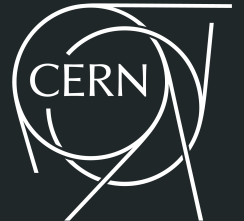


# Atmospherics events in the VD geometry

---

P. Granger - CERN



# Generation

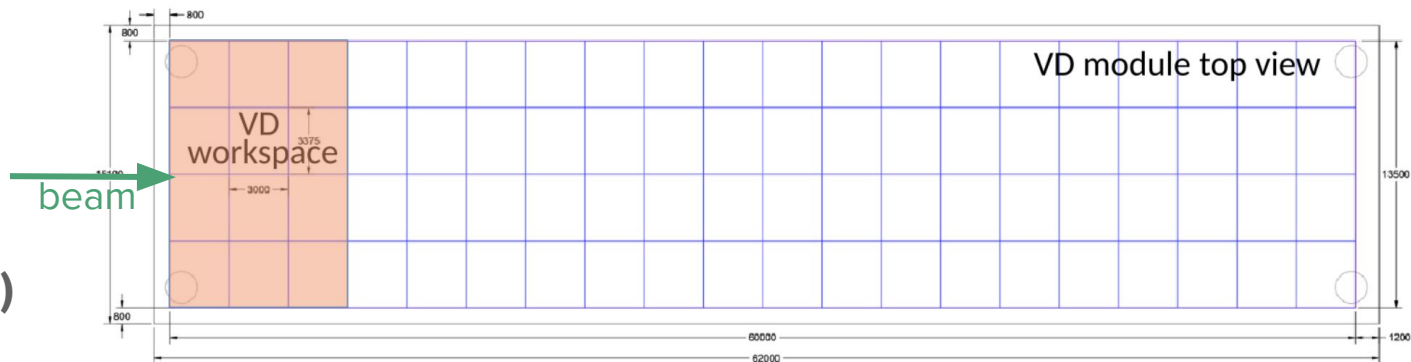
`prodgenie_atmnu_max_weighted_randompolicy_dunevd10kt_1x8x6_3view_30deg.fcl`

- Same config as for the latest atm. HD production
- Reweightable power spectrum flux
- All flavours generated according to their xsec ratios

## 1x8x6 geometry

- Full width (13m)
- 15% length (9m)
- **Half height (6.5m)**

↓  
Not ideal for up-going atmospheric



# Reconstruction

```
standard_reco2_atmos_dunevd10kt_1x8x6_3view_30deg.fcl
```

- HD atmospheric Pandora config (including vertex CNN)
- VD beam flavour CVN
- Energy reconstruction
- Direction reconstruction (same algos as HD)

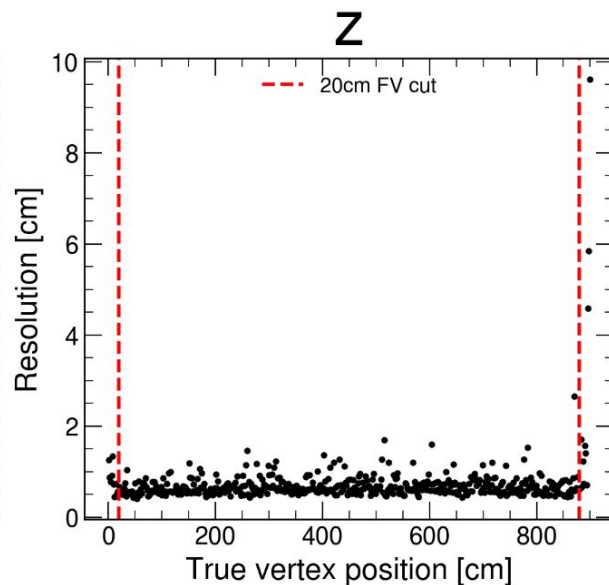
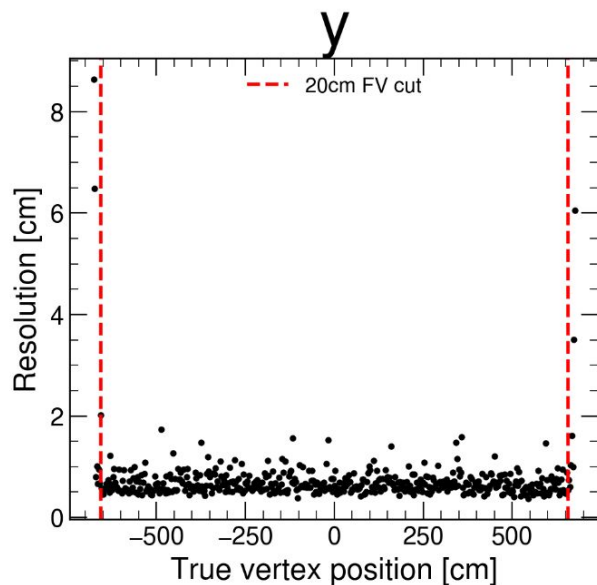
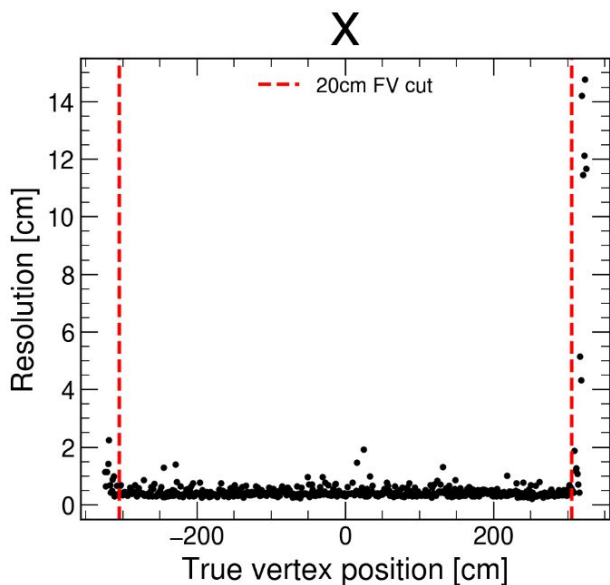
Available test sample (~250k events) at:


```
/pnfs/dune/persistent/users/pgranger/atm_VD_test/reco2/
```

