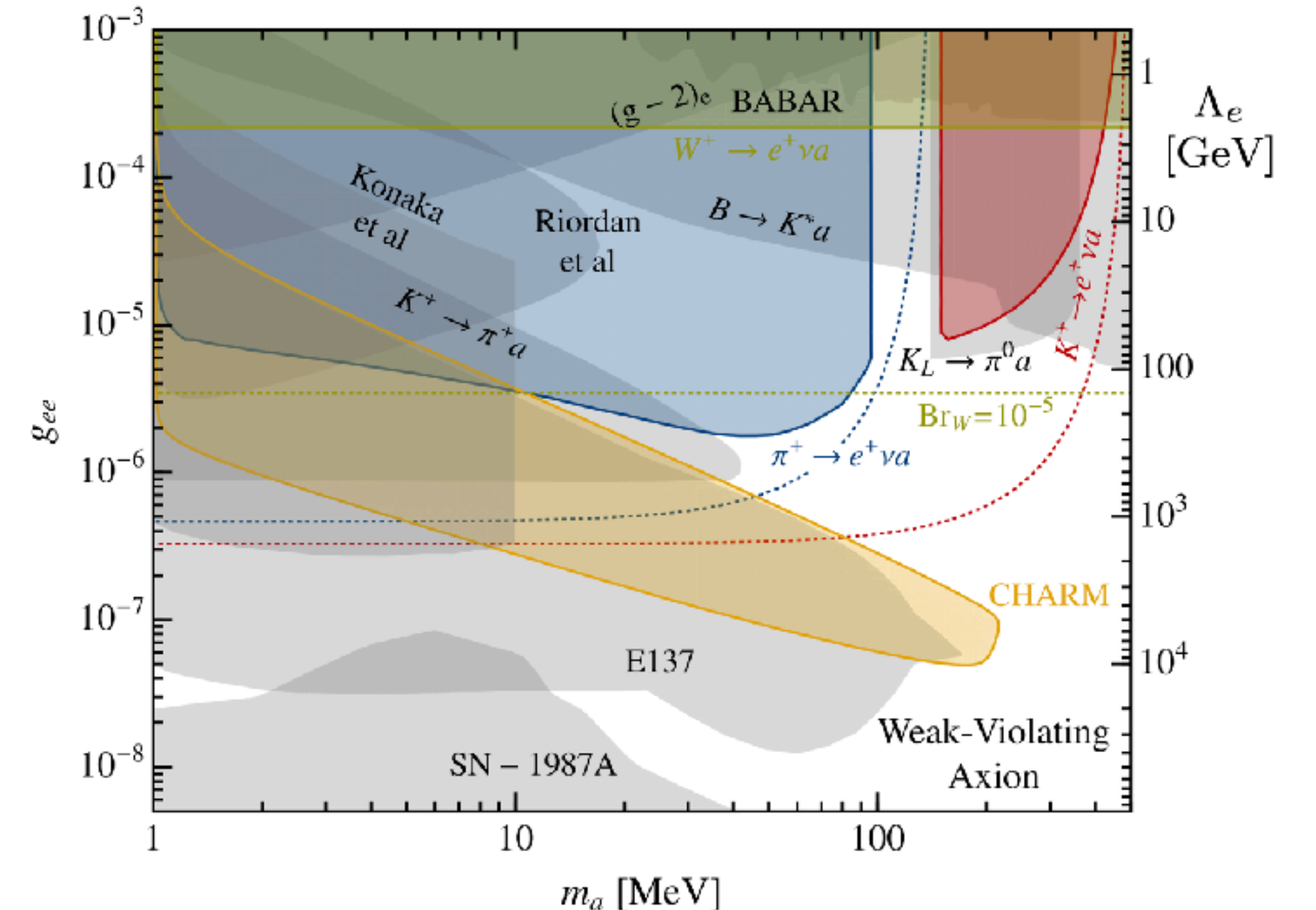
**This is an exception: underlying current is not gauged, so it is ok!**

[W. Altmannshoffer et al, arXiv:2209.00665](https://arxiv.org/abs/2209.00665)

In this case, three-body decays of the pion are the dominant source of these ALPs at accelerators.

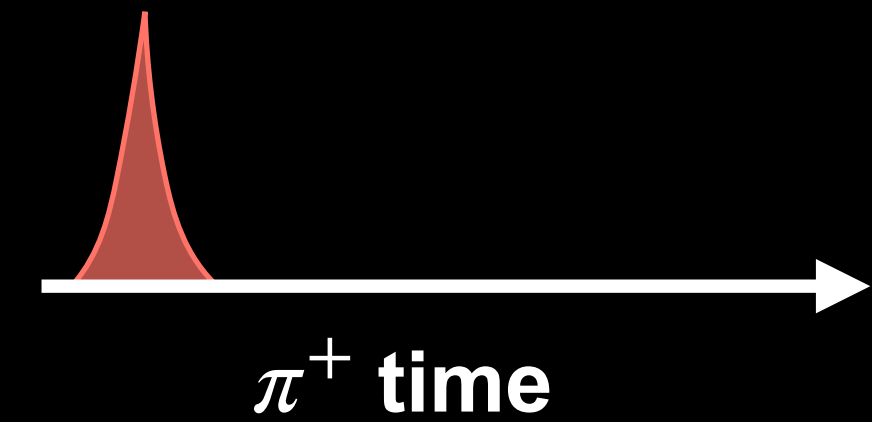


$$\frac{\mathcal{B}(P^+ \rightarrow \ell^+ \nu_\ell a)}{\mathcal{B}(P^+ \rightarrow \ell'^+ \nu_{\ell'})} \simeq \frac{1}{1536\pi^2} \frac{m_P^4}{m_\ell^2 m_{\ell'}^2} \left(1 - \frac{m_{\ell'}^2}{m_P^2}\right)^{-2} \times \left[ (g_{\ell\ell} - \bar{g}_{\ell\ell} + g_{\nu\ell})^2 f_0(x_P) + \frac{16m_\ell^2}{m_P^2} g_{\ell\ell}^2 f_1(x_P) \right].$$



# Long-lived particles at spallation sources

## “Weak-violating” axion-like-particle (WV ALP)

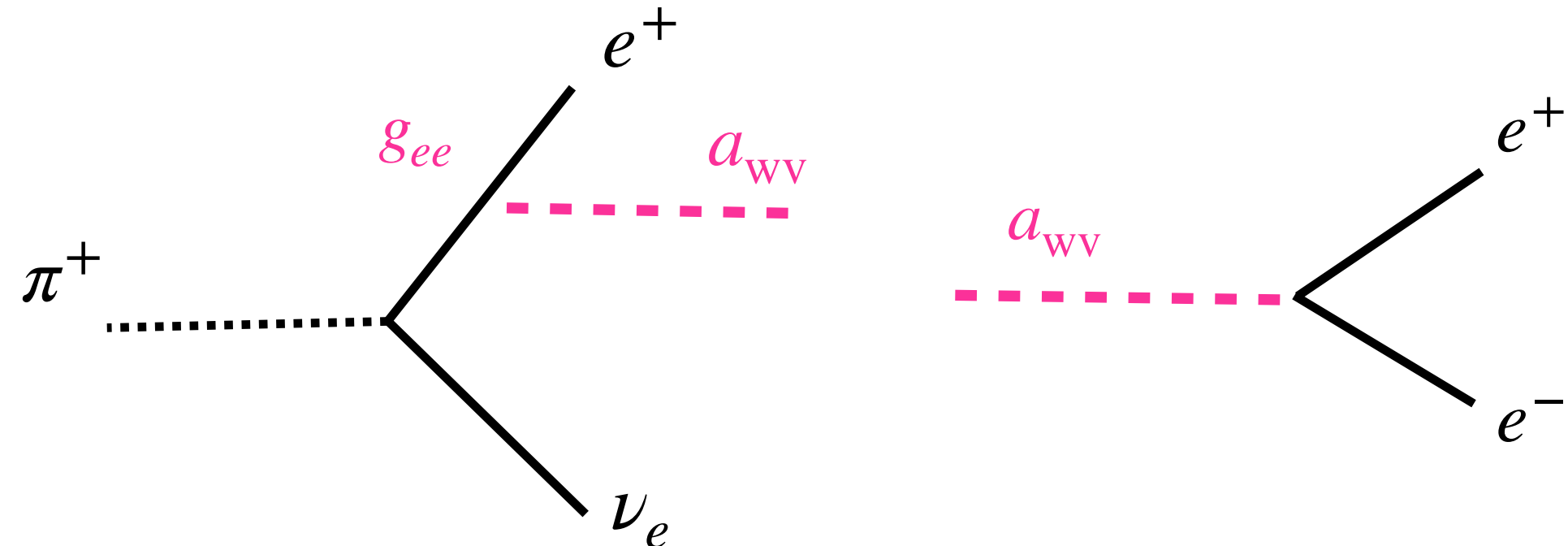


### Light goldstone boson that probes exotic electron couplings

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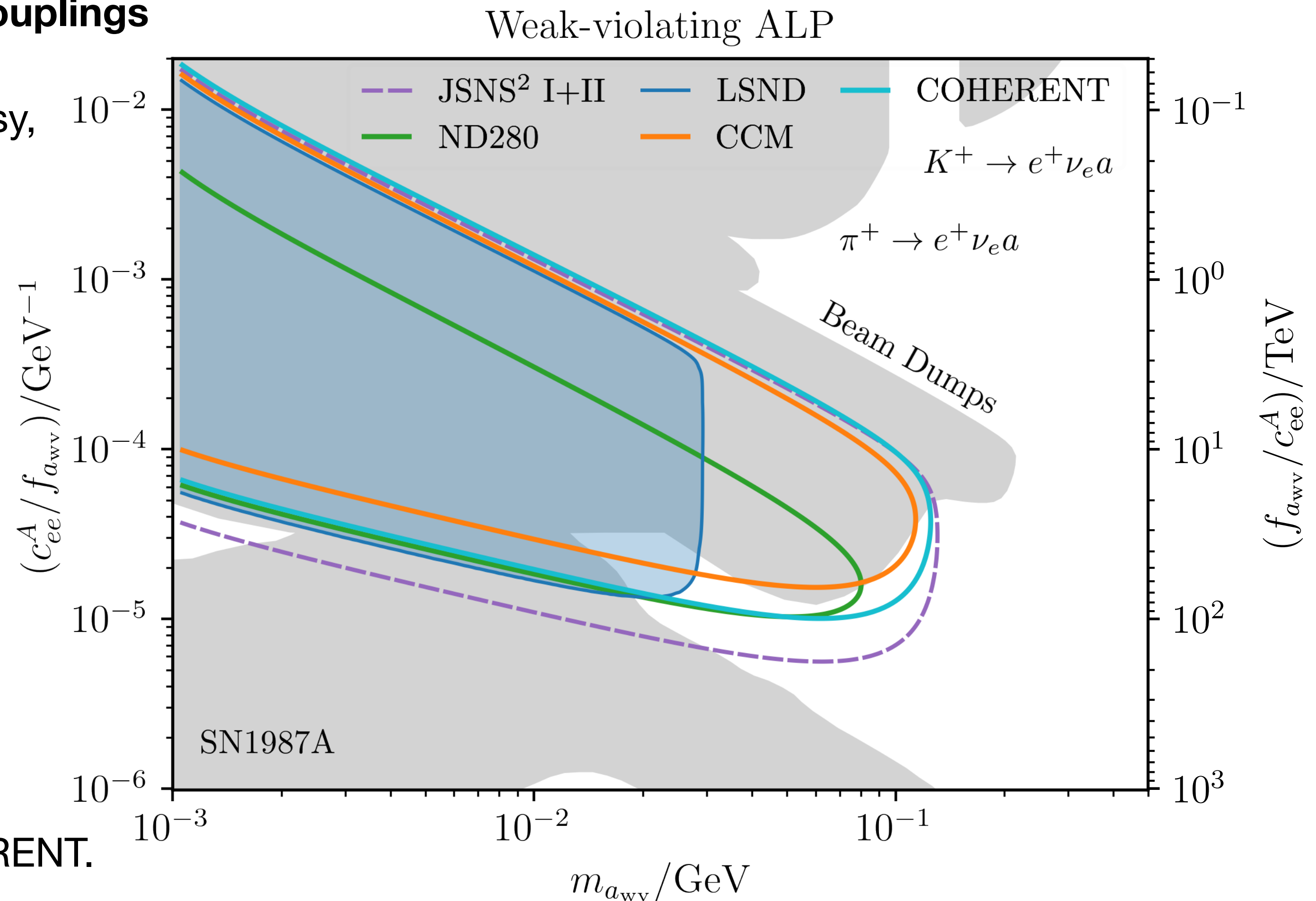
[W. Altmannshoffer et al, arXiv:2209.00665](#)

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Complementary coverage by ND280, CCM, and COHERENT.

