

# Future Long-baseline Neutrino Oscillations ~ View from Asia ~

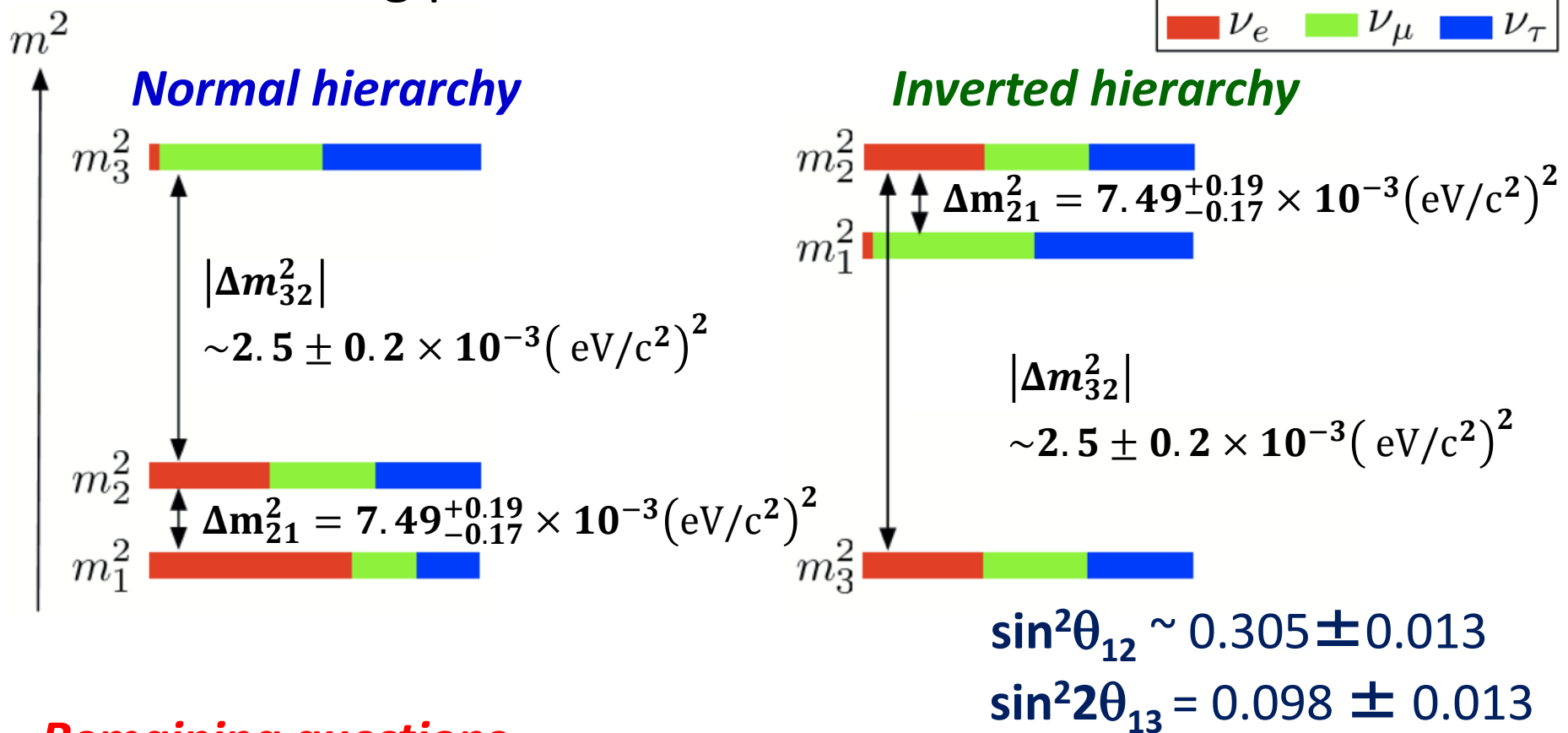
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( Kamioka, ICRR, UTokyo )

Used material in this presentation

Kindly Provided by

J-PARC accelerator group,  
KEK neutrino group,  
Hyper-Kamiokande working group, and  
MOMENT group.

# Neutrino mixing parameter measurements



## Remaining questions

1)  $\theta_{23}$  is really  $45^\circ$  or  $< 45^\circ$  or  $> 45^\circ$  ?

Current uncertainty of  $\sin^2 \theta_{23}$  is still large  $\sim 10\%$  level

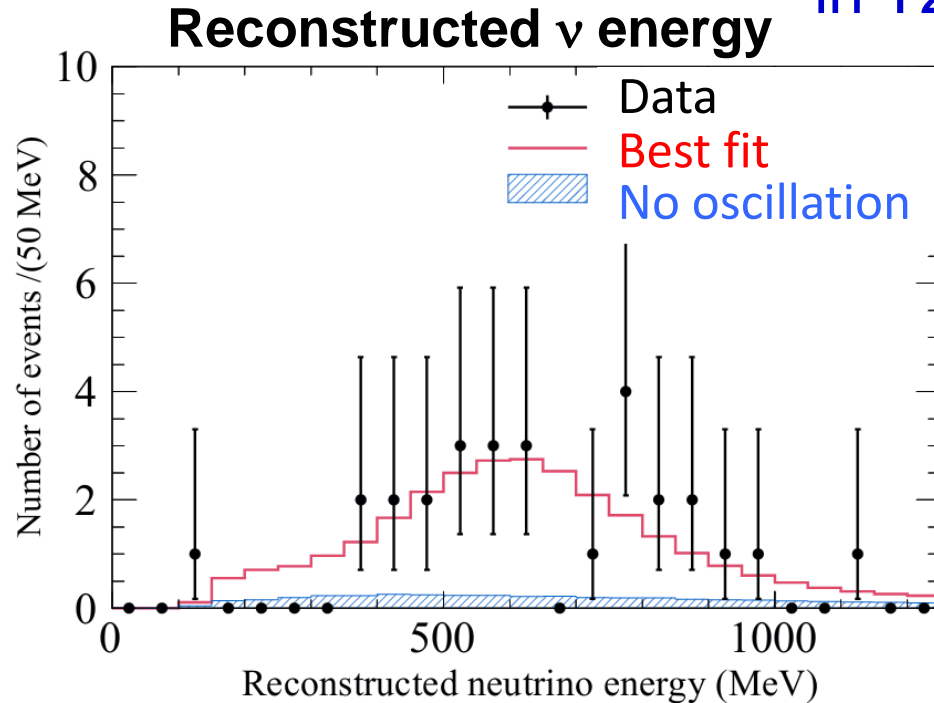
$$\sin^2 \theta_{23} = 0.514 \pm 0.055 \quad (\text{T2K 2014})$$

2) CP is violated or not (  $\delta = 0$  or not ) ?

3) Mass hierarchy  $\sim$  which is heavier ? (  $\Delta m_{32}^2 > 0$  or  $< 0$  ? )

# Recent results from Reactor & T2K

$\nu_e$  appearance clearly observed  
in T2K

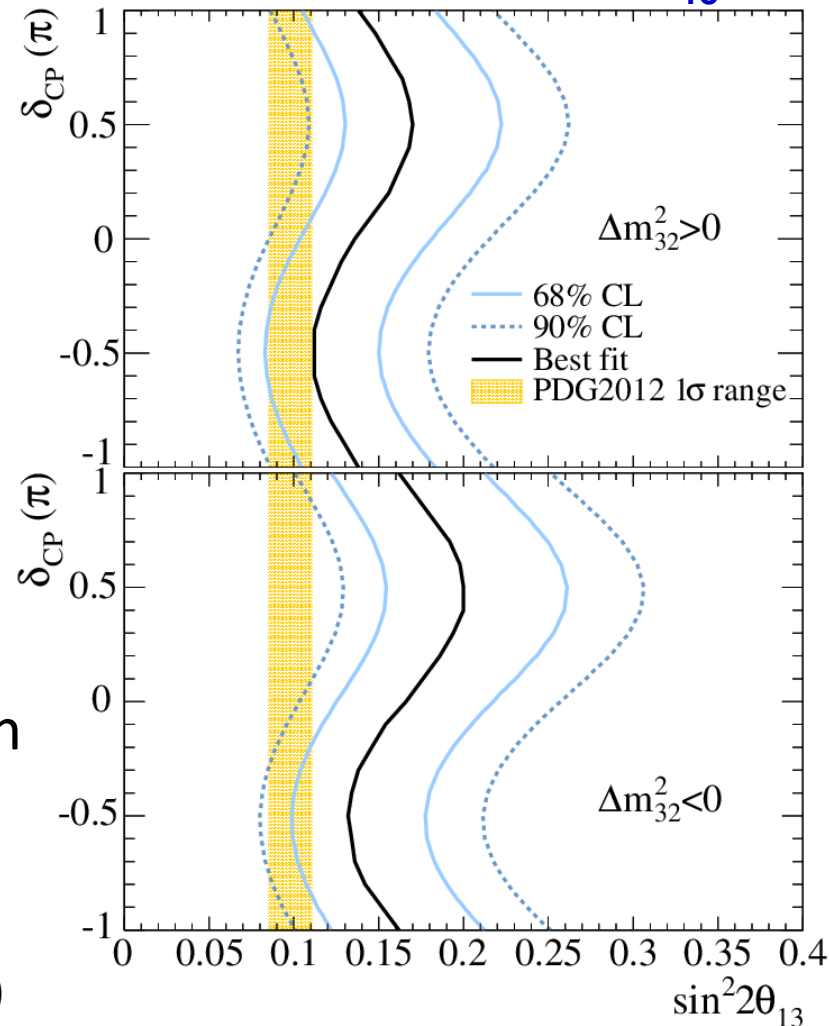


Extracted  $\sin^2 2\theta_{31}$  is slightly larger than the ones from reactor experiments

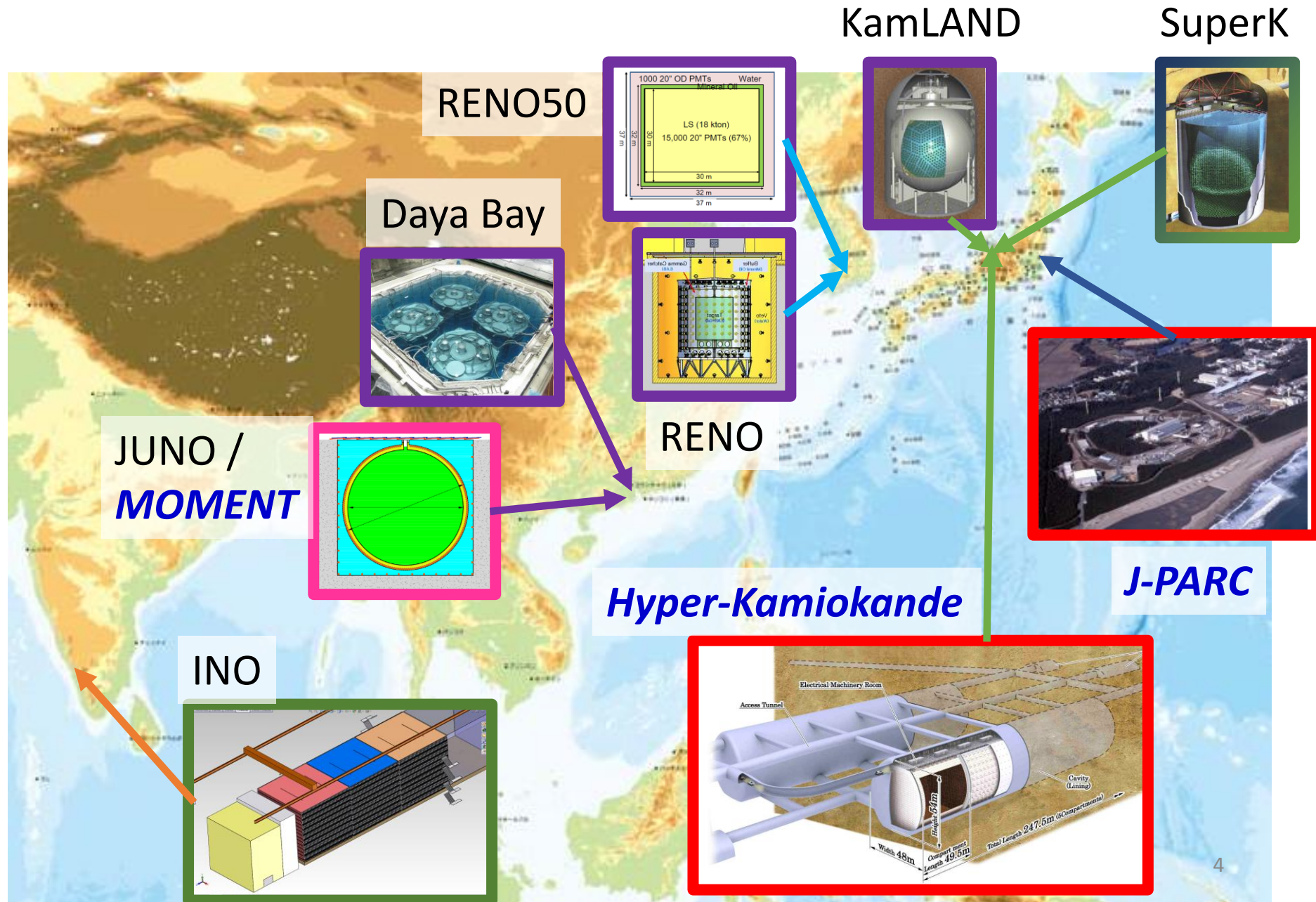
$$\sin^2 2\theta_{13} = \begin{matrix} 0.140^{+0.038}_{-0.032} & (\text{normal hierarchy}) \\ 0.170^{+0.044}_{-0.037} & (\text{inverted hierarchy}) \end{matrix}$$

**Indication of non-zero  $\delta_{CP}$ ?**

**Allowed region of  $\delta_{CP}$   
for each  $\sin^2 2\theta_{13}$**



# Neutrino experiments & related facilities in Asia



*J-PARC neutrino beam  
& Hyper-Kamiokande*



## Hyper-Kamiokande



One of the most powerful beamlines in operation  
and further intensity upgrade ( >750kW ) is undergoing.

