

DECam Study of El Gordo

DES-LSST Workshop
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What is El Gordo?



A “Fat” Cluster at $z=0.87$



Image credit: J. Hughes, F. Menanteau, NASA

Why is “El Gordo” Interesting?

A pink elephant?

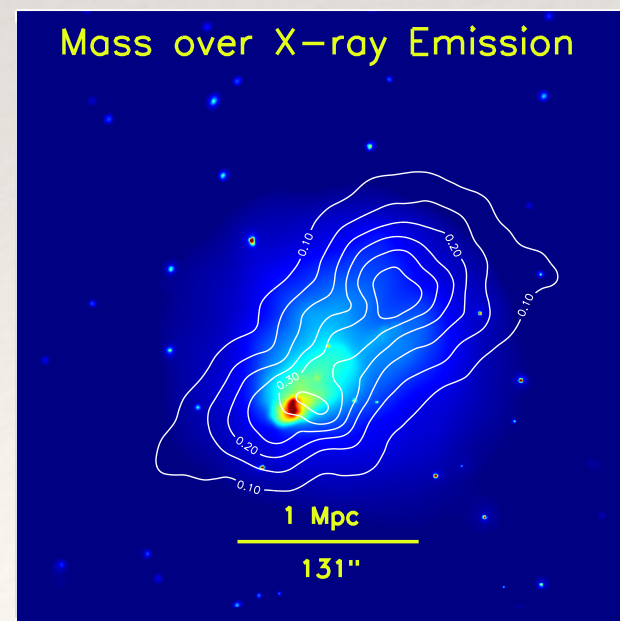
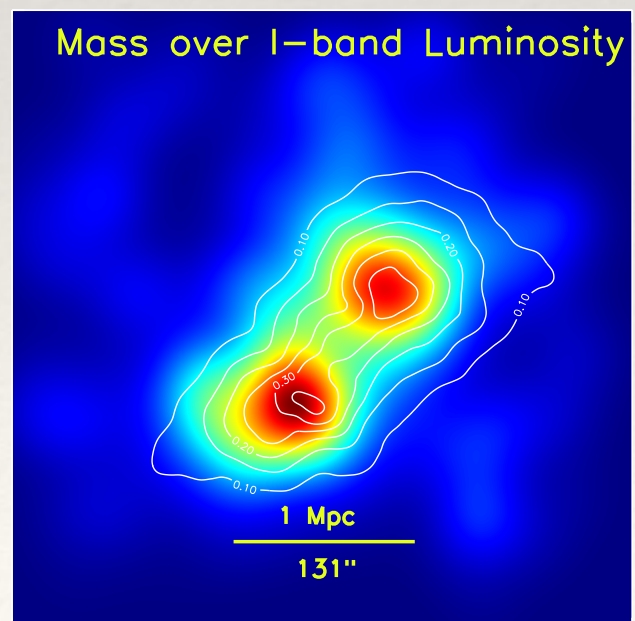
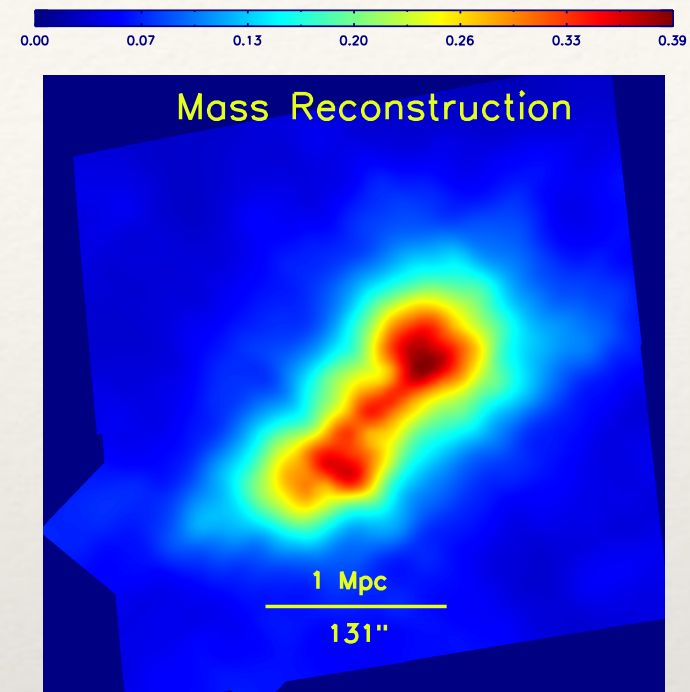
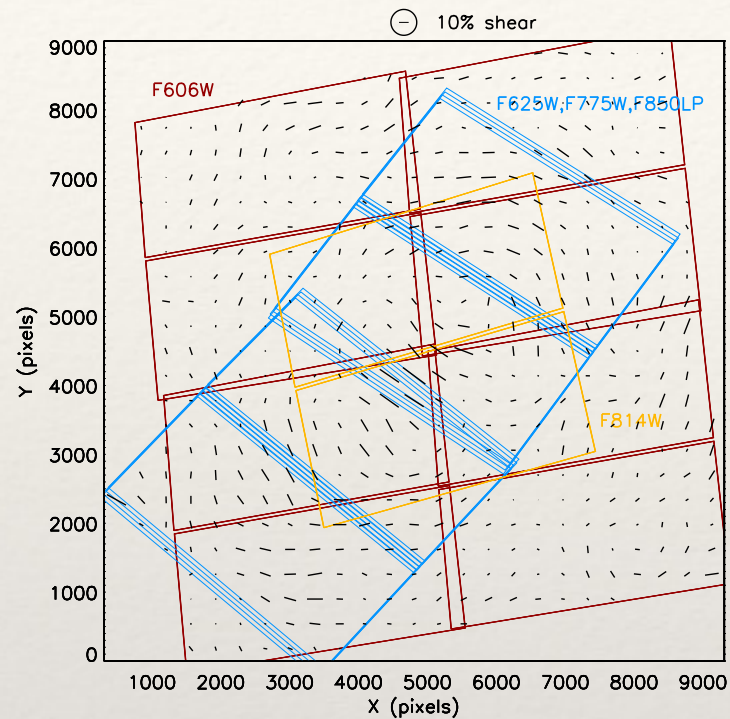


Another “bullet”?

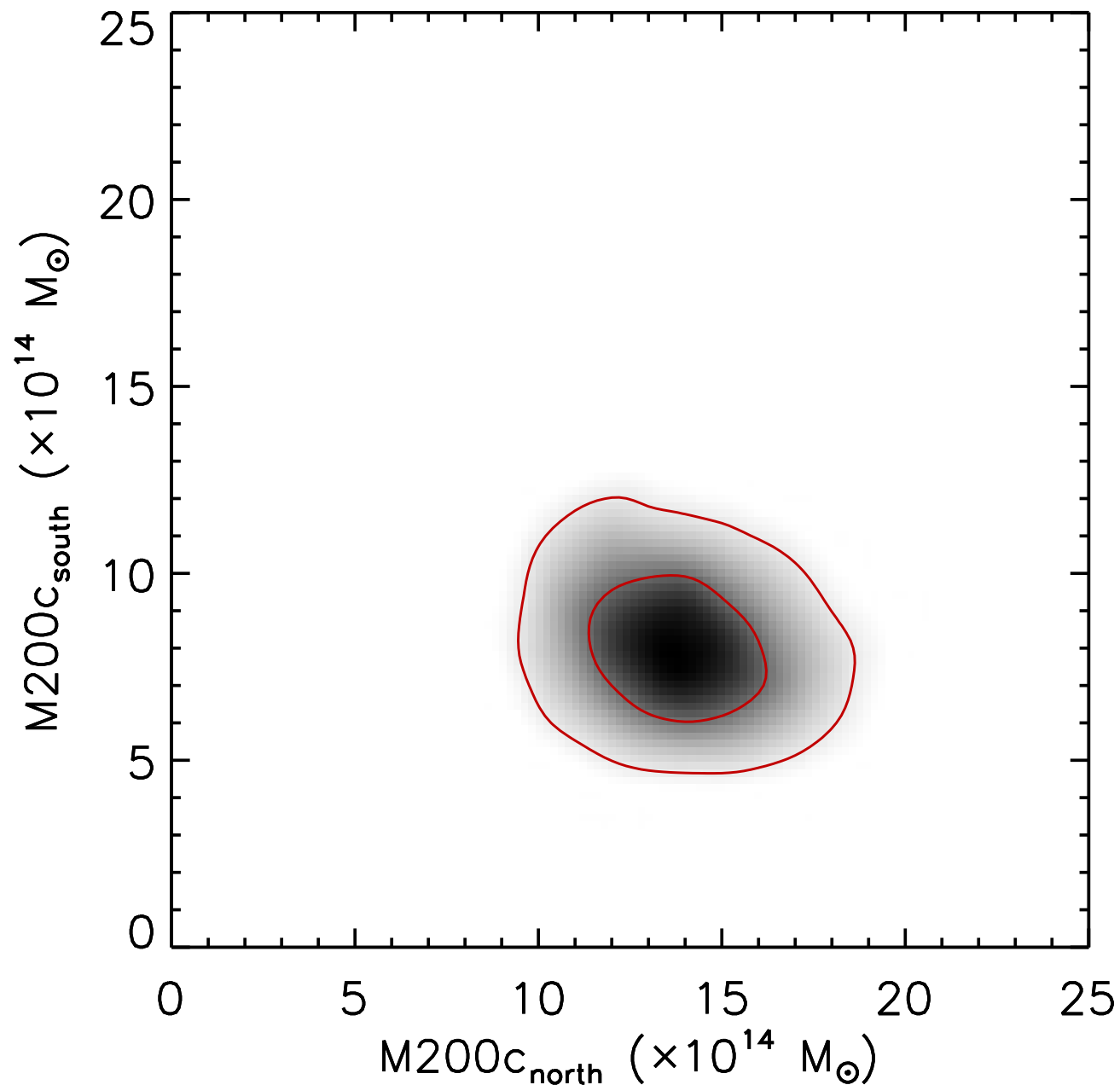


We need to know where and how much dark matter is.

HST Weak-Lensing



How Fat is El Gordo?



