

Mono-X Searches at CMS and ATLAS



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Outline

- LHC and ATLAS/CMS
- Introduction/Motivation

- Mono-jets
- Mono-photons
- Mono-W/Z

• Final notes



[&]quot;Particles, particles, particles."

The results are put in terms of searches for

- Dark Matter
- Large Extra Dimensions and unparticles
- Supersymmetry
- Invisible Higgs

LHC Performance (2010-2012)



...will come back in 2015 with 13-14 TeV collisions



CMS & ATLAS

Recorded Luminosity [pb⁻¹/0.1]

ATLAS Online Luminosity

10 15

√s = 8 TeV, ∫Ldt = 20.8 fb⁻¹, <µ> = 20.7

_____√s = 7 TeV, ∫Ldt = 5.2 fb⁻¹, <µ> = 9.1

About 5 fb-1 collected by CMS and ATLAS @ 7 TeV More than 21 fb-1 collected by CMS and ATLAS @ 8 TeV

challenging pile up conditions for the physics analysis



CMS Integrated Luminosity, pp, 2012, $\sqrt{s}=$ 8 TeV

Search for Dark Matter



Evidence for Dark Matter

The rotation of the stars around the center of the galaxies is not consistent with the amount of mass observed (L/M ratio)_{SUN}

Spherical dark matter halo



Large distortion of the imagines of distant galaxies due to gravitation lensing → indication of DM in galaxy clusters

Gravitational Lensing

Collisions of clusters of galaxies .

Considered the ultimate demonstration of the presence of Dark Matter since this does not involve Newton's Law

Planck (20 March 2013) arXiv:1303.5062v1



Dark Matter Candidates

- Neutrinos ? (Ω_vh² < 0.0067 @ 95%CL)
- Sterile Neutrinos
- Axions
- SUSY particles
 - Lightest neutralino
 - Sneutrinos
 - Gravitinos
 - Axinos
- KK states (UED)
- Wimpzillas
- - •••••

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General requirements

- Electrically Neutral ("dark")
- Stable (lifetime larger than age of the Universe)
- Massive and Weakly interacting $(\Omega_{\rm CDM} \, h^2 \sim 0.1)$

→WIMPS

Note: No reason DM should be made out of a single component (neutrinos exist)

WIMP Pair Production at Hadron Colliders

At colliders (LHC) WIMPs can be produced in pairs leading to "nothing to detect" in the final state

Such events are tagged via the presence of an energetic jet or a photon (or a W/Z) from initial state radiation



\rightarrow Mono-jet, Mono-photon, Mono-W/Z







Rather spectacular and distinctive signature to search for new physics





Good agreement with SM

Effective Theory

J. Goodman et al., Phys.Rev.D82:116010,2010

(model independent approach)

Effective Lagrangian approach (contact interaction) with parameters M_* (Λ) and m_{γ}

 $M_*^2 \sim M^2/g_1g_2$ [M > 2 m_{\chi}, $g_1g_2 < (4\pi)^2$] assuming the interaction is mediated by a heavy particle with mass M and couplings g_1 and g_2

Different operators are considered with different structures and here χ will be taken as Dirac fermions

Important note:

Not clear whether the effective approach under- or over-estimates the cross sections since this depends on the details of the unknown UV limit of the theory

Strictly speaking theory only applicable D11 gg s when M is much larger than the energy scale present in the reaction $[Q^2 << (4\pi M_*)^2, m_{\gamma} < 2\pi M_*]$



Name	Initial state	Type	Operator		
D1	qq	scalar	$rac{m_q}{M_\star^3}ar\chi\chiar q q$		
D5	qq	vector	$rac{1}{M_\star^2}ar\chi\gamma^\mu\chiar q\gamma_\mu q$		
D8	qq	axial-vector	$rac{1}{M_\star^2}ar\chi\gamma^\mu\gamma^5\chiar q\gamma_\mu\gamma^5 q$		
D9	qq	tensor	$rac{1}{M_\star^2}ar\chi\sigma^{\mu u}\chiar q\sigma_{\mu u}q$		
D11	gg	scalar	$rac{1}{4M_\star^3}ar\chi\chilpha_s(G^a_{\mu u})^2$		





WIMP-nucleon cross section

JHEP09(2012)094 JHEP 04 (2013) 075



Different operators contribute either to spin-dependent or spin-independent WIMP-nucleon cross sections



