

NOvA Analysis Status

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NOvA Operational Readiness Review

October 28, 2014



A spin through NOvA's analysis structure

- A distributed effort. Organization:

NC and ν_e CC

ν_μ CC

Exotics

Interaction “channels”
(multiple physics topics in each)

Reconstruction

Detector simulation

Beam data and simulation

Calibration and alignment

Computing, production, software

Support and infrastructure

Data quality

Triggers

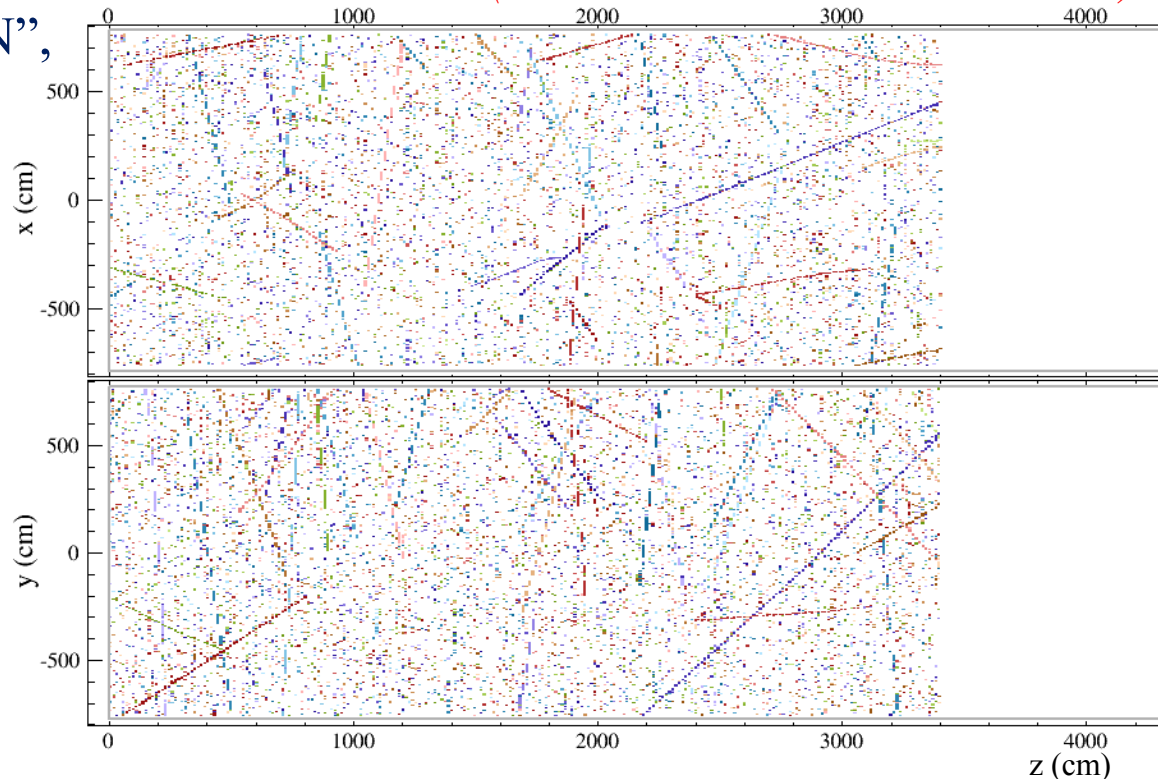
Data collection

- Plus lots of **activity at the group boundaries**
- Focusing here on items directed at the **3-flavor $\nu_\mu \rightarrow \nu_\mu$ and $\nu_\mu \rightarrow \nu_e$ analyses**

Isolating individual interactions

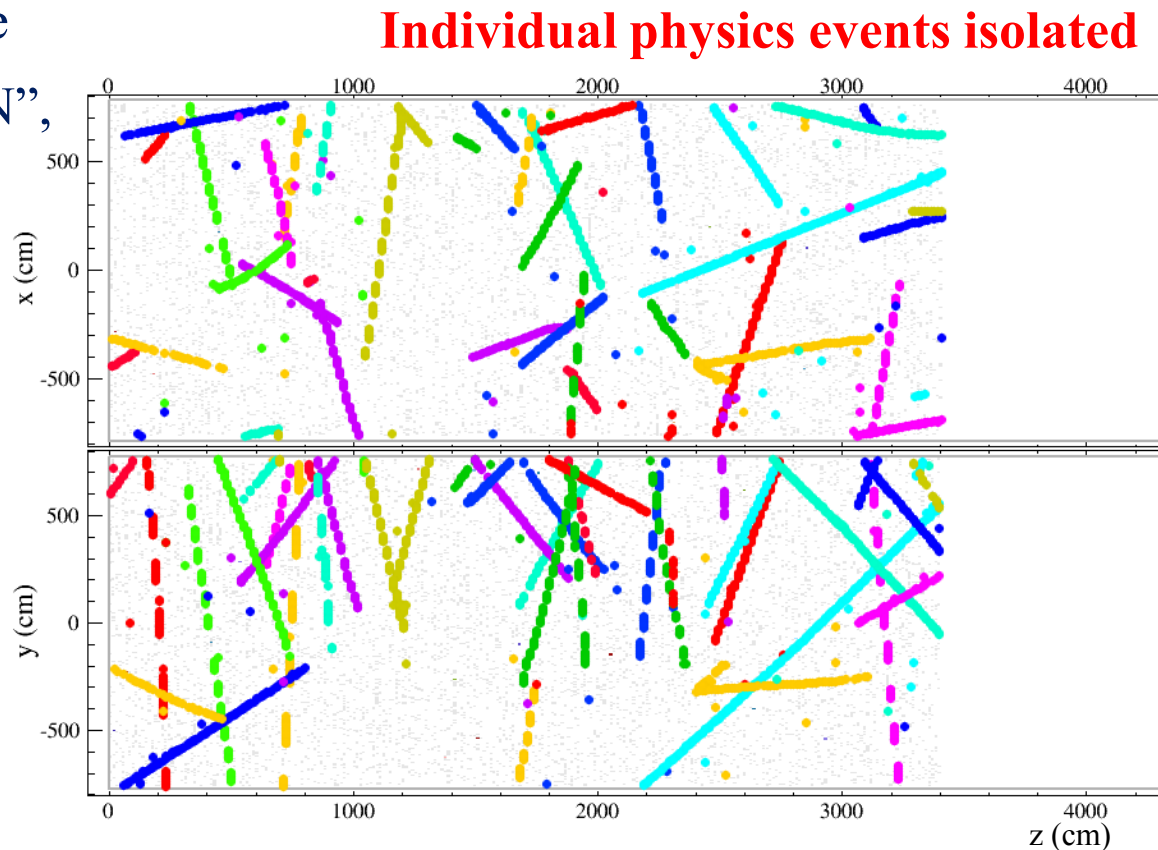
- A standard trigger in the Far Detector (FD) records 550 μs of activity:
 - *hundreds of **noise hits** (since we keep the DAQ thresholds as low as possible)*
 - *about **50 cosmic rays***
 - *and rarely, a **neutrino interaction***
- **Isolating these components** is the first step of the offline analysis sequence
- Algorithm based on “DBSCAN”,
M. Ester *et al.* (1996)
- Looks for causally connectable **clusters in space/time**, also using knowledge of how noise hits behave
- *According to FD simulation:*
Avg. completeness: 99.3%
Avg. purity: 99.5%
(Actually improved beyond this now...)

A NuMI trigger recorded in April
(detector still under construction)



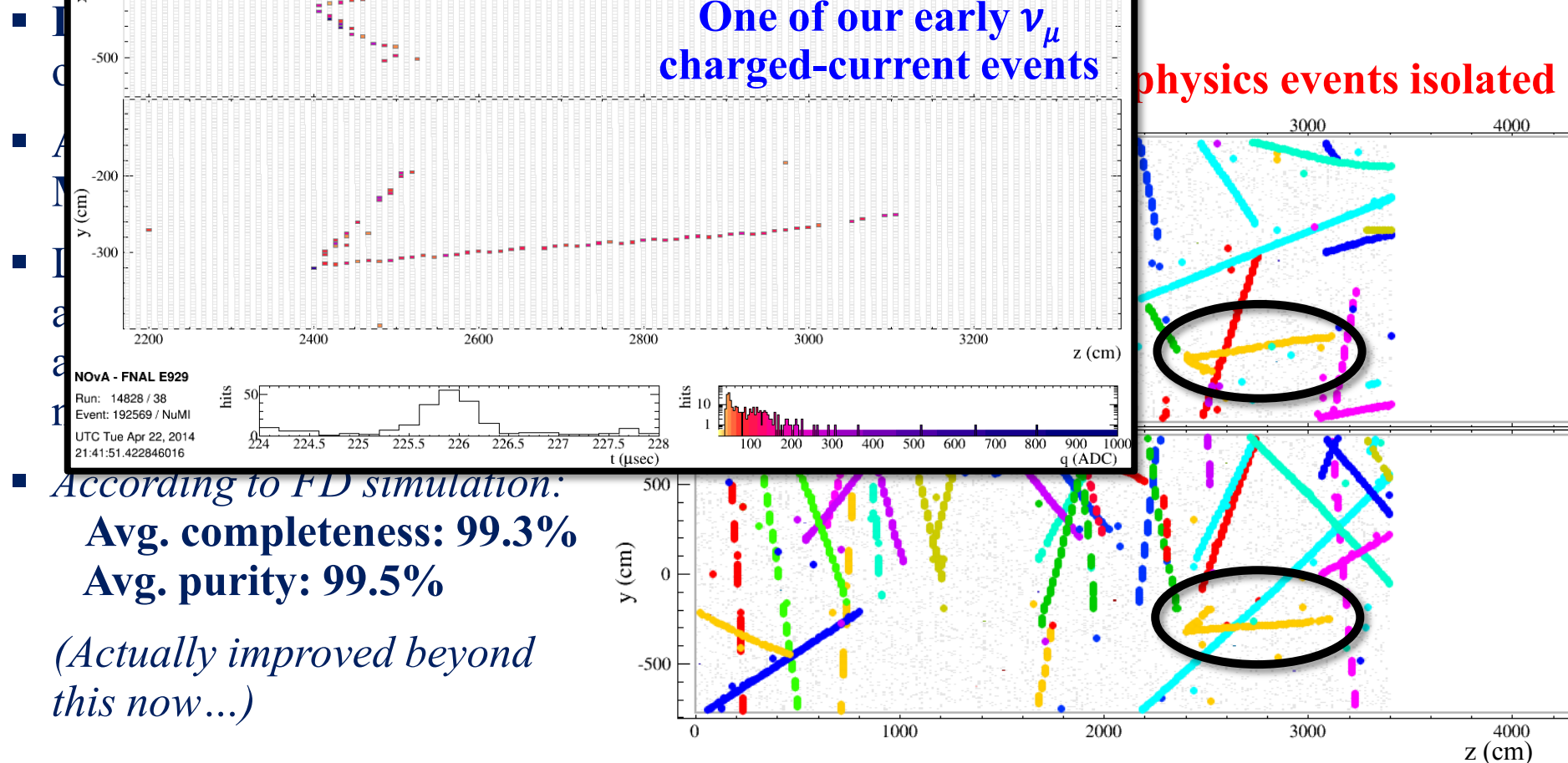
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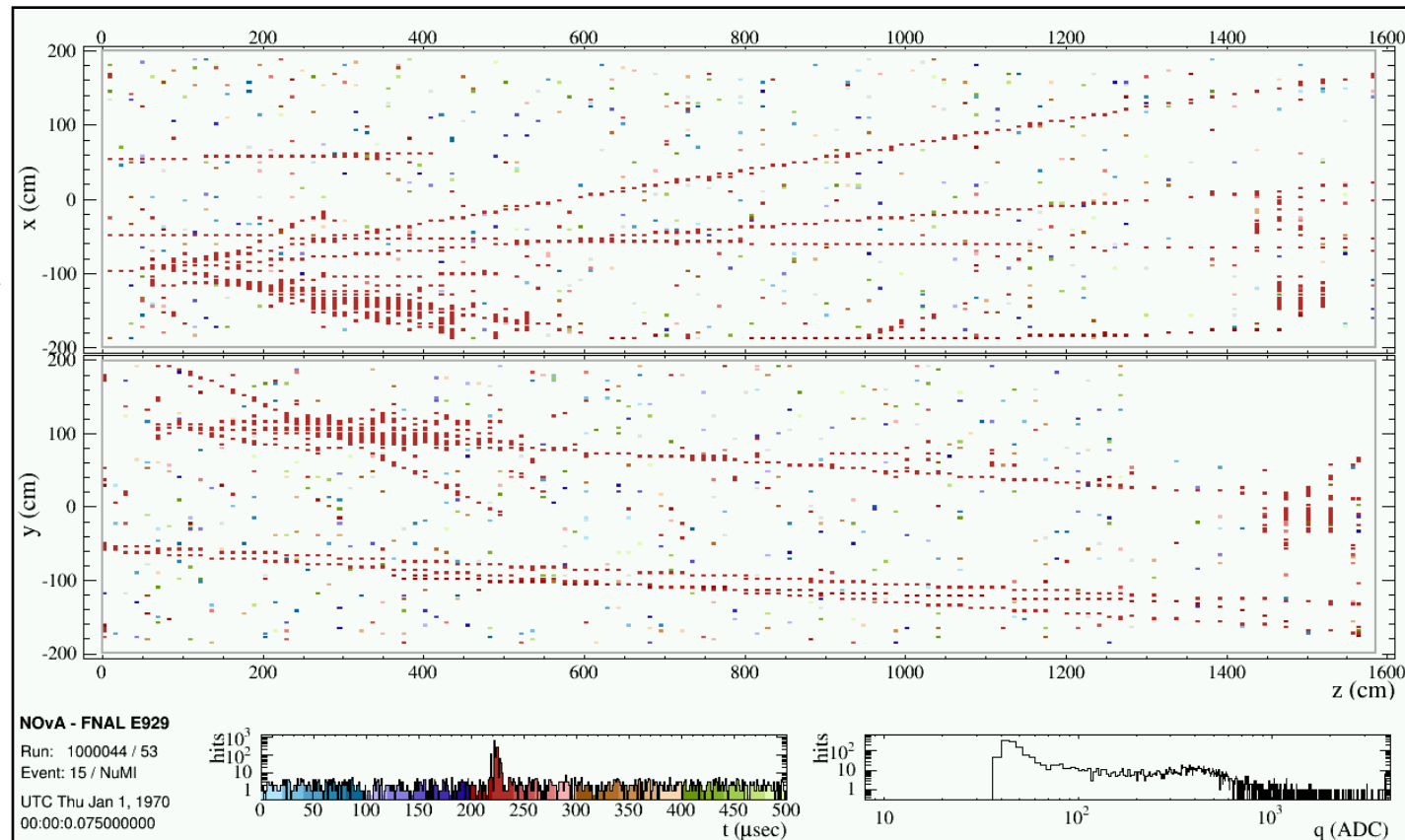
- A standard trigger in the Far Detector (FD) records 550 μs of activity: (as low as possible)



