### Post talk comments

#### Comments

- Generally seemed positive.
- Main comments from Jake:
  - Consider dividing templates into energies too.
  - Some magic about the energy slice which might skip unfolding.
  - Potentially some confusion about MC/data discrepancies, still communicating.
- Started looking through tech note, still trying to understand the fit minimisation
- Planning to chat with Jake soon

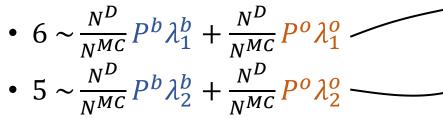
## Fitting discussion

#### Fitting method

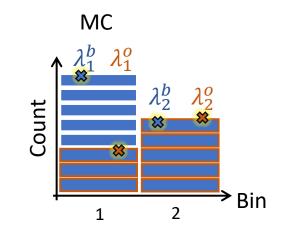
- The fit uses (python) <u>Minuit's template fit</u>, using <u>Dembinski and Abdelmotteleb</u> method.
- D. and A.'s method approximates the <u>Beeston-Barlow method</u>.
  - Henceforth, will discuss pure Beeston-Barlow, trusting the D. and A. method is sensible
- Methods can also deal with weighting the MC templates (no longer integer)
  - Currently only considering unweighted templates

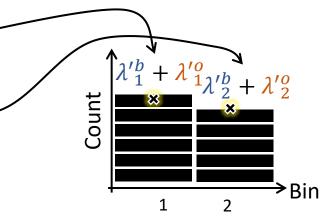
#### Fitting method

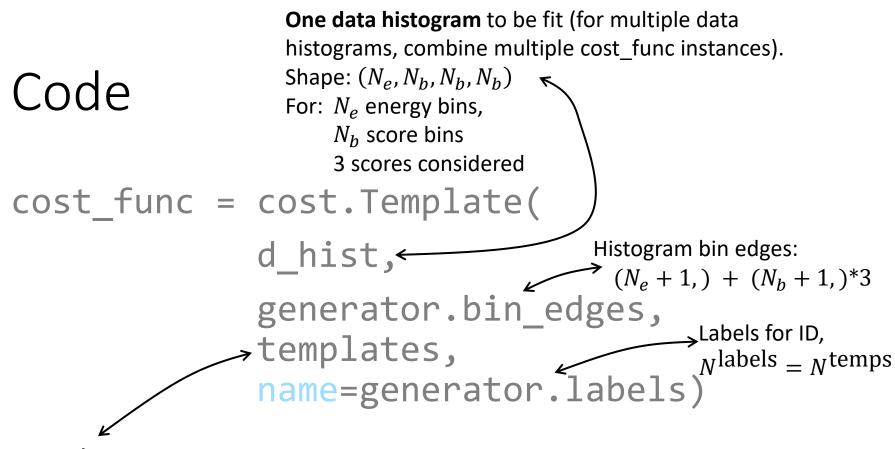
- Example fit 2 bins, 2 channels
- MC sample has counts (8, 5)<sup>b</sup>, (3, 5)<sup>o</sup>.
- Data has counts (6, 5)
- From MC, create  $\lambda_1^b, \lambda_2^b, \lambda_1^o, \lambda_2^o$ Eqs. 17 and 2 of <u>BB</u> • Note:  $\lambda_2^{b/o} = N^{MC} - \lambda_1^{b/o}$ 
  - Compare data:



• We want data yields  $N^D P^b$ ,  $N^D P^o$ 





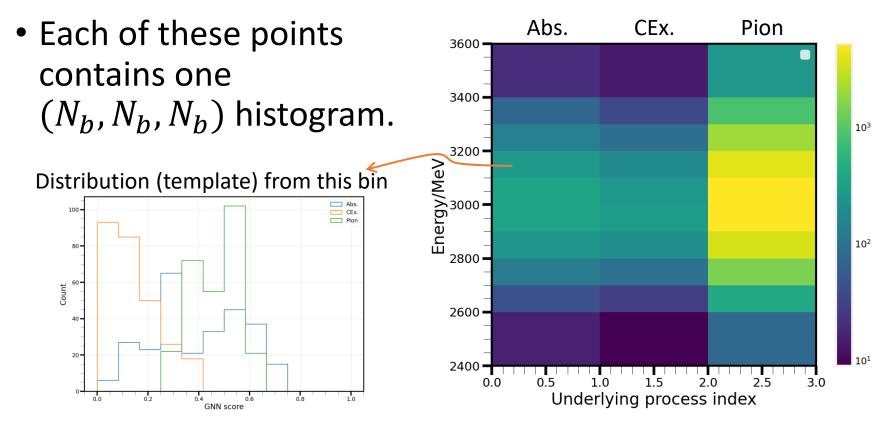


List of  $N^{\text{temps}}$  histograms as templates. There will be  $N^{\text{temps}}$  yields given by the fit, one for each templates

Shape:  $[(N_e, N_b, N_b, N_b)] * N^{\text{temps}}$ (Each template has the same shape as the data histogram, but there are  $N^{\text{temps}}$  in the list)

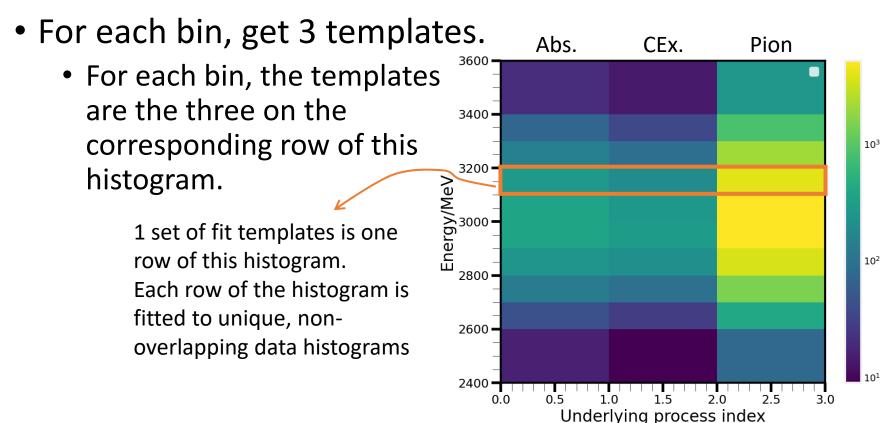
#### Fitting options

• 2D histogram displays the total count of events as a function of energy and underlying process.



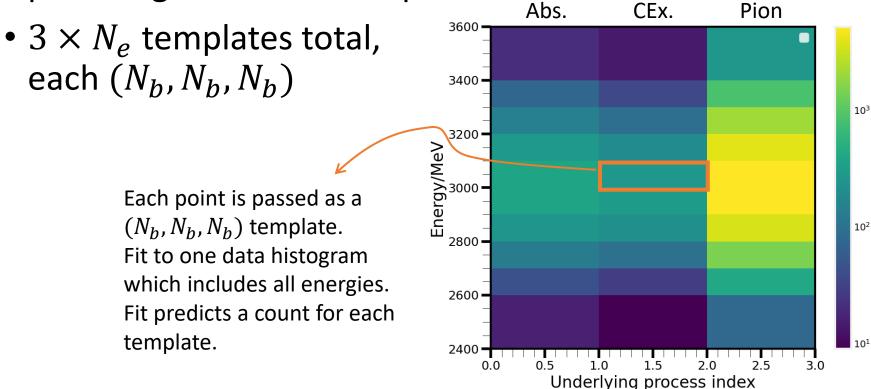
#### Fitting options – current

• Current idea, do  $N_e$  separate fits, each to one data histogram, shape  $(N_b, N_b, N_b)$ .



#### Fitting options – free-for-all

 A valid (but poor) fitting option would be to do one fit to all data (N<sub>b</sub>, N<sub>b</sub>, N<sub>b</sub>), where each energy and process gets its own template.