World largest water Cherenkov detector ( fid. vol. 560 kt. )

## Powerful combination

***to search for the lepton sector CP violation!*** 6

# Hyper-Kamiokande with J-PARC neutrino beam

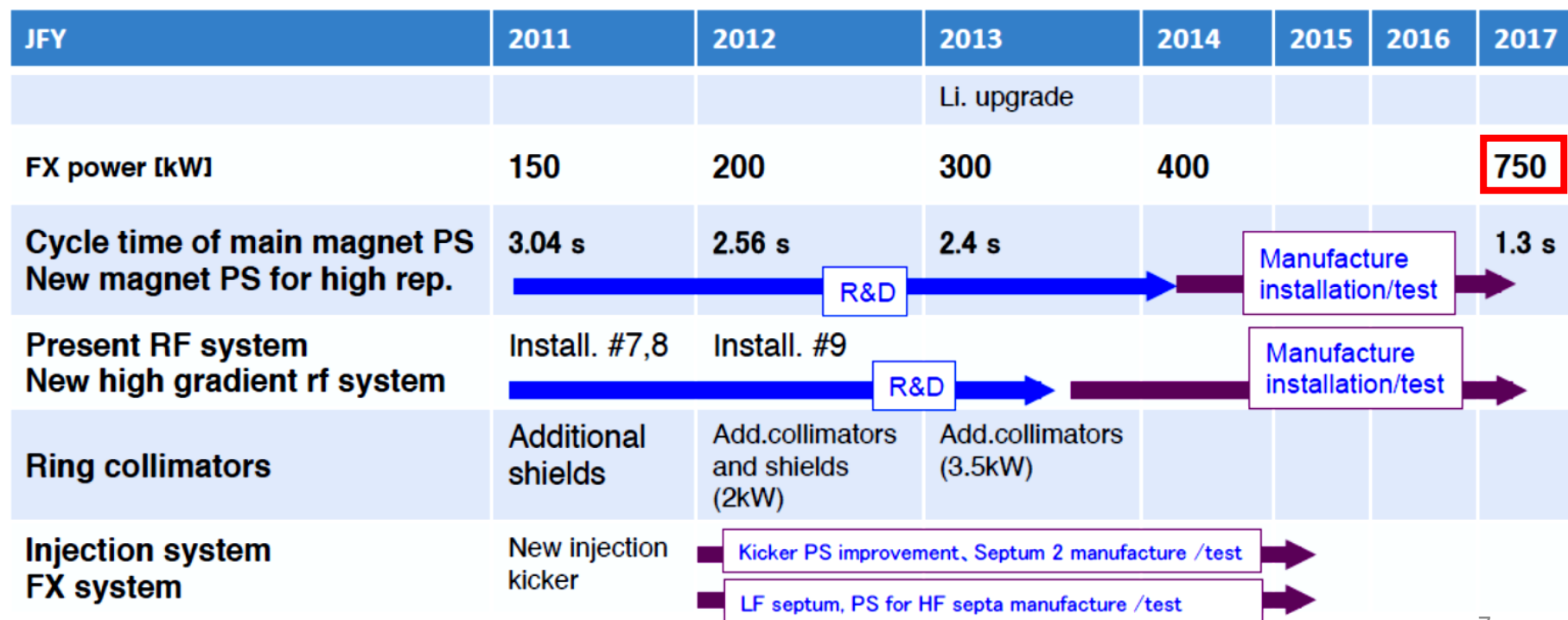
## *J-PARC Accelerator & neutrino beamline*

Proton intensity up to **750 kW** has been planned  
from the beginning.

Expected to achieve in 3 years.

*Further upgrade ( > 1MW ) is under study.*

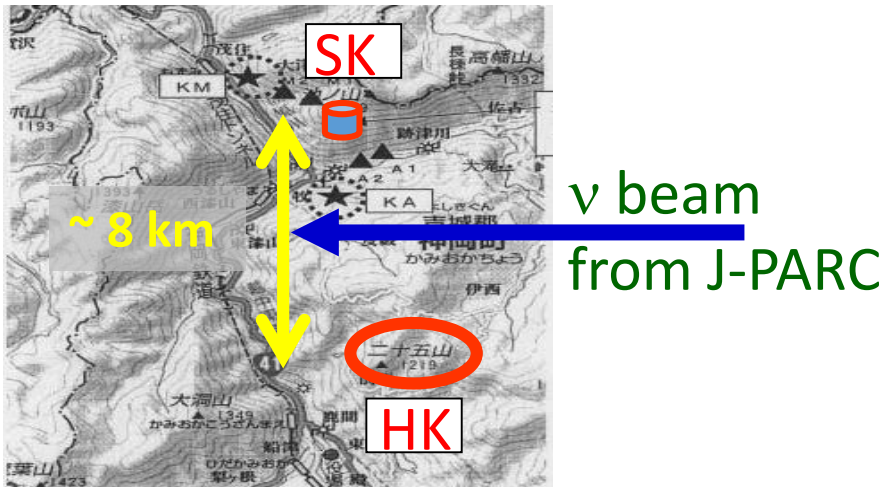
All the beamline components are designed to accept full 750 kW beam.  
Some unreplaceable components are designed to accept > MW beam.



# Hyper-Kamiokande with J-PARC neutrino beam

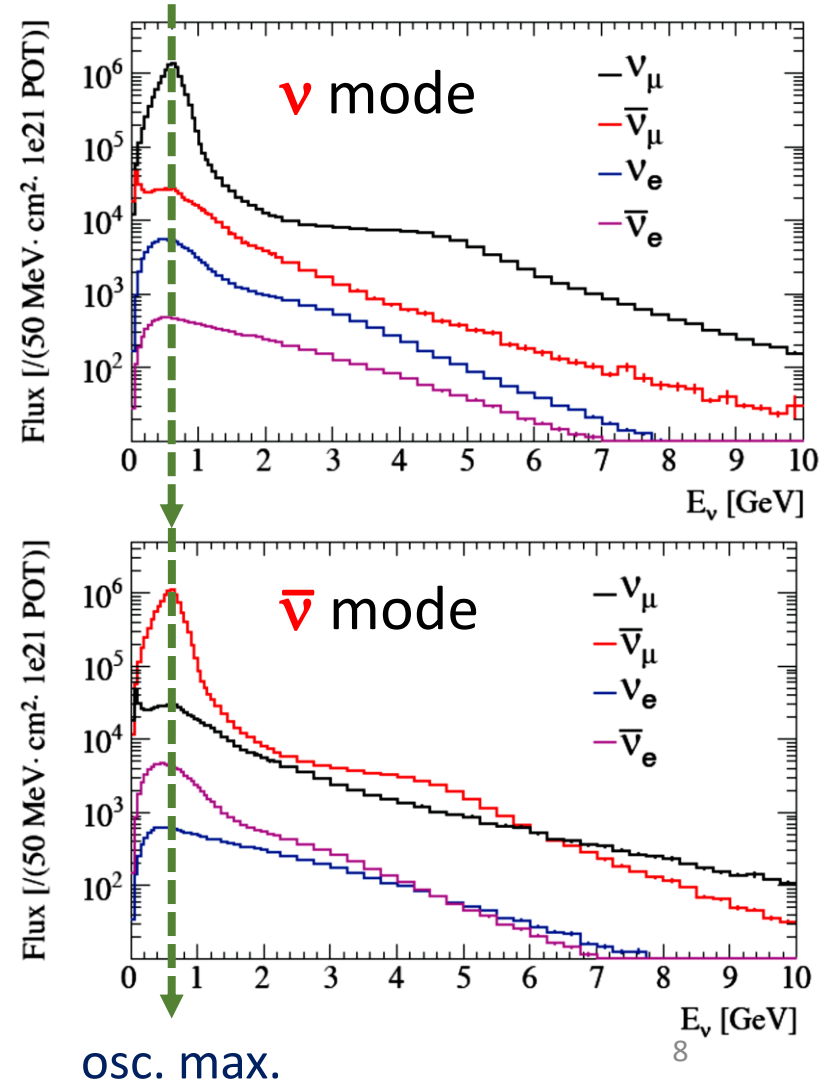
## *Off-axis narrow band $\nu$ / $\bar{\nu}$ beam*

Optimized for the neutrino oscillation experiments in Kamioka.  
( Both for SK and Hyper-K. )



*Designed to have same  
off axis angle at both sites.*

Small  $\nu_e$  /  $\bar{\nu}_e$  background  
around the flux peak



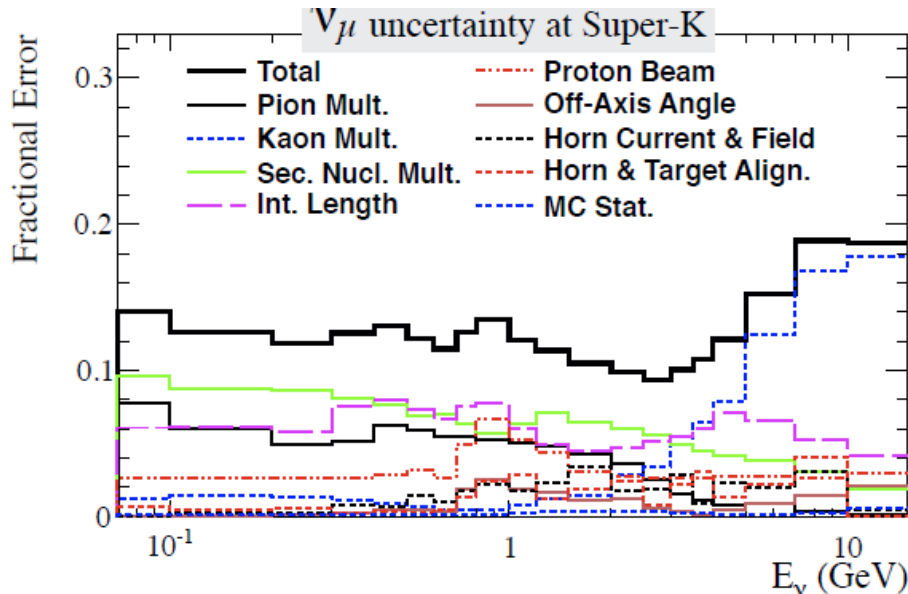


# Hyper-Kamiokande with J-PARC neutrino beam

## *Off-axis narrow band beam*

Extensive studies have been performed to understand the characteristics of the neutrino beam.

1) Dedicated hadron production experiments  
( as a part of NA61-Shine )



2) Uncertainties on the ratio  
( near/far ratio ) is  
*less than 2%*  
near the flux peak.

*& Further reduction is expected.*

2) Established beam monitoring scheme.

Primary, secondary (  $\mu$  ) and  $\nu$  themselves.