LSND more on par since  $a_{wv}$  is more energetic.



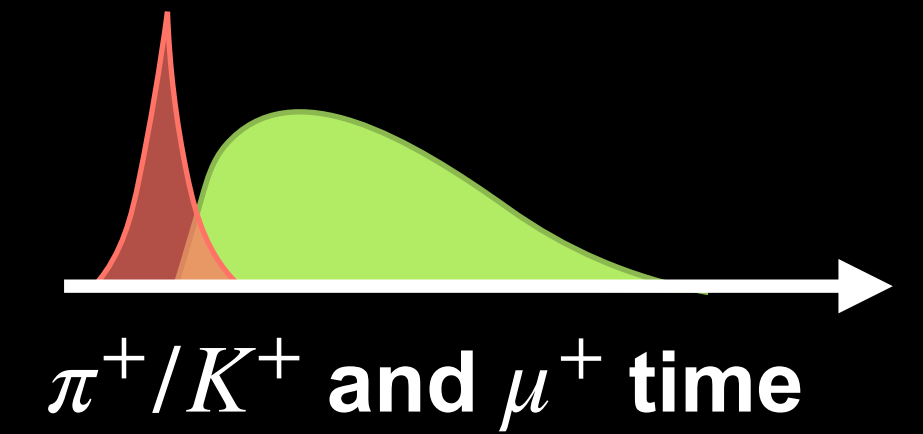
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- 1) Spallation (neutron) sources: what and why?
- 2) Opportunities with neutrino detectors
- 3) Long-lived particles below  $K$ ,  $\pi$ , and  $\mu$  masses

Heavy neutral leptons (low-scale seesaw)

# Long-lived particles at spallation sources

## Heavy neutral leptons

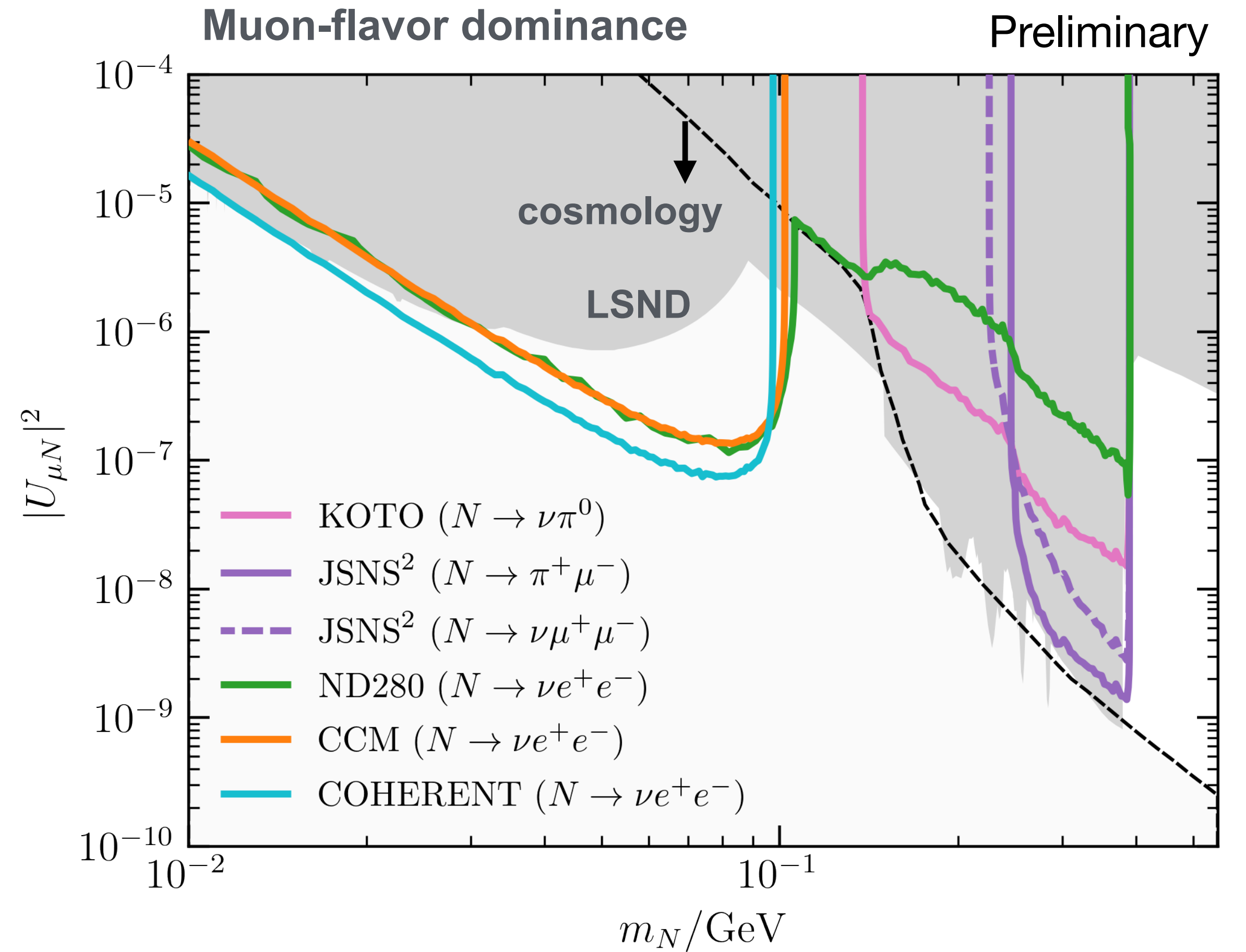
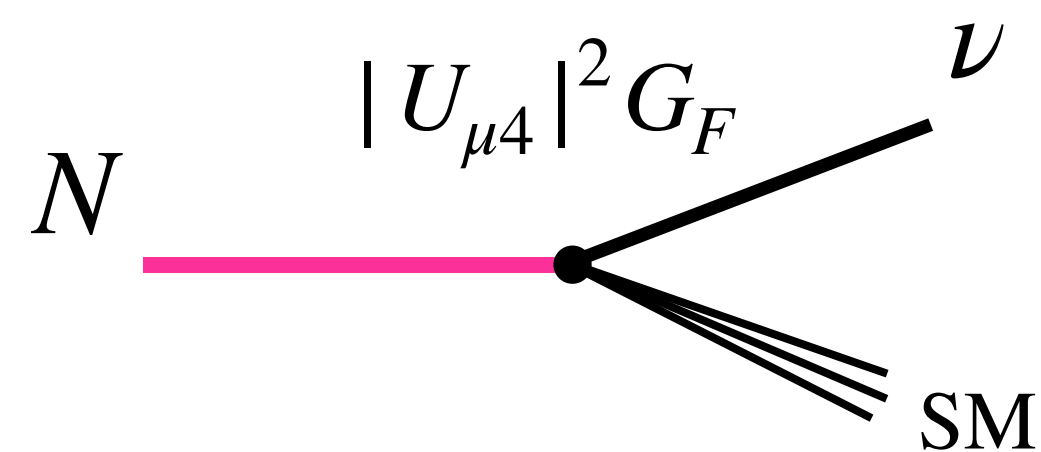


### Low-scale neutrino mass model

Improvement over LSND because of the stringent signal selection criterion to fake  $\nu - e$  scattering.

Most final states are relevant for  $K^+$  parentage so they have limited sensitivity.

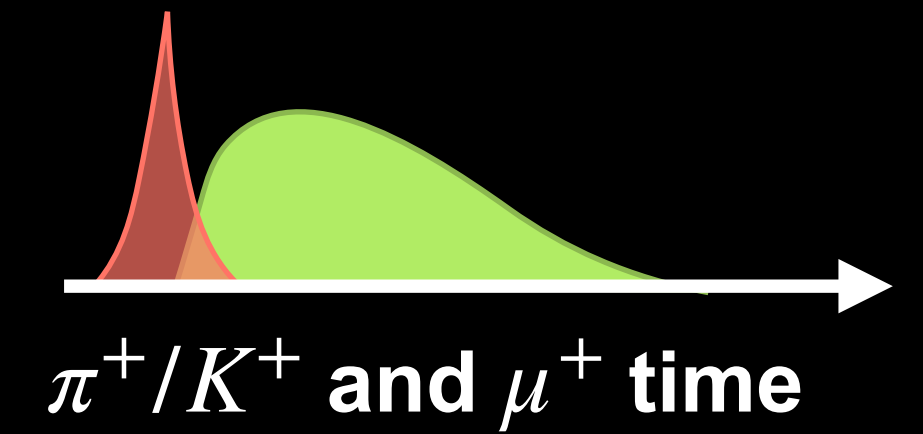
Competition with LSND and meson peak searches.



LSND limit derived in Y. Ema, Z. Liu, K. Lyu, M. Pospelov, [arXiv:2306.07315](https://arxiv.org/abs/2306.07315)

# Long-lived particles at spallation sources

## Heavy neutral leptons

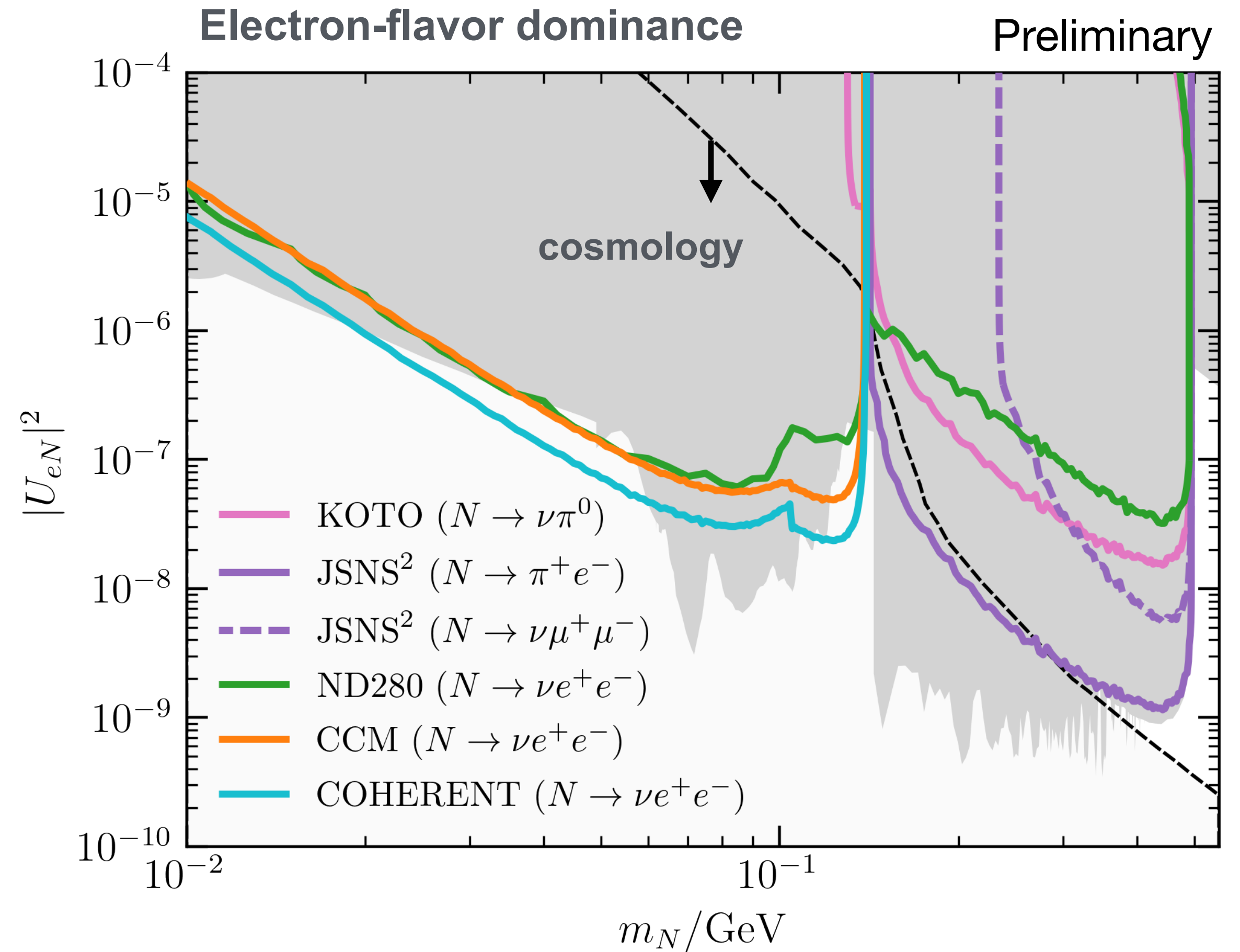
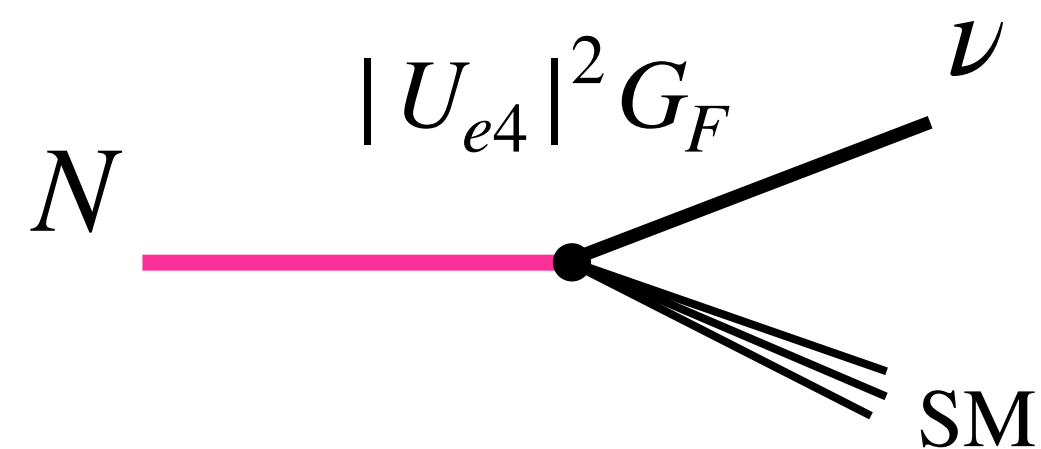