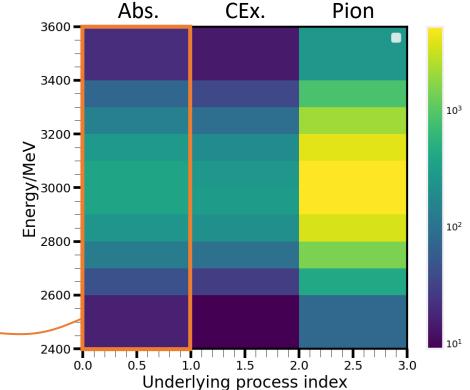


#### Fitting options – energy fixed

- An attempt at simultaneous energy fitting could use one data histogram, which includes energy bins: (N<sub>e</sub>, N<sub>b</sub>, N<sub>b</sub>, N<sub>b</sub>).
- 3 templates total, each (N<sub>e</sub>, N<sub>b</sub>, N<sub>b</sub>, N<sub>b</sub>)
- Bad, since this doesn't allow the energy shape to change

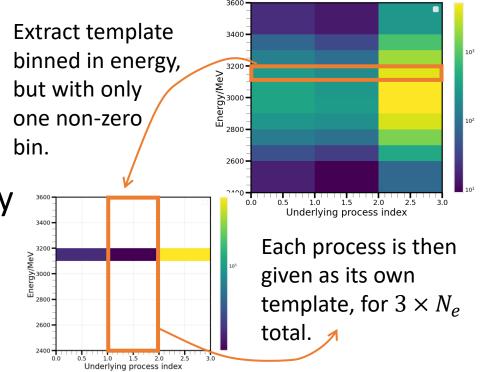
**Relative fractions of events** 

fixed in by templates – bad! <



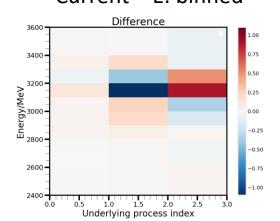
#### Fitting options – energy binned

- Use one data histogram, including energy bins:  $(N_e, N_b, N_b, N_b)$ . But separate templates for each energy bin.
- $3 \times N_e$  templates total, each  $(N_e, N_b, N_b, N_b)$
- Each template has non-zero values in exactly one of the indices across the first dimension (N<sub>e</sub>).



#### Energy binned vs. separate fits

- Use 50% MC as template, 50% as "data".
  - Not done any energy weighting.
- Performed current fit (separate fits for each energy)
- Perform the energy binned (final option mentioned).
- Investigated the difference between the two: Current – E. binned Current / E. binned - 1



[ 3.62256938e-05 2.47275383e-03 2.55603376e-04 [-1.63322683e-04 1.58220131e-03 -7.34721345e-05 [ 3.74767953e-05 3.63326015e-04 -2.24963122e-05 [ 9.47627484e-05 3.42696054e-04 5.75051774e-06

[ 9.47627484e-05	3.42696054e-04	5.75051774e-06]
[ 8.28801557e-05	7.26401418e-04	-2.93631169e-05]
[ 1.00984362e-04	1.45238574e-03	-6.34326830e-05]
[ 4.74770175e-04	-8.51413460e-03	2.29216628e-04]
[-4.59448680e-04	-6.49467485e-03	2.36643514e-04]
[ 9.60154333e-04	2.06738156e-03	-9.37032226e-05]
[ 4.94567951e-04	3.48525444e+01	-2.39613045e-04]]

#### Other options – energy unfolding

• In the energy fitting method, templates are picked by the same binning as the y-axis

