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## Observation of $\mathrm{v}_{\mu} \rightarrow \mathrm{v}_{\mathrm{T}}$ oscillation with OPERA

On behalf of the OPERA Collaboration
 S.Dusini - INFN Padova

A
XXVI International Conference on Neutrino Physics and Astrophysics


## Neutrino appearance

Neutrino oscillation pioneered via neutrino disappearance Super-KAMIOKANDE, MACRO....
...and for long time the disappearance dominated the scene SK, SNO, MINOS, KamLAND, Borexino....

Hard life for appearance:
Solar scale:
$v_{e} \rightarrow v_{\mu}$ : below threshold for $\mu$ production
Atmospheric scale: experimentally difficult
$v_{\mu} \rightarrow v_{\mathrm{e}}$ : subleading (T2K)
$v_{\mu} \rightarrow v_{T}$ : with cosmic ray neutrinos (SK) statistical separation from large background
$v_{\mu} \rightarrow v_{T}$ : LBL beam neutrinos (OPERA) with tau lepton identification on an event by event basis

The neutrino appearance is a key observation to establish the neutrino oscillation phenomenon.

## CNGS beam

Beam requirements

1) high neutrino energy,
2) long baseline,
3) high beam intensity,


High energy beam optimized to maximize tau production

$$
P\left(v_{\mu} \rightarrow v_{T}\right) \sim O(1) \%
$$

## $\tau$ detection <br> $\nu_{\mu} \xrightarrow{\text { oscillation }}$ $\nu_{\tau}+N \rightarrow \tau^{-}+X$

The separation of the $v_{\tau} C C$ from the dominant $v_{\mu}$ interactions event-by-event, of the peculiar decay topology of the $\tau$.


Hybrid detector
micrometric resolution


Emulsion Cloud Chamber

~150000 bricks (1.25 kton) + electronic detectors


## Collected data and status of the analysis

| Year | P.O.T. <br> $\left(10^{19}\right)$ | SPS <br> Eff. | Beam <br> days | v interactions |
| :--- | :---: | :---: | :---: | :---: |
| 2008 | 1.7 | $61 \%$ | 123 | 1931 |
| 2009 | 3.53 | $73 \%$ | 155 | 4005 |
| 2010 | 4.09 | $80 \%$ | 187 | 4515 |
| 2011 | 4.75 | $79 \%$ | 243 | 5131 |
| 2012 | 3.86 | $82 \%$ | 257 | 3923 |
| Total | 17.97 | $77 \%$ | 965 | 19505 |


$80 \%$ of the design
$\longrightarrow ~ 87 \%$ predicted in the bricks
Scanning strategy
Bricks ordered by the probability to contain the neutrino interaction 2008-2009 : analysis of the $1^{\text {st }}$ and $2^{\text {nd }}$ most probable brick 2010-2012 : analysis of the $1^{\text {st }}$ brick ( $2^{\text {nd }}$ brick postponed)

## Vertex location and decay topology Search

The first two steps of the analysis chain are:

1. location of neutrino interaction
2. search of decay topologies (e.g. large Impact Parameter-IP)



Full MC simulation including all steps of the scanning procedure followed in the scanning labs.


## $v_{e}$ search

## $\mathrm{v}_{\mathrm{e}}$ beam contamination $\sim 0.9 \%$


$v_{e}$ searched in 505 ( $\sim 50 \%$ full statistic) neutrino interaction without the muon in the final state
Extension to full statistic in progress

|  |  | $\mathrm{E}<20 \mathrm{GeV}$ |
| :--- | :---: | :---: |
| Candidate $\mathrm{v}_{\mathrm{e}}$ | 19 | 4 |
| Expected | $19.8 \pm 2.8(\mathrm{sys})$ | 4.6 |
| $\sin ^{2}\left(2 \theta_{\text {new }}\right)<7.2 \times 10^{-3}(90 \% \mathrm{CL})$ |  |  |
| $\sin ^{2}\left(2 \theta_{13}\right)<0.44(90 \% \mathrm{CL})$ |  |  |