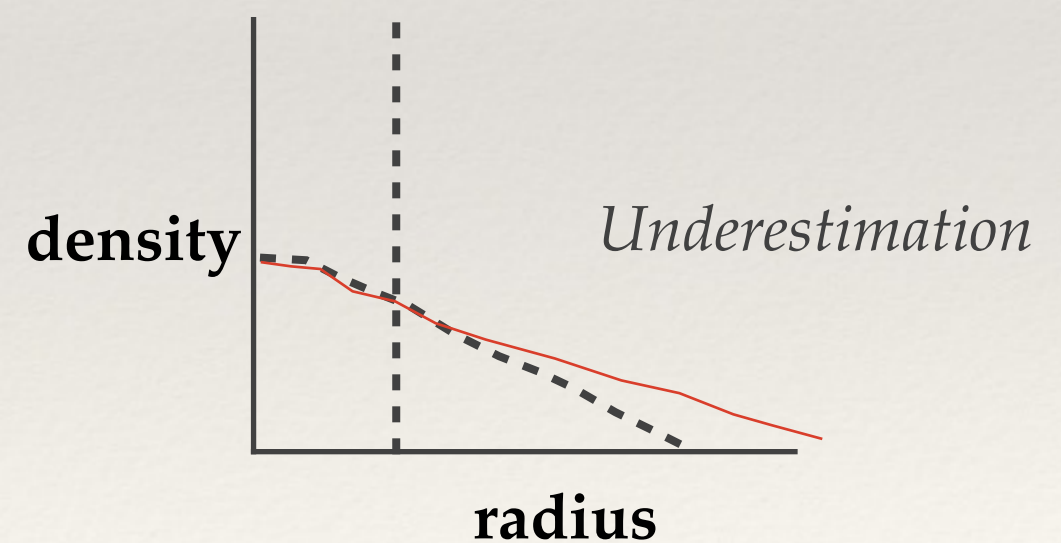
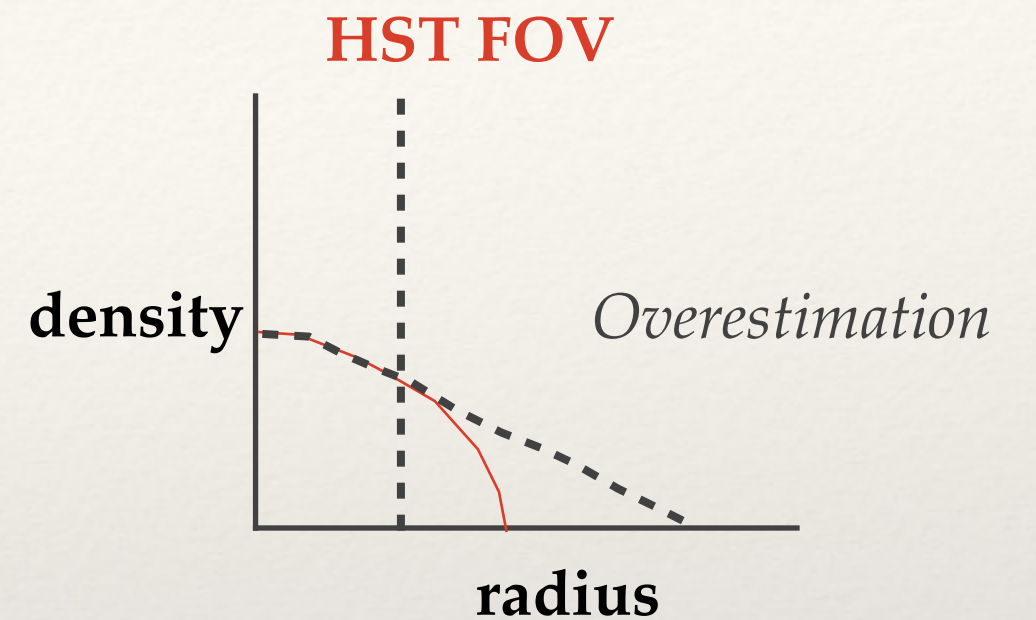
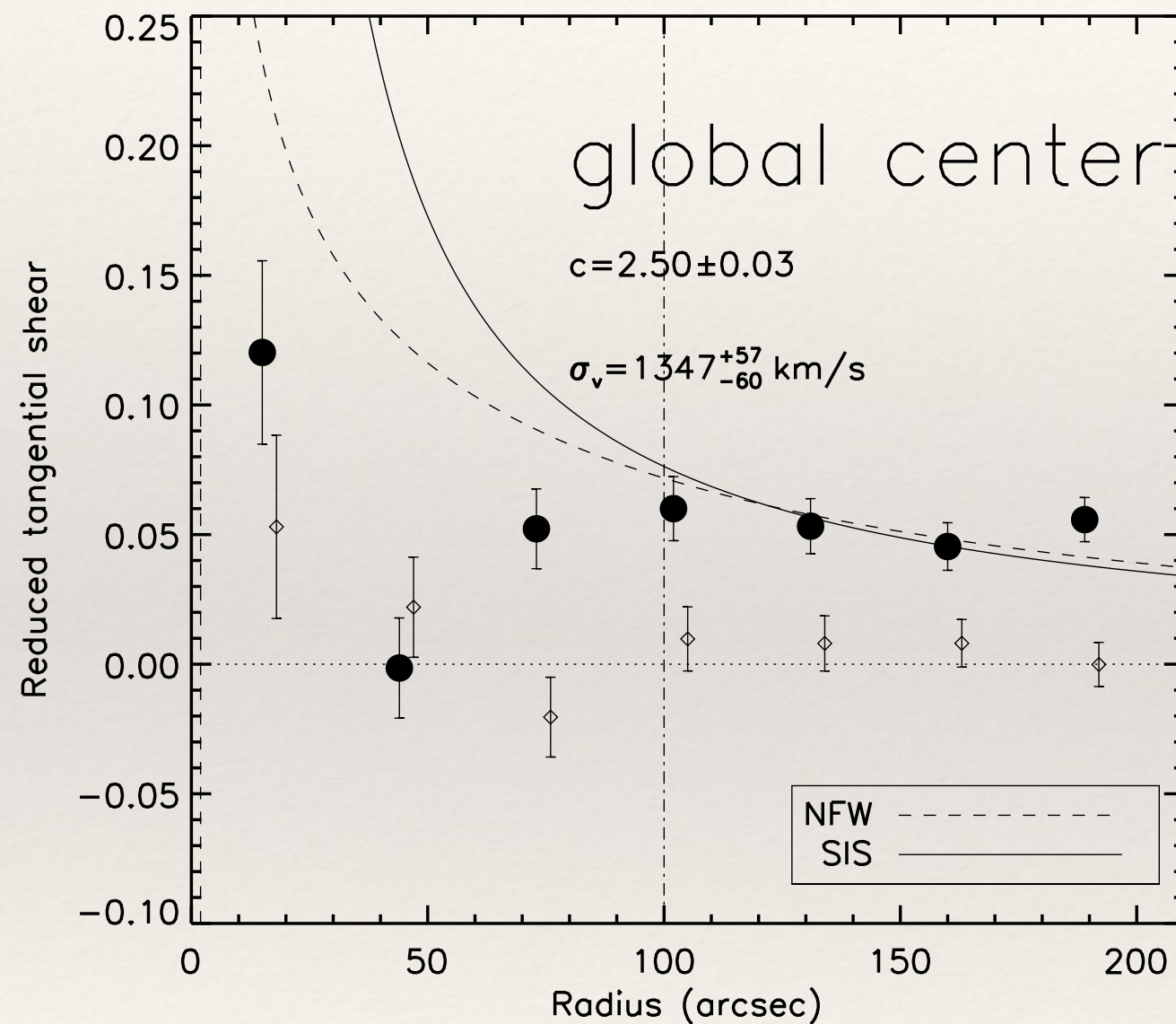
NW Halo: $(1.6 \pm 0.2) \times 10^{15} M_{\odot}$

SW Halo: $(0.9 \pm 0.2) \times 10^{15} M_{\odot}$

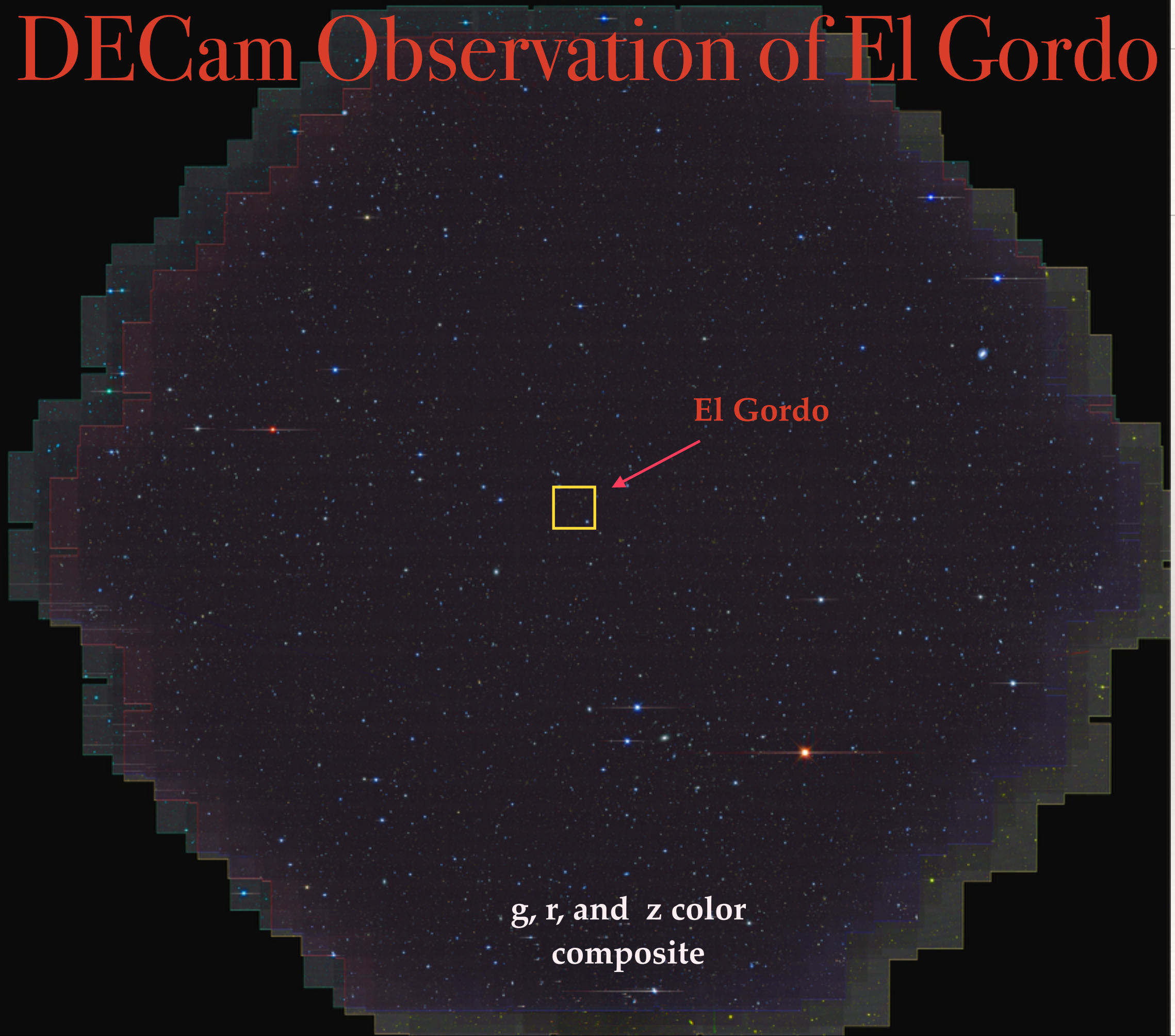
Together: $(3.1 \pm 0.6) \times 10^{15} M_{\odot}$

