# Dark forces and Higgs exotic decays

### Stefania Gori

Perimeter Institute for Theoretical Physics

BSM Higgs Workshop @ LPC Fermilab

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## Dark Matter and the Higgs

- The nature of dark matter is still a mystery
- ◆The WIMP miracle and EWSB point towards the same scale: the electroweak scale
- Maybe the Higgs is (one of) the best way to have access to DM?
- The Higgs can easily couple to DM:  $\frac{1}{\Lambda} H H^\dagger \bar{\chi} \chi, H H^\dagger S^2$

Very much motivated to look for the Higgs decaying to invisible!

#### In good shape:

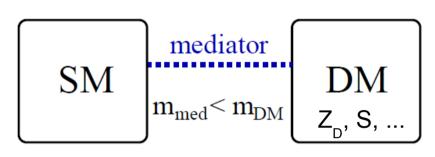
ATLAS 1402.3244, Zh,  $Z\rightarrow II$ , CMS PAS HIG-13-018, Zh,  $Z\rightarrow II$ , CMS PAS HIG-13-028, Zh,  $Z\rightarrow bb$ , CMS-PAS-HIG-13-013 VBF Higgs

Zhou, et al. tth, 1408.0011

## Secluded Dark Matter

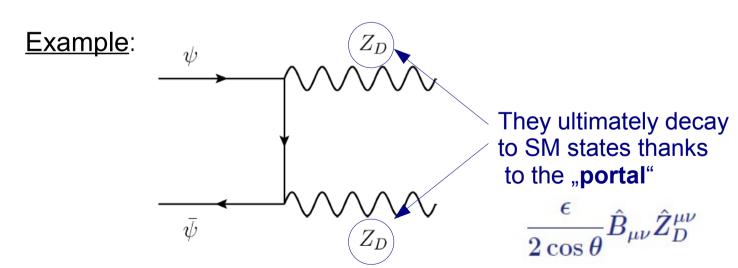
### Something beyond this simple framework?

DM does not interact "directly" with our SM world, but only "indirectly"



Pospelov et.al. 0711.4866 Feldman et al 0702123

Suppression of DM direct detection signals, but still possible to have a thermal DM candidate



How to probe this setup in non-DM experiments? New Higgs measurements?

## Secluded DM & Visible Matter

### Scalar portal

$$V(H,S)\supset \zeta |S|^2|H|^2$$

Arising in extended Higgs sector models: NMSSM, ...

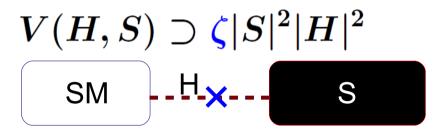
### Vector portal

$$\mathcal{L}\supsetrac{\epsilon}{2\cos heta}\hat{B}_{\mu
u}\hat{Z}_{D}^{\mu
u}$$
SM  $Z_{
u}$ 

Realized e.g. in Hidden valley models, ...

### Secluded DM & Visible Matter

### Scalar portal



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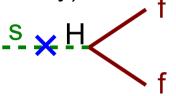
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#### After EWSB:

If also S aquires a VEV, S and H mix and S is unstable  $(\Theta \ge 10^{-6} \text{ to have a})$ 





 $Z_D$  decays **proptly** as long as  $\epsilon \gtrsim O(10^{-4})$ 

### Secluded DM & Visible Matter

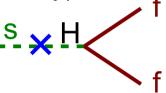
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#### **After EWSB:**

If also S aquires a VEV, S and H mix and S is unstable  $(\Theta \ge 10^{-6} \text{ to have a}$ prompt decay):



Jessie's talk

### Vector portal

$$\mathcal{L} \supset rac{\epsilon}{2\cos heta} \hat{B}_{\mu
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Realized e.g. in Hidden valley models, ...

#### This talk

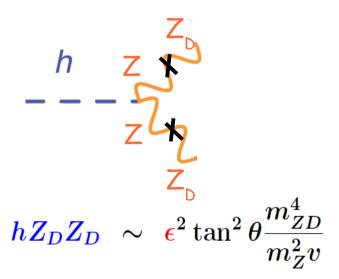


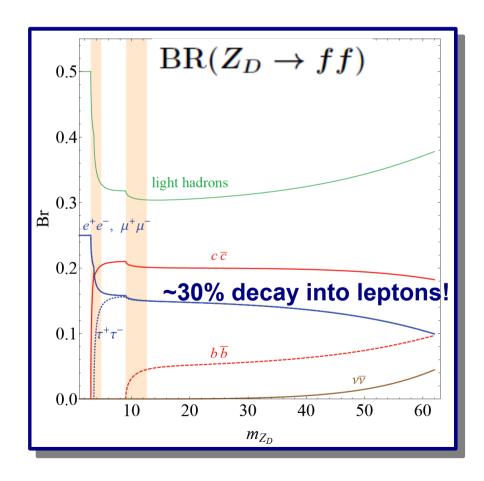
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## The Higgs connection

