Ab initio methods for nuclei Lecture I

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NuSTEC Training in Neutrino-Nucleus Scattering Physics FNAL, October 21-29, 2014

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Outline

★ Lecture I: Nuclear Many-Body Theory

- Disclaimer
- Basic facts on nuclear forces
- Untying the Gordian knot of Nuclear Physics
- The nuclear hamiltonian
- Introduction to Nuclear Many-Body Theory (NMBT)
- ★ Lecture II: Nucleon Green's function and nuclear response at low to moderate momentum transfer
- ★ Lecture III: Electron and neutrino cross section in the impulse approximation and beyond

Disclaimer

- ★ Bottom line: there is no such thing as a *ab initio* method to describe the properties of atomic nuclei.
- ★ In the low-energy regime, the fundamental theory of strong interactions (QCD) is nearly intractable already at the level required for the description of hadrons, let alone nuclei
- ★ Nuclei are described in terms of effective degrees of freedom, protons and neutrons, and effective interactions, mainly meson exchange processes
- As long as their size is small compared to the relative distance, treating nucleons as individual particles appears to be reasonable



★ Nucleons behave as non relativistic particles, the dynamics of which are described by the hamiltonian

$$H = \sum_{i=1}^{A} \frac{\mathbf{k}_i^2}{2m} + \sum_{i < j=1}^{A} \mathbf{v}_{ij} + \dots ,$$

where \mathbf{v}_{ij} is nucleon-nucleon (NN) interaction potential, and the ellipses refer to the possible occurrence of forces involving more than two nucleons (to be discussed at a later stage)

★ The main qualitative features of the potential v_{ij} can be deduced from nuclear systematics (binding energies, charge-density distributions, energy spectra ...)

[†]Paradigm: a phylosophical or theoretical framework of any kind (Merriam-Webster).

Binding energies and charge-density distributions

★ The observation that the nuclear binding energy per nucleon is roughly the same for A> 20, its value being ~ 8.5 MeV, suggests that the range of the NN interaction is short compared to the nuclear radius.

★ The observation that the charge-density in the nuclear interior is constant and independent of A indicates that the NN forces become strongly repulsive at short distance



- ★ The spectra of mirror nuclei, e.g. ³⁵₁₈Ar and ³⁵₁₇Cl are identical, up to small electromagnetic corrections
- Nuclear forces exhibit *charge independence*, which is a manifestation of a more general property: *isotopic invariance*



- ★ Neglecting the small mass difference, nucleons can be seen as two states of the same particle, the nucleon, specified by their *isospin*, $\tau_3 = \pm 1/2$.
- ★ The force acting between two nucleons depends on the total isospin of the pair, T = 0 or 1, but not on its projection T_3 .

Untying the Gordian[‡] knot of nuclear physics

- ★ In principle, the form of the potential may be accurately determined through a fit to the large database of nuclear properties.
- ★ The calculations needed to obtain these quantities necessarily involve approximations, casting a strong bias on the underlying models of nuclear interactions.
- ★ The inextricable tie between the uncertainty associated with the nuclear hamiltonian and that arising from the solution of the nuclear many-body problem can be severed determining the nuclear hamiltonian from the properties of *exaxtly solvable* few-nucleon systems.



^{*}A metaphor for an apparently intractable problem solved by thinking "out of the box".

The NN force: Yukawa's theory (AD 1935)

- * NN interaction mediated by a particle of mass $\mu \sim 1 \text{ fm}^{-1} = 200 \text{ MeV}$, to be later identified with the π -meson, or pion
- ★ The pion, discovered in 1947, is a pseudoscalar (spin-parity 0⁻) particle of mass $m_{\pi} \approx 140 \text{ MeV}$
- ★ The three charge states of the pion, π^{\pm} and π^{0} , form the isospin triplet π
- * Simplest πN interaction lagrangian compatible with the observation that NN interactions conserve parity

$$\mathcal{L}_Y = ig\overline{N}\gamma^5\tau N\pi$$



The one-pion-exchange (OPE) potential

★ Potential extracted from the non relativistic reduction of the NN amplitude, at 2nd order in \mathcal{L}_Y