Very strong limits for SD processes



Within the assumption of the validity of the effective theory the LHC results are competitive to direct detector experiments (particularly relevant at $m_{\chi} < 10$ GeV) Large sensitivity in case of D11 (gg initiated)

Mono-photons

Phys. Rev. Lett.108 261803 (2012)



 $P_t^{\gamma} > 145 \text{ GeV}, |\eta^{\gamma}| < 1.44$, isolated $E_t^{miss} > 130 \text{ GeV}$ Veto on leptons, isolated tracks, jets

Good agreement with SM

Background dominated by the irreducible $Z\gamma (\rightarrow vv\gamma)$ contribution followed by photon fakes and non-collision background, plus other small contributions

5 fb⁻¹

CMS

 QCD-jet fakes data driven using EM-enriched sample with loose photon requirements

• Time distribution of the calorimeter energy deposit used to estimate non-collision background

Source	Estimate	
Jet Mimics Photon	11.2 ± 2.8	
Beam Halo	11.1 ± 5.6	
Electron Mimics Photon	3.5 ± 1.5	
Wγ	3.0 ± 1.0	
γ +jet	0.5 ± 0.2	
$\gamma\gamma$	0.6 ± 0.3	
$Z(\nu \bar{\nu})\gamma$	45.3 ± 6.9	
Total Background	75.1 ± 9.5	
Total Observed Candidates	73	





3 x10⁻⁴⁰ cm² (10⁻³⁹ cm²) are excluded for spin –dependent (spin-independent) operators

(*) not confirmed by Xenon100 and LUX results





CMS-PAS-EXO-12-048



Same strategy as in the 7 TeV analysis

Good agreement with SM predictions

E,^{miss} > 200 GeV $p_{\tau}(j1) > 110 \text{ GeV}$ $N_{iet}(p_T > 30 \text{ GeV}) < 3$

Δφ (j1,j2) < 2.5 Lepton vetoes

Mono-jets 8 TeV

/ 20 GeV/c

19.5 fb⁻¹







8 TeV data improves previous limits on Λ (M_{*}) by about 150 GeV (approaching the 1 TeV scale)



For $m_{\chi} \sim 10$ GeV the experiment excludes: WIMP-nucleon x-section > 1.2 x10⁻³⁹ cm² at 90% CL (vector operator) WIMP-nucleon x-section > 4.2 x10⁻⁴¹ cm² at 90% CL (axial vector operator)



For $m_{\chi} \sim 10$ GeV the experiment excludes: WIMP-nucleon x-section > 1.2 x10⁻³⁹ cm² at 90% CL (vector operator) WIMP-nucleon x-section > 1.1 x10⁻⁴⁴ cm² at 90% CL (scalar)



Light Mediator

Exploring the scenario with a vector-like coupling and a light mediator with given M and Γ and different DM masses

Μ





CMS-PAS-EXO-12-048

For M > few x 100 GeV the EFT is adequate and somehow conservative in the bounds on Λ (note however the effective couplings become large)

For M < 100 GeV the collider bounds are weakened





Mono-W (90% CL limits)

CMS-PAS-EXO-13-004



	$\xi = -1$		$\xi = 0$		$\xi = +1$	
M_{χ}	Λ [TeV]	$\sigma_{p\chi}$ [cm ²]	Λ [TeV]	$\sigma_{p\chi}$ [cm ²]	Λ [TeV]	$\sigma_{p\chi}$ [cm ²]
1	1.02	1.3×10^{-40}	0.72	2.3×10^{-40}	0.32	1.5×10^{-39}
10	0.10	4.4×10^{-40}	0.72	7.3×10^{-40}	0.32	4.6×10^{-39}
300	0.85	1.0×10^{-39}	0.60	1.7×10^{-39}	0.26	1.1×10^{-38}
1000	0.33	4.4×10^{-38}	0.24	7.6×10^{-38}	0.11	3.6×10^{-37}

cross section enhancement when interference is constructive : $\xi = -1$



		$\xi = -1$		$\xi = 0$		$\xi = +1$	
Γ	M_{χ}	Λ [TeV]	$\sigma_{p\chi}$ [cm ²]	Λ [TeV]	$\sigma_{p\chi}$ [cm ²]	Λ [TeV]	$\sigma_{p\chi}$ [cm ²]
Γ	1	0.95	3.5×10^{-41}	0.68	5.5×10^{-40}	0.31	2.7×10^{-38}
Γ	10	0.97	1.0×10^{-40}	0.68	1.7×10^{-39}	0.31	9.0×10^{-38}
	300	0.86	1.9×10^{-40}	0.60	3.2×10^{-39}	0.30	1.3×10^{-37}
	1000	0.44	3.0×10^{-39}	0.31	4.6×10^{-38}	0.16	1.6×10^{-36}