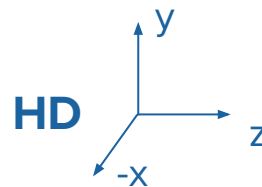
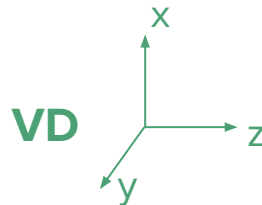
Analysis file used for all later plots (from `atmoAnalysis` module):

```
/pnfs/dune/persistent/users/pgranger/atm_VD_test/reco2/ana_sum.root
```

# Containment

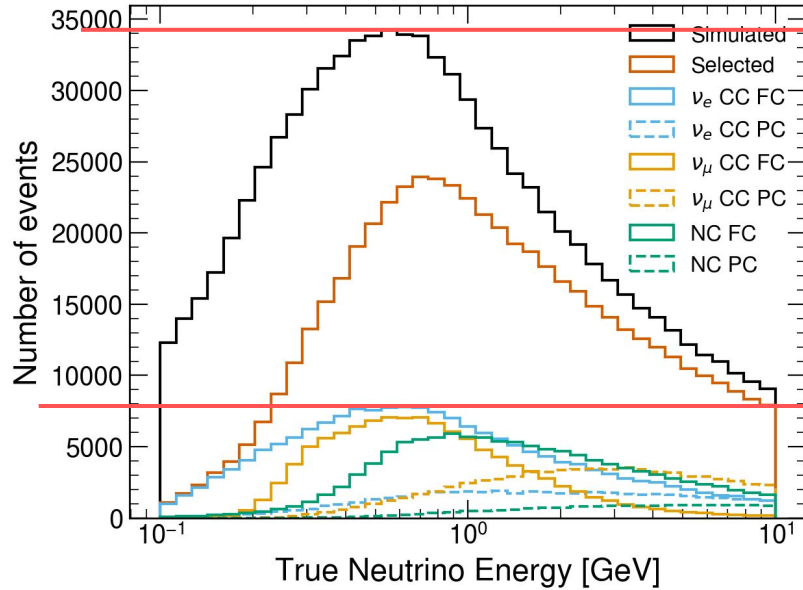


- 20cm fiducialization on all borders
-  Up direction is x in VD

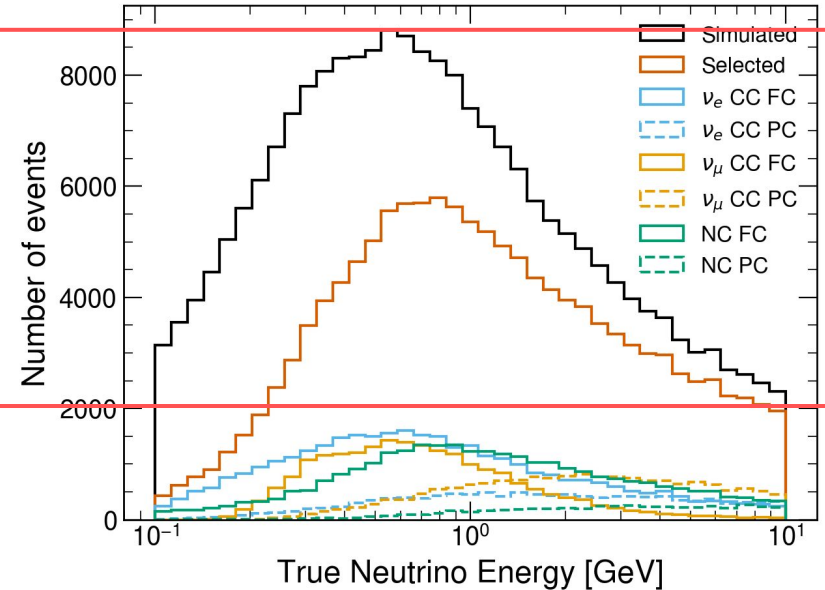