**Consistent with velocity
dispersion and X-ray
temperature**

Limitation of HST WL



DECam Observation of El Gordo

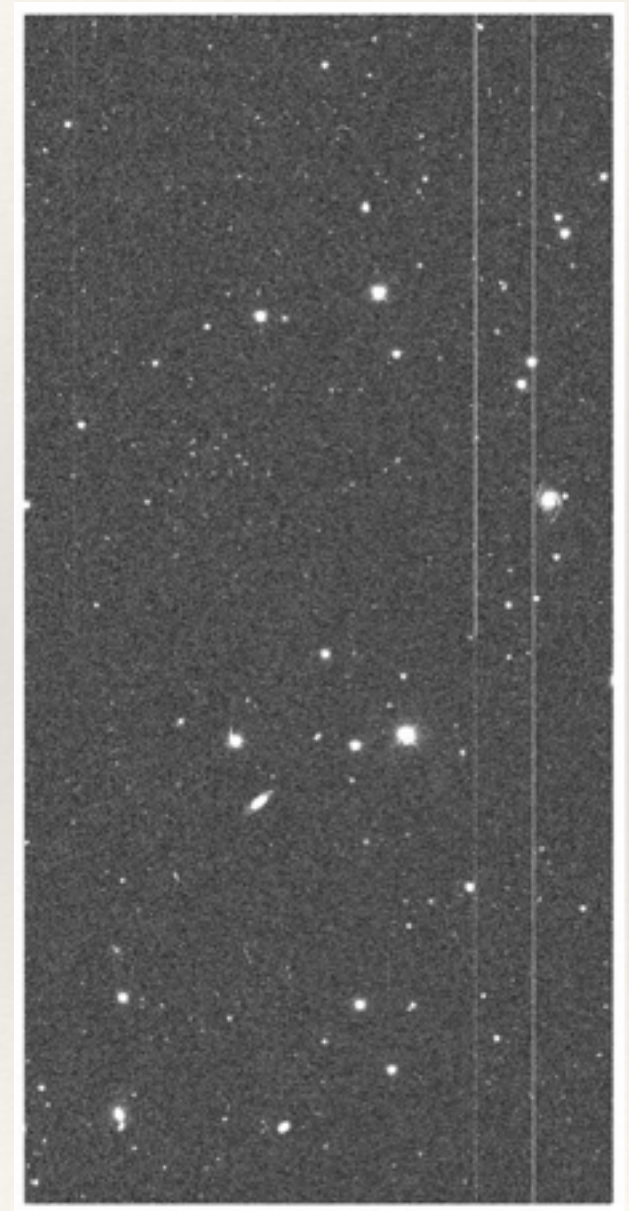
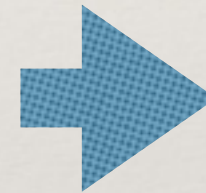
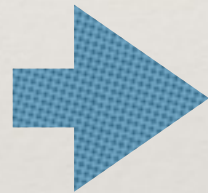
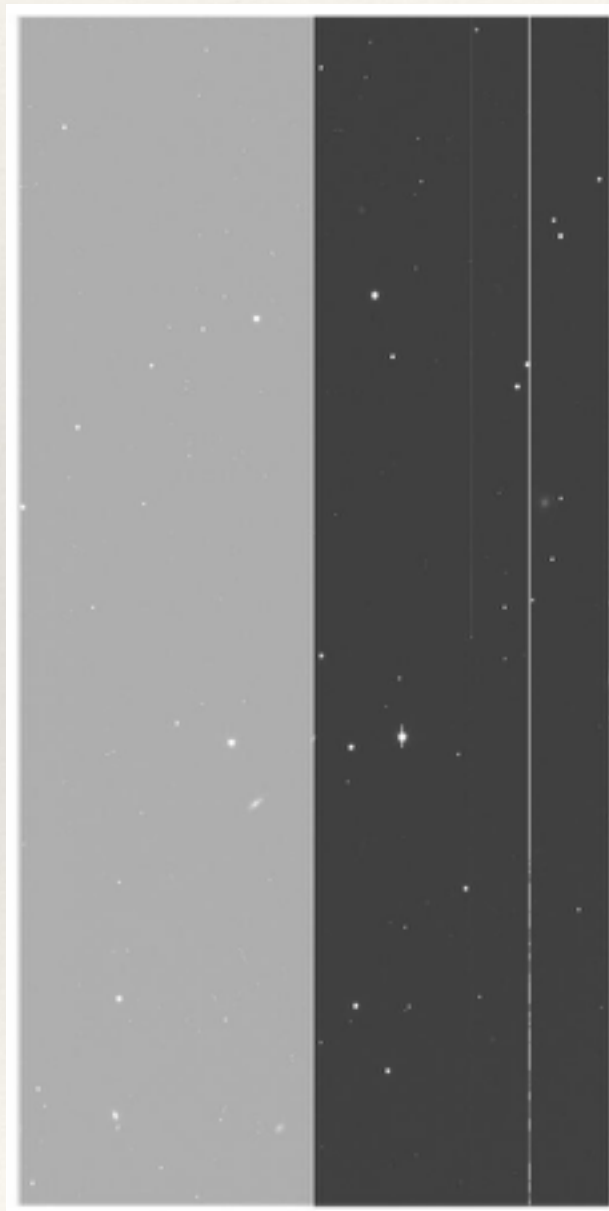




DECam Data of El Gordo

	Exp Time	Seeing
g	900 s	~1.05"
r	1800 s	~0.94"
i	900 s	~0.87"
z	900 s	~0.85"
y	1400 s	~0.84"

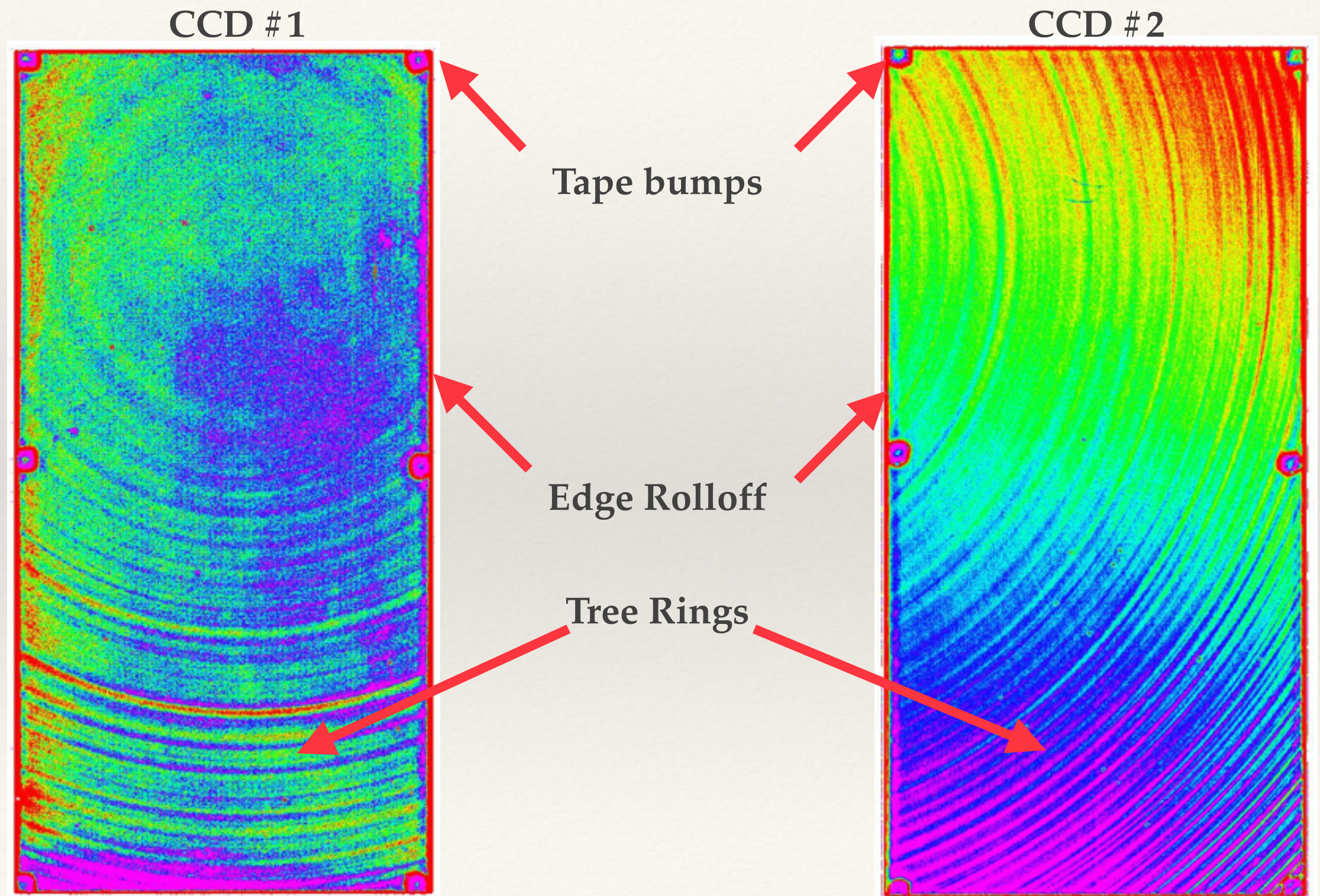
DECam Image Reduction - Bias Subtraction/Trim