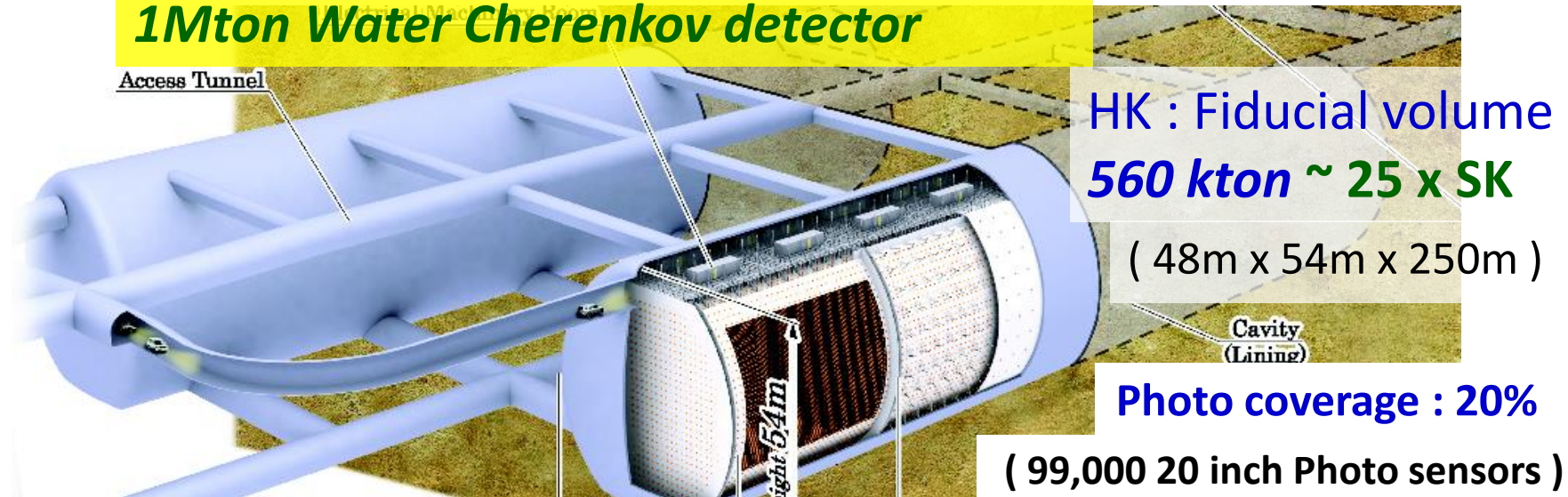
➡ *Properties of the neutrino beam is well understood.*

# Hyper-Kamiokande

## *Far detector “Hyper-Kamiokande”*

What is not sufficient in SK? => **~ Statistics = target mass ~**

### **1Mton Water Cherenkov detector**



Maximum utilization of resources and experiences in SK

~ Use established technology for the long term operation  
to achieve physics goal in timely manner.

### **Broad science programs**

- 1) Accelerator neutrinos from J-PARC**
- 2) Atmospheric, Solar, Super Nova and cosmic neutrinos
- 3) Nucleon decay searches etc....

# Physics in Hyper-Kamiokande

## *Neutrino oscillation*

- Accelerator neutrinos
- Atmospheric neutrinos
- Solar neutrinos

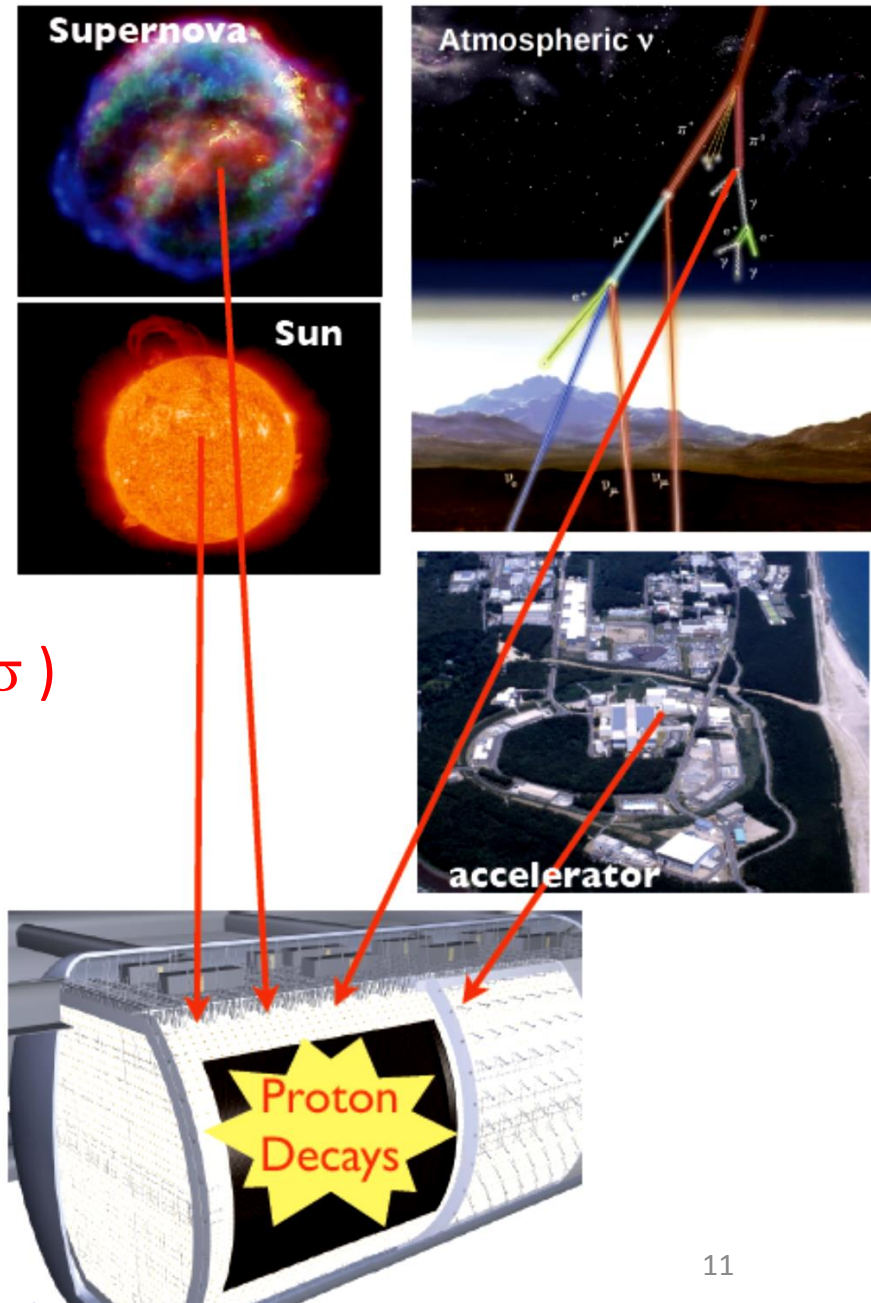
## *Proton decay search ~ GUT*

Possible discovery with  $\sim 10$  times better sensitivities than SK

$$\begin{aligned} p &\rightarrow e^+ + \pi^0 : \sim 5.7 \times 10^{34} \text{ years} \\ p &\rightarrow K^+ + \bar{\nu} : \sim 1.2 \times 10^{34} \text{ years} \end{aligned} \quad (3\sigma)$$

## *Neutrino astrophysics*

- Super nova burst neutrino  
Expected  $\sim 200,000$   $\nu$  events from SN @ 10kpc
- Relic SN neutrinos  
Expected several hundreds of events



# Hyper-Kamiokande with J-PARC neutrino beam

## *Water Cherenkov detector*

- Well established technology

Past experiences in the long term operation

~ Need O(10) years of operation ~

- Continuous improvements in the data analysis

( Event reconstruction & background rejections )

One particular example ~ *New  $\pi^0$  rejection method* ~

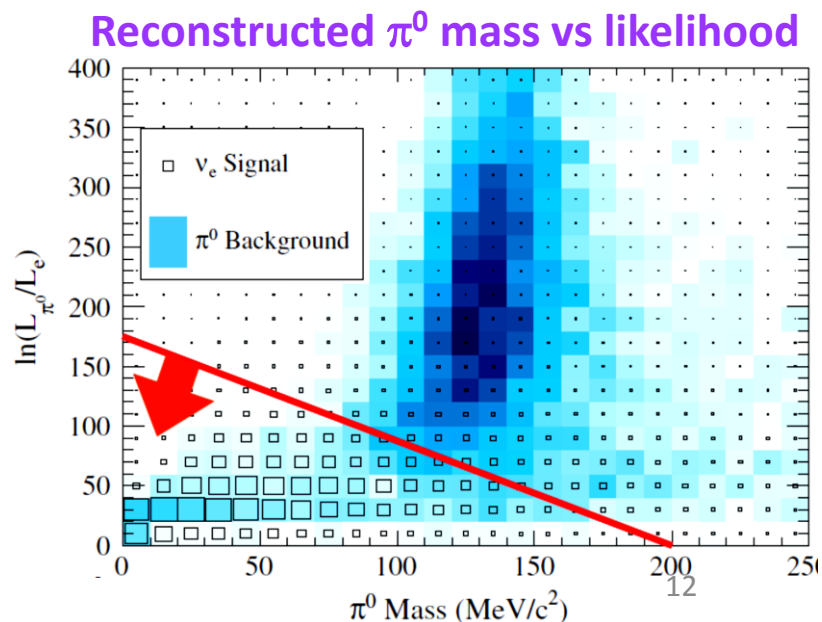
$\pi^0$  used to be one of the dominant background sources

in  $\nu_e$  appearance search.

In the latest T2K analyses,

**only 23%** is coming from  $\pi^0$ s!

\*) 74% of background events  
are beam intrinsic  $\nu_e$ .





# Hyper-Kamiokande project

Selected one of the **27 'top projects'**

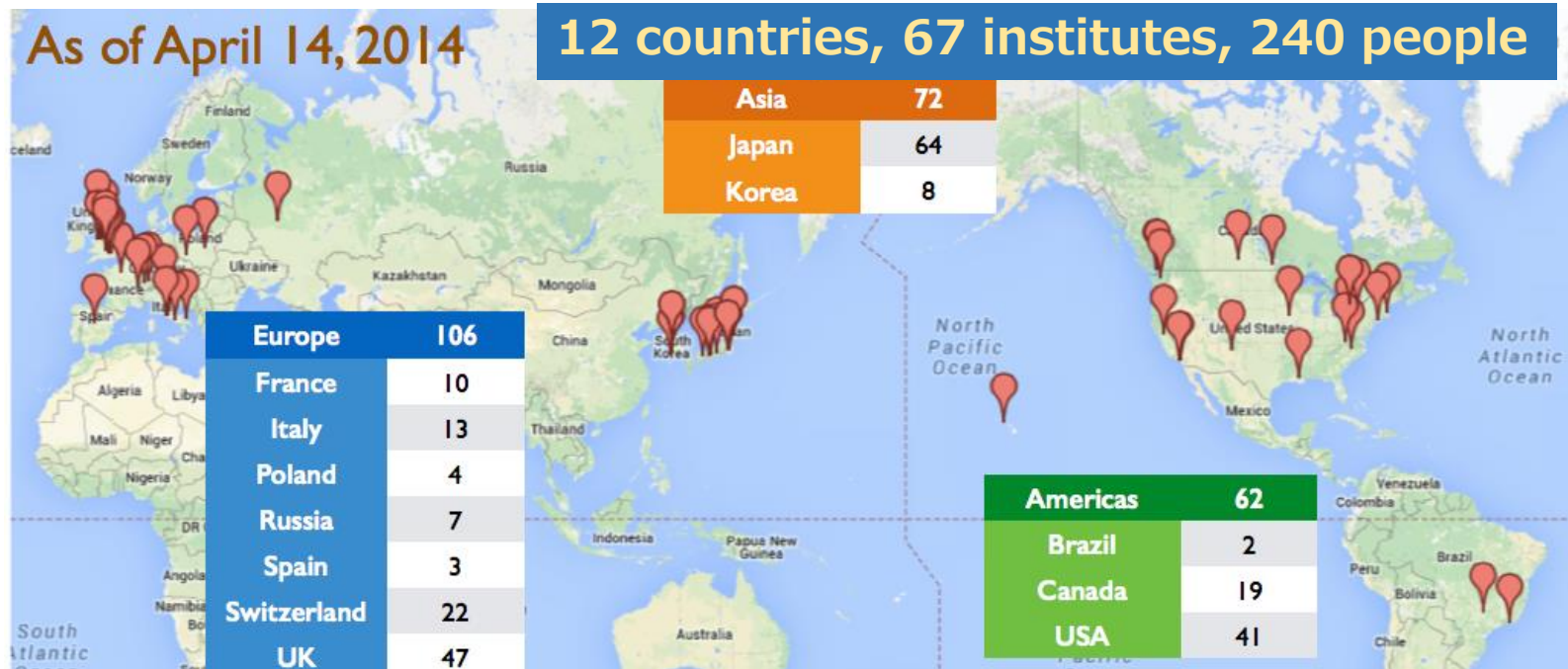
in 'Japanese master plan for large scale research projects'  
by Science Council of Japan

International working group was formed ( <http://www.hyperk.org> )