# Containment

HD



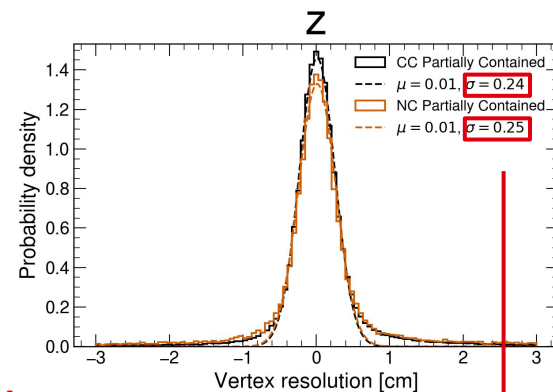
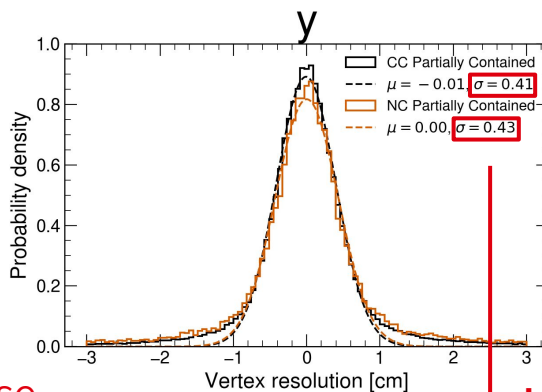
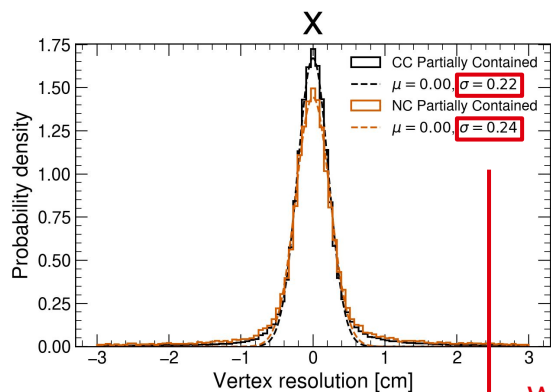
VD



Lower containment in the simulated VD geometry (as expected) by smaller size

# Vertex reconstruction

HD

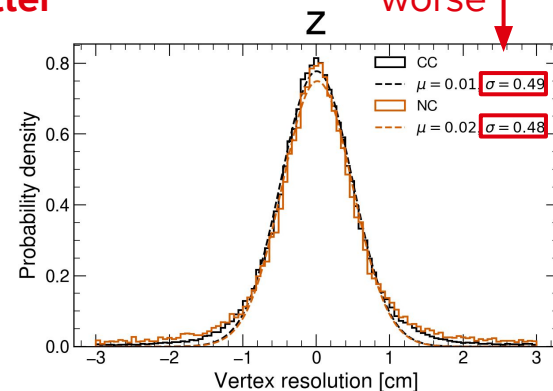
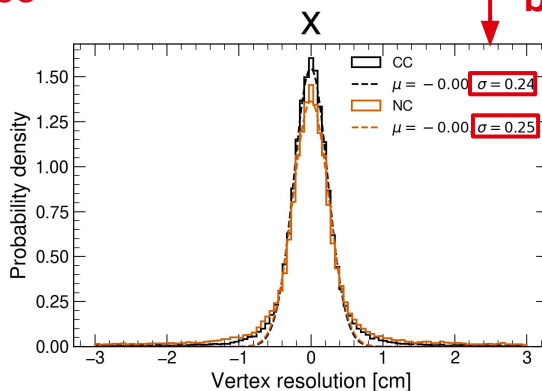
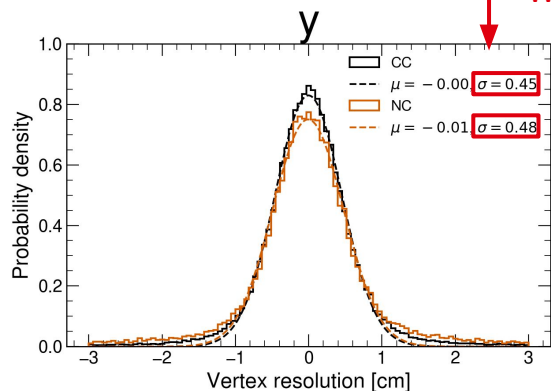


