

# Report from SAND Calibration WG

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**for the SAND Calibration WG**

**SAND General Meeting**  
**January 14, 2025**

# Outline

1. Updates on the studies for ECAL Calibration (R.D'Amico – P.Gauzzi)
2. Grain Calibration with muons (A.Surdo)
3. Conclusions

# ECAL calibration in SAND

Calibration constants cell by cell determined with cosmic muons and muons from beam

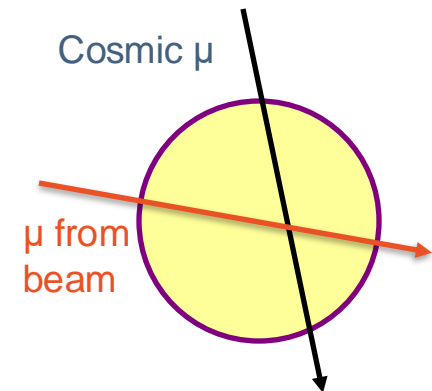
MIPs from cosmic rays: muon flux at surface  $\sim 0.02 \mu/(s \text{ cm}^2)$

$\Rightarrow \sim 10^4 \mu/s$  on ECAL ( $\Rightarrow 100 \text{ Hz}$  of “golden mips” in KLOE)

- Underground reduction of **a factor of about 100**  $\Rightarrow \sim 100 \mu/s$  on ECAL (no selection)
- Rough estimate by rescaling the KLOE numbers  $\Rightarrow 1 \text{ day (24 hrs): } \sim 10 \text{ evts/cell}$
- Relaxing the “golden mip” selection: in few days  $\sim 10^3 \text{ evts/cell}$

MIPs from beam (rock, magnet and Fe yoke,  
upstream ECAL modules)

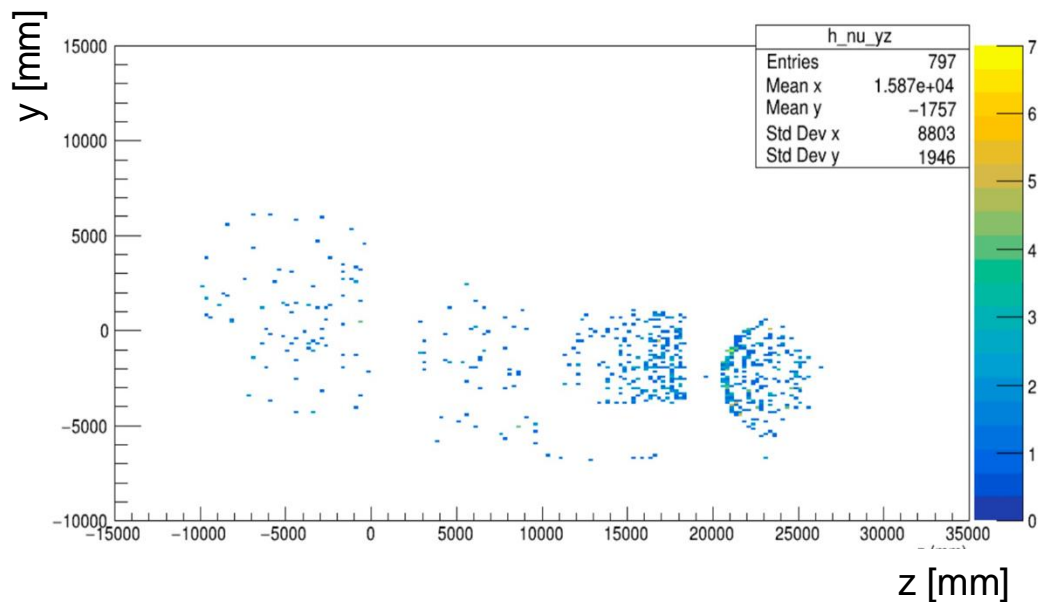
- We need also muons from beam for the modules around the median plane and for the endcaps



# MIPs from beam

- Generation of 25000  $\nu_\mu$  events in the generation window = DUNE\_ND\_HALL (X and Y in  $\sim -6.0 - 6.0$  m) and to cut at  $Z > -10$  m  $\Rightarrow$  797 events with at least 1 cluster from  $\mu$  in the ECAL

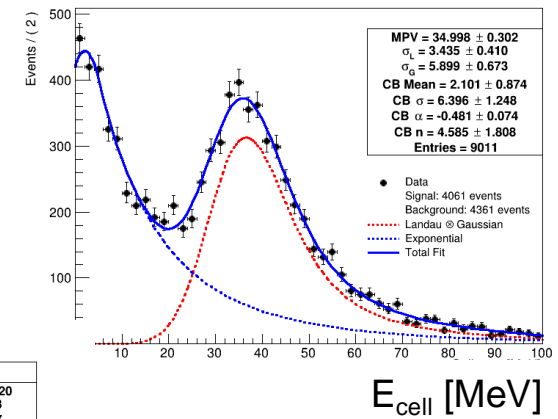
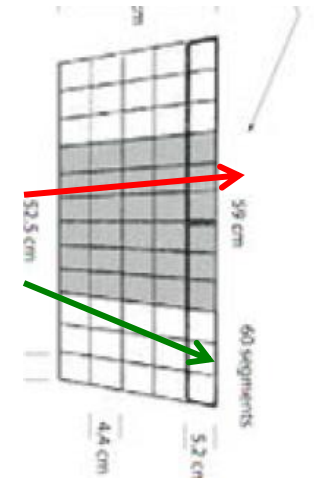
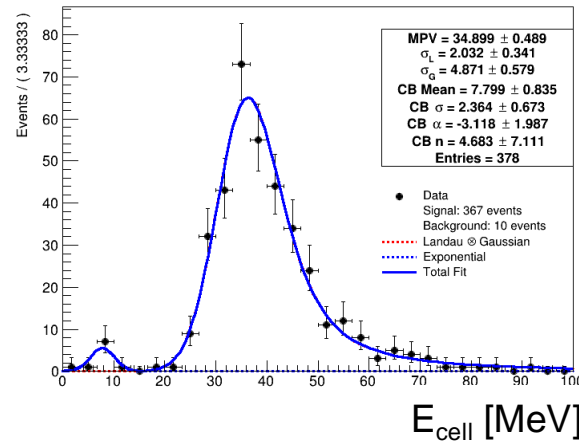
Vertices	
Rock	104
Fe Yoke	224
ECAL upstream modules	86
TMS	278
Cryostat/Solenoid	28
Others	57



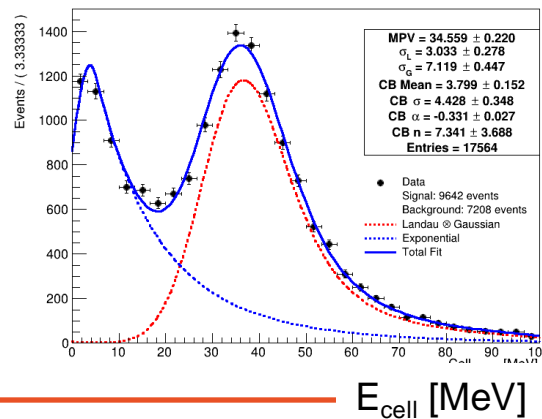
- $\sim 800$  evts in 30 spills means  $\sim 26$  muons/spill
- $2 \times 10^6$  good muons in 24 hours of beam

# MIPs from beam

- **Golden mips:** all the cluster cells in the same column
  - Low statistics
  - Clean distribution
  - Good peak fit
- 
- **Less stringent selection:** at least 3 cells in the same column
  - Peak still clear

