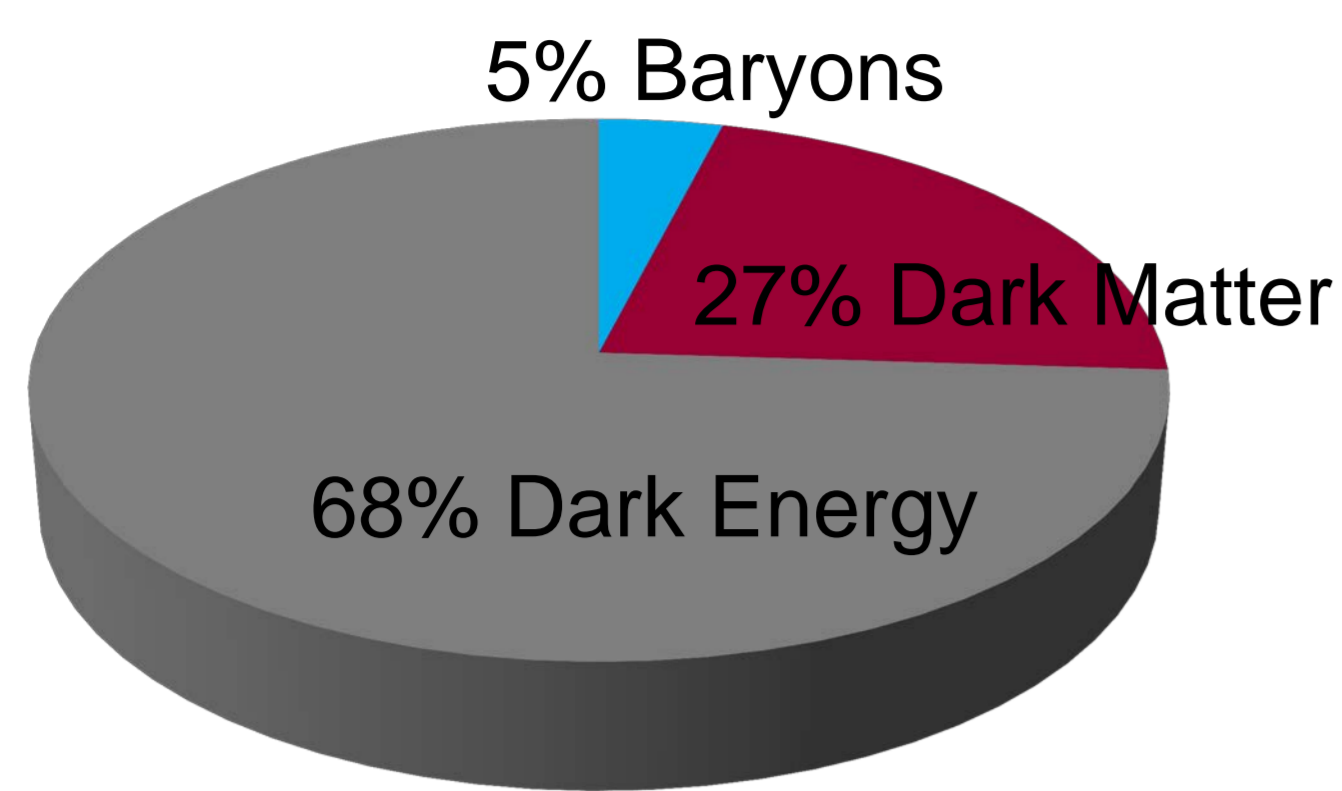


KeV Neutrino Search in Tritium β -Decay

S. Mertens and T. Lasserre

Motivation

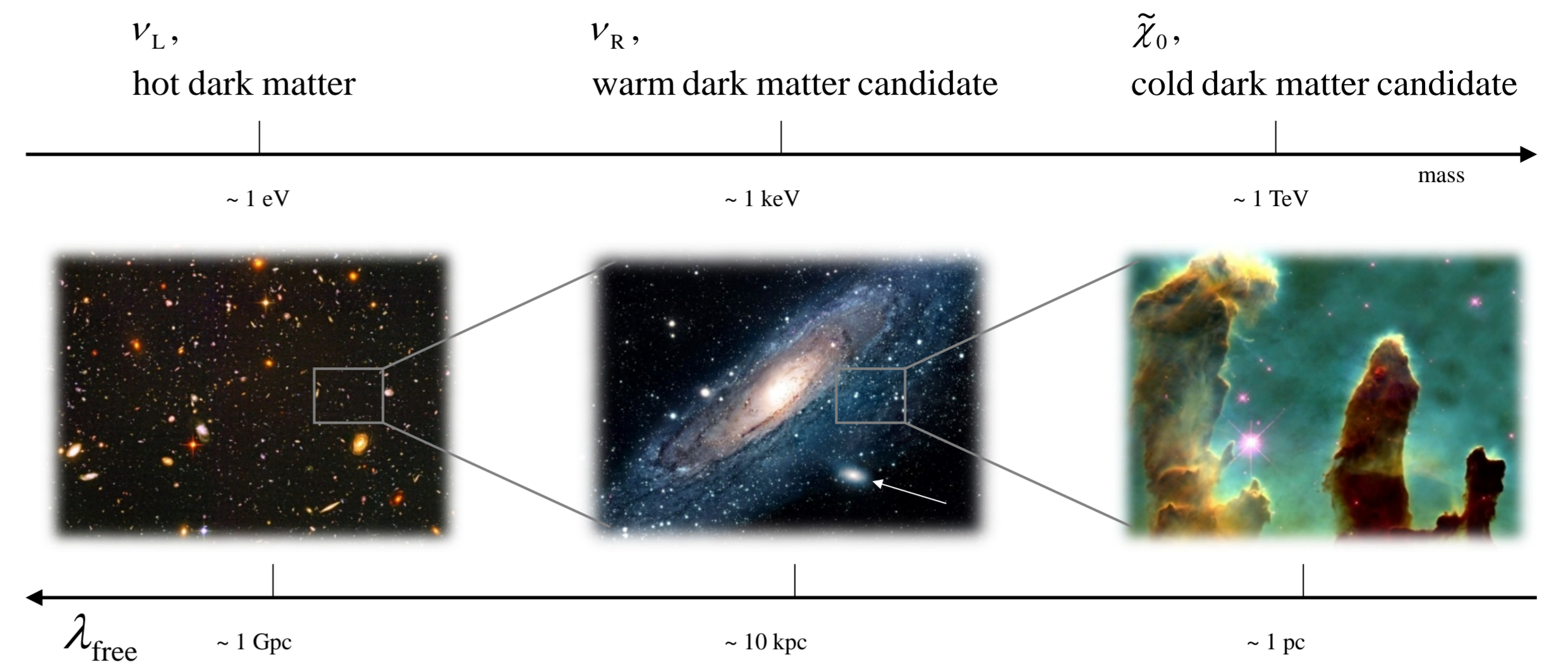


$$\begin{pmatrix} u \\ d \end{pmatrix}_L, \begin{pmatrix} c \\ s \end{pmatrix}_L, \begin{pmatrix} t \\ b \end{pmatrix}_L$$

$$\begin{pmatrix} u_R \\ d_R \end{pmatrix}, \begin{pmatrix} c_R \\ s_R \end{pmatrix}, \begin{pmatrix} t_R \\ b_R \end{pmatrix}$$

$$\begin{pmatrix} \nu_e \\ e \end{pmatrix}_L, \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L, \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L$$