And in the ND...

- ND event pile-up easily handled
- Cosmic ray rate much reduced (50 Hz in ND vs. 100,000 Hz in FD)
Instead, neutrino interactions in the rock send particles (mostly muons) into the detector

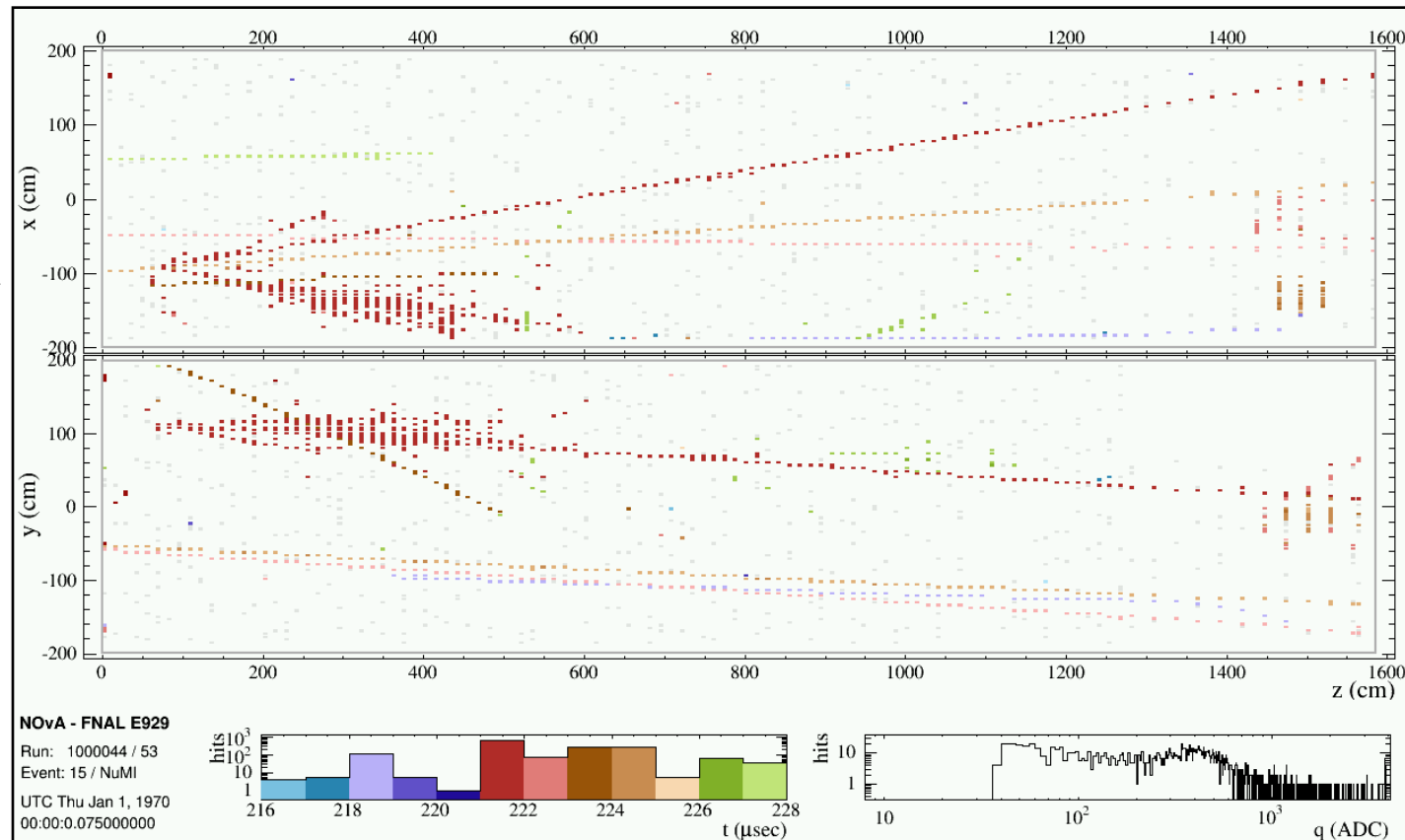
*5×10^{13} ppp ND event
(i.e., design intensity)
shown here, from
simulation*



