

Reliability of Predictions for the SSA in Drell-Yan

John Collins (Penn State)

- Issues in theory of TMD factorization and the Sivers sign-reversal
- Accuracy of phenomenology and quantitative predictions

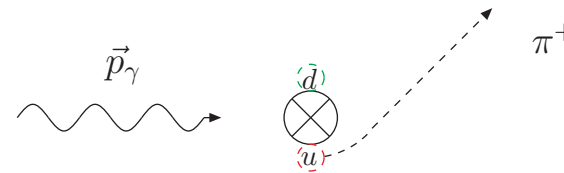
Especially about dilution of Sivers asymmetry by evolution to higher Q

However, you should review history of SSA, e.g., in $pp \rightarrow \pi X$ before evaluating theoretical predictions.

Justification of TMD factorization ($q_T \ll Q$), Sivers sign-reversal

- Full proof with details about Wilson lines: JCC's "Foundations of Perturbative QCD". [Check!](#)
- Verification in low-order graphs for SIDIS and DY
[Brodsky, Hwang, Schmidt, PL B530, 99 (2002); Brodsky, Hwang, Schmidt, NP B642, 344 (2002); Collins, Qiu, PRD 75, 114014 (2007; Brodsky et al., arXiv:1304.5237.)]

- Burkardt model gives sign reverse:

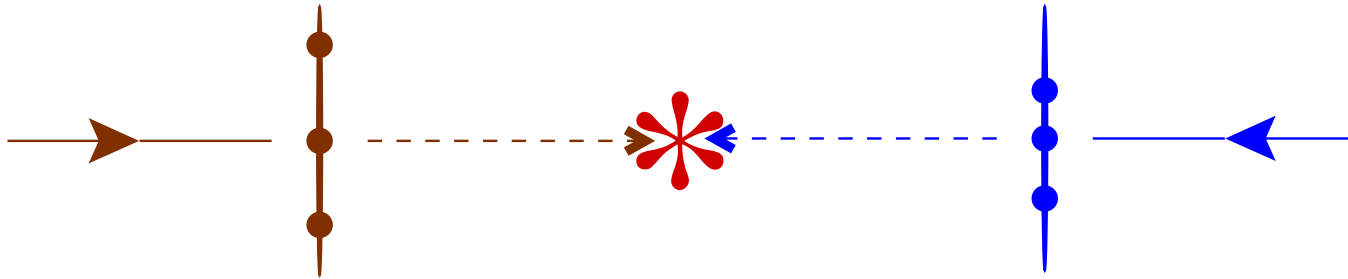


Force on outgoing quark v. incoming antiquark in color field in spinning proton.
E.g., Burkardt, arXiv:1009.5442

- Lattice QCD verifies Sivers sign-reversal in suitable correlation function.

→ N.B. Better understanding of size of effects of power corrections is needed

TMD factorization for Drell-Yan with $q_T \ll Q$



$$\frac{d\sigma}{d^4q d\Omega} = \frac{2}{s} \sum_j \frac{d\hat{\sigma}_{j\bar{j}}(Q, \mu, g(\mu))}{d\Omega} \int d^2\mathbf{b}_T e^{i\mathbf{q}_{hT} \cdot \mathbf{b}_T} \tilde{f}_{j/A}(x_A, \mathbf{b}_T; \zeta_A, \mu) \tilde{f}_{\bar{j}/B}(x_B, \mathbf{b}_T; \zeta_B, \mu)$$

+ poln. terms + high- q_T term + power-suppressed

with: $\zeta \simeq (2 \times \text{parton energy})^2, \quad \zeta_A \zeta_B = Q^4.$

TMD factorization, and evolution

$$\frac{d\sigma}{d^4q d\Omega} = \frac{2}{s} \sum_j \frac{d\hat{\sigma}_{j\bar{j}}(Q, \mu, g(\mu))}{d\Omega} \int d^2\mathbf{b}_T e^{i\mathbf{q}_{hT} \cdot \mathbf{b}_T} \tilde{f}_{j/A}(x_A, \mathbf{b}_T; \zeta_A, \mu) \tilde{f}_{\bar{j}/B}(x_B, \mathbf{b}_T; \zeta_B, \mu)$$