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# Long-lived particles at spallation sources

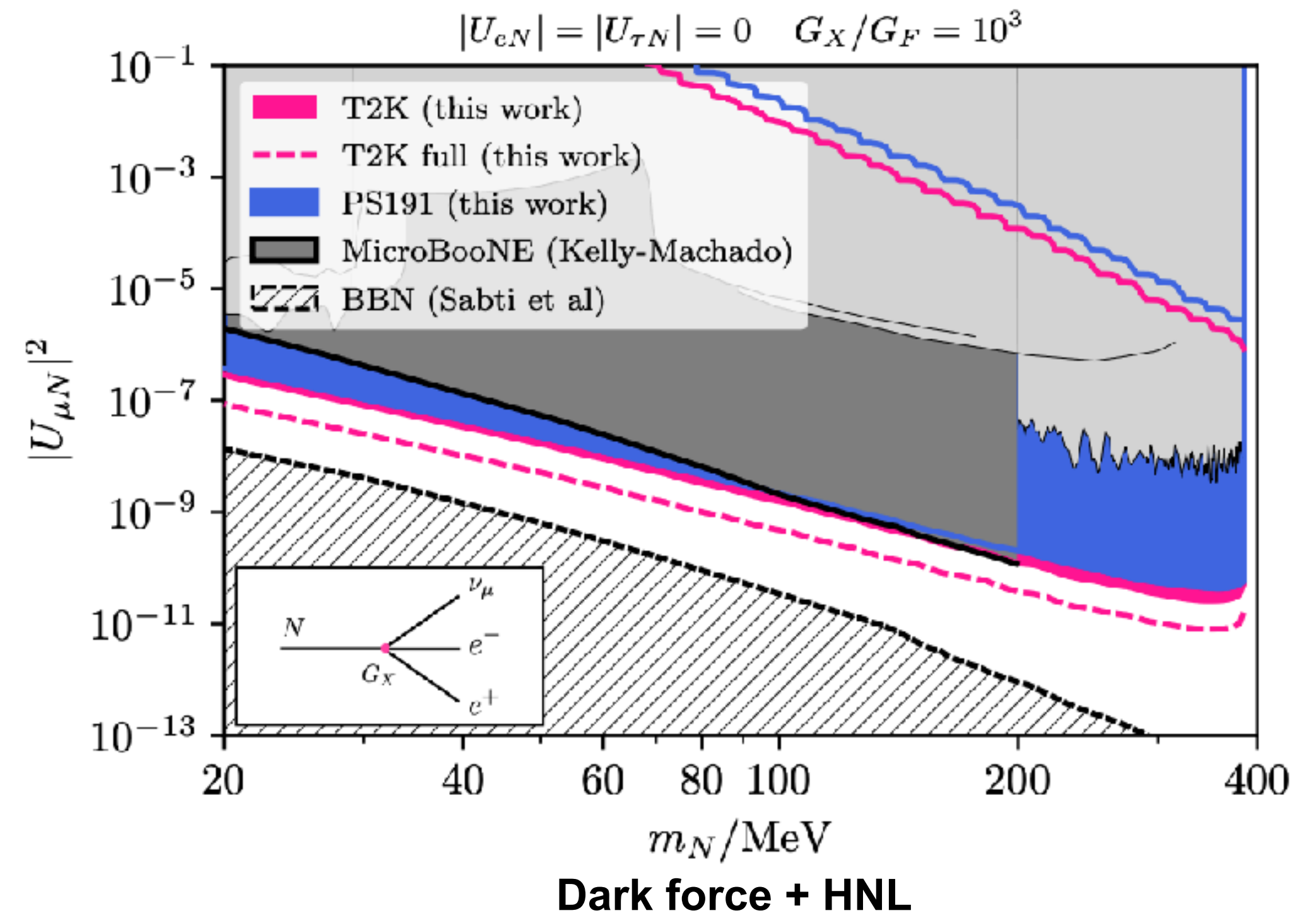
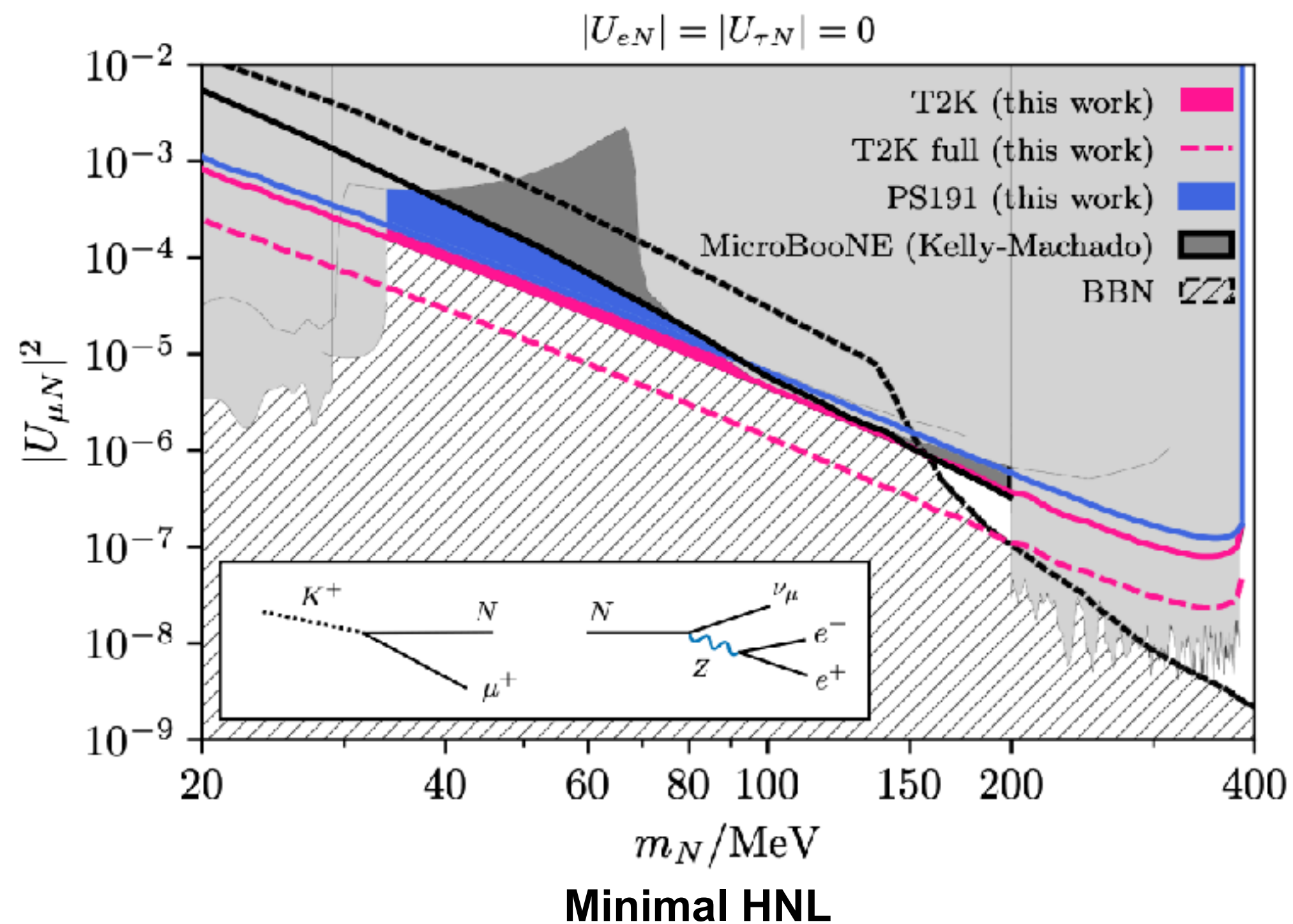
## Heavy neutral leptons — avoiding cosmological limits?



Cosmological limits typically make the sub-100 MeV region less interesting in minimal HNL models.

If new forces exist (e.g., magnetic moments or dark photons), decay-in-flight limits on LLPs quickly become the most important.

**Spallation sources are useful for “non-minimal” HNL models. For example:**





# Outline:

**1) Spallation (neutron) sources: what and why?**

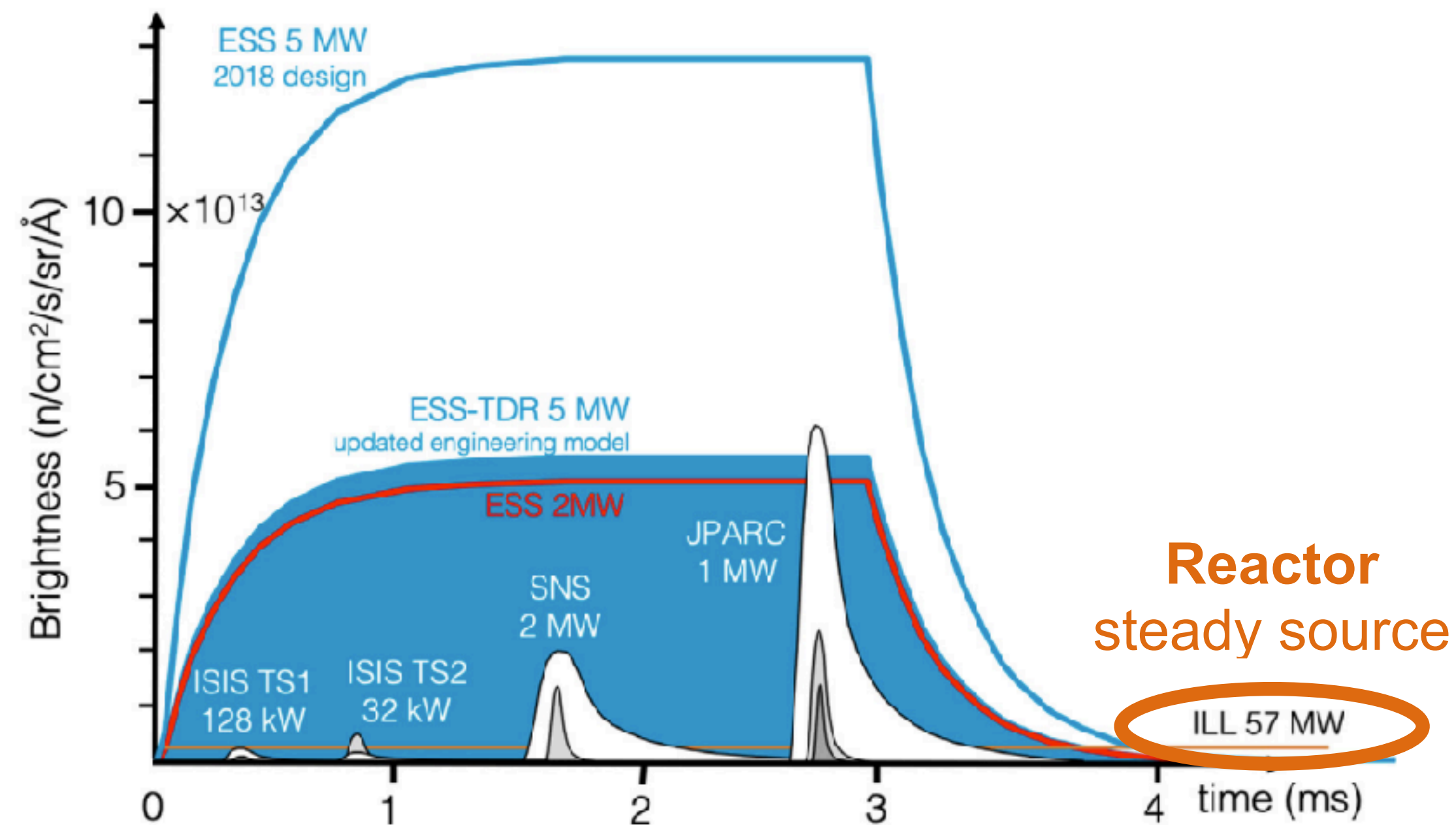
**2) Opportunities with neutrino detectors**