Bias Subtraction

Residual Bias Subtraction

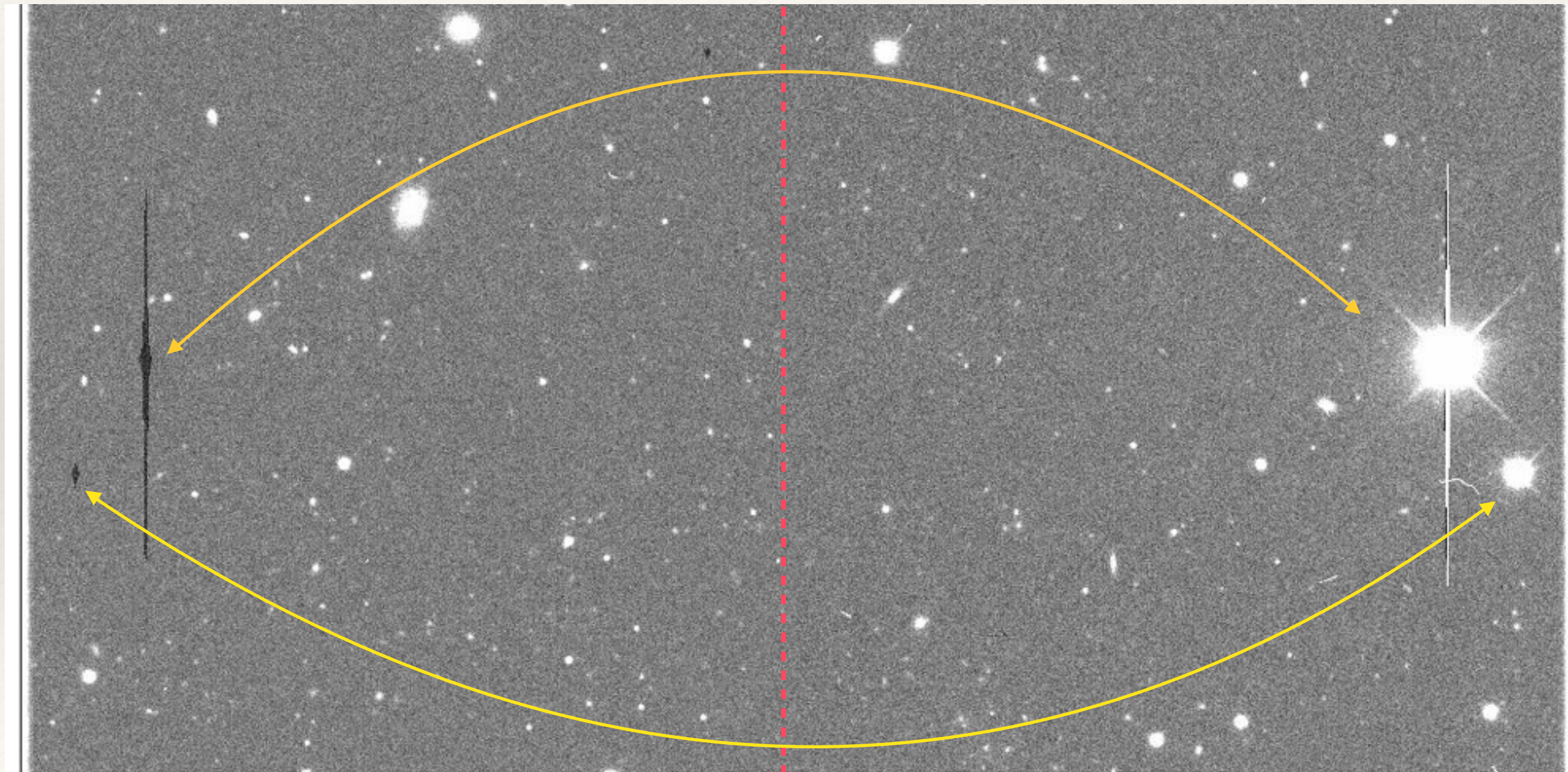
DECam Image Reduction - Sky Flat



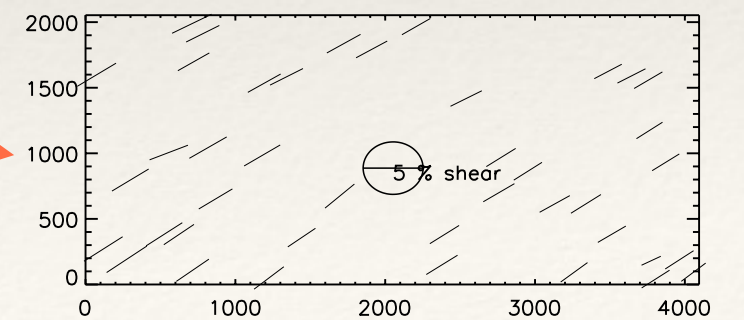
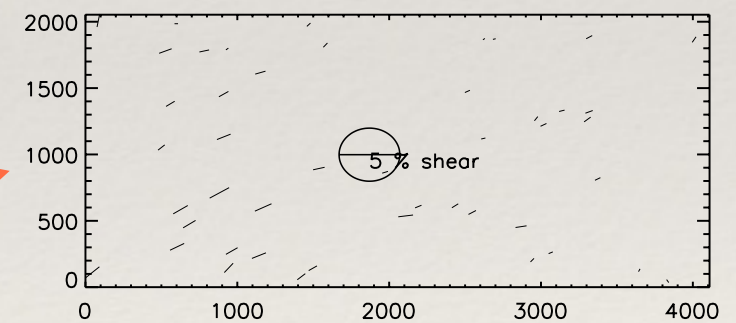
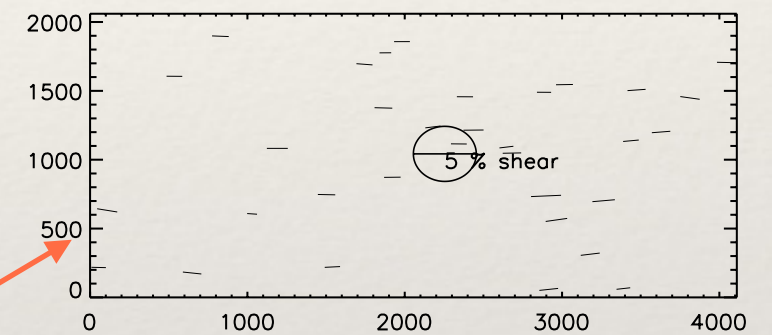
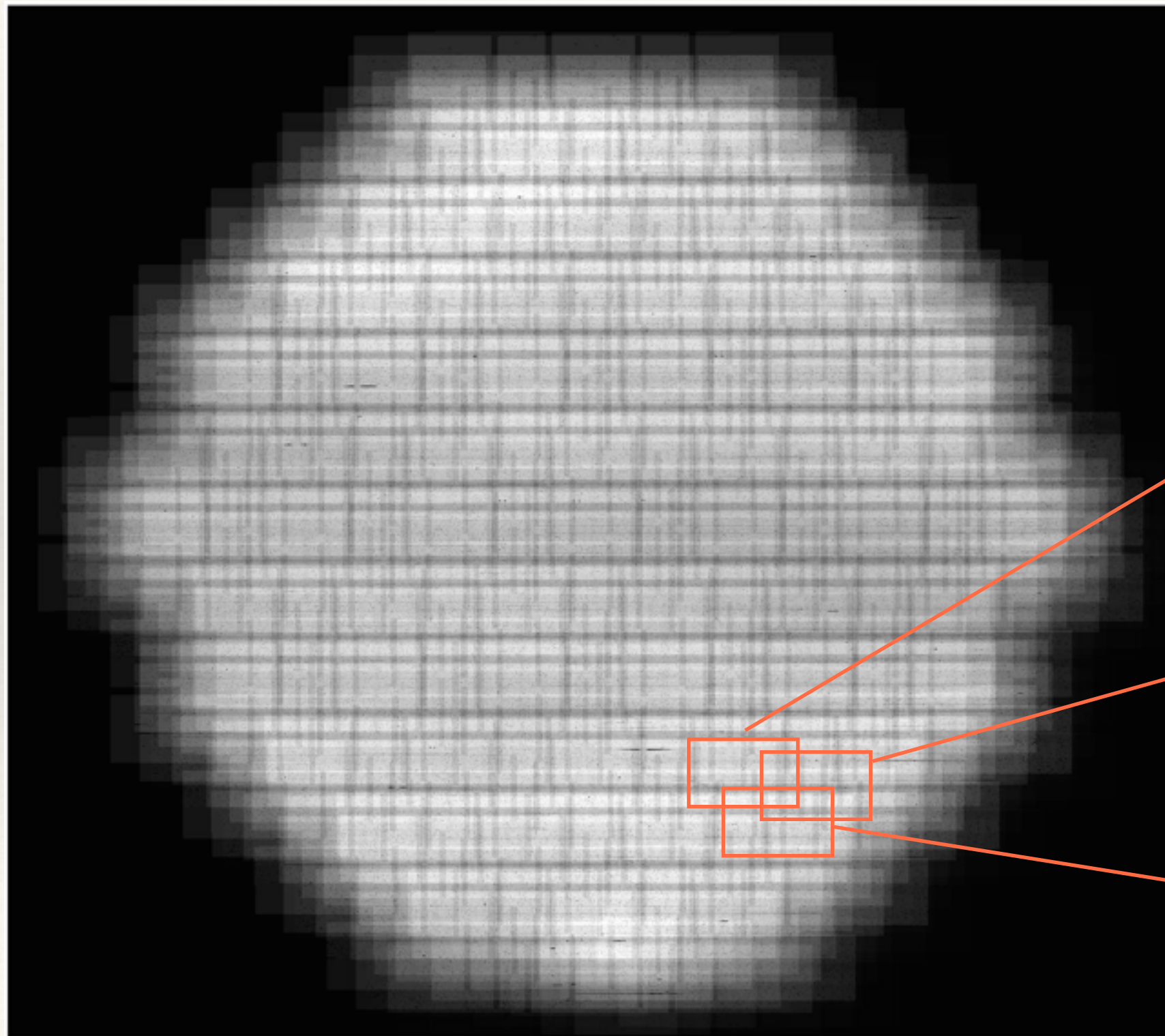
DECam Image Reduction - Distortion Correction

- Reference star catalog: 2MASS.
- Third order polynomial.
- Assume that the frames from the same CCD share identical geometric distortion. However, each frame is allowed to shear and rotate.
- Photometric calibration is done with matching stars.