$$\begin{aligned} \mathbf{v}_{\pi} &= \frac{g^2}{4m^2} \left(\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2 \right) (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\nabla}) (\boldsymbol{\sigma}_2 \cdot \boldsymbol{\nabla}) \frac{\mathrm{e}^{-m_{\pi}r}}{r} \\ &= \frac{g^2}{(4\pi)^2} \frac{m_{\pi}^3}{4m^2} \frac{1}{3} (\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2) \left\{ \left[(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) + S_{12} \left(1 + \frac{3}{x} + \frac{3}{x^2} \right) \right] \frac{\mathrm{e}^{-x}}{x} \right. \\ &- \frac{4\pi}{m_{\pi}^3} (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) \delta^{(3)}(\mathbf{r}) \right\} , \\ &S_{12} &= \frac{3}{r^2} (\boldsymbol{\sigma}_1 \cdot \mathbf{r}) (\boldsymbol{\sigma}_2 \cdot \mathbf{r}) - (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) , \end{aligned}$$

- ★ Note that the potential is *spin dependent* and *non sperically symmetric*
- ★ For $g^2/4\pi \approx 14$ the above potential provides a reasonable description of NN scattering in states of high angluar momentum, driven by long-range interactions

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Phenomenological potential models

★ Phenomenological potentials describing the *full* NN interaction can be written in the form

$$\mathbf{v} = \mathbf{v}_S + \mathbf{v}_I + \tilde{\mathbf{v}}_{\pi}$$

where \tilde{v}_{π} is the OPE potential, stripped of the δ -function contribution

★ State-of-the-art NN potential models include momentum-dependent and charge-symmetry breaking terms. The widely used ANL v₁₈ potential is written in the form

$$\mathbf{v}_{12} = \sum_{p=1,18} \mathbf{v}^{(p)}(r) O_{12}^{(p)}$$

 $O_{12}^{(p)} = [\mathbb{1}, (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2), S_{12}, \mathbf{L} \cdot \mathbf{S}, \mathbf{L}^2, \mathbf{L}^2(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2), (\mathbf{L} \cdot \mathbf{S})^2] \otimes [\mathbb{1}, (\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2)]$ $[\mathbb{1}, (\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2), S_{12}] \otimes T_{12} , \text{ and } (\boldsymbol{\tau}_{z1} + \boldsymbol{\tau}_{z2})$ $T_{12} = \frac{3}{r^2} (\boldsymbol{\tau}_1 \cdot \mathbf{r}) (\boldsymbol{\tau}_2 \cdot \mathbf{r}) - (\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2)$

Phenomenological approach (continued)

- ★ The phenomenological potentials reproduce the two-nucleon data, for both bound and scattering states, by construction
- ★ Phase shifts extracted from NN scattering data

★ Differential cross section in the proton-neutron channel



The NN potential in the ${}^{1}S_{0}$ channel



★ Chiral perturbation theory provides an alternative scheme, allowing to derive the two- and three-nucleon potentials within a framework preserving the symmetries of QCD.

Three-nucleon interactions

- ★ Interactions involving more two nucleons arise as a consequence of the internal structure of the participating particles
- ★ The main contribution to the three nucleon forces comes from the Fujita-Miyazawa mechanism
- ★ Phenomenological three-nucleon potentials, written in the form

 $V_{ijk} = V_{ijk}^{2\pi} + V_{ijk}^N$

are determined through a fit to the properties of the three-nucleon system



The nuclear many-body problem

★ The starting point for the description of nuclear properties within the Nuclear Many-Body Theory is the solution of the Schrödinger equation

 $H|\Psi_n\rangle = E_n|\Psi_n\rangle$

★ Quantum Monte Carlo results are available for $A \le 12$.



Nuclear Many-Body Theory (NMBT)

★ In principle, more complex calculations (i.e. involving different observables and heavier nuclei) may be performed in perturbation theory, setting

$$H=H_0+H_I,$$

 H_0 being the hamiltonian describing A non interacting nucleons.

★ Problem: due to the nature of the NN potential, the matrix element of the perturbaton between stets belonging to the base of eigenstates of H_0

$\langle n_0 | H_I | m_0 \rangle$, $H_0 | n_0 \rangle = \mathcal{E}_n | n_0 \rangle$

turn out to be *large*. Perturbative expansions are useless.