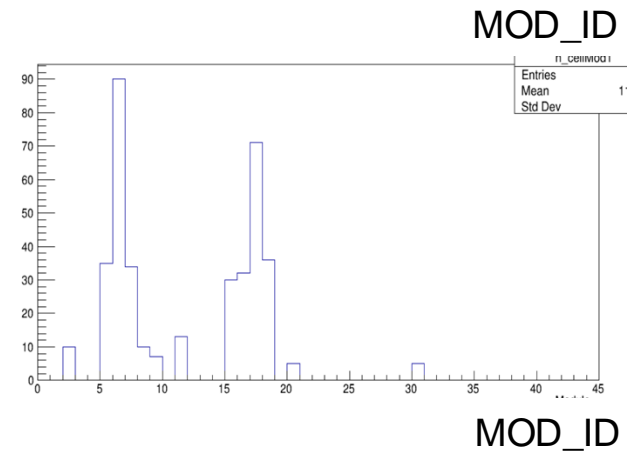
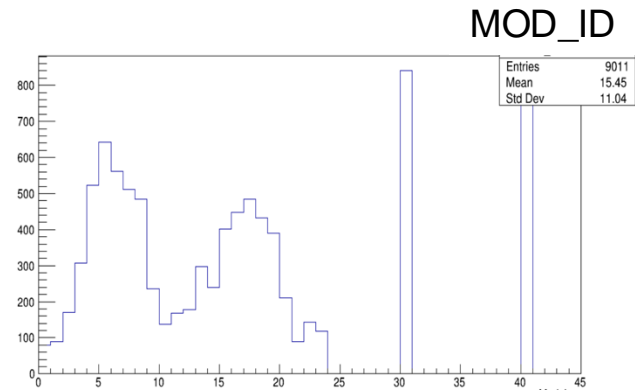
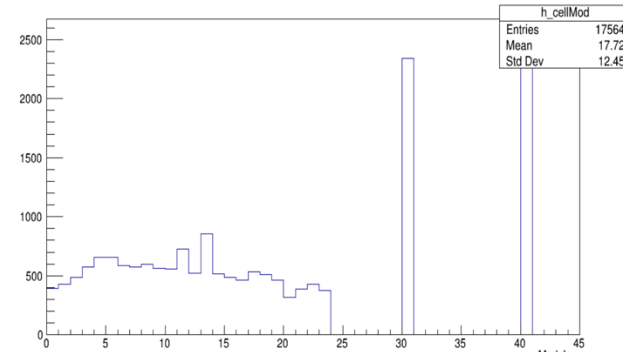


- No conditions on muon clusters



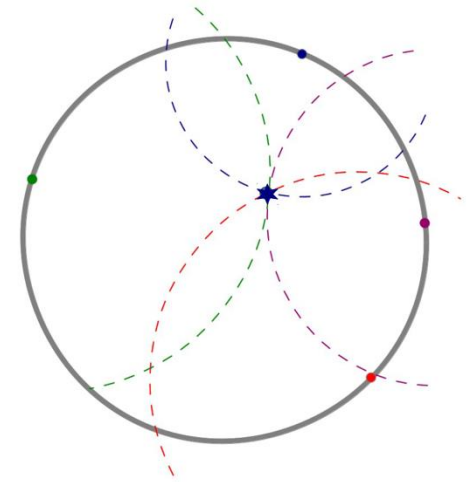
# MIPs from beam

- Occupancy:
  - No conditions on muon clusters
  - At least 3 cells in one column
  - Golden mip selection



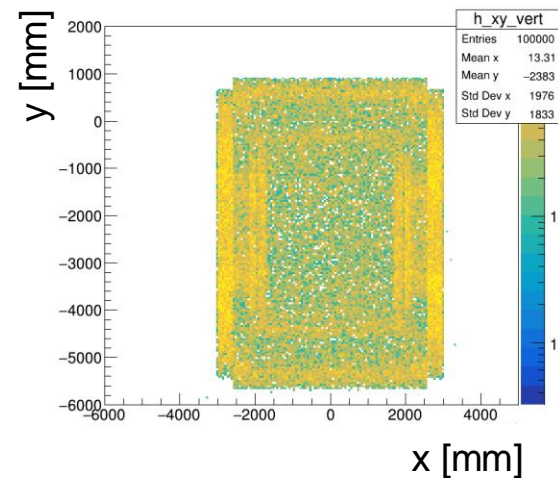
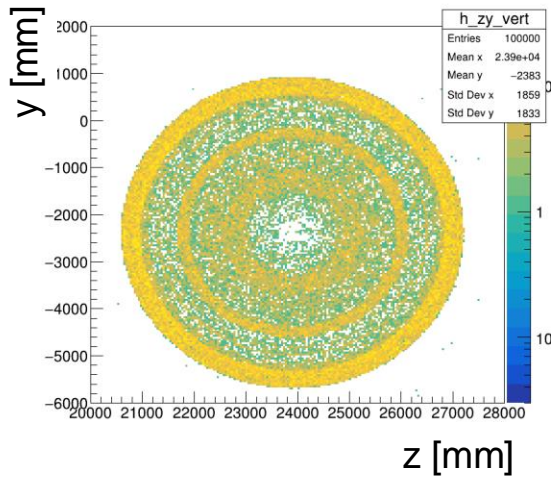
# Energy scale calibration in SAND

- $\gamma$ 's from  $\pi^0$  decays: invariant mass reconstruction (need a vertex from the tracker)
- $\gamma$  + electrons:  $\sim 30\%$  of photons from  $\pi^0$  convert in the tracker  
 $\Rightarrow \sim 50\%$  of  $\pi^0$  have at least one  $\gamma \rightarrow e^+e^-$  (from DUNE-doc-13262 A Near Detector for DUNE)
- High energy electrons from  $\nu_e$  interactions  $\Rightarrow$  need the momentum measurement in the tracker
- Possibility to exploit  $K^0 \rightarrow \pi^0 \pi^0 \rightarrow 4\gamma$
- From a naive rescaling of  $K^0 \rightarrow \pi^+ \pi^- \Rightarrow O(10^5)$  evts in 5 years of FHC data-taking
- Reconstruct a vertex with the ECAL only, back-propagating each of the 4 photons, but the times of the ECAL cells must be very well aligned

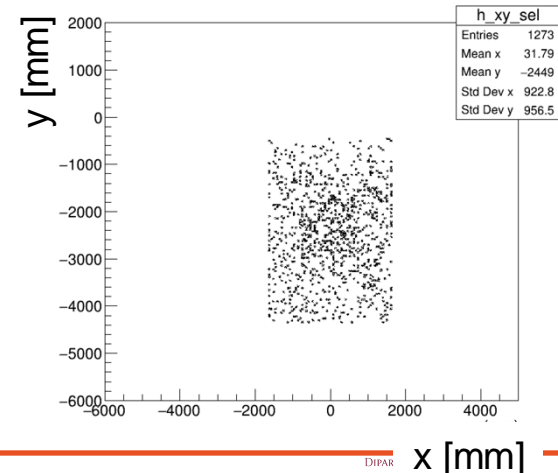
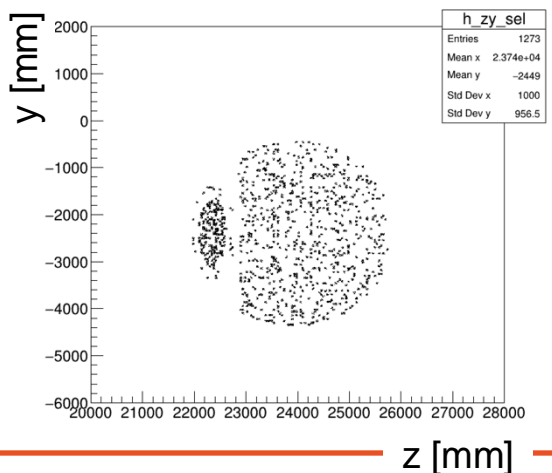