## Oscillation results

| variable | $\tau \rightarrow 1 h$ | $\tau \rightarrow 3 h$ | $\tau \rightarrow \mu$ | $\tau \rightarrow e$ |
| :---: | :---: | :---: | :---: | :---: |
| lepton-tag |  | No $\mu$ or $e$ at the primary vertex |  |  |
| $z_{\text {dec }}(\mu \mathrm{m})$ | $[44,2600]$ | $<2600$ | $[44,2600]$ | $<2600$ |
| $p_{T}^{\text {miss }}(\mathrm{GeV} / c)$ | $<1^{\star}$ | $<1^{\star}$ | $/$ | $/$ |
| $\phi_{l H}(\mathrm{rad})$ | $>\pi / 2^{\star}$ | $>\pi / 2^{\star}$ | $/$ | $/$ |
| $p_{T}^{2 r y}(\mathrm{GeV} / c)$ | $>0.6(0.3)^{*}$ | $/$ | $>0.25$ | $>0.1$ |
| $p^{2 r y}(\mathrm{GeV} / c)$ | $>2$ | $>3$ | $>1$ and $<15$ | $>1$ and $<15$ |
| $\theta_{\text {kink }}(\mathrm{mrad})$ | $>20$ | $<500$ | $>20$ | $>20$ |
| $m, m_{\text {min }}\left(\mathrm{GeV} / \mathrm{c}^{2}\right)$ | $/$ | $>0.5$ and $<2$ | $/$ | $/$ |

Kinematical selection cuts kept fixed since beginning of the experiment.

Data sample: 2008/09: $1^{\text {st }}$ and $2^{\text {nd }}$ probable brick 2010/11/12 : $1^{\text {st }}$ probable brick 5522 events analysed

Expected $2.1 \pm 0.4$
$\left(\Delta m_{23}{ }^{2}=2.32 \times 10^{-3} \mathrm{eV}^{2}, \theta_{23}=\pi / 4\right)$ Observed 4

$3^{\mathrm{rd}} \nu_{\tau}$ candidate $(\tau \rightarrow \mu)(2013)$
PHYSICAL REVIEW D 89 (2014) 051102(R)



$$
41
$$

IT upstream module $\underbrace{E}_{x}$

Negative muon measured in the muon spectrometer


Neutrino 201 First measurement of the lepton charge in appearance mode





|  | Values | Selection | Kinematics of |
| :---: | :---: | :---: | :---: |
| P daughter ( $\mathrm{GeV} / \mathrm{c}$ ) | $6.0+2.2$ | >2 |  |
| Kink $P_{+}(\mathrm{GeV} / \mathrm{c})$ | $0.82+0.30$ | $>0.6$ | $3^{\text {rd }} \nu_{\tau}$ candidate$(\tau \longrightarrow h)$ |
| $P_{+}$at 1ry ( $\mathrm{GeV} / \mathrm{c}$ ) | $0.55+0.30$ | < 1.0 |  |
| Phi (degrees) | $166+2$ | $>90$ |  |
| Kink angle (mrad) | $137 \pm 4$ | > 20 |  |
| Decay position (um) | $1090 \pm 30$ | < 2600 | dova 12 |

## Values Selection

Decay position ( $\mu \mathrm{m}$ )
$1090 \pm 30$
< 2600

Kinematics of
$3^{\text {rd }} \nu_{\tau}$ candidate
( $\tau \rightarrow h$ )

## Background to

 $\mu^{-}, e, h, 3 h$
## Production of charmed particles in CC interactions (affect all decay channels) <br> 

MC tuned on CHORUS data (cross section and fragmentation functions), validated with measured charm events in OPERA

FLUKA + test beam data (OPERA bricks exposed to pion beams)

MC tuned on old measurements on lead form factor + dedicated test beam (in progress)

## Data sample:

2008/09: 398 ( $0 \mu$ events) +1553 ( $1 \mu$ events)
2010/11/12: 582 ( $0 \mu$ events) +2153 ( $1 \mu$ events)
The expected signal and background is normalized to the number of located events

$$
n^{0 \mu}\left(\nu_{\tau}^{C C}\right)=\frac{\left\langle\sigma\left(\nu_{\tau}^{C C}\right)\right\rangle}{\left\langle\sigma\left(\nu_{\mu}^{C C}\right)\right\rangle} \frac{\left\langle\epsilon^{0 \mu}\left(\nu_{\tau}^{C C}\right)\right\rangle}{\left\langle\epsilon^{0 \mu}\left(\nu_{\tau}^{C C}\right)\right\rangle+\alpha\left\langle\epsilon^{0 \mu}\left(\nu_{\tau}^{N C}\right)\right\rangle} n^{0 \mu} \quad \alpha=\frac{N C}{C C}
$$

| Decay <br> channel | Expected signal <br> $\Delta \mathrm{m}_{23}{ }^{2}=2.32 \mathrm{meV}^{2}$ | Total <br> background | Observed |
| :--- | :---: | :---: | :---: |
| $\tau \rightarrow \mathrm{h}$ | $0.4 \pm 0.08$ | $0.033 \pm 0.006$ | 2 |
| $\tau \rightarrow 3 \mathrm{~h}$ | $0.57 \pm 0.11$ | $0.155 \pm 0.03$ | 1 |
| $\tau \rightarrow \mu$ | $0.52 \pm 0.1$ | $0.018 \pm 0.007$ | 1 |
| $\tau \rightarrow \mathrm{e}$ | $0.61 \pm 0.12$ | $0.027 \pm 0.005$ | 0 |
| Total | $2.1 \pm 0.42$ | $0.23 \pm 0.04$ | 4 |