~ *Wide variety of physics attracts many people from the world* ~

*Next open meeting : July 19-22, 2014 in Vancouver*

<http://indico.ipmu.jp/indico/conferenceDisplay.py?confId=34>

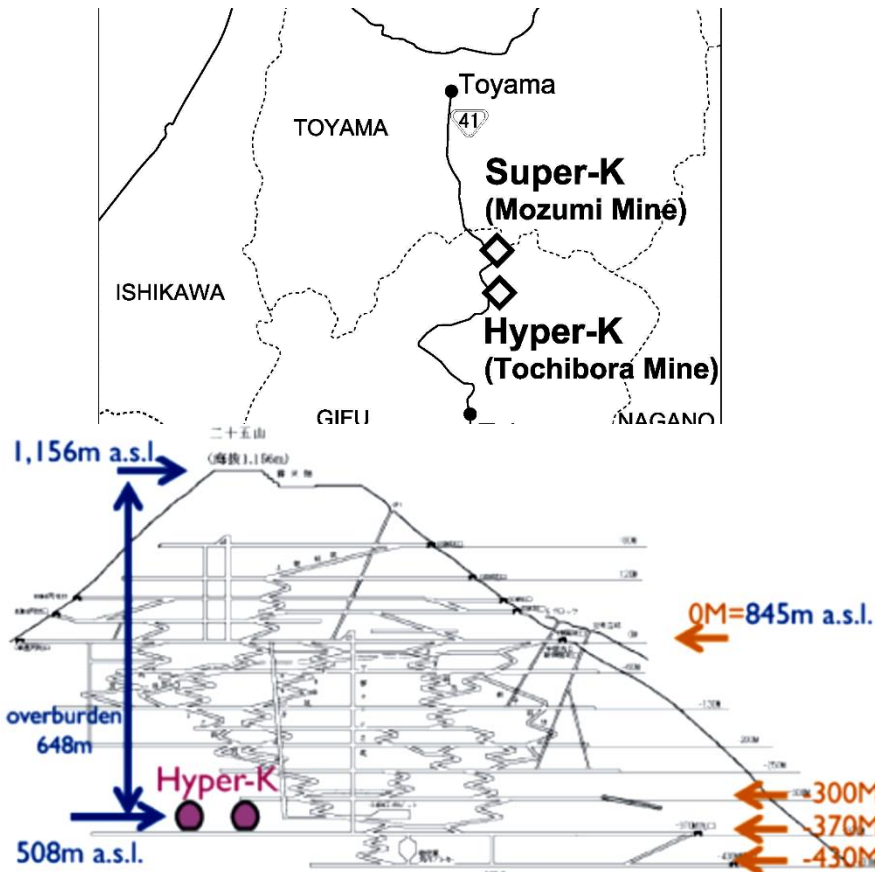




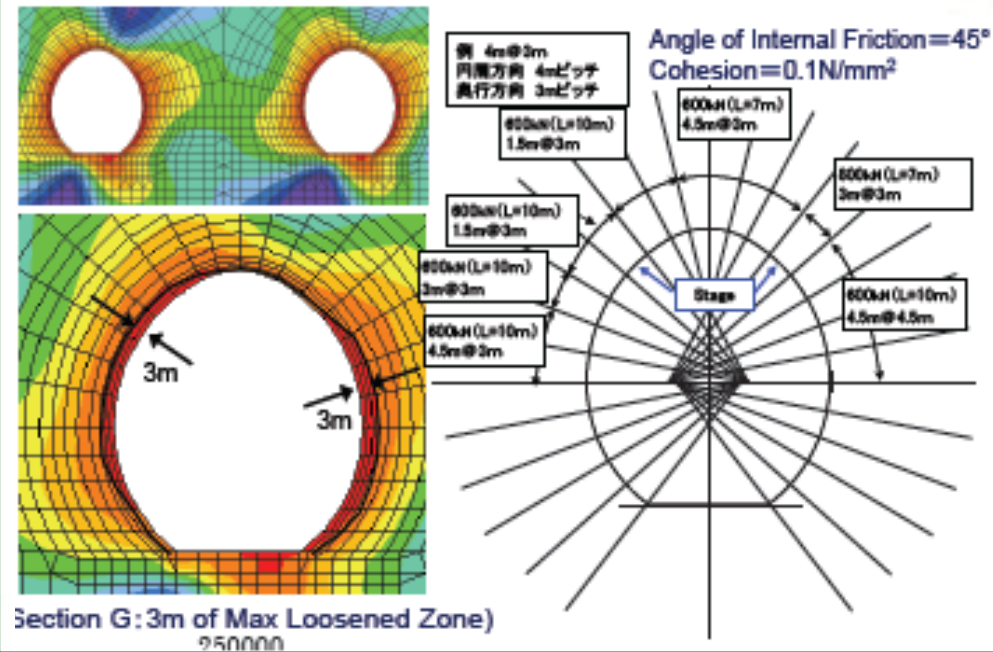
# Hyper-Kamiokande detector

## Is it possible to construct such gigantic detectors?

## Candidate site : Tochibora mine in Kamioka



## Design of the cavern and its supporting structures



➡ Based on the geological survey and analyses,  
the cavern and the supporting structures were designed.

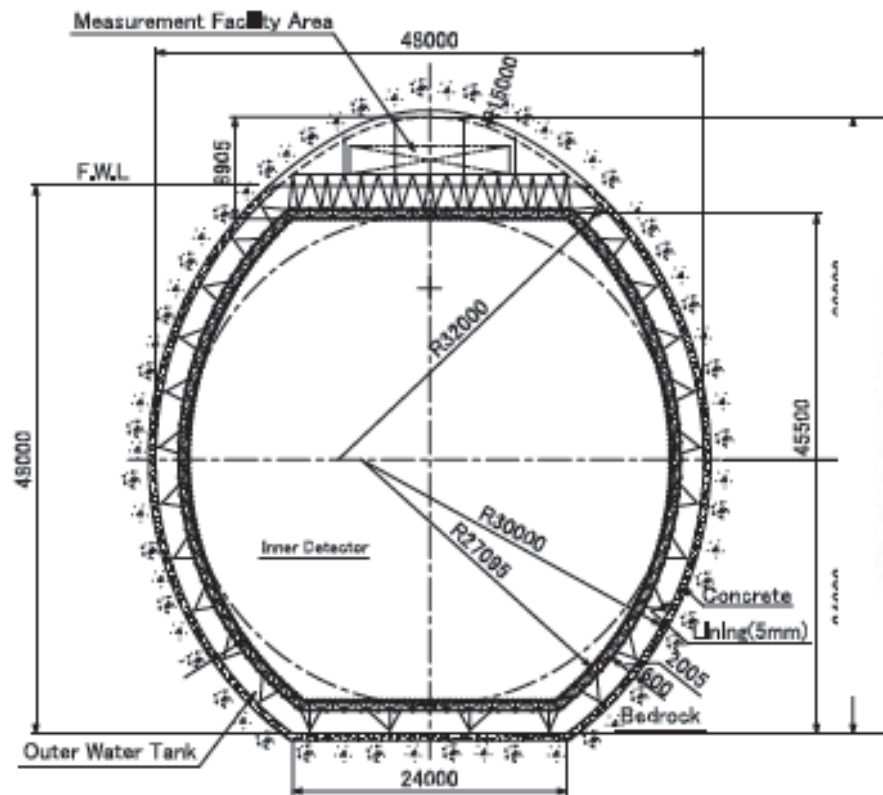
***Possible to construct HK Caverns with existing technology.*** <sup>14</sup>

# Hyper-Kamiokande detector

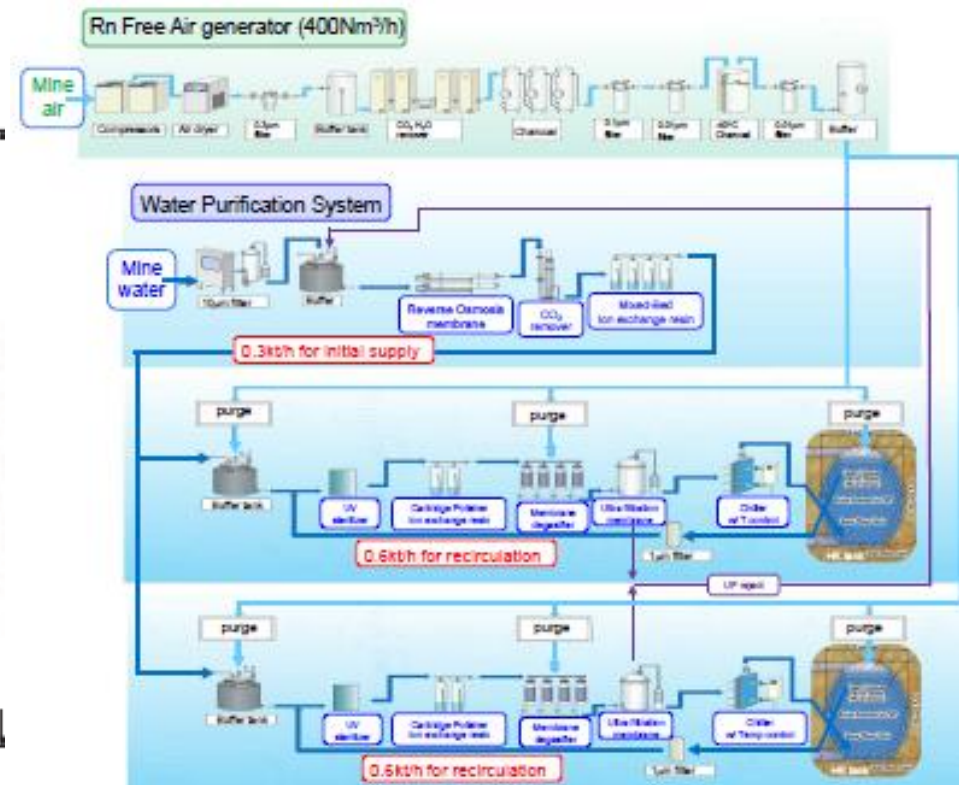
## Design of the detector structure

Incl. PMT supports

### CROSS SECTION



## Schematic diagram of water purification system for HK detector



➡ **Baseline design of the detector is finished based on the past experiences in SK.** <sup>15</sup>

# Hyper-Kamiokande detector ~ Further improvements ~

## *Photo sensors ~ R&D to improve the detector performance*

Better timing resolution ~ better vertex resolution

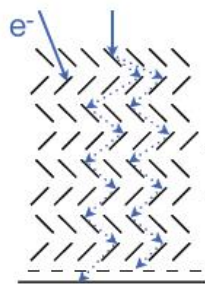
Higher quantum efficiency

### *Baseline (reference)*

20" Super-K PMT



Venetian blind dynode



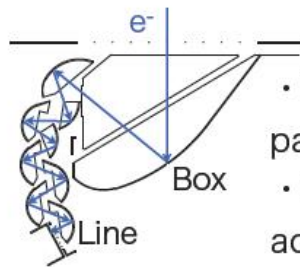
- Various drift path
- Might miss dynode

### *Candidates (R&D phase)*

20" Box&line PMT



Box&line dynode

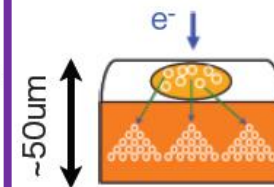


- Unique drift path
- Large acceptance

20" HPD



Avalanche diode (AD)



- Short drift path
- High first step gain ( $\times 1600$ )

Quantum eff.

22%

30%

30%

Collection eff.

80%

93%

95%

Timing res. ( FWHM )