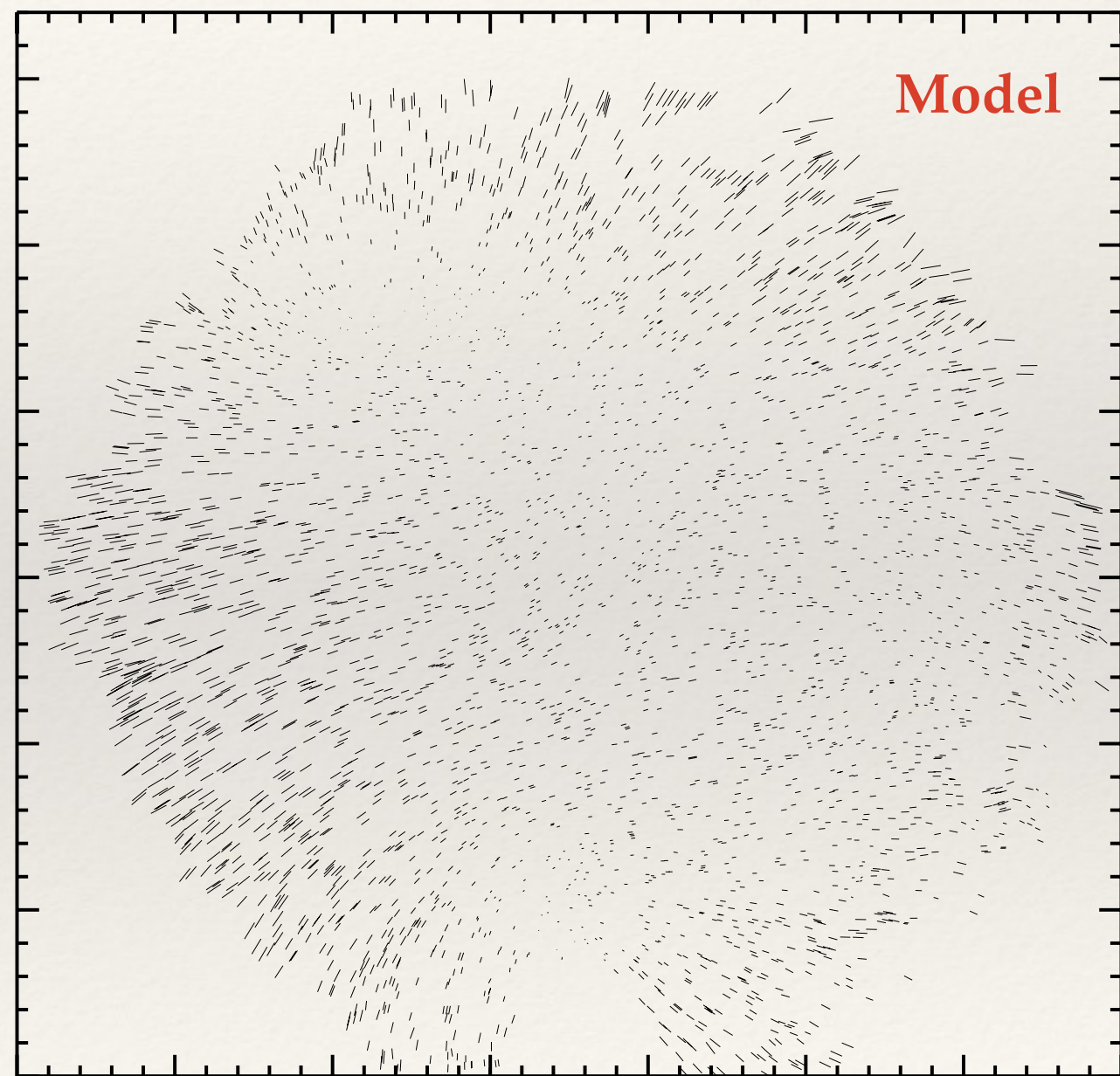
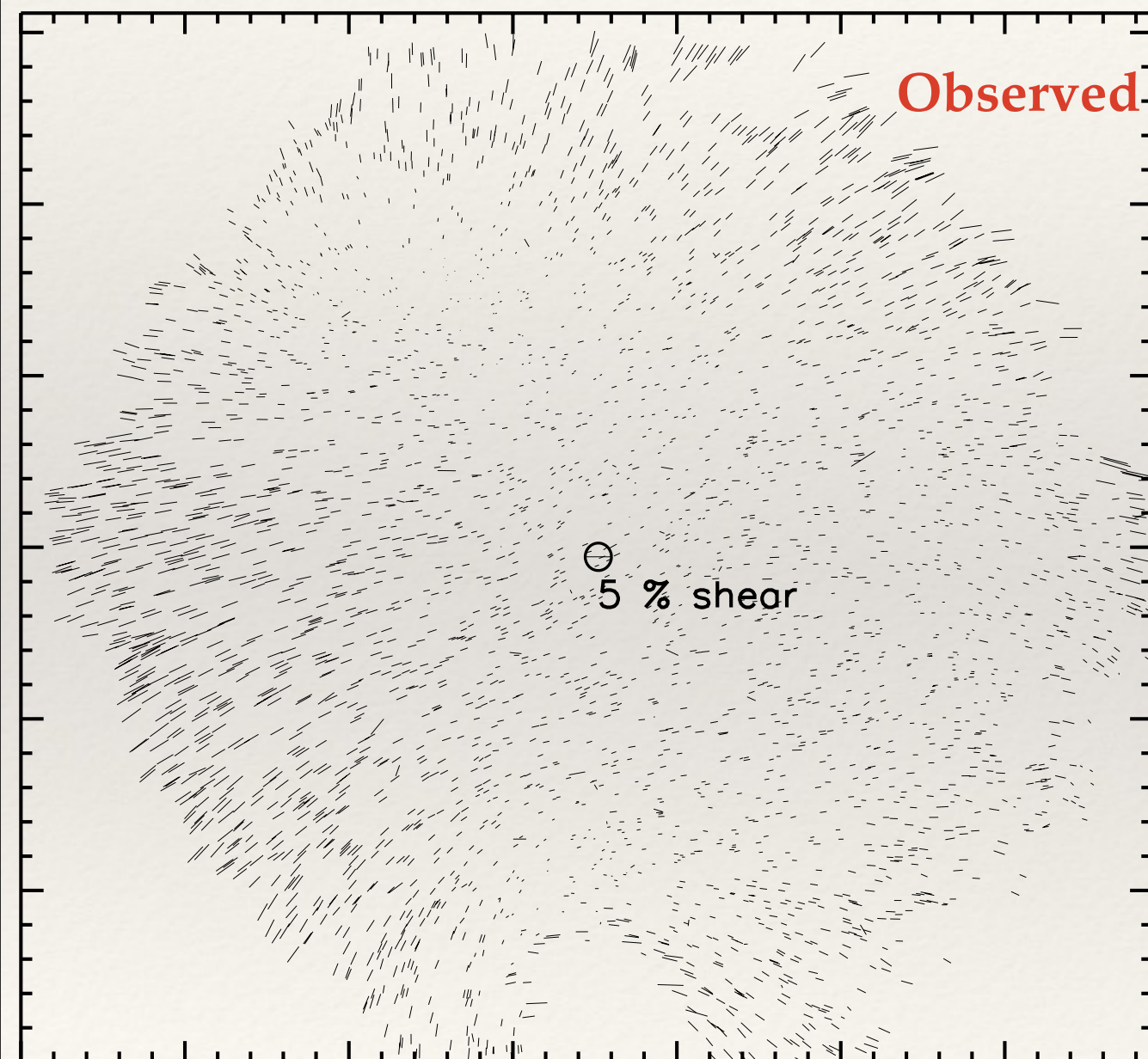
DECam Image Reduction - Crosstalks