worse

better

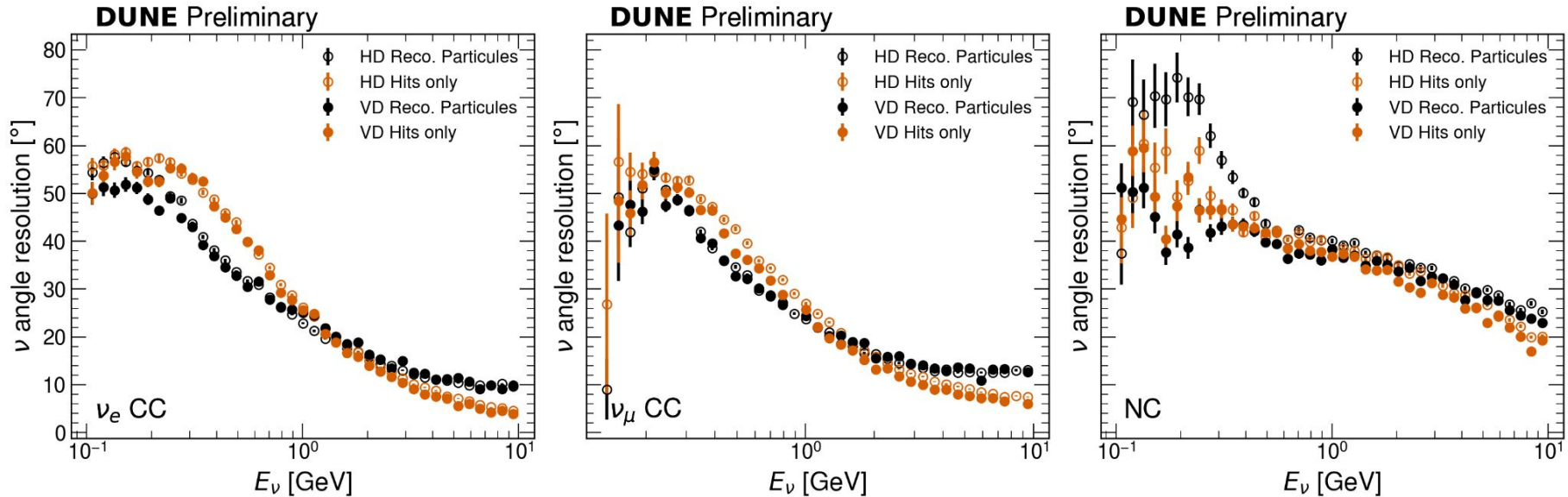
worse

VD



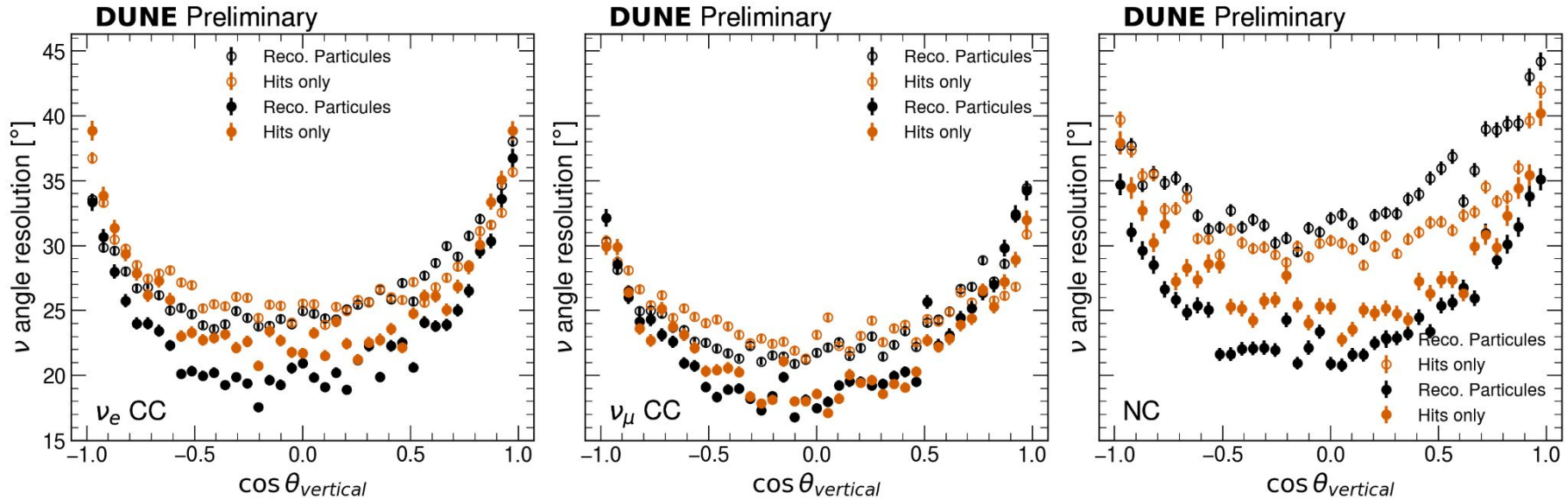
Very good perfs. (HD atm. CNN, no VD retraining) — Improved resolution in vert. dir.