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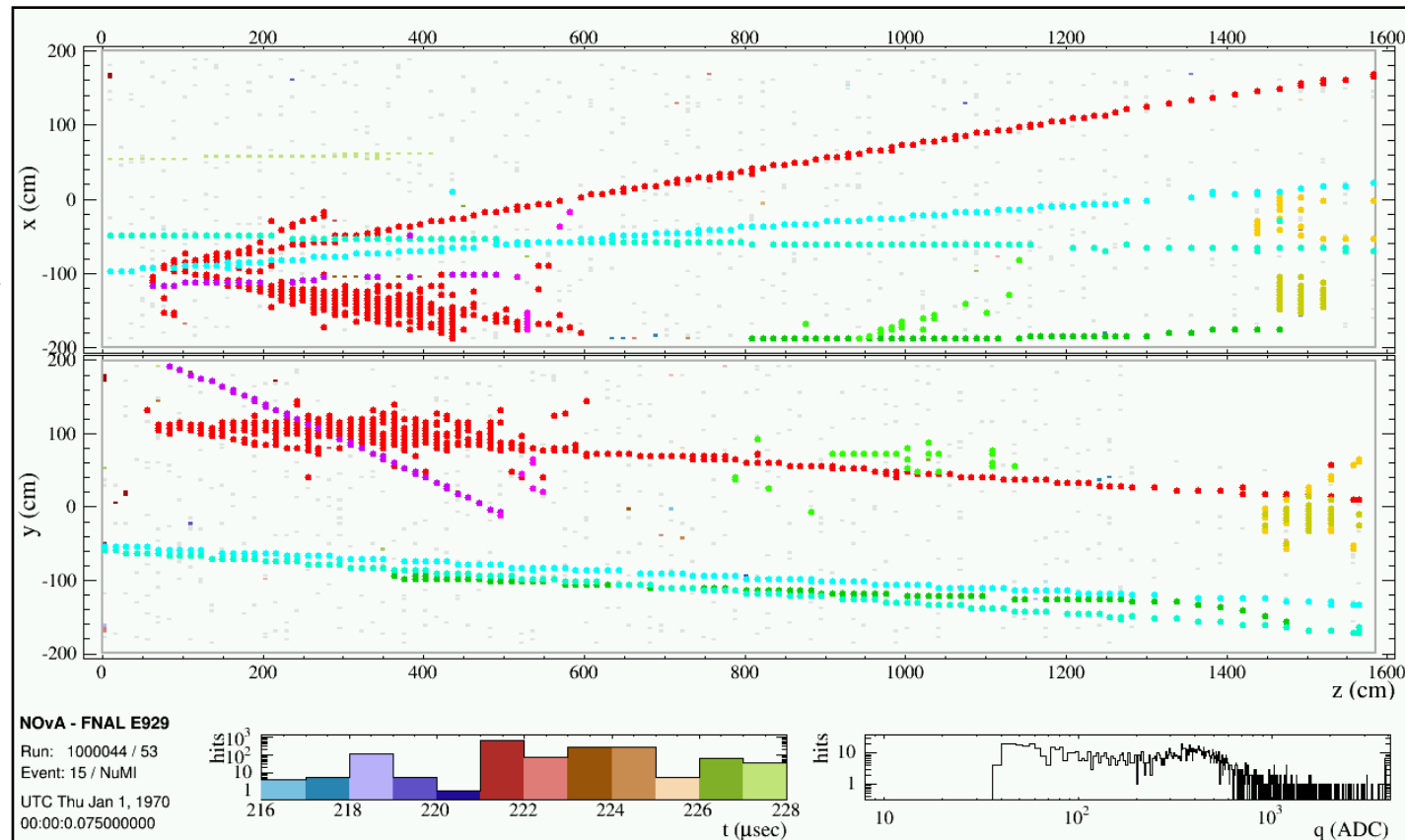
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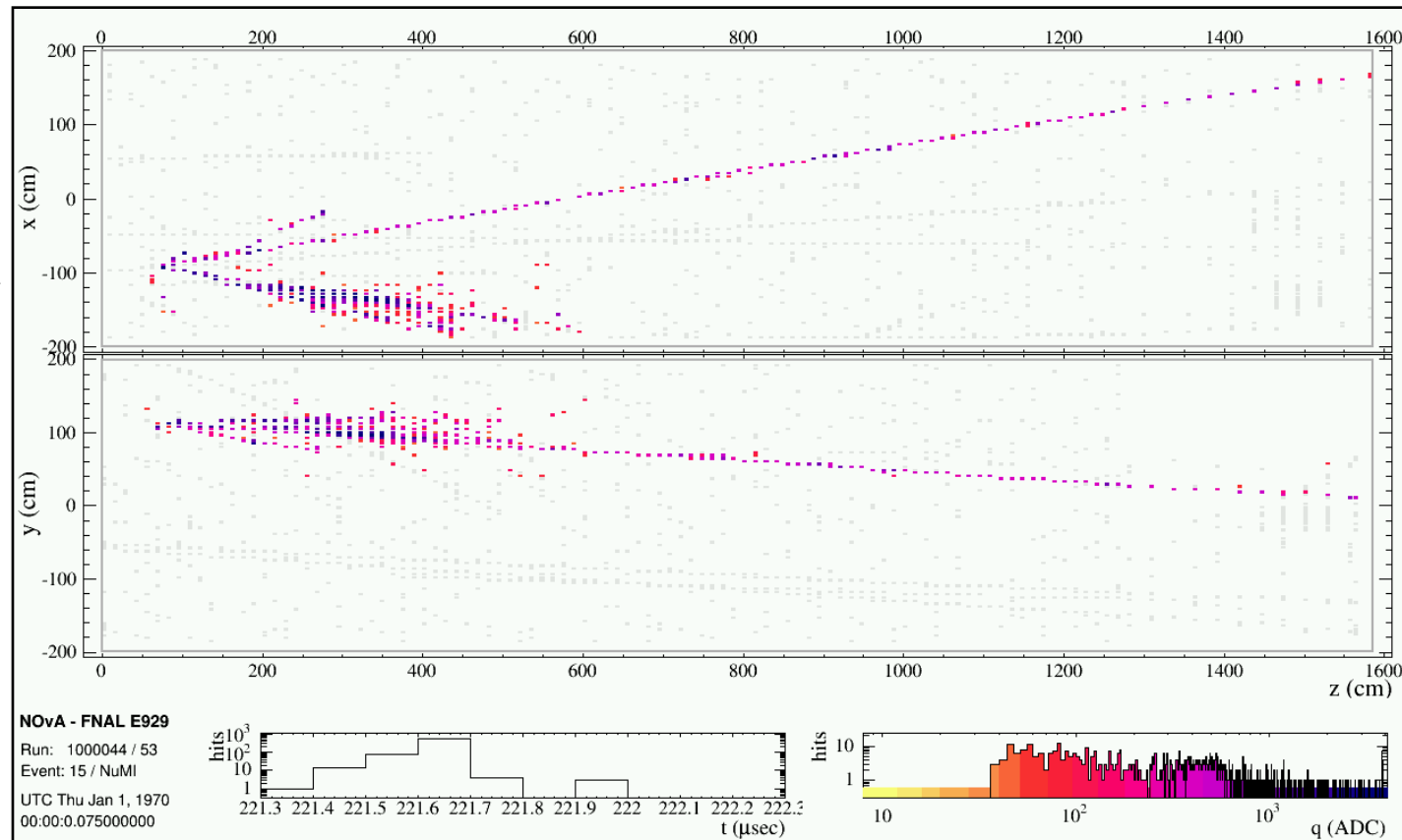
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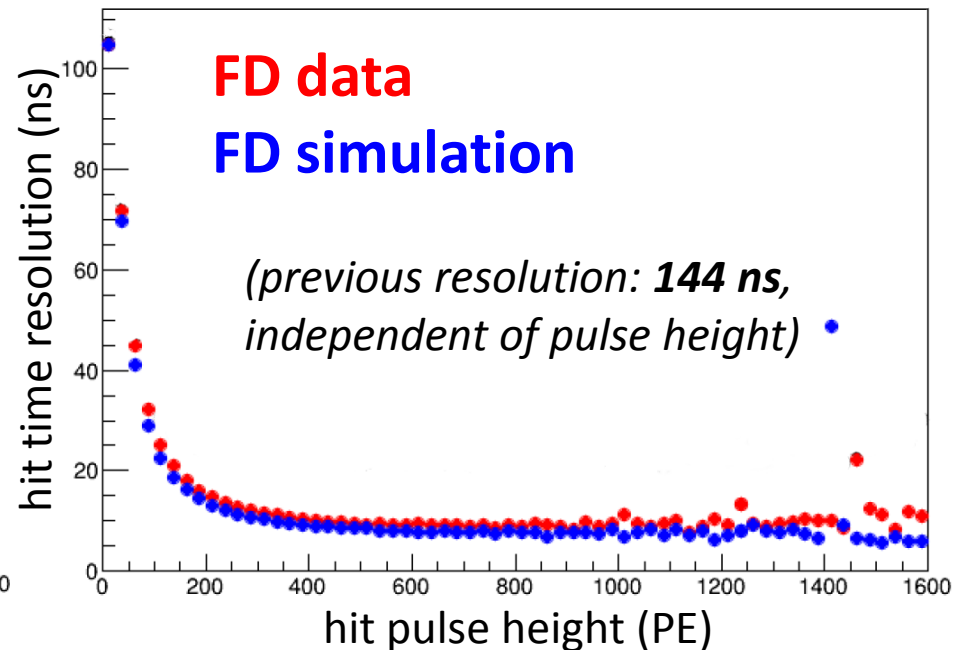
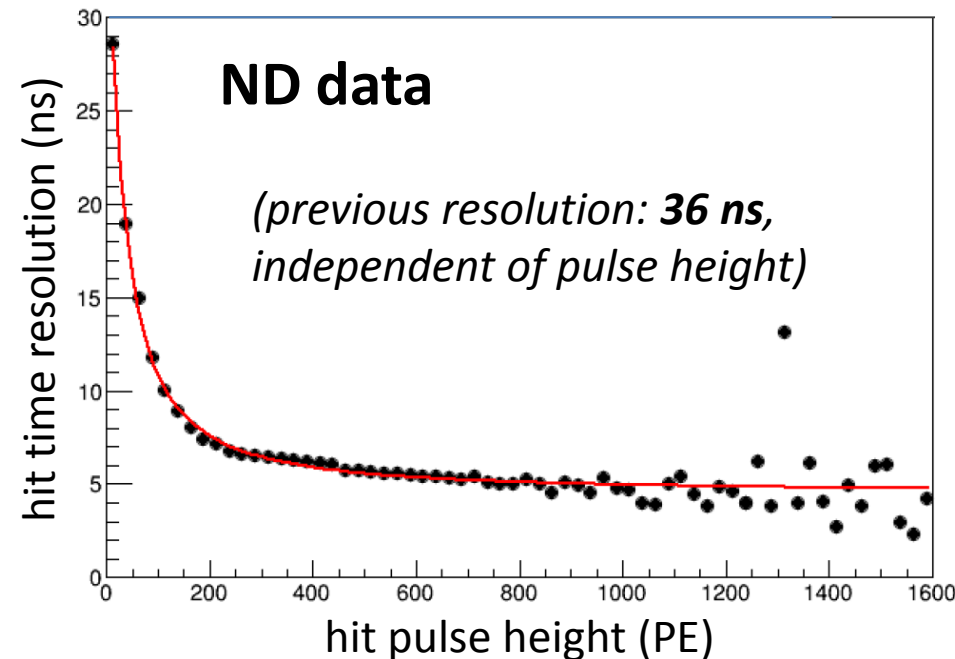
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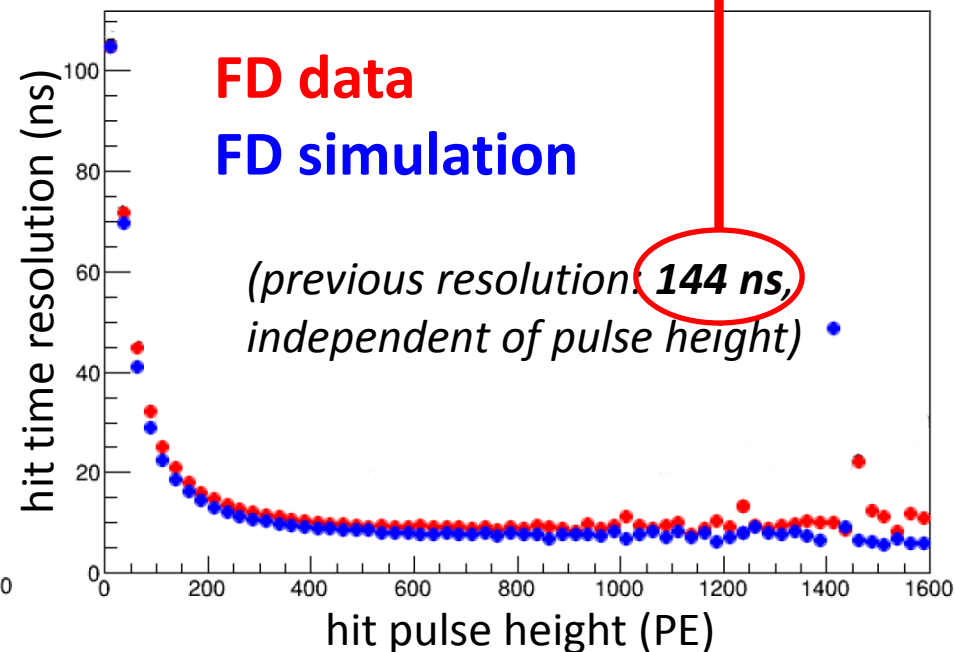
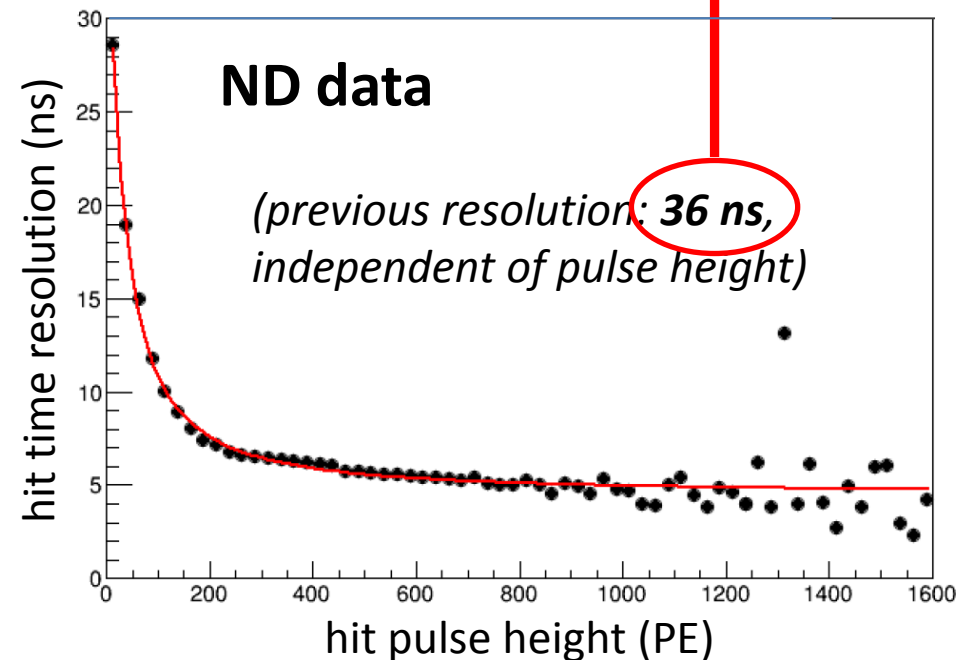
Timing resolution

- **Newly deployed firmware** leads to substantial improvement in timing resolution
- **Fully incorporated** into calibration procedures, simulation packages, and analysis software
- Benefitting event clustering and opening new lines of analysis



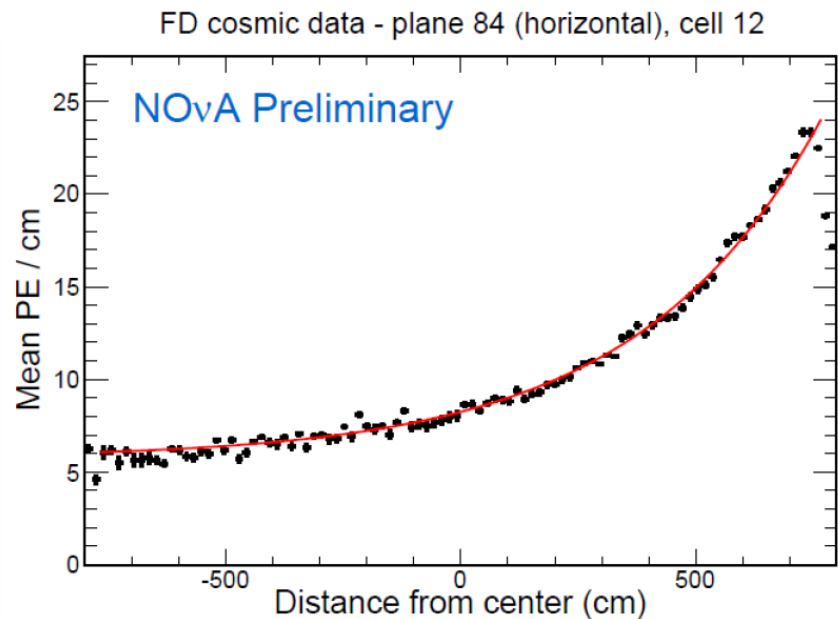
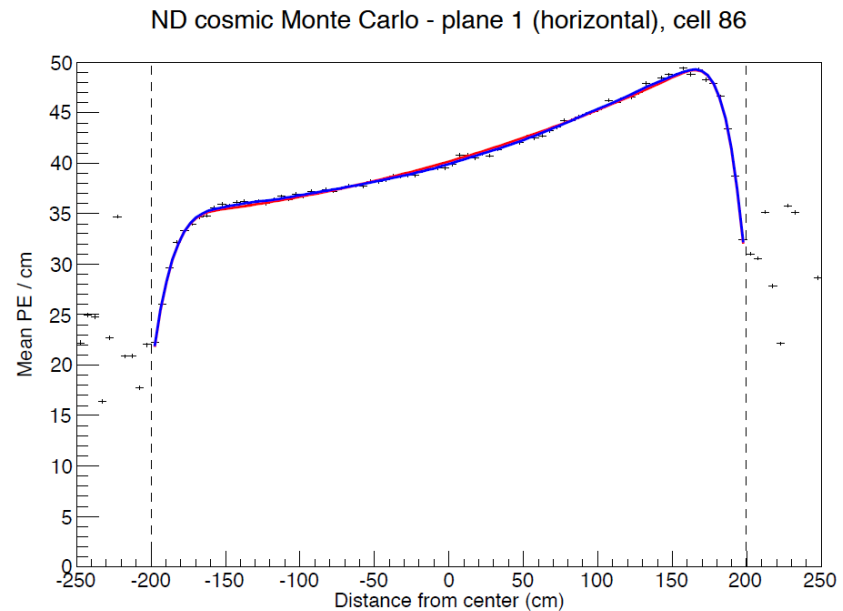
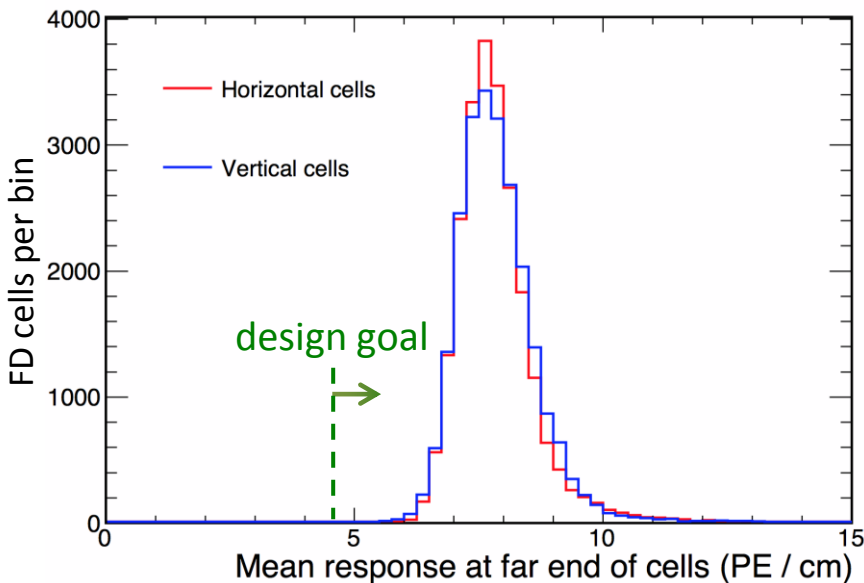
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Energy calibration