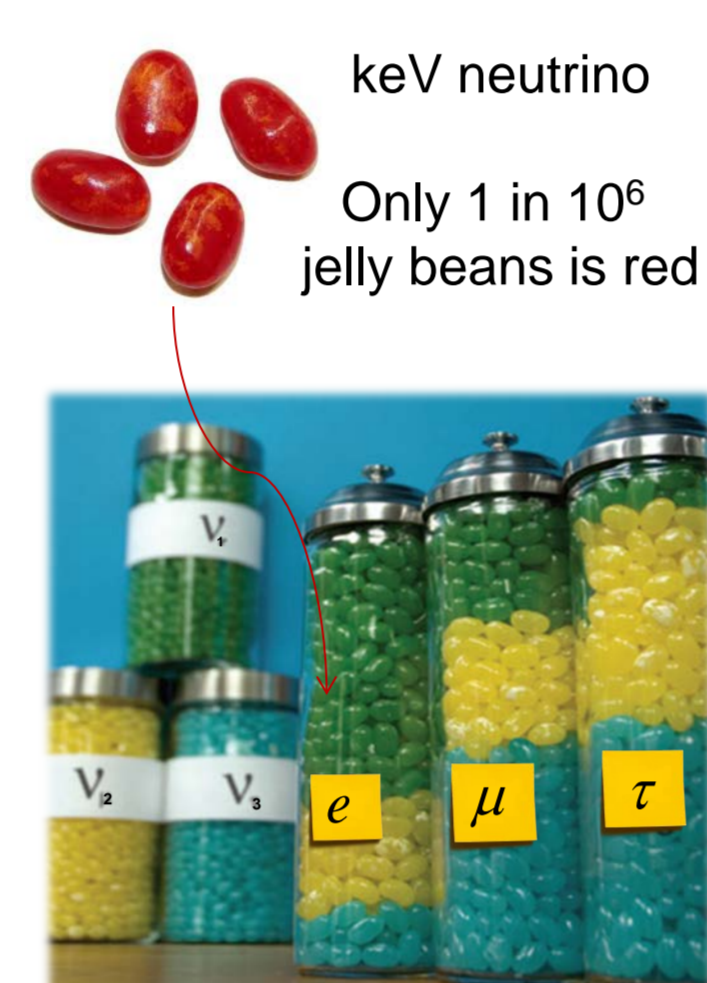
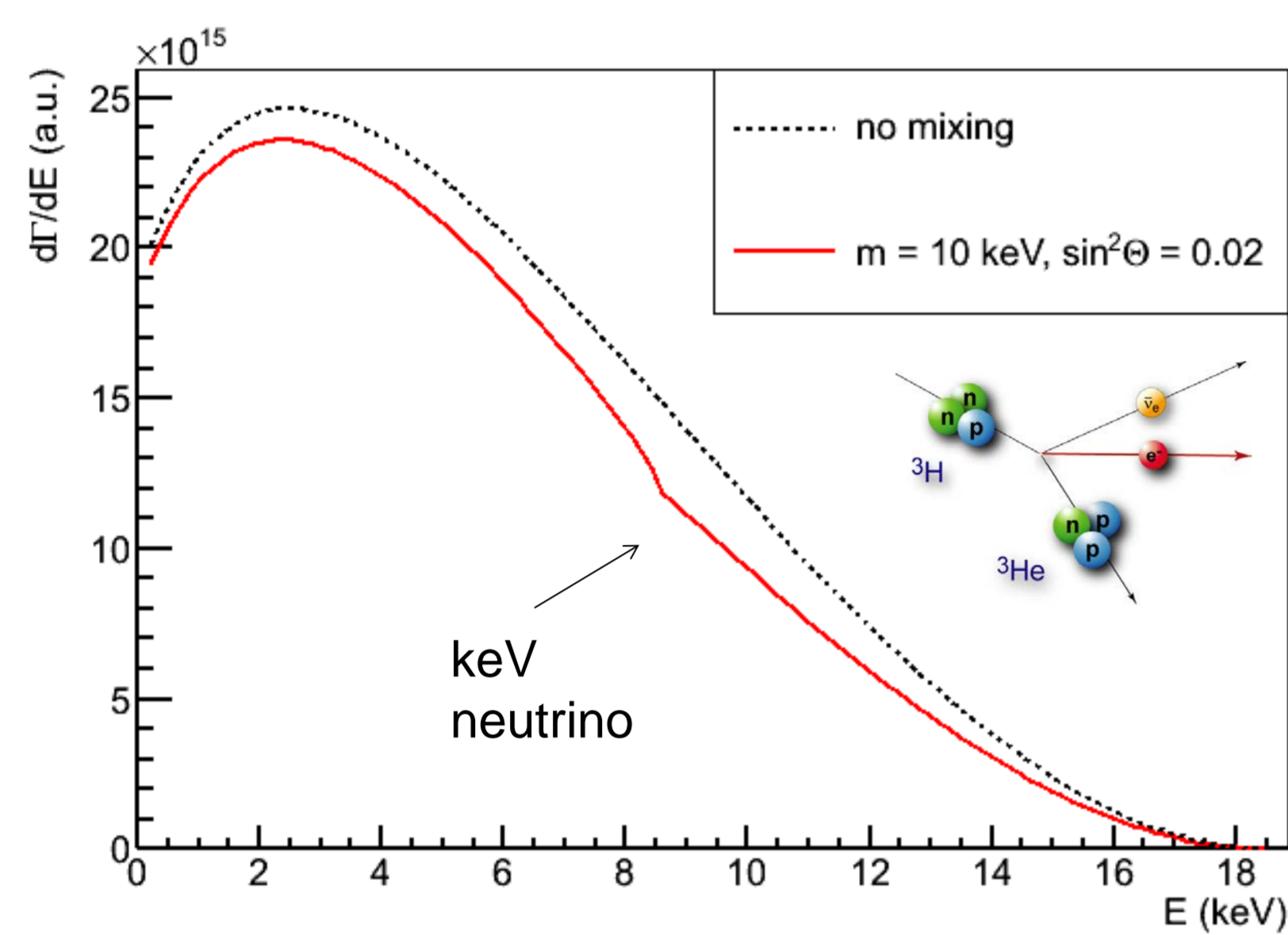
$$\begin{pmatrix} - & - & - \\ e_R & \mu_R & \tau_R \end{pmatrix}$$