#### Minimal model:

Kinetic mixing is the only link between the dark and the SM sectors

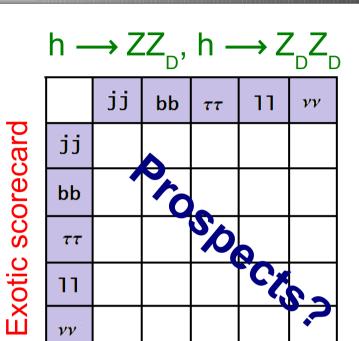




# The final aim

	h –	$\rightarrow$ Z	$Z_{D}$	h –	$\rightarrow \overline{Z}$	
		jj	bb	ττ	11	νν
ard	jj		<u> </u>			
orec	bb		10			
Exotic scorecard	ττ			2.0	2	
otic	11				C	0 -
Ж	νν					•

## The final aim

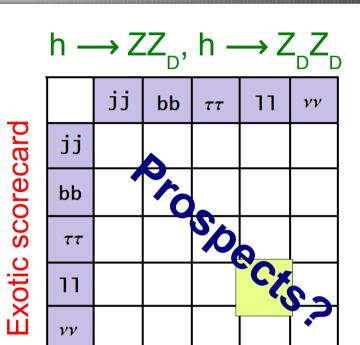


Initial attempt for  $h \longrightarrow Z_D Z_D$  (by theorists)

	Projected/Current			
Decay	$2\sigma$ Limit	Produc-		Limit on
Mode	on $\mathrm{BR}(\mathcal{F}_i)$	tion	$\frac{\mathrm{BR}(\mathcal{F}_{\mathrm{i}})}{\mathrm{BR}(non-SM)}$	$\frac{\sigma}{\sigma_{\rm SM}} \cdot {\rm BR(non\text{-}SM)}$
$ \mathcal{F}_i $	7+8 [14] TeV	Mode		7+8 [14] TeV
jjjj	> 1	W	0.25	> 1
	$[0.1^*]$			$[0.4^*]$
$\ell\ell\ell\ell$	$4 \cdot 10^{-5}$	G	0.09	$4 \cdot 10^{-4}$
$jj\mu\mu$	0.002 - 0.008	G	0.15	0.01 - 0.06
	$[(5-20)\cdot 10^{-4}]$			[0.003 - 0.01]
$bar{b}\mu\mu$	$(2-7)\cdot 10^{-4}$	G	0.015	0.01 - 0.05
	$[(0.6-2)\cdot 10^{-4}]$			[0.003 - 0.01]

Curtin, Essig, SG, Jaiswal, Katz, Liu, Liu, Mckeen, Shelton, Strassler, Surujon, Tweedie, Zhong, 1312.4992

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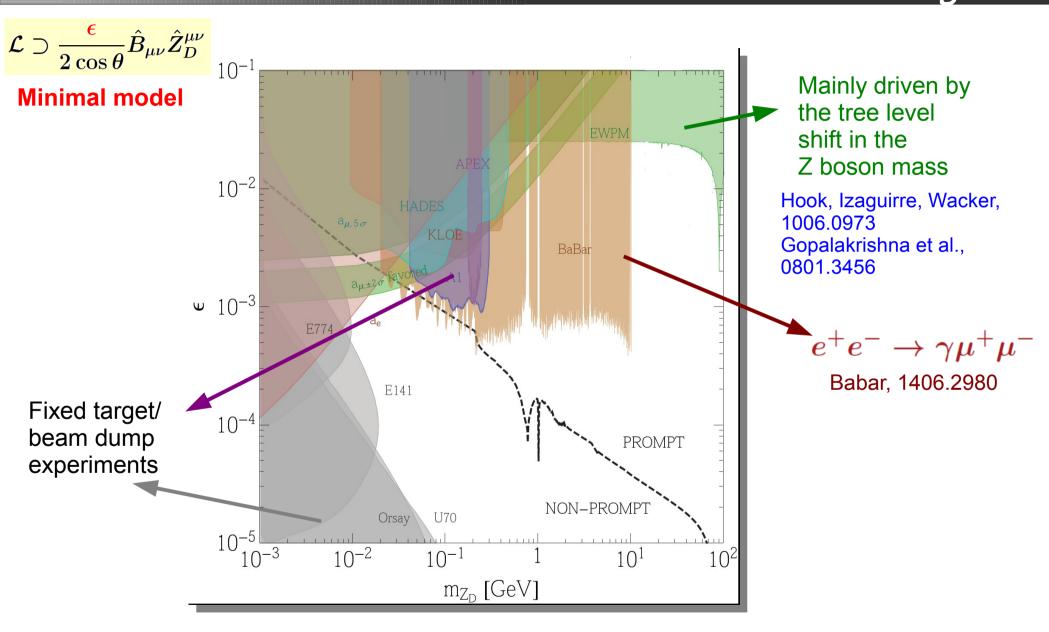


