

BEYOND THE STANDARD MODEL

Lecture II

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THE HIERARCHY PROBLEM

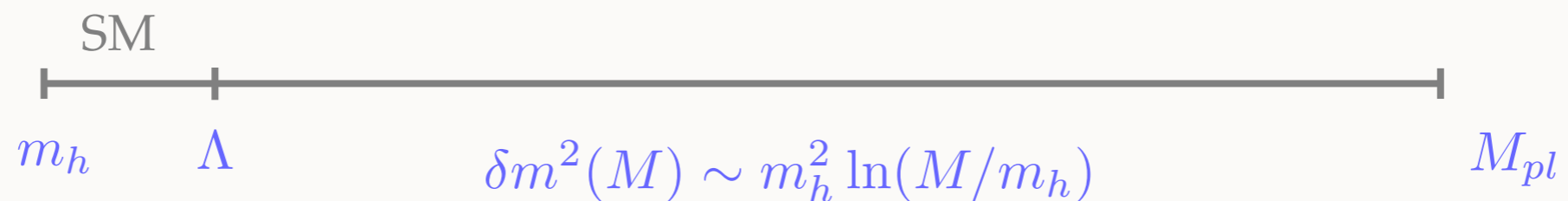
- Natural expectation for scalar fields: $m_s \sim \Lambda$
- natural EWSB needs **new physics near TeV**



- but this new physics must be **special**: theory above Λ must be free of quadratic divergences

THE HIERARCHY PROBLEM

- Idea 1: cancellation of quadratic divergences



- new physics closely related to SM:



THE HIERARCHY PROBLEM

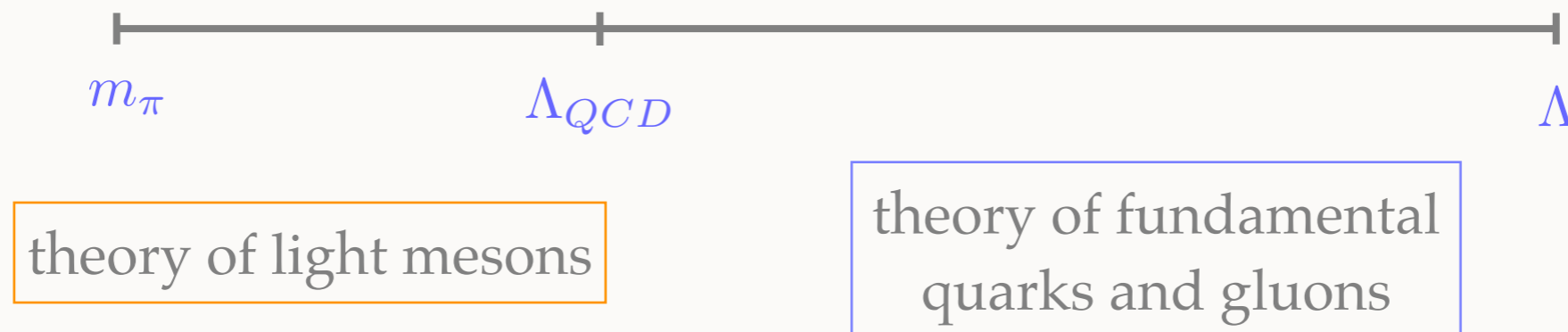
- Complete solution: cancellation must be **exact**
 - This requires a **lot** of new states!
 - **symmetry** to relate couplings of NP to those of the SM
 - e.g.: **SUSY**
 - If there is no symmetry, then cancellation is **accidental** and will break down at higher scales: defers hierarchy problem

THE HIERARCHY PROBLEM

- Idea 2: get rid of the problematic operator

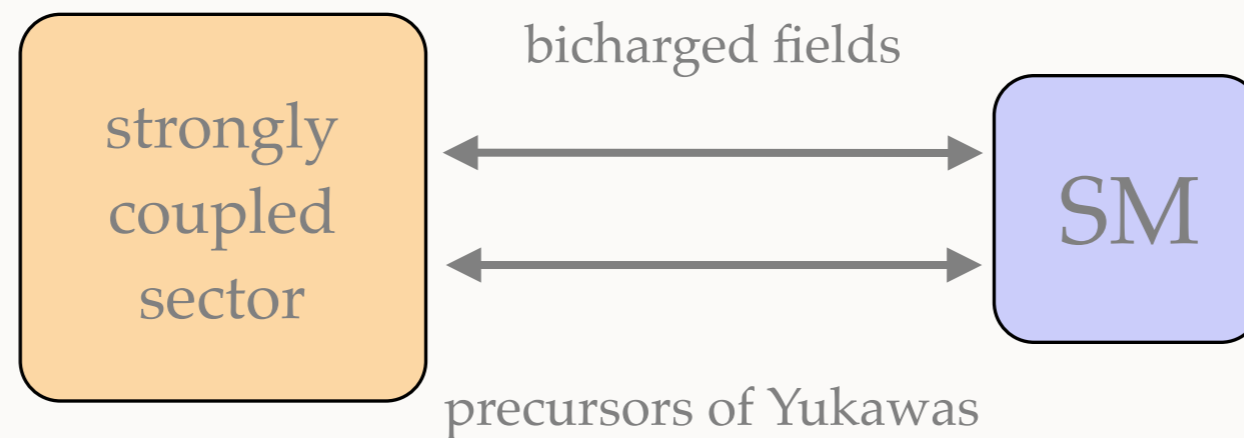


- Analogy: QCD



THE HIERARCHY PROBLEM

- In these models the Higgs is a **composite state**

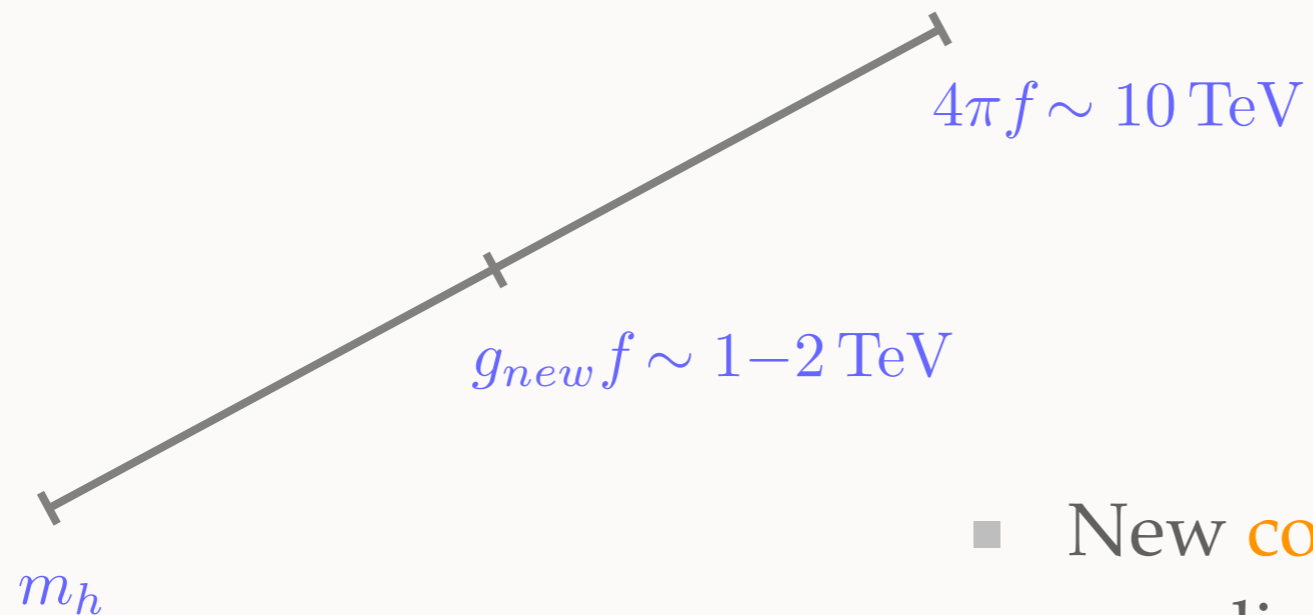


- Confinement triggers chiral symmetry breaking
 - SM-like Higgs is a **pseudo-Nambu-Goldstone** boson

$$\begin{array}{c} | \text{-----} | \text{-----} | \\ m_h \qquad g_{new} f \qquad 4\pi f \end{array}$$

THE HIERARCHY PROBLEM

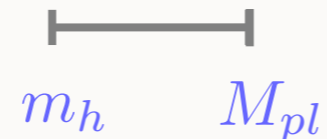
- Minimal set of collider signatures:



- New strongly coupled partons
- New **composite resonances**, coupling most strongly to W, Z, H, t
- Deviations in Higgs properties: $\mathcal{O}(v^2/f^2) \sim 10\%$
 - additional goldstones?

THE HIERARCHY PROBLEM

- Idea 3: *no running*

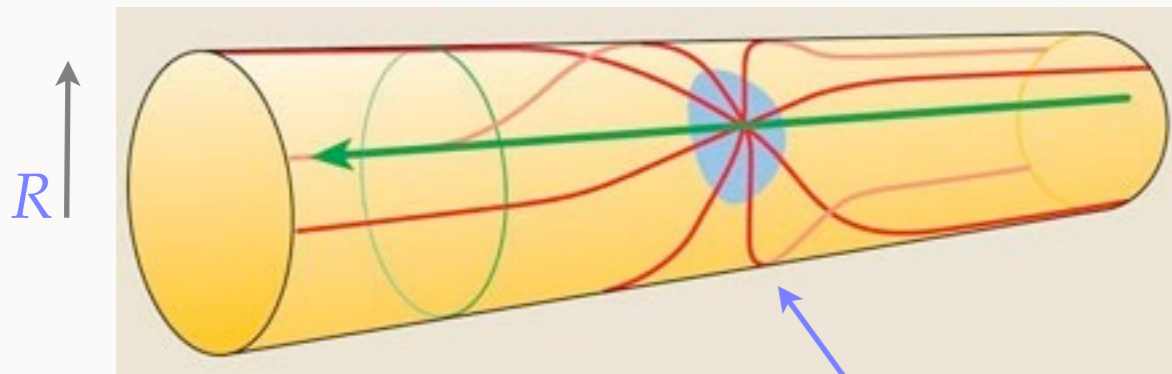


m_h M_{pl}

- apparent weakness of gravity compared to SM forces is an *illusion* due to *geometry of spacetime*
 - SM particles are inherently 4D (string theory makes this plausible)
 - various models: flat, warped, ...