**3) Long-lived particles below  $K$ ,  $\pi$ , and  $\mu$  masses**

*Bonus) Thoughts on future and next-generation facilities*

# Future facilities

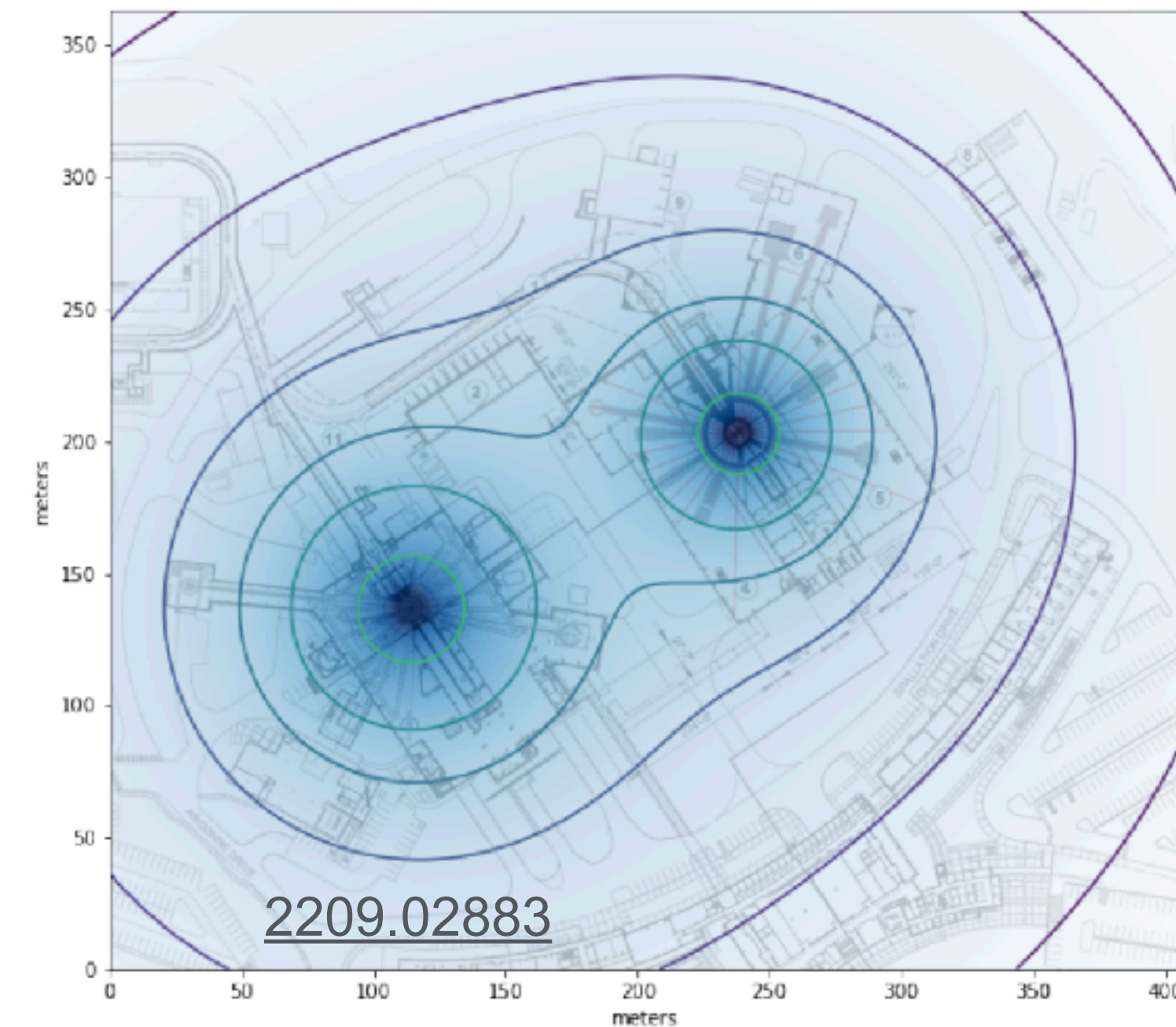
## European Spallation Source a “long” jump in intensity (rotating target)



<https://doi.org/10.1016/j.nima.2020.163402>

5 MW (1-2 GeV  $p^+$ ) — huge intensity, but worse background rejection.

## Second Target Station at SNS More neutrinos (rotating target)



First target would receive 2.0 MW at 45 Hz  
 Second target would receive 0.7 MW at 15 Hz  
**Neutrino rate is actually the same**

# Summary

Thank you for listening!

Matheus Hostert ([mhostert@g.harvard.edu](mailto:mhostert@g.harvard.edu))

- 1) Spallation targets are a very messy environment... but move a bit further out and build sufficiently large-volume detectors, and *extremely rare processes* from LLPs could appear.
- 2) Shown a non-exhaustive list of long-lived particle (LLP) models that can be constrained with existing spallation sources and detectors. Usually less minimal to survive other limits below  $\pi$  and  $\mu$  masses.
- 3) A clear application for a well-shielded, low-density, large-volume, and fast detector close to the source.

Magnificent CEvNS 2024

Not all about POTs and volume: background rejection, timing, and people-power.  
Lots of stones are left unturned.

**Build bigger and away from the neutrino alley? Lower density CCM?**

**The future is bright and I look forward to the new searches!**



# Back-up slides

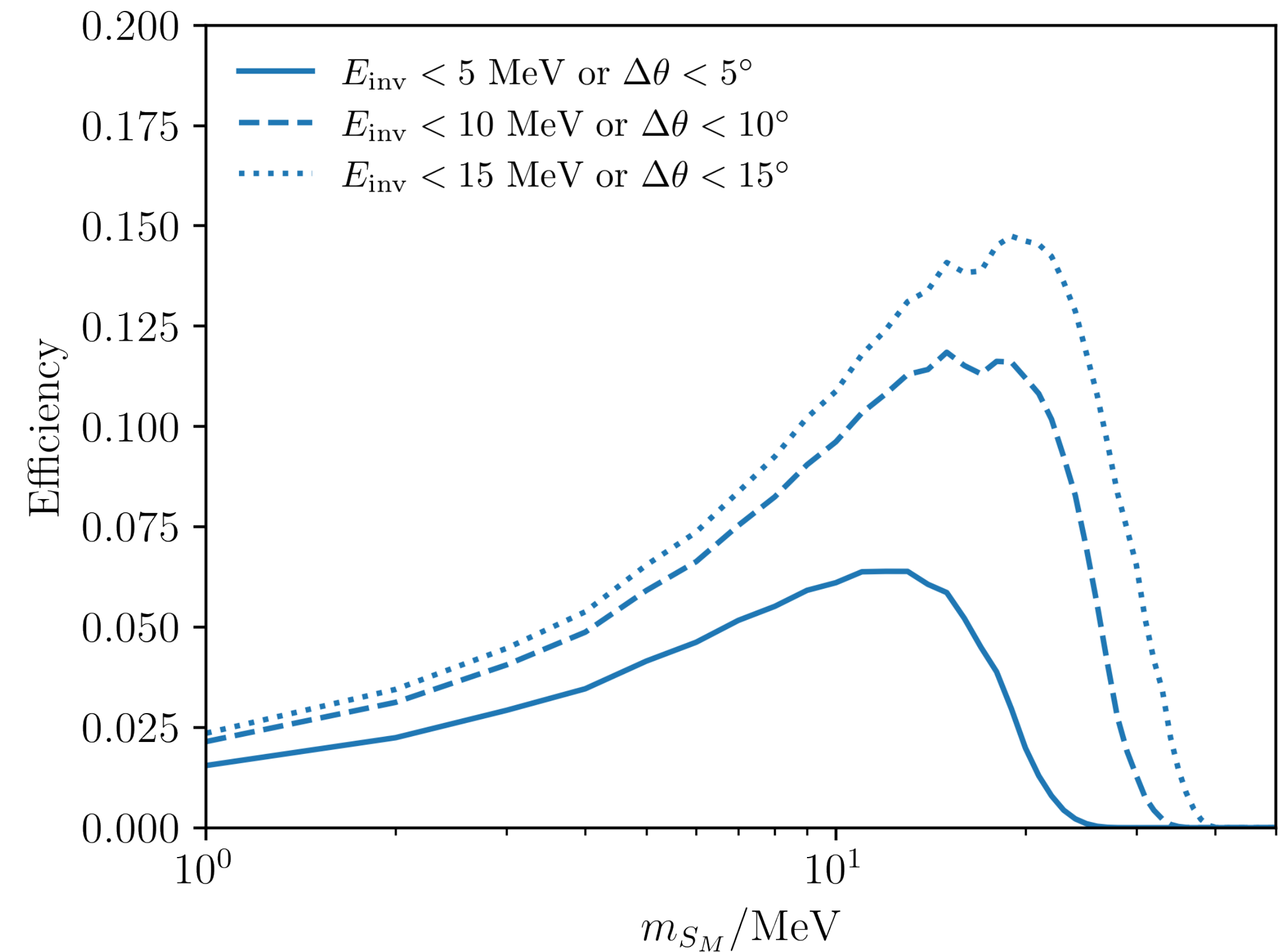
# CEvNS — Coherent Elastic Neutrino Nucleus Scattering

Three examples for requiring  $\gamma\gamma \leftrightarrow e^-$  mis-identification:

- 1)  $E_{e_{inv}} < 5 \text{ MeV}$  or  $\theta_{ee} < 5^\circ$
- 2)  $E_{e_{inv}} < 10 \text{ MeV}$  or  $\theta_{ee} < 10^\circ$
- 3)  $E_{e_{inv}} < 15 \text{ MeV}$  or  $\theta_{ee} < 15^\circ$  (strongest limit)

All events must also satisfy signal selection criterion:

$$18 \text{ MeV} < E_{\text{vis}} < 50 \text{ MeV} \text{ and } \cos \theta_{\text{vis}} > 0.9$$