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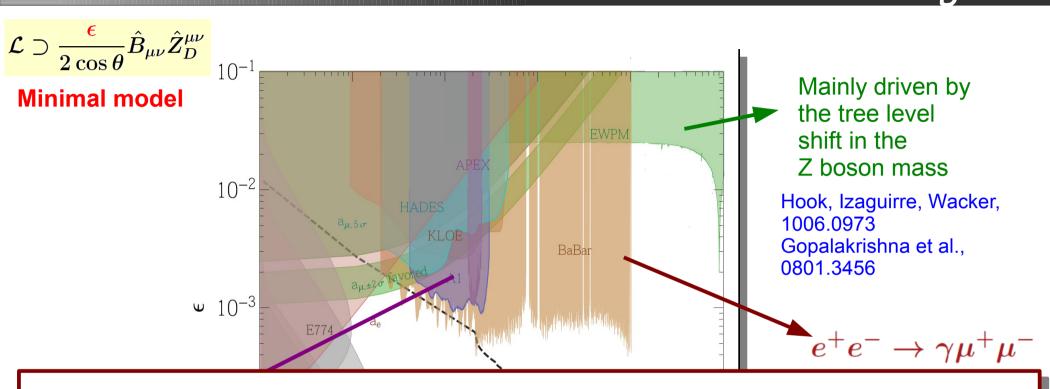
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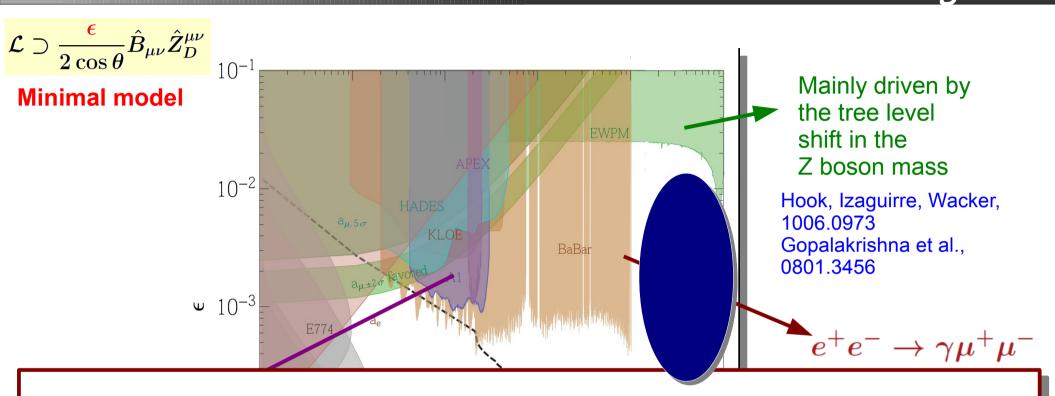


The ultimate goal: probing as much parameter space as possible

A little theory bias: in GUT theories, the kinetic mixing operator is generated at one loop

$$\epsilon \sim rac{g_1 g'}{16 \pi^2} \log \left(rac{M_1}{M_2}
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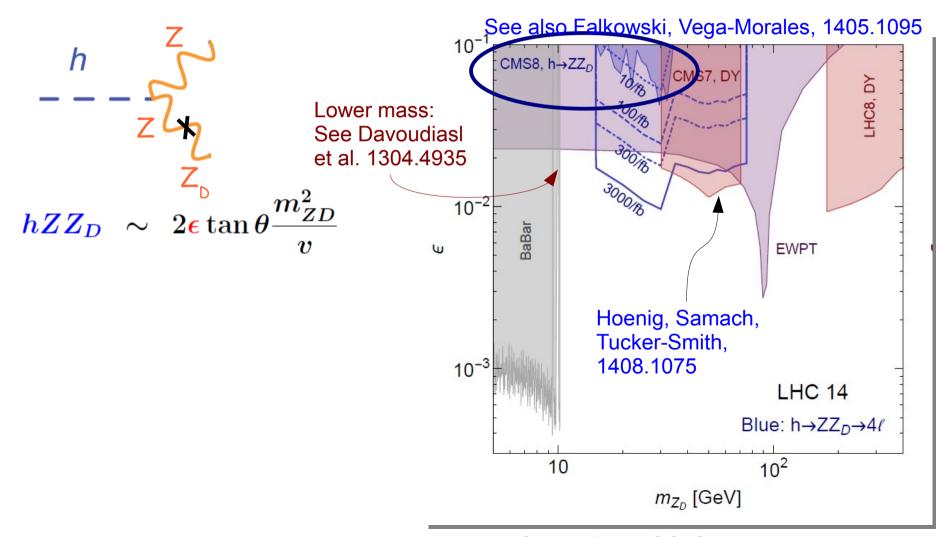
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# The Higgs and the $Z_{D}$ (1)