# Direction reconstruction



- Slightly better VD direction reconstruction performance (despite lower containment)
- Significant VD gain at low-E using reco. particles.
- VD gain at high-E using the calorimetric method

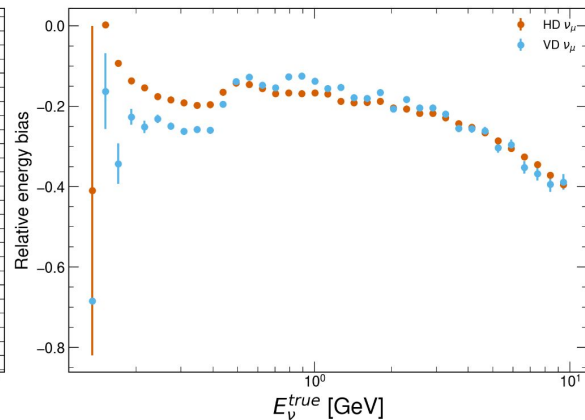
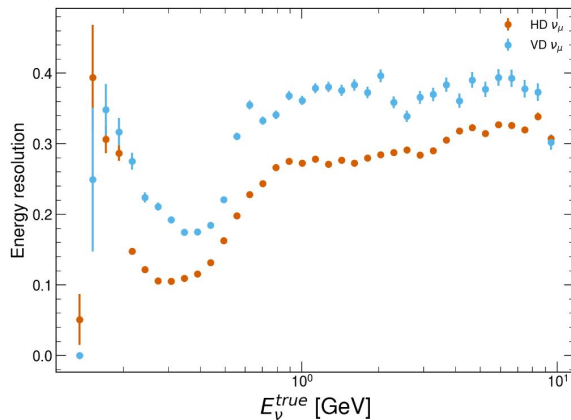
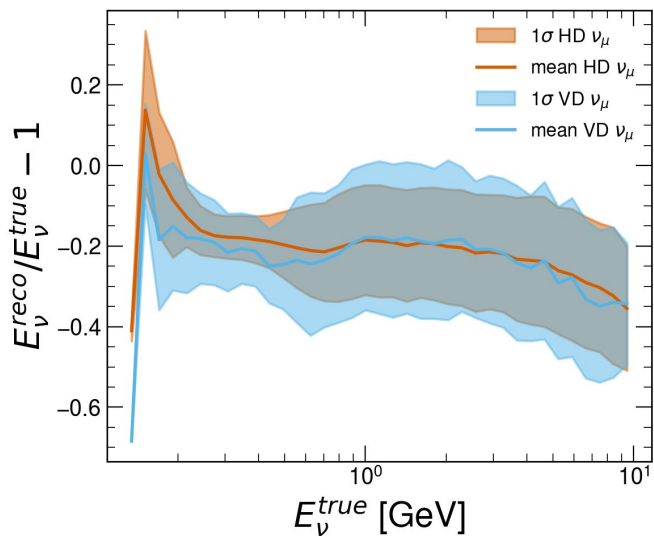
# Direction reconstruction



- Worse VD performance for vertically aligned tracks (along drift)
- Better performance for other tracks
- Higher dependency of the performance on the neutrino direction