In the case of SI interactions and ξ = -1, 0 the mono-W limits compete with mono-jets



Extra dimensions Mono-jets and mono-photons



Search for SUSY

Very light gravitino (GMSB) Stop in compressed scenario





Invisible Higgs Mono-W/Z











Final Notes

- Very successful LHC operations during the last 3 years : more than 26 fb-1 of data on tape for ATLAS/CMS (7 TeV & 8 TeV)
- Mono-X final states demonstrated to be rather sensitive channels in several searches for physics beyond SM including
 - Dark Matter, Extra Dimensions, SUSY, Higgs...
- Within the effective lagrangian framework the LHC DM searches are rather competitive for low WIMP masses
 - The validity of the EFT approach, the proper comparison with the results from direct detection experiments, the use of simplified models with light mediators, and the Higgs portal, are hot topics in an active field.
- Searches continue with the full 2012 dataset but a new discovery might eventually require more energy and more data (coming in 2015)



"Just checking."

..and more data bring new things and maybe a direct access to DM



Final Notes

More energy and more data !

El LHC will almost double the centre-of-mass energy in 2015 8 TeV → 13 TeV

We expect a significant data sample already in 2015 (10-30 fb⁻¹..my guess)





Ready for a new discovery ?

Backup Slides



pp collisions at 7 & 8 TeV





Weak scale for $\chi\chi$ annihilation cross section

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Mono-W (90% CL limits) a



CMS-PAS-EXO-13-004



WIMP-WIMP annihilation



$$\begin{aligned} \sigma_V v_{\rm rel} &= \frac{1}{16\pi\Lambda^4} \sum_q \sqrt{1 - \frac{m_q^2}{m_\chi^2}} \left(24(2m_\chi^2 + m_q^2) + \frac{8m_\chi^4 - 4m_\chi^2 m_q^2 + 5m_q^4}{m_\chi^2 - m_q^2} v_{\rm rel}^2 \right), \\ \sigma_A v_{\rm rel} &= \frac{1}{16\pi\Lambda^4} \sum_q \sqrt{1 - \frac{m_q^2}{m_\chi^2}} \left(24m_q^2 + \frac{8m_\chi^4 - 22m_\chi^2 m_q^2 + 17m_q^4}{m_\chi^2 - m_q^2} v_{\rm rel}^2 \right). \end{aligned}$$

ATLAS √s = 7 TeV, 4.7 fb⁻¹, 95%CL Annihilation rate $\langle \sigma v \rangle$ for $\chi \overline{\chi} \rightarrow qq$ [cm³ / s] 10⁻¹⁹ $2 \times (\text{ Fermi-LAT dSphs } (\chi \chi)_{Majorana} \rightarrow b\overline{b})$ 10⁻²⁰ D5: $q\overline{q} \rightarrow (\chi \overline{\chi})_{\text{Dirac}}$ 10⁻²¹ D8: $q\overline{q} \rightarrow (\chi \overline{\chi})_{\text{Dirac}}$ 10⁻²² -1 σ_{theory} 10⁻²³ 10⁻²⁴ 10⁻²⁵ Thermal relic value 10⁻²⁶ 10⁻²⁷ 10⁻²⁸ 10⁻²⁹ 10³ 10² 10 WIMP mass m_{γ} [GeV]

Results can be expressed in terms of limits on WIMP-WIMP annihilation cross section (assuming the interaction is dominated by a given operator)

For a given operator, WIMPS are required to have a minimum mass to meet the annihilation rate (and therefore the proper relic abundance)

Alternatively, for light WIMPS more than a single process is needed



Modest (~10%) improved with respect to 7 TeV limits (due to Backg. MC statistics limitations)

http://arxiv.org/abs/1310.8214





CMS-PAS-EXO-12-048

Unparticles

Howard Georgi, Phys Rev Lett. 98, (2007) 221601.



Monojet results also interpreted in term of unparticle production Limits are set in the scale dimension (d_u) vs mass scale (Λ_{u}) plane