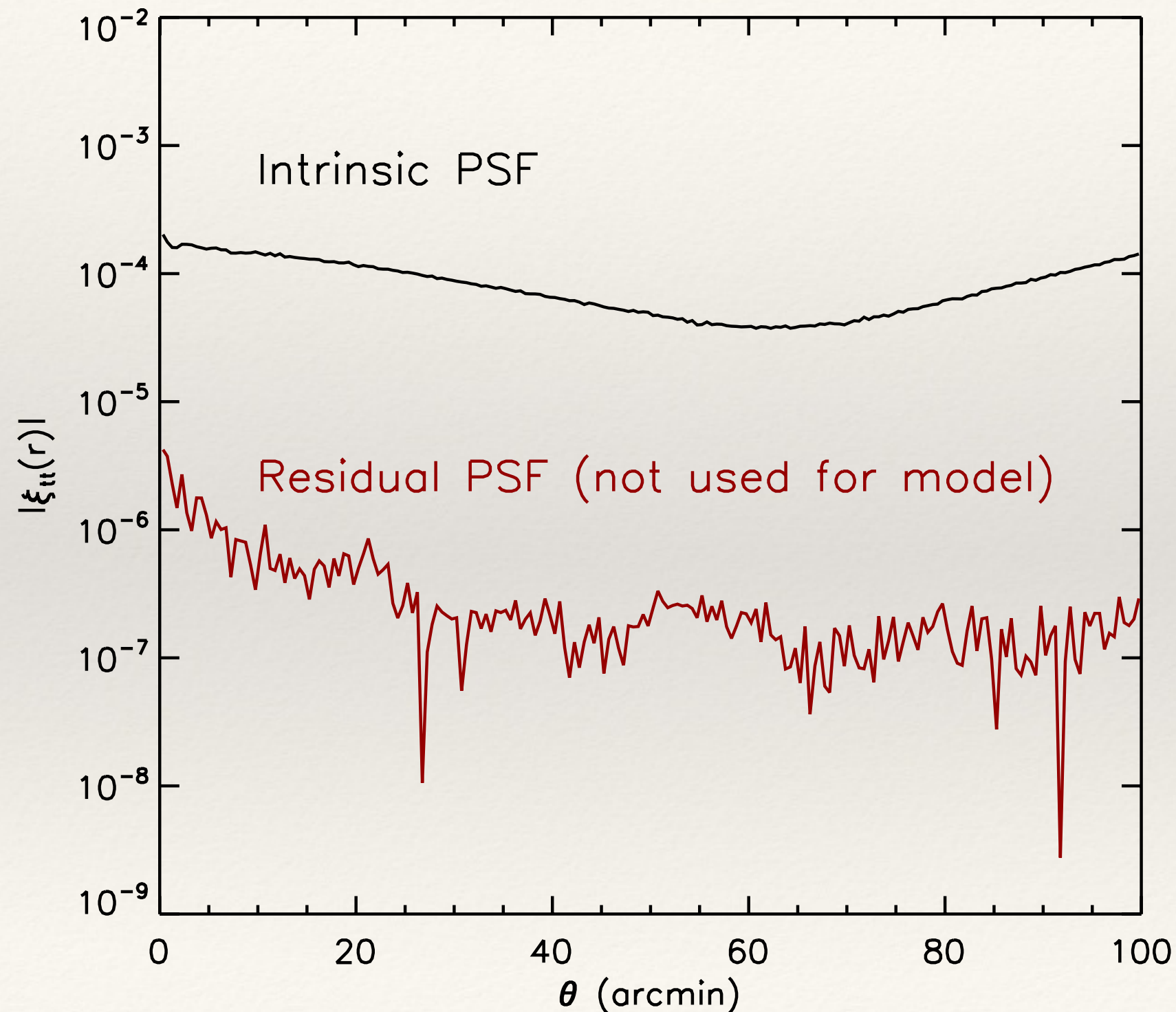
PSF Modeling Through PCA



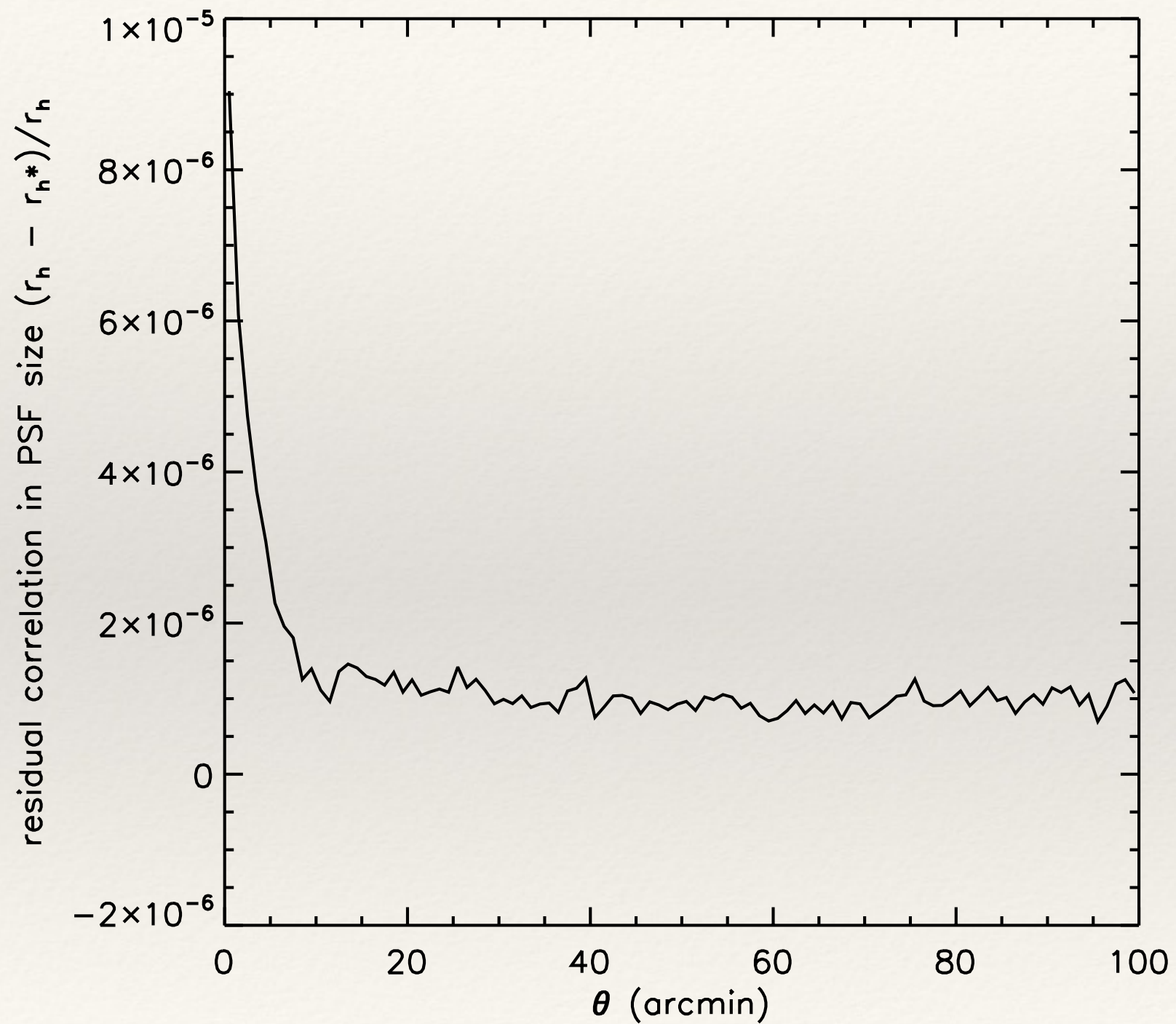
PSF Modeling - Continued



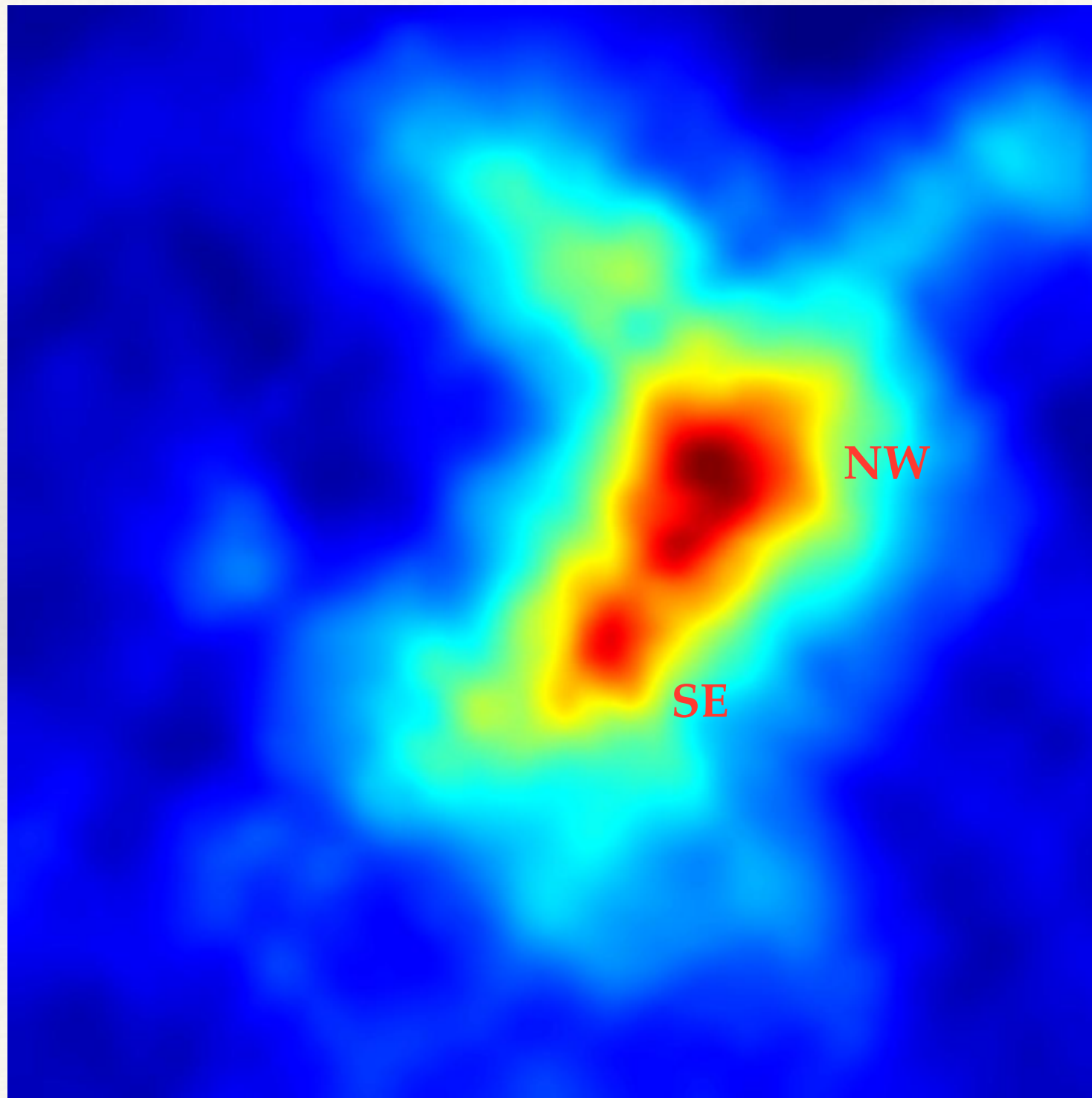
PSF Ellipticity Correlation



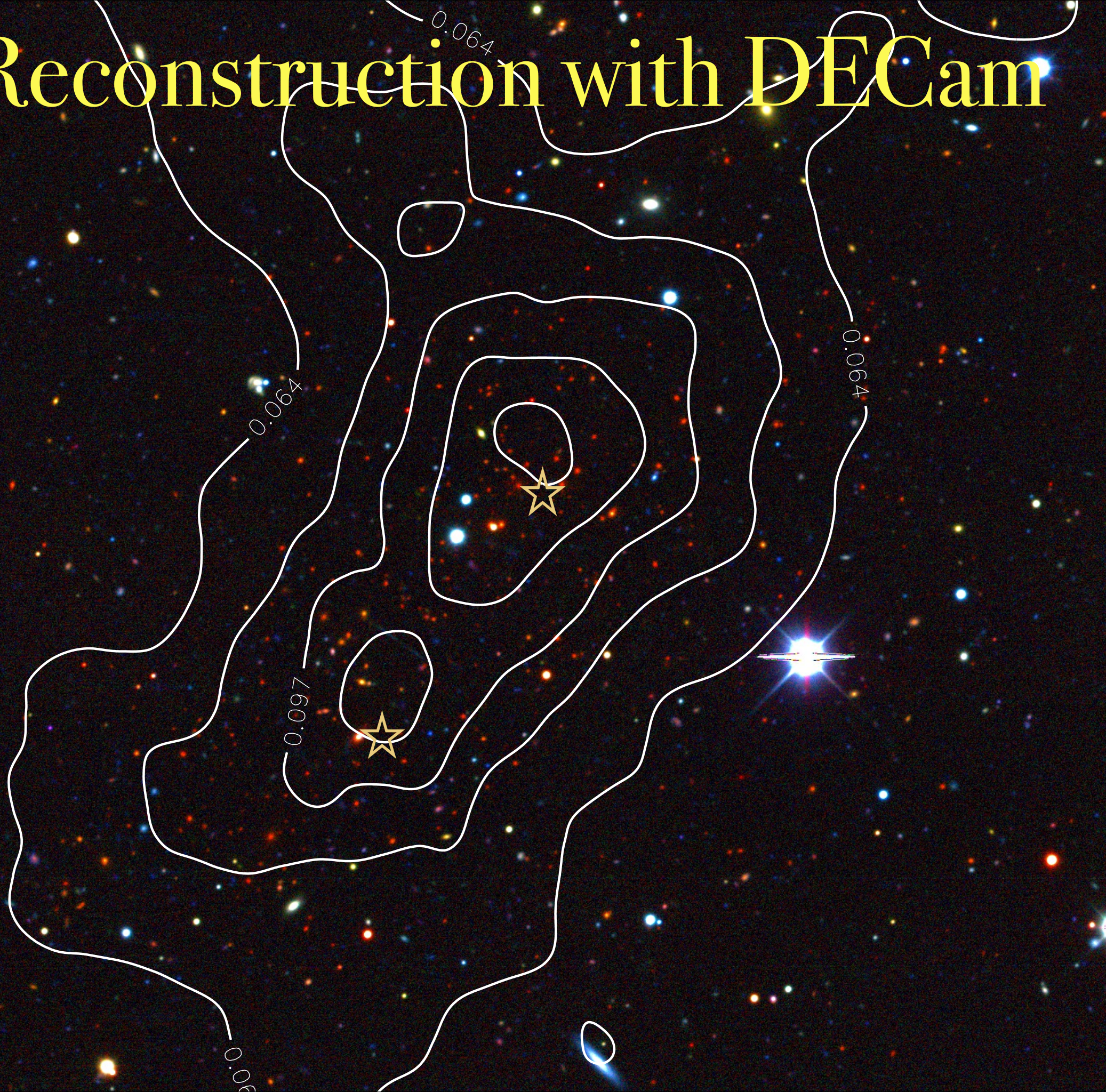
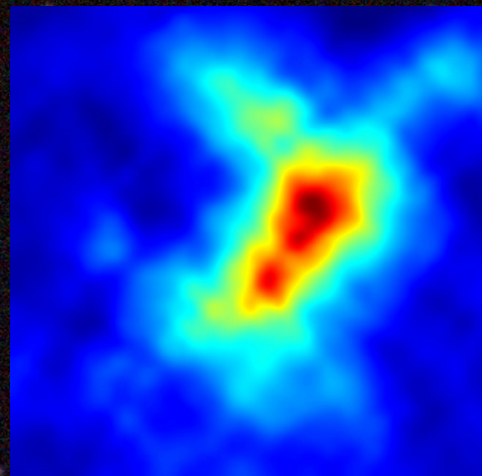
RESIDUAL PSF SIZE