5.5 nsec

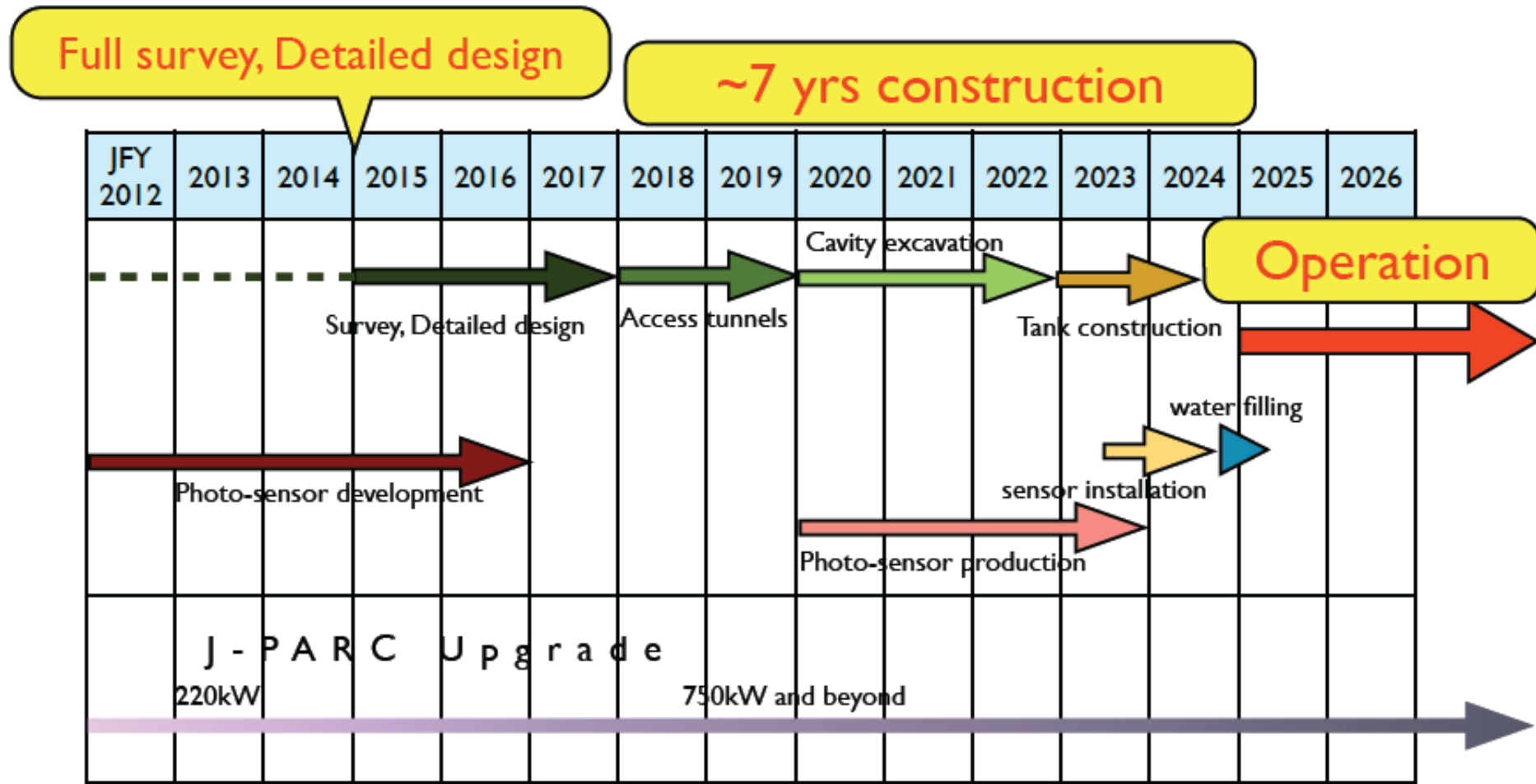
2.7 nsec

1 nsec





# Hyper-Kamiokande project ~ Notional Timeline ~



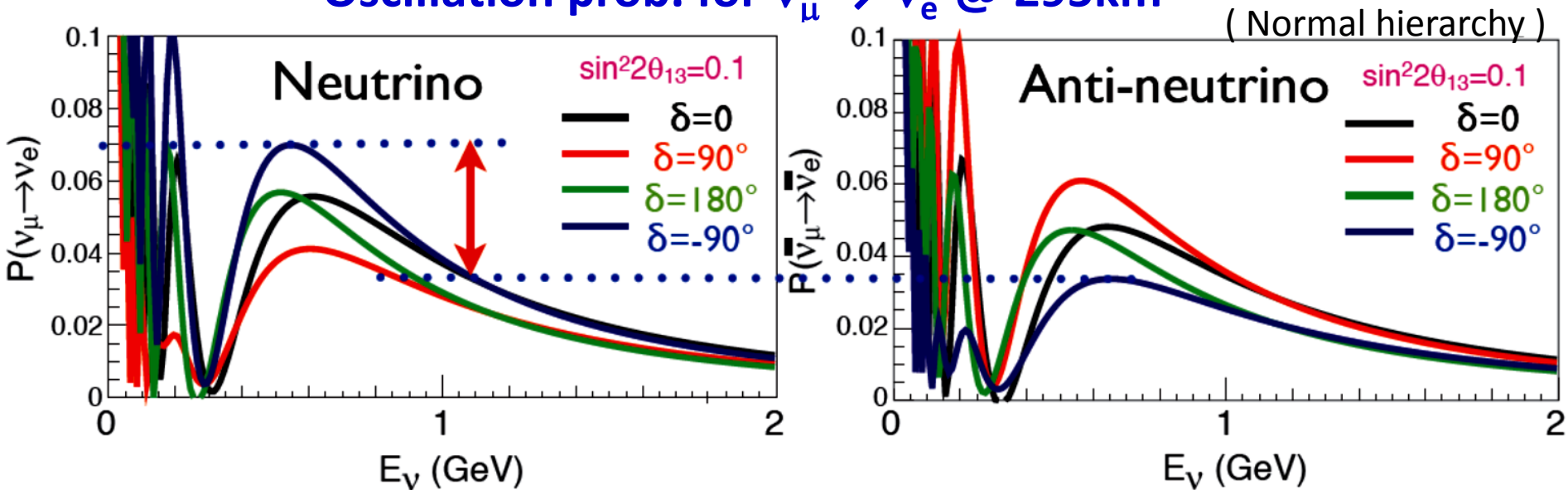
- 2015 Full survey, Detailed design (3 years)
- 2018 Excavation start (7 years)
- 2025 Start operation



# Neutrino physics of LBL J-PARC & HK ~ Determination of CP $\delta$

CP-non conservation term in osc. prob.  $\propto \sin\theta_{13}\sin\delta$  (  $\sin^2 2\theta_{13} \sim 0.1$  )  
( sign of  $\delta$  for anti-neutrino is different from neutrino )

## Oscillation prob. for $\nu_\mu \rightarrow \nu_e$ @ 295km



→ Hyper-Kamiokande + J-PARC neutrino beam  
~ 3000  $\nu_e$  & ~ 2000  $\bar{\nu}_e$  signal events are expected, when  $\delta = 0$   
(  $7.5 \times 10^7$  MW·sec )

*Measurements of  $\delta$  by comparing oscillations of  $\nu$  and  $\bar{\nu}$ .*

At maximum CP violation, ~25% difference from  $\delta=0$  case.

# Neutrino physics of LBL J-PARC & HK ~ Determination of CP $\delta$

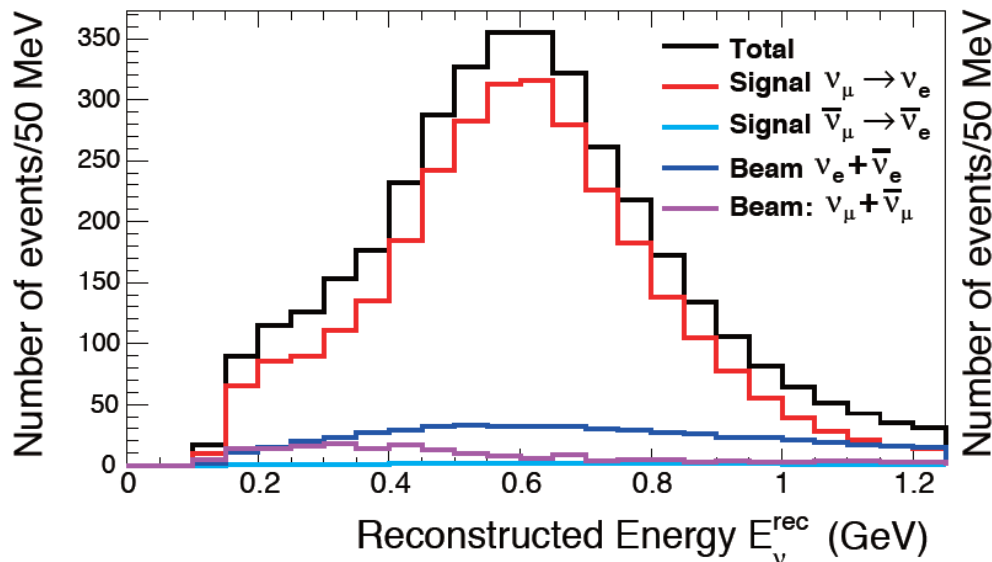
- Expected # of events for  $\sin^2 2\theta_{13} = 0.1$ ,  $\delta = 0$  and NH  
(  $7.5 \times 10^7 \text{ MW}\cdot\text{sec}$  )

	Signal ( $\nu\mu \rightarrow \nu_e \text{ CC}$ )	Wrong sign appearance	$\nu\mu/\bar{\nu}\mu$ CC	beam $\nu_e/\bar{\nu}_e$ contamination	NC
$\nu$	3,016	28	11	523	172
$\bar{\nu}$	2,110	396	9	618	265

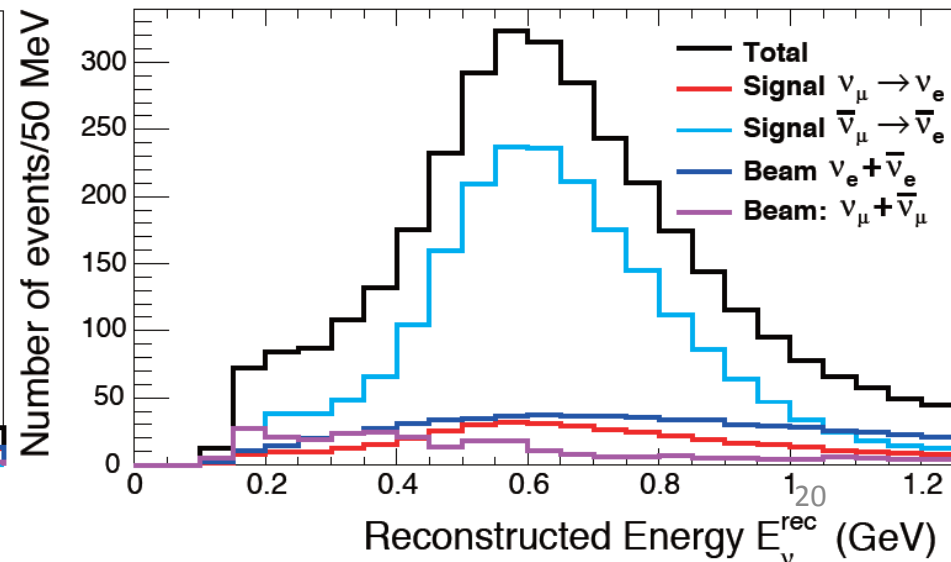
*NC (  $\pi^0$  ) is not the dominant background already.*

- Reconstructed energy of neutrino for candidate events

Neutrino mode beam



Anti neutrino mode beam



# Neutrino physics of LBL J-PARC & HK ~ Determination of CP $\delta$

## *Systematic error Errors used in the sensitivity studies*

~ Realistic estimation of the errors based on the experiences ~

	$\nu$ mode		anti- $\nu$ mode	
	$\nu e$	$\nu \mu$	$\nu e$	$\nu \mu$
Flux&ND	3.0	2.8	5.6	4.2
XSEC model	1.2	1.5	2.0	1.4
Far Det. +FSI	0.7	1.0	1.7	1.1
<b>Total</b>	<b>3.3</b>	<b>3.3</b>	<b>6.2</b>	<b>4.5</b>

(T2K 2014)	
$\nu e$	$\nu \mu$
2.9	2.7
4.7	4.9
3.5	5.6
6.8	8.1

Reduction of errors in the XSEC models

New measurements of neutrino interactions

Improved theoretical modeling

Reduction of errors in the far detector + Final state interactions

Increased statistics of atmospheric  $\nu$  control sample in HK

New near ( intermediate ) detectors with  $H_2O$  target

( incl. Water Cherenkov detector )

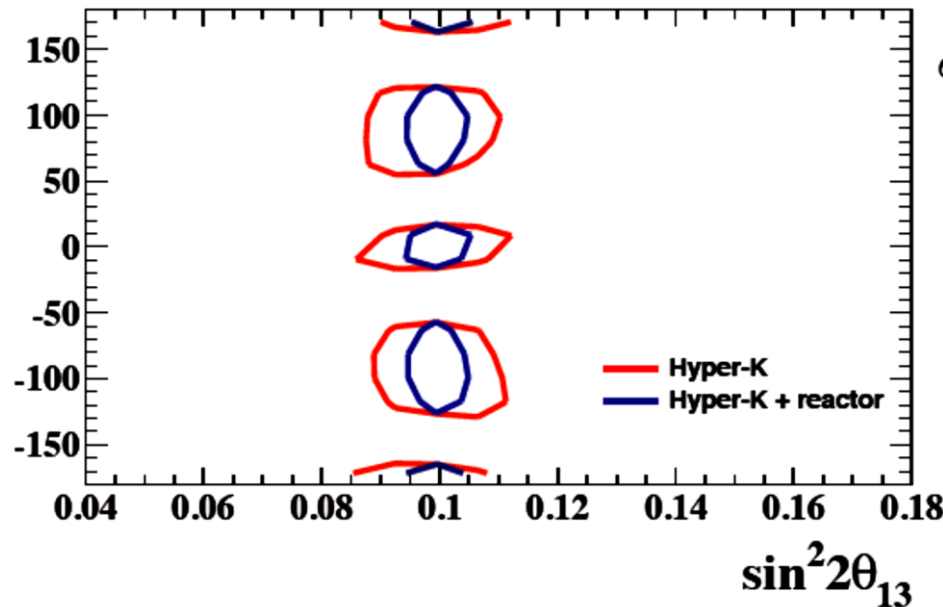
# Neutrino physics of LBL J-PARC & HK ~ Determination of CP $\delta$

Use both # of observed events

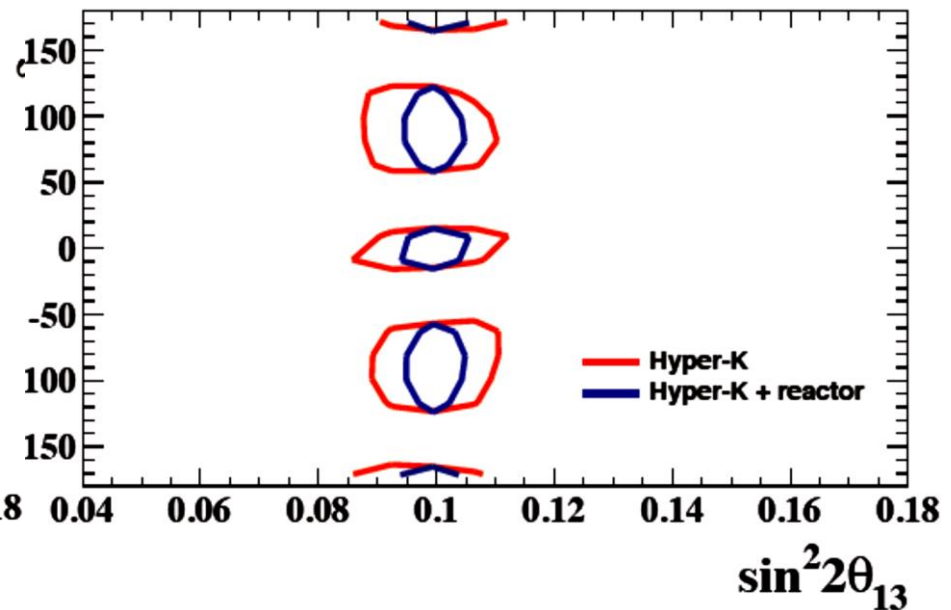
and reconstructed energy spectra of  $\nu$  and  $\bar{\nu}$ .

( @  $7.5 \times 10^7 \text{ MW}\cdot\text{sec}$ ,  $\nu$ :  $\bar{\nu}=1:3$  )

*Normal mass hierarchy*



*Inverted mass hierarchy*

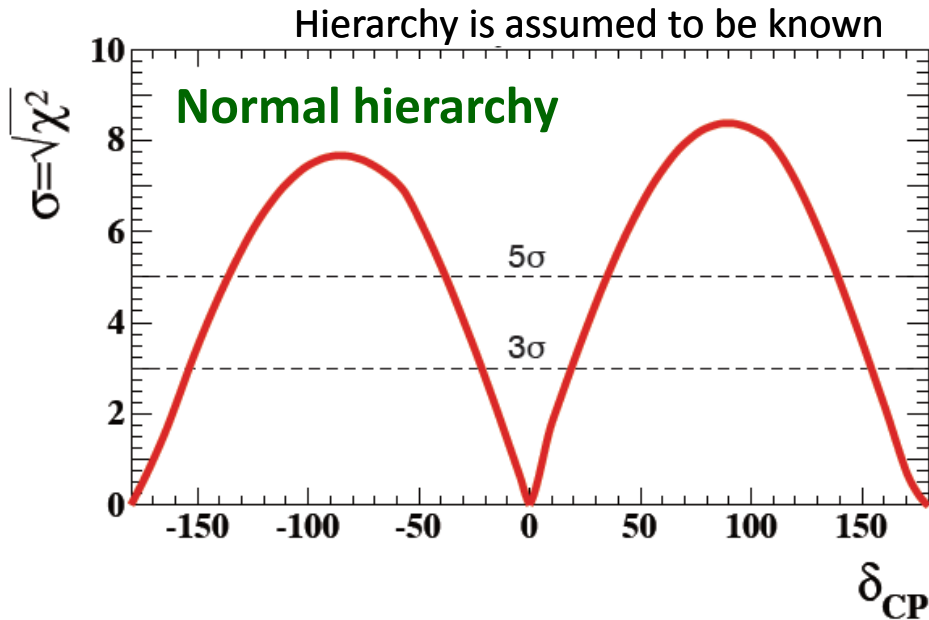


*Determination power of CP  $\delta$  parameter*

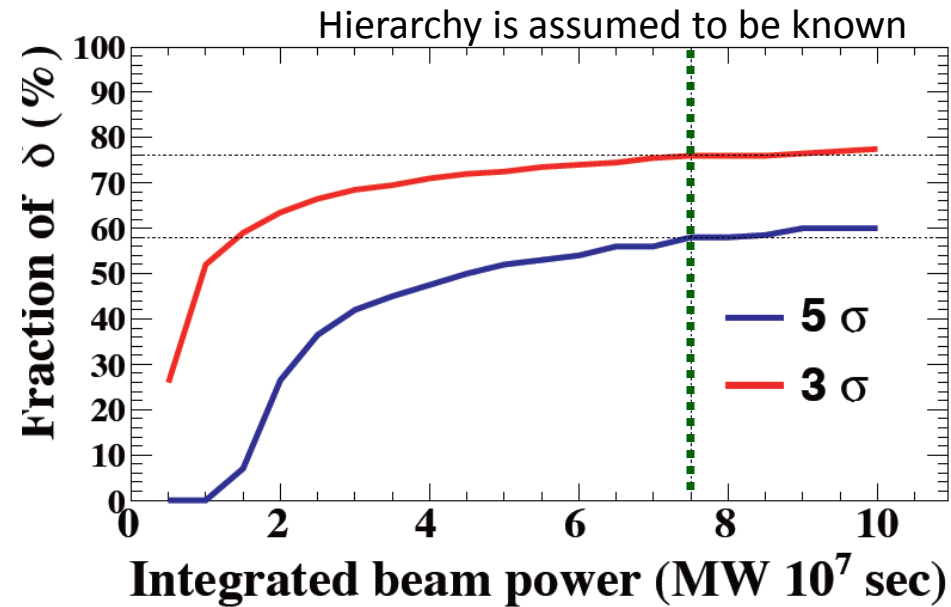
*$1\sigma$  error of  $\delta$  is expected to be  $8^\circ \sim 19^\circ$ .*

# Neutrino physics of LBL J-PARC & HK ~ Determination of CP $\delta$

*Sensitivity ~ Exclusion of  $\sin\delta = 0$*   
*(  $7.5 \times 10^7$  MW·sec )*



*Fraction of  $\delta$*   
*~ Exclusion of  $\sin\delta = 0$*



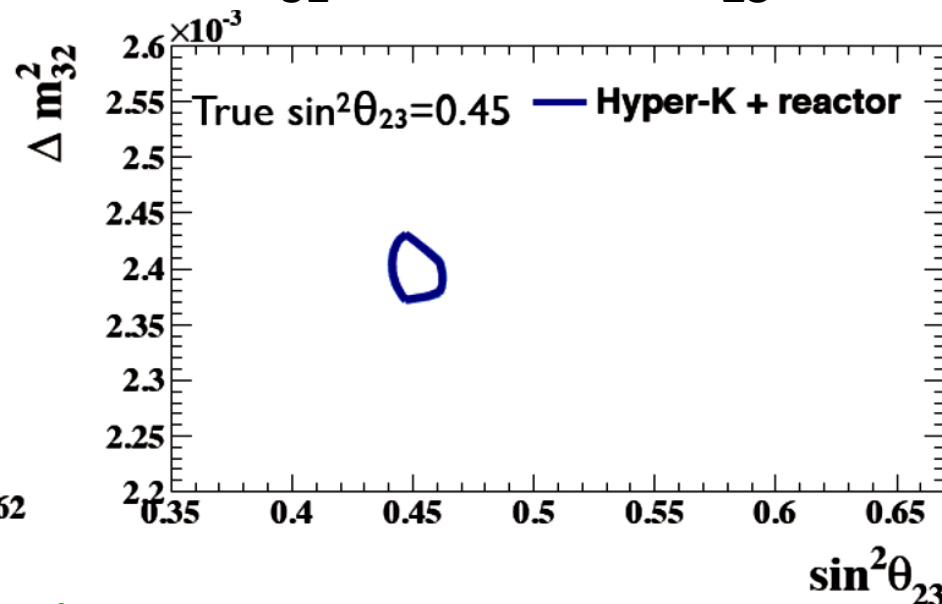
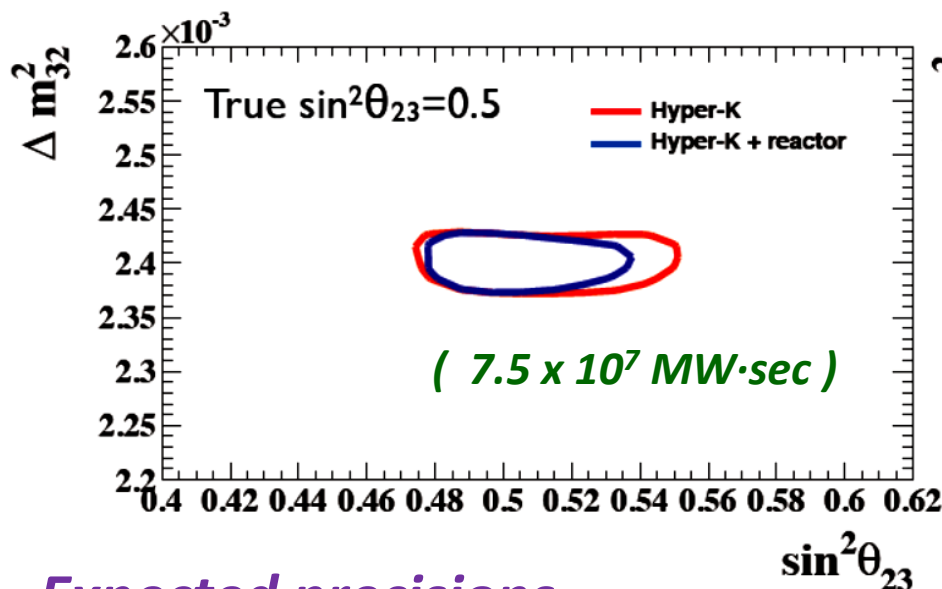
*Exclusion of  $\sin\delta = 0$*

*76% of  $\delta$  at 3 $\sigma$  level and 58% of  $\delta$  at 5 $\sigma$  level*  
*with realistic systematic error estimations.*



# Neutrino physics of LBL J-PARC & HK

~ Measurements of  $|\Delta m_{32}^2|$  and  $\sin^2 2\theta_{23}$



*Expected precisions* (  $7.5 \times 10^7 \text{ MW}\cdot\text{sec}$  )

$$\text{True } \Delta m_{32}^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

$$\pm 0.015 \times 10^{-3} \text{ eV}^2 \text{ ( 0.6\% )}$$

$$\text{True } \sin^2 \theta_{23} = \begin{cases} 0.55 \\ 0.50 \text{ ( Full mixing )} \\ 0.45 \end{cases}$$

$$\pm 0.009 \text{ ( 1.6\% )}$$

$$\pm 0.015 \text{ ( 3\% )}$$

$$\pm 0.006 \text{ ( 1.3\% )}$$

( T2K 2014 results

$$\Delta m_{32}^2 = 2.51 \pm 0.10 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} = 0.514 \pm 0.055 \text{ )}$$

*Large improvements & good chance to identify non-maximal mixing.*

## *MOMENT*

*~ a muon decay medium baseline neutrino facility ~*

MOMENT ~ a muon decay medium baseline neutrino facility

***Future accelerator based neutrino oscillation experiment in China***

Intended to study CP violation in the lepton sector

*Lowering the energy to largely reduce backgrounds  
from pion productions.*

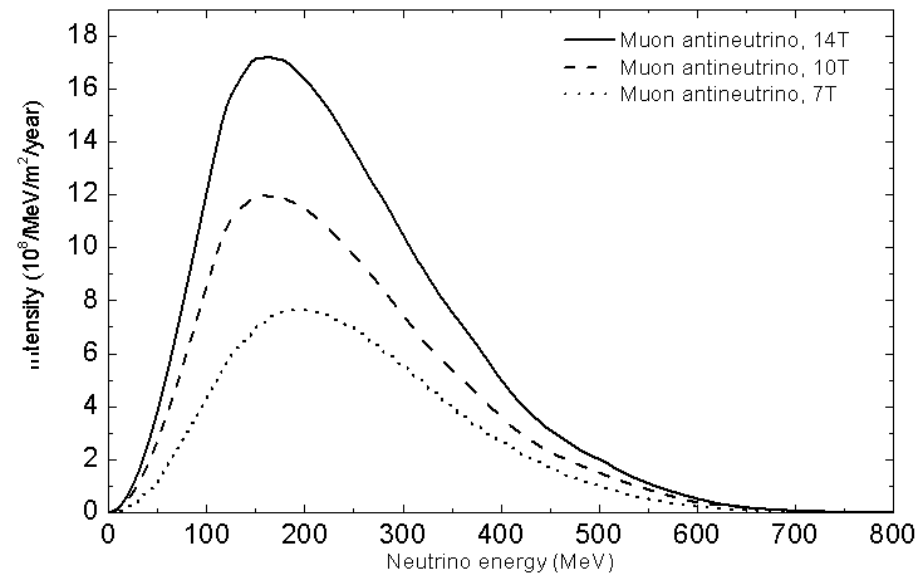
Planned **baseline is ~ 150km**

**Mean beam** energy is tuned **~ 240 MeV.**

Neutrino beam from **muon decays**  
( Conventional neutrino beam  
uses  $\nu$  from pion decays )

Expected neutrino flux  
( Depends on the level of  
the pion capture field at target )

14T Field :  **$4.7 \times 10^{11}$**   $\nu/\text{m}^2/\text{year}$   
( 7T Field :  $2.1 \times 10^{11}$   $\nu/\text{m}^2/\text{year}$  )

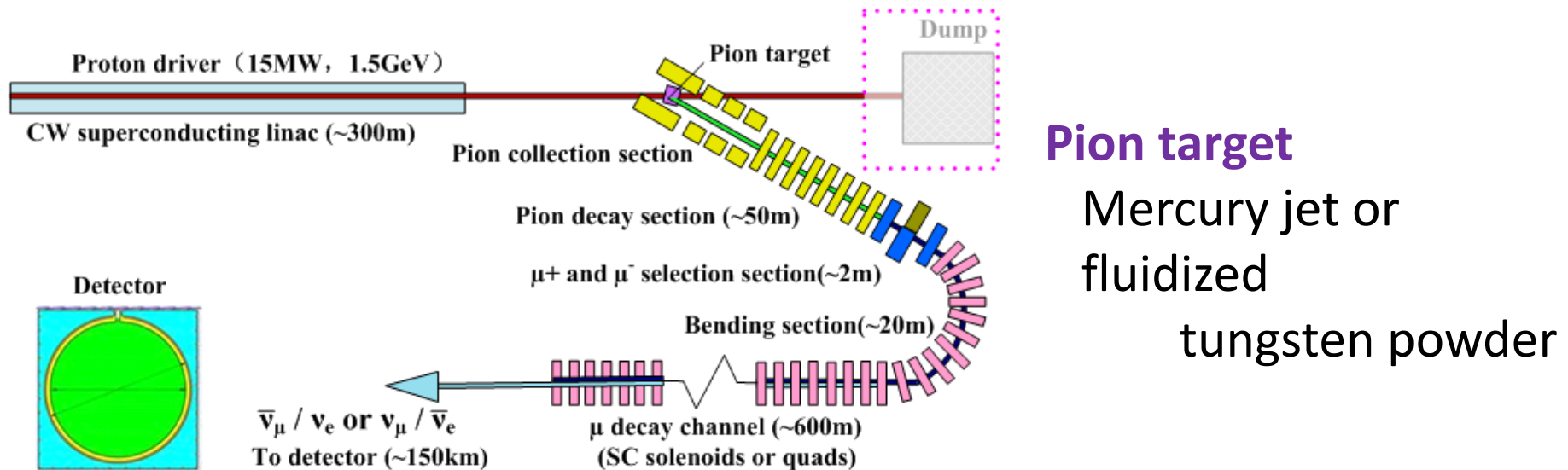


MOMENT ~ a muon decay medium baseline neutrino facility

Use extremely intense proton beam ( **15MW** )

from CW Superconducting LINAC

**Accelerator** : Use the same technology but different design  
for China-ADS (a funded R&D project)



**Detector** : \* *Need charge identification* \*

Various possibilities are considered

Water Cherenkov detector with Gd ( like GAZOOKS )

Magnetized Iron detector ( MIND )

Magnetized liquid Argon detector

Summary     *~ It's time to start new experiments  
to explore CP violation in the lepton sector. ~*

## J-PARC neutrino beam with Hyper-Kamiokande

- Utilize existing intense neutrino beamline with upgrades  
~ Well understood and under control.

- Feasible Gigantic Water Cherenkov detector

Established detector technology ~ Feasible

Proven excellent performance in physics analyses

Proven long term stability

Further reduction of the cost

& performance improvements

Realized by new technology and analysis methods

( *Will be confirmed with the test detectors and SK. )*

Broad physics opportunities

Various sources of neutrinos, nucleon decay etc..

**MOMENT** ~ a muon decay medium baseline neutrino facility

Different configuration of a new experiment with

a new design of intense neutrino beam production



Fin.

# Introduction

Neutrino oscillation ~ discovered in 1998 & extensively studied.

Flavor mixing & non-zero neutrino mass

~ Beyond the standard model ~

- Parameters
- 3 oscillation angles( $\theta_{12}, \theta_{23}, \theta_{13}$ )
  - 2 mass differences( $\Delta m^2_{12}, \Delta m^2_{32}$ )
  - **1 CP phase ( $\delta$ )**

PMNS Matrix ( $U_{\alpha i}$ )

$$|v_\alpha\rangle = \sum U_{\alpha i} |v_i\rangle$$

*Weak*

*Mass eigenstates*

$$s_{ij} = \sin \theta_{ij}, c_{ij} = \cos \theta_{ij}$$

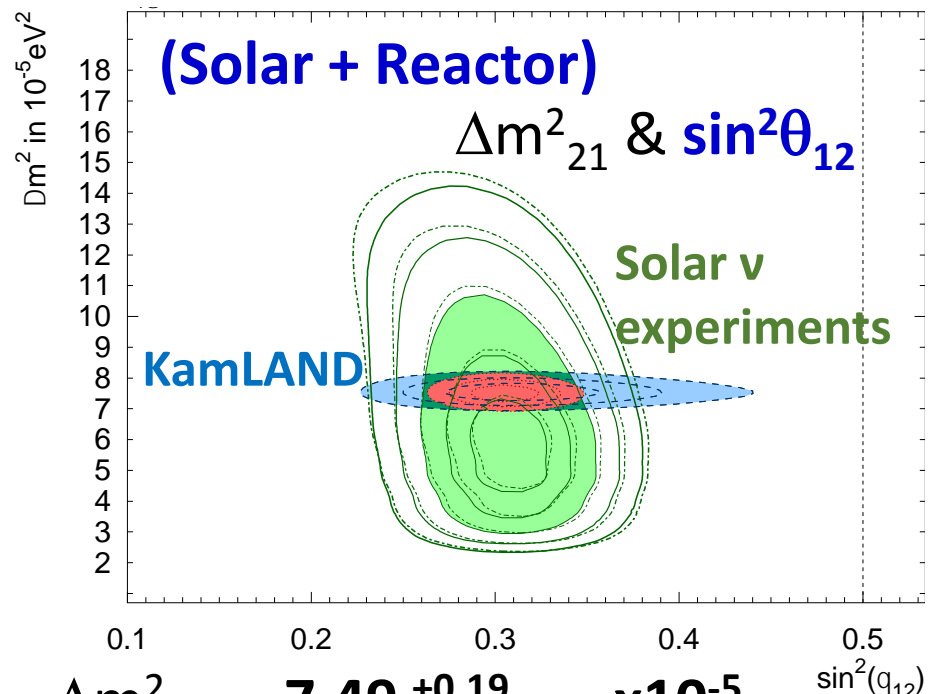
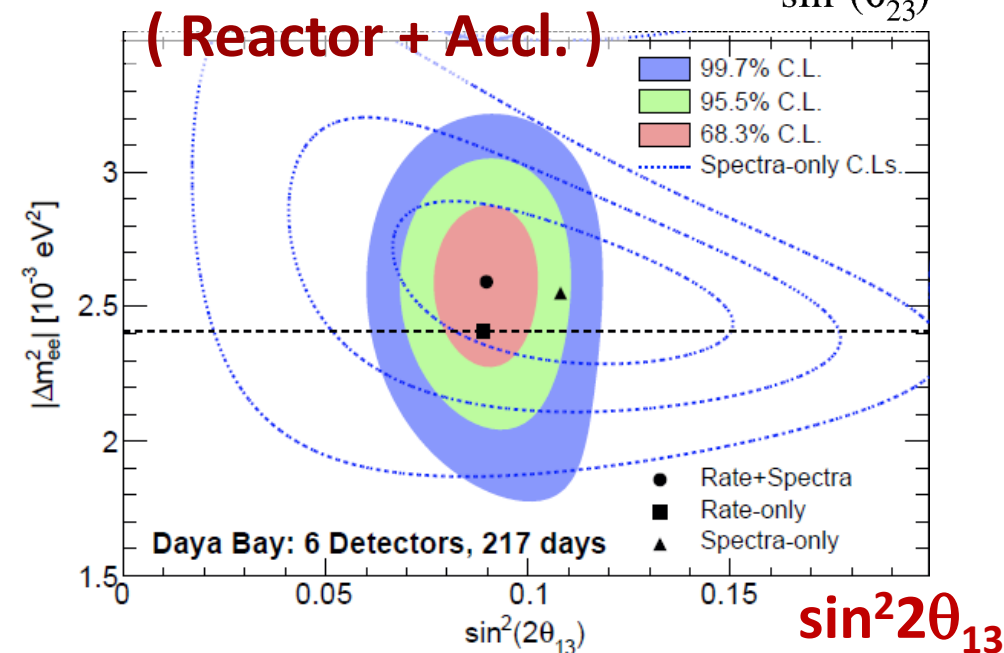
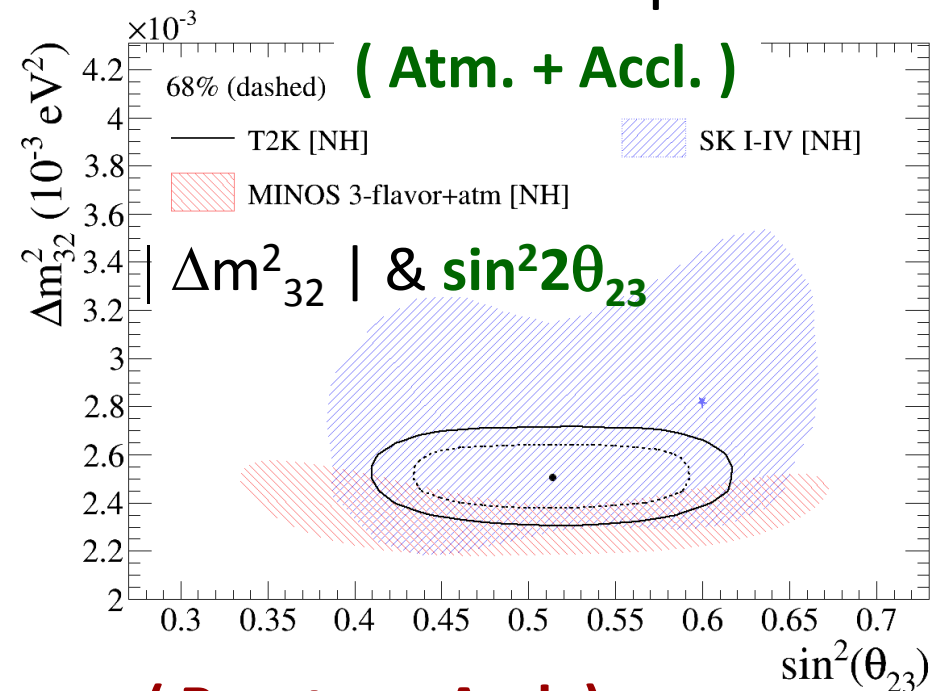
$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix}$$