- **Biggest effect** that needs correction is **attenuation** in the WLS fiber
- A local regression **corrects any residuals** beyond the basic functional form
- **Light level requirements at end of cell are well met**



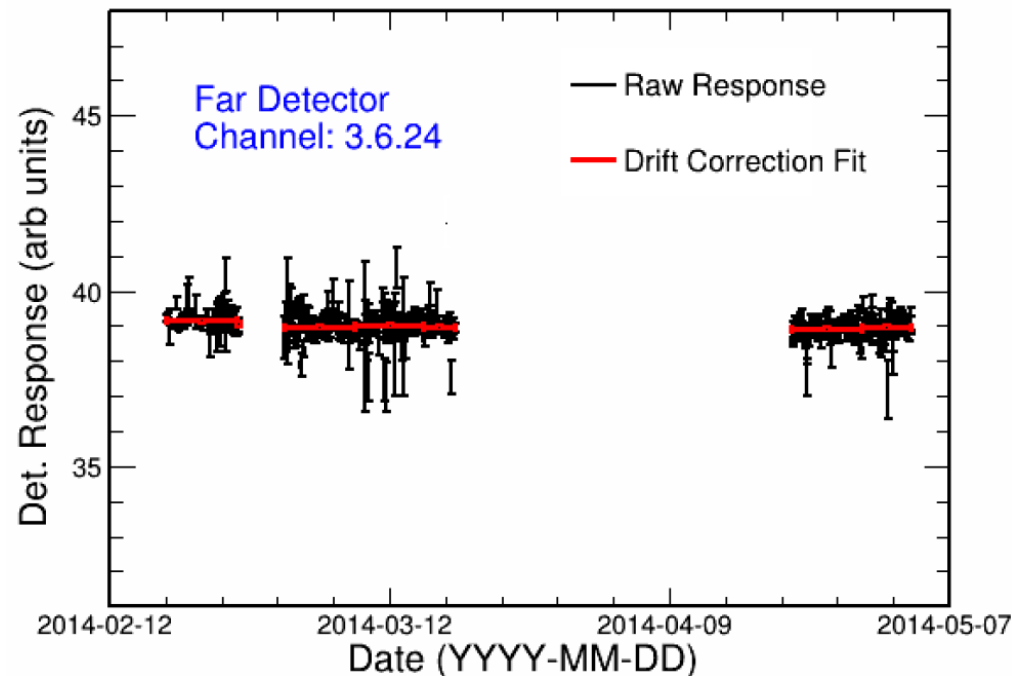
To APD

1560 cm

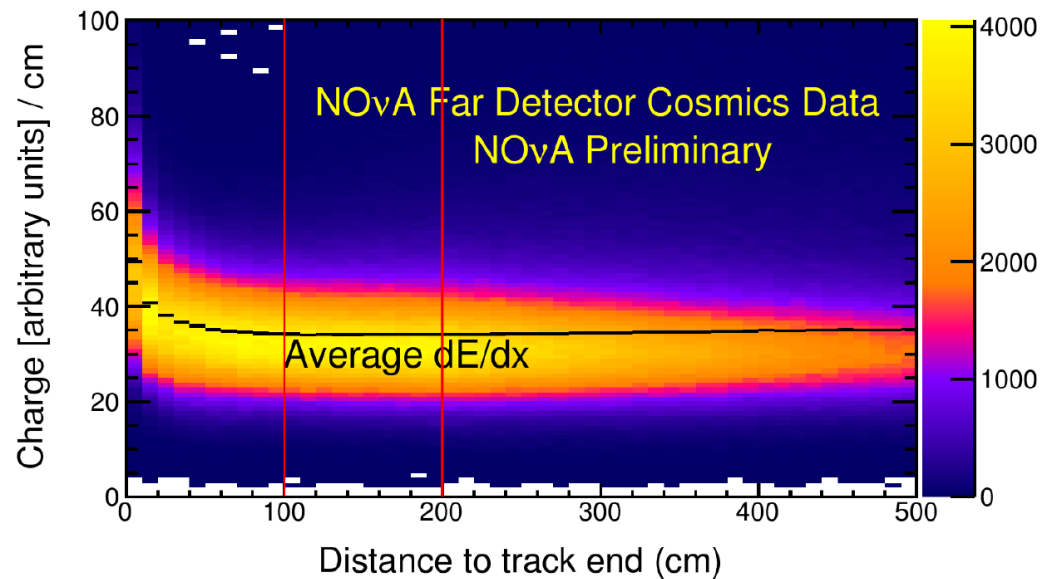
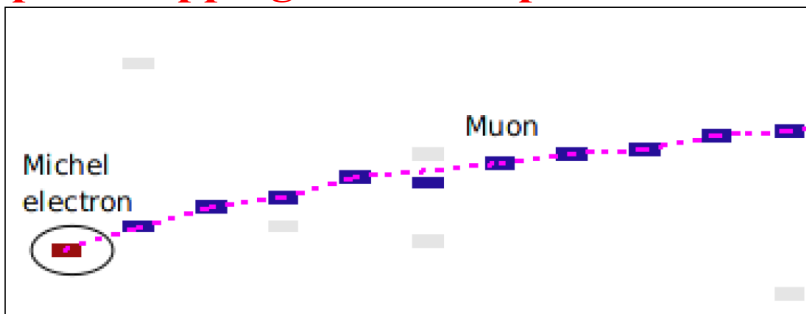
4 cm × 6 cm

Energy calibration

- A **drift correction** handles changes in charge response over time (included that due to swapped hardware)
- **Stopping muons** provide a standard candle for setting the **absolute energy scale**

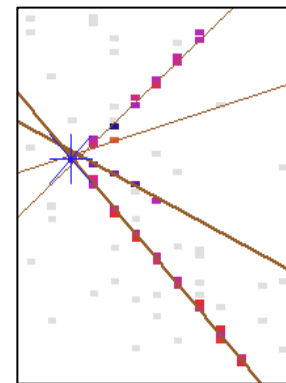
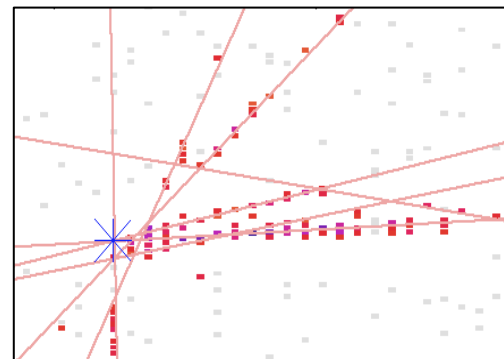
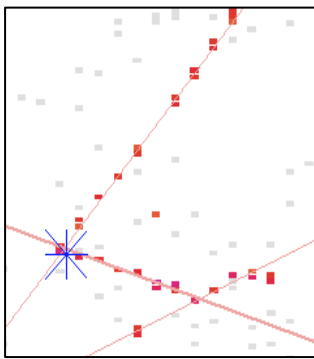


Michel electron tag yields very pure stopping muon sample

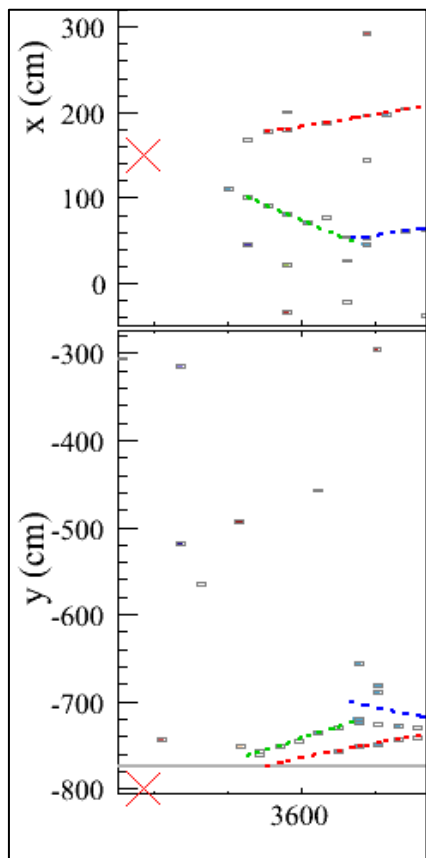


Event vertexing

- Find lines of energy depositions using a Hough transform.



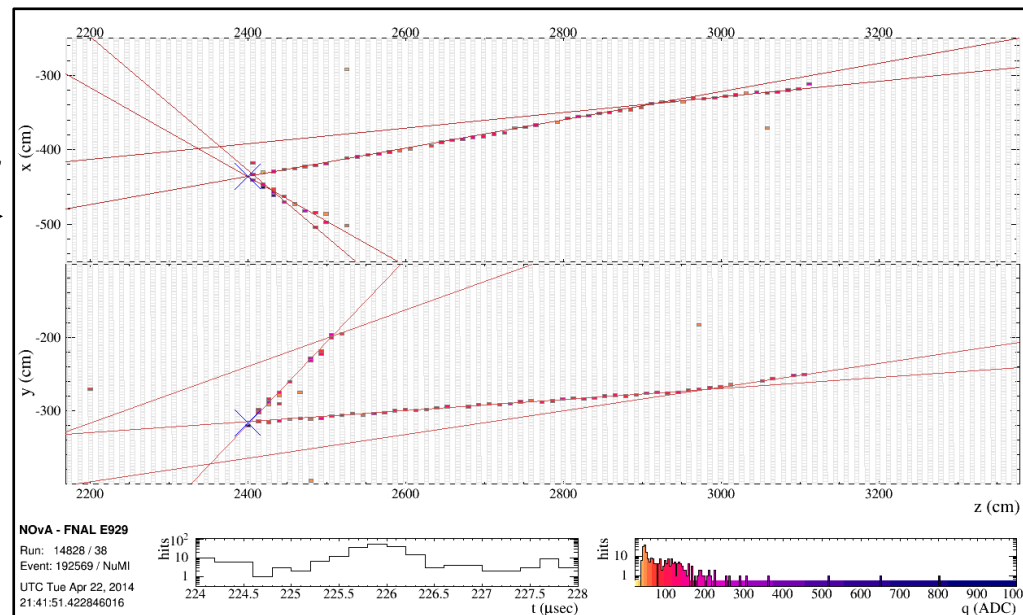
Examples...



The FD ν_μ CC data event from earlier



A vertex outside the detector



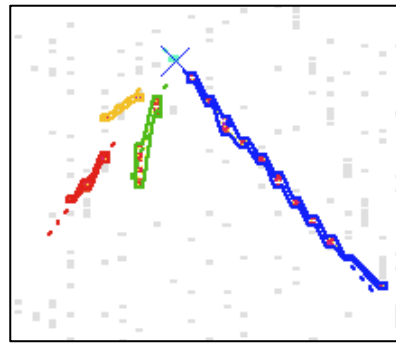
Vertex resolution for **charged-current events: 11 cm**
Vertex resolution for **neutral-current events: 29 cm**

→ Note: this is one of several *vertexing algorithms* implemented

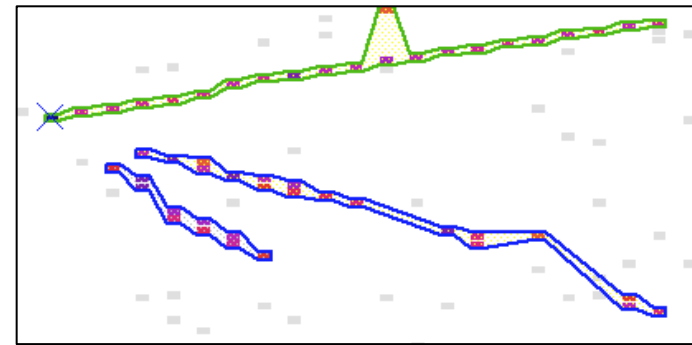
Prong clustering

- Given a seed vertex, look for **clusters in angular space** around it.
- Energy depositions are also used in the clustering metric, and hits can be shared between the resulting “prongs”.
- Prongs in each view are matched based on topology and dE/dx to form 3D objects.

Prongs are drawn here by outlining the cells that belong to them.