THE HIERARCHY PROBLEM

- consider n flat extra dimensions:



$$V(r) = \frac{m}{M_{Pl(4+n)}^{n-2}} \frac{1}{R^n} \frac{1}{r}$$

but for $r > R$, field propagation is 4-dimensional

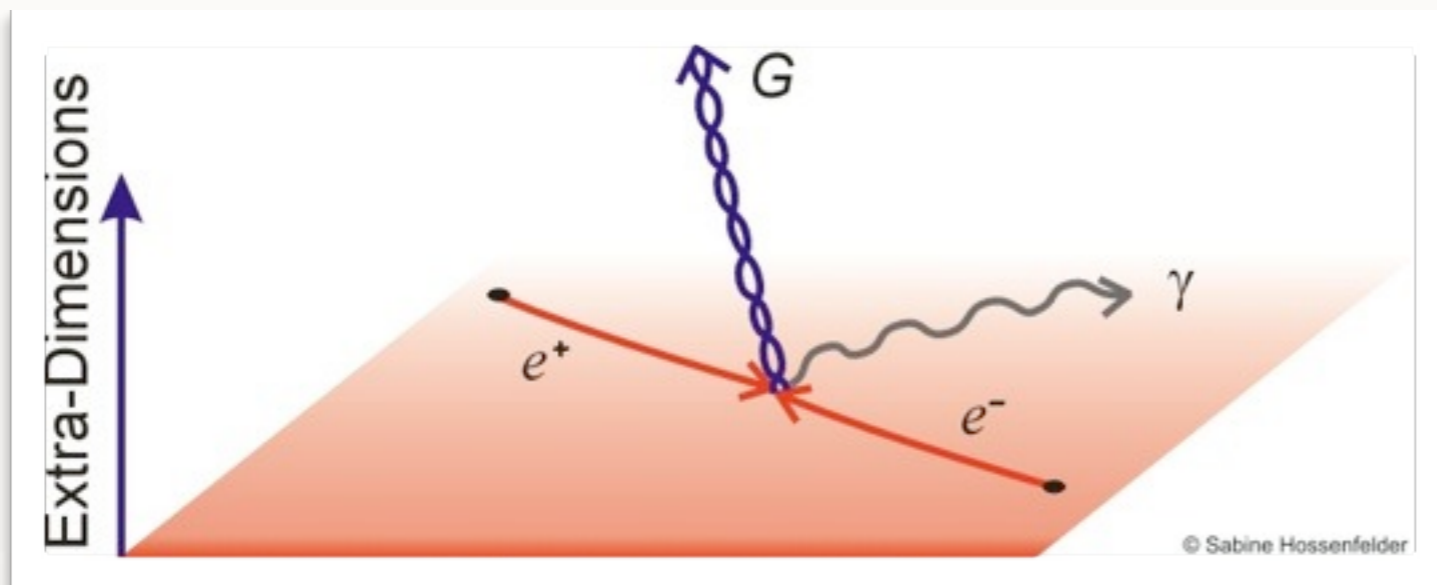
short distance behavior is $4+n$ -dimensional

$$V(r) = \frac{m}{M_{Pl(4+n)}^{n-2}} \frac{1}{r^{n+1}}$$

- if **volume of extra dimensions is large** in units of Planck length, 4D gravity appears weak: $M_{Pl(4+n)} R^n = M_{Pl(4)}$

THE HIERARCHY PROBLEM

- Some collider signatures:
 - production of Kaluza-Klein gravitons:



mono-(jet, W, Z, photon)
recoiling against
large MET

- ...maybe quantum black holes?!

THE HIERARCHY PROBLEM

- I am going to focus on SUSY
 - In my view: most compelling as a theory of the universe
 - Many consequences of SUSY as applied to the hierarchy problem are **qualitatively similar** to those of other models
 - **partner particles** for SM
 - **parity** symmetry leading to **dark matter** candidates (**MET**)
 - collider searches for heavy states with SM charges

SUPERSYMMETRY

- Theory of 1 complex scalar + 1 Weyl fermion:

$$\mathcal{L} = \partial_\mu \phi \partial^\mu \phi^* + i\bar{\psi} \gamma^\mu \partial_\mu \psi$$

- invariant under **supersymmetry** transformation:

$$\delta\phi = \bar{\epsilon}\psi \qquad \delta\psi = -i\epsilon\gamma^\mu \partial_\mu \phi$$

- two SUSY variations yield a **translation**:

$$[\delta_1, \delta_2] \phi = -i\bar{\epsilon}_2 \gamma^\mu \epsilon_1 \partial_\mu \phi$$

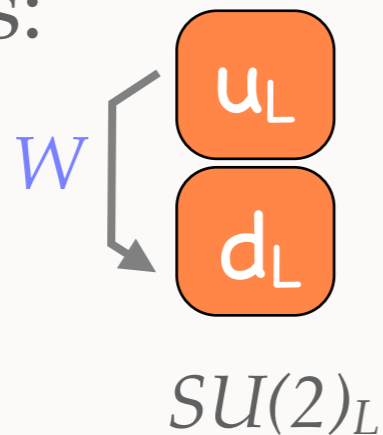
- recall $\delta\phi = a^\mu \partial_\mu \phi$: generated by **momentum**

SUPERSYMMETRY

- SUSY is thus inherently intertwined with spacetime (**Poincare**) symmetry
 - SUSY: a statement about **background spacetime**
 - we can't pick and choose a subsector of the universe to supersymmetrize
 - the kinds of **representations** of SUSY that we can have depend on particle's Lorentz quantum numbers, in particular, on their **spin**.

SUPERSYMMETRY

- Multiplets:

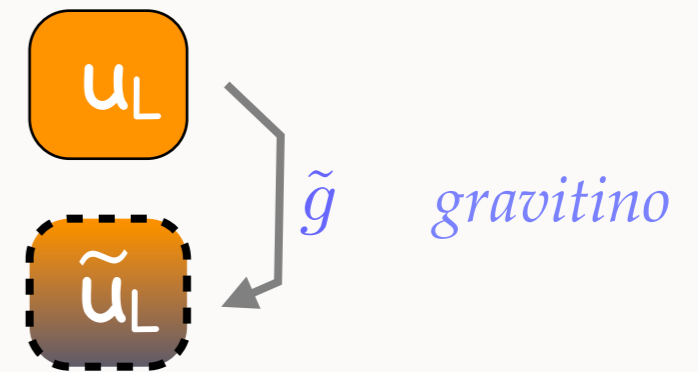


- supermultiplets: particle and superpartner

- fermion - sfermion (\tilde{u}_L, u_L)

- gauge boson - gaugino (\tilde{B}, B_μ)

- Higgs boson - higgsino (H_u, \tilde{H}_u)



chiral multiplets

vector multiplets

SUPERSYMMETRIC THEORIES

- Supersymmetry restricts possible interactions
 - Analogy: EWSB
 - Below scale of EWSB, u_R, u_L seem to have quantum numbers allowing Dirac mass term: $m_u u_R u_L$
 - But forbidden under underlying $SU(2)_L \times U(1)_Y$ - need $\frac{y_u v}{\sqrt{2}} u_R u_L$
 - from the parent interaction $y_u H u_R Q_L$
 - which also yields the interaction $\frac{y_u}{\sqrt{2}} h u_R u_L$

SUPERSYMMETRIC THEORIES

- SUSY relates Yukawa interactions $H Q_L u_R$ to quartic scalar couplings $|H|^2 |\tilde{Q}_L|^2, \dots$
- useful compact formalism: **superpotential**

$$\text{superfields} \xrightarrow{\quad} W = y_u Q_L H u_R + \dots \quad \begin{array}{l} \text{renormalizable} \\ \text{interactions are cubic} \end{array}$$

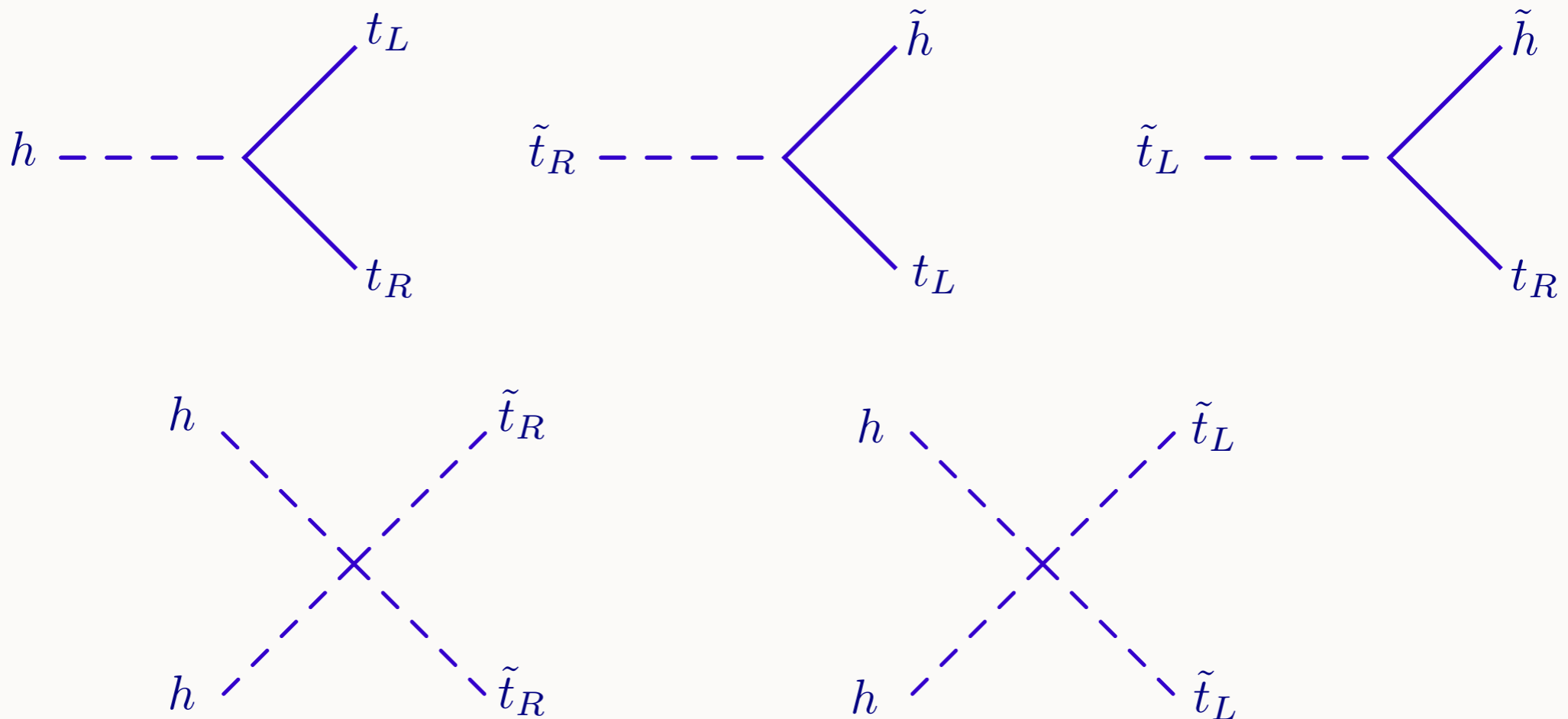
- determines all supersymmetric interactions between chiral multiplets:

$$\mathcal{L}_{Yuk} = -W_{ij} \psi_i \psi_j$$

$$V(\phi) = |W_i|^2$$

SUPERSYMMETRIC THEORIES

- Thus one cubic superpotential term $y Q_L H u_R$ encodes



SUPERSYMMETRIC THEORIES

- SM Yukawas:

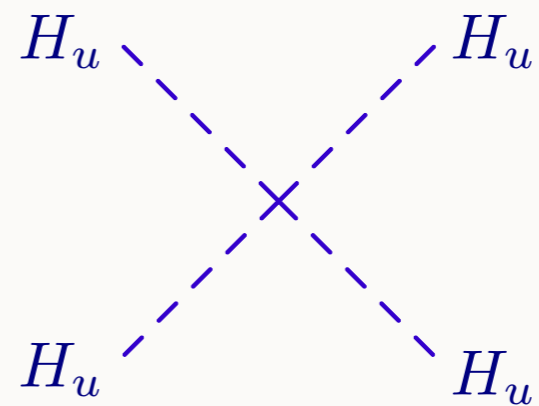
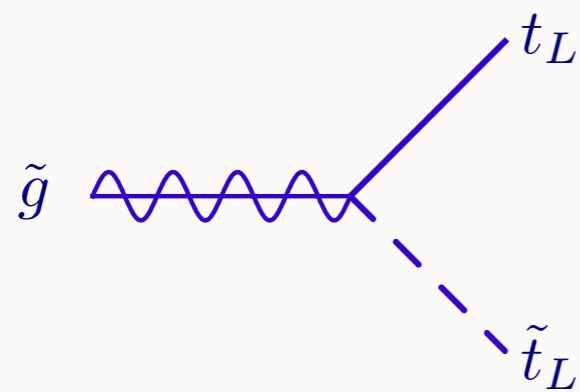
$$\mathcal{L}_{SM} = y_d Q_L H d_R + y_\ell L_L H e_R + y_u Q_L H^c u_R$$