- ★ Two possible solutions:
 - Redefine the perturbing hamiltonian
 - Redefine the basis

Isospin-symmetric nuclear matter

- ★ Isospin-symmetric nuclear matter (SNM) can be thought of as a giant nucleus, with equal numbers of protons and neutrons interacting through nuclear forces only.
- ★ The understanding of SNM, besides being a useful intermediate step towards the description of real nuclei, is needed to develop realistic models of neutron star matter.
- ★ The calculation of the properties of SNM is greatly simplified by translational invariance
- ★ Basis states

$$|n_0\rangle = \frac{1}{\sqrt{\mathbf{A}!}} \mathrm{det}\{\varphi_{\mathbf{k}\sigma\tau}(\mathbf{r})\} \ , \ \varphi_{\mathbf{k}\sigma\tau}(\mathbf{r}) = \frac{1}{\sqrt{V}} \mathrm{e}^{\mathbf{k}\cdot\mathbf{r}} \chi_{\sigma}\eta_{\tau} ,$$

where V is the normalisation volume, while χ and η are the Pauli spinors belonging to the spin and isospin space, respectively.

 \star In the ground state the momenta of the occupied states fulfill

$$|\mathbf{k}| < k_F = (3\pi^2 \rho/2)^{1/3}$$
, $\rho = A/V$

G-matrix perturbation theory

★ Replace the bare NN potential with the G - matrix, describing NN scattering in the nuclear medium

$$\mathbf{v} \to G(e) = \mathbf{v} - \mathbf{v} \, \frac{Q}{e} \, G = \mathbf{v} \, \Omega$$



- ★ The expansion in powers of matrix elements of the operator $\zeta = 1 \Omega$ turns out to be convergent
- ★ Rate of convergence not fully established
- ★ Treatment of three-nucleon forces involves non trivial problems

Correlated Basis Function (CBF) perturbation theory

★ Replace the basis states of the non interacting system with the set of correlated states

$$|n\rangle = \frac{F|n_0\rangle}{\langle n_0|F^{\dagger}F|n_0\rangle}^{1/2}$$

$$F = S \prod_{j>i} f_{ij} \quad , \quad f_{ij} = \sum_p f^{(p)}(r_{ij})O^{(p)}_{ij} \quad , \quad [f_{ij}, f_{jk}] \neq 0$$

★ Perturbing hamiltonian defined in terms of matrix elements in the correlated basis

 $H = H_0 + H_I$

 $\langle m|H_0|n\rangle = \delta_{mn}\langle m|H|n\rangle$, $\langle m|H_I|n\rangle = (1 - \delta_{mn})\langle m|H|n\rangle$

★ If the correlated states have large overlaps with the true eigenstates of the hamniltonian, the perturbative expansion in powers of H_I is rapidly convergent

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Cluster expansion and FHNC summation scheme

- ★ The calculation of matrix elements of many-body operators between correlated states involves prohibitive difficulties
- ★ The cluster expansion formalism (consider the ground state expectation value of the hamiltonian, as an exmple)

$$\langle 0|H|0\rangle = \frac{k_F^2}{2m} + \sum_n (\Delta E)_n$$

 $(\Delta E)_n$ is the contribution arising from subsystem (clusters) consisting of *n* nucleons

★ The terms of the cluster expansion are represented by diagrams, that can be classified according to their topological structure and summed to all orders solving a set of integral equations, called Fermi-Hyper-Netted-Chain (FHNC) equations ★ The shapes of the correlation functions f^(p)(r) are determined solving a set of Euler-Lagrange equations, resulting from the minimization of the hamiltonian expectation value in the correlated ground state

 $E_V = \min \langle 0|H|0 \rangle \ge E_0$



- ★ CBF has been widely employed to study both structure and dynamics of nuclear matter and nuclei: the available results (to be discussed in the next lectures) include
 - ▷ Dynamic response to scalar and electromagnetic interactions at low to moderate momentum transfer ($q \leq 400 \text{ MeV}$)
 - Green's functions
 - ▷ Electron and neutrino cross sections in the impulse approximation (IA) regime ($q \gtrsim 600 \text{ MeV}$)

- ★ In spite of the fact that no truly *ab initio* approach is available, a consistent description of a variety of nuclear properties can be obtained from approaches based on effective degrees of freedom and effective interactions.
- ★ Highly realistic nuclear hamiltonian can be derived from the analysis of the properties of *exactly solvable* few-nucleon systems.
- ★ The formalism of many-body theory has reached the degree of maturity required for the treatment of nuclear structure and dynamics based on realistic hamiltonian.