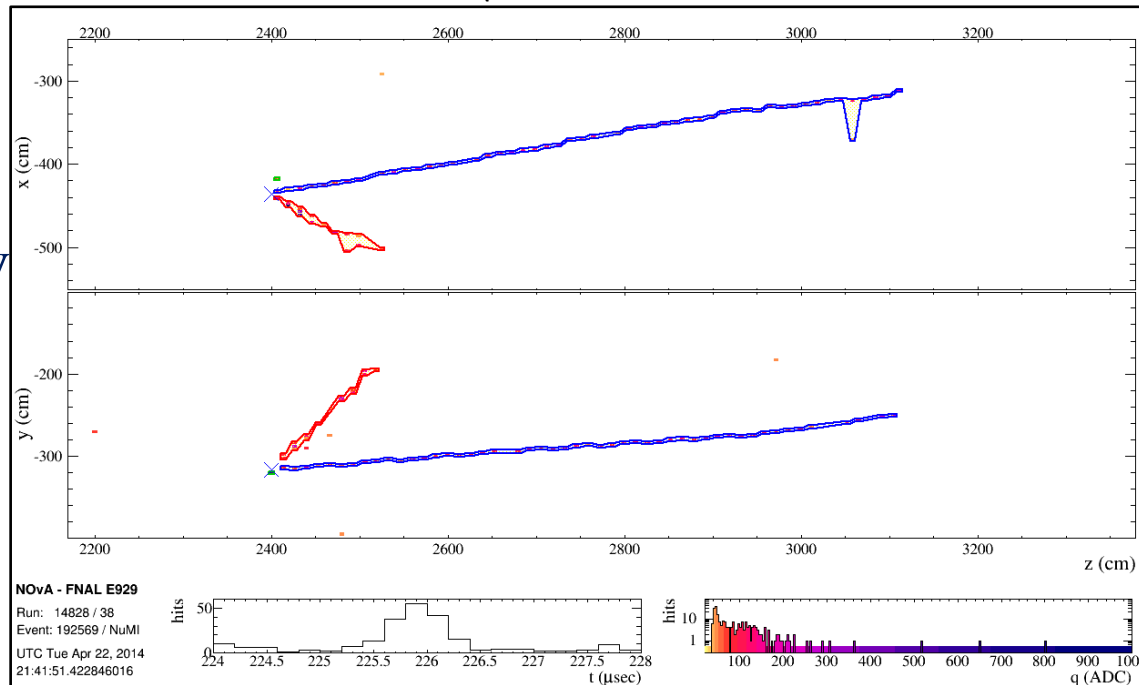


Cosmic ray neutron event in the FD data



Simulated ν_μ CC event

FD ν_μ CC data event



Tracking

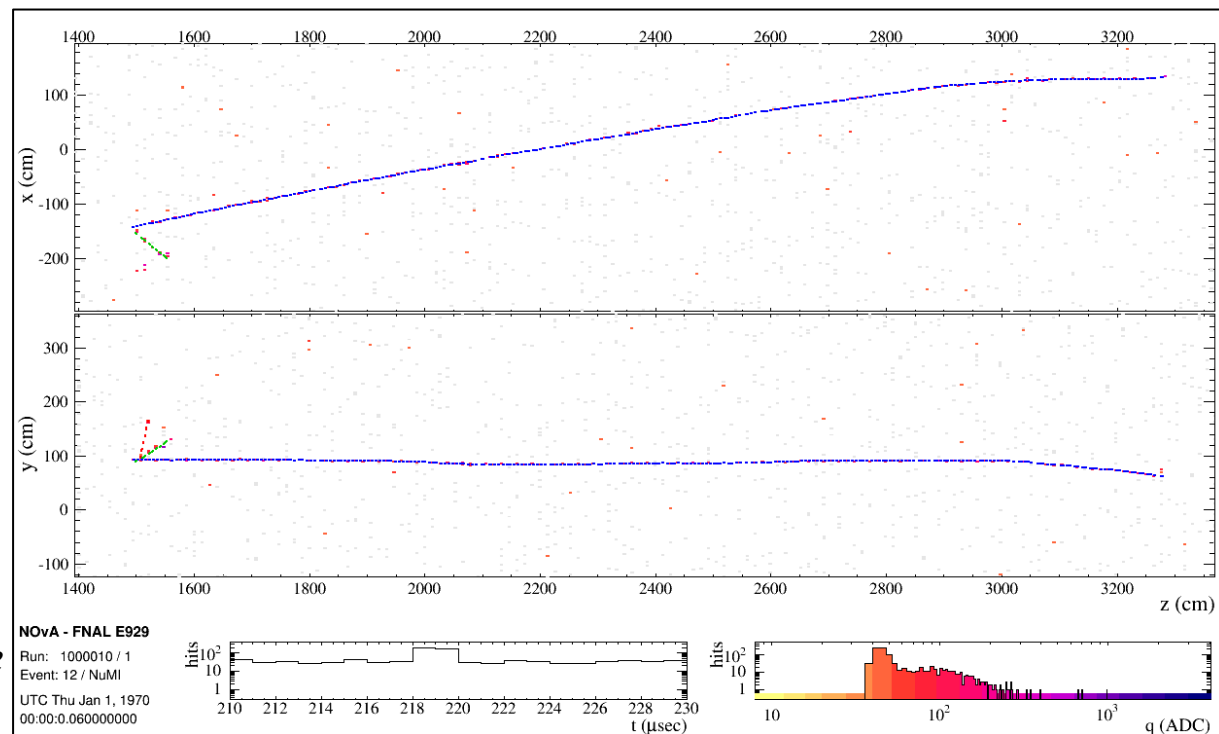
- **Multiple trackers** developed and in use. *Two primary trackers:*
 - ***Cosmic ray tracker:*** lightweight, very fast, good for large calibration samples and online tools
 - ***Kalman filter tracker:*** more detailed, traces scattering for accurate energy, direction measurement.

(simulated event)

Resulting **energy resolution**
for ν_μ CC events (driven in
large part by the track energy)

quasielastic: 4.5%
non-quasielastic: 6.0%

*Reconstructed tracks are
drawn here as colored lines.*

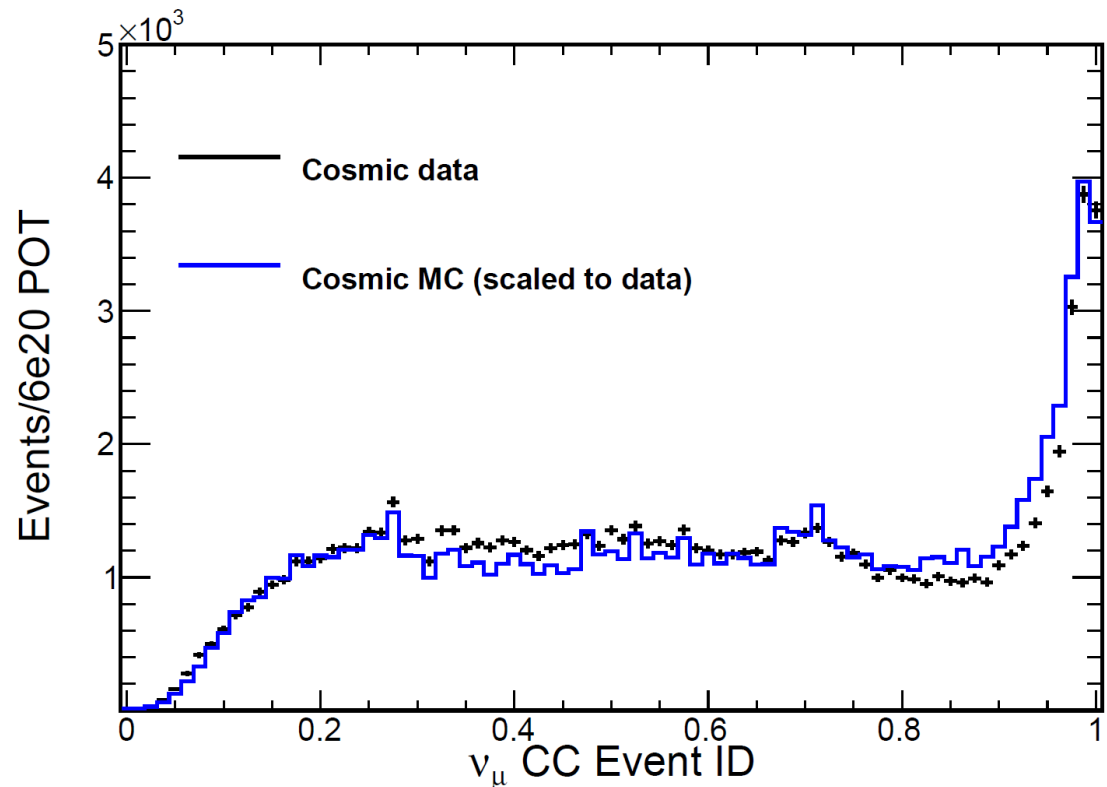


Identifying ν_μ CC events

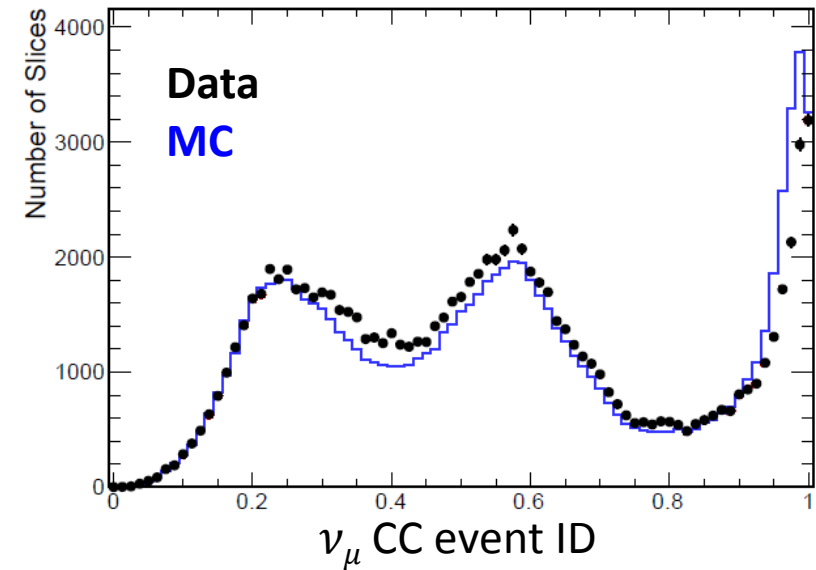
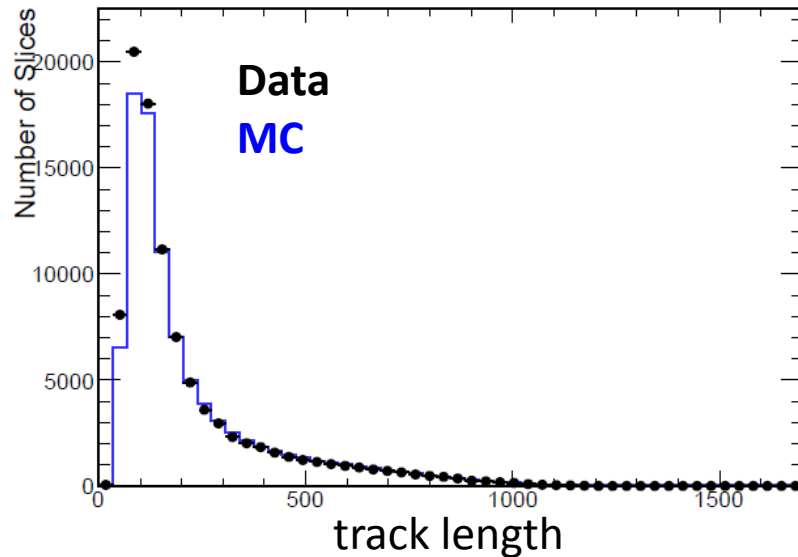
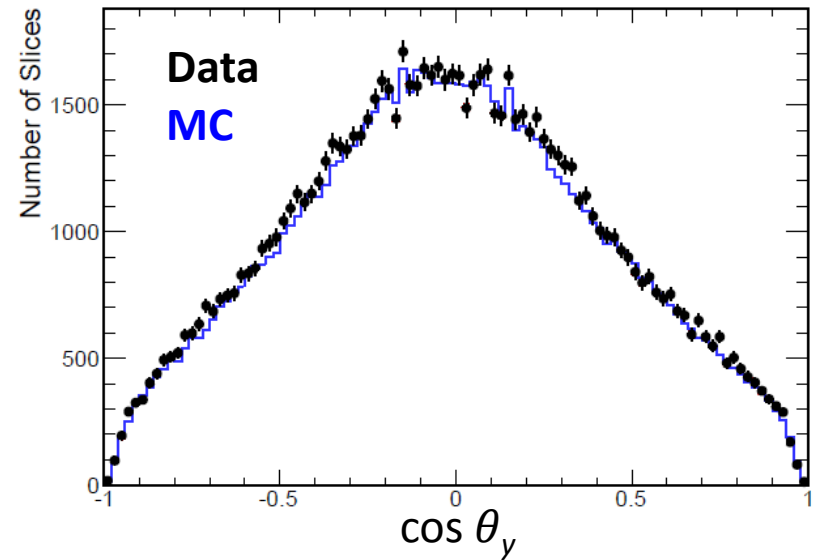
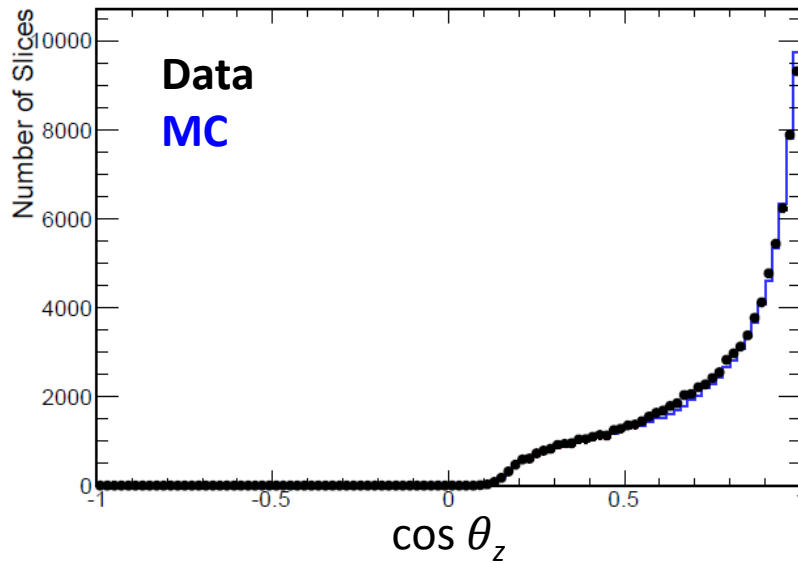
- PID the **muon track candidate** using a **k NN classifier** that considers:
 dE/dx , track scattering, track length, extent of hadronic activity
- Further classify into quasielastic and non-quasielastic samples based on track multiplicity and kinematics
QE and non-QE samples fitted separately in the analysis