Right handed (sterile) neutrinos are theoretically well motivated
Sterile neutrinos can mix with active neutrinos, forming new mass eigenstates
Neutrinos in the keV mass range could contribute to cold and warm dark matter

Relic Neutrinos affect formation of structure in the early universe

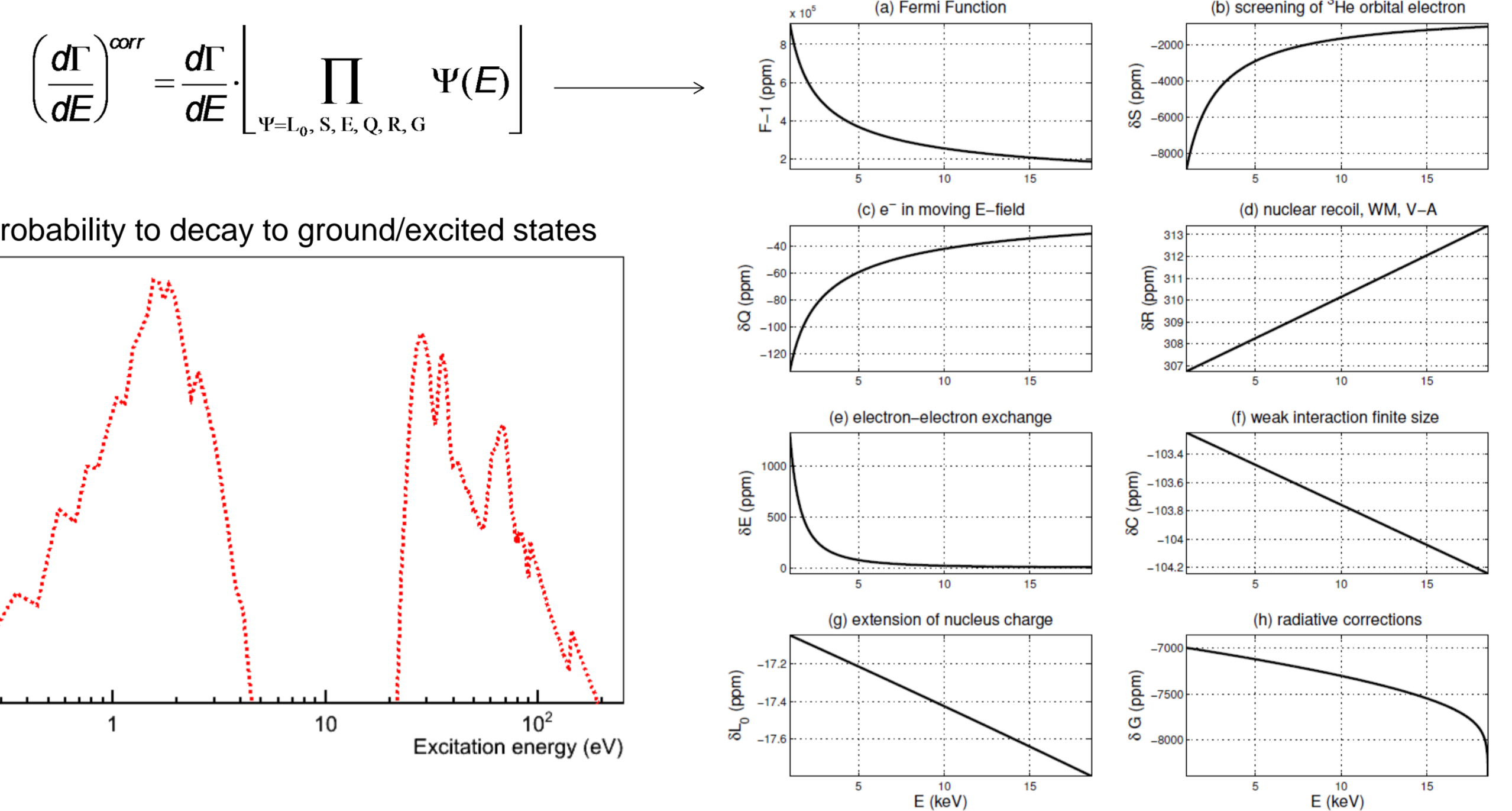
Imprint of keV Neutrinos on β -spectrum



