# CP violation reach in the next decade and beyond

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#### **CP** violation

Like in the quark sector mixing can cause CP violation

 $P(\nu_{\alpha} \to \nu_{\beta}) - P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) \neq 0$ 

The size of this effect is proportional to

$$J_{CP} = \frac{1}{8}\cos\theta_{13}\sin2\theta_{13}\sin2\theta_{23}\sin2\theta_{12}\sin\delta$$

but the asymmetry

$$\frac{P(\nu_{\alpha} \to \nu_{\beta}) - P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta})}{P(\nu_{\alpha} \to \nu_{\beta}) + P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta})} \propto \frac{1}{\sin 2\theta_{13}}$$

The experimentally most suitable transition to study CP violation is  $\nu_e \leftrightarrow \nu_\mu$ .

#### **Consequences for experiments**

- need to measure 2 out of  $P(\nu_{\mu} \to \nu_{e})$ ,  $P(\bar{\nu}_{\mu} \to \bar{\nu}_{e}), P(\nu_{e} \to \nu_{\mu}) \text{ and } P(\bar{\nu}_{e} \to \bar{\nu}_{\mu})$
- need more than 1 energy and/or 1 baseline
- large  $\theta_{13}$  implies small CP asymmetries  $\Rightarrow$  need for small systematics

Ultimately, the combination of large exposure  $\gg 100$  kt MW yr with percent-level systematics will be needed – see Ken's talk.

 $\nu_{\rm e}/\nu_{\mu}$  x-sections



#### PH, M. Mezzetto, T. Schwetz arXiv:0711.2950

Appearance experiments using a (nearly) flavor pure beam can **not** rely on a near detector to predict the signal at the far site!

Large  $\theta_{13}$  most difficult region.

## **QE** energy reconstruction



Nuclear effects change the relation between true neutrino energy and lepton energy

Lalakulich, Mosel, arXiv:1208.3678.

Inferring the CP phase from QE spectrum seems quite difficult

Not obvious that near detectors alone can solve this problem.

#### **Nuclear effects**



arXiv:1307.1243

$$N_i^{\text{test}}(\alpha) = \alpha \times N_i^{QE} + (1 - \alpha) \times N_i^{QE - like}$$

where  $\alpha = 0$  corresponds to perfectly know nuclear effects and  $\alpha = 1$  to entirely unknown nuclear effects in the fit.

## **CP** precision and systematics

We specifically simulate near and far detectors We use common assumptions for all experiments on

- cross sections split into QE, RES and DIS for each flavor and neutrinos and antineutrinos
- cross section ratios between e and  $\mu$  flavors for QE, RES and DIS and neutrinos and antineutrinos
- fiducial volume and near/far extrapolation errors

We use experiment type specific errors for

- fluxes
- beam backgrounds
- detector backgrounds

### **Systematics I**



Nuclear effects NOT included Near detector crucial for new physics searches NOvA<sup>+</sup> higher risk from systematics Current  $\Delta\delta$  is 30-35° Fogli *et al.*, 2012

#### arXiv:1209.5973

### **Systematics II**



Disappearance data can play the role of near detector if three flavor framework is assumed

Scaling with luminosity is strongyl affected by systematics

arXiv:1209.5973

# Luminosity scaling



Extrapolating superbeam performances beyond several 100 kt MW years is entirely dependent on the assumptions on systematics!

LBNE10 – 70 kt MW yr LBNE – 238 kt MW yr LBNE + Project X – 782 kt MW yr T2HK –3920 kt MW yr NuMAX+ 34kt – 1020 kt MW yr

### LBNE – CP sensitivity



### **Comparison of CP sensitivities**



# Summary

- New facilities are indispensable to fully exploit the discovery of neutrino oscillation
- LBNE10, LBNE and LBNE + Project X provide a staged program to discover CP violation with increasing reach.
- LBNE, in particular, together with Project X is competitive in performance
- Eventually systematics issue, which currently are not well understood, will limit the sensitivity of pion-decay based beams