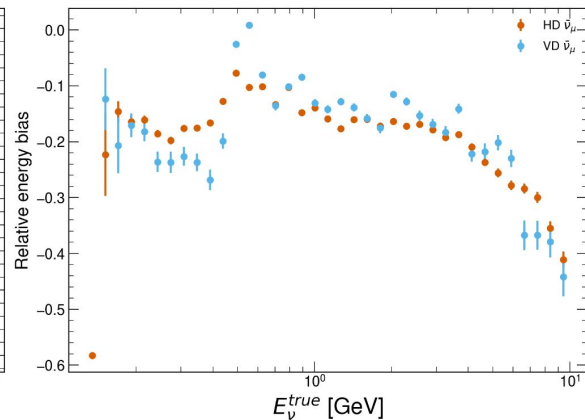
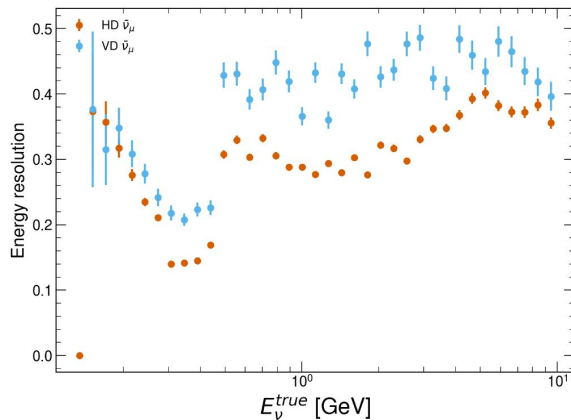
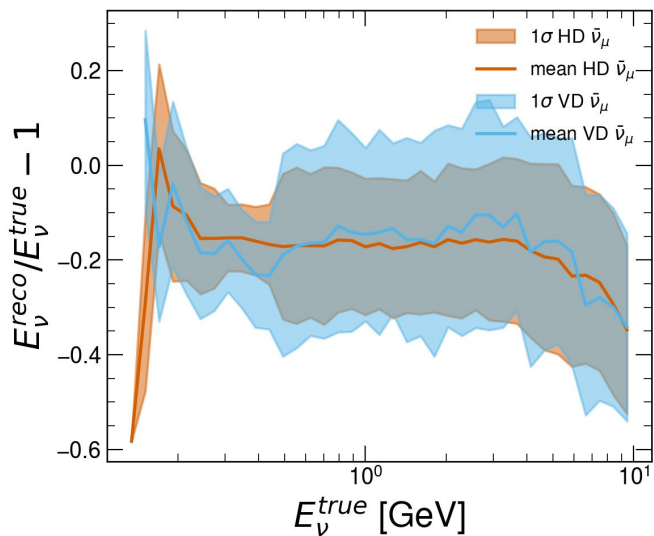


# Energy reconstruction (numu)



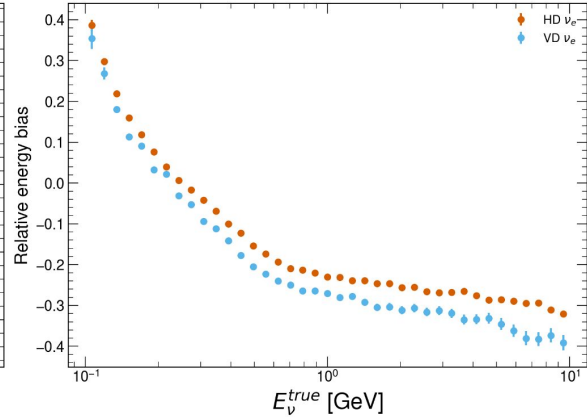
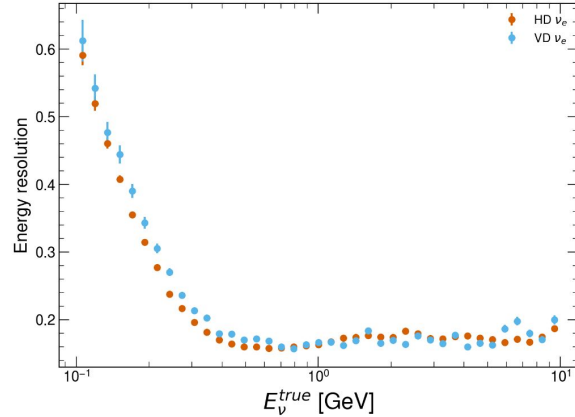
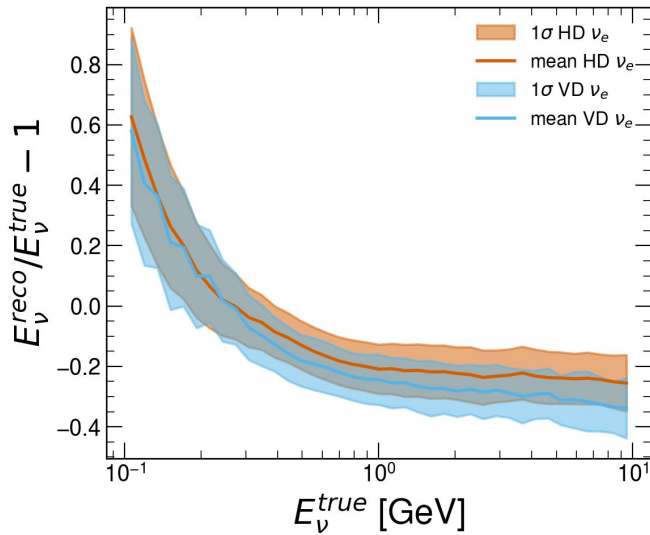
- Just using the reconstructed hadronic + leptonic sum => **no scaling correction**
- Currently better performance of HD => can be due to containment, no retuning of the MCS method (different pitch and resolution)