$$\frac{d\Gamma}{dE} \propto \cos^2(\theta) \frac{d\Gamma}{dE}(m_{\nu_{light}}) + \sin^2(\theta) \frac{d\Gamma}{dE}(m_{\nu_{heavy}})$$

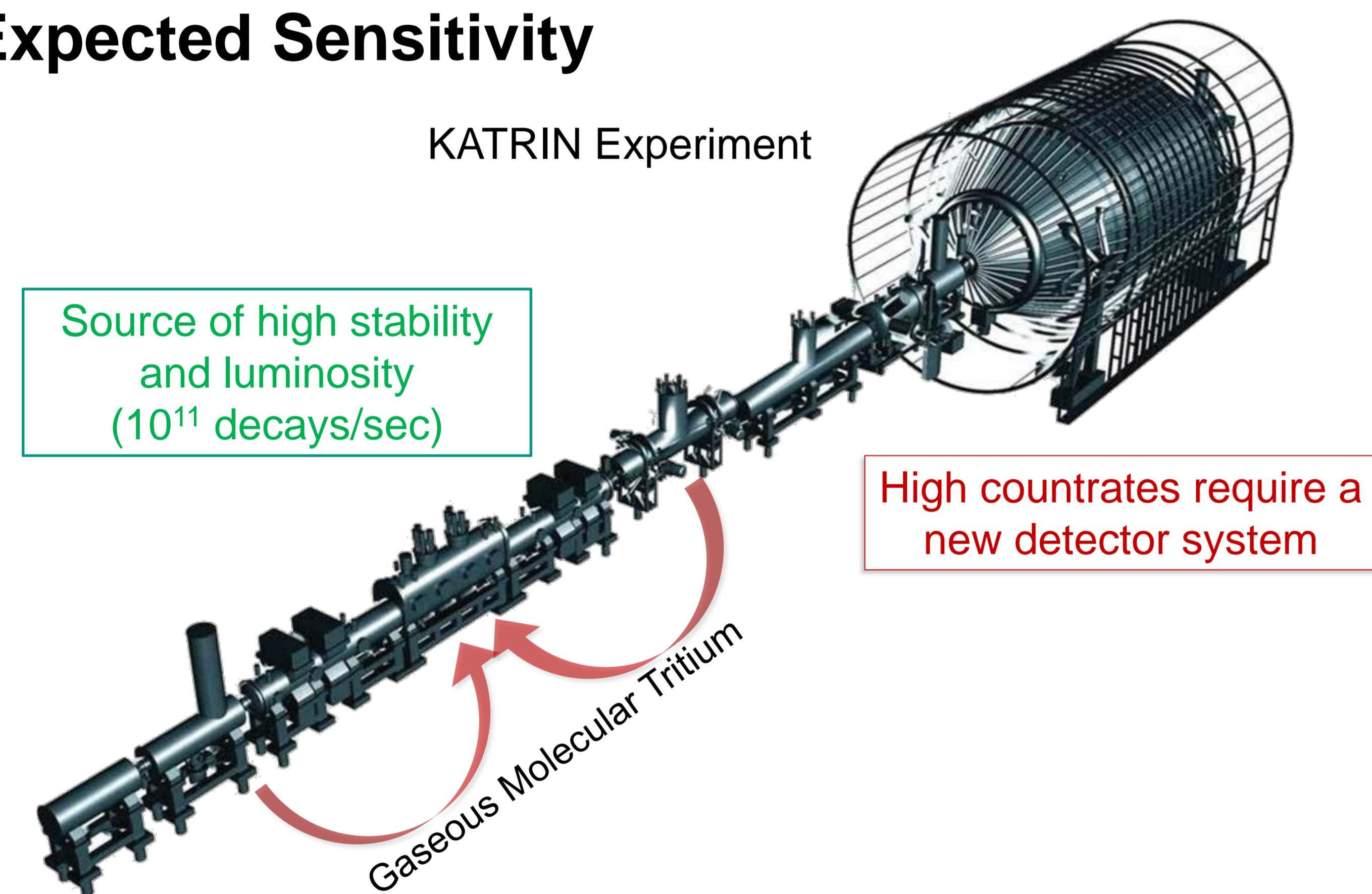
A neutrino in the keV range leaves a kink signature in the β -spectrum
Cosmological constraints: $\sin^2\theta < 10^{-7}$, $2 \text{ keV} < m_{\nu_{heavy}} < 50 \text{ keV}$