# MC sample

- Generated 100000  $\nu_\mu$  events with vertices in in the SAND volume (TOP\_VOLUME = volSAND), POTs  $\sim 10^{17} \Rightarrow \sim 30$  min of beam



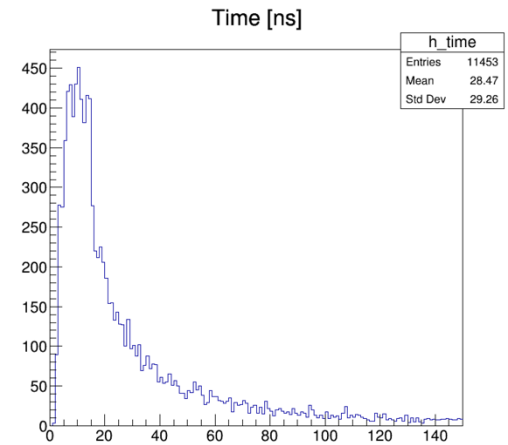
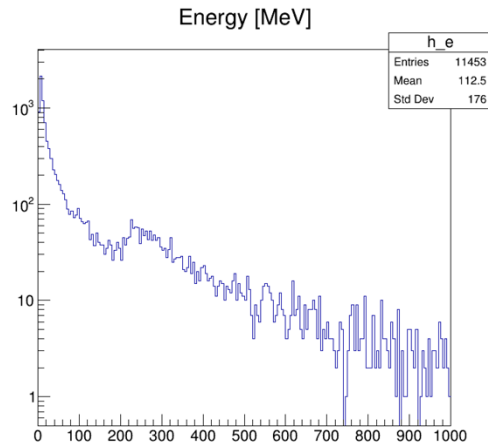
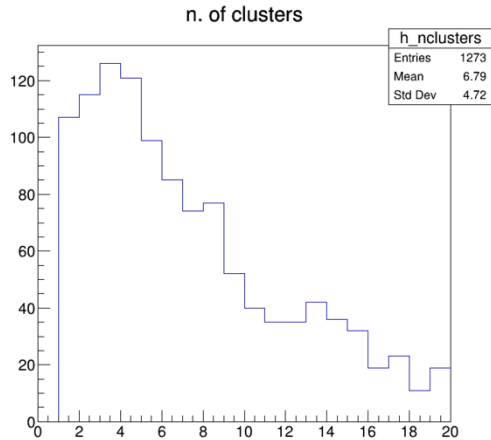
- Select vertices inside the ECAL  $\Rightarrow$  1273 events





# MC sample

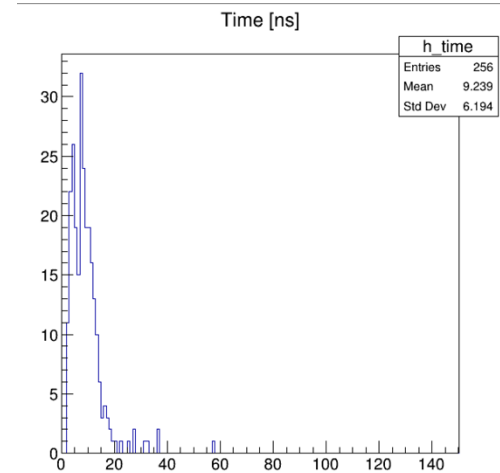
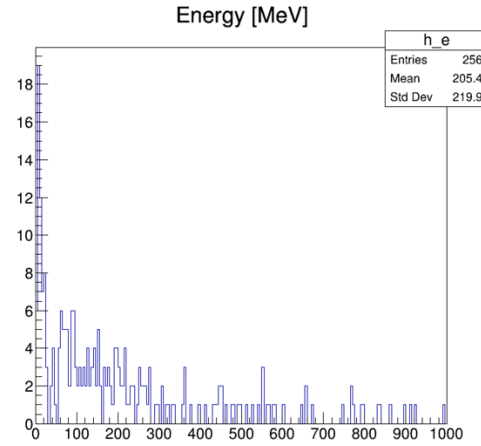
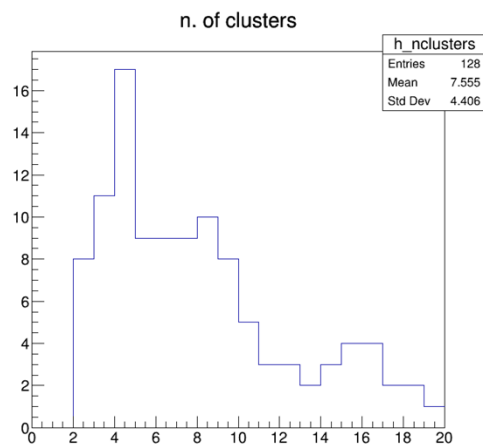
- Clusters in the ECAL



Cluster particle (main contribution)								
$e^+/e^-$	$\mu^+/\mu^-$	$\gamma$	$K_L$	$\pi^+/\pi^-$	$K^+/K^-$	n	p	$\Lambda$
360	1260	2105	56	3033	128	3016	1491	4

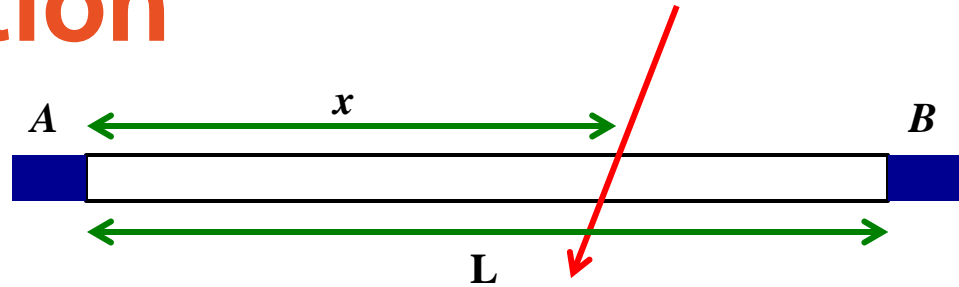
# Photons from $\pi^0$ 's

- Look at the parent of the cluster particle: select events with 2 photons from a  $\pi^0$  decay  $\Rightarrow$  128 events (256 clusters)



- Energy not well calibrated because I used an old version of SANDRECO
- Work in progress (waiting for the new version)

# Time calibration



$$t = \frac{1}{2}(t_A + t_B) - \frac{L}{2v} - t_0 - t_G^0$$

$$x = \frac{1}{2}v(t_A - t_B) - \Delta t_0$$

$$t_0 = \frac{1}{2}(t_A^0 + t_B^0)$$

$$\Delta t_0 = \frac{1}{2}(t_A^0 - t_B^0)$$

- Alignment of times in the ECAL requires to determine the  $t_0$ 's cell by cell
- Select straight tracks ( $p > 6$  GeV) with 2 clusters, connecting as much as possible different regions of the ECAL
- Also in this case we could use beam muons together with cosmics
- Global offset  $\Rightarrow t_G^0$  to be determined