+ poln. terms + high- q_T term + power-suppressed

CSS evolution:

$$\frac{\partial \ln \tilde{f}_{f/H}(x, b_T; \zeta; \mu)}{\partial \ln \sqrt{\zeta}} = \tilde{K}(b_T; \mu) \quad (\text{with } \zeta = Q^2)$$

- Evolve to remove logarithms in perturbative regions “resummation”.
- Parameterize intrinsically non-perturbative part:
 - Large b_T in TMD pdfs;
 - Large b_T in CSS kernel $\tilde{K}(b_T)$, or corresponding function in other formalisms.
- Non-perturbative region in b_T . Strikman & Weiss [JHEP 01 (2013) 163] argue for two non-perturbative scales:
 $1 \text{ fm} = 5 \text{ GeV}^{-1}$ (confinement); $0.3 \text{ fm} = 1.5 \text{ GeV}^{-1}$ (chiral condensate).
- Also use relation of TMD pdfs to integrated pdfs at small- b_T .

One solution: Factorization with fixed TMD pdfs

$$\begin{aligned}
 \frac{d\sigma}{d^4q d\Omega} &= \frac{2}{s} \sum_j \frac{d\hat{\sigma}_{j\bar{j}}(Q, \mu_Q, g(\mu_Q))}{d\Omega} \int d^2\mathbf{b}_T e^{i\mathbf{q}_{h_T} \cdot \mathbf{b}_T} \times \\
 &\times \tilde{f}_{f/A}(x_A, \mathbf{b}_T; m^2, \mu_0) \tilde{f}_{\bar{f}/B}(x_B, \mathbf{b}_T; m^2, \mu_0) \\
 &\times \left(\frac{Q^2}{m^2} \right)^{\tilde{K}(b_T; \mu_0)} \times \exp \left\{ \int_{\mu_0}^{\mu_Q} \frac{d\mu'}{\mu'} \left[2\gamma(g(\mu'); 1) - \ln \frac{Q^2}{(\mu')^2} \gamma_K(g(\mu')) \right] \right\} \\
 &+ \text{polarized terms} + \text{large } q_{h_T} \text{ correction, } Y + \text{p.s.c.}
 \end{aligned}$$

- where:
- $\mu_Q \propto Q$
 - blue \Leftrightarrow non-perturbative and/or non-“resummed” logarithms

N.B. Perturbative analysis/prediction of TMD pdfs at small b_T also used

What affects shape in q_T at $q_T \ll Q$?

We have q_T -independent factor times

$$\begin{aligned} & \int d^2 \mathbf{b}_T e^{i \mathbf{q}_{hT} \cdot \mathbf{b}_T} \tilde{f}_{j/A}(x_A, \mathbf{b}_T; m^2, \mu_0) \tilde{f}_{\bar{j}/B}(x_B, \mathbf{b}_T; m^2, \mu_0) \left(\frac{Q^2}{m^2} \right)^{\tilde{K}(b_T; \mu_0)} \\ &= \int d^2 \mathbf{b}_T e^{i \mathbf{q}_{hT} \cdot \mathbf{b}_T} \tilde{f}_{j/A}(x_A, \mathbf{b}_T) \tilde{f}_{\bar{j}/B}(x_B, \mathbf{b}_T) e^{\tilde{K}(b_T; \mu_0) \ln(Q^2/m^2)} \end{aligned}$$

