# **EWPO** in the pMSSM

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- 1. Introduction
- 2. Our tool: MasterCode
- 3. Some pMSSM results
- 4. Conclusions

# 1. Introduction

Comparison of observables with theory:

Precision data:  

$$M_W, \sin^2 \theta_{\text{eff}}, a_\mu, M_h$$
  $\leftrightarrow$  SM, MSSM , ...  
 $\downarrow$ 

Test of theory at quantum level: Sensitivity to loop corrections, e.g.  $\boldsymbol{X}$ 



BSM: limits on  $M_X$ 

Very high accuracy of measurements and theoretical predictions needed

Precision observables in the SM and the MSSM  $M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $M_h$ ,  $(g-2)_{\mu}$ , b physics, ...

**A)** Theoretical prediction for  $M_W$  in terms

### Evaluate $\Delta r$ from $\mu$ decay $\Rightarrow M_W$

One-loop result for  $M_W$  in the SM: [A. Sirlin '80], [W. Marciano, A. Sirlin '80]

$$\Delta r_{1-\text{loop}} = \Delta \alpha - \frac{c_{W}^2}{s_{W}^2} \Delta \rho + \Delta r_{\text{rem}}(M_H)$$
$$\sim \log \frac{M_Z}{m_f} \sim m_t^2 - \log (M_H/M_W)$$
$$\sim 6\% \sim 3.3\% \sim 1\%$$

Precision observables in the SM and the MSSM  $M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $M_h$ ,  $(g-2)_{\mu}$ , b physics, . . .

**A)** Theoretical prediction for  $M_W$  in terms

**B)** Effective mixing angle:

$$\sin^2 \theta_{\text{eff}} = \frac{1}{4 |Q_f|} \left( 1 - \frac{\operatorname{Re} g_V^f}{\operatorname{Re} g_A^f} \right)$$

Higher order contributions:

$$g_V^f \to g_V^f + \Delta g_V^f, \quad g_A^f \to g_A^f + \Delta g_A^f$$

**Corrections to**  $M_W$ ,  $\sin^2 \theta_{\text{eff}} \rightarrow \text{approximation via the } \rho$ -parameter:

 $\rho$  measures the relative strength between neutral current interaction and charged current interaction

$$\rho = \frac{1}{1 - \Delta \rho} \qquad \Delta \rho = \frac{\Sigma_Z(0)}{M_Z^2} - \frac{\Sigma_W(0)}{M_W^2}$$

(leading, process independent terms)

 $\Delta \rho$  gives the main contribution to EW observables:



 $\Delta \rho^{\text{SUSY}}$  from  $\tilde{t}/\tilde{b}$  loops > 0  $\Rightarrow M_W^{\text{SUSY}} \gtrsim M_W^{\text{SM}}$ ,  $\sin^2 \theta_{\text{eff}}^{\text{SUSY}} \lesssim \sin^2 \theta_{\text{eff}}^{\text{SM}}$ 

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# SM result for $M_W$ and $\sin^2 \theta_{\text{eff}}$ :

- full one-loop
- full two-loop
- leading 3-loop via  $\Delta\rho$
- leading 4-loop via  $\Delta \rho$

# Our MSSM result for $M_W$ and $\sin^2 \theta_{\text{eff}}$ :

- full SM result (via fit formel)
- full MSSM one-loop (incl. complex phases)
- all existing two-loop  $\Delta\rho$  contributions
- $\Rightarrow$  non- $\Delta \rho$  one-loop and  $\Delta \rho$  two-loop contributions sometimes non-negligible!

### The W boson mass

Experimental accuracy:

Today: LEP2, Tevatron:  $M_W^{\text{exp}} = 80.385 \pm 0.015 \text{ GeV}$ ILC:

$$\delta M_W^{ extsf{exp,ILC}} = 5-6$$
 MeV

Theoretical accuracies:  $\Rightarrow$  see my next talk!