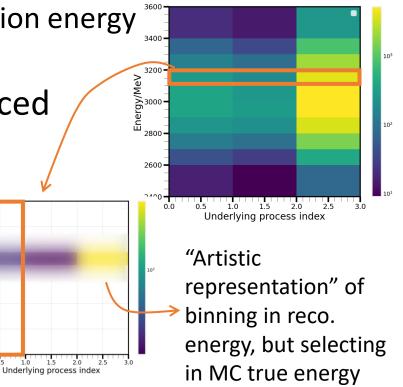
2800

2600-

2400

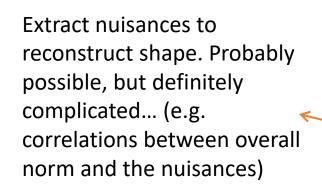
0.5

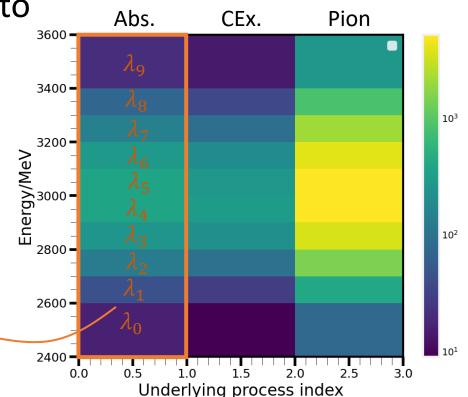
- In this case, beam instrumentation energy rather than interaction energy.
- The histogram could be produced from MC truth interaction energy, but split into 3600 templates via reco. 3400 3200 Interaction energy 3200 WeV/WeV



# Other options – using the nuisances

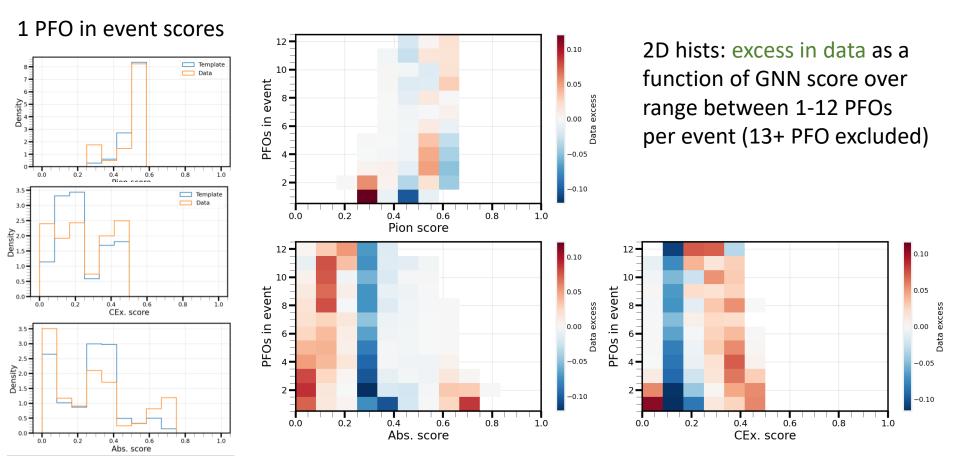
- The fit must produce nuisances per bin of the fit for each template.
- In principle, we could try to extract these and "manually" reconstruct the energy binning





#### **PFO count variation - comparison**

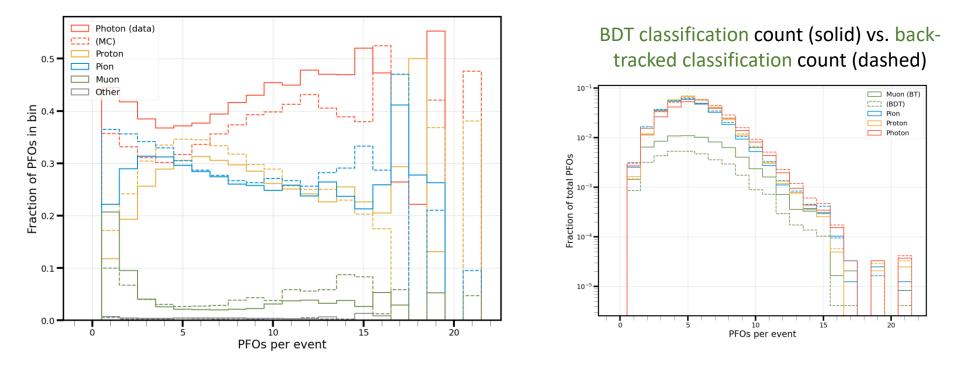
 Plots compare all MC events (not split by true process) vs. data events.