#### Minimal model:

Kinetic mixing is the only link between the dark and the SM sectors



# The Higgs and the $Z_{\rm D}$ (1)

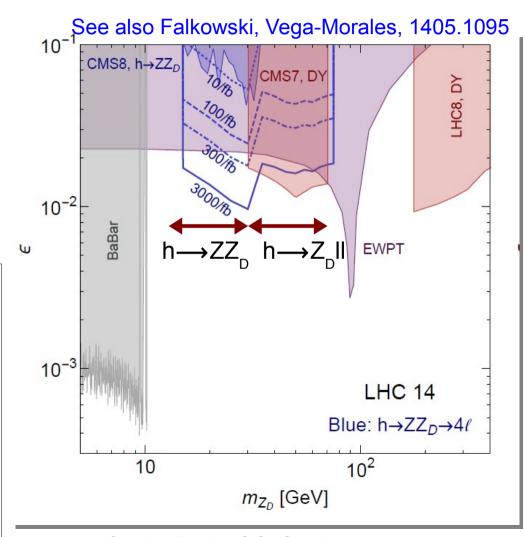
#### Minimal model:

Kinetic mixing is the only link between the dark and the SM sectors

$$120\,{\rm GeV} < m_{4\ell} < 130\,{\rm GeV}$$

Bump hunt in the SFOS dilepton invariant mass most away from m<sub>2</sub>: M<sub>2</sub>

$$|M_2 - m_{Z_D}| < \ \left\{ egin{array}{ll} 0.02 M_2 & ext{(electrons)} \ 2.5 (0.026 \; {
m GeV} + 0.013 M_2) & ext{(muons)} \end{array} 
ight.$$

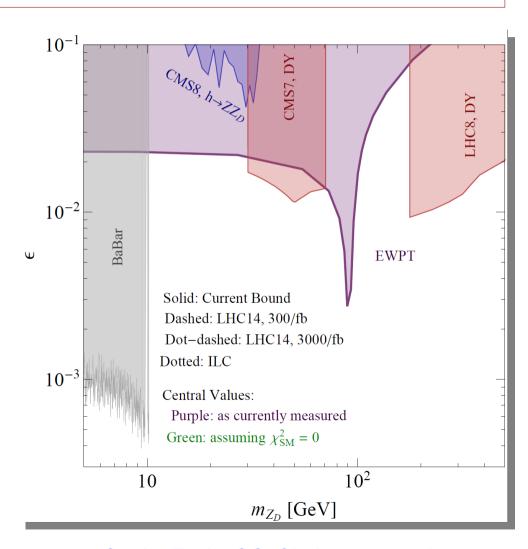


## Complementarity with EWPMs

#### Minimal model:

Kinetic mixing is the only link between the dark and the SM sectors

- Effects on the Z phenomenology!
  - ◆ Tree level shift in the <u>Z mass</u>;
  - Modification of the <u>Z couplings</u>



## Complementarity with EWPMs

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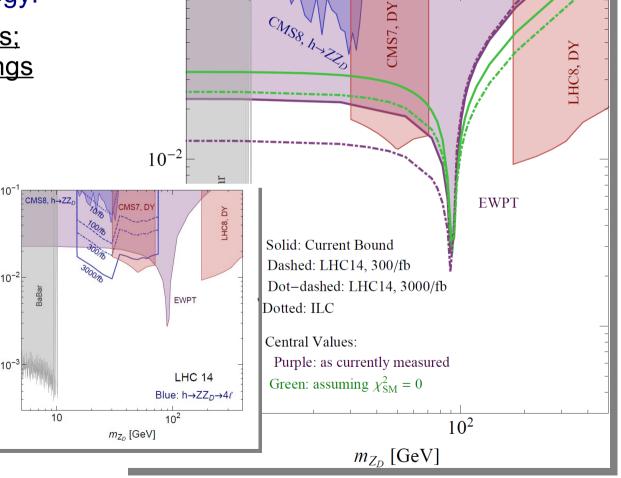
10-