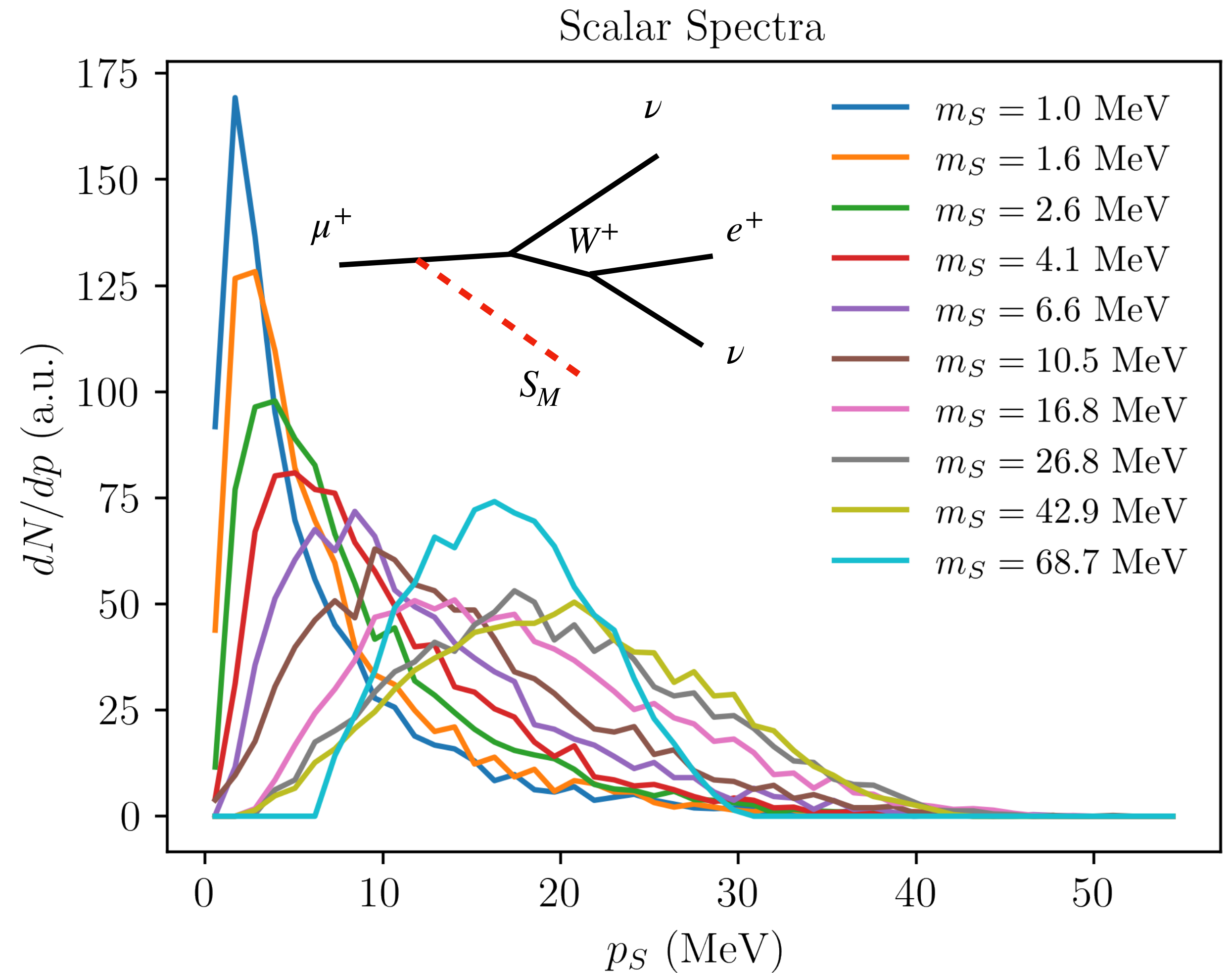
# CEvNS — Coherent Elastic Neutrino Nucleus Scattering

Three examples for requiring  $\gamma\gamma \leftrightarrow e^-$  mis-identification:

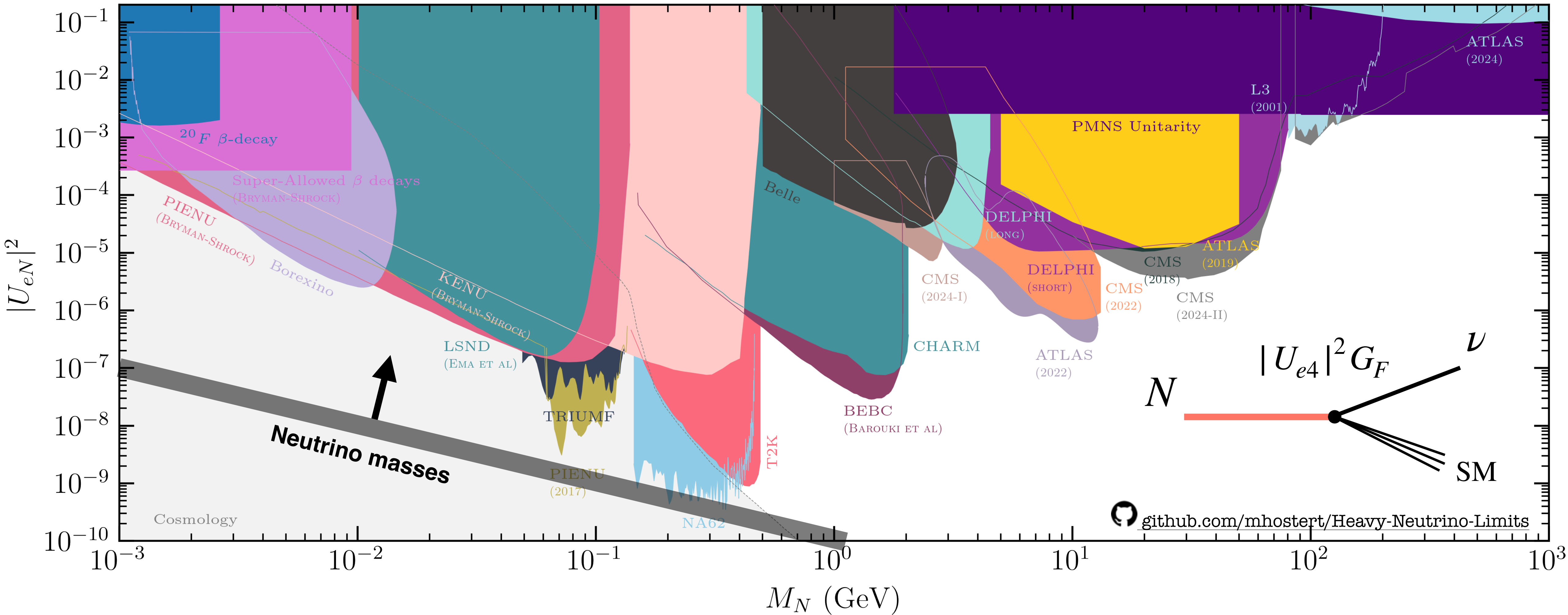
- 1)  $E_{e_{inv}} < 5 \text{ MeV}$  or  $\theta_{ee} < 5^\circ$
- 2)  $E_{e_{inv}} < 10 \text{ MeV}$  or  $\theta_{ee} < 10^\circ$
- 3)  $E_{e_{inv}} < 15 \text{ MeV}$  or  $\theta_{ee} < 15^\circ$  (strongest limit)

All events must also satisfy signal selection criterion:

$$18 \text{ MeV} < E_{\text{vis}} < 50 \text{ MeV} \text{ and } \cos \theta_{\text{vis}} > 0.9$$



# Heavy Neutral Leptons



# Axion-like particles with kaon decay at rest

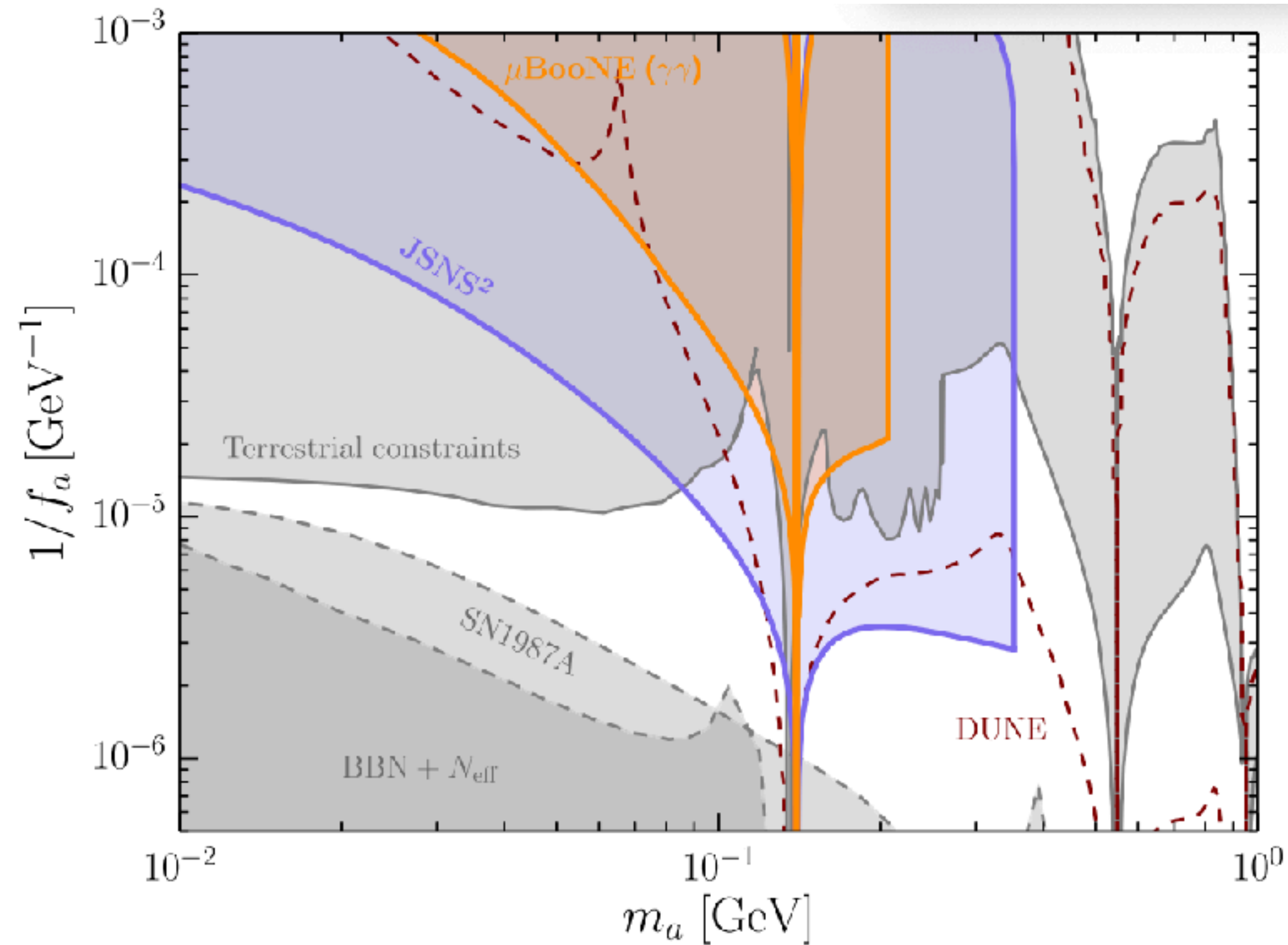


FIG. 2. (*Codominance*) The same figure as Fig. 1 but with  $c_{WW} = c_{BB} = c_{GG}$ . The sensitivity for  $m_a \ll m_\pi$  is worse than the gluon dominance case since  $|c_{\gamma\gamma}^{\text{eff}}| \ll 1$  for this specific choice of the parameters.