- But only superfields, not their complex conjugates, can appear in W : cannot be supersymmetrized
- Must introduce **two** Higgs doublets H_u, H_d
 - also fixes up quantum consistency of MSSM: anomaly cancellation
- SUSY quadratic Higgs potential terms from $W = \mu H_u H_d$

SUPERSYMMETRIC THEORIES

- What about gauge interactions?
 - Gauge invariance uniquely dictates interactions of gauge bosons with charged particles
 - SUSY relates these to **gaugino interactions** and new **scalar quartics**,

$$\mathcal{L}_{new} = -\sqrt{2}g(\phi^* t^a \psi)\lambda^a + H.c. - \frac{g^2}{2}(\phi^* t^a \phi)^2$$



SUPERSYMMETRIC MSSM

- This gives us the SUSY-preserving part of the MSSM:

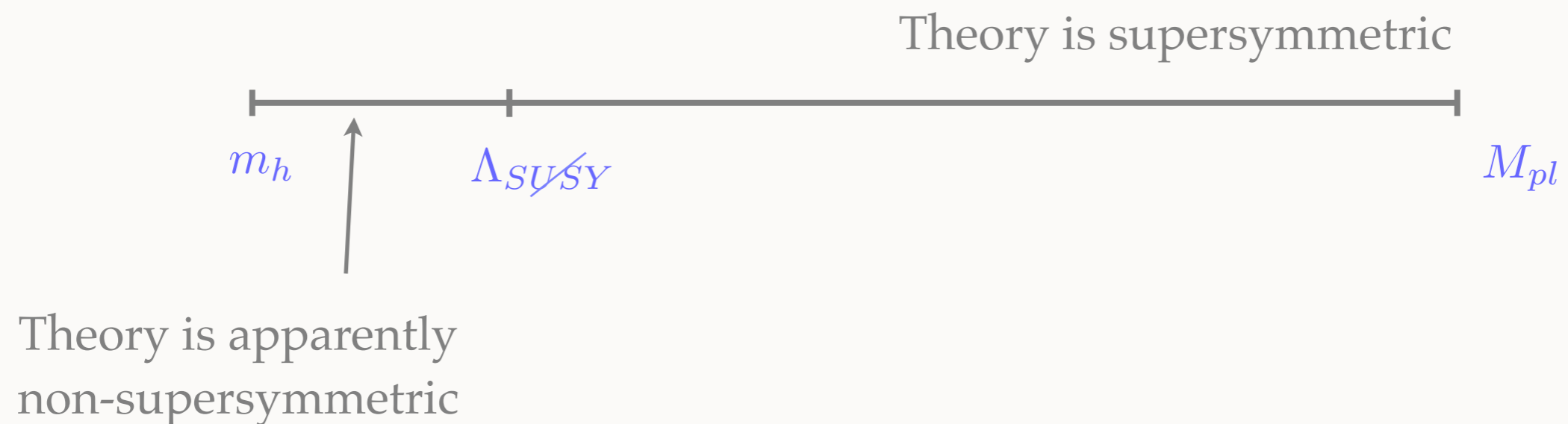
particles	sparticles
$\begin{pmatrix} u_L \\ d_L \end{pmatrix}$ u_R d_R	$\begin{pmatrix} \tilde{u}_L \\ \tilde{d}_L \end{pmatrix}$ \tilde{u}_R \tilde{d}_R
$\begin{pmatrix} \nu_L \\ e_L \end{pmatrix}$ e_R ν_R	$\begin{pmatrix} \tilde{\nu}_L \\ \tilde{e}_L \end{pmatrix}$ \tilde{e}_R $\tilde{\nu}_R$
H_u H_d	\tilde{H}_u \tilde{H}_d
g_μ^a W_μ^a B_μ	\tilde{g}^a \tilde{W}^a \tilde{B}

- Extremely predictive!
- More than double the particles of the SM
- Fewer parameters

- Of course, SUSY is broken in nature...

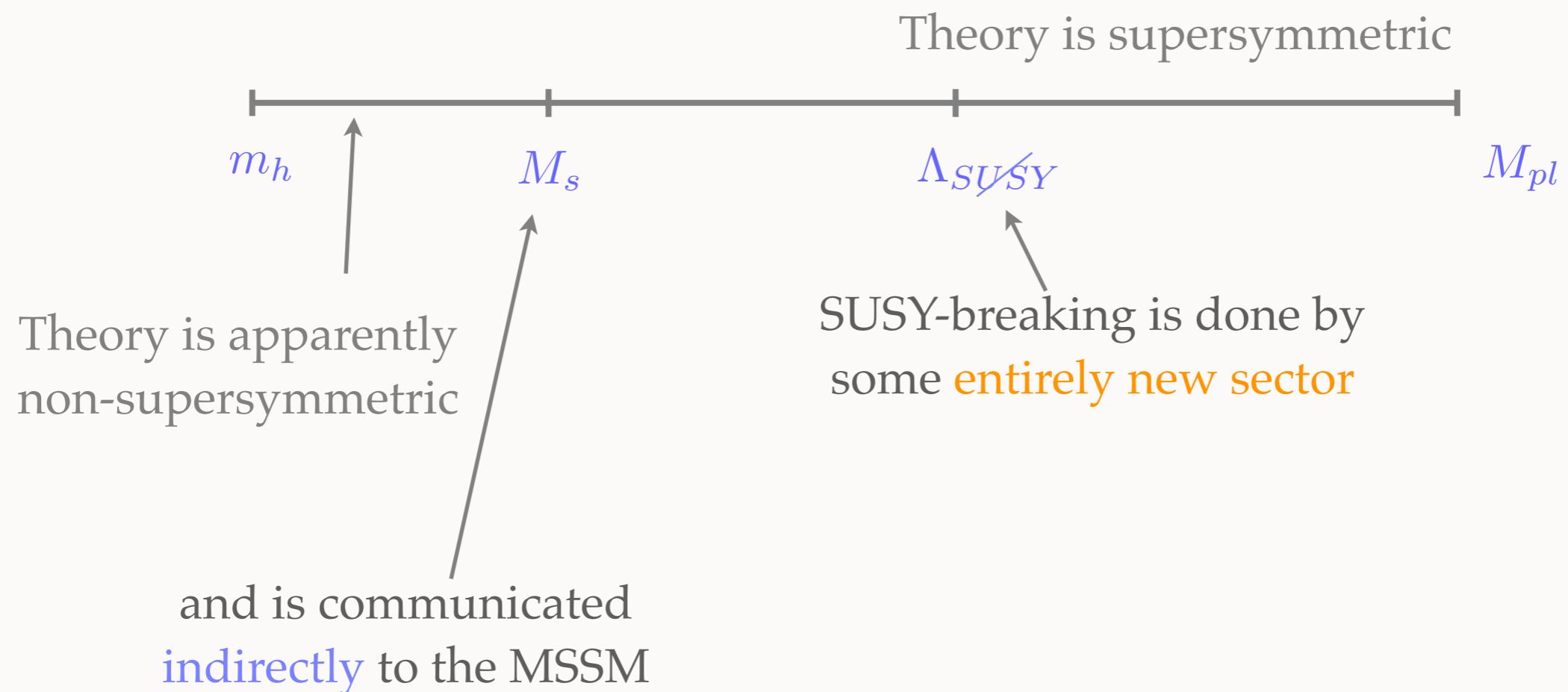
SUSY BREAKING

- How can we break SUSY without spoiling the solution to the hierarchy problem?
 - Must break SUSY *spontaneously*



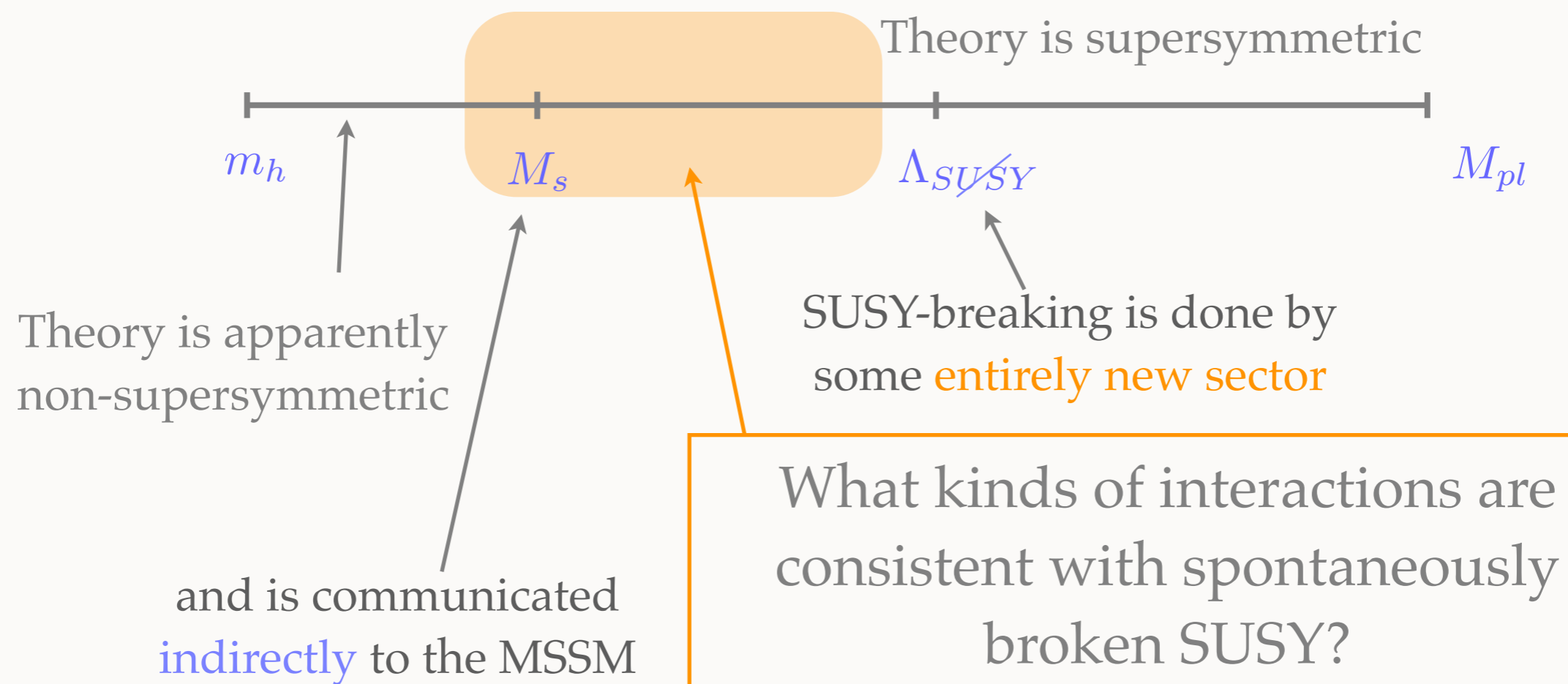
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SUSY BREAKING

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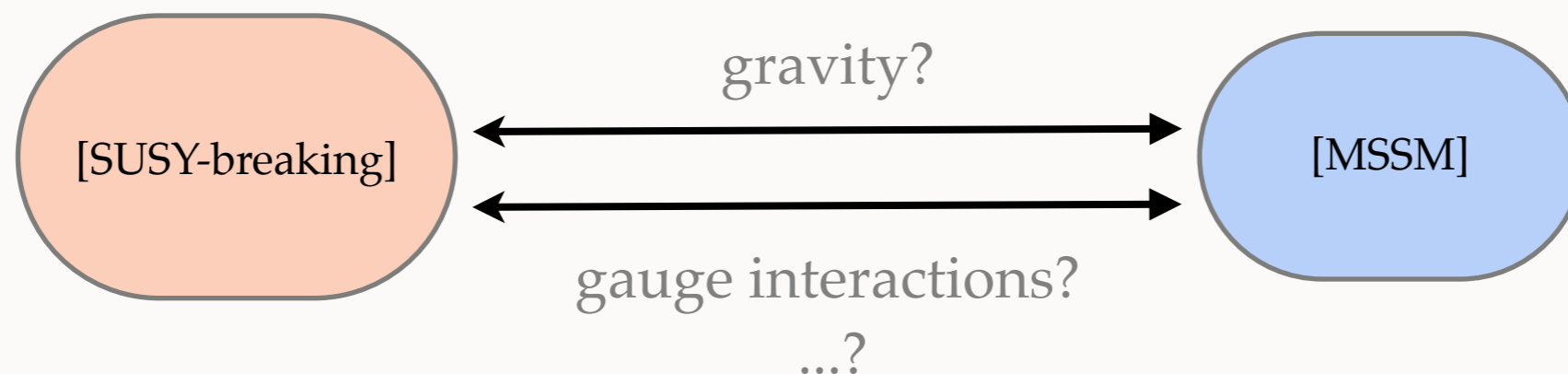