- Effects on the Z phenomenology!
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In the future, a global fit will give stronger constraints

	Now	Future HL-LHC
$\delta m_W$	$15~{ m MeV}$	$5~{ m MeV}$
$\delta m_h$	$0.24~{ m GeV}$	$0.05~{ m GeV}$
$\delta m_t$	$0.76~{ m GeV}$	$0.2~{ m GeV}$
$\delta\Deltalpha_{ m had}^{(5)}$	$10\cdot 10^{-5}$	$4.7\cdot 10^{-5}$
$\sigma\Deltalpha_{ m had}$	10 · 10 °	4.7 • 10

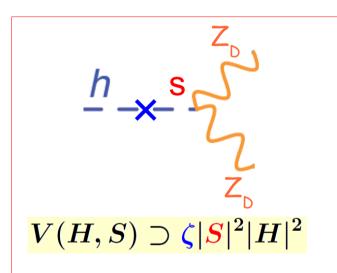
Babar expectation



# The Higgs and the $Z_{\rm D}$ (2)

#### **Next to Minimal model:**

Kinetic mixing and scalar portal are the link between the dark and the SM sectors. The scalar S is responsible of breaking U(1)'



$$\mathcal{L}\supsetrac{m{\epsilon}}{2\cos heta}\hat{B}_{\mu
u}\hat{Z}_{D}^{\mu
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Responsible for the decay of  $Z_D$  back to the SM

#### Free parameters:

$$m_{Z_D},\,m_s,\,\epsilon,\, \kappa'\equiv \zetarac{m_h^2}{|m_h^2-m_s^2|}$$

A very clean 4l signal with two equal mass resonances!

Bound from recasting CMS-ATLAS h→ZZ\*→4I:

$$BR(h \rightarrow Z_D Z_D \rightarrow 4I) \le few*10^{-5}$$
, at best

Curtin, et al. 1312.4992

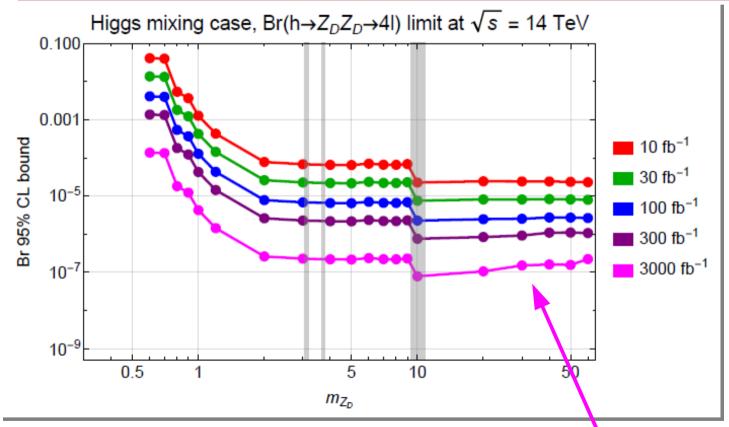
We can for sure do much better!

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# The Higgs and the $Z_{N}$ (3)

#### **Next to Minimal model:**

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HL is crucial since the search is almost background free/ statistically limited

This corresponds to set a bound on the Higgs mixing with another scalar at the level of

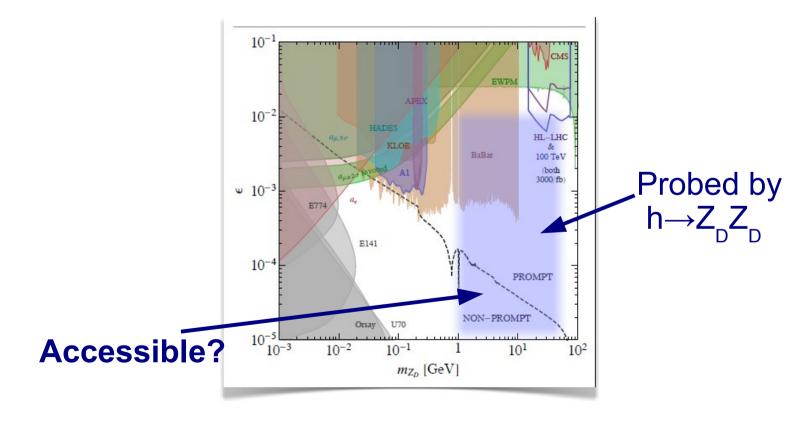
 $\kappa' \sim \text{few} \cdot 10^{-5}$ 

# A non-promptly decaying Z

#### **Next to Minimal model:**

Kinetic mixing and scalar portal are the link between the dark and the SM sectors. The scalar S is responsible of breaking U(1)'

What if the kinetic mixing is very small  $\epsilon \lesssim 10^{-4}-10^{-5}$  and Z<sub>D</sub> does not decay promptly?

