Single-photo electron amplitudes and baseline RMS in ProtoDUNE-HD Federico Galizzi – Milano Bicocca





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In this presentation

Outline

- An input for Jose's simulation based on data
- Spe amplitude estimates
 - 9th July calibration: APA1 runs: 27901, 27902, 27903, 27904
 - 30th July calibration: APA 2,3,4 runs: 28368, 28369, 28370, 28371, 28489, 28491, 28492, 28493, 28494, 28495
 - Remember: signal amplitude changes with the overvoltage and the electronic chain gain (we can play with DAPHNE's gain)
- Baseline noise (RMS)
 - 8th July runs: APA 1,2 (run 27877) APA 3,4 (run 27877)
 - **Remember**: Noise level can vary **only** for DAPHNE's gain (if we exclude enrvironmental interferences)
- Choice of the runs: they are the ones used to tune DAPHNE's gain at the beginnig of August (see my <u>CM talk</u>) and they reflects the configuration used for many coscmic ray runs



Example from last week data



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SPE amplitudes

Waveform selection

- Cut waveforms with a signal in the pre-LED baseline
 - To void baseline mis-evaluation
 - Acceptance range $\pm 20 ADCs$
- Cut waveforms with large pulses in coincidence
 - To ensure I'm integrating only the LED signal
- About 10% loss of statistics

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• Ok, since I believe the results uncertainty is not statistically domitated







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SPE candidate

Selection

- After fitting the calibration histogram with a multi gaussian function^{*}, we select the waveforms whose histogram is within $\pm 1\sigma$ from the corresponding peak
- We observed certain correlation between the average number of detected photons and the estimated amplitude
 - Considering the possible 0 pe and 2 pe waveforms gives very little correction
 - There could be a subtle correlation in the fit parameters that bias the selection (?)
- Errors:

$$\sigma^{2} = \sigma^{2}_{Bsl_rnd} + \sigma^{2}_{Bsl_Eval} + \sigma^{2}_{Fit_sel}$$

- $\sigma_{Bsl_rnd} = \frac{Noise RMS}{(\#Spe \ candidates)^{0.5}} [tipically \frac{\sim 4 \ ADC}{10^3 \ candidates}]$
- σ_{Bsl_eval} = Evaluation of the baseline (could be a bias?)
- σ_{Fit_sel} = selection not perfectly centered in the 1pe peak

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Insight The LED intensity – Spe amplitude correlation

Note: it's a correlation in the results, it's not phyisical



Runs 28368-69-70-71 Yellow points are from run 28371 – mask 12





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RMS of the baseline

A noise-level indication

- Dedicated runs with SiPM biased below their brekdown
- Full-baseline subraction
- Projection and RMS evaluation
- Looks pretty gaussian









Conclusions

... and "to dos"

- We have the data to characterize the signal response and the noise level
 - For the SPE shape (spe template) look at <u>Henrique</u>'s tau slow studies or <u>Maritza's</u> template.
 - We also have noise FFTs
- Improve the analysis
- Analyses for other calibrations

... It's a long way (youtube)







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