ν_μ CC event ID for
cosmic rays,
FD data and MC

*Current MC tuning
already shows
good agreement*



ν_μ CC tracking & event ID distributions, ND beam data and MC
Again, good level of agreement from the current tuning of MC.
(Investigating residual differences is a central part of current activity.)



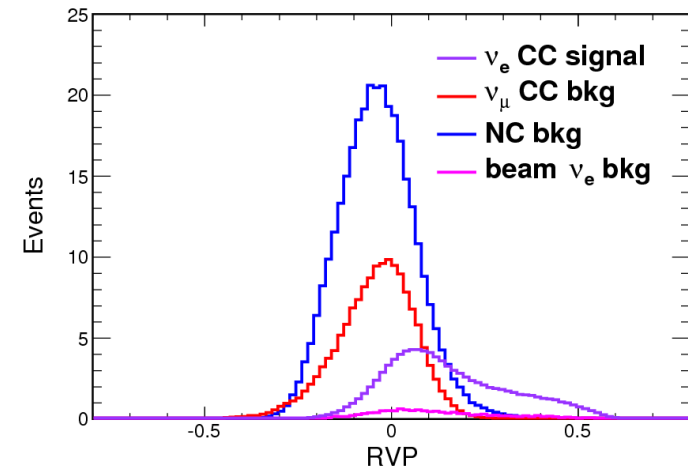
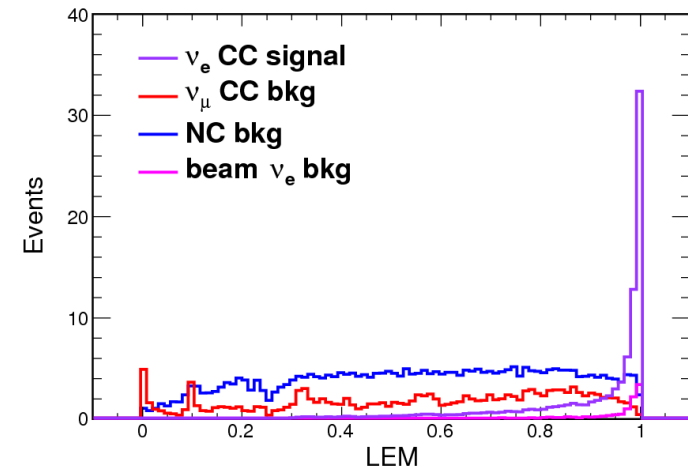
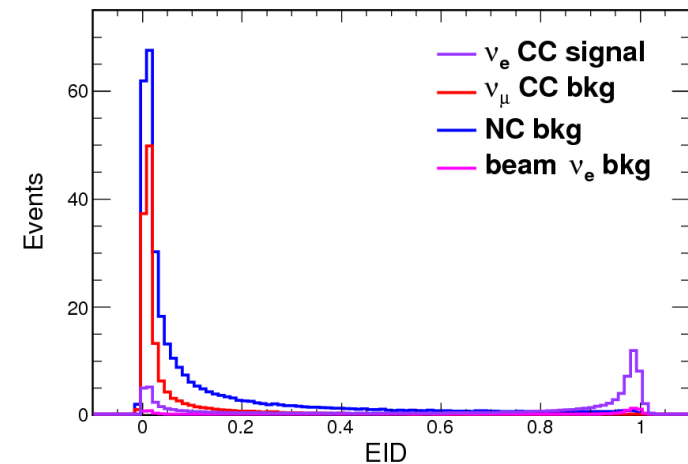
Identifying ν_e CC events

Three very different approaches implemented and supported

EID: Neural network that uses as input EM shower likelihoods, vertex activity, event kinematics, and topology variables

LEM: Library event matching algorithm that compares events to a large library of known events, with the comparisons based directly on the pattern of detector hits

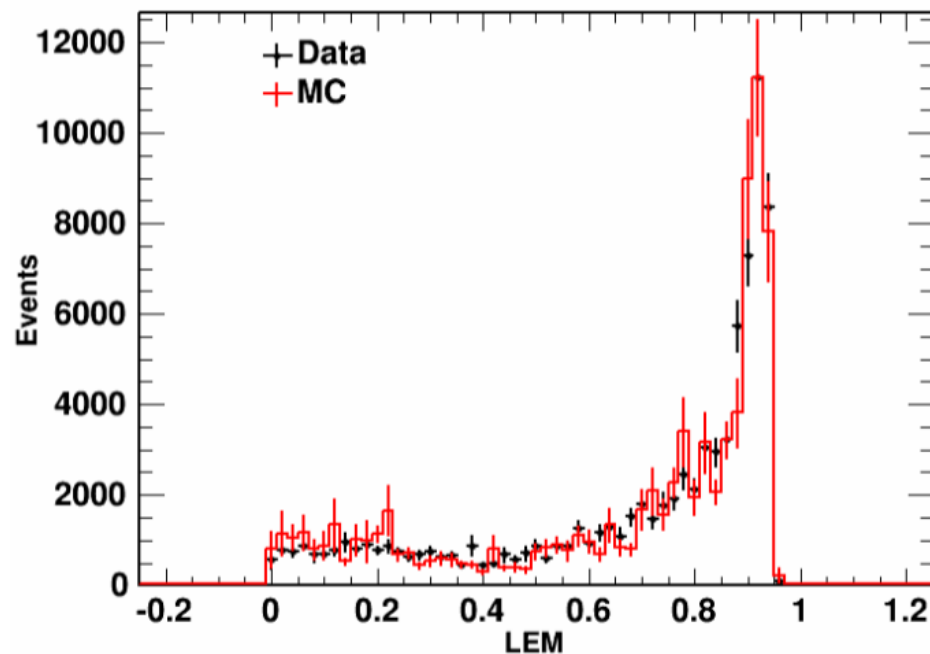
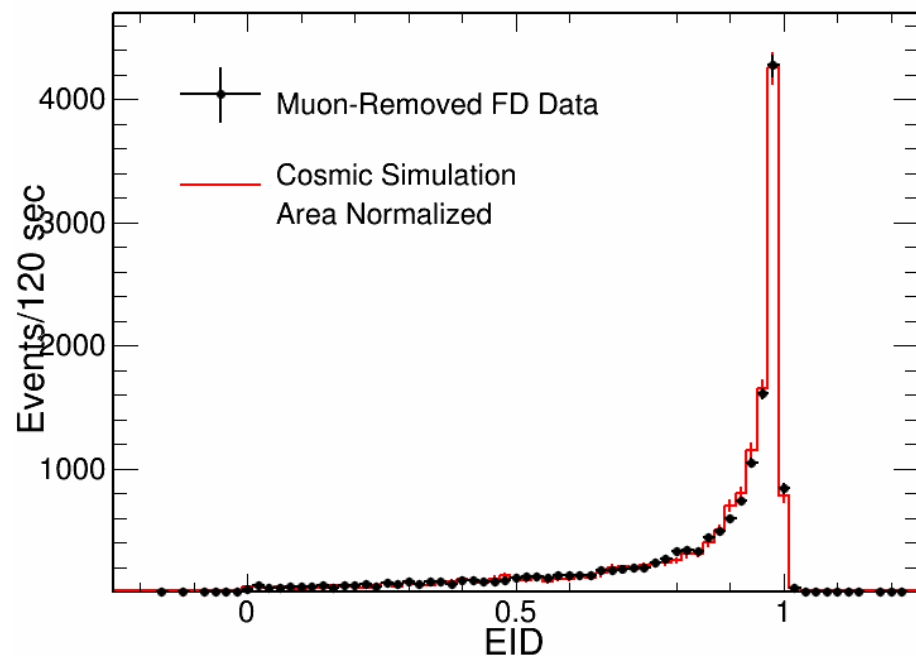
RVP: Boosted decision tree that uses as input a selection of topological and kinematic variables. (Serves as the “simple” ID algorithm.)



EM shower sample

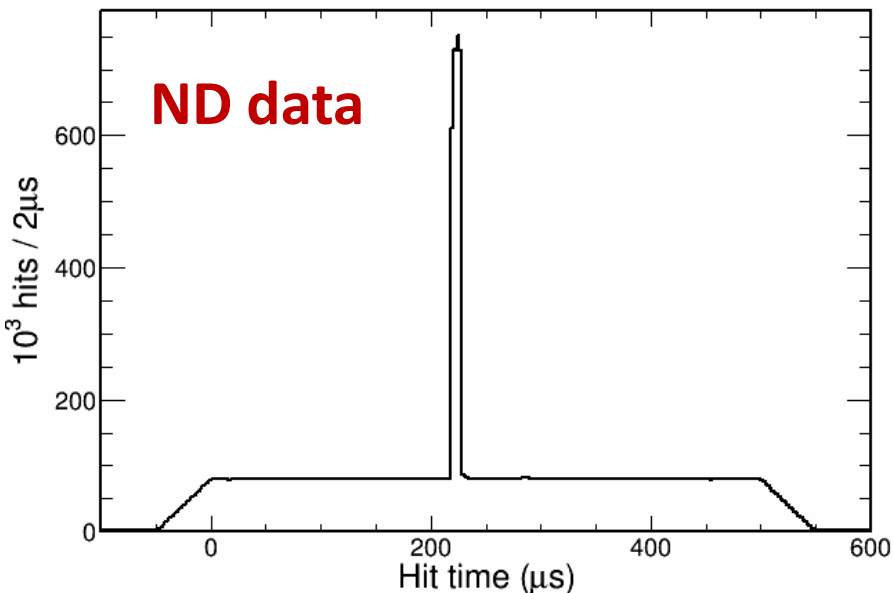
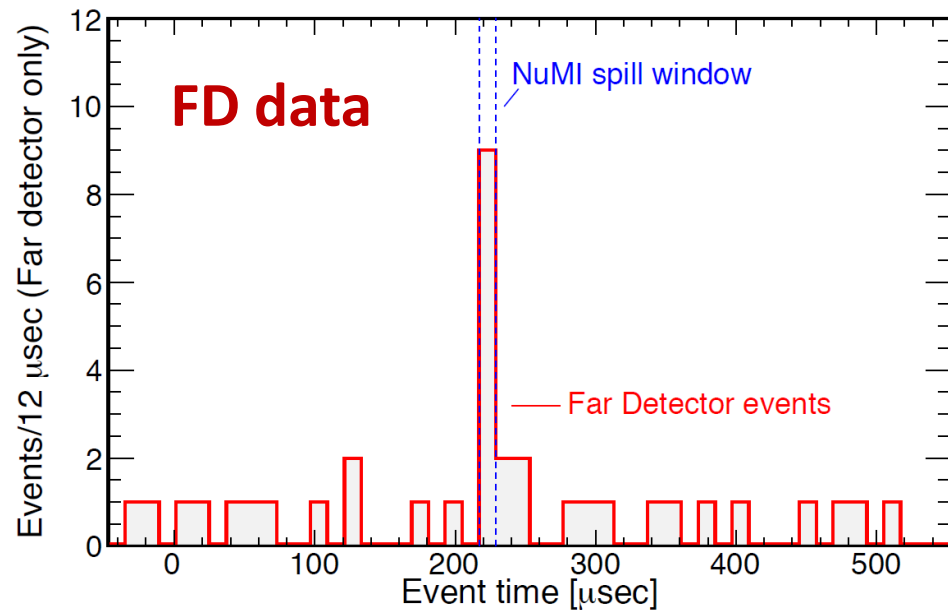
- As a check of **EM shower modeling** and **PID performance**, isolate cosmic ray **bremsstrahlung showers** by removing the parent muon hits
- Data / MC agreement is excellent
- Primary ν_e CC PID distributions shown below