SUSY BREAKING

- Spontaneous SUSY-breaking yields a **supertrace sum rule** at tree level

$$\text{Tr } m_S^2 - 2\text{Tr } m_F^2 + 3\text{Tr } m_V^2 = 0$$

- To get acceptable spectra, must break SUSY in a **hidden sector**:



SUSY BREAKING

- This induces the “**soft SUSY-breaking**” Lagrangian:

$$\mathcal{L}_{soft} = -\frac{1}{2} \left(M_3 \tilde{g} \tilde{g} + M_2 \tilde{W} \tilde{W} + M_1 \tilde{B} \tilde{B} + H.c. \right)$$

masses for superpartners only

$$-\tilde{Q}_L^* M_Q^2 \tilde{Q}_L - \tilde{u}_R^* M_u^2 \tilde{u}_R - \tilde{d}_R^* M_d^2 \tilde{d}_R - \tilde{L}_L^* M_L^2 \tilde{L}_L - \tilde{e}_R^* M_e^2 \tilde{e}_R$$

trilinear couplings: one for each super-potential term

$$-\left(A_u \tilde{u}_R \tilde{Q}_L H_u + A_d \tilde{d}_R \tilde{Q}_L H_d + A_e \tilde{e}_R \tilde{L}_L H_d + H.c. \right)$$

and same in the Higgs sector

$$-m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (b H_u H_d + H.c.)$$

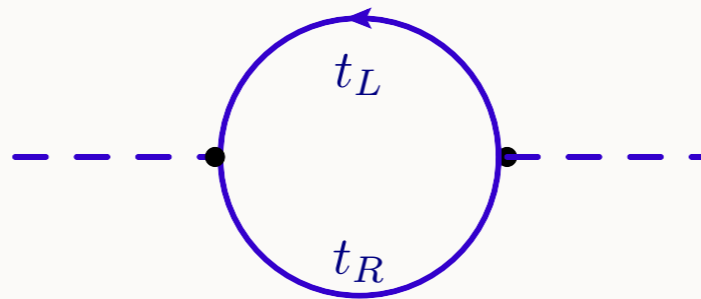
- over **100 free parameters!**

SUSY BREAKING

- All soft SUSY-breaking terms are **dimensionful**
 - sensible: less important in the UV than in the IR
 - must vanish as $\Lambda_{SUSY} \rightarrow 0$
 - no changes in the **renormalizable** couplings \Rightarrow leading divergence cancellations **unaffected**
 - **no new quadratic divergences** are generated

SUSY BREAKING

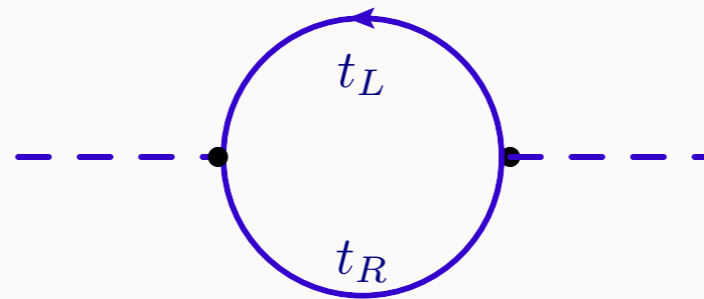
- Let's do an explicit example: top and stop loops



$$-i \delta m_h^2|_{top} = -2N_c y_t^2 \int \frac{d^4 k}{(2\pi)^4} \frac{k^2 + m_t^2}{(k^2 - m_t^2)^2}$$

SUSY BREAKING

- Let's do an explicit example: top and stop loops



overall minus sign
from fermion loop

logarithmic
divergence

$$-i \delta m_h^2|_{top} = -2N_c y_t^2 \int \frac{d^4 k}{(2\pi)^4} \frac{k^2 + m_t^2}{(k^2 - m_t^2)^2}$$

color degrees
of freedom

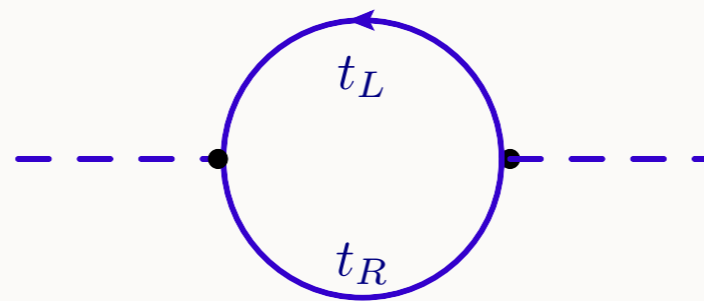
vertices

two propagators

quadratic
divergence

SUSY BREAKING

- Let's do an explicit example: top and stop loops

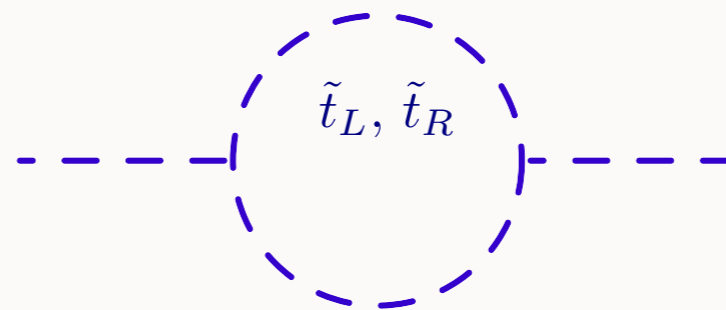
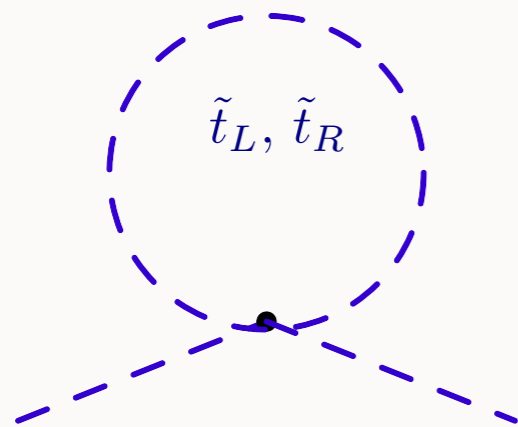


$$-i \delta m_h^2|_{top} = -2N_c y_t^2 \int \frac{d^4 k}{(2\pi)^4} \frac{k^2 + m_t^2}{(k^2 - m_t^2)^2}$$

$$\delta m_h^2|_{top} = -\frac{3y_t^2}{8\pi^2} \left(\Lambda^2 - 3m_t^2 \ln \left(\frac{\Lambda^2 + m_t^2}{m_t^2} \right) + \dots \right)$$

SUSY BREAKING

- Let's do an explicit example: top and stop loops



$$\mathcal{L}_{stop} = -\frac{1}{2}(v+h)^2 (|\tilde{t}_L|^2 + |\tilde{t}_R|^2)$$

SUSY-preserving

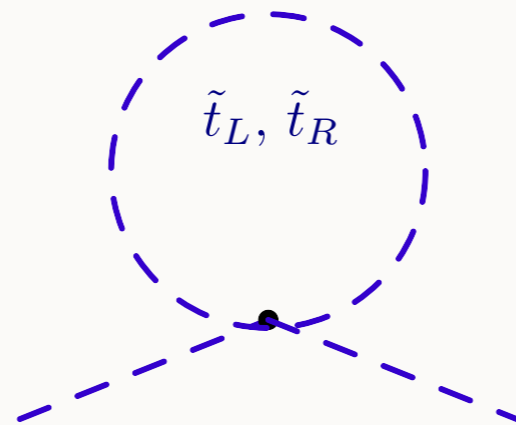
$$+m_R^2|\tilde{t}_R|^2 + m_L^2|\tilde{t}_L|^2$$

SUSY-breaking

- In general also SUSY-breaking contribution to trilinears

SUSY BREAKING

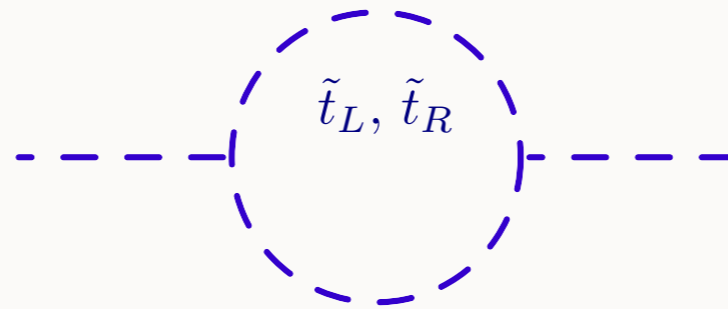
- Let's do an explicit example: top and stop loops



$$\delta m_h^2|_{stop 1} = \frac{3y_t^2}{16\pi^2} \left(2\Lambda^2 - m_L^2 \ln \left(\frac{\Lambda^2 + m_L^2}{m_L^2} \right) - m_R^2 \ln \left(\frac{\Lambda^2 + m_R^2}{m_R^2} \right) + \dots \right)$$

SUSY BREAKING

- Let's do an explicit example: top and stop loops



$$\delta m_h^2|_{stop 2} = -\frac{3y_t^2}{8\pi^2} \left(m_t^2 \ln \left(\frac{\Lambda^2 + m_L^2}{m_L^2} \right) - m_t^2 \ln \left(\frac{\Lambda^2 + m_R^2}{m_R^2} \right) + \dots \right)$$

No quadratic divergences:
dimensionally impossible

SUSY-breaking trilinears:
mt \rightarrow more general function

SUSY BREAKING

- Let's do an explicit example: top and stop loops
 - Add everything up:

$$\delta m_h^2|_{top} = -\frac{3y_t^2}{8\pi^2} \left(\Lambda^2 - 3m_t^2 \ln \left(\frac{\Lambda^2 + m_t^2}{m_t^2} \right) + \dots \right)$$

$$\delta m_h^2|_{stop1} = \frac{3y_t^2}{16\pi^2} \left(2\Lambda^2 - m_L^2 \ln \left(\frac{\Lambda^2 + m_L^2}{m_L^2} \right) - m_R^2 \ln \left(\frac{\Lambda^2 + m_R^2}{m_R^2} \right) + \dots \right)$$

$$\delta m_h^2|_{stop2} = -\frac{3y_t^2}{8\pi^2} \left(m_t^2 \ln \left(\frac{\Lambda^2 + m_L^2}{m_L^2} \right) - m_t^2 \ln \left(\frac{\Lambda^2 + m_R^2}{m_R^2} \right) + \dots \right)$$

Quadratic divergence cancels independently of soft breaking terms

Exact SUSY: log divergence cancels too

SUSY BREAKING

- So about those >100 free parameters...
 - **Tremendous** constraints from **flavor, CP**
 - flavor structure can't be arbitrary: SUSY flavor problem
 - **Top-down**: specific mediation mechanisms impose **characteristic relationships** between soft parameters
 - gauge mediation, gravity mediation, anomaly mediation, ...
 - **Bottom-up**: CP-preserving, nearly flavor-symmetric sector
 - “**pMSSM**”: a mere **20** parameters