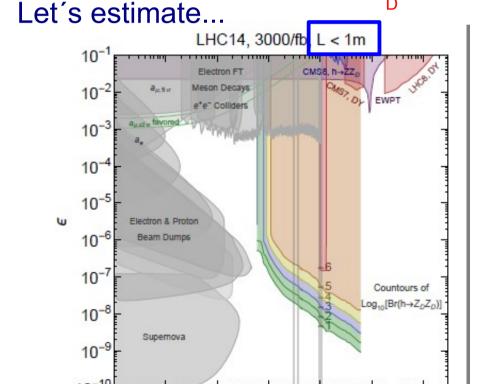


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What if the kinetic mixing is very small  $\epsilon \lesssim 10^{-4}-10^{-5}$  and Z<sub>n</sub> does not decay promptly?



10

 $m_{Z_D}$  [GeV]

A similar signal efficiency should be achievable for macroscopic decay lenghts contained in the inner detector

$$BR_{eff} = BR(h \to Z_D Z_D) P(L, \sqrt{s}, m_{Z_D}, \epsilon)$$

The BR we measure

Probabilty for Z<sub>D</sub> to decay inside the detector

## Going to higher energy?

What we will gain on the Higgs exo. decays going to higher energy?

Huge productions!

	√S=14 TeV	√S=33 TeV	√S=100 TeV
ggF	50.4 pb	178 pb	740 pb
VBF	4.4 pb	17 pb	82 pb
WH	1.6 pb	4.7 pb	16 pb
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НН	.034 pb	.2 pb	1 pb

Higgs cross section working group

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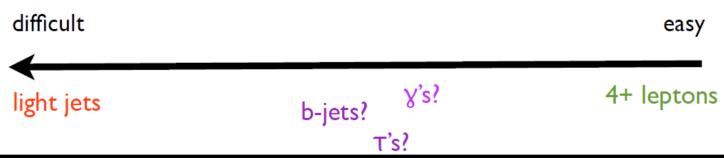
#### 1. Difficult decay modes

Benefit from having "accessible" Higgs production in association with tops, Z bosons, ...

#### 2. Clean decay modes

At the high-lumi LHC, we cannot expect to be able to put bounds on  $BR_{exo} \le 10^{-6}-10^{-7}$ 

because of the lack of statistics



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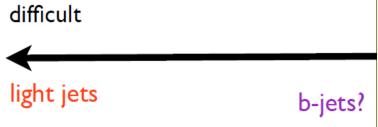
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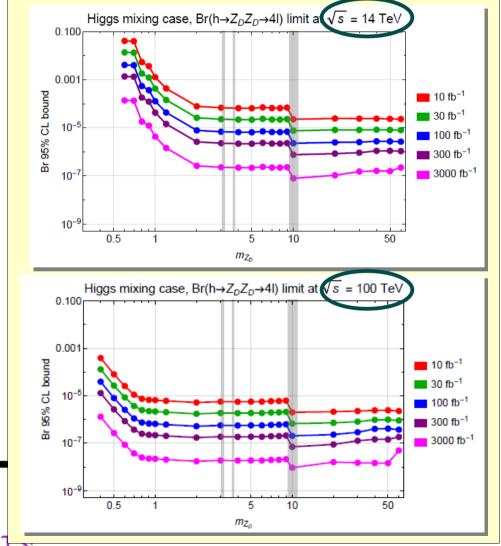
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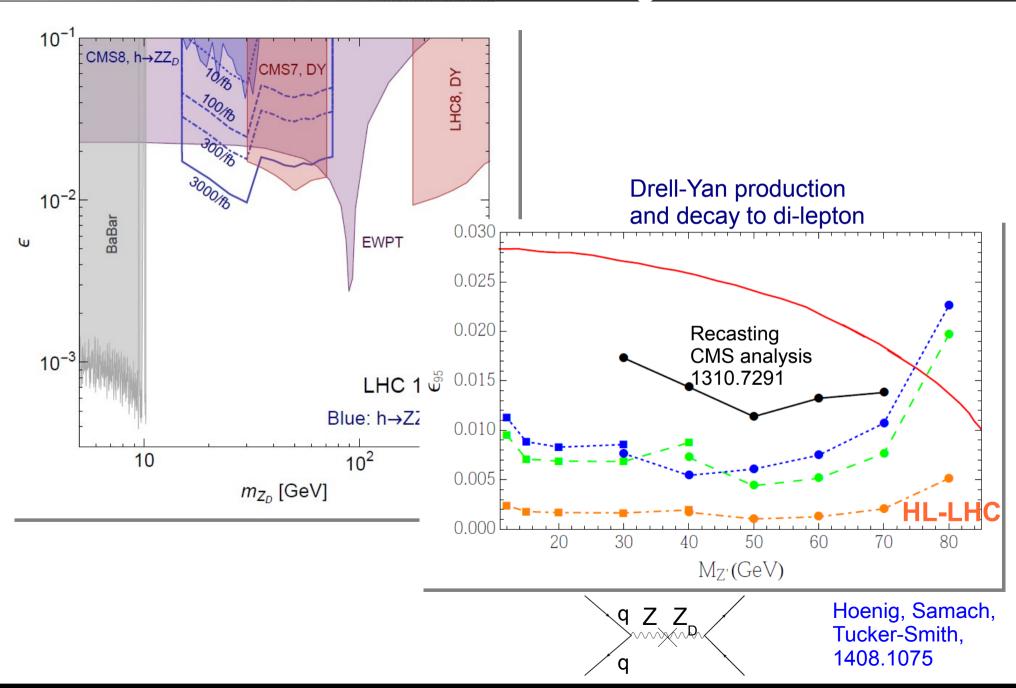