PID distributions from FD data EM showers and equivalent MC

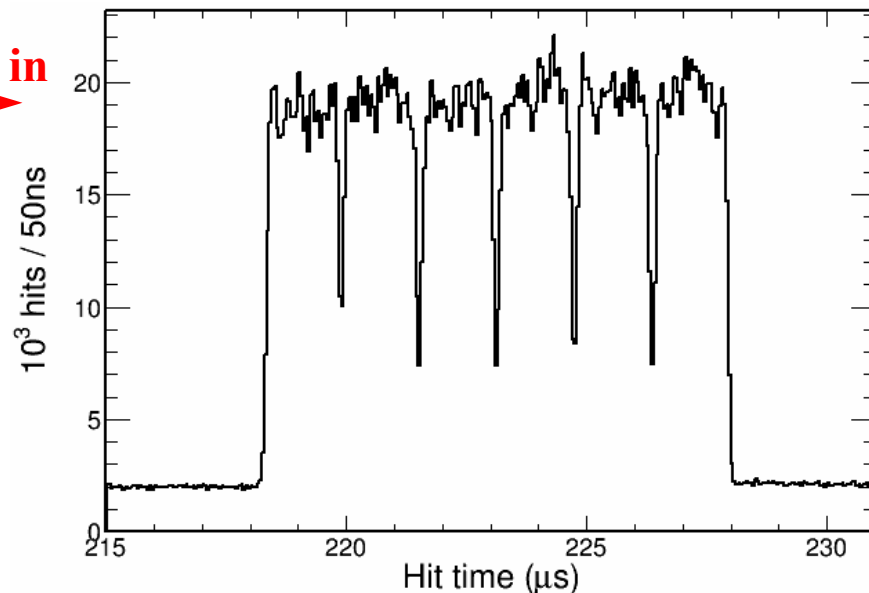


Rejecting cosmic rays in the FD

- **Beam timing** gives us a 10^5 head start ($10\ \mu\text{s}$ spill every 1.3 s)
- *At right: FD neutrino peak*
- *Below: ND neutrino peak*
Spill structure visible in zoomed version.



zoom in
→



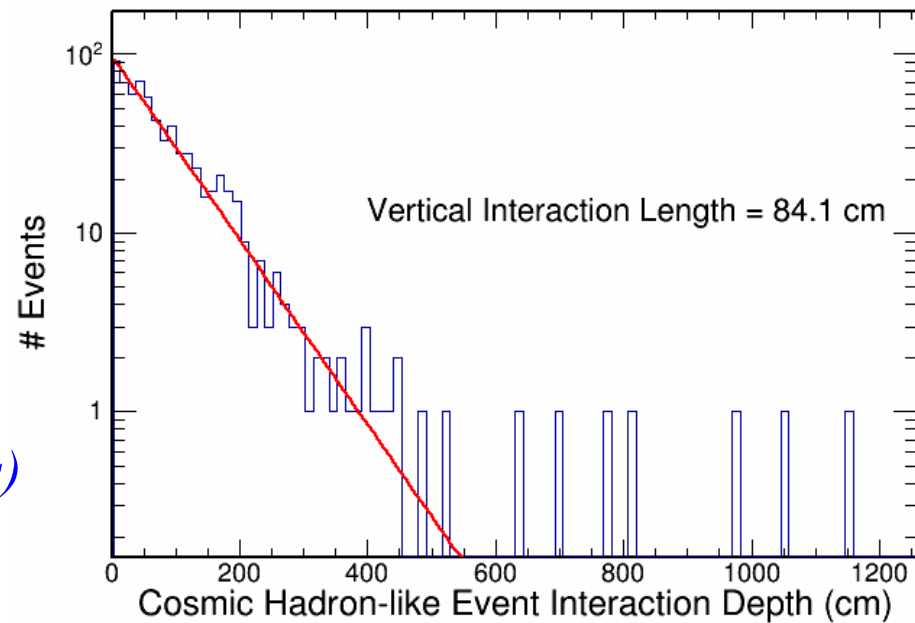
An additional 10^7 rejection from...

ν_e CC case: Cut events...

- ...*too near the detector edges*
- ...*with high p_T/p*
- ...*with poor ν_e CC PID values*

For nominal **1-yr exposure**
(6×10^{20} p.o.t. at 700 kW), **expect:**

14 signal (“avg.” osc. pars)
6 beam bkg.
<1 cosmic bkg.



40M-to-1 cosmic rejection *demonstrated* with FD data

*Above: penetration of neutrons
into the top of the detector
(largely removed by p_T/p cut)*

ν_μ CC case: Cut events...

- ...*whose tracks project too near to the detector edges*
- ...*with cosmic-like muon directions*

For nominal **1-yr exposure**, **expect:**

75 signal (after disappearance)
4 beam bkg.
~1 cosmic bkg.

20M-to-1 cosmic rejection *demonstrated* with FD data

Treatment of detector variations

- **If a portion of the detector is down** for some reason (*in the past*: initial detector construction; *in the future*: hardware swaps, etc.), the **rest of the detector continues to operate** thanks to DAQ “partitioning”.
 - *The analysis code responds appropriately to this (for example, geometry functions report dynamically where the effective edges of the detector are.)*
- A **table of misbehaving channels** is updated on a subrun basis
 - *Event reconstruction code references this table*
- Our large **MC samples** use a **run-by-run matched suite** of bad channel tables, detector configurations (*i.e.*, live portions), and exposures.

Putting all the pieces together

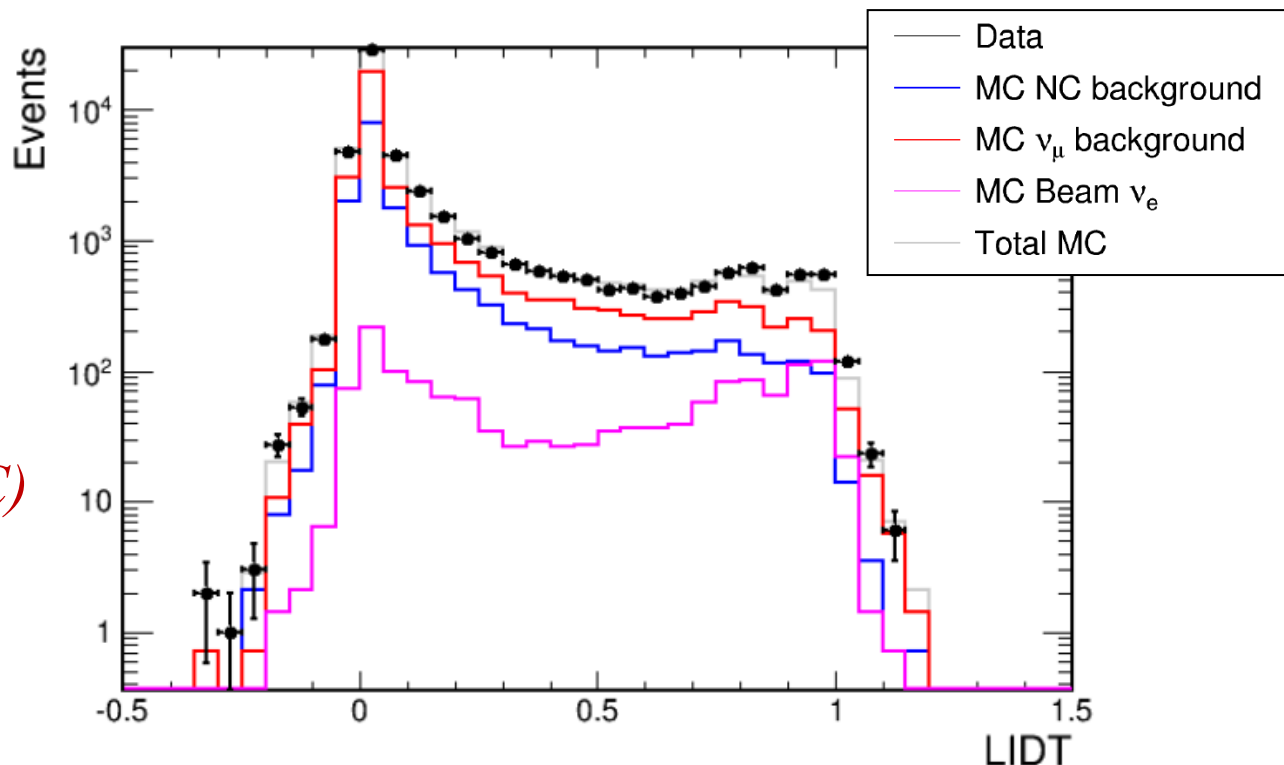
- File production concludes with **Common Analysis Files (CAFs)**
→ *Light-weight ROOT files for end-game analysis work.*
- **CAFAna**: a suite of utilities both for **executing common tasks** with CAFs and for **constructing full oscillation analyses**.
→ *In use throughout collaboration*
- **Uniform treatment** of systematic uncertainties, cuts, protons-on-target accounting, spectrum comparisons, oscillation weighting, ...
- The CAF/CAFAna infrastructure is well-documented, well-supported
→ *Including a live tutorial session that was recorded and posted (audio and screen capture), plus demo exercises to follow along. Great for getting new folks up to speed quickly.*

Brief examples of the end-game tools

- Apply selection to the ND to determine: $\Sigma_{\text{flavors}} (\text{flux} \times \sigma \times \text{eff.})$

ν_e CC event selection
applied to
ND data and MC

(Great agreement in MC)



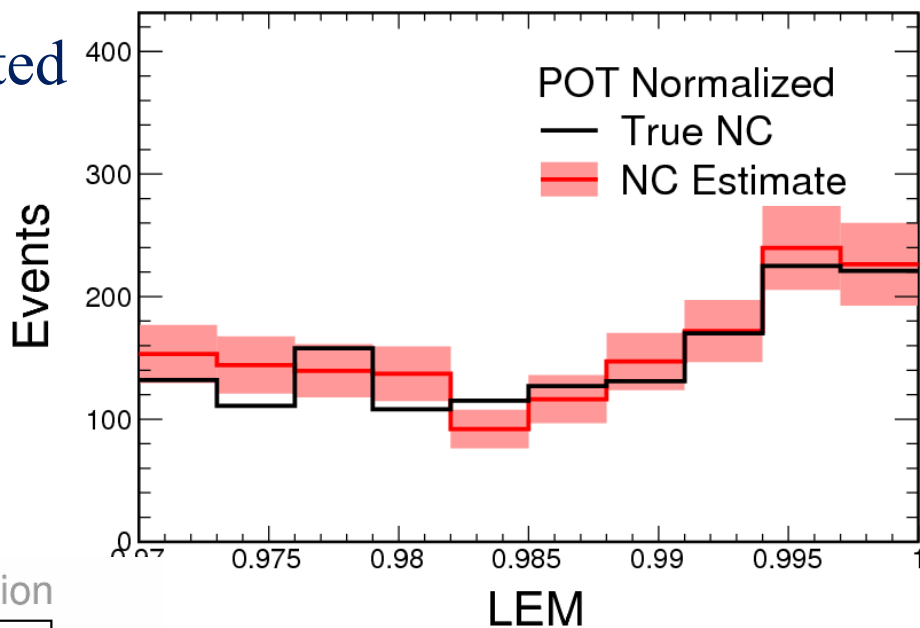
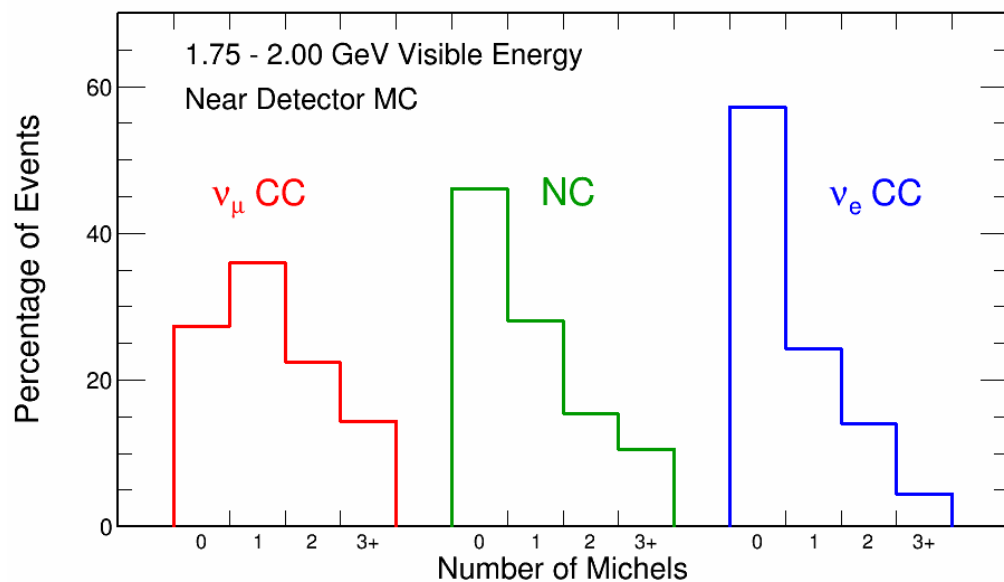
- Use “**decomposition**” techniques to break apart the summed sample
→ *Need to flavor-separate sample since oscillations affect each piece differently when converting from ND to FD prediction.*