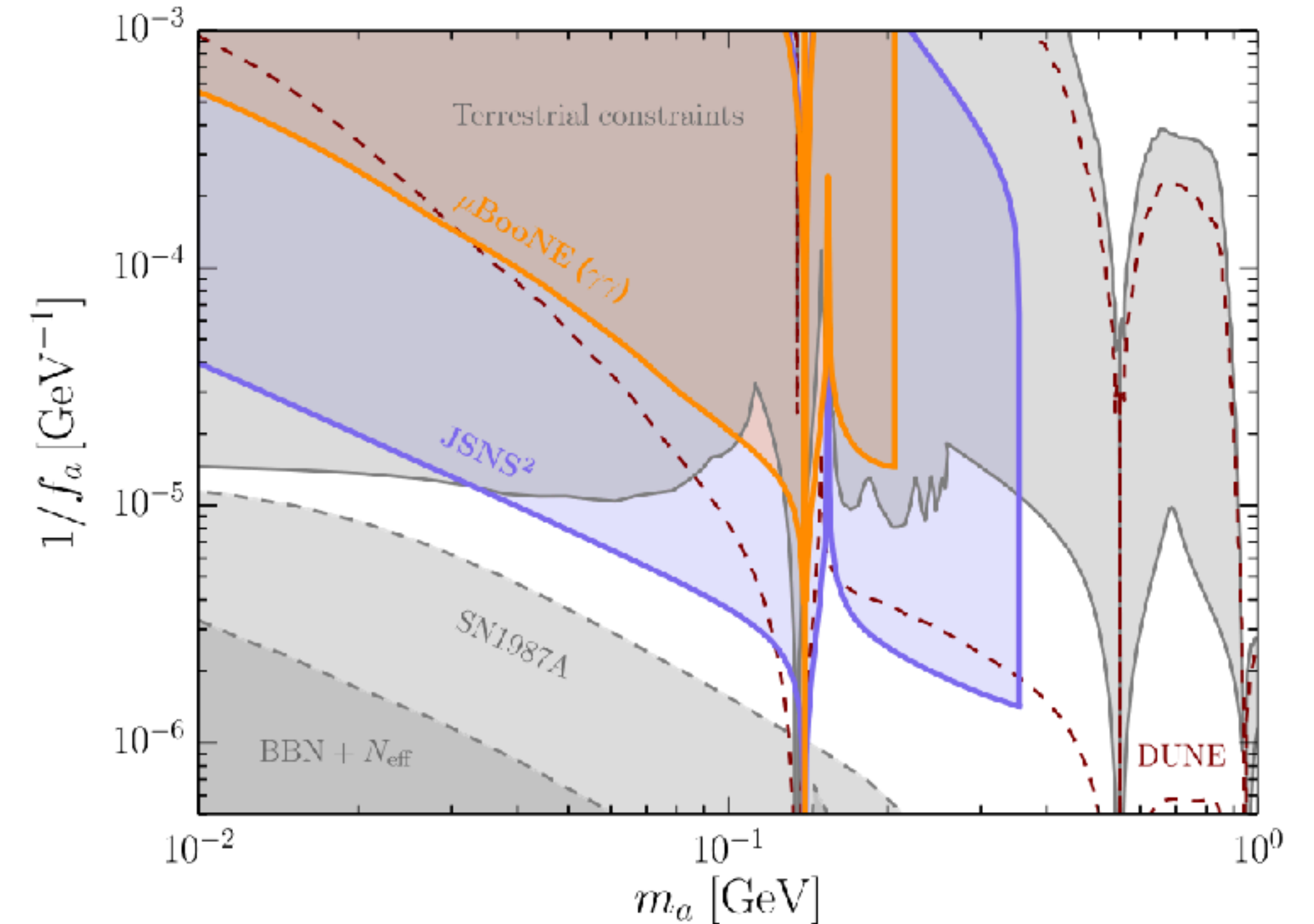
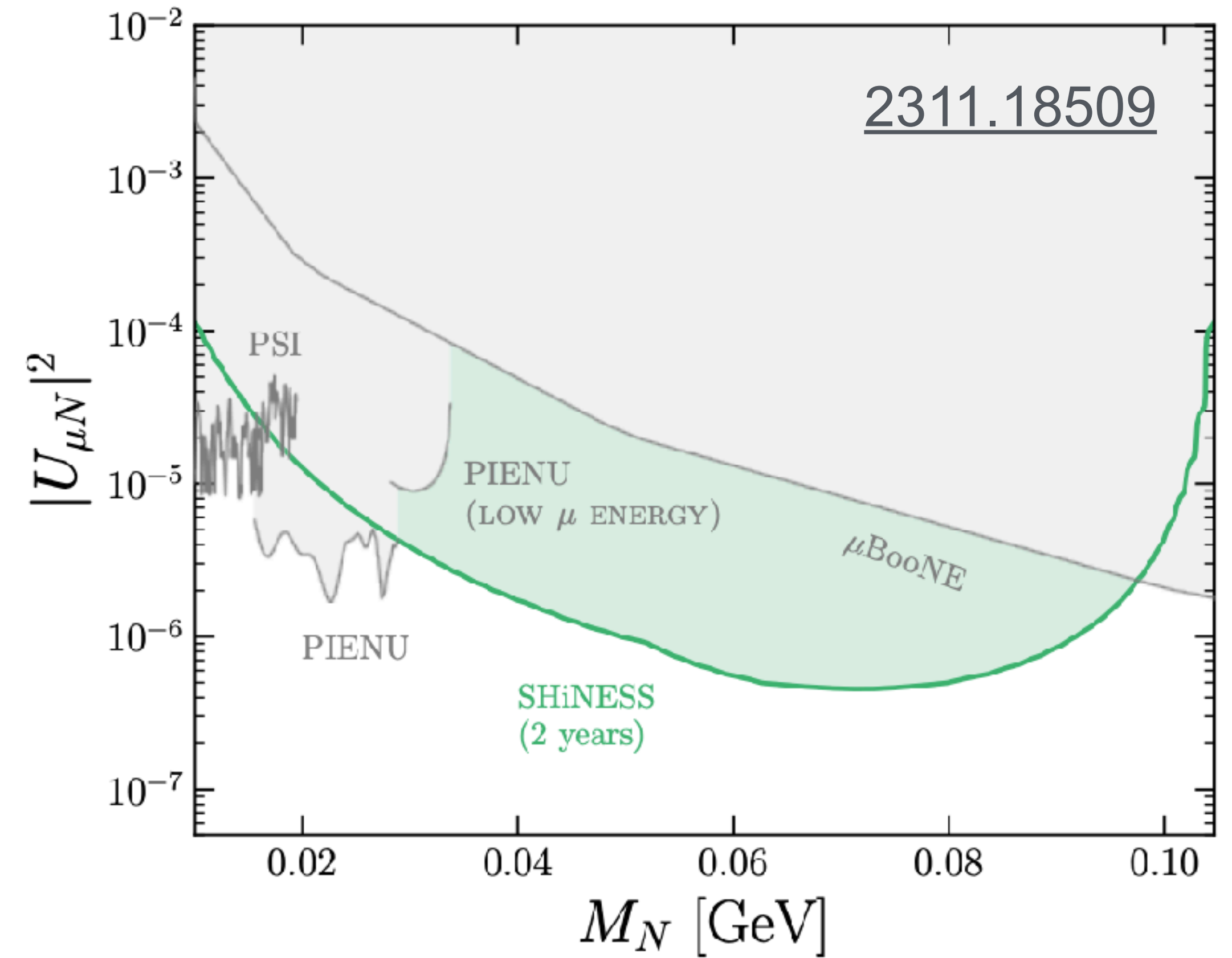
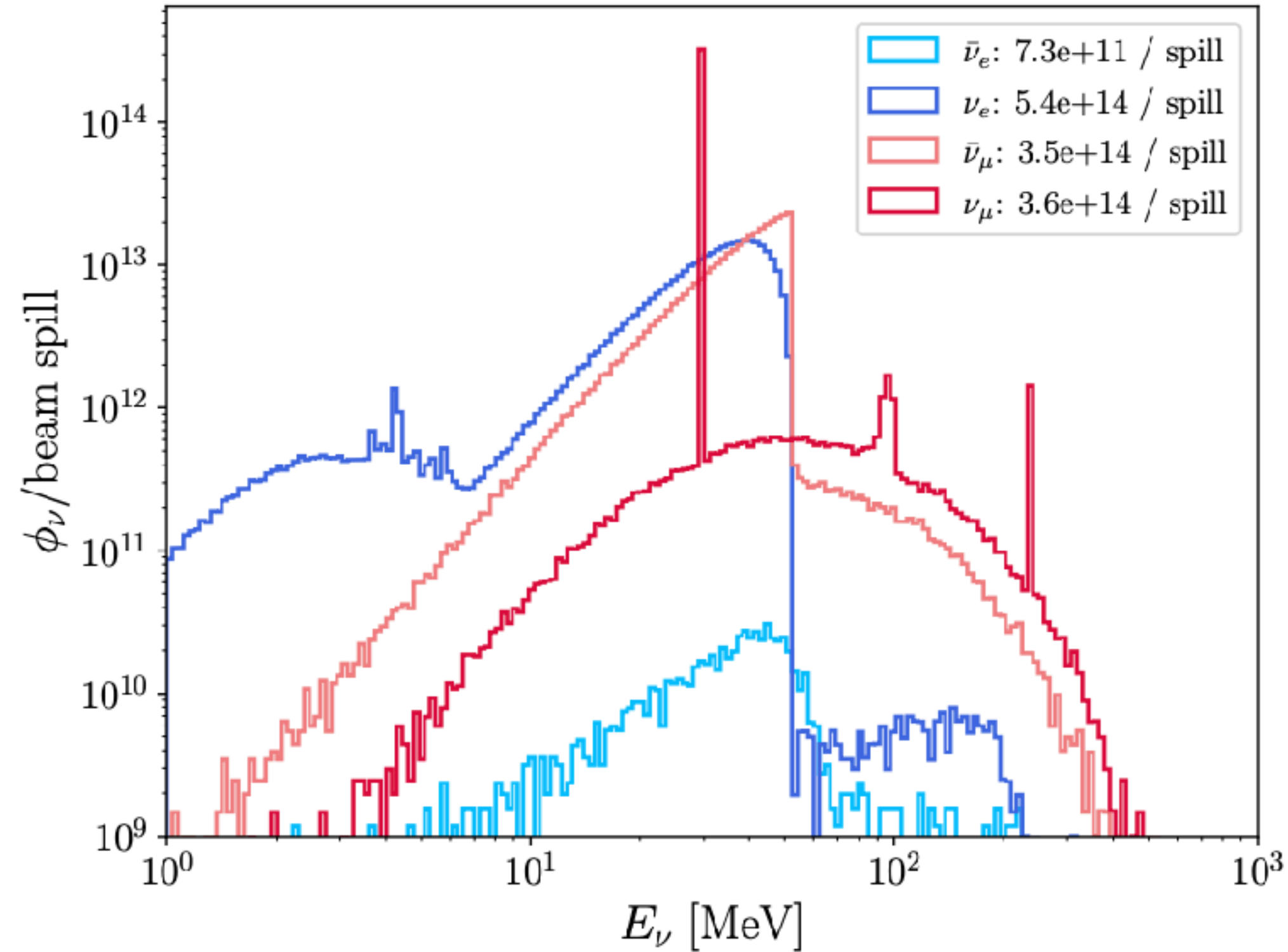


FIG. 1. (*Gluon dominance*) Sensitivities of MicroBooNE and JSNS<sup>2</sup> compared with existing limits and other projected sensitivities when all couplings are induced by a gluon coupling  $c_{GG}$  at a high scale. The MicroBooNE sensitivity is cut at 210 MeV