Effective weak leptonic mixing angle:  $\sin^2 \theta_{eff}$ 

Experimental accuracy:

Today: LEP, SLD:  $\sin^2 \theta_{eff}^{exp} = 0.23153 \pm 0.00016$ 

ILC/GigaZ:

 $\delta \sin^2 \theta_{\text{eff}}^{\text{exp,ILC}} = 0.000013$ 

Theoretical accuracies:  $\Rightarrow$  see my next talk!

# What is the pMSSM?

Particle content of the MSSM:

$$\begin{bmatrix} u, d, c, s, t, b \end{bmatrix}_{L,R} \begin{bmatrix} e, \mu, \tau \end{bmatrix}_{L,R} \begin{bmatrix} \nu_{e,\mu,\tau} \end{bmatrix}_{L} \quad \text{Spin } \frac{1}{2} \\ \begin{bmatrix} \tilde{u}, \tilde{d}, \tilde{c}, \tilde{s}, \tilde{t}, \tilde{b} \end{bmatrix}_{L,R} \begin{bmatrix} \tilde{e}, \tilde{\mu}, \tilde{\tau} \end{bmatrix}_{L,R} \begin{bmatrix} \tilde{\nu}_{e,\mu,\tau} \end{bmatrix}_{L} \quad \text{Spin } 0$$

$$g \quad \underbrace{W^{\pm}, H^{\pm}}_{\tilde{g}} \qquad \underbrace{\gamma, Z, H_1^0, H_2^0}_{\tilde{\chi}_{1,2}^{\pm}} \qquad \text{Spin 1 / Spin 0}$$

$$\tilde{g} \quad \tilde{\chi}_{1,2}^{\pm} \qquad \tilde{\chi}_{1,2,3,4}^0 \qquad \text{Spin } \frac{1}{2}$$

most general case:  $\Rightarrow$  105 new parameters: masses, mixing angles, phases

How many free parameters do you want?

- what is relevant for (your) physics case?
- what can you handle reliably?
- is your set really representative?

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- what is relevant for (your) physics case?
- what can you handle reliably?
- is your set really representative?
- $\Rightarrow$  many options:
- pMSSM7, pMSSM8, ...
- pMSSM19, pMSSM20, ...
- pMSSM105

# Further questions:

- do you scan? do you scatter?(How many points are reliable?)
- MCMC to find "best-fit" points? do you find them?
- do you pick "some benchmarks"?
- does your choice cover everything (relevant)?

Snowmass: 24 points with  $\mathcal{O}(20)$  free parameters (real)

[M. Cahill-Rowley, J. Hewett, A. Ismail, M. Peskin, T. Rizzo, arXiv:1305.2419 [hep-ph]]



⇒ collaborative effort of theorists and experimentalists
[Buchmüller, Cavanaugh, De Roeck, Dolan, Ellis, Flächer, SH, Isidori, Marouche, Martinez
Santos, Olive, Rogerson, Ronga, de Vries, Weiglein]

Über-code for the combination of different tools:

- Über-code original in Fortran, now re-written in C++
- tools are included as subroutines
- compatibility ensured by collaboration of authors of "MasterCode" and authors of "sub tools" /SLHA(2)
- sub-codes in Fortran or C++
- $\Rightarrow$  evaluate observables of one parameter point consistently with various tools

cern.ch/mastercode

# Status of the "MasterCode":

- one model: (MFV) MSSM (see next section)
- tools included:
  - *B*-physics observables [*SuFla*]
  - more *B*-physics observables [*SuperIso*]
  - Higgs related observables,  $(g-2)_{\mu}$  [FeynHiggs]
  - Electroweak precision observables [FeynWZ]
  - Dark Matter observables [MicrOMEGAs, DarkSUSY]
  - for GUT scale models: RGE running [SoftSusy]
- $\Rightarrow$  all most-up-to-date codes on the market!
- added:  $\chi^2$  analysis code [*Minuit*]
- currently being implemented:
  - Higgs constraints (for  $\chi^2$  contributions . . . ) [HiggsBounds]
- planned: inclusion of more tools / more models

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 $\Rightarrow$  crucial for precision!