RG to deal with logarithms, segregation of non-perturbative information:

$$\begin{aligned} \tilde{K}(b_T; \mu_0) &= \tilde{K}(b_T; \mu_b) + \int_{\mu_0}^{\mu_b} \frac{d\mu'}{\mu'} \gamma_K(g(\mu')) \\ &= \tilde{K}(b_*; \mu_b) + \int_{\mu_0}^{\mu_b} \frac{d\mu'}{\mu'} \gamma_K(g(\mu')) + \tilde{K}(b_T; \mu_b) - \tilde{K}(b_*; \mu_b) \\ &= \tilde{K}(b_*; \mu_b) + \int_{\mu_0}^{\mu_b} \frac{d\mu'}{\mu'} \gamma_K(g(\mu')) - g_K(b_T; b_{\max}) \end{aligned}$$

where (CSS):

- $\mathbf{b}_* = \mathbf{b}_T / \sqrt{1 + b_T^2/b_{\max}^2}$, $\mu_b = C_1/b_*$,

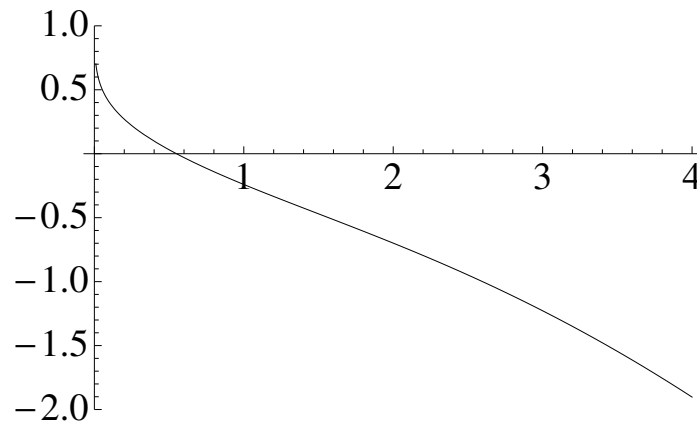
- red \Leftrightarrow (treated as) non-perturbative

Evolution shifts TMD to smaller b_T as Q increases

We have q_T -independent factor times

$$\begin{aligned} & \int d^2\mathbf{b}_T e^{i\mathbf{q}_{hT}\cdot\mathbf{b}_T} \tilde{f}_{j/A}(x_A, \mathbf{b}_T; m^2, \mu_0) \tilde{f}_{\bar{j}/B}(x_B, \mathbf{b}_T; m^2, \mu_0) \left(\frac{Q^2}{m^2}\right)^{\tilde{K}(b_T; \mu_0)} \\ &= \int d^2\mathbf{b}_T e^{i\mathbf{q}_{hT}\cdot\mathbf{b}_T} \tilde{f}_{j/A}(x_A, \mathbf{b}_T) \tilde{f}_{\bar{j}/B}(x_B, \mathbf{b}_T) e^{\tilde{K}(b_T; \mu_0) \ln(Q^2/m^2)} \end{aligned}$$

$\tilde{K}(b_T)$ is a decreasing function of b_T

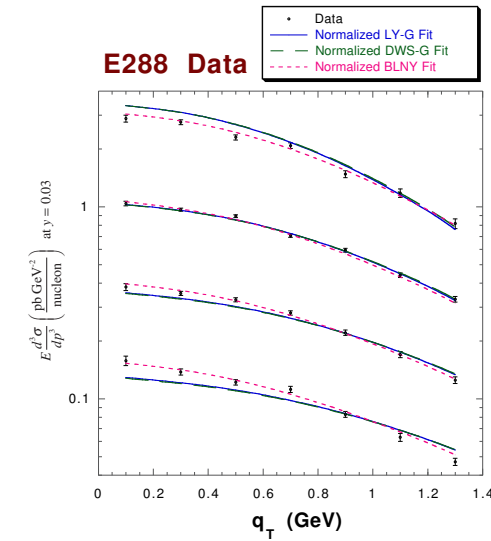
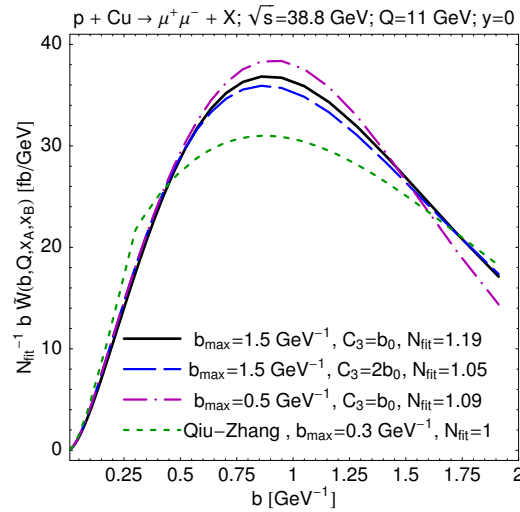


