SUSY scenarios we're not looking for* but should be

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Disclaimer

Many important "loopholes" are covered in existing or planned searches

- RPV
- Reduced MET compressed spectra, "stealth" SUSY,...
- Long decay chains
- NLSP \rightarrow gravitino + X

This talk will discuss a few areas that may benefit from more attention

A personal selection, no attempt at completeness

The Big Picture



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125 GeV Higgs SM-like couplings







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Naturalness = no tuning of parameters Simplicity = minimality of structure & assumptions

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So what about the quadrupole?



Tuning!







Outline

- Tuned SUSY
- Natural SUSY

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 - R-parity violation
 - Hidden sector dark matter

Tuned SUSY

Occam's razor:

"Entities must not be multiplied beyond necessity"



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MSSM with one tuning is arguably the <u>simplest</u> explanation of particle physics data

Implications:

- SUSY most likely "just around the corner"
- Keep looking for standard SUSY signals

Natural SUSY

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Requirements:

 $\mu \lesssim 200 \text{ GeV}$ $m_{\tilde{t}} \lesssim 500 \text{ GeV} \longleftrightarrow$ $m_{\tilde{g}} \lesssim 1.5 \text{ TeV}$

EWino searches, stop, sbottom searches

Natural SUSY

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But also: Higgs sector beyond MSSM

 $V_{\text{eff}} = m_H^2 |H|^2 + \lambda |H|^4 \quad \Rightarrow (126 \text{ GeV})^2 = 2\lambda v^2$

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Need additional contributions to quartic

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- Non-decoupling D-terms

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- NMSSM
- Non-decoupling D-terms
- ... or additional sources of EWSB
 - "Induced" EWSB

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Motivate searches for additional Higgs bosons with significant mixing to 126 GeV state

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Higgs sector

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TYPE 1 2HDM: LHC A \rightarrow Zh vs. couplings, m_A=300 GeV TYPE 2 2HDM: LHC A \rightarrow Zh vs. couplings, m_A=300 GeV $t_{\beta} = 100$ $t_{\beta} = 100$ Zh disc. 14 TeV 300 fb⁻¹ ● 5σ A→Zh disc. 14 TeV 300 fb⁻ $t_{\beta} = 50$ Zh disc. 14 TeV 3000 fb⁻¹ 5σ A→Zh disc. 14 TeV 3000 fb⁻¹ disc 33 TeV 3000 fb-5σ A→Zh disc. 33 TeV 3000 fb⁻¹ allowed 14 TeV 300 fb-1 95% coupl. allowed 14 TeV 300 fb-1 pl. allowed 14 TeV 3000 fb⁻¹ 95% coupl. allowed 14 TeV 3000 fb⁻ $t_{\beta} = 10$ $t_{\beta} = 10$ $t_{\beta} = 5$ $t_{\beta} = 5$ $t_{\beta} = 1$ $t_{\beta} = 1$ -0.2 0.2 0.0 -0.10-0.050.05 -0.40.4 0.00 0.10 $\cos(\beta - \alpha)$ $\cos(\beta - \alpha)$

Example: $A \rightarrow Zh \rightarrow (II)(bb)$ or (TT)

Brownson, Kukartsev, Narain, Heintz, Stupak, Craig Snowmass 2013

0.10

Also: $H \rightarrow ZZ \rightarrow (II)(II)$



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Other examples deserving further study: H/A \rightarrow tt, bb, $\tau\tau$ H \rightarrow hh H \rightarrow A Z \rightarrow ZZh

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Cosmology motivates displaced vertices Barry, Graham, Rajendran arXiv:1310.3853



UDD destroys baryon asymmetry in early universe ⇒ requires low scale baryogenesis



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NLSP = chargino



NLSP = neutralino



 \Rightarrow displaced vertex



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 \Rightarrow exploding track



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SUSY WIMP:

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In the absence of signals in direct detection and collider searches, look at plausible mechanisms beyond WIMPs \Rightarrow new signals?

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• SuperWIMP (Feng, Rajaraman, Takayama 2003)

LOSP dominates universe, then decays

$$\Omega_X = \frac{m_X}{m_{\rm LOSP}} \Omega_{\rm LOSP}$$

X = gravitino or ...

(Hall, Jedamzik, March-Russell, West 2009)

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 $m_X \gtrsim \mathrm{keV}$ (structure formation) \Rightarrow motivates highly displaced decays

(Kaplan, ML, Zurek 2009)

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visible hidden

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Constraining/measuring dark matter mass possible

Conclusions

- Impressive breadth and depth of SUSY searches
- Both tuned and natural versions of SUSY are still plausible and important to test
- BMSSM Higgs and displaced vertices are important