Two statistical method:

- Fisher combination of single channel p-value
- Likelihood ratio

$$
p \text {-value }=1.03 \times 10^{-5} \text { of no oscillation }
$$

no oscillation excluded at $4.2 \sigma \mathrm{CL}$

## First measurement of $\Delta m^{2}{ }_{32}$ with tau appearance

$N_{\nu_{\tau}} \propto \int \phi(E) \sin ^{2}\left(\frac{\Delta m_{32}^{2} L}{4 E}\right) \epsilon(E) \sigma(E) d E$

$$
\propto\left(\Delta m_{32}^{2}\right)^{2} L^{2} \int \phi(E) \epsilon(E) \frac{\sigma(E)}{E^{2}} d E
$$

OPERA Off-peak
L/<E> ~ $43 \mathrm{Km} / \mathrm{GeV}$
$(\mathrm{L} /\langle E\rangle)_{\text {peak }} \sim 500 \mathrm{Km} / \mathrm{GeV}$
strong dependence on $\Delta m^{2} \rightarrow$ measure $\Delta m^{2}$ with counting experiment
$90 \% \mathrm{CL}$ intervals on $\Delta \mathrm{m}^{2}{ }_{32}$ assuming $\sin 2\left(2 \theta_{23}\right)=1$
Feldman\&Cousin

$$
[1.8-5] \times 10^{-3} \mathrm{eV}^{2}
$$



## Sterile neutrinos

Tau appearance in the presence of sterile neutrino (3+1)
Solar driven oscillation neglected $\Delta_{21} \sim 0$

$$
\begin{array}{rr}
\sim \text { standard oscillation } & \text { pure exotic oscillation } \\
\boldsymbol{P}_{\nu_{\mu} \rightarrow \nu_{\tau}}=4\left|U_{\mu 3}\right|^{2}\left|U_{\tau 3}\right|^{2} \sin ^{2} \frac{\Delta_{31}}{2}+4\left|U_{\mu 4}\right|^{2}\left|U_{\tau 4}\right|^{2} \sin ^{2} \frac{\Delta_{41}}{2}
\end{array}
$$

Profile likelihood using Tau rate only

$$
\Delta m_{32}^{2}=2.32 \times 10^{-3} \mathrm{eV}^{2}
$$

$90 \% \mathrm{CL}$ bounds on $\mathrm{U}_{\tau 4}$ and $\mathrm{U}_{\mu 4}$
interference
terms


Two extreme values $(\pi / 2,3 \pi / 2)$ of


$$
\begin{aligned}
& +8 \Re\left[U_{\mu 4}^{*} U_{\tau 4} U_{\mu 3} U_{\tau 3}^{*}\right] \sin ^{2} \frac{\Delta_{31}}{2} \sin ^{2} \frac{\Delta_{41}}{2} \\
& +4 \Im\left[U_{\mu 4}^{*} U_{\tau 4} U_{\mu 3} U_{\tau 3}^{*}\right] \sin \Delta_{31} \sin ^{2} \frac{\Delta_{41}}{2}
\end{aligned}
$$




Choosing a particular representation (same as MINOS)

$$
\boldsymbol{U}=\boldsymbol{R}_{\mathbf{3 4}}\left(\boldsymbol{\theta}_{\mathbf{3 4}}\right) \boldsymbol{R}_{\mathbf{2 4}}\left(\boldsymbol{\theta}_{\mathbf{2 4}, \boldsymbol{\delta}_{\mathbf{2}}}\right) \boldsymbol{R}_{14}\left(\theta_{14}\right) \boldsymbol{R}_{\mathbf{2 3}}\left(\boldsymbol{\theta}_{\mathbf{2 3}}\right) \boldsymbol{R}_{\mathbf{1 3}}\left(\boldsymbol{\theta}_{\mathbf{1 3}}, \boldsymbol{\delta}_{\mathbf{1}}\right) \boldsymbol{R}_{12}\left(\theta_{12}, \delta_{3}\right)
$$