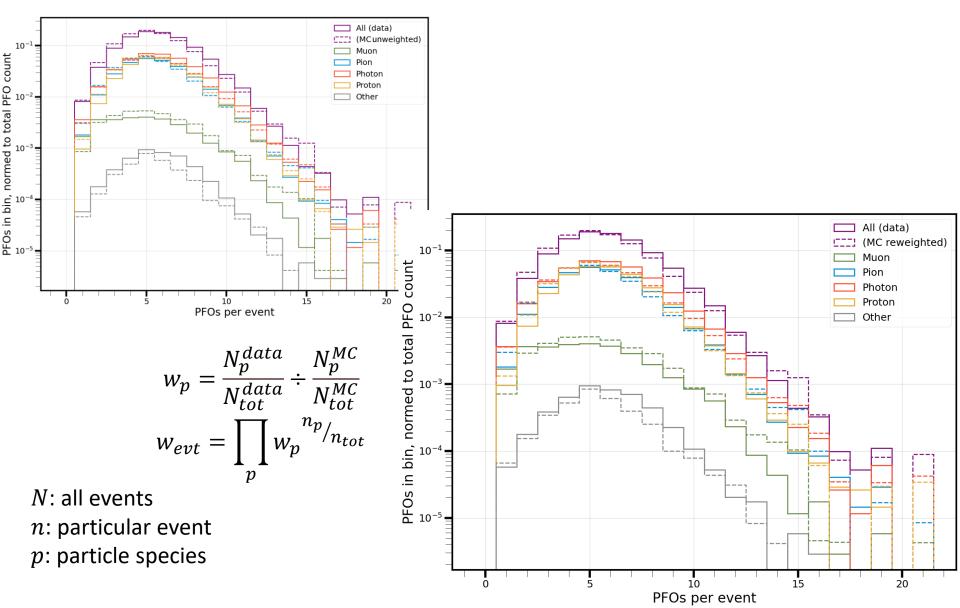
Dennis Lindebaum | Fit performance of GNN scores

#### **Particle content**

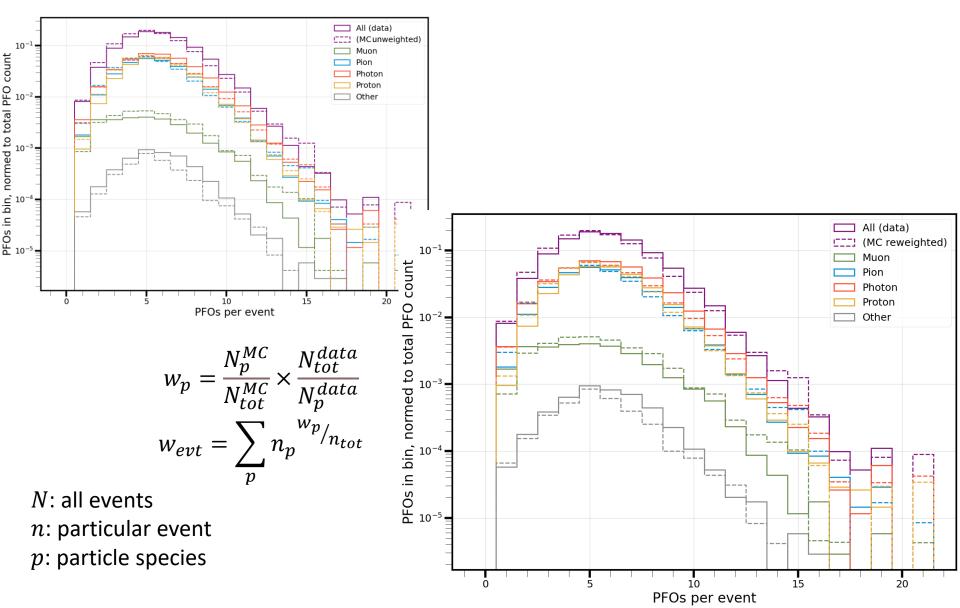
- MC vs. data discrepancy could be caused by mismodelling of the species expected from nuclear events.
- Use a simple BDT (same BDT used for PID in the full network) to estimate proportions of particles in MC vs. data.