Particle, Nuclear, Atomic and Molecular Corrections

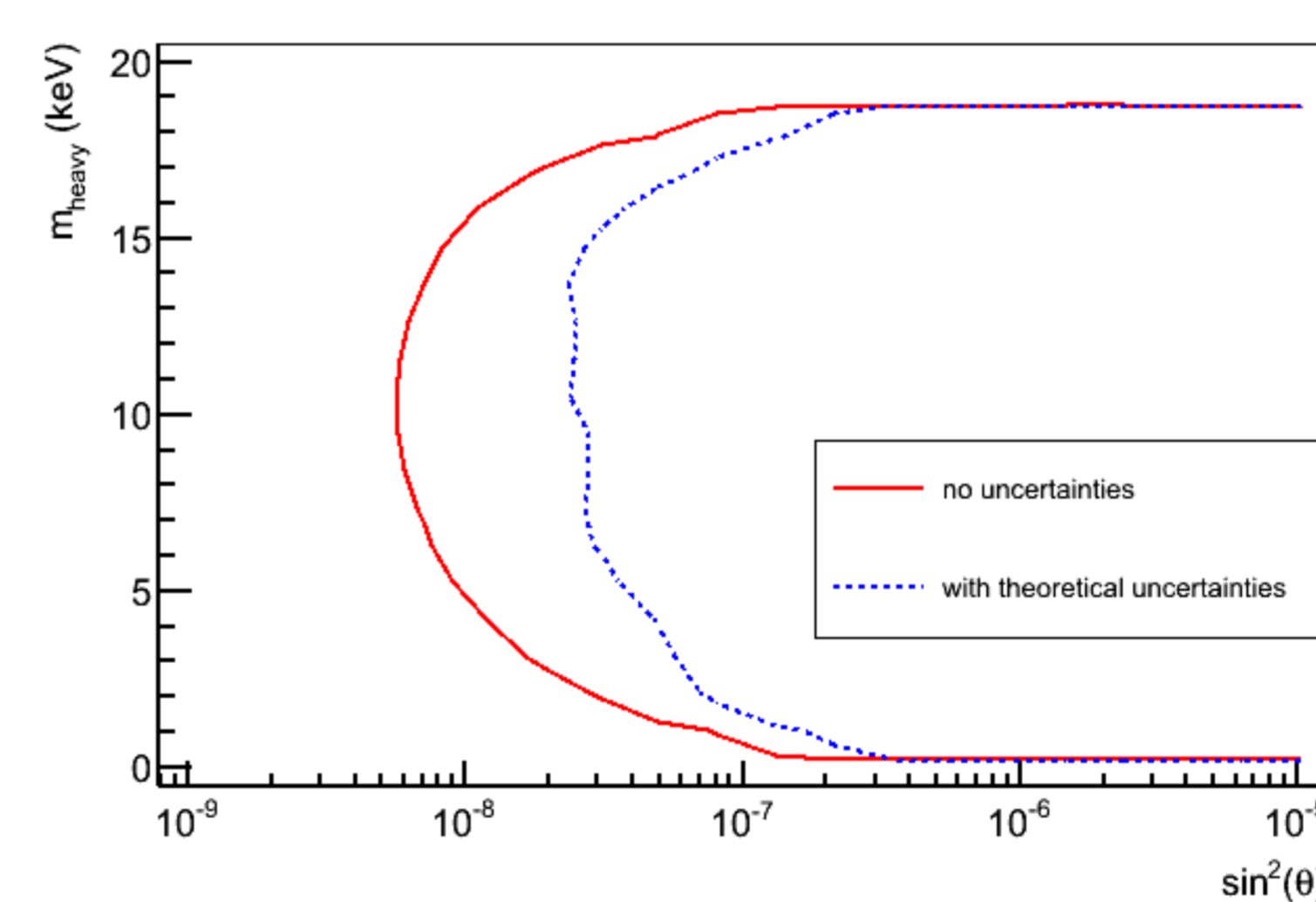


Non-negligible (10-10000 ppm) but smooth corrections
Decay to excited states: largest and least known correction