# Time calibration

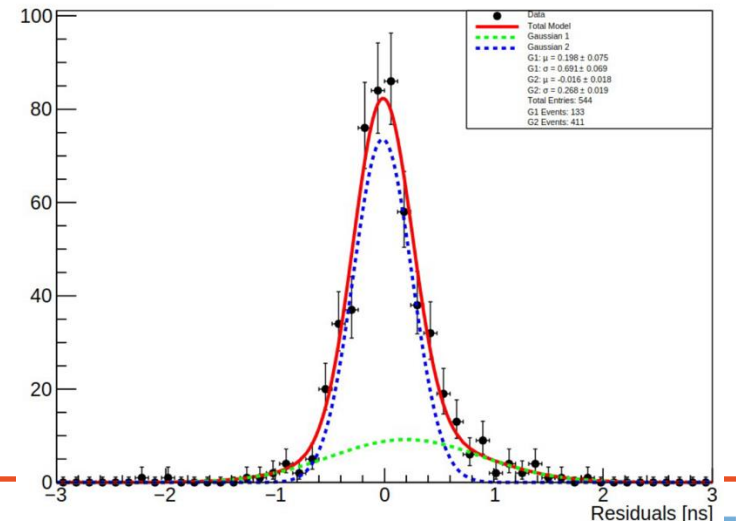
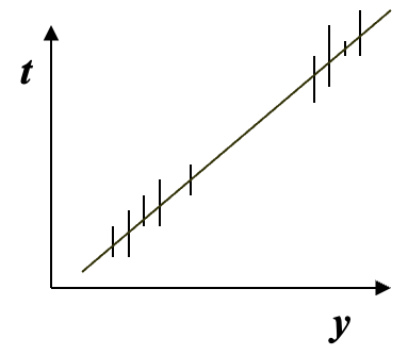
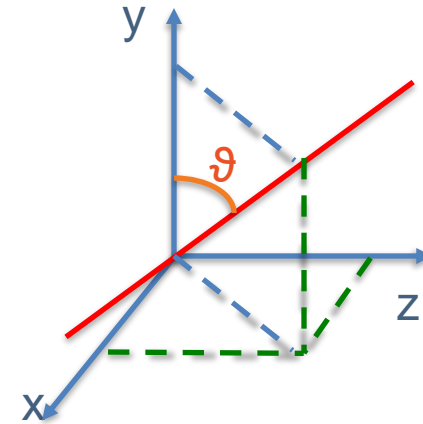
- 3D linear fit of the straight track to get the  $\vartheta$  angle
- Linear fit:  $t$  vs  $y$  (or  $z$ ), at least 5 + 5 points / track

$$t = T_0 + \frac{y}{c \cos \theta}$$

$$T_0 = \frac{\sum_i (t_i - t_i^0 - \frac{y_i}{c \cos \vartheta}) E_i}{\sum_i E_i} \quad (\text{Energy-weighted average})$$

- Histograms of the residuals (one histo. per cell)

$$\Delta y_i = t_i - t_i^0 - T_0 - \frac{y_i}{c \cos \vartheta}$$



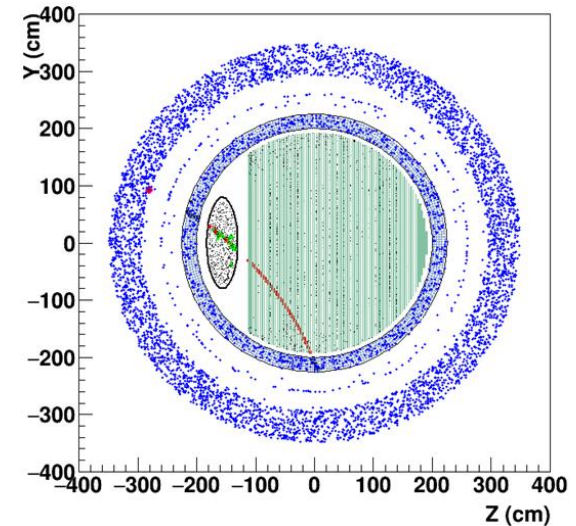
- The center of the distribution is the correction to the  $t_i^0$
- Iterate the procedure:** re-run the ECAL reconstruction and clustering with the  $t_i^0$ 's updated and go to step 1.

# Next steps (ECAL)

- Generate few  $\times 10^6$  events for more statistics of muons from beam (waiting for the implementation of the last version of the ECAL Digitization which includes the real Endcap geometry)
- Continue the study of  $\gamma$ 's from  $\pi^0$  decays for the absolute energy scale
- Start the discussion on ideas for the global  $t_0$  determination
- Other items:
  - Generate events from beam flux
  - Study cosmic muons with MC

# GRAIN calibration with muons

Muon from  $\nu$  interaction in the yoke and crossing GRAIN



✓ Most obvious process to be considered:

## MIPs crossing the LAr volume

- muons from the beam interaction outside GRAIN
- cosmic ray muons

✓ Specific energy loss for a generic material:  $\langle dE/dx \rangle \sim 2 \text{ MeV}/(\text{g}\cdot\text{cm}^{-2})$

Can be estimated from MC simulation or measured from experimental data.

For LAr:

$dE/dL \sim 2.5 \text{ MeV}/\text{cm} \Rightarrow N_0 \sim 10^5 \text{ ph}/\text{cm}$  Photon emission per unitary pathlength  
(assuming  $f \sim 4 \cdot 10^4 \text{ ph}/\text{MeV}$ )

➤ The relation between muon Pathlength and Energy loss exploited to get knowledge of energy deposit in LAr, to be related to the amount of detected photons

# Muon beams to test GRAIN calibration

## Simulation of different muon beams crossing GRAIN with:

- Monochromatic muons (1 GeV)
- Different impact points on GRAIN surface
- Different beam directions and path-lengths inside the LAr

## .. in order to study the dependence on:

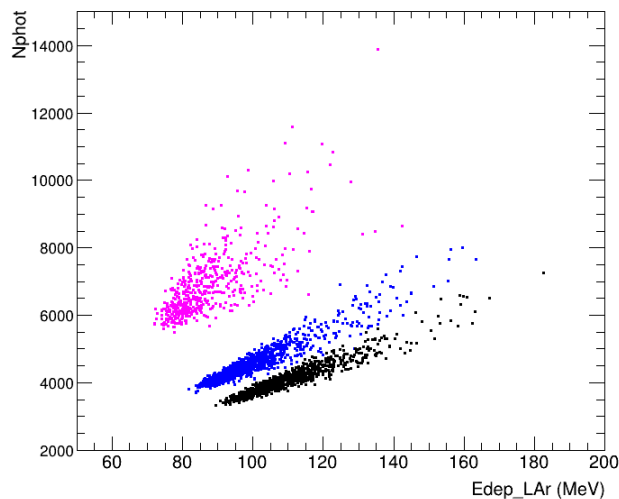
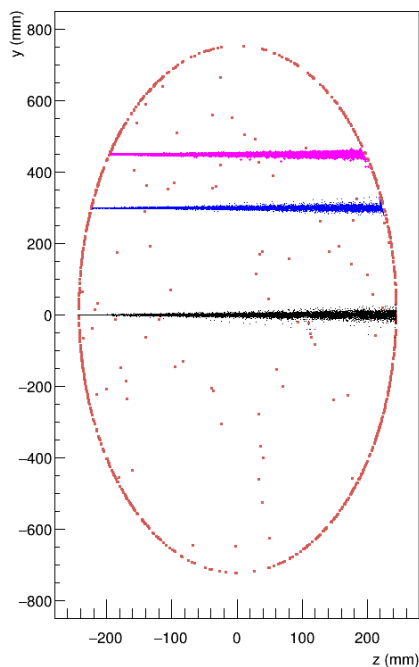
- track location and distances from the cameras (geometric acceptance)
- path-length inside GRAIN (energy deposit) for different track orientations
- ...

