#### Booster Neutrino Beamline Flux Prediction

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- Proton delivery:
  - 8 GeV protons from Booster
  - Average rate up to 5Hz
  - 4.2e12 PPP
- Horn:
  - Neutrino mode +170kA
  - Antineutrino mode -170kA





## Experiments

- Past/Current
  - MiniBooNE
  - MicroBooNE
  - SciBooNE
  - SciBath
- Future:
  - LAr1ND
  - ICARUS
  - CENNS/Captain

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## Neutrino flux prediction

- Geant4 based MC used to
   predict the flux
- Hadron production cross sections tuned to external data



	$ $		$\overline{\nu}_{\mu}$ $3.26 \times 10^{-11}$	
Flux $(\nu/\mathrm{cm}^2/\mathrm{POT})$				
Frac. of Total		93.6%		5.86%
Composition	$\pi^+$ :	96.72%	<b>π</b> <sup>-</sup> :	89.74%
	$K^+$ :	2.65%	$\pi^+ \rightarrow \mu^+$ :	4.54%
	$K^+ \rightarrow \pi^+$ :	0.26%	$K^{-}$ :	0.51%
	$K^0 \rightarrow \pi^+$ :	0.04%	$K^0$ :	0.44%
	$K^0$ :	0.03%	$K^0 \rightarrow \pi^-$ :	0.24%
	$\pi^-  ightarrow \mu^-$ :	0.01%	$K^+ \to \mu^+$ :	0.06%
	Other:	0.30%	$K^- \rightarrow \pi^-$ :	0.03%
			Other:	4.43%
	$\nu_e$		$\overline{ u}_e$	
Flux $(\nu/\mathrm{cm}^2/\mathrm{POT})$	2.	$.87 \times 10^{-12}$		$3.00 \times 10^{-13}$
Frac. of Total		0.52%		0.05%
Composition	$\pi^+ \rightarrow \mu^+$ :	51.64%	$K_L^0$ :	70.65%
	$K^+$ :	37.28%	$\pi^-  ightarrow \mu^-$	19.33%
	$K_L^0$ :	7.39%	$K^-$ :	4.07%
	$\pi^+$ :	2.16%	$\pi^-$ :	1.26%
	$K^+ \rightarrow \mu^+$ :	0.69%	$K^- \rightarrow \mu^-$ :	0.07%
	Other:	0.84%	Other:	4.62%

## **Pion production**

- Sanford-Wang fits:
  - HARP (thin target)
    - 8.89GeV p on Be target
    - P = 0.75 6.5 GeV/c, $\theta = 30 - 210 \text{ mrad}$
  - E910
    - 6.4, 12.3, 17.5 GeV/c
    - $P=0.4 5.6 \text{ GeV/C}, \theta = 18 400 \text{ mrad}$
- Fits done both for pi+ and pi-

Phys. Rev. D79, 072002 (2009)



## Kaon production

- Feynman scaling based parameterisation used to fit world K+ production data
- Sanford-Wang fits to K0s production data from BNL E910 and KEK Abe et al.



# Kaons in BNB

- Kaon production further constrained by SciBooNE measurements
- Found production to be 0.85+-0.12 relative to the global fit to kaons (with 30% error)





## Wrong signs

- Significant numu component in antineutrino mode (~16%)
- Not constrained well by HARP/E910
- In-situ measurement in MiniBooNE
  - Numu CCQE angular fit
  - $CC\pi^+$  rate
  - µ<sup>-</sup> capture





## Flux uncertainty

- Propagate uncertainties using many MC worlds to build error matrices that capture correlations between bins of neutrino observables
  - spline fits through HARP data
  - kaon fits
  - Hadron cross sections on Be and Al
  - Horn focusing
  - POT counting
- For numu/numubar CCQE measurement resulting flux uncertainty was at 9-10% level

#### **Reinteractions in BNB**

 At Booster energies ~90% of pions contributing to neutrino flux from primary p+Be interactions



## HARP thick target analysis



DATA	Beam radius cut (reduce the edge effect)	P.O.T
MB100	0.4 cm	622791
MB50	0.4 cm	814749
Empty	0.4 cm	475776
Be5	1.0 cm	13070000
Empty	1.0 cm	1990000

# **Revisiting HARP**

- BNB MC correctly models reinteractions – see talk by A. Wickremasinghe at flux workshop (09/22)
  - Extrapolation to thick target agrees at 1% level
- Analysis using alternative (Extended Sanford Wang) parametrisation fitted to HARP only data results in 2% change in neutrino flux
  - Within the expected systematic error



## Single detector

- MiniBooNE  $\nu_{_{\!\!\!\!\!\!0}}$  appearance analysis further constrains flux by simultaneously fitting  $\nu_{_{\!\!\!\!\!\!\!\!\!\!\!}}$  and  $\nu_{_{\!\!\!\!\!\!\!\!\!\!\!}}$ 



## Multiple detectors

- SciBooNE/MiniBooNE muon (anti)neutrino disappearance analysis
- Flux/cross-section uncertainties cancel, detector do not



## Multiple detectors

- LAr1ND will serve as a near detector for MicroBooNE and any future Far Detector experiment
- Compared to Sci/MiniBooNE expect better cancellation of detector systematics



#### Low energy neutrinos at BNB



 Looking at neutrinos near target hall



# Low energy neutrino physics

- Core collapse supernova physics
  - Only a couple of v-N cross sections are measured in SN energy range
  - Important for SN neutrino detection
  - Understanding the supernova explosion process
- Coherent Elastic Neutrino Nucleus Scattering (CENNS)
  - Has not been observed yet
  - Background for dark matter searches
  - Neutron form factor
  - Neutrino magnetic moment
  - Test week mixing angle
  - Non Standard Model Interactions
- Sterile neutrinos and neutrino oscillations
  - L/E with lower energy -> observe oscillation in detector

Slide courtesy J. Yoo

## Multipurpose beamline



- Search for Dark Matter in beam off-target mode
- Reduce neutrino flux by ~50 times (background for DM search)
- Collected 1.9e20 POT



## Summary

- 12 years of experience running with BNB
- Well understood neutrino flux
  - Constrained using external hadron production measurements
  - Years of running experiments
- Many experiments using the BNB facility
- Exciting future experimental program ahead

# Backup

## Stable running

• 2 target/horn assemblies

/POT × 10 -17

160

140

120

100

80

60

40

20

• 96 million pulses with 1<sup>st</sup> horn

6.27e+20 v POT

1.13e+21 <sup>7</sup> POT

31/Dec/04

01/Jan/04

- 367 million pulses with 2<sup>nd</sup> horn
- Neutrino flux remained extremly stable over 12 years of running
- No degradation in flux or energy spectrum