$\Delta_{21} \sim 0$ (solar oscillation) $\mathrm{s}_{14} \sim 0$ (reactor anomaly)
$\rightarrow \delta_{1}=0$

$$
U=\left[\begin{array}{lll}
U_{e 1} & U_{e 2} & - \\
U_{\mu 1} & U_{\mu 2} & -s_{14} s_{13} e^{-i \delta_{1}} \frac{c_{14} s_{13}}{s_{24} e^{-i \delta_{2}}+c_{13} s_{23} c_{24}}-\frac{s_{14}}{c_{14} s_{24}} e^{-i \delta_{2}} \\
U_{\tau 1} & U_{\tau 2} & -s_{14} c_{24} s_{34} s_{13} e^{-i \delta_{1}}-c_{13} s_{23} s_{34} s_{24} e^{i \delta_{2}}+c_{13} c_{23} c_{34} \\
U_{s 1} & U_{s 2} & -s_{14} c_{24} c_{34} s_{13} e^{-i \delta_{1}}-c_{13} s_{23} c_{34} s_{24} e^{c_{2}}-c_{13} c_{23} c_{24} s_{34} \\
c_{14} c_{24} c_{34}
\end{array}\right]
$$

Effective mixing



$$
\Delta m_{41}^{2}=1 \mathrm{eV}^{2}
$$



## Conclusions

- OPERA has recorded neutrino interaction equivalent to $\sim 1.8 \times 10^{20}$ pot delivered by CNGS beam from 2008 to 2012 ( $80 \%$ of nominal)
- $4 \nu_{\tau}$ candidates observed so far with a background of 0.23 event.
- No oscillation hypothesis excluded at $4.2 \sigma$


## Observation of $\nu_{\tau}$ appearance

- First measurement of $\Delta m^{2} 31=[1.8-5.0] \times 10^{-3} \mathrm{eV}^{2}(90 \% \mathrm{CL})$ for $\sin ^{2}\left(2 \theta_{23}\right)=1$ using neutrino appearance
- Constrain on sterile neutrinos: first limits on $\left|U_{\mu 4}\right|^{2}\left|U_{T 4}\right|^{2}$ from direct measurement of $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillation


## Thank you for your attention



## Backup

## Visible energy of all the candidates

Sum of the momenta of charged particles and $\gamma$ 's measured in emulsion


## Hadron interaction background

Estimated with Fluka MC and validated with test beam data (OPERA bricks exposed to pion beams)

Background to $\tau \rightarrow h=3.09 \times 10^{-5} /$ located events $v$ $\tau \rightarrow 3 \mathrm{~h}=1.5 \times 10^{-5} /$ located events

Hadron interaction rate suppressed by
 search of large angle tracks produced by nuclear fragments




## Large angle muon scattering background



No measurements except an upper limit from scattering on Cu: S.A. Akimenko et al., NIM A243 (1986) 518 ( $<10^{-5}$ in lead). $10^{-5}$ rate used
Plan to revise this number by an experimental measurement with emulsion


## Track features



|  |  | $\Delta \mathrm{Z}(\mu \mathrm{m})$ | $\delta \theta_{\text {RM }}(\mathrm{mrad})$ | IP ( $\mu \mathrm{m}$ ) | IP Resolution ( $\mu \mathrm{m}$ ) | Attachment | $M=0.59_{-0.15}^{+0.20} \mathrm{GeV} / \mathrm{c}^{2}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\gamma 1$ | To 1ry | 676 | 21.9 | 2 | 8 | OK |  |
| $\gamma 2$ | To 1ry | 7176 | 9.2 | 33 | 43 | OK | Not a single $\pi^{0}$ |
|  | To 2ry | 6124 | 9.2 | 267 | 36 | Excluded |  |

## $1^{\text {st }} v_{T}$ candidate $(\tau \rightarrow h)(2010) \bigcirc$ Beam view


$2^{\text {nd }} \nu_{\tau}$ candidate (2012)
JHEP 11 (2013) 036 $(\tau \rightarrow 3 h)$


Kinematics of $3^{\text {rd }} \nu_{\tau}$ candidate $(\tau \rightarrow \mu)$







Kink angle (mrad)
decay length ( $\mu \mathrm{m}$ )

| $\mathrm{P} \mu(\mathrm{GeV} / \mathrm{c})$ | $2.8 \pm 0.2$ |
| :--- | :---: |
| $\mathrm{P}+(\mathrm{MeV} / \mathrm{c})$ | $690 \pm 50$ |
| $\varphi$ (degrees) | $154.5 \pm 1.5$ |

## Search for highly ionizing particles in hadron interactions

Hadron interactions background can be reduced by increasing the detection efficiency of protons and nuclear fragments emitted in the cascade of intra-nuclear interactions and in nuclear evaporation process


Up stream of interaction point
$>$ Specific tool for scanning
$>$ Validation on the test-beam sample of hadronic interactions $>$ No highly ionizing particle found in OPERA $v_{\tau}$ candidate S.Dusini - INFN Padova

## Oscillation Project with Emulsion tRacking Apparatus