Scintillation light photons propagated in LAr and collected by the photo-sensor system through *OptMen* code

Simulation of the **Lens-camera setup** with proper SiPM-PDE and Electronics

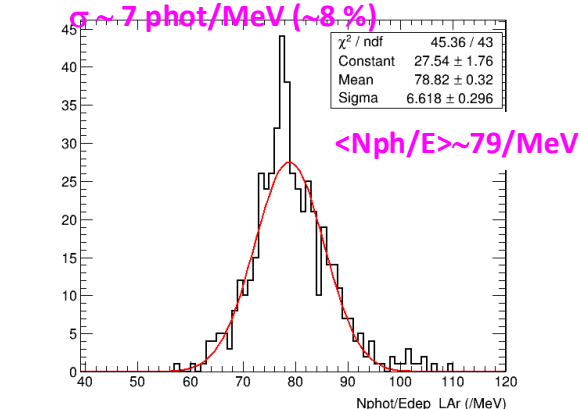
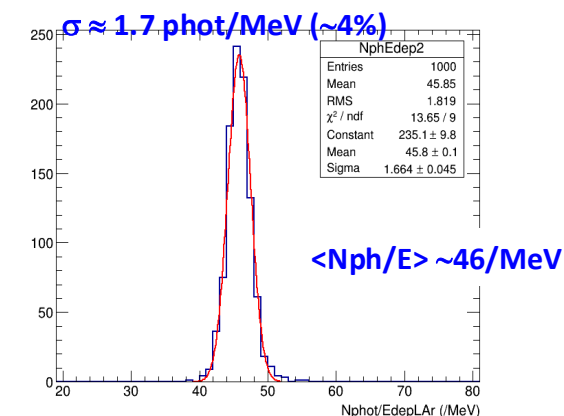
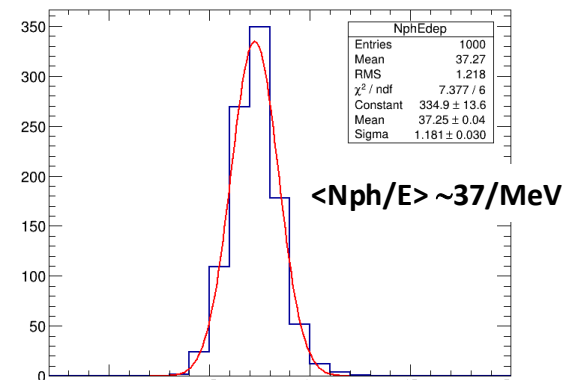
# Simulation of Muon beams

- 1) Beam of 1 GeV muons **along Z axis** ( $X=Y = 0$ )
- 2) Beam of 1 GeV muons parallel to Z axis, at  **$Y = 30$  cm**
- 3) Beam of 1 GeV muons parallel to Z axis, at  **$Y = 45$  cm**



**Large spread for Beam\_3** in the correlation btw EdepLAr and Nphot!

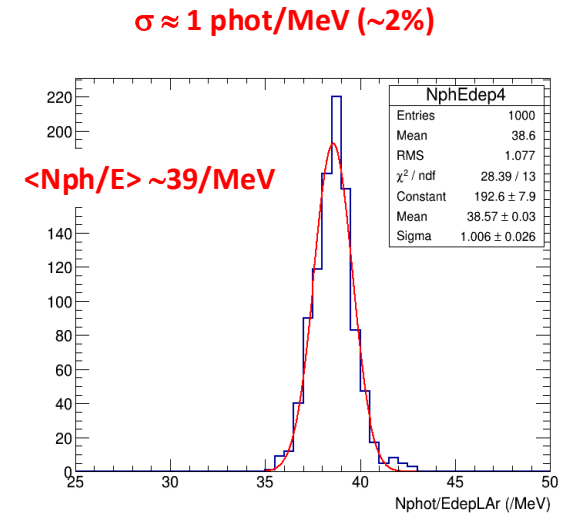
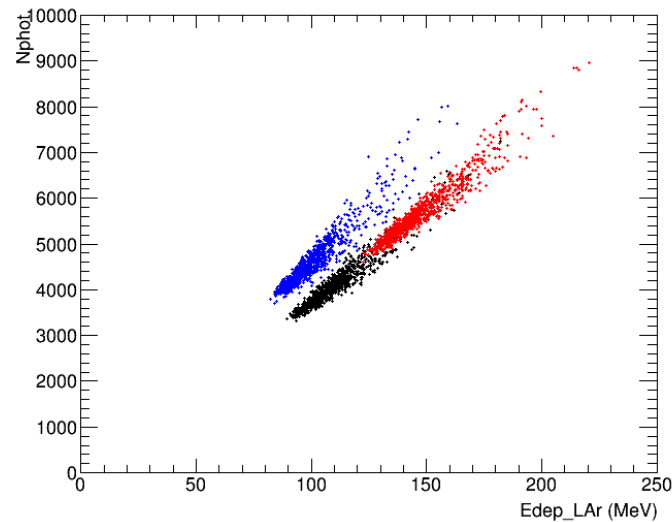
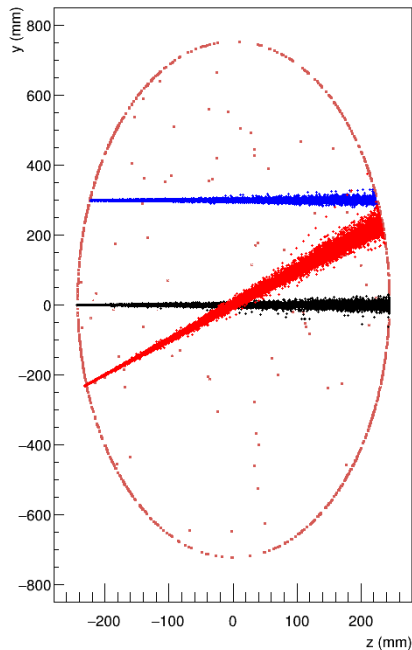
... due to very near cameras?