So it shifts TMD to smaller b_T and broadens the q_T distribution . . .

Evolution in q_T v. b_T

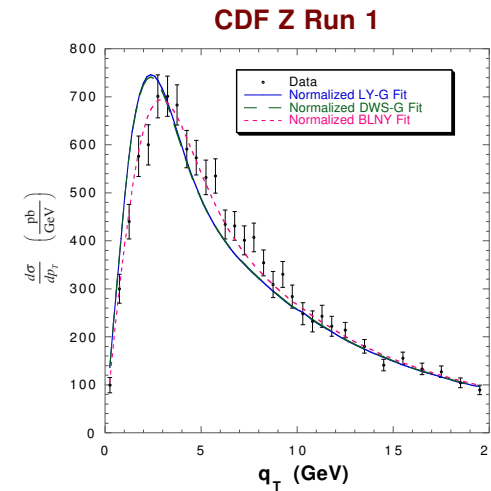
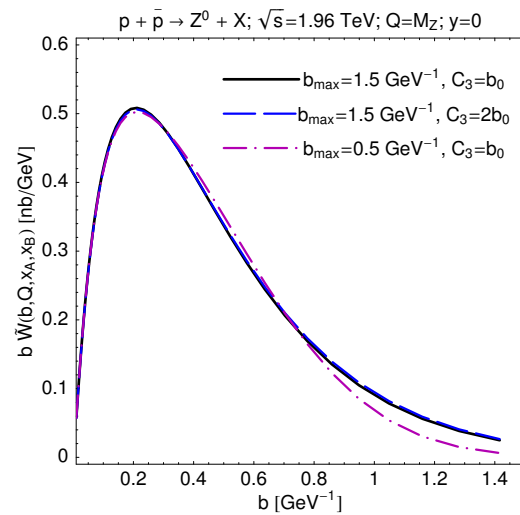
$$Q = 11 \text{ GeV}$$

$$\sqrt{s} = 38.8 \text{ GeV}$$



$$Q = m_Z$$

$$\sqrt{s} = 1960 \text{ GeV}$$



Konychev & Nadolsky, PLB 633, 710 (2006)

Landry et al., PRD 67,073016 (2003)