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(Some) Electroweak precision observables in the MasterCode

- ( $\rightarrow$  as for blue band analysis, except  ${\sf \Gamma}_W)$
- 1.  $M_W$  (LEP/Tevatron)
- 2.  $A^e_{LR}$  (SLD)
- 3.  $A^b_{\mathsf{FB}}$  (LEP)
- 4.  $A_{\mathsf{FB}}^c$  (LEP)
- 5.  $A_{\mathsf{FB}}^l$
- 6.  $A_b, A_c$
- **7**.  $R_b, R_c$
- 8.  $\sigma_{had}^0$

# $\Rightarrow$ largest impact: (1), (2), (3)

(2), (3), (4), ...  $\Rightarrow \sin^2 \theta_{\text{eff}}$ 

(Some) B/K physics observables in the MasterCode

- 1.  $BR(b \rightarrow s\gamma)$  (MSSM/SM)
- 2. BR( $B_s \rightarrow \mu^+ \mu^-$ )
- **3**. Δ*M*<sub>s</sub>
- 4.  $R(\Delta M_s/\Delta M_d)$
- 5.  $BR(B_u \rightarrow \tau \nu_{\tau})$  (MSSM/SM)
- 6. BR( $B \to X_x \ell^+ \ell^-$ )
- 7.  $BR(K \rightarrow \ell \nu)$  (MSSM/SM)
- 8. BR( $\Delta M_K$ ) (MSSM/SM)
- $\Rightarrow$  largest impact: (1) and (2)

– anomalous magnetic moment of the muon:  $(g-2)_{\mu}$ 

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Higgs physics observables in the MasterCode

- lightest Higgs mass:  $M_h$
- effective mixing angle:  $\alpha_{eff}$  , especially for  $\sin^2(\beta \alpha_{eff})$

– anomalous magnetic moment of the muon:  $(g-2)_{\mu}$ 

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Dark Matter observables in the MasterCode

- CDM density:  $\Omega_{\chi}h^2$
- Direct detection cross section:  $\sigma_p^{SI}$  (prediction; not incl. in the fit yet)

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Dark Matter observables in the MasterCode

- CDM density:  $\Omega_{\chi}h^2$
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# SM parameters

- top mass:  $m_t$
- -Z boson mass:  $M_Z$
- hadronic contribution to fine structure constant:  $\Delta \alpha_{had}$

Simple strategy:

- 1. Take the 24 "Snowmass pMSSM-SLHA files"
- 2. Pipe them through MasterCode
- 3. Check results! Do we learn anything?

Results for  $M_W$ :



#### $\Rightarrow$ mostly ok, some not, some better, . . .





#### $\Rightarrow$ draw your own conclusions . . .

## Spectrum point # 8:



#### $\Rightarrow$ draw your own conclusions . . .



#### $\Rightarrow$ mostly ok, same "problematic" points





#### $\Rightarrow$ mostly ok, same "problematic" points





#### $\Rightarrow$ different pattern

Sven Heinemeyer Snowmass: Seattle EF WS, 07/01/'13



Results for  $A^e_{LR}(SLD)$ :

Total  $\chi^2$  values:



### $\Rightarrow$ Some points clearly disfavored







### Comparison to our scan:



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#### $\Rightarrow$ deviation from line can indicate non- $\Delta \rho$ contributions

# 4. Conclusinos

- EWPO can give valuable information about SM, BSM Best:  $M_W$ ,  $\sin^2 \theta_{\text{eff}}$ ,  $(g-2)_{\mu}$ , ...
- SUSY corrections mainly via  $\Delta \rho$ But also non  $\Delta \rho$  can be relevant
- Our tool: MasterCode
   Only code that contains the most precise EWPO evaluation in the MSSM

# • pMSSM?

Snowmass: 24 points with O(20) free parameters (real) [1305.2419 [hep-ph]] Sufficient? Representative?  $\Rightarrow$  check (again) EWPO!

• Result: Most points ok, about 5 clearly disfavored Some more (non-negligible) variation among the rest Extreme spectra!?