# Simulation of Muon beams

## 4) Beam of 1 GeV inclined muons crossing the center



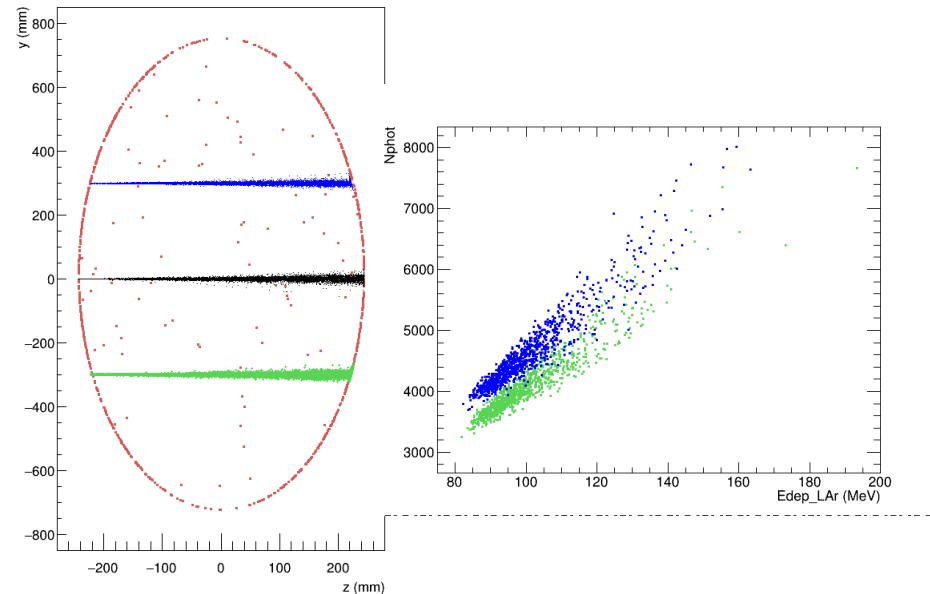
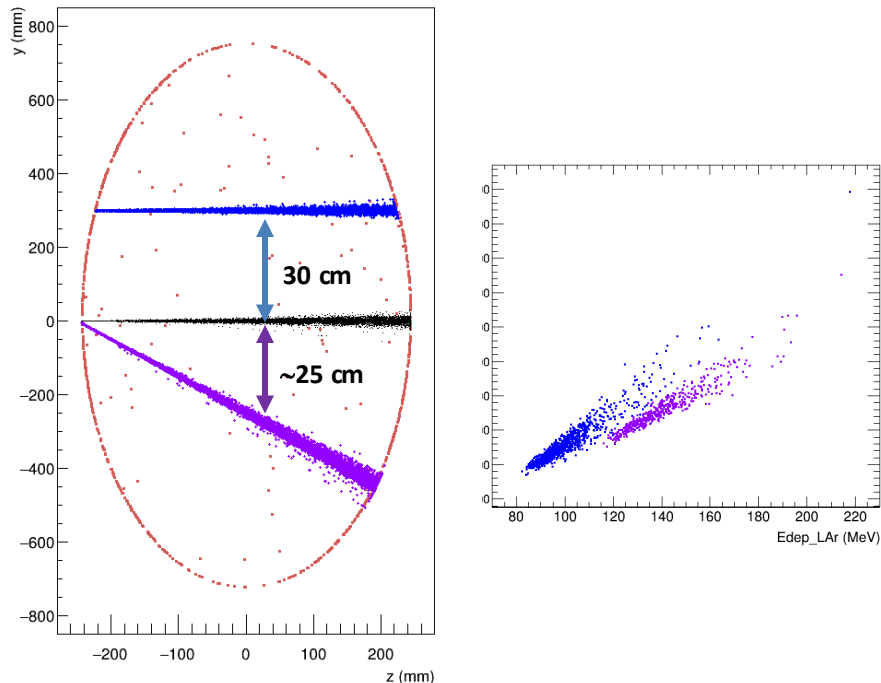
It is remarkable that the beams through GRAIN center (black and red) give aligned correlations one each other

⇒ The distance from GRAIN center could be a parameter which affects calibration curves

# Simulation of Muon beams

5) Beam of 1 GeV inclined muons, distant from the center

6) Beam of 1 GeV horizontal muons, ad  $Y = -30\text{cm}$  (to be compared with Beam\_2)



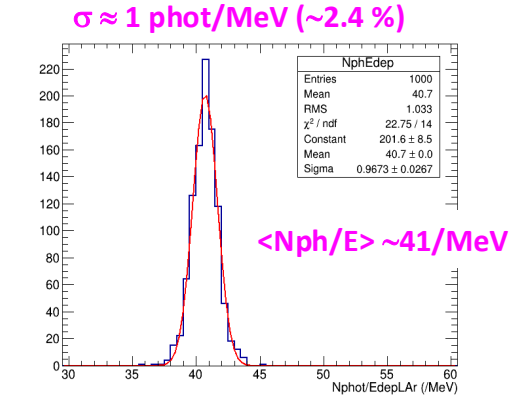
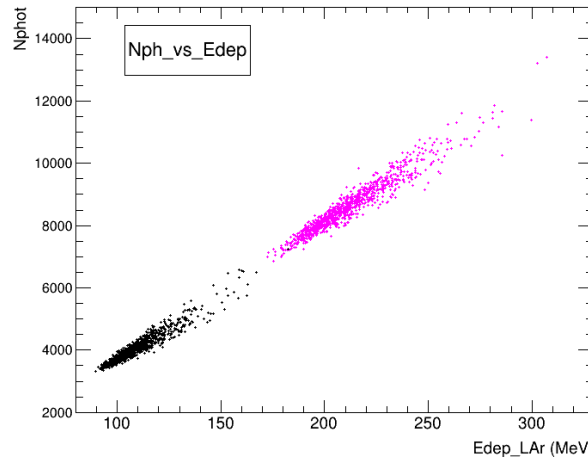
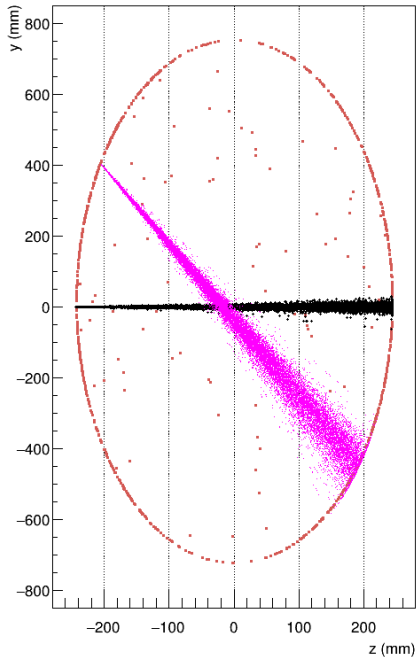
**Beams 1 and 5 (blue and violet) give not perfect aligned correlations one each other (distance from center not equal and camera-layouts not asymmetric ..)**

**Difference due to the asymmetry btw Top and Bottom camera layouts (14 on the Top, 7 on the Bottom)**

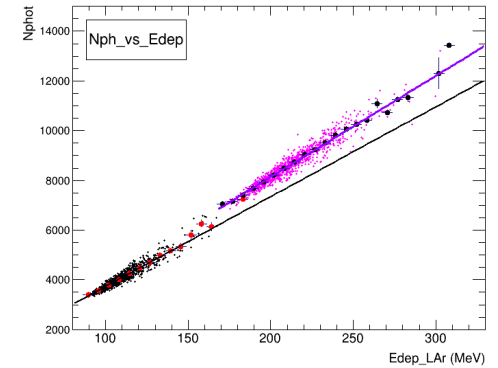
# Simulation of Muon beams

## 7) Beam of 1 GeV very inclined muons, crossing GRAIN center

⇒ comparison with Beam 1



Again, the beams through GRAIN center give aligned correlations one each other (despite very different directions)



## Conclusions (Simulation of Muon beams in GRAIN)

- ✓ The test of the procedure with simulated muon beams shows that the calibration method with muons could work, in principle
- ✓ For muon pathlengths inside GRAIN not far from the center, a tight correlation (i.e. calibration curve) between EdepLAr and Nphot is obtained
- ✓ Larger spreads observed for tracks very near to the cameras ..
- ✓ It is remarkable that the calibration curves seem aligned for the same distance from GRAIN center (→ a parameter to be used)
- ✓ Observed the effect of the asymmetry Top/Bottom in the Lens-camera system