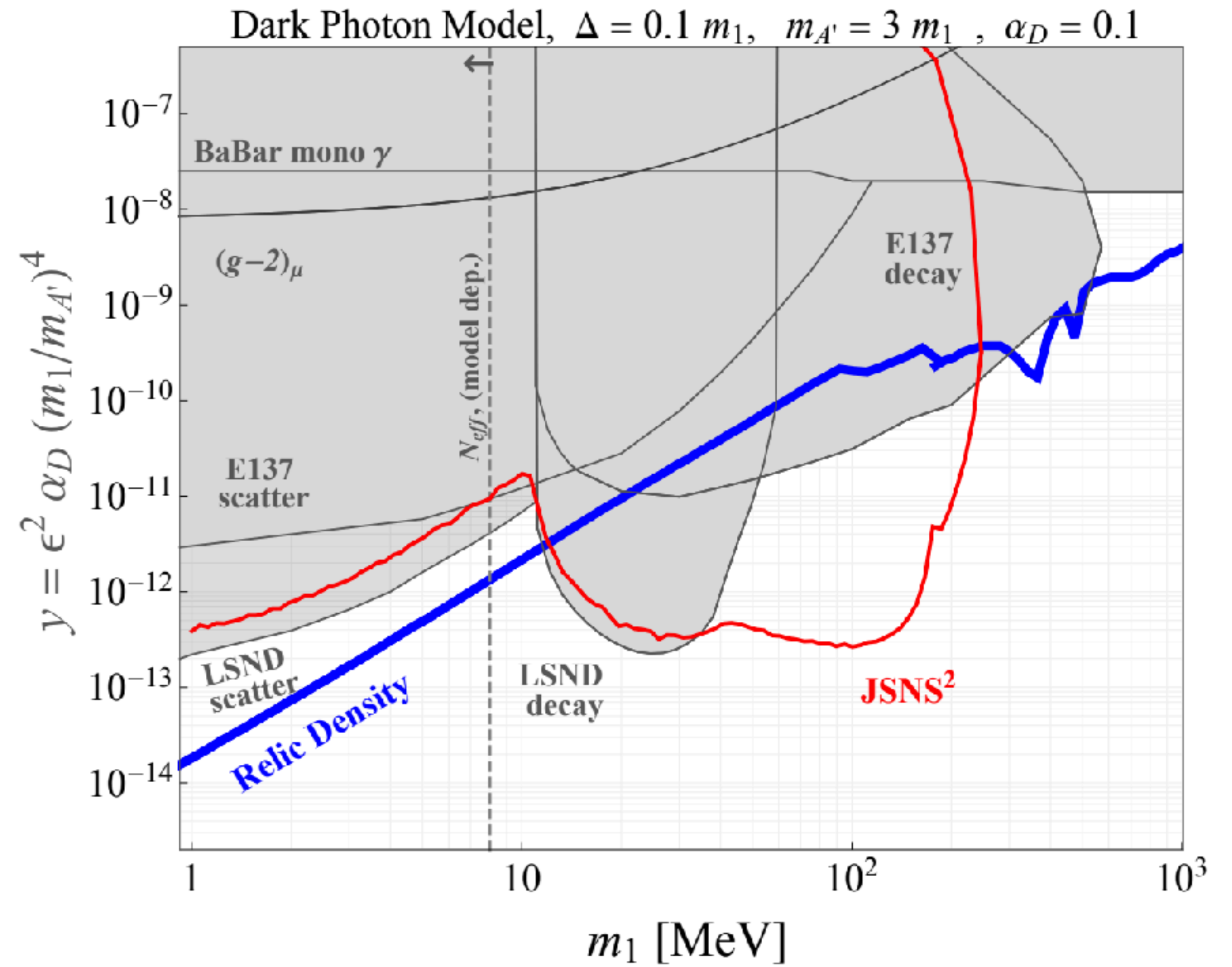
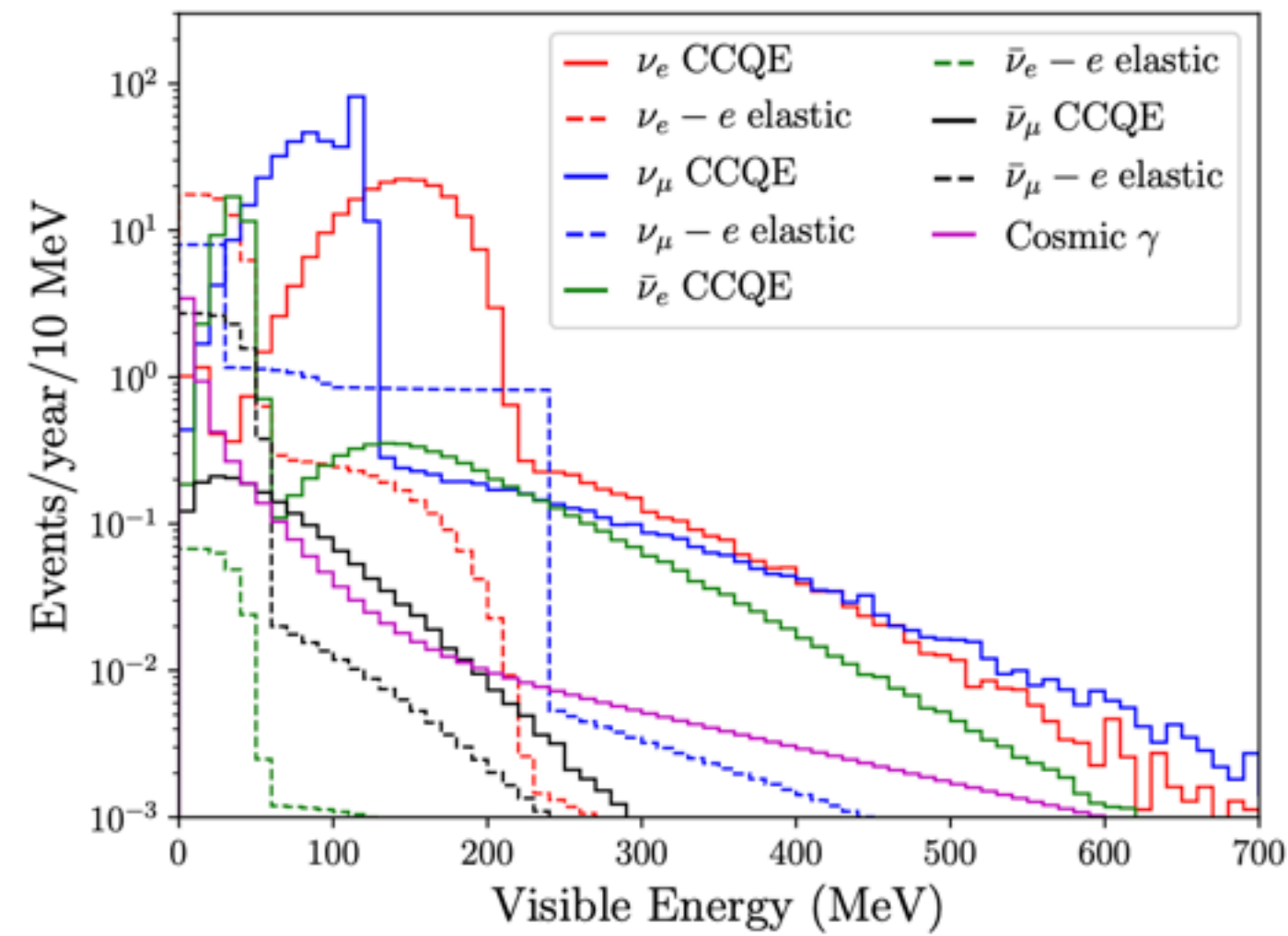
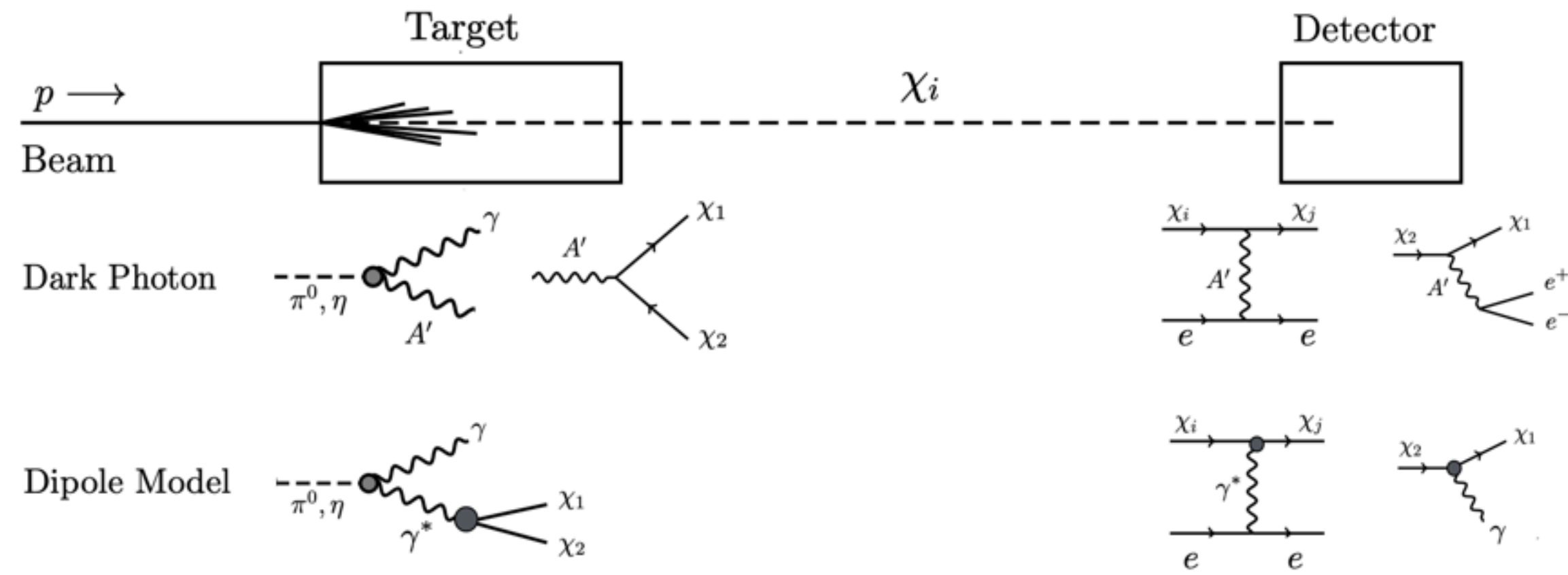
# European Spallation Source: Hidden Neutrinos

**SHiNESS**



(b) Muon mixing.

# Inelastic Dark Matter



J.R. Jordan, Y. Kahn, G. Krnjaic, M. Moschella, and J. Spitz. [arXiv:1806.05185](https://arxiv.org/abs/1806.05185)