R-PARITY

- Unlike in the SM, we cannot write down all interactions allowed by gauge symmetries:

$$W = \mu H_u H_d + Y_u Q_L H_u u_R + Y_d Q_L H_d d_R + Y_e L_L H_d e_R \\ + \hat{\mu} H_u L_L + \lambda'' u_R d_R d_R + \lambda' Q_L L_L d_R + \lambda L_L L_L e_R$$

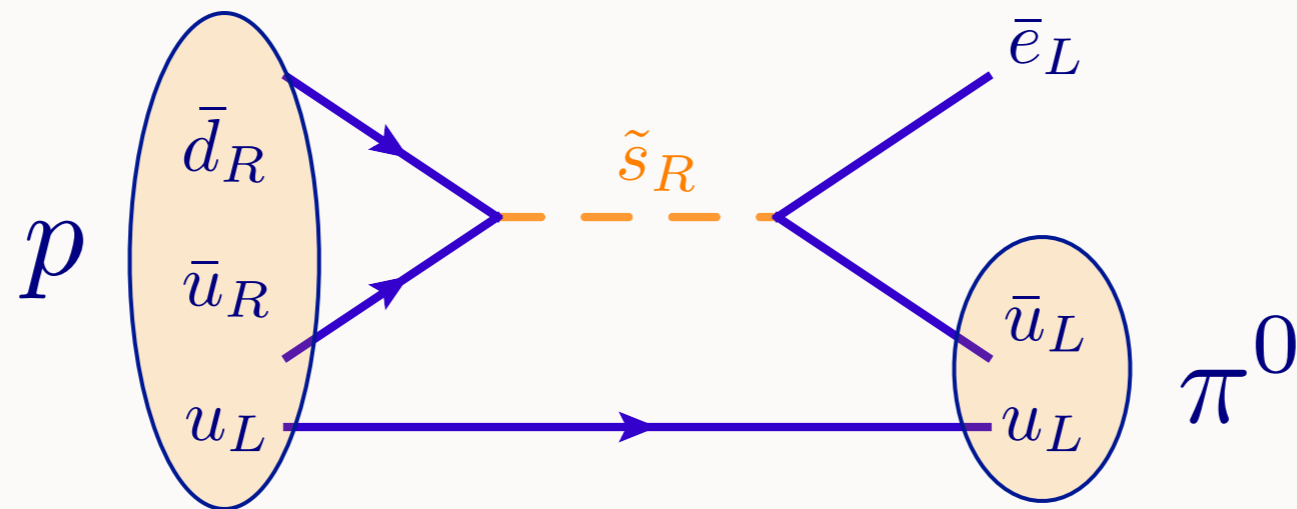
violates B

violates L

- Leads to whole tensors of new B and L -violating couplings:
 - e.g. Yukawas, $\lambda''_{112}(u_R d_R) \tilde{s}_R$, $\lambda'_{112} \tilde{s}_R (e_L u_L)$

R-PARITY

- Catastrophic proton decay:



- B, L violating Yukawa couplings must be **extremely** small:

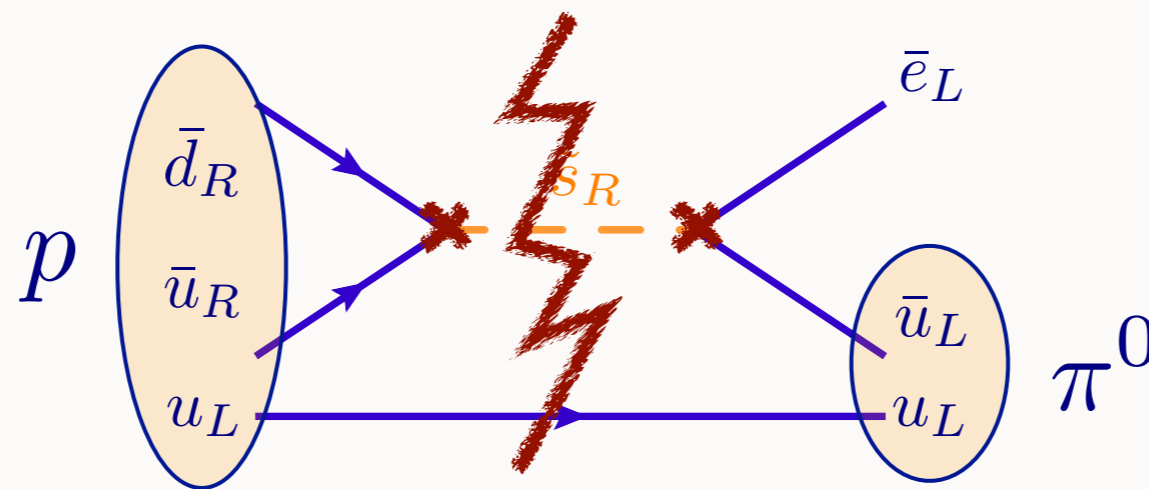
$$\Gamma \sim \frac{|\lambda''_{112}\lambda'_{112}|^2 m_p^5}{m_{\tilde{s}}^4} < 10^{34} \text{ years}$$

R-PARITY

- Easy solution: impose a new **global** symmetry:

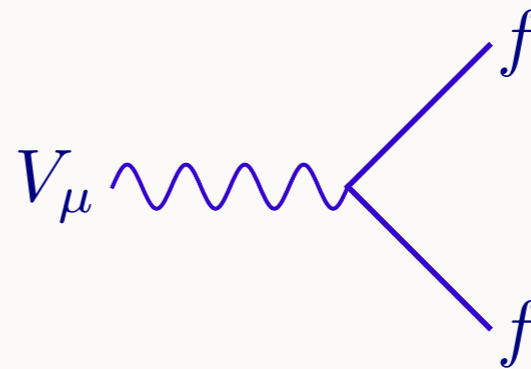
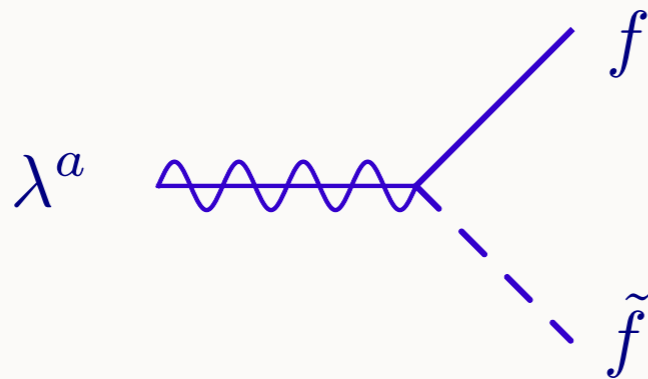
$$W = \mu H_u H_d + Y_u Q_L H_u u_R + Y_d Q_L H_d d_R + Y_e L_L H_d e_R \\ + \hat{\mu} H_u L_L + \lambda'' u_R d_R d_R + \lambda' Q_L L_L d_R + \lambda L_L L_L e_R$$

- impose **matter parity**: $P_M = (-1)^{3(B-L)}$



R-PARITY

- Gauge interactions:



- define **R-parity**:

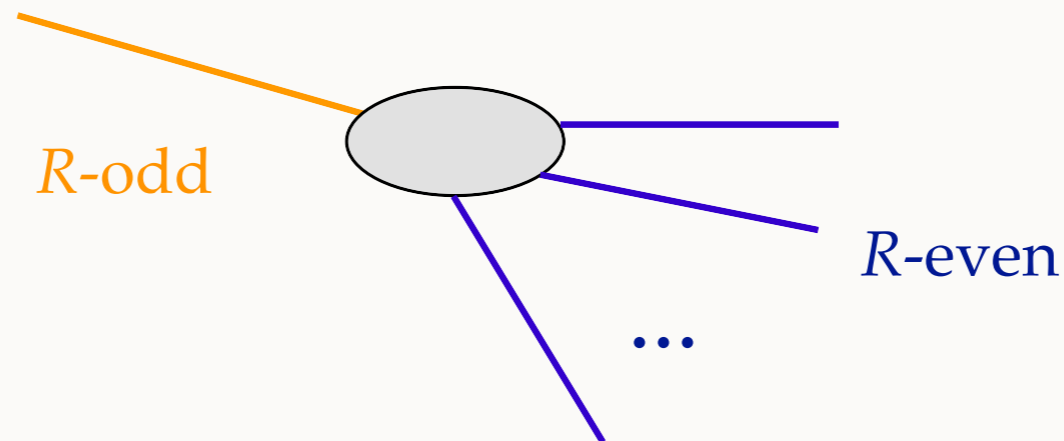
$$P_R = (-1)^{3(B-L)+2s}$$

- exactly the same! but
 - easier to see consequences
 - natural in SUSY

even	odd
f (spin 1/2)	\tilde{f} (spin 0)
V (spin 1)	\tilde{V} (spin 1/2)
H (spin 0)	\tilde{H} (spin 1/2)

R-PARITY

- Immediate consequence: lightest superpartner is **stable**



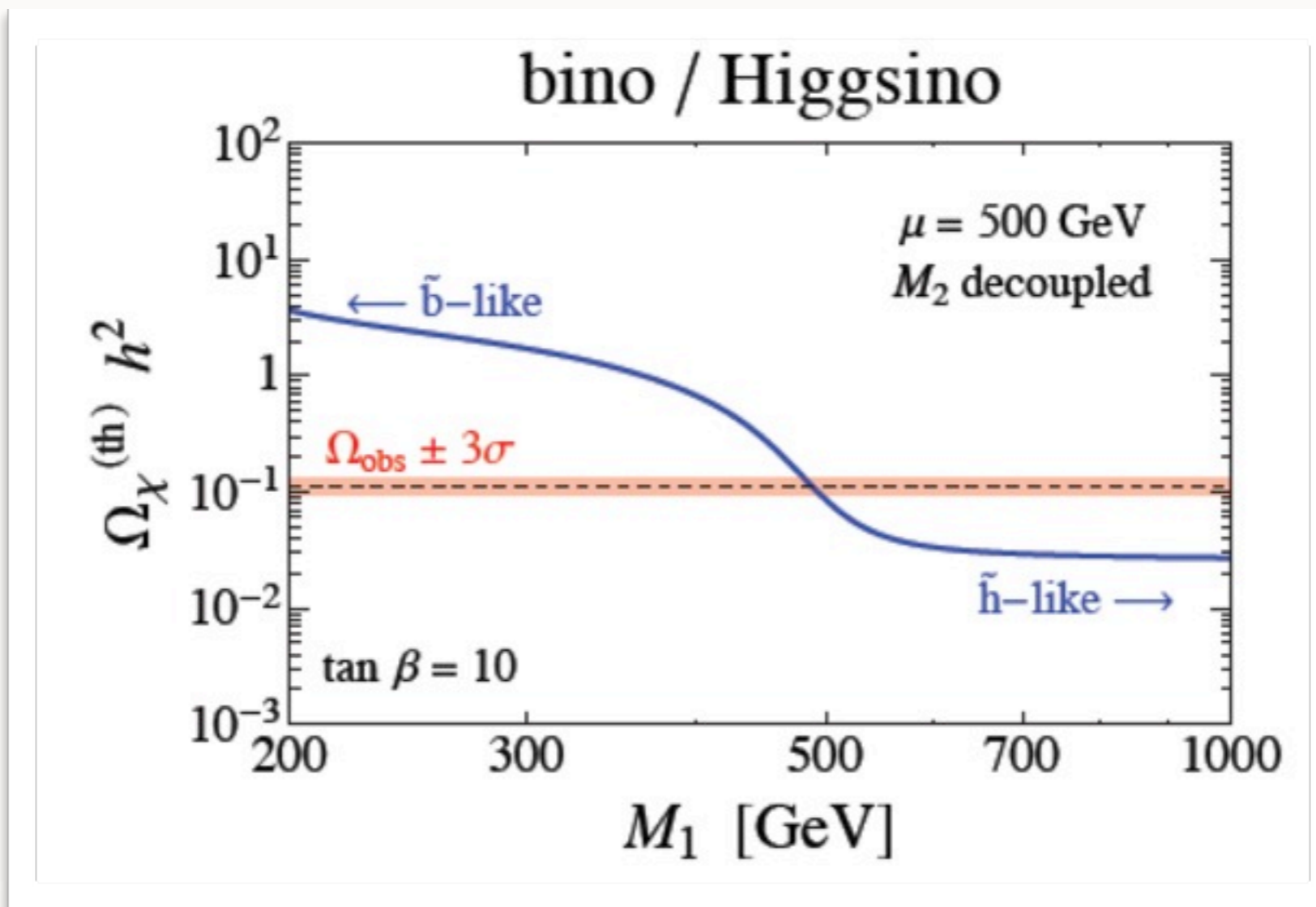
- This significantly restricts the spectrum:
 - lightest superpartner must be **neutral**
 - and must not **over-close the universe**