### Conclusions & lessons

### Great opportunity to test new forces using Higgs decays!

- ◆ Interesting possibility of testing Z<sub>D</sub> gauge bosons from dark sectors using new Higgs boson decays
- Access to regions of parameter space very well theoretically motivated
- ◆ We will incredibly benefit from the <u>High-Luminosity LHC</u>, since the <u>signature is very clean</u>
- ◆ Going to higher energy, we will gain in rate:
  100TeV collider can set bounds BR(h→Z<sub>D</sub>Z<sub>D</sub>→4I) ~ few\*10<sup>-8</sup>!

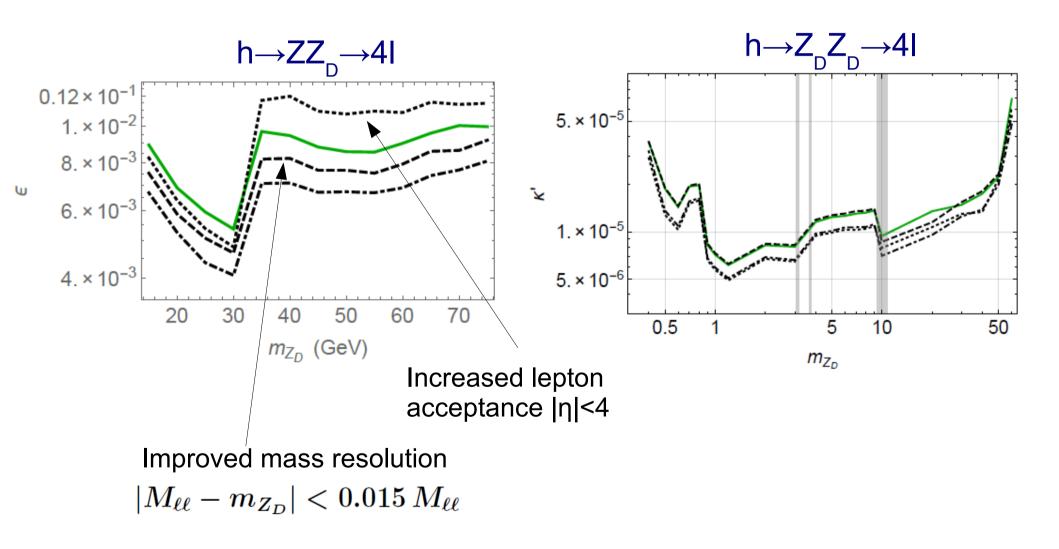
# Drell-Yan production of Z



Backup

# Importance of detector design

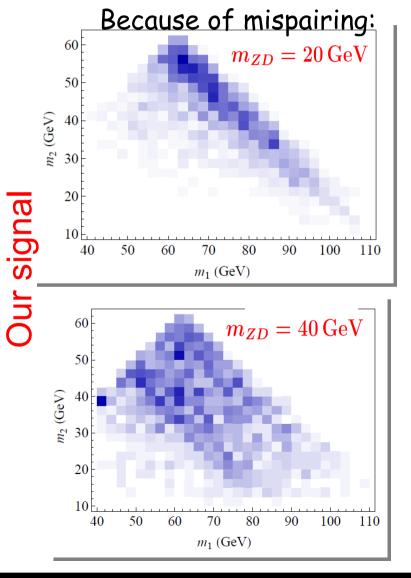
Reach at 100 TeV, with 3000 fb<sup>-1</sup>