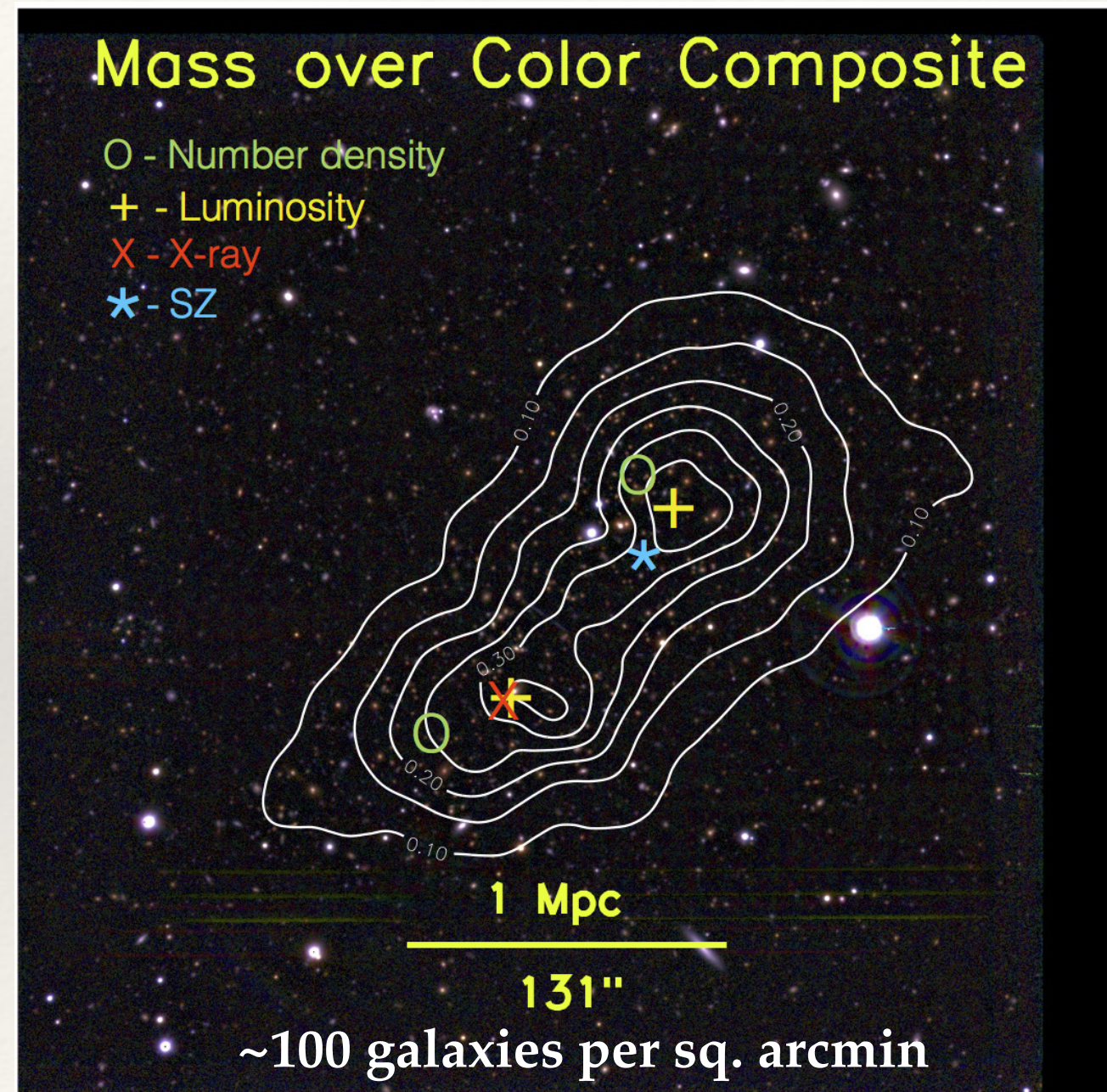
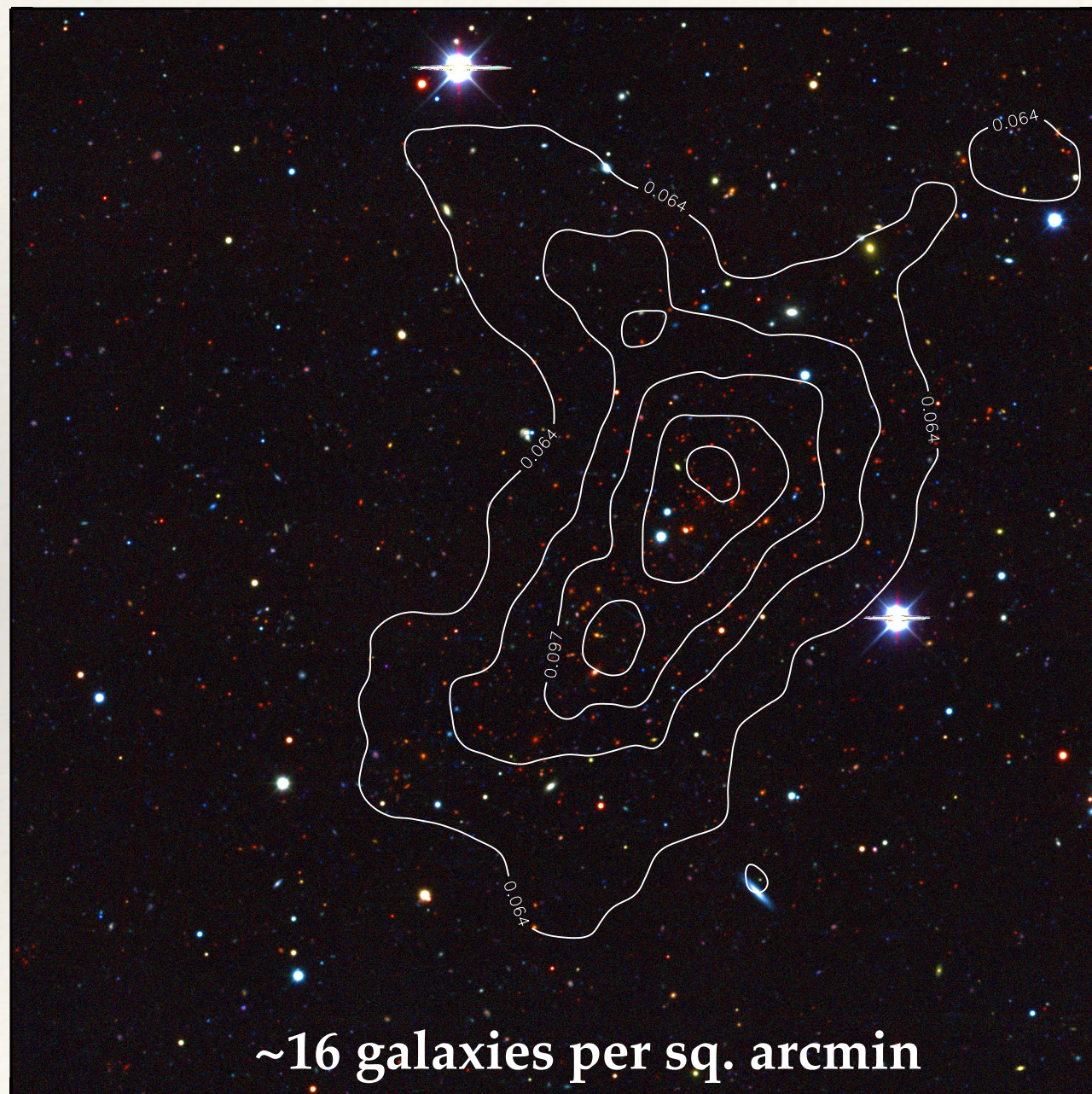
Mass Reconstruction