# Conclusions

- **ECAL:**
  - Studies of Energy and Time calibration with muons in progress
  - Studies of Energy scale calibration with  $\gamma$ 's from  $\pi^0$  decays started
  - Next step: strategy for  $t_0$  global determination
- **GRAIN:**
  - Studies on calibration with muons in progress
  - In particular the dependence of the response on track location and path length inside GRAIN and on the distance from the cameras is studied
- **Other items to be addressed in the future:**
  - Calibration of the inner tracker
  - Intercalibration among subdetectors (timing)
  - Organize the software for calibration and define a place for calibration constants

# Spares

# SAND Calibration WG

- Calibration: from detector signals to physical variables
  - ECAL: energy, time and positions of the particles
  - GRAIN: tracks, time, energy, ....
  - Tracker : r-t relations, track momentum,  $dE/dx$  for PID, ....
  - Timing alignment among the subdetectors
- Define a strategy for each subdetector:
  - Sources: cosmics, particles from beam, ...
  - Choose suitable processes (given the expected fluxes of particles in the detector, e.g. for the ECAL: cosmic  $\mu$ 's as MIPs, MIPs from the beam, electrons and photons ....)
- Set a calibration procedure (Which level of precision ? How much time expected ?)
- Reference people: ECAL - P.Gauzzi, GRAIN: A.Surdo, Tracker: .....
- Next meeting: Thursday, January 16, at 3:30 p.m. CET (8:30 a.m. CT)
- WG mailing list: [dune-nd-sand-calibration@fnal.gov](mailto:dune-nd-sand-calibration@fnal.gov)

# ECAL calibration

MIPs from beam (rock, magnet and Fe yoke, upstream ECAL modules)

- $\sim 1.5 \times 10^3 \mu/\text{spill}$  (1 spill = 9.6  $\mu\text{s}$  every 1.2 s) without any selection

Cut	ECAL		Rock muons		Magnet events	
	Events	$\varepsilon$ (%)	Events	$\varepsilon$ (%)	Events	$\varepsilon$ (%)
No cut	2.23	100.0	1447.26	100.000	50.82	100.000
$\mu$ in ECAL FV	2.23	100.0	12.73	0.880	18.92	37.229
STT & ECAL hits	1.63	72.9	6.05	0.420	3.443	6.775
NN cut	1.56	95.5	0.10	0.007	0.07	0.136

Table 40: Number of events per spill (9.6  $\mu\text{s}$ ,  $7.5 \times 10^{13}$  pot) and selection efficiency for the signal from  $\nu_\mu$  CC in the front barrel ECAL and the backgrounds from rock muons and magnet events.

(from DUNE-doc-13262, A Near Detector for DUNE)

- By requiring hits in the STT and ECAL  $\Rightarrow \sim 11$  muons/spill



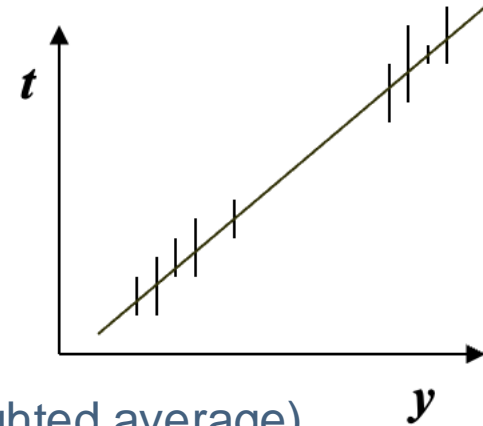
# Time calibration

2. Linear fit:  $t$  vs  $y$  ( at least 5 + 5 points / track)

$$t = T_0 + \frac{y}{c \cos \theta}$$

$$T_0 = \frac{\sum_i (t_i - t_i^0 - \frac{y_i}{c \cos \vartheta}) E_i}{\sum_i E_i}$$

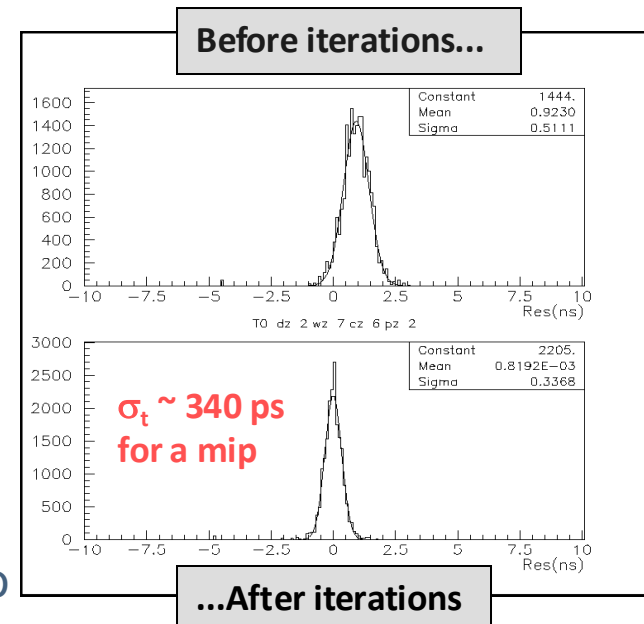
(Energy-weighted average)



3. Histograms of the residuals (one per cell)

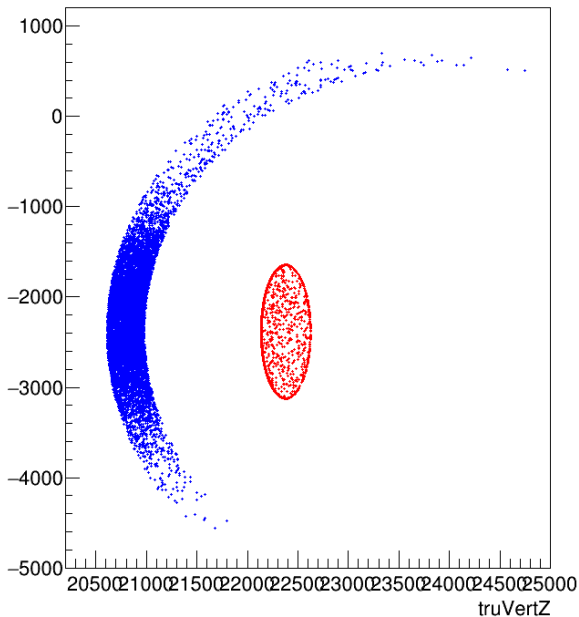
$$\Delta y_i = t_i - t_i^0 - T_0 - \frac{y_i}{c \cos \vartheta}$$

- The center of the distribution is the correction to the  $t_i^0$
- **Iterate the procedure:** re-run the ECAL reconstruction and clustering with the  $t_i^0$ 's updated and go to step 1.
- Stop when the corrections are compatible with zero