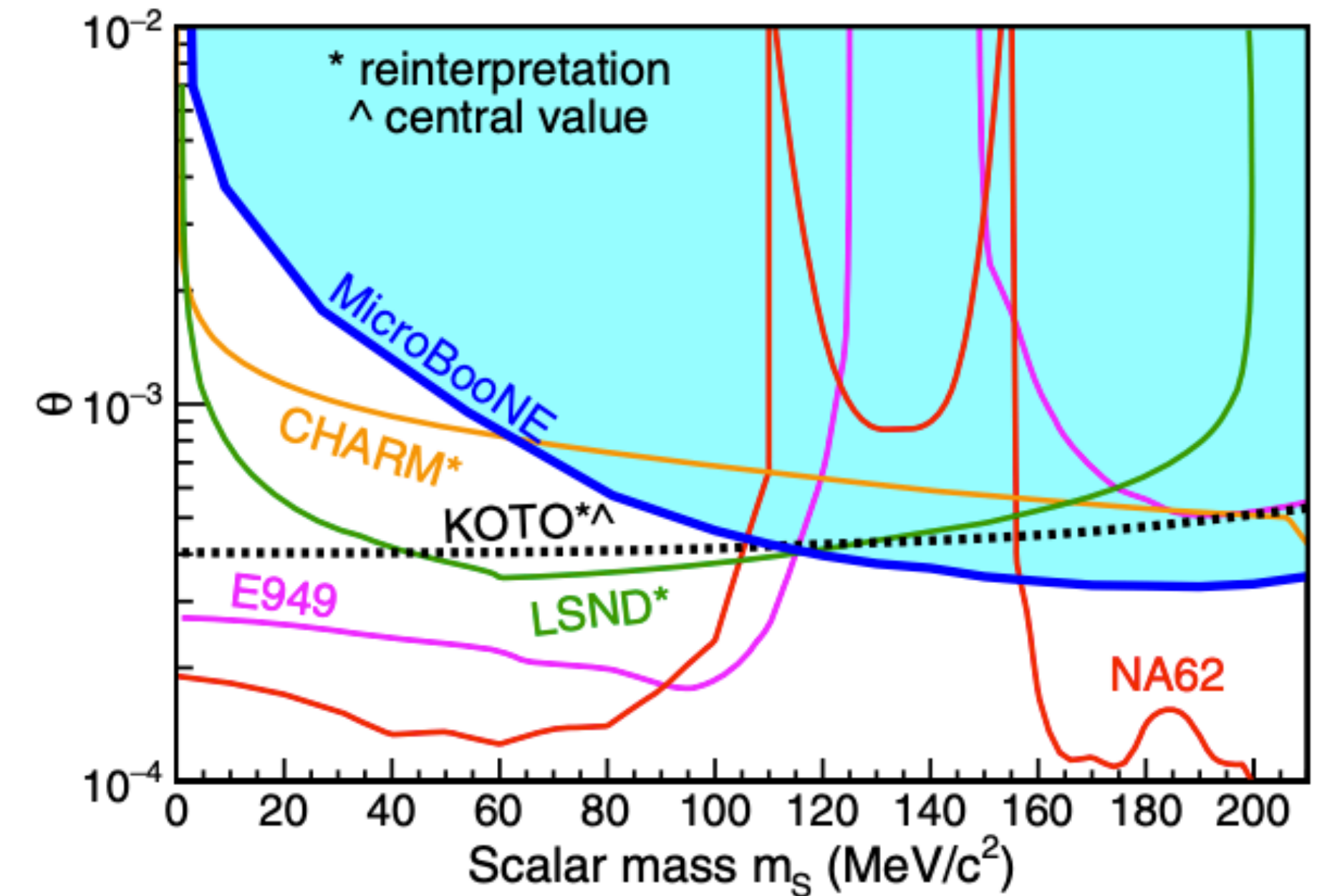
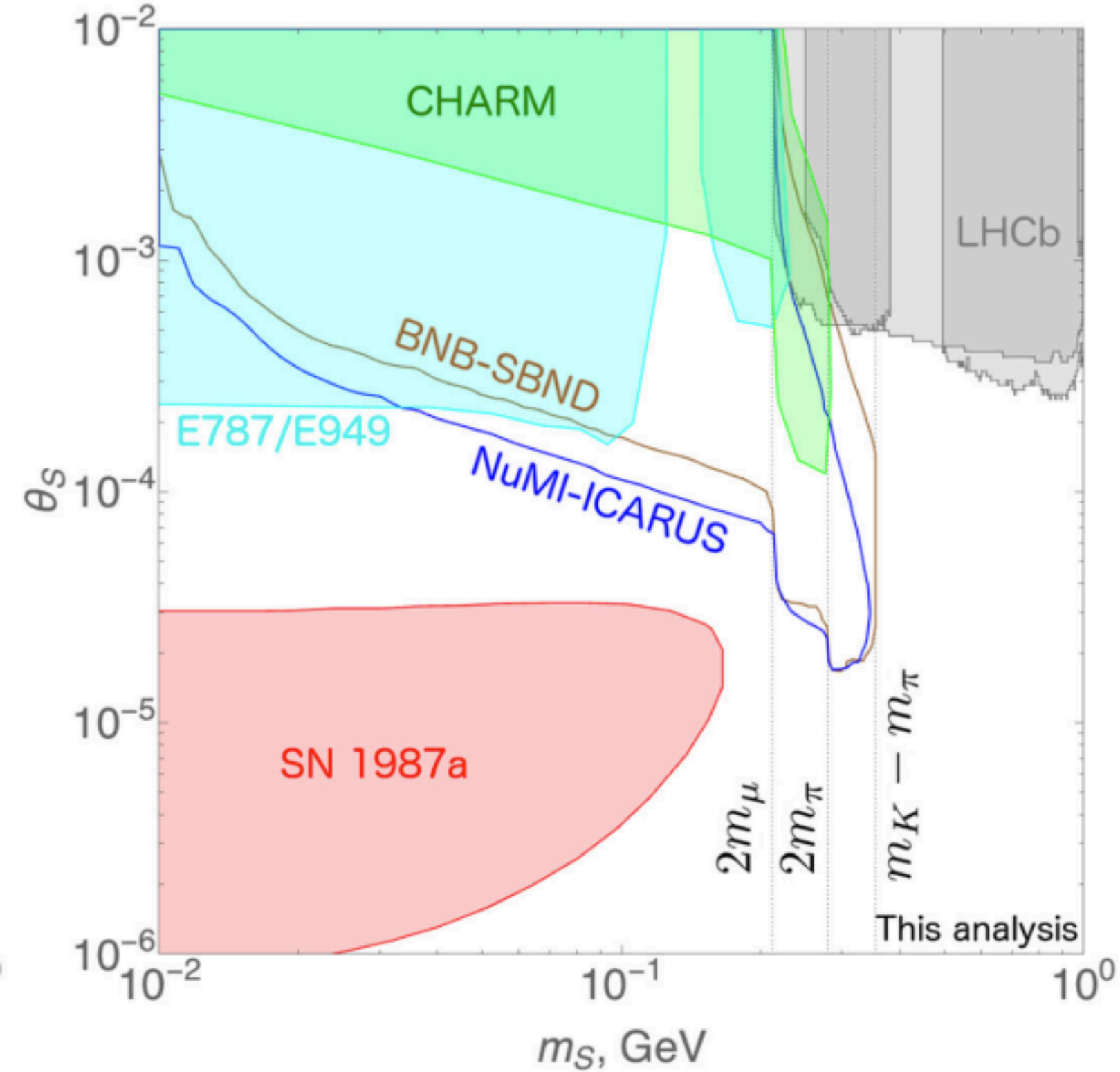
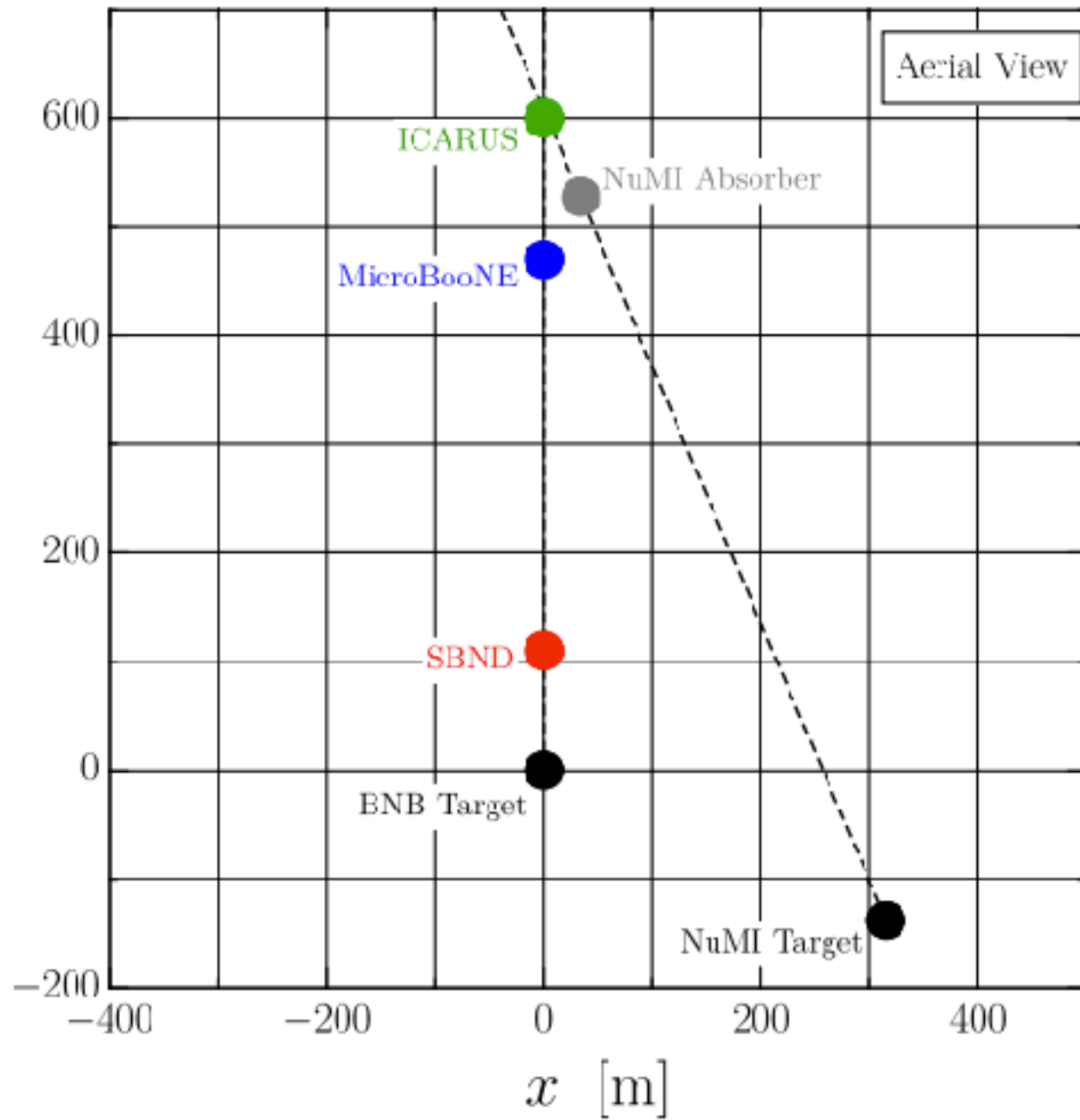
# MicroBooNE and KDAR from NuMI absorber

Fermilab's fortunate coincidence:

SBN program detectors are close to a KDAR source:

Batell, Berger, Ismail [PRD.100.115039](#)

MicroBooNE [PRL.127.151803](#)



About 0.13 KDAR per NuMI POT.

TABLE I. Parameters of subsystems for CEvNS detection.

Nuclear target	Detector Technology	Target Mass (kg)	Distance from source	Energy threshold (keV <sup>†</sup> )	Deployment dates
CsI[Na]	Scintillating crystal	14	20 m	5	2015-2019
Ar	Single-phase LAr*	24	29 m	20	2016-2021
Ge	HPGe PPC <sup>‡</sup>	18	22 m	<5	2022
NaI[Tl]	Scintillating Crystal	3500	22 m	13	2022
Ar	Single-phase LAr*	750	29 m	20	2025
Ge	HPGe PPC <sup>‡</sup>	50	22 m	<5	2025
CsI	CsI+SiPM arrays at 40 K	10~15	20 m	1.4	2025

Finished Planned, \*liquid argon, ‡p-type point-contact, †nuclear recoil energy, approximate threshold

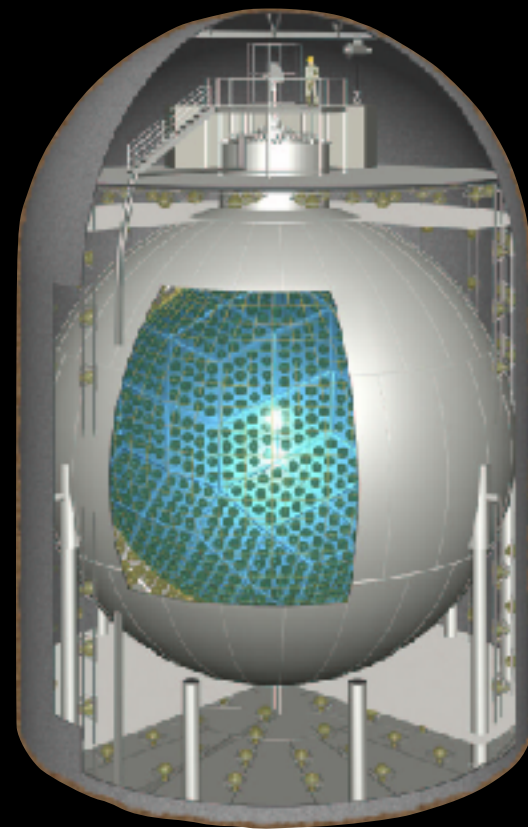
TABLE II. Additional detectors that broaden the physics reach of COHERENT.

Name	Detector Technology	Main purpose	Deployment dates
NaIvE	185 kg NaI[Tl] crystals	Measure $\nu_e$ + I CC cross section & beam-related backgrounds	2016 - present
MARS	scintillation panels interleaved with Gd-painted foils	Measure beam-related neutrons in Neutrino Alley	2017 - present
NIN	liquid scintillator cells in lead and iron shields	Measure neutrino-induced neutrons (NIN) in lead & iron	2015 - present
D <sub>2</sub> O	heavy water Cherenkov detector	Measure neutrino flux precisely & $\nu_e$ +O inelastic cross section	2022
LAr TPC	liquid argon time-projection chamber	Measure $\nu_e$ +Ar inelastic cross section	2025

Current Planned

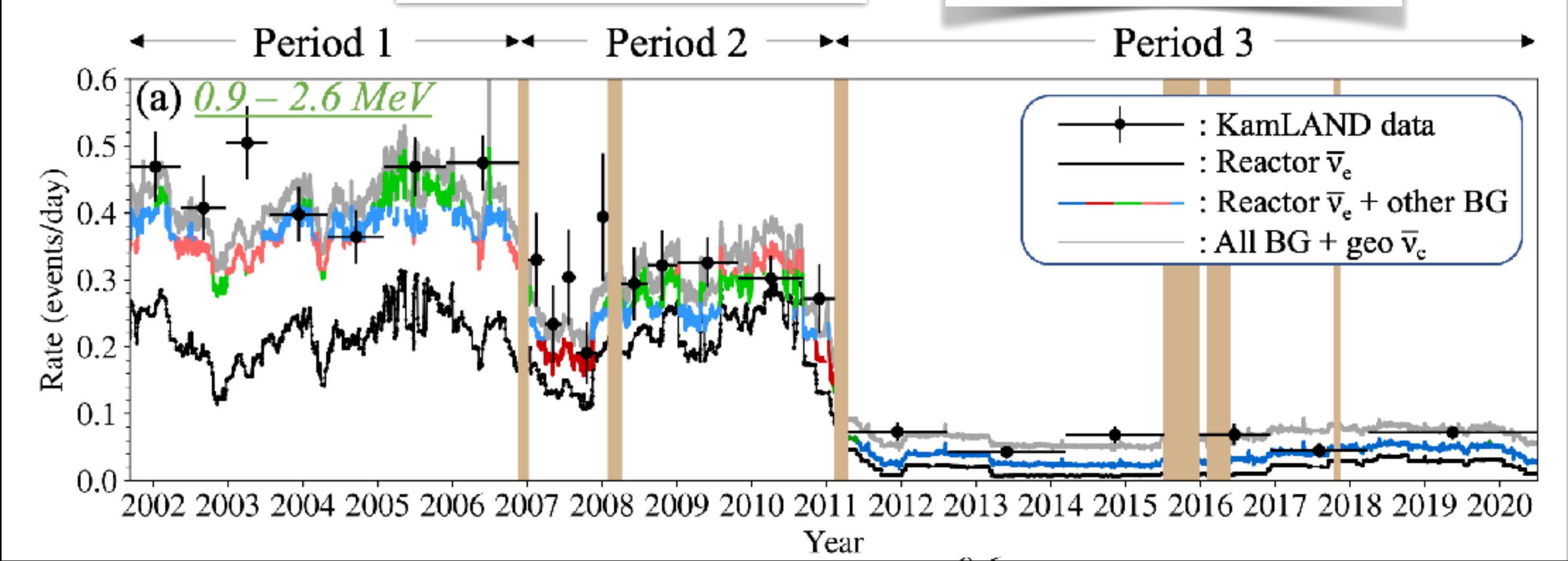
# Exothermic sources have their drawback...

Impact of the reactor shutdown in Japan on the KamLAND neutrino experiment



Purification periods

Tōhoku Tsunami Fukushima disaster



KamLAND coll. [10.1029/2022GL099566](https://arxiv.org/abs/10.1029/2022GL099566)