# Energy reconstruction (numubar)



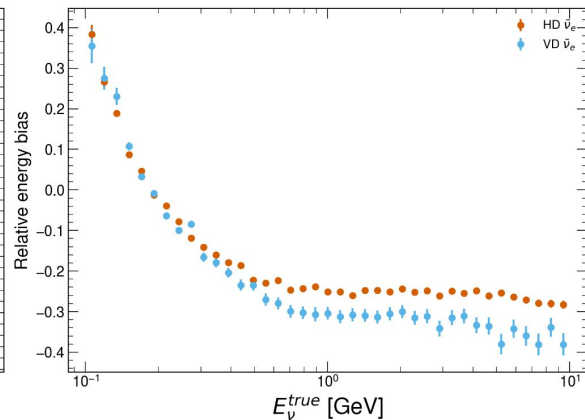
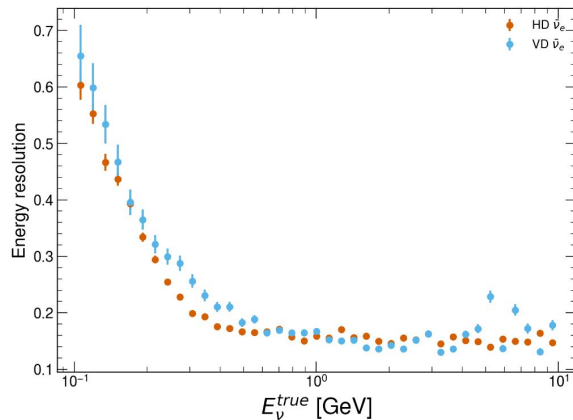
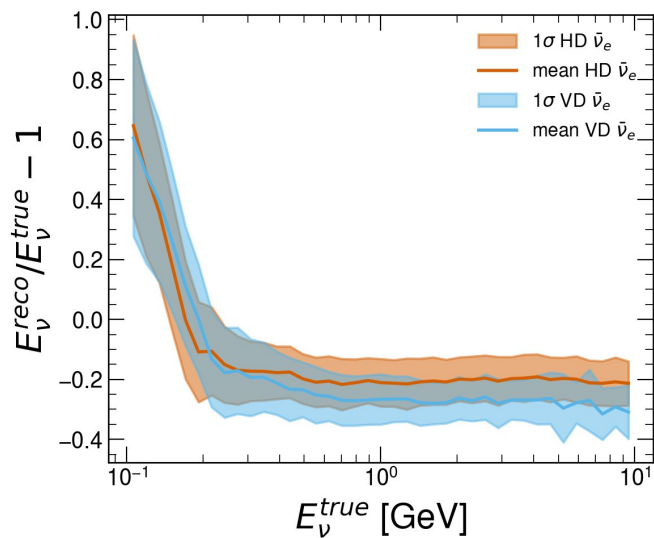
- Just using the reconstructed hadronic + leptonic sum => **no scaling correction**
- Currently better performance of HD => can be due to containment, no retuning of the MCS method (different pitch and resolution)

# Energy reconstruction (nue)



- Just using the reconstructed hadronic + leptonic sum => **no scaling correction**
- More similar resolution but stronger bias => probably different scalings required for HD and VD

# Energy reconstruction (nuebar)



- Just using the reconstructed hadronic + leptonic sum => **no scaling correction**
- More similar resolution but stronger bias => probably different scalings required for HD and VD

# Conclusion

Just a quick presentation to say that:

- Atmospheric events can be simulated and reconstructed in VD 1x8x6 geometry with relevant fcls in dunesw
- **Without requiring any further work:**
  - Good performance of HD vertexing CNN on VD sample. Better vertex resolution in vertical direction => good for atmospheric
  - Better vertical direction resolution performance in VD thanks to the different orientation
- **Several things to investigate:**
  - Degraded energy resolution => Possibly due to lower containment, no adjustment of reco methods (MCS), ...
  - Flavour CVN => Can the HD atm. one be used directly?

*Relevant topic as VD will be the first installed detector => many things to look at for whoever is interested*

# Backup

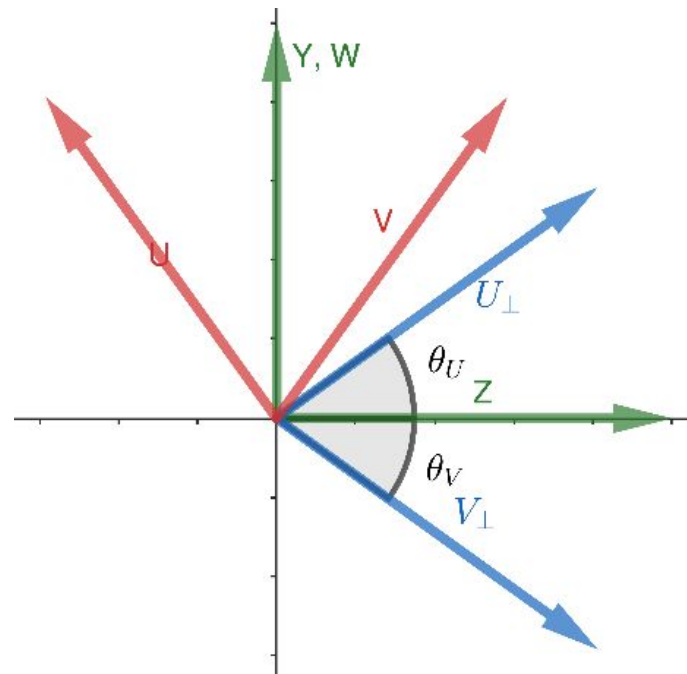
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# Using the recob::Hit objects