Mass Reconstruction with DECAM

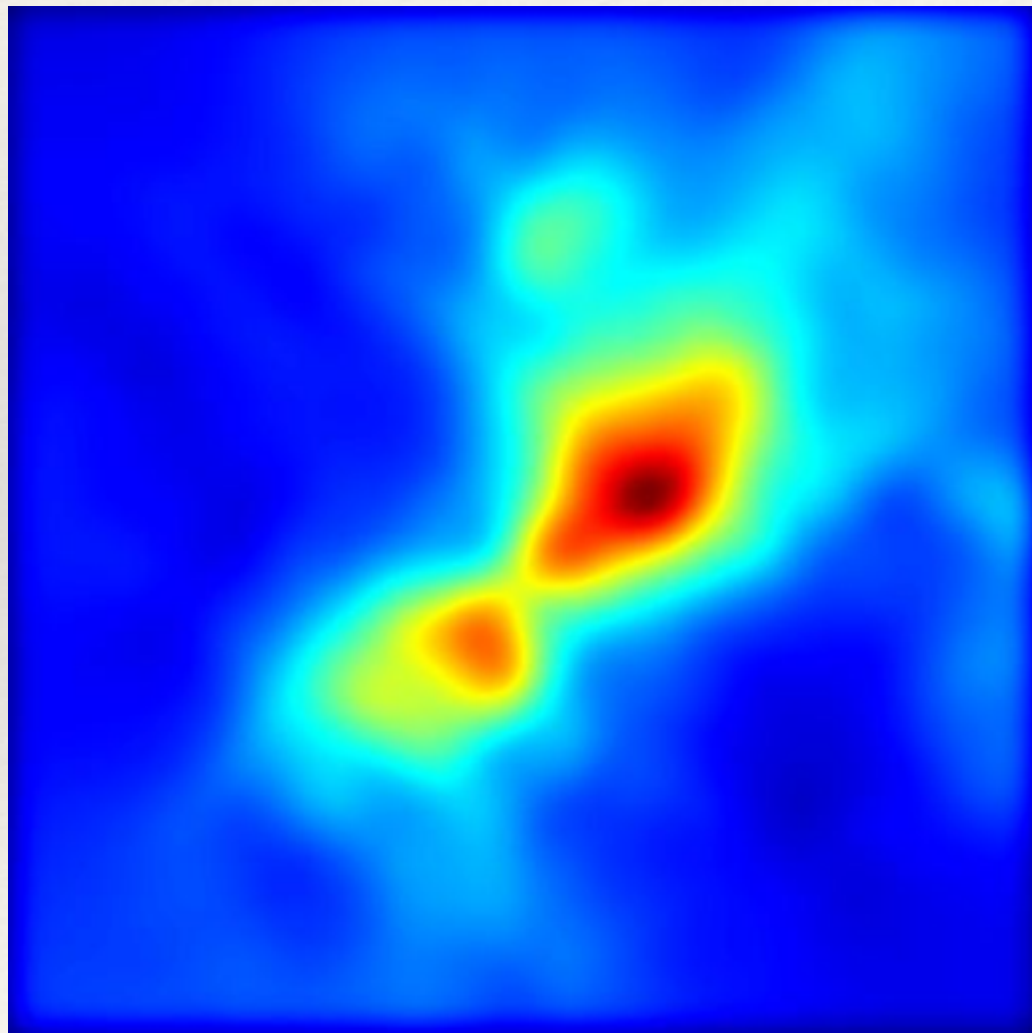


DECam vs. HST

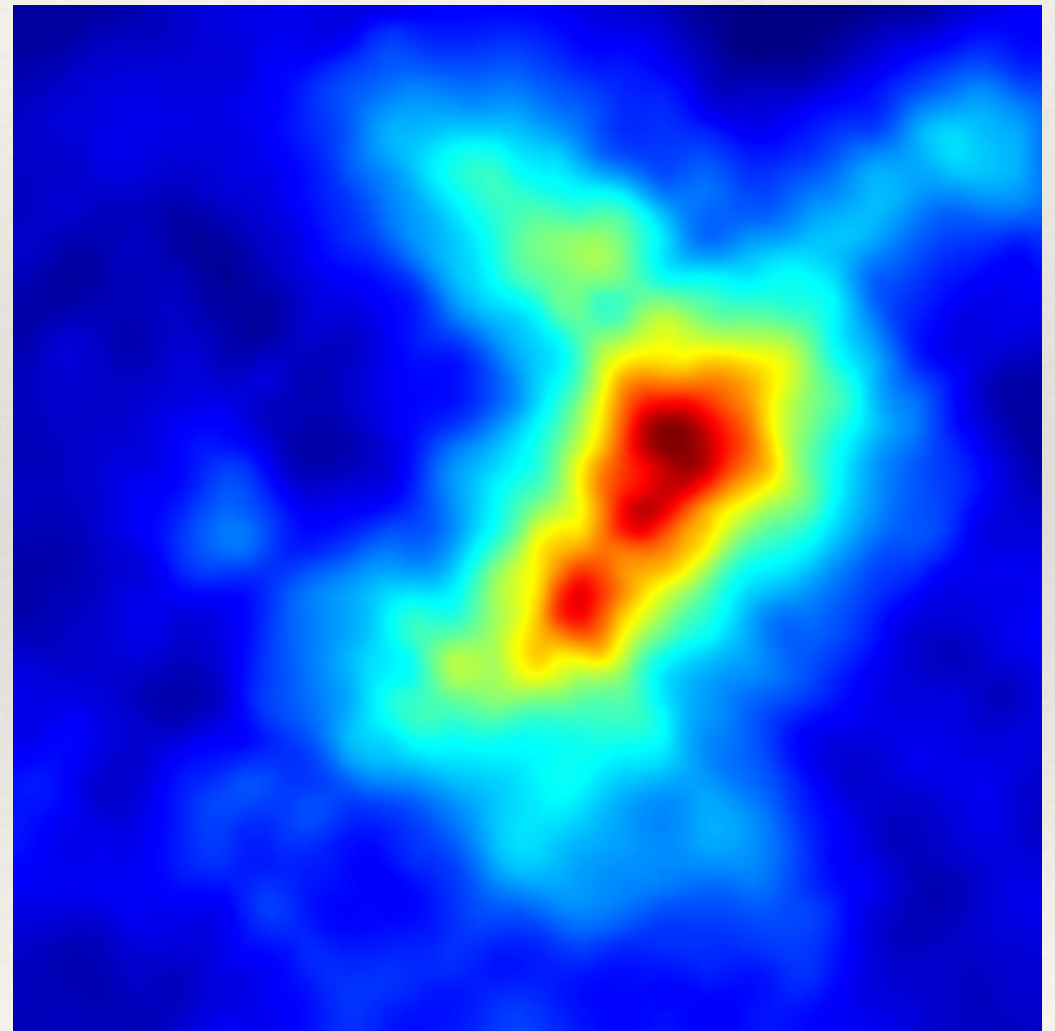


After Throwing Away ~80% of Galaxies from the HST Sources

Degraded HST
(~16 galaxies per sq. arcmin)

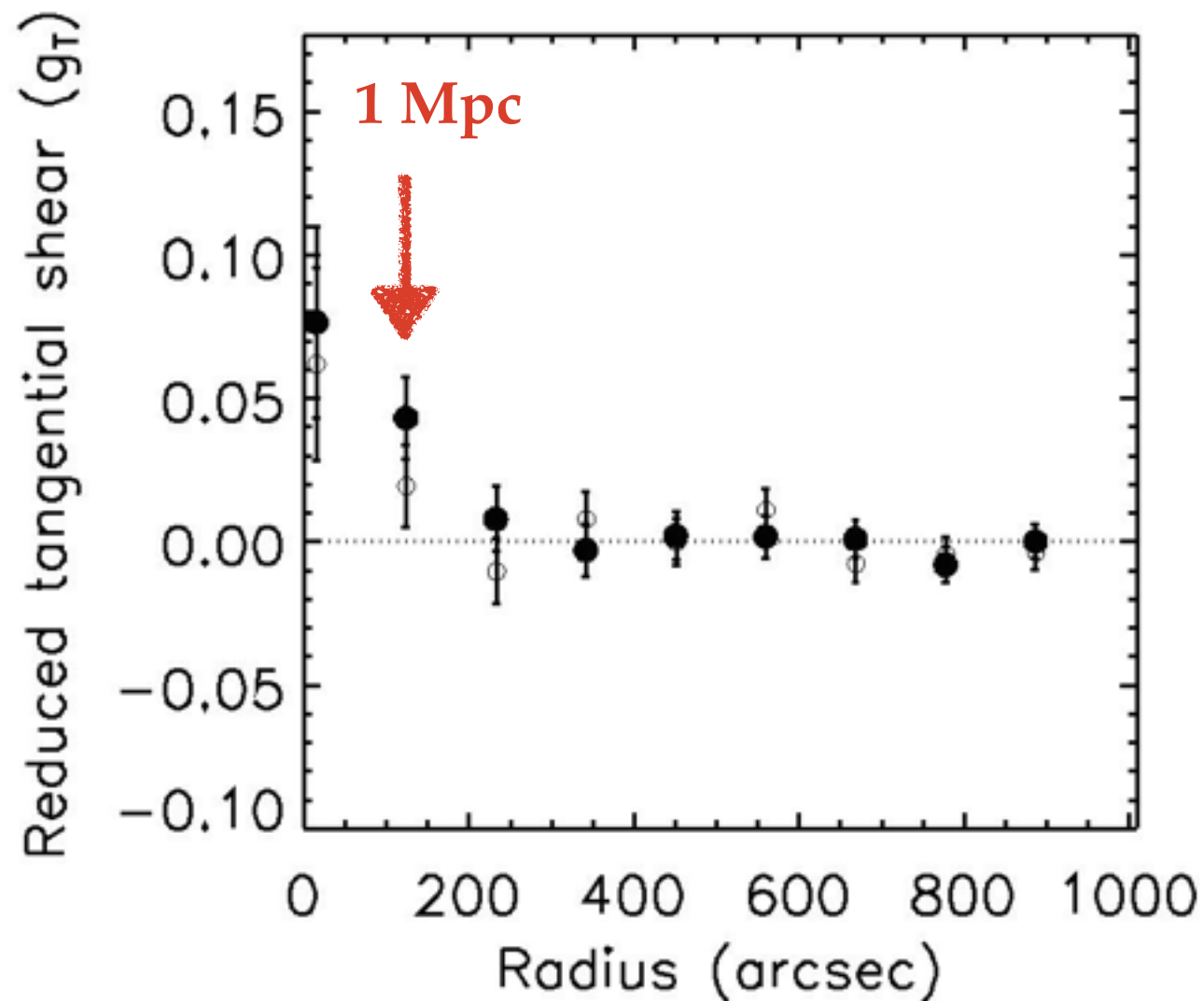


DECam

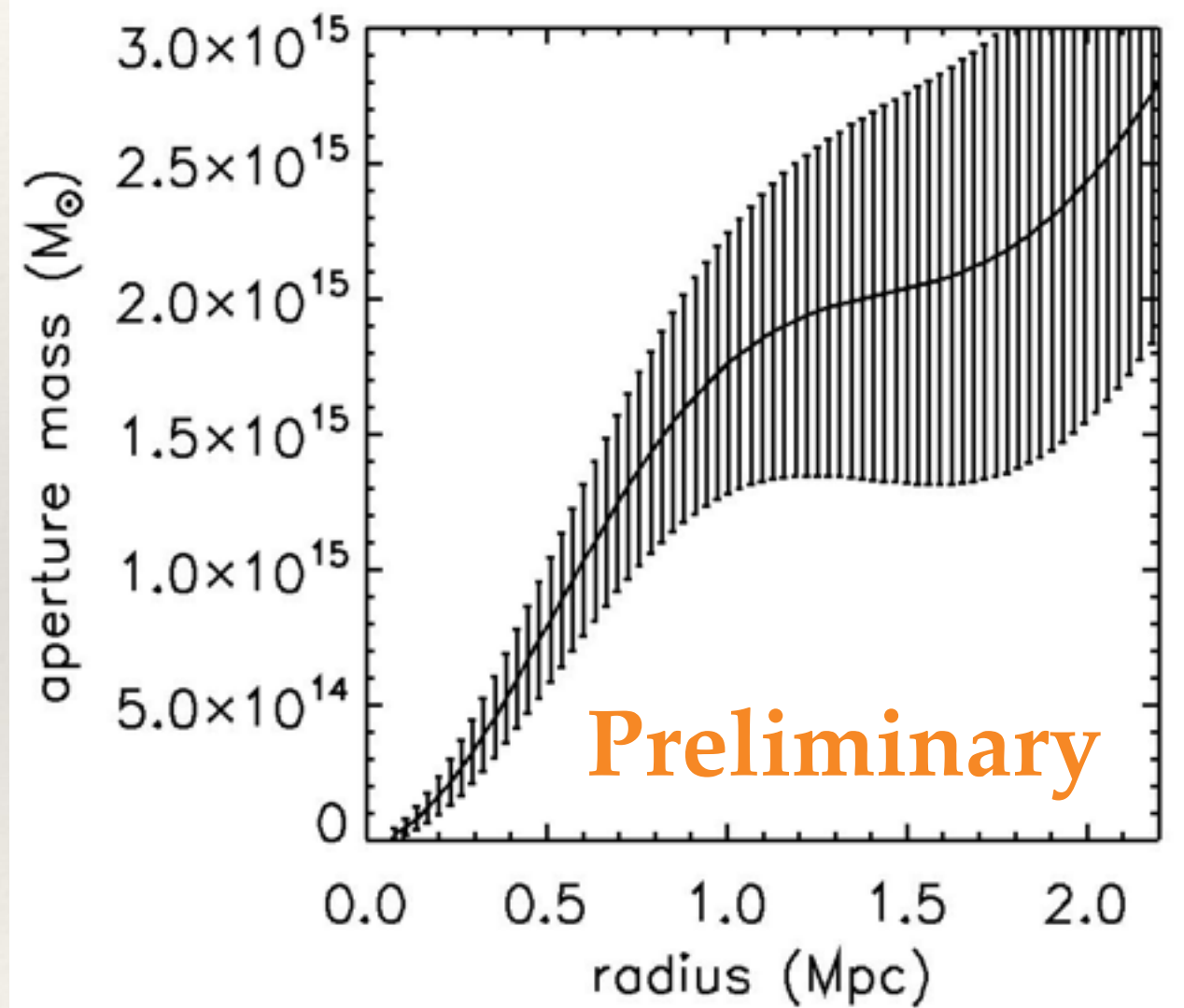


Mass of El Gordo

Tangential Shear



Aperture Mass Densitometry



Things to Address

- ❖ Depth variation due to lensing
- ❖ Astrometric effects of tree rings and tape bumps
- ❖ Crosstalks
- ❖ Shear calibration

Conclusions

- ❖ Lensing signal of El Gordo at $z=0.87$ is clearly detected by DECam.
- ❖ The DECam mass reconstruction resolves the two mass clumps of El Gordo, consistent with the HST study.
- ❖ No significant lensing signal is seen beyond $r>1$ Mpc, perhaps due to either Truncation of mass profile or lensing-induced depth variation.
- ❖ The projected mass within $r<1$ Mpc with DECam WL is consistent with the HST parametric model.