R-PARITY: DARK MATTER

- Lightest Supersymmetric Particle is an attractive DM candidate:
 - electroweak interactions, electroweak scale mass
 - Possible candidates:
 - neutralinos $\tilde{B}, \tilde{W}^3, \tilde{h}_u, \tilde{h}_d$
 - sneutrinos $\tilde{\nu}_L, \tilde{\nu}_R$
 - the devil is in the details

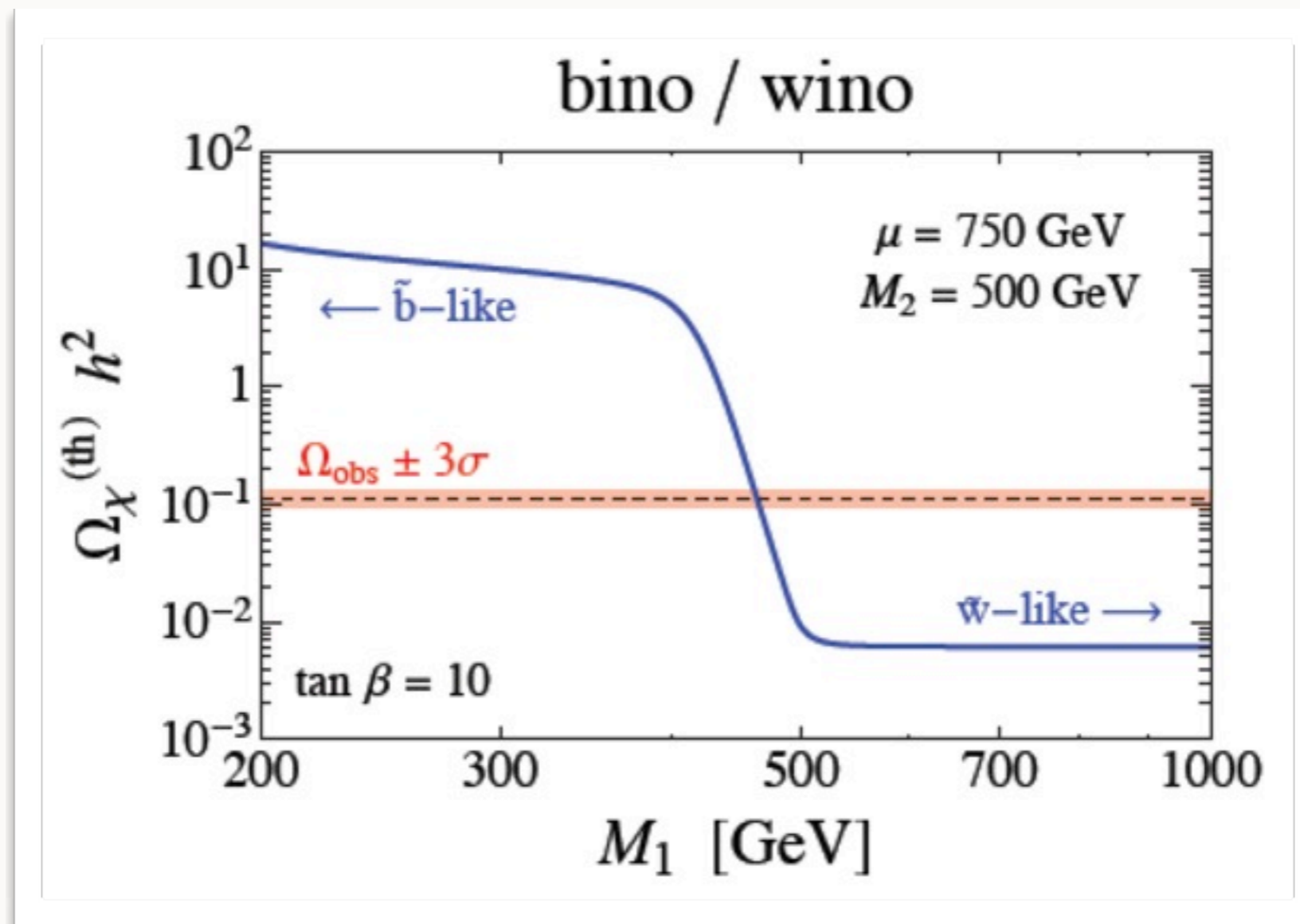
R-PARITY: DARK MATTER

- Relic abundance delicate function of spectrum:



R-PARITY: DARK MATTER

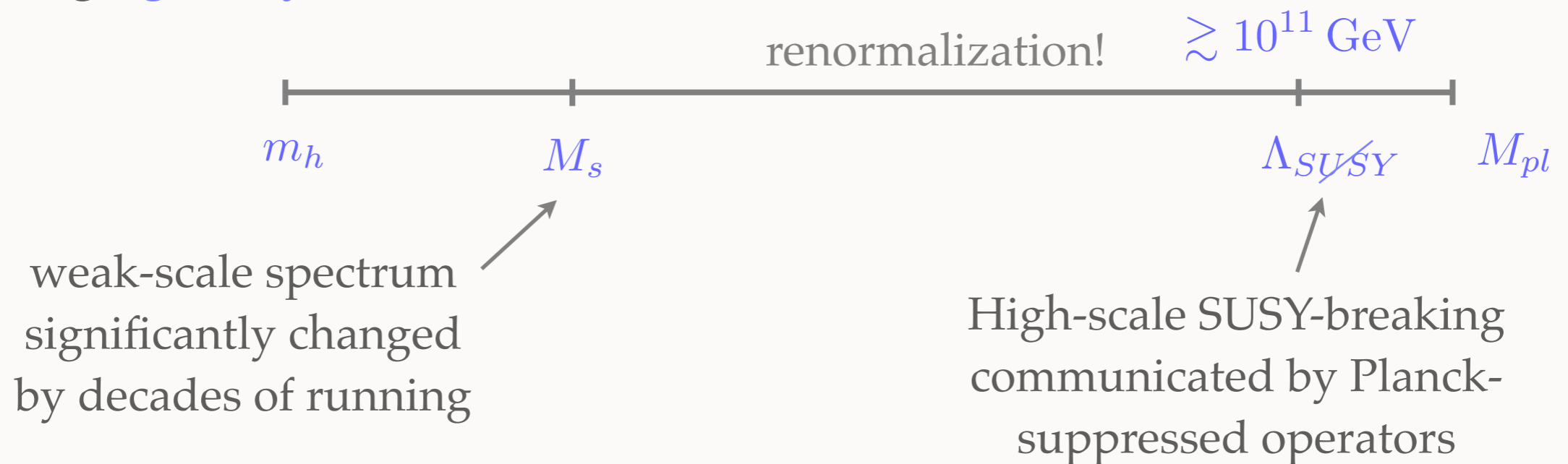
- Relic abundance delicate function of spectrum:



MSSM SPECTRA

- **High-scale** SUSY breaking

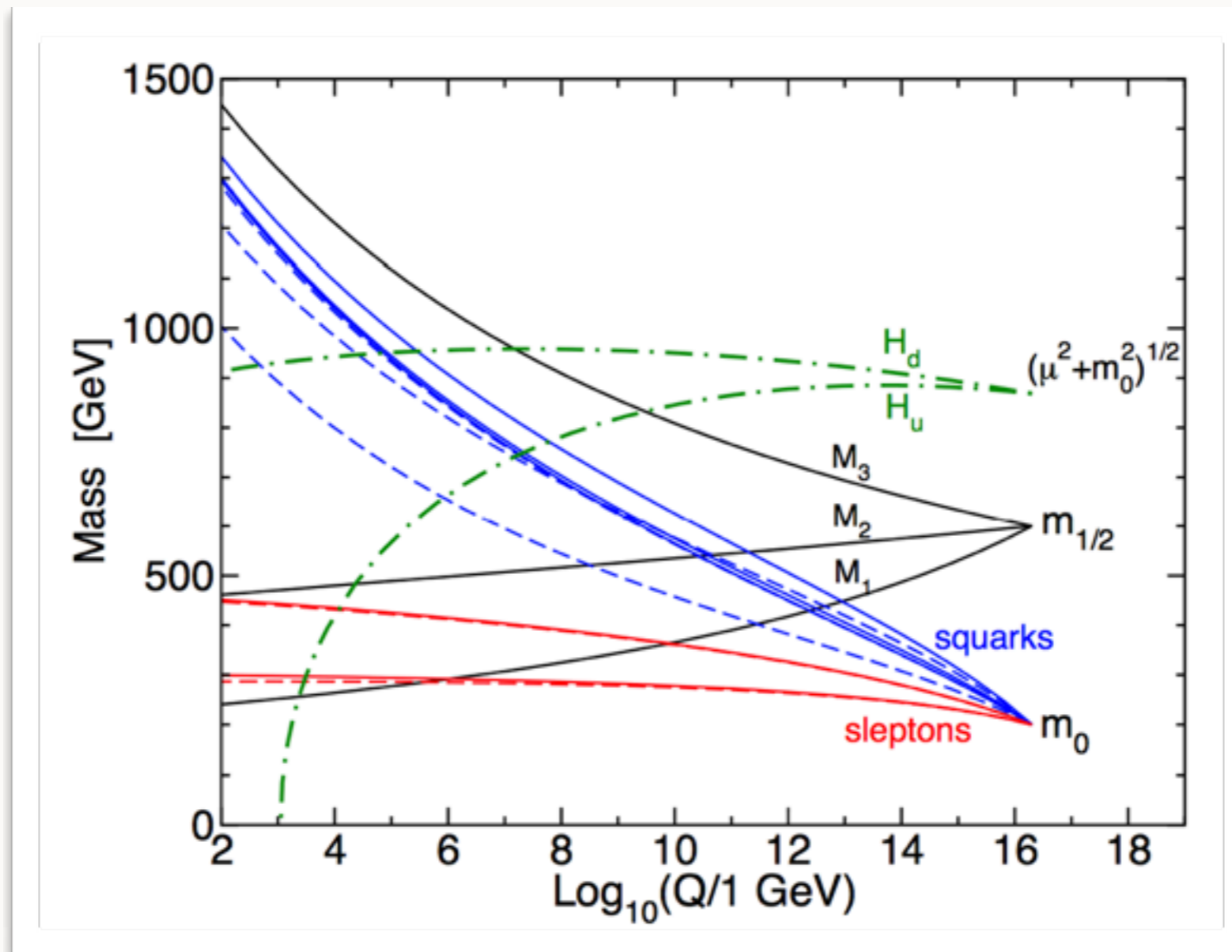
- e.g.: gravity-mediated



- **gravitino mass:** $m_{3/2} = \frac{\Lambda_{SUSY}^2}{M_{Pl}}$ sets scale for soft masses

MSSM SPECTRA

- Effects of RG evolution:



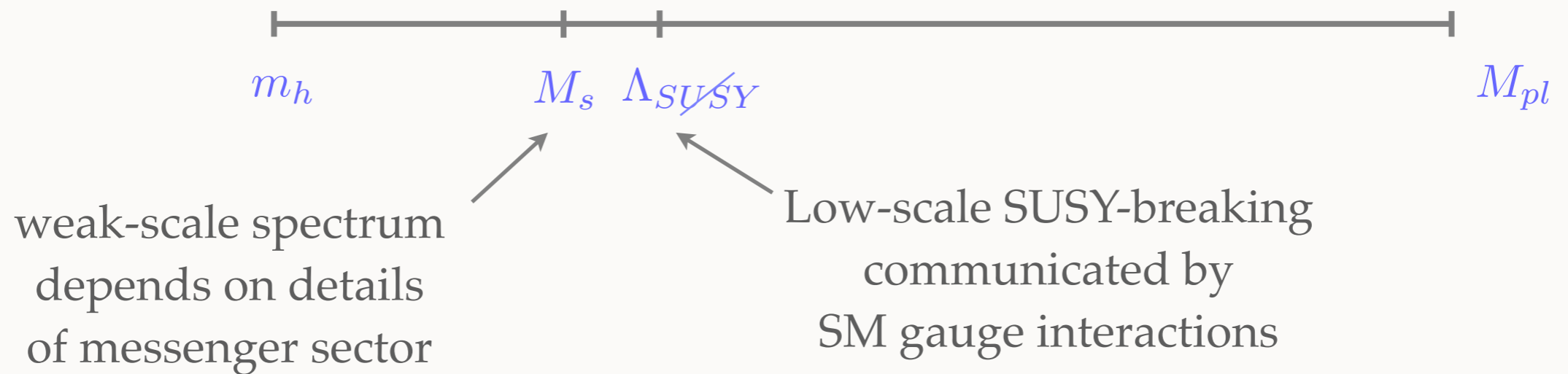
MSSM SPECTRA

- Simplest gravity-mediated model: **mSUGRA**
 - 100 parameters \rightarrow 5
 - useful toy model! But highly simplified
- Biggest issue: **flavor**
 - Gravitational interactions don't care about flavor \Rightarrow anarchic flavor structure
 - Straightforward mediation requires sfermions $m \gtrsim 100$ TeV
 - (Maybe this is our universe? Not quite natural....)

MSSM SPECTRA

- **Low-scale** SUSY breaking

- e.g.: gauge-mediated $\gtrsim 10 \text{ TeV}$

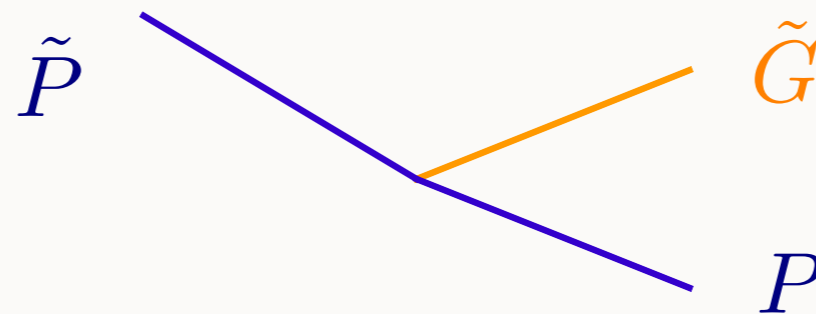


- **gravitino mass:** $m_{3/2} = \frac{\Lambda_{SUSY}^2}{M_{Pl}}$

- **soft masses:** $m_{soft} \sim \frac{\alpha}{4\pi} \Lambda_{SUSY}$

MSSM SPECTRA

- Gravitino is the LSP
 - Cosmology very different - no more neutralino dark matter
 - Now **charged** superpartners can be the **NLSP**:



- Decay of NLSP to gravitino can be **prompt** or **displaced**
- Big plus: neatly solves **flavor** problem

MSSM SPECTRA

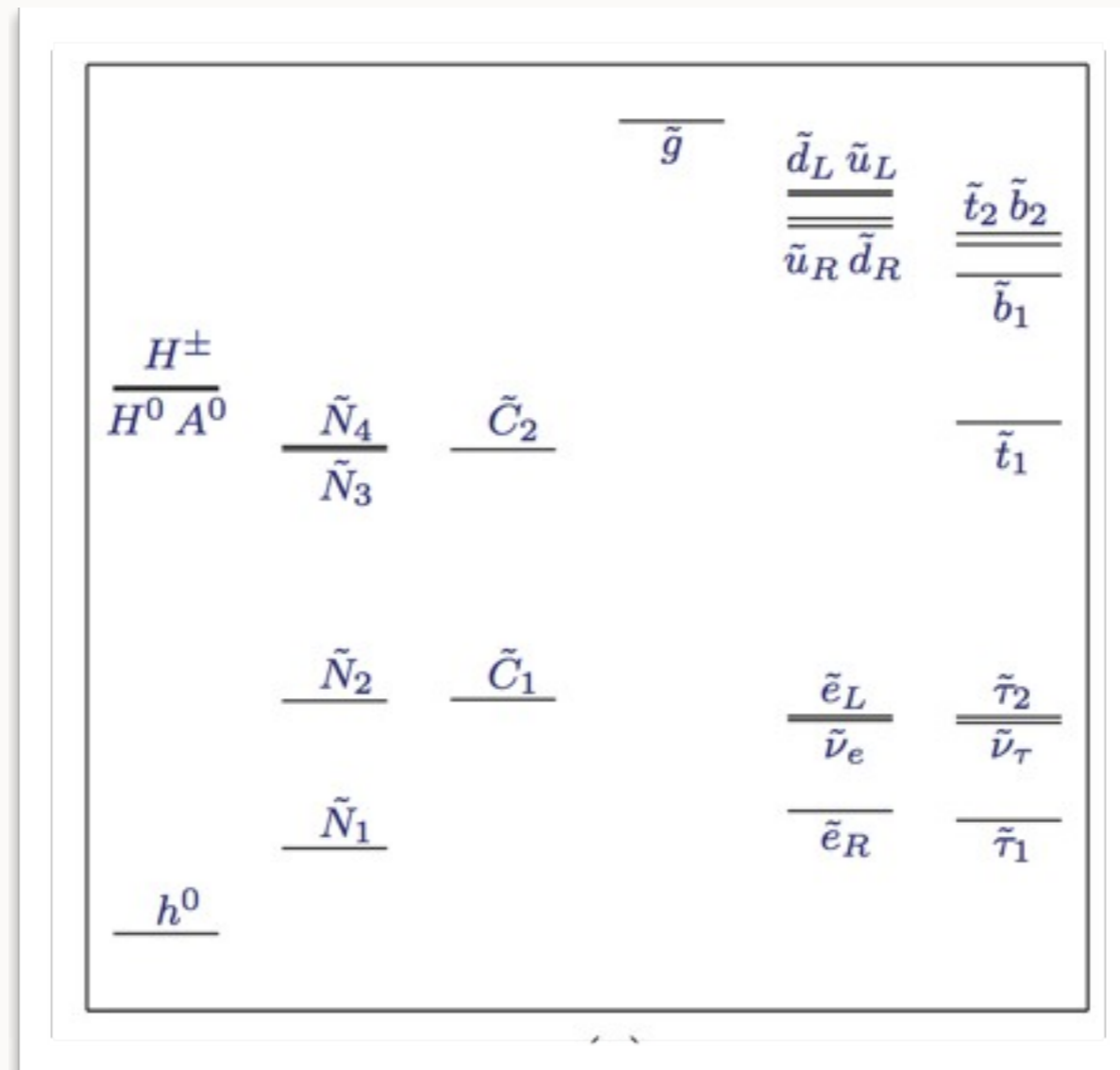
- But...
 - adding charged mediators can make it tricky to maintain gauge unification
 - new cosmological problems with non-thermally produced stable gravitinos
 - Biggest disadvantage: hard to accommodate $m_h = 125 \text{ GeV}$

MSSM SPECTRA

- Have never yet found a completely convincing top-down model of SUSY-breaking
- Since experimental signatures extremely sensitive to detailed spectrum, also important to consider **bottom-up** approaches to make sure bases are covered