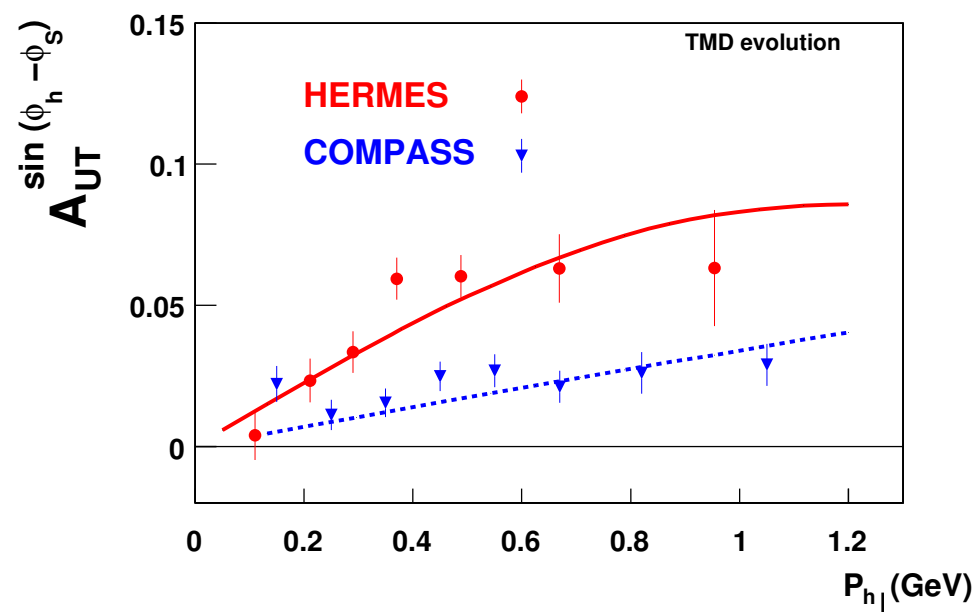
- N.B.
- x -dependence of k_T shape in TMD pdfs
 - Sensitivity to non-perturbative region goes away at large Q

Data used in fits, etc

Data	Q	x	ff	pdf	Sivers	\tilde{K}	
HERMES (SIDIS)	$\sqrt{2.4}$ GeV	0.04–0.3	✓	✓	✓		Torino
COMPASS (SIDIS)	$\sqrt{3.8}$ GeV	0.01–0.3					Predict
E288 (DY)							
$\sqrt{s} = 27.4$ GeV	5–9 GeV	0.18–0.33		✓		✓	KN, BLNY
E605 (DY)							
$\sqrt{s} = 38.8$ GeV	7–18 GeV	0.18–0.46		✓		✓	KN, BLNY
R209 (DY)							
$\sqrt{s} = 62$ GeV	5–11 GeV	0.08–0.18		✓		✓	KN, BLNY
D0, CDF (DY)							
$\sqrt{s} = 1.8$ TeV	m_Z	0.05		✓		✓	KN, BLNY

Prediction for COMPASS v. HERMES

[Aybat, Prokudin, Rogers, PRL 108, 242003 (2012)]



On basis of

- Unpolarized DY for: $g_K(b_T)$ in BLNY fit (with $b_{\max} = 0.5 \text{ GeV}^{-1} = 0.1 \text{ fm}$)
- HERMES (Torino fit) for: fragmentation fns. unpolarized TMD pdf, Sivars
- But *not* non-perturbative unpolarized TMD pdfs of BLNY

N.B. Shape looks inaccurate

Problem with BLNY fit

[Landry et al., PRD 67,073016 (2003)]

- Non-perturbative factor (large b_T):

$$\exp\left\{-b_T^2 [0.21 + 0.68 \ln(Q/3.2 \text{ GeV}) - 0.126 \ln(100x_Ax_B)]\right\}$$

- Coefficient of b_T^2 becomes negative when Q is small and x_Ax_B large.

E.g., $Q = 3.2 \text{ GeV}$ and $x_A = x_B = 0.3$,

or $Q = \sqrt{2.4} \text{ GeV}$ (HERMES) and $x_A = x_B = 0.1$,

- So fit is not applicable beyond range of fairly small b_T relevant for the fitted data.

- But $b_{\max} = 0.5 \text{ GeV}^{-1} = 0.1 \text{ fm}$

- Konychev & Nadolsky, PLB 633, 710 (2006) use $b_{\max} = 1.5 \text{ GeV}^{-1} = 0.3 \text{ fm}$.
They get

$$g_K(b_T) = \frac{0.16}{2}b^2 = 0.08b^2, \quad \text{instead of} \quad g_K(b_T) = \frac{0.68}{2}b^2 = 0.34b^2$$

with better fit

- BLNY result \implies overestimate of evolution of Sivers from HERMES to COMPASS and Polarized DY @ Fermilab

Sun & Yuan [arXiv:1304.5037v1]

