

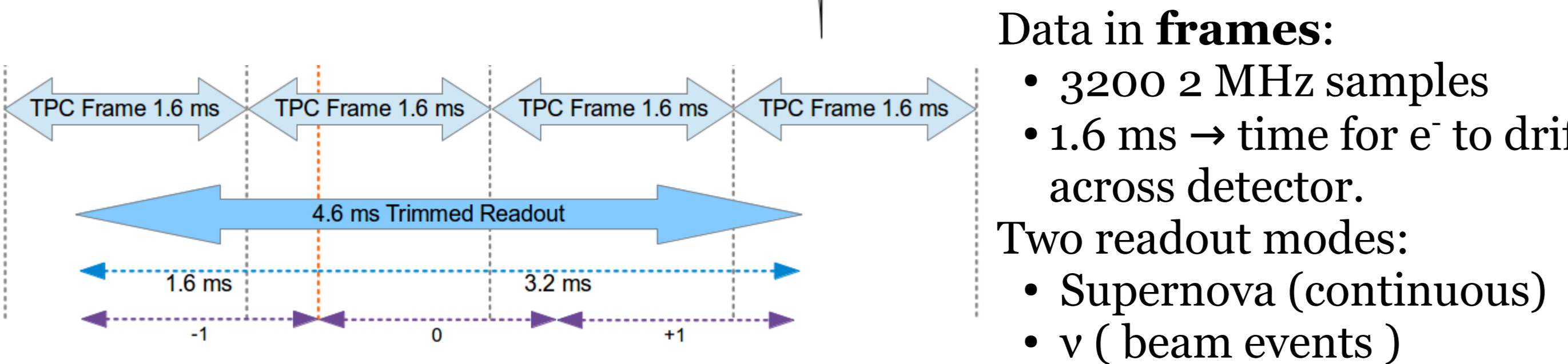
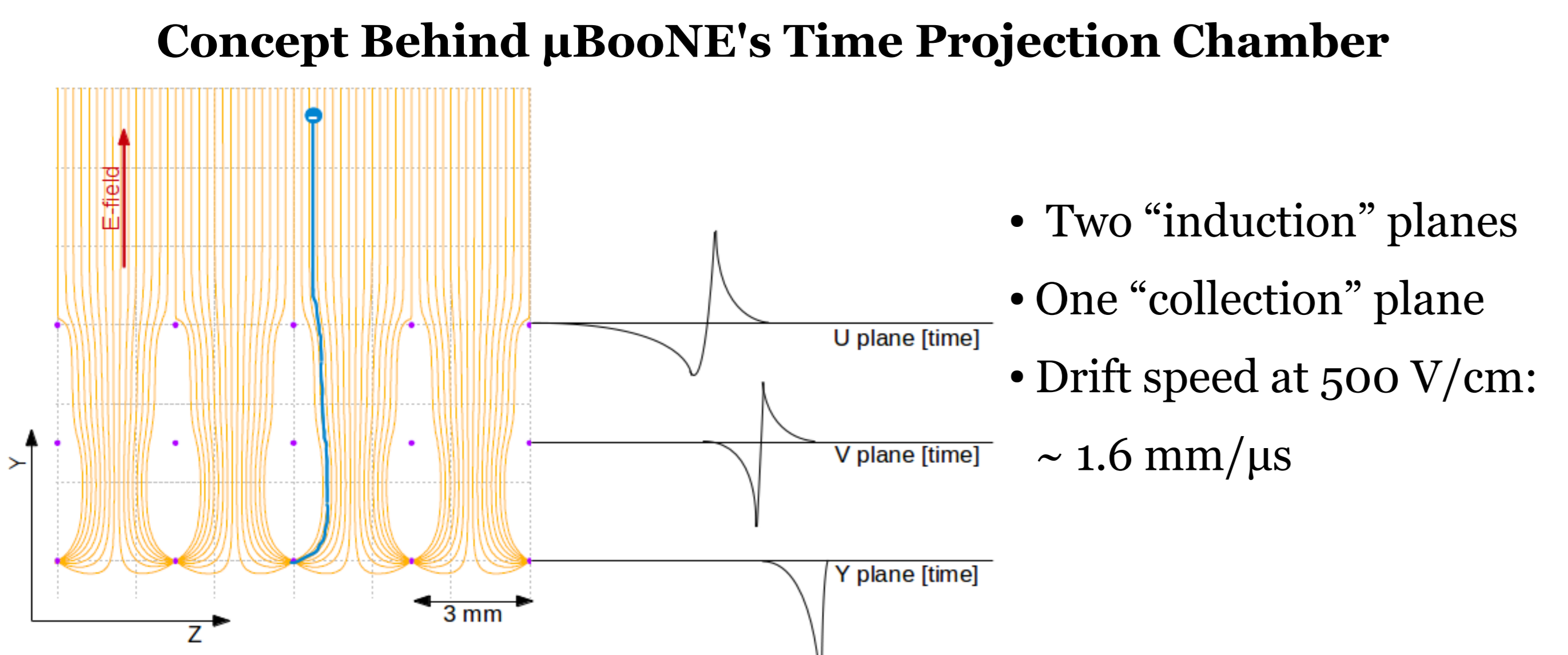
The μ BooNE Experiment