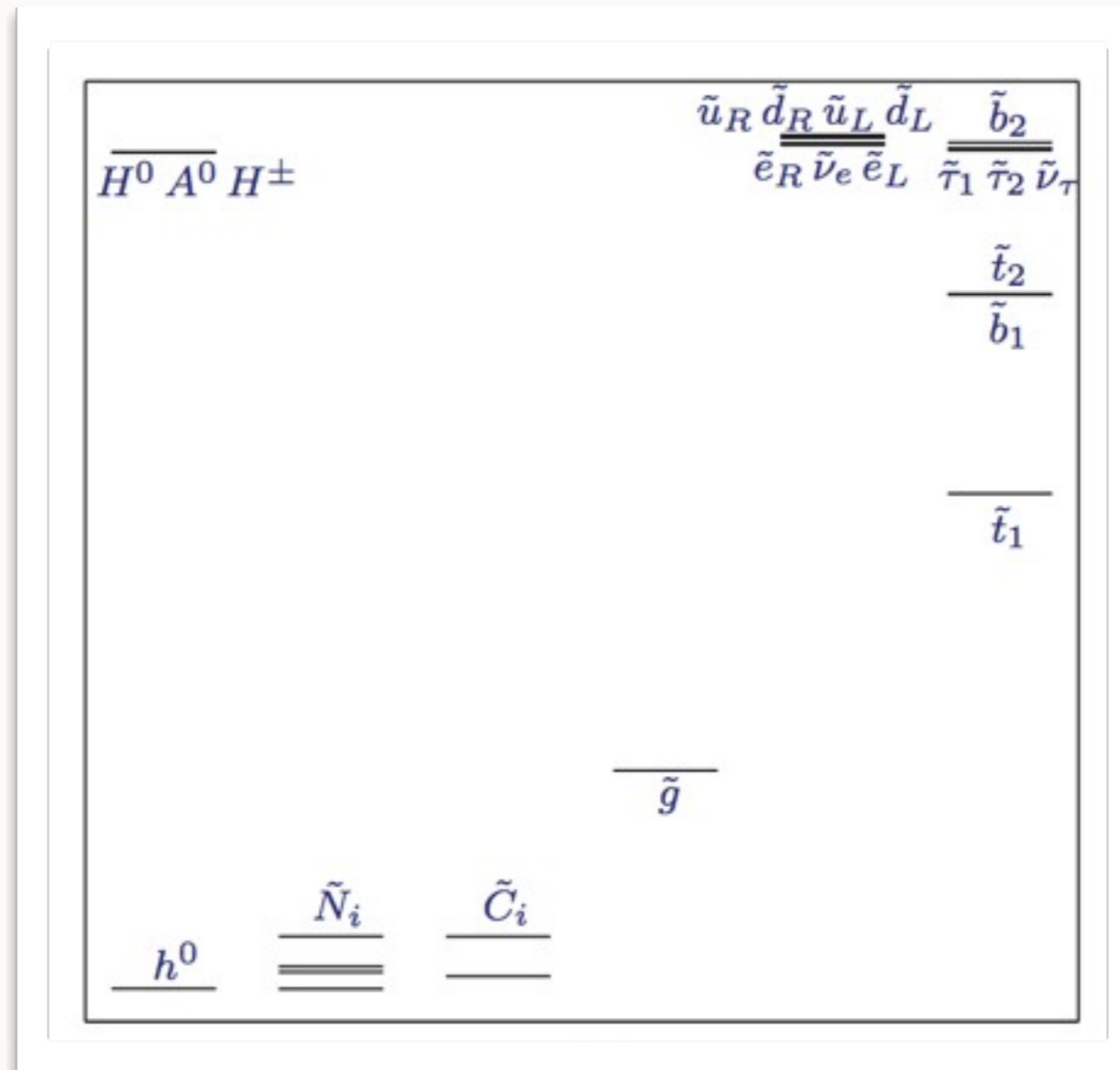
MSSM SPECTRA

- Example gravity-mediated spectrum



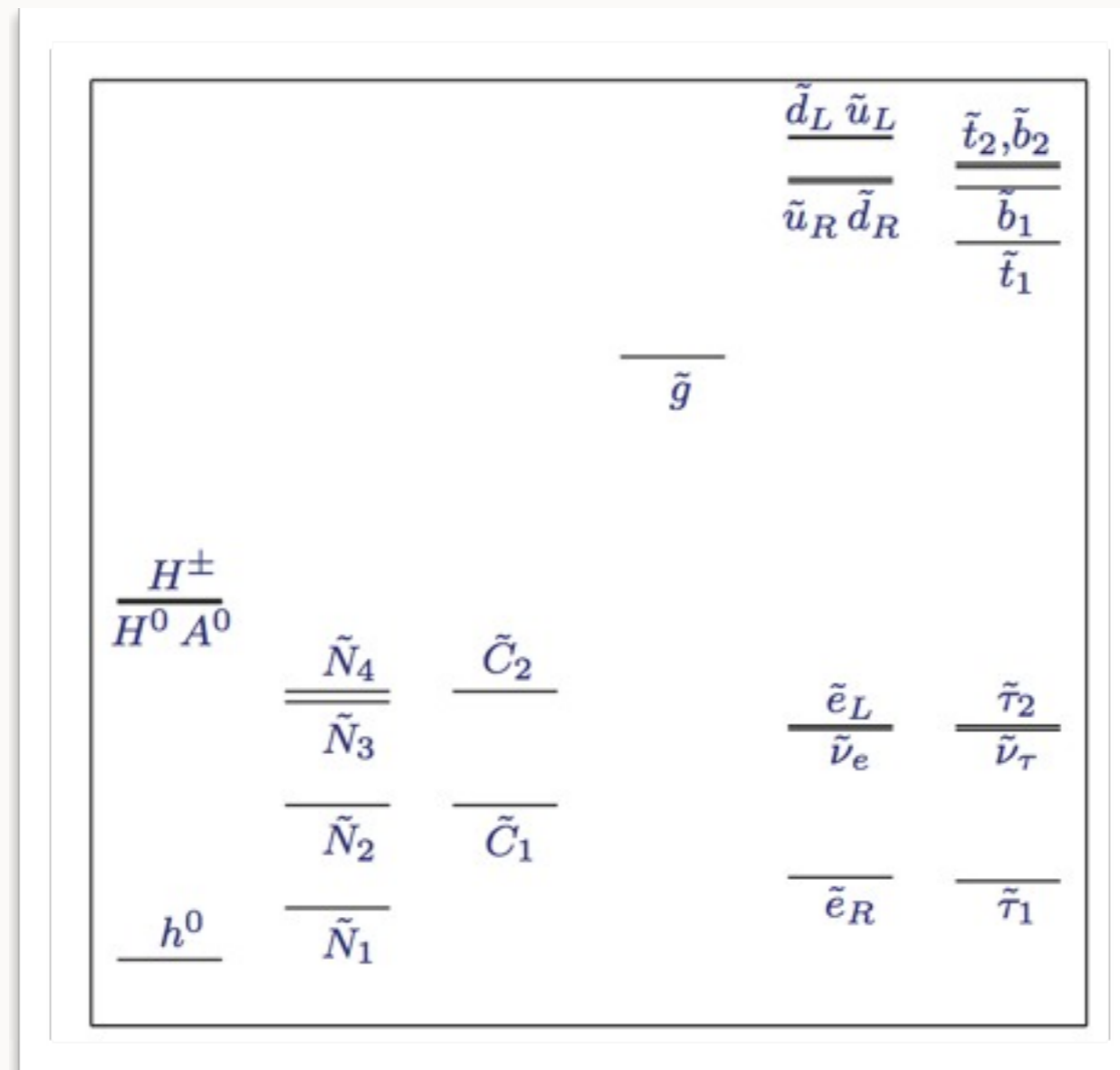
MSSM SPECTRA

- Example gravity-mediated spectrum



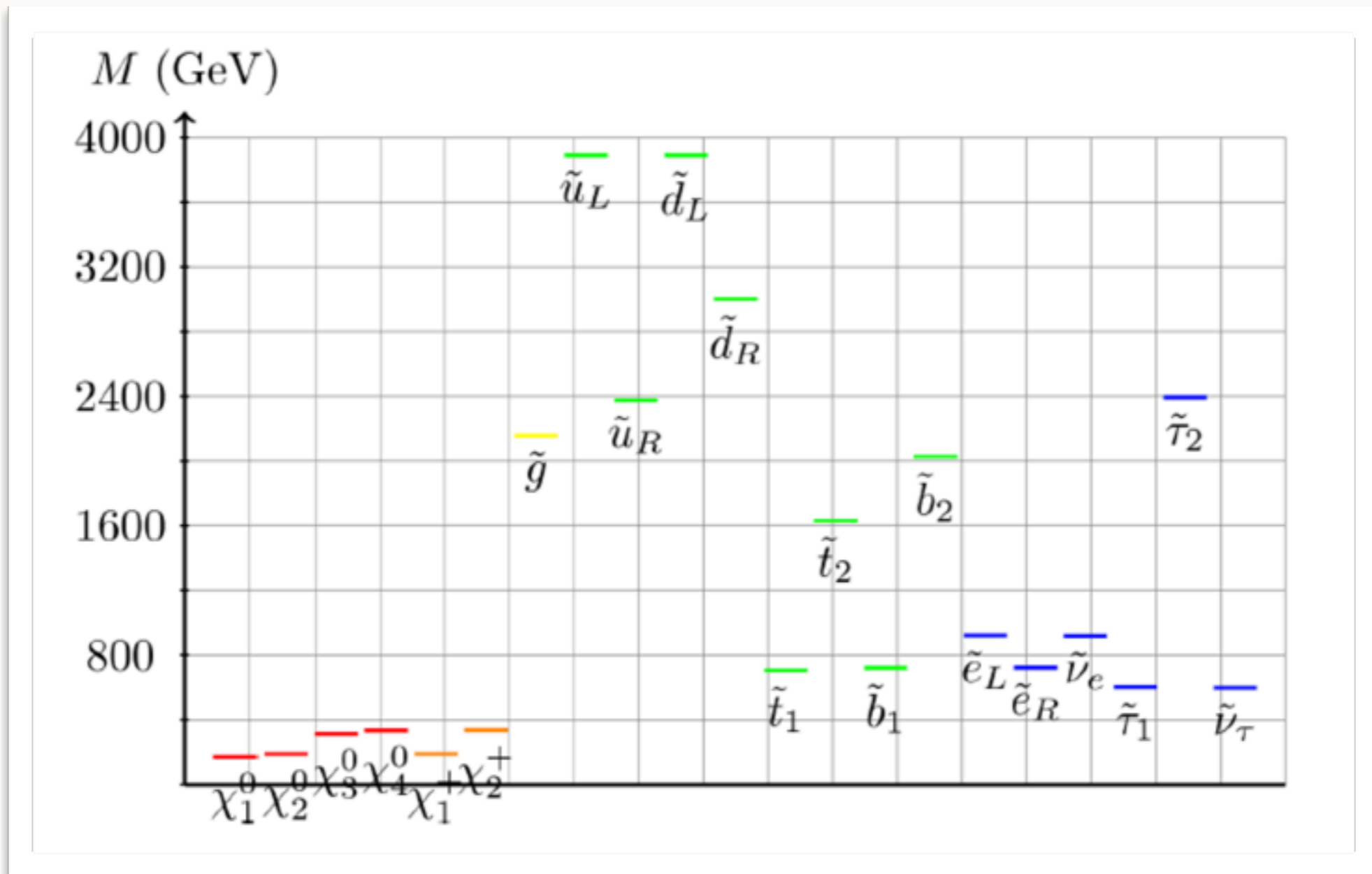
MSSM SPECTRA

- Example gauge-mediated spectrum



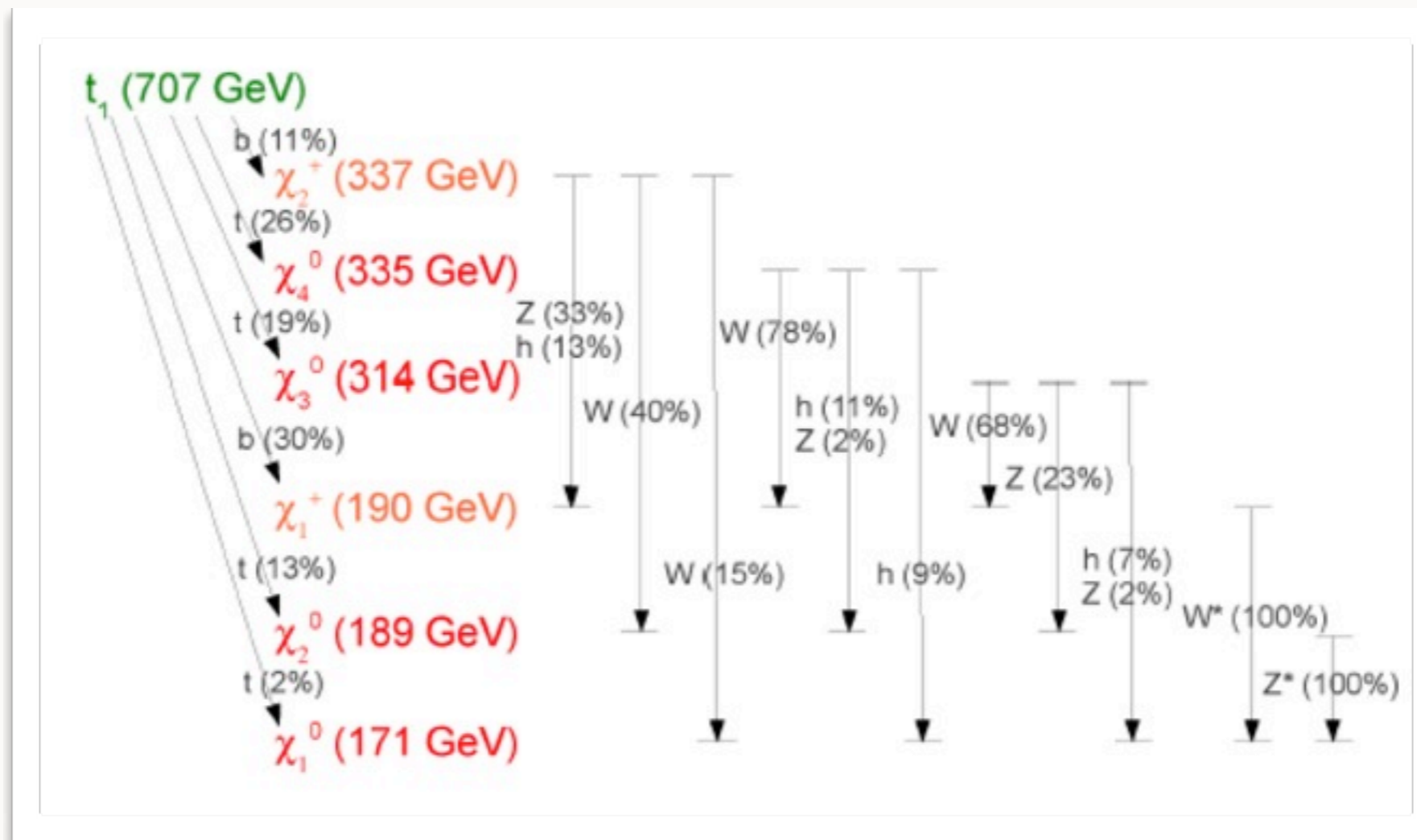
MSSM SPECTRA

- Example bottom-up spectrum



MSSM SPECTRA

- Rich spectrum means complicated decays:



MSSM SIGNALS

- Two questions to end on:
 - Given enormous complexity and variability of signals, how can we best design SUSY searches at colliders?
 - What should we make of lack of BSM signals so far?