ND decomposition in the ν_e analysis

- Two primary methods implemented in the CAFAna framework

Muon-removal: Remove μ 's from ν_μ CC events to form an NC-like events. From these, infer NC mis-ID contribution in selected sample. \longrightarrow

NOvA Simulation



Michel electrons: Statistically decompose the ND sample based on differences in Michel electron rates among ν_μ CC, ν_e CC, and NC events. \longleftarrow

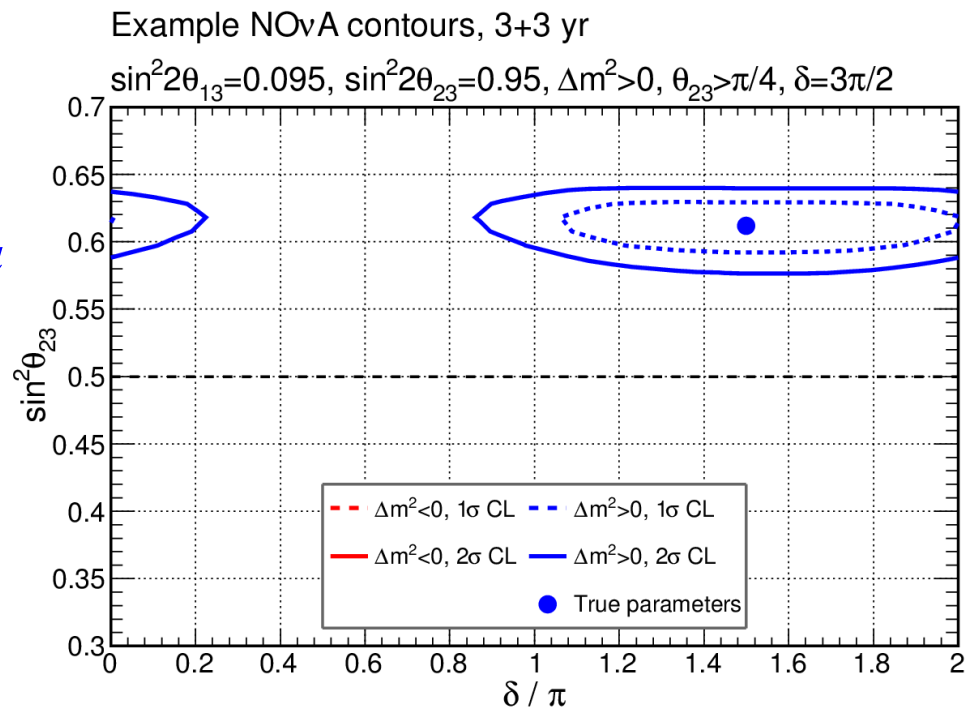
Final fits

- A flexible **ND-to-FD extrapolation** toolkit is built in CAFAna
- The resulting FD predictions feed an **oscillation fitting framework**
- *A few noteworthy features:*
 - Seamlessly **float, freeze, or marginalize any oscillation parameters**, with or without penalty terms
 - Seamlessly **combine NOvA analyses and/or external results**. Important for folding in known constraints (e.g., reactor, T2K)
 - Statistically **rigorous contours** via Feldman-Cousins unified approach
 - Systematics via **nuisance parameters**

At right: example contours from CAFAna for one point in parameter space

Simultaneously break ν_3 **flavor degeneracy** (θ_{23} octant), determine **mass hierarchy**, and constrain **CP phase δ** .

(Full TDR exposure of 36×10^{20} p.o.t.)



Summary

- **End-to-end analysis tools in place**
- Calibration, reconstruction, particle ID, fitting tools are **operational on ND and FD data**
- Surface operation: cosmic ray rejection **demonstrated with FD data**
- *Current analysis frontiers include:*
 - Understanding residual data/MC differences
Several MC improvements already developed
 - Assessing systematic uncertainties
Should not a major factor until we gain significant exposure
 - **Most critical: integrating protons-on-target !!**

NOvA is ready to produce the next round of significant results in neutrino physics.

The world is watching, waiting.

Extras

NOvA

A broad physics scope

Using $\nu_\mu \rightarrow \nu_e$, $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$...

- Determine the ν mass hierarchy
- Determine the θ_{23} octant
- Constrain δ_{CP}

Using $\nu_\mu \rightarrow \nu_\mu$, $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$...

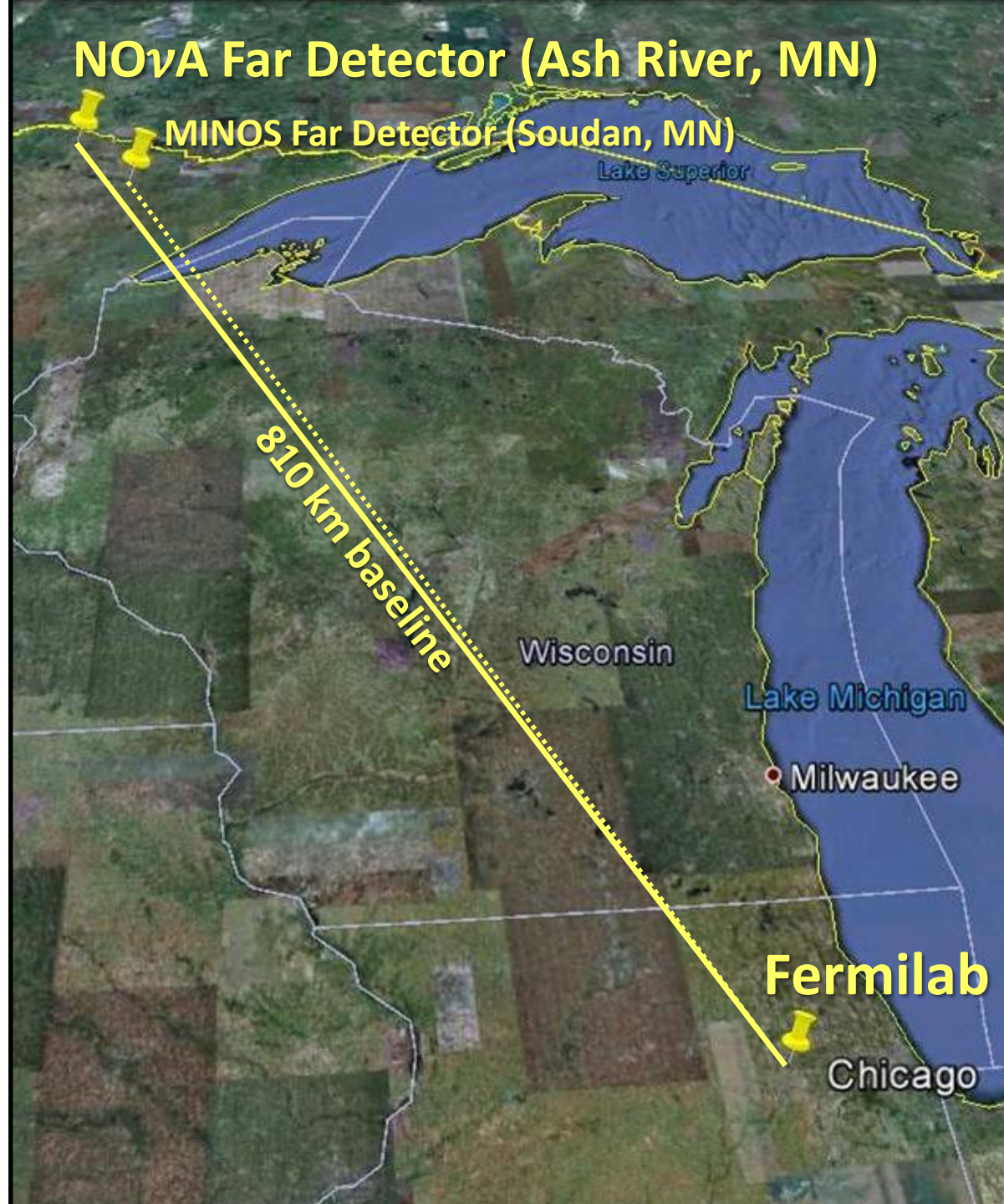
- **Atmospheric parameters:**
precision measurements of θ_{23} ,
 Δm_{32}^2 . (*Exclude $\theta_{23}=\pi/4$?*)
- **Over-constrain** the atmos. sector
(*four oscillation channels*)

Also ...

- Neutrino cross sections at the NOvA Near Detector
- Sterile neutrinos
- Supernova neutrinos
- Other exotica

NOvA Far Detector (Ash River, MN)

MINOS Far Detector (Soudan, MN)



Wisconsin

Lake Michigan

Milwaukee

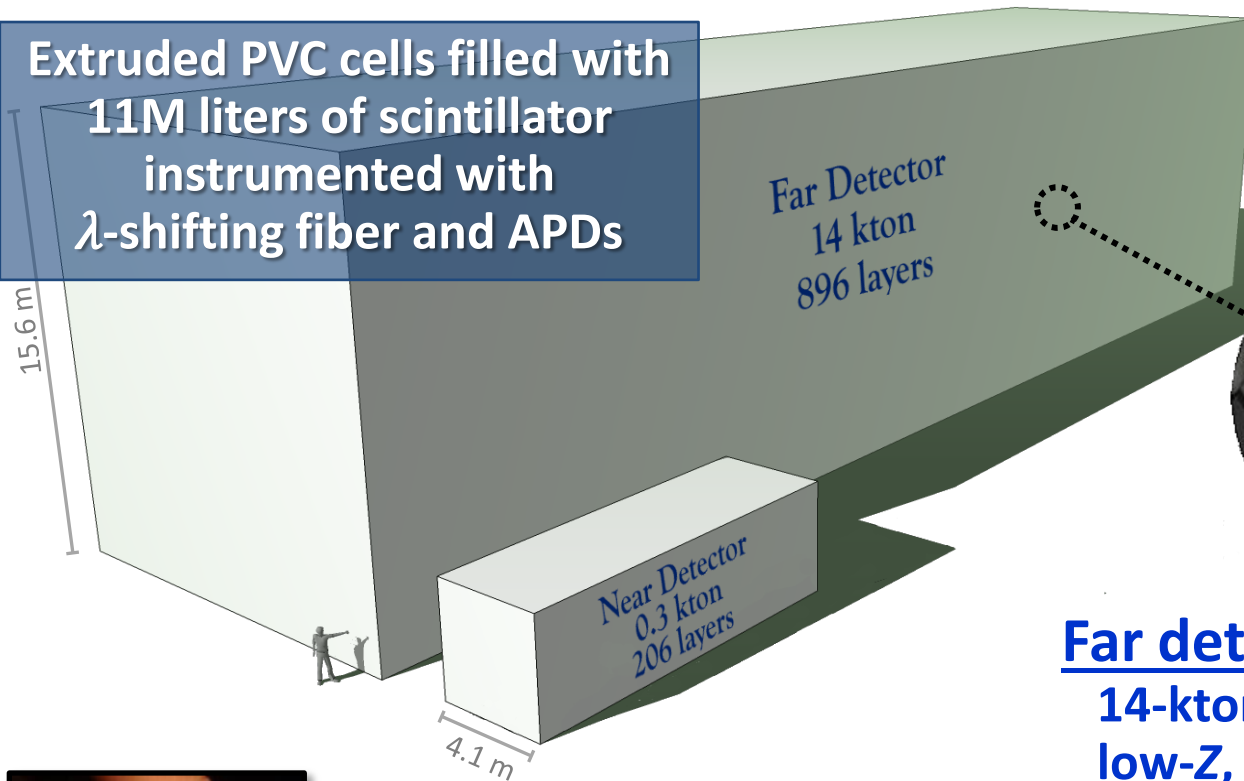
Fermilab

Chicago

810 km baseline

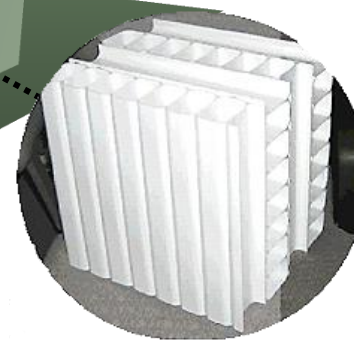
NOvA detectors

Extruded PVC cells filled with
11M liters of scintillator
instrumented with
 λ -shifting fiber and APDs



A NOvA cell

To APD



1560 cm

4 cm × 6 cm

Far detector:

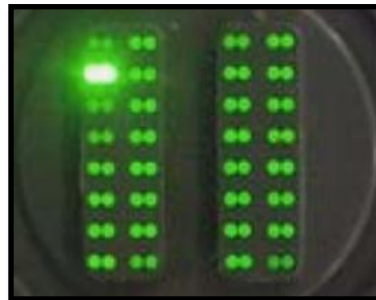