#### **Reweighting events**



#### **Reweighting events**



#### **After weighting**

If the re-weighting accounts to the MC/data discrepancy, the MC/reweighted difference should match the MC/data difference.

0.8

0.8

1.0

1.0

12

10.

2.

12.

10 -

8-

6-

2 0.0

PFOs in event

0.0

0.2

0.2

0.4

0.4

0.6

0.6

Abs. score

Pion score

PFOs in event

1.0-

0.8-

0.6-

0.4 -0.2 -

12

10-

8-

6-

4 -

0.0

0.2

0.4

CEx. score

0.6

0.8

1.0

PFOs in event

0.02

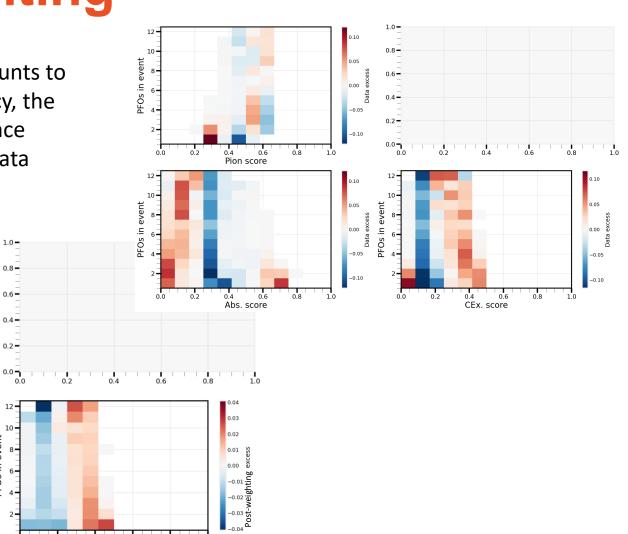
0.01

0.00

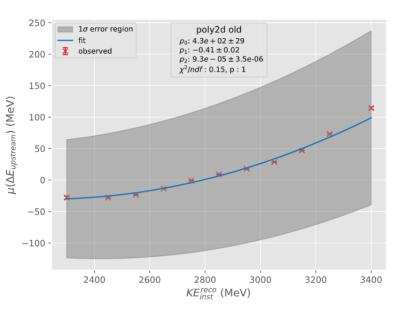
0.02

Post-weighting excess 0.00

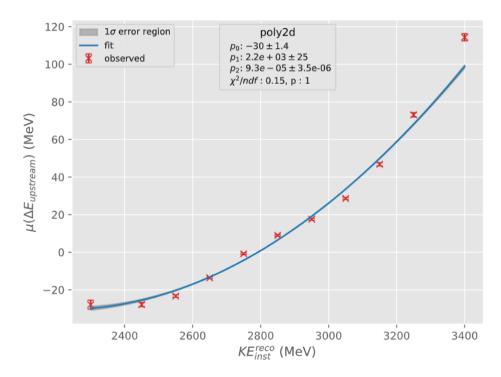
Post-weighting exces



#### Upstream correction fit



Recall previously, fit of upstream energy correction had these excessive errors (left). Changing the equation fixes this: Left:  $p_2x^2 + p_1x + p_0$ Below:  $p_2(x - p_1)^2 + p_0$ 



#### Upstream correction

- Systematic

   offset between
   Gaussian mean
   (black) and
   arithmetic
   mean (blue)
- Not seen in 2GeV
- Scrapers?