- They use evolution factor

$$\exp \left\{ -2C_F \int_{Q_0}^Q \frac{d\mu}{\mu} \frac{\alpha_s(\mu)}{\pi} \left[\ln \left(\frac{Q^2}{\mu^2} \right) + \ln \left(\frac{Q_0^2 b_T^2}{c_0^2} \right) - \frac{3}{2} \right] \right\}$$

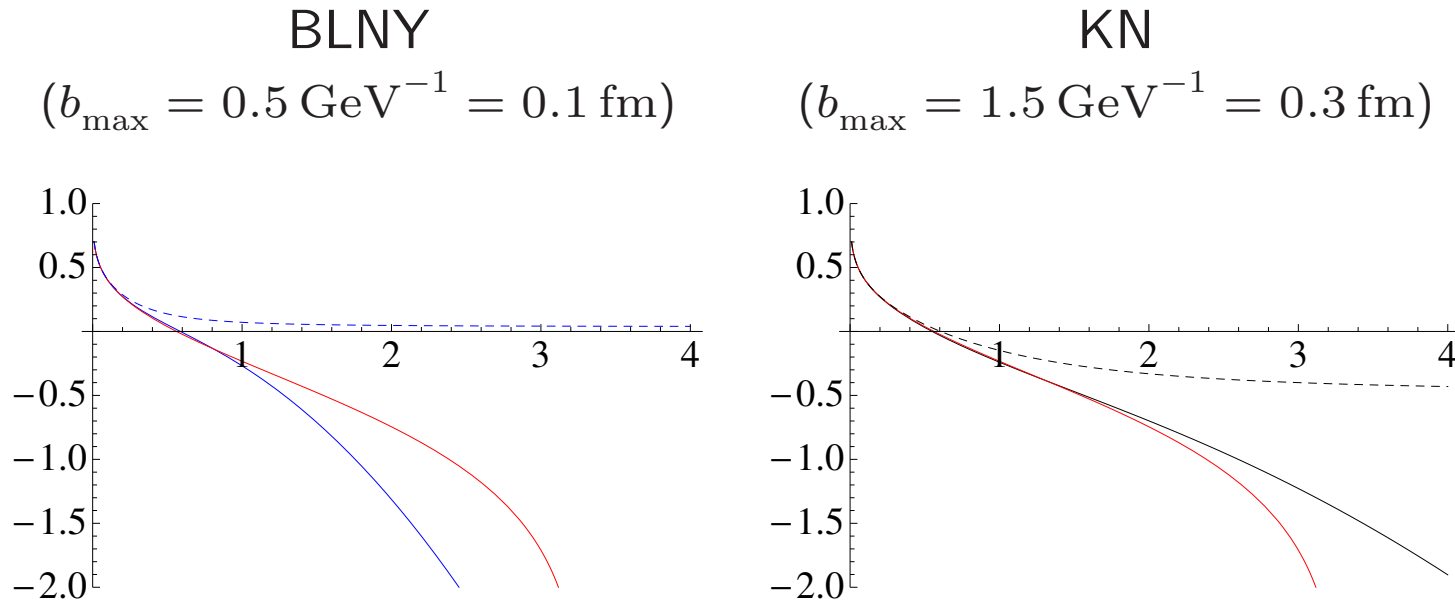
[with $c_0 = 2e^{-\gamma_E}$]

- Logarithms of b_T not resummed, no non-perturbative function
- Hence effective evolution kernel is

$$\frac{d \ln \tilde{\sigma}}{d \ln Q^2} = -C_F \frac{\alpha_s(Q)}{\pi} \ln \left(\frac{Q_0^2 b_T^2}{c_0^2} \right) + \text{terms independent of } b_T$$