**Atmospheric  
& Accelerator**

**Reactor  
& Accelerator**

**Solar  
& Reactor**

# Neutrino oscillation parameter measurements



$$\Delta m^2_{21} = 7.49^{+0.19}_{-0.17} \times 10^{-5} \text{ (eV/c}^2\text{)}^2$$

$$|\Delta m^2_{32}| \sim 2.5 \pm 0.2 \times 10^{-3} \text{ (eV/c}^2\text{)}^2$$

$$\sin^2 2\theta_{23} > 0.9$$

$$\sin^2 \theta_{12} \sim 0.305 \pm 0.013$$

$$\sin^2 2\theta_{13} = 0.098 \pm 0.013$$

Daya-Bay collab. PRL 112 (2014) 061801

$$\sin^2 2\theta_{13} = 0.090^{+0.008}_{-0.009}$$

# Next step ~ Prospects of the current experiment

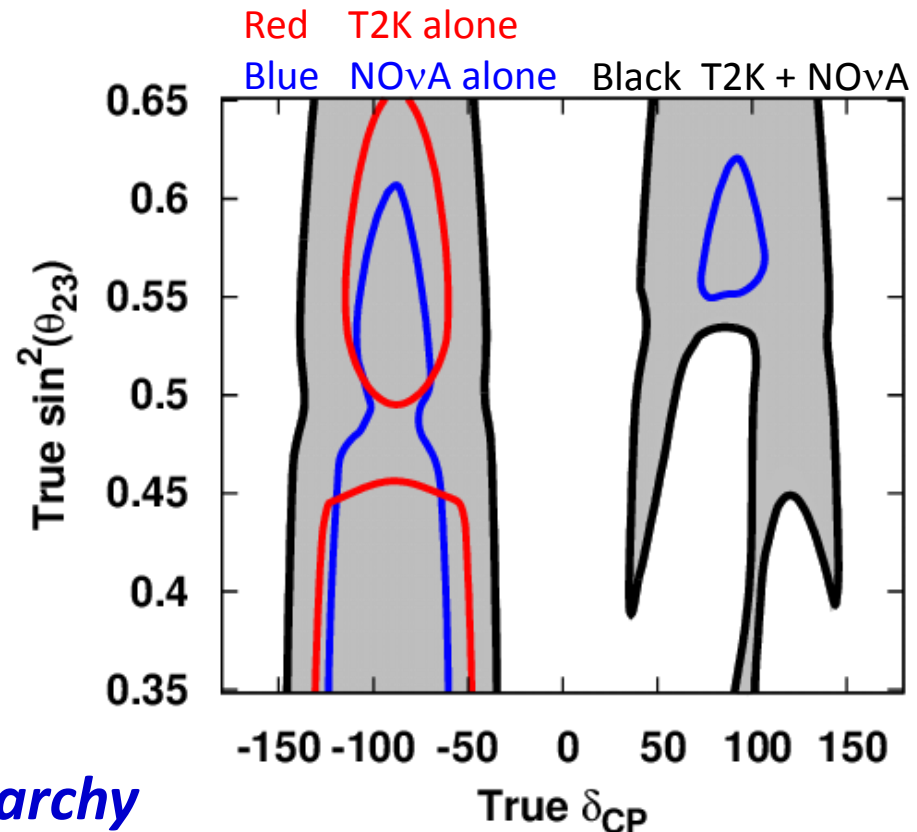
- Current generation experiments ( T2K and NOvA ) are expected to lead the study of the CP violation in the lepton sector.
- However, the expected sensitivity of CPV is  $\sim 2\sigma$  level



*Next generation experiments  
are essential  
for the study of CPV  
and mass hierarchy*

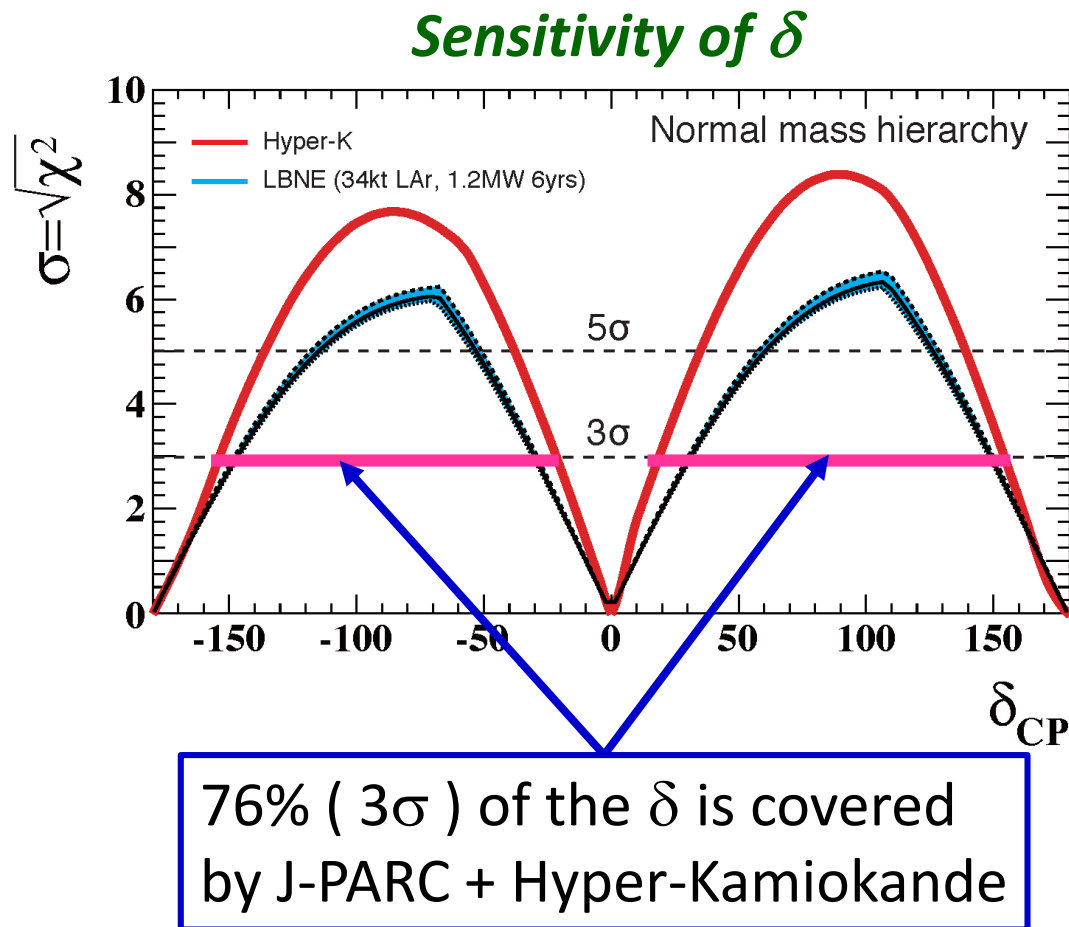
Both T2K / NOvA : full POT ( POT fraction : 50%  $\nu$  + 50% anti- $\nu$ )  
Assuming 5% (10%) normalization uncertainty on signal (background)  
Assuming true:  $\sin^2 2\theta_{13}=0.1$ ,  $\Delta m^2_{32}=2.4 \times 10^{-3} \text{ eV}^2$   
 $\theta_{13}$  constrained by  $\delta(\sin^2 2\theta_{13}) = 0.005$

*Region where  $\sin \delta=0$   
can be excluded by 90% C.L.  
Normal hierarchy*



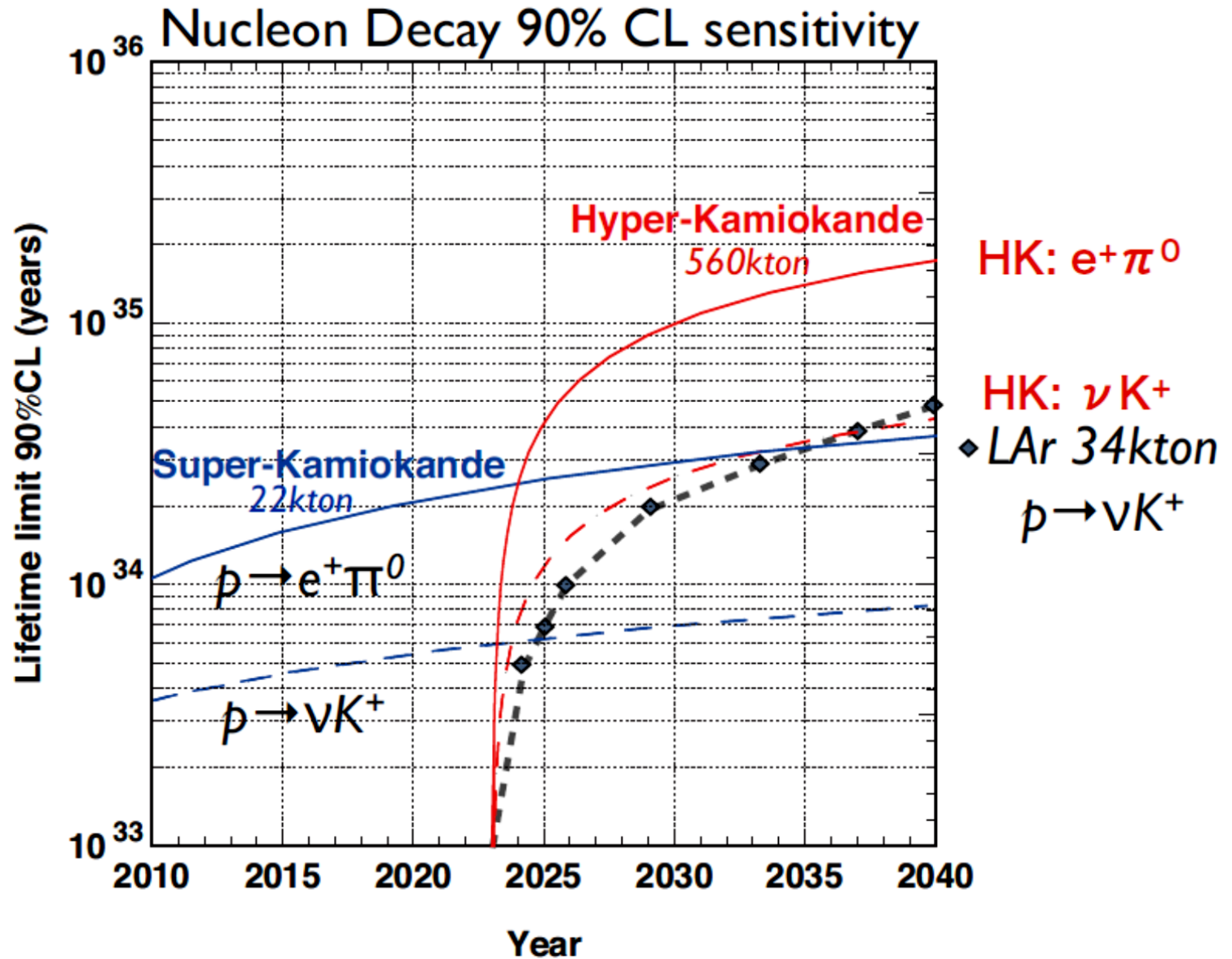
## Excerpts from the P5 report

For a long-baseline oscillation experiment, based on the science Drivers and what is practically achievable in a major step forward, we set as the goal **a mean sensitivity to CP violation of better than  $3\sigma$**  (corresponding to 99.8% confidence level for a detected signal) **over more than 75%** of the range of possible values of the unknown CP-violating phase  $\delta_{CP}$ .

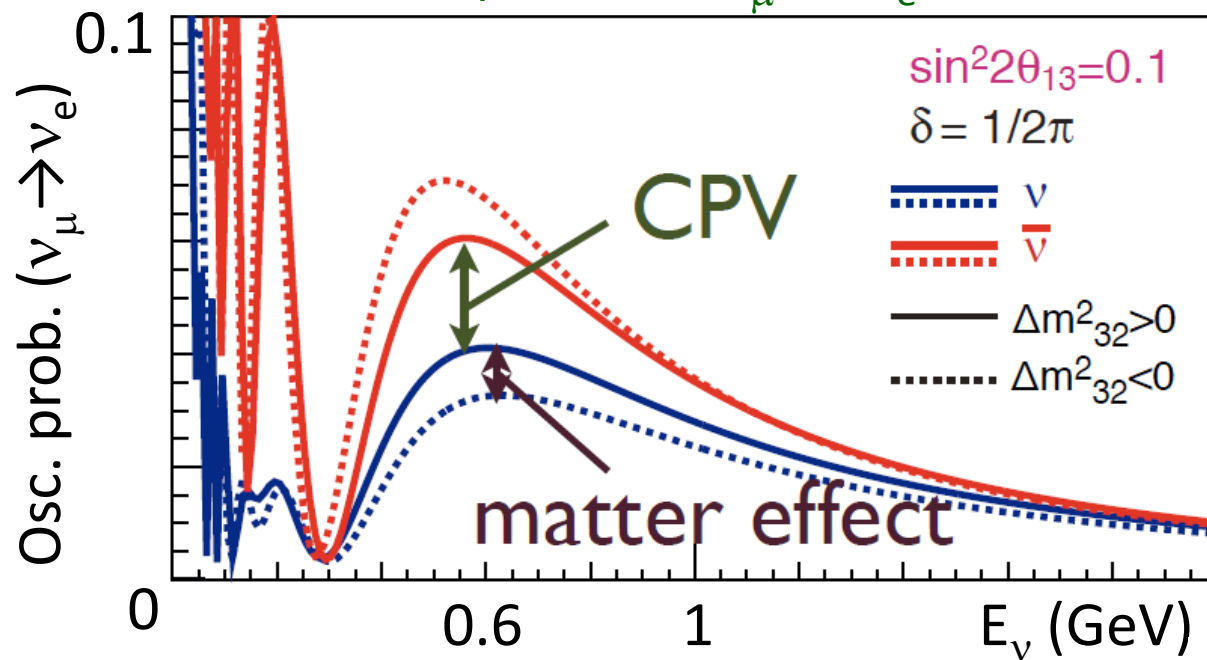




# Sensitivity of proton decays



# Oscillation prob. for $\nu_\mu \rightarrow \nu_e$ @ 295km



Hierarchy is assumed to be known