Backup

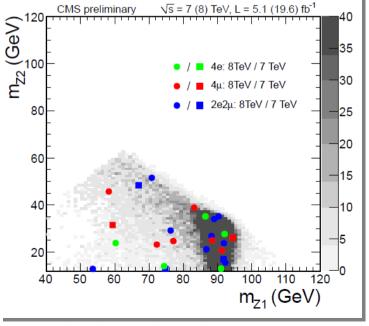
# Setting bounds on ZZ: present

Bounds coming from SM h  $\rightarrow$  ZZ\* $\rightarrow$  4l searches at the LHC CMS PAS HIG-13-002, ATLAS-CONF-2013-013



$$40 \text{ GeV} < m_{_1} < 120 \text{ GeV},$$
  
 $12 \text{ GeV} < m_{_2} < 120 \text{ GeV}$ 

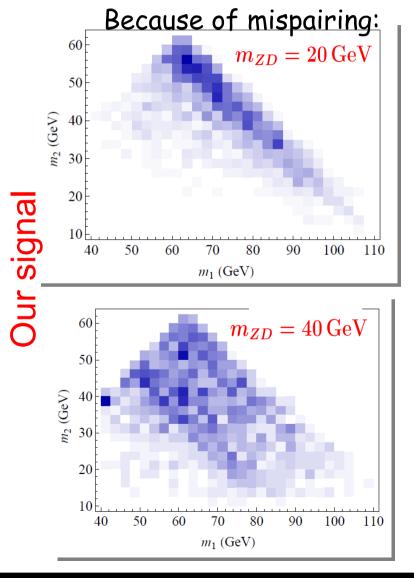
#### To compare to the experimental data:



**CMS PAS HIG-13-002** 

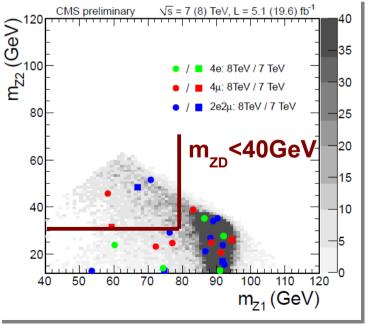
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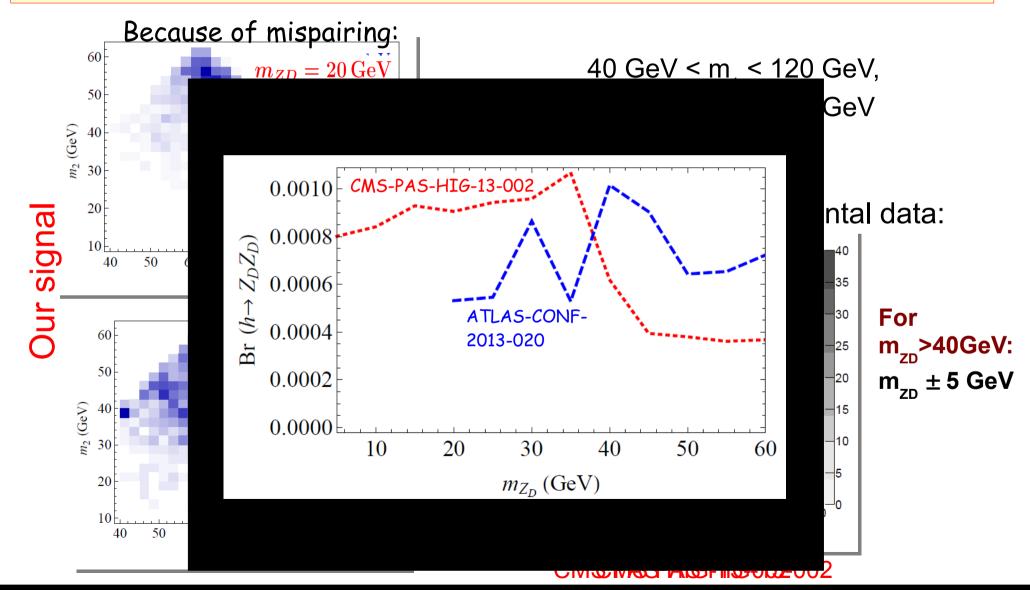


**CMS PAS HIG-13-002** 

For  $m_{ZD} > 40 \text{GeV}$ :  $m_{ZD} \pm 5 \text{ GeV}$ 

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