which is Q -dependent

Plots of $\tilde{K}(b_T, \mu_0 = 2 \text{ GeV})$ I



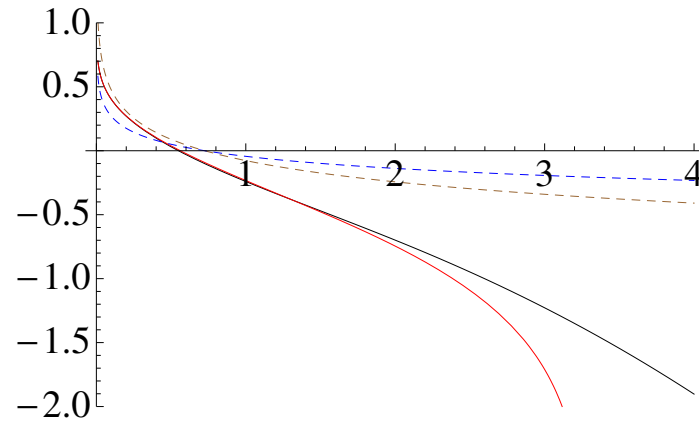
Code:

Red	With RG improvement
Blue	BLNY (Landry et al., PRD 67,073016 (2003))
Blue dotted	Perturbative part of BLNY
Black	KN (Konychev & Nadolsky, PLB 633, 710 (2006))
Black dotted	Perturbative part of KN

N.B. $4 \text{ GeV}^{-1} = 0.8 \text{ fm}$

Plots of $\tilde{K}(b_T)$ II

KN v. Sun-Yuan



Code: Red With RG improvement
 Black KN
 Brown dashed Sun-Yuan at 2 GeV
 Blue dashed Sun-Yuan at 10 GeV

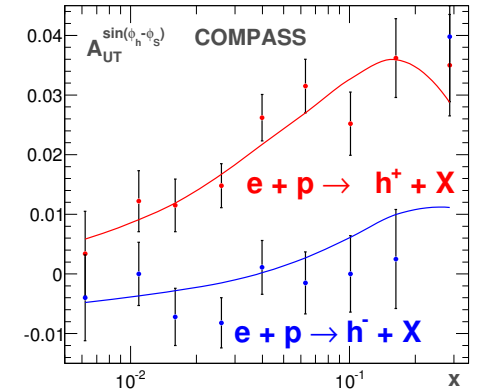
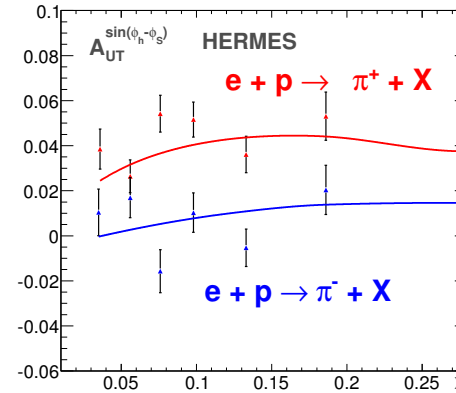
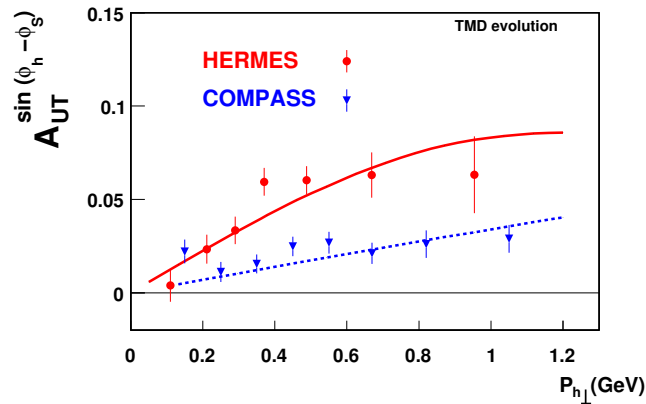
N.B. $4 \text{ GeV}^{-1} = 0.8 \text{ fm}$

Are the results consistent?

Aybat et al.