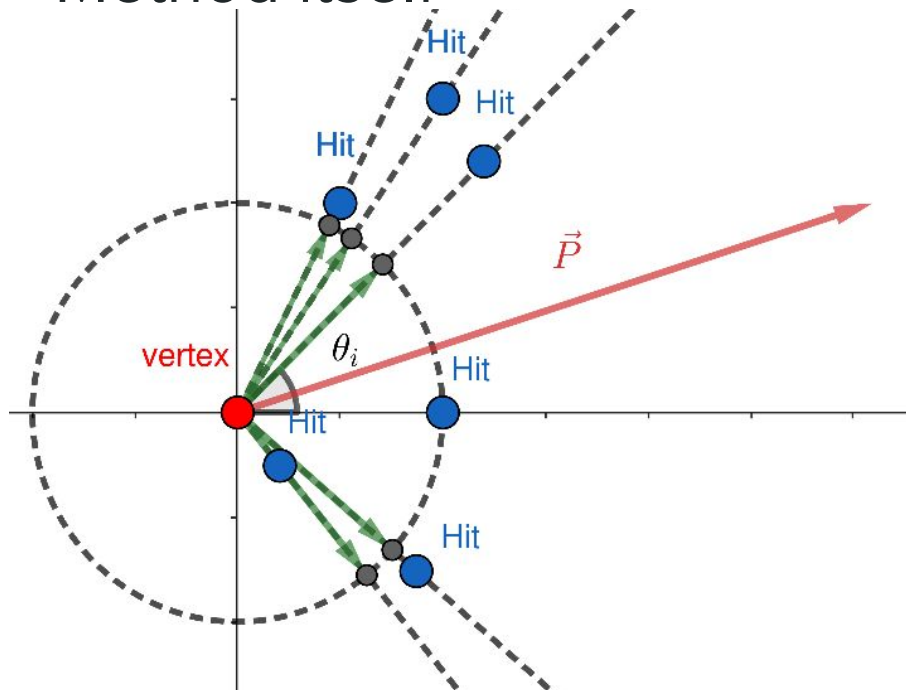
- All the recob::Hit objects are sorted in the 3 views (U, V, W) and put back in the view coordinate system (with  $V^- = U^+$  and  $U^- = V^+$ )

$$\begin{pmatrix} e_{\vec{U}_\perp} \\ e_{\vec{V}_\perp} \end{pmatrix} = \begin{pmatrix} -\sin \theta_U & \cos \theta_U \\ -\sin \theta_V & \cos \theta_V \end{pmatrix} \begin{pmatrix} e_y \\ e_z \end{pmatrix}$$

- The coordinates are shifted so that the vertex lies at (0, 0) in all the views.



## Method itself



$$\vec{P} = \sum_i \left( w_i \frac{x_i \vec{e}_x + y_i \vec{e}_y}{\sqrt{x_i^2 + y_i^2}} \right) = \sum_i w_i e^{j\theta_i}$$

'y' here is the view direction, not the real 'y'

Hits  $(x_i, y_i, w_i)$ :

- $x_i$  is the time position
- $y_i$  is the view we consider
- $w_i$  is the number of ADCs of the hit

Method:

- We get the direction of all the hits with respect to the vertex. We sum all of them weighted by the number of ADCs.
- Should give the average KE vector in this view



## Combining the 2D infos back to 3D

- We measure  $\vec{P}_{\text{view}} = P_{\text{view}} e^{i\theta_{\text{view}}}$  for each view.
- We only use the  $\theta_{\text{view}}$  information, which should be the only relevant for the direction (and allows to not care about views calibration)
  
- We have an excess of measurements with respect to what we want to reconstruct. We make a fit in spherical coordinates to use all of them:

$$\chi^2 = \sum_{\text{view}} \left( \theta_{\text{view}}^{\text{obs}} - \theta_{\text{view}}^{\text{fit}} \right)^2$$

$$\vec{P}_{\text{fit}} = \begin{pmatrix} \sin \theta \cos \phi \\ \sin \theta \sin \phi \\ \cos \theta \end{pmatrix}$$

e.g.  $\theta_W^{\text{fit}} = \arctan 2(\cos \theta, \sin \theta \cos \phi)$

# Calorimetric direction reconstruction: demo