Expected Sensitivity



Spectral fit approach

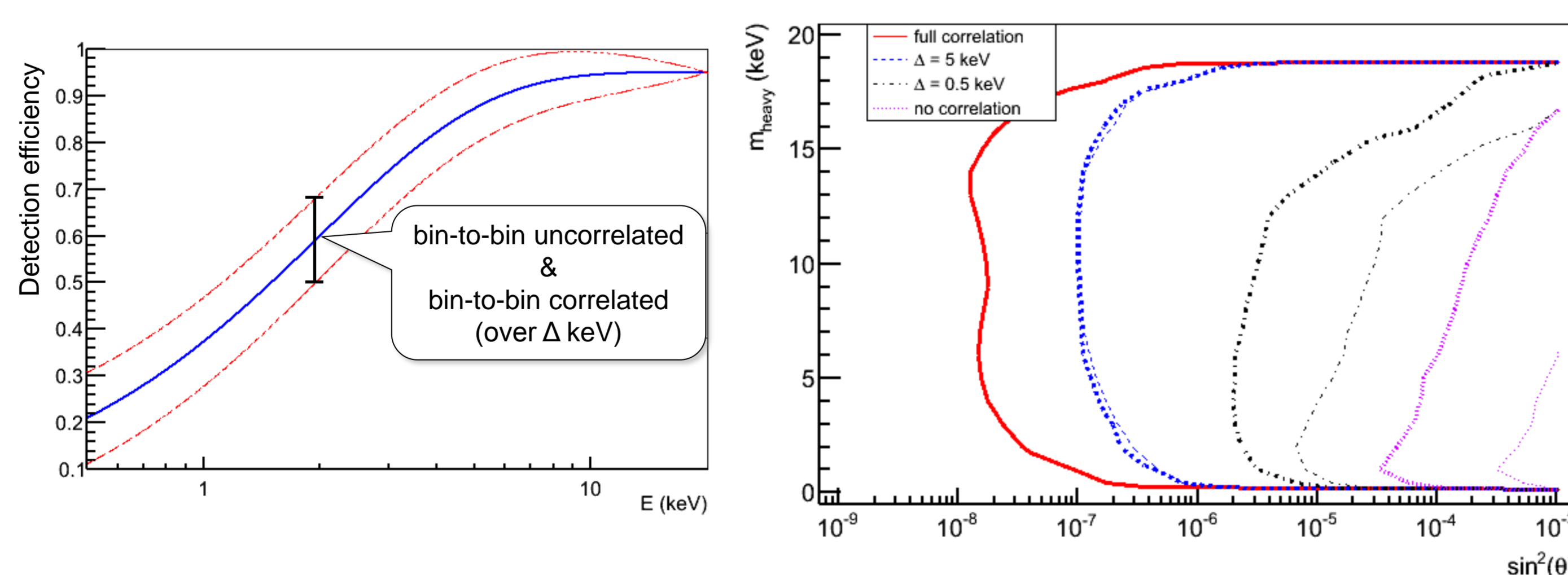


Fit β -decay spectrum

Leave theoretical corrections free to fake a keV neutrino signal

Smooth corrections do not prevent detecting a kink-signature down to $\sin^2\theta > 10^{-7}$

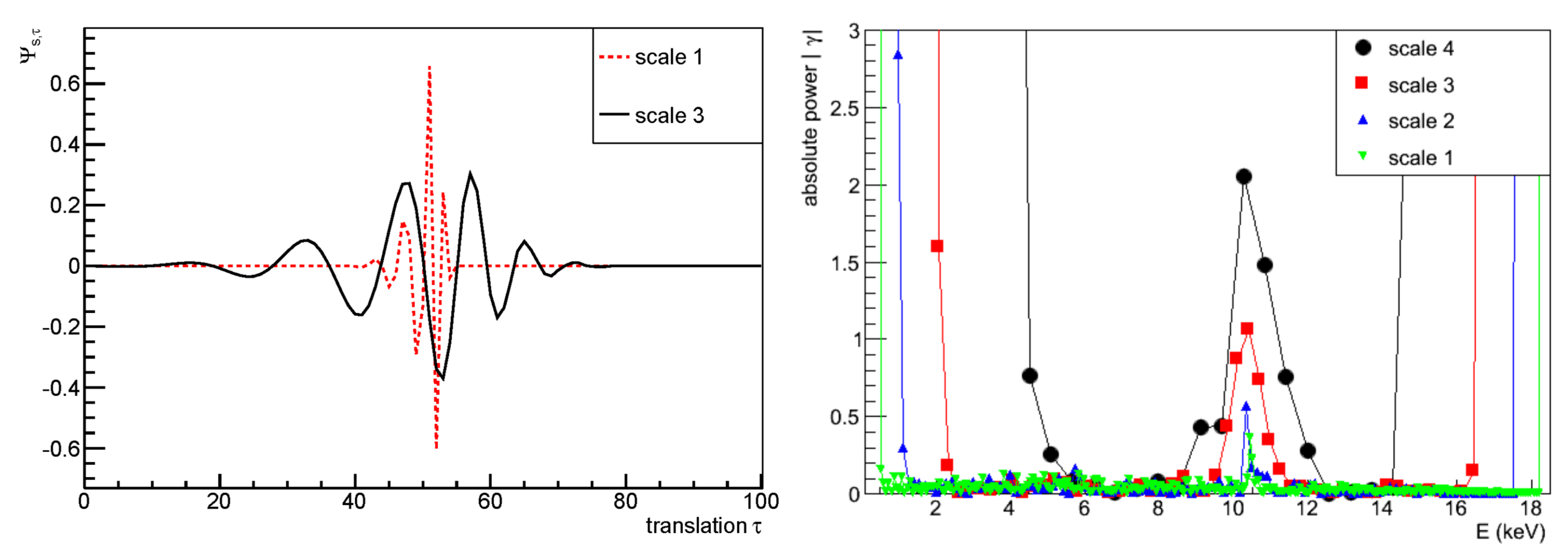
Impact of experimental uncertainties



Construct realistic covariance matrix to investigate experimental uncertainties in a generic way

Uncorrelated errors affect the sensitivity: require calibrations, simulations, and good understanding of correlations

Wavelet approach



Use wavelet transformation to detect „kink“ feature in the spectrum

Kink search independent of exact knowledge of the spectrum shape

Good energy resolution ($\approx 100 \text{ eV}$) required