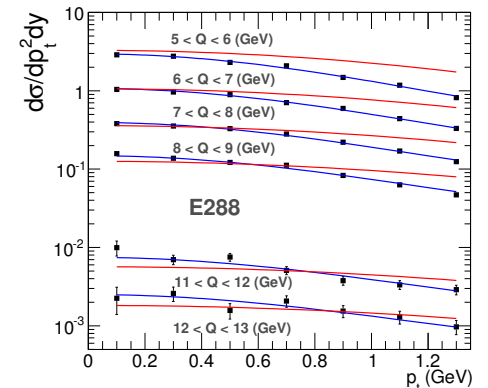
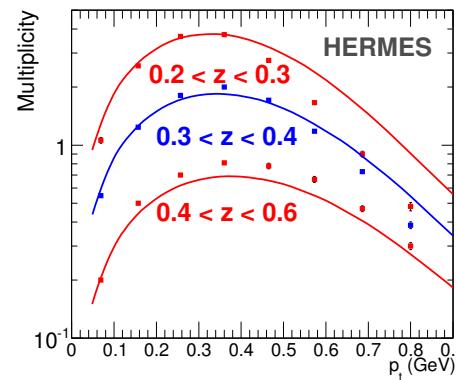
Sun & Yuan

Sivers



unpol.

Evolution from higher energy fit



- Slower evolution $Q = \sqrt{2.4} \text{ GeV}$ to $Q = 9 \text{ GeV}$; faster evolution above $Q = 9 \text{ GeV}$
- Roughly: Relevant b_T : $3\text{--}4 \text{ GeV}^{-1}$ for HERMES $1\text{--}2 \text{ GeV}^{-1}$ at $Q = 11 \text{ GeV}$
- Does $\tilde{K}(b_T)$ flatten at $b_T \gtrsim 2 \text{ GeV}^{-1}$?

Conclusions

- $A_N(\text{DY})$ reduced substantially compared with HERMES
- Evolution, perturbative and non-perturbative, is controlled by $\tilde{K}(b_T, \mu_0)$
- But it's used at larger b_T than where “non-perturbative” content was fit
- Probably Aybat-Prokudin-Rogers too pessimistic
- But Sun-Yuan probably too optimistic, and don't allow for known physics issues
- We need global fit/calculation of $\tilde{K}(b_T, \mu_0)$
- New fit mustn't violate agreement with any existing data and principles
- Intrinsic TMD: Probably exponential rather than Gaussian better at large b_T
- Perhaps even constant for \tilde{K} at large b_T
- We also need to evaluate effect of power corrections in factorization

EXTRA SLIDES

Power corrections (“higher twist”)

- Factorization derived up to errors suppressed by a power of Q
- SSA (Sivers-type):
 - Graph-by-graph leading power if $q_T \ll Q$
 - Suppressed by M/q_T at large q_T . (“Twist-3”)
 - But gluon radiation (evolution) dilutes small q_T SSA as Q increases
 - Need answer to: Does that uncover power corrections, or are these also diluted?
- Need answer to: How accurate is factorization when Q is not large?

What form for large b_T ?

- Standard:

- $e^{-\text{const} \times b_T^2}$ in TMD pdf
- $e^{-\tilde{K} \ln Q^2} \sim e^{-\text{const} \times b_T^2 \ln Q^2}$ in evolution

Coefficients significantly non-zero according to fits: Landry et al. PRD 67,073016 (2003), and Konychev & Nadolsky, PLB 633, 710 (2006)

- But: Euclidean correlation functions in QFT are usually e^{-mb_T}
- KN & BLNY fits are at relatively large Q (10 GeV up), and hence determine non-perturbative functions up to $b_T \lesssim 2 \text{ GeV}^{-1} = 0.4 \text{ fm}$
- But to get evolution from HERMES ($Q \sim \sqrt{2.4} \text{ GeV}$), we need the non-perturbative functions at larger b_T : Extrapolation v. theoretical motivation, . . .
- Need to retry fits with better forms at large b_T .
- *Un*polarized HERMES v. higher energy DY should be enough.