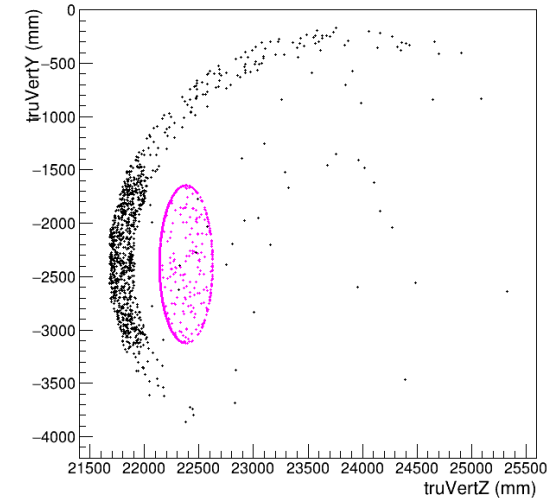
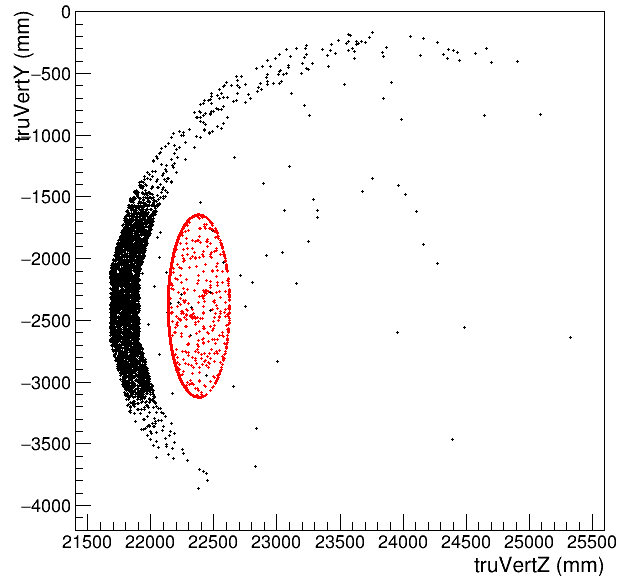


# Events with the muon entering GRAIN

## Vertex in the Magnet yoke



## Vertex in ECAL



Total interaction events: 200,000

Muons entering GRAIN:  $\sim 6,000$  (3%)

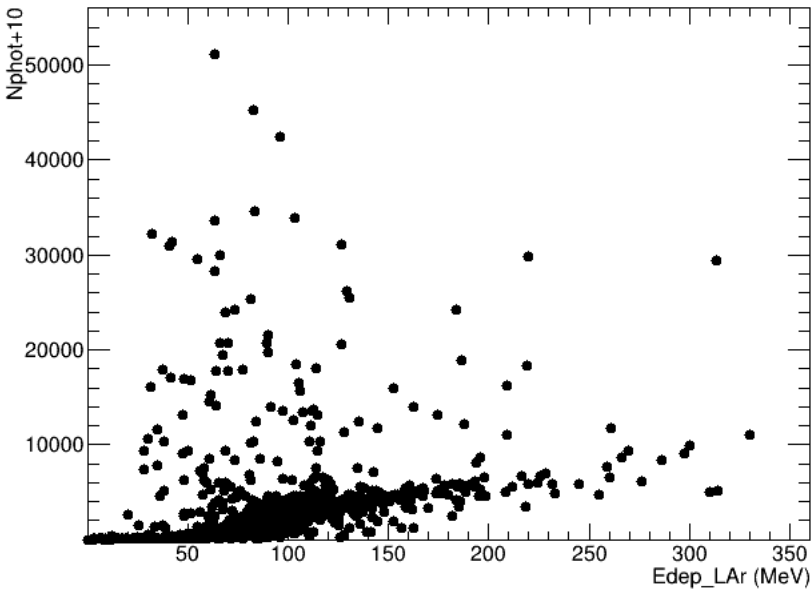
↳ Clean muons:  $\sim 1,500$  (0.8%)

Total interaction events: 441,000

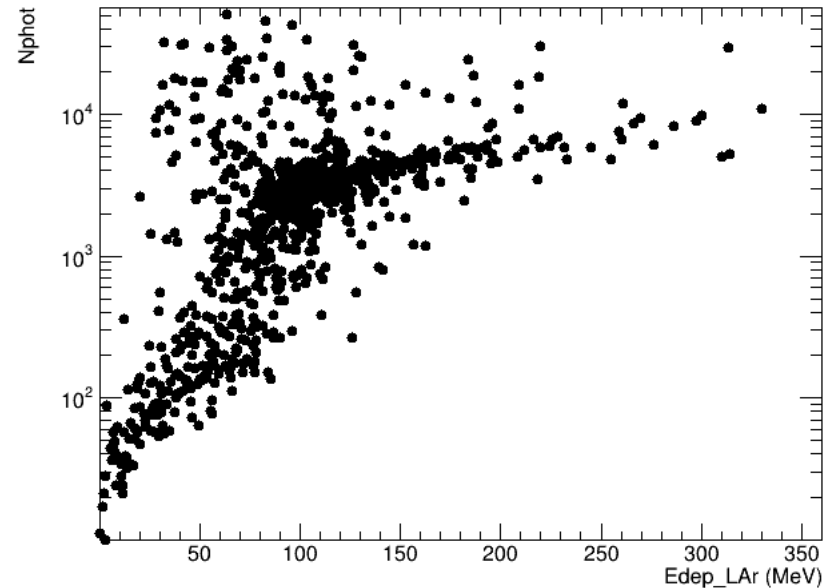
Muons entering GRAIN:  $\sim 13,000$  (3%)

↳ Clean muons:  $\sim 10,000$  (2.3%)

# Correlation btw detected photons and deposited energy



Log scale for  $N_{\text{phot}}$



- Not a so narrow correlation
- Possible effects from track position vs geometrical acceptance

Apparently, different behaviours ?

# Expected muon flux from the beam and CRs

- ✓ Different contributions of the target masses in SAND for beam neutrinos

(from DUNE-doc-13262, A Near Detector for DUNE)

Table 1.29: Total number of ( $\nu_\mu + \bar{\nu}_\mu + \nu_e + \bar{\nu}_e$ ) CC+NC events expected within a single beam ( $9.6 \mu\text{s}$ ,  $7.5 \times 10^{13}$  POT) in the various detector components for both the FHC and RHC beam

Detector element	Mass	FHC	RHC
Magnet	511 t	68.9	36.6
ECAL	100 t	13.5	7.2
LAr+STT	8.2 t	1.1	0.59
STT fiducial volume	5.5 t	0.74	0.39
Total	619.2	83.5	44.39

Table 1.34: Number of events per spill ( $9.6 \mu\text{s}$ ,  $7.5 \times 10^{13}$  POT) and selection efficiency for the signal from  $\nu_\mu$  CC in the front barrel ECAL and the backgrounds from rock muons and magnet events.

Cut	ECAL		Rock muons		Magnet events	
	Events	$\epsilon$ (%)	Events	$\epsilon$ (%)	Events	$\epsilon$ (%)
No cut	2.23	100.0	1447.26	100.000	50.82	100.000
$\mu$ in ECAL FV	2.23	100.0	12.73	0.880	18.92	37.229
STT & ECAL hits	1.63	72.9	6.05	0.420	3.443	6.775
NN cut	1.56	95.5	0.10	0.007	0.07	0.136

- Further contribution from rock  $\mu$ 's ( $\sim 1.7/\text{spill}$ ) ...

## ✓ Contribution from Cosmic Rays ...

CR Muon flux at surface  $\sim 0.01 \mu/(\text{s cm}^2)$  +  
 underground reduction of  $\sim 100$   
 Effective area of GRAIN for  $<60^\circ$  CR muons:  
 $\sim 3 \times 10^4 \text{ cm}^2 \Rightarrow \sim 3 \mu/\text{s}$  are expected to cross GRAIN

Drawback: smaller acceptance by the tracker for a precise track reconstruction

**Main contribution only if inter-spill DAQ were ON**

- From the interaction rate /spill in Magnet yoke and ECAL, a quite low number of clean muons are expected to cross GRAIN per spill ( $\leq 1 \mu/\text{spill}$ )