μ BooNE: 170 Ton Liquid Argon Time Projection Chamber neutrino detector

- On Booster Neutrino Beamline at FermiLab (~ 1 GeV neutrinos)
 - R&D - Investigate miniBooNE “low-energy excess” - Cross Sections
 - Potential to observe neutrino events from a supernova within our galaxy
 - TPC will give us morphology of neutrino interactions
 - Signal from TPC and PMTs → powerful and exciting detector!
-
- TPC data will be read out in two modes: “neutrino” and “supernova”:
 - neutrino: trigger on beam events
 - Supernova: record data continuously and wait for alert
→ Recording all data requires lossy data compression

How μ BooNE's TPC Works

- Ionization electrons drift in 500 V/cm field towards three wire-planes.
- Induced current on wires gives us signal:
- Digitize at 2 MHz
 - Three planes + PMTs → 3D track reconstruction
 - Total of 8256 wires. Wire separation is 3 mm.
 - Amount of charge deposited → calorimetric reconstruction & particle ID



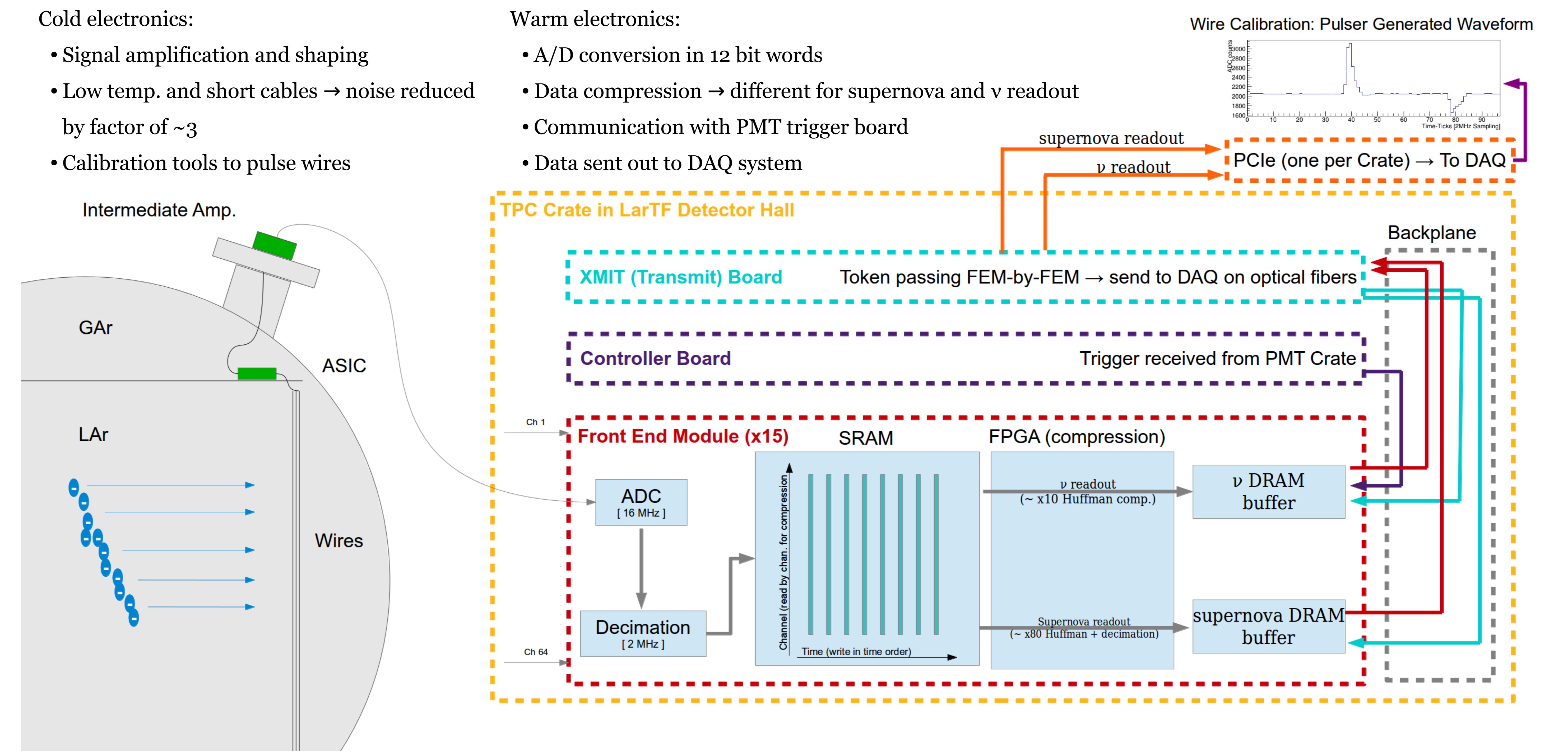
“ ν ” readout mode:

- Store data when a trigger is received from the PMT Crate
→ For more, see poster by David Kaleko at Neutrino14
- Many triggers: Beams, PMTs, Calibration, Laser

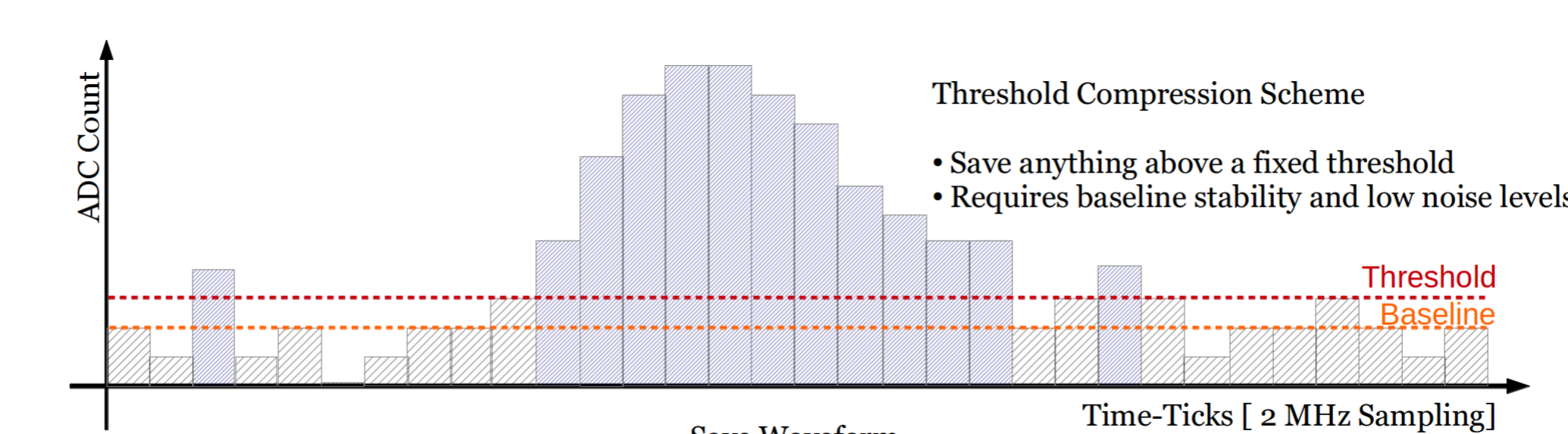
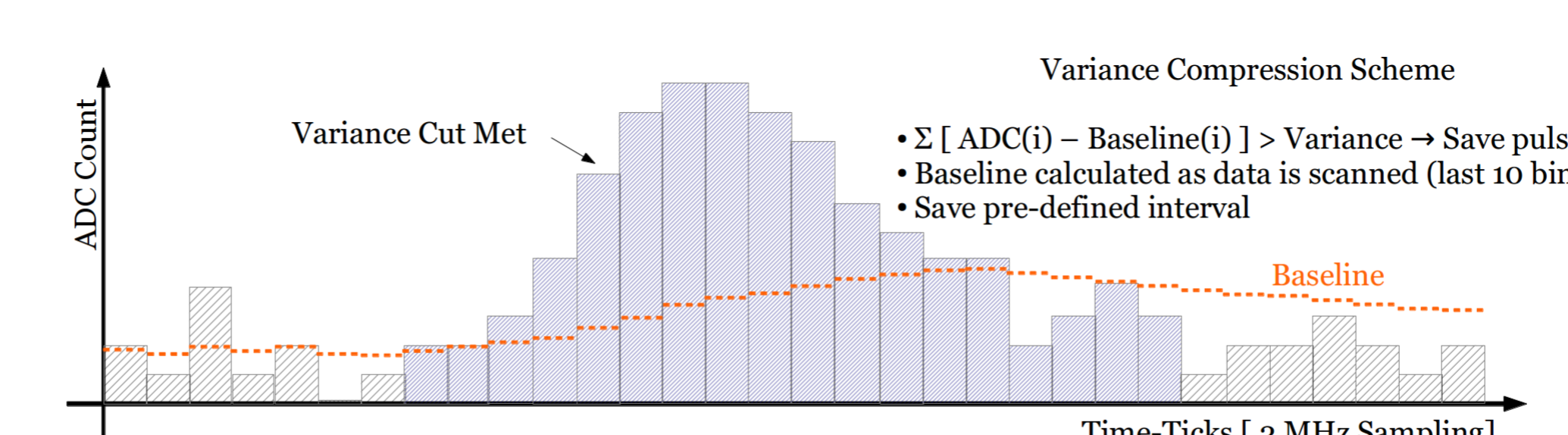
1/2500 Beam spills with ν -event. Require beam + PMT trigger → 1/20

Data Rates:
4.8 ms x 2 MHz samples x 2 bytes/sample x 8256 wires.
BNB (10Hz) → ~ 1.5 GB/s. BNB+PMT trigger → ~ 12 MB/s. Huffman → ~ 1.2 MB/s

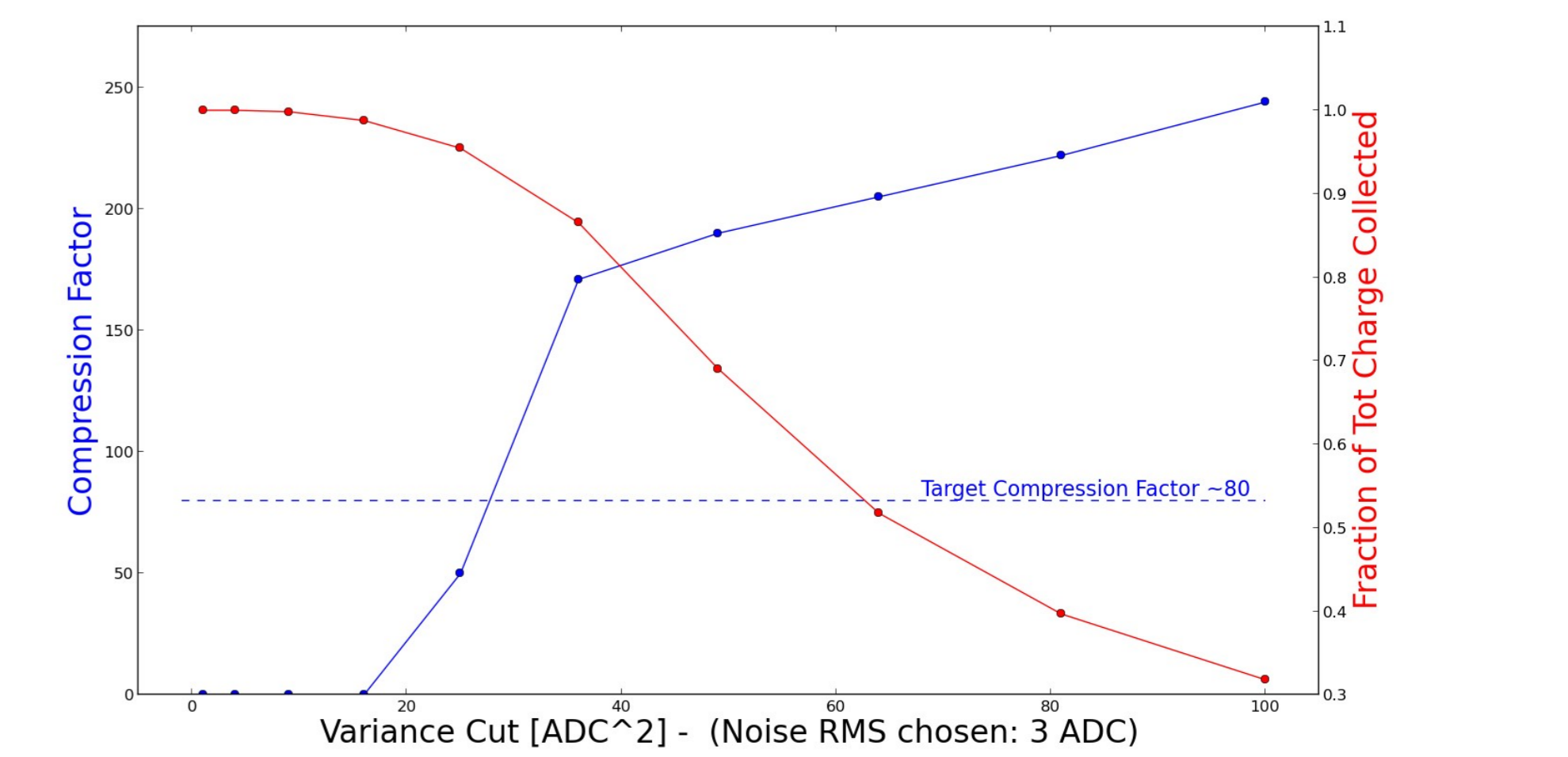
μ BooNE's Time Projection Chamber Readout Chain



Supernova Readout and Data Compression

- Compression “on the fly” → must be fast!
 - Leave out quiet regions & save interesting pulses
 - Interested in 10s of MeV electron tracks from supernova
- Compression Scheme will depend on:
- Baseline stability + Noise levels + Signal Shape
→ Optimize based on Detector
- 
- Threshold Compression Scheme
- Save anything above a fixed threshold
 - Requires baseline stability and low noise levels
- 
- Variance Compression Scheme
- Variance Cut Met
 - $\Sigma [ADC(i) - Baseline(i)] > \text{Variance} \rightarrow \text{Save pulse}$
 - Baseline calculated as data is scanned (last 10 bins)
 - Save pre-defined interval
- μ BooNE will be sensitive to supernova neutrino events
Channel: $\nu_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$
 - Expect ~ 10s of events for a supernova within our galaxy.
 - Hard to trigger on → Instead wait for SuperNova Early Warning System (SNEWS)
- Need to:
- Record data continuously
 - Hold on to it for ~ 1 hour
- How much Data?
Sample @ 2 MHz x 2 bytes/data-word x 8256 wires → 33 GB/sec
 - Must compress by ~ x80 to be within experiment's data-writing limits

Study of Variance Scheme on Simulated Events



- Compression Factor: Tot Data / Saved Data
- Calculated from cosmics (rate ~ 5 KHz)
- Fraction of Charge Collected
- Looked at simulated 10 MeV e^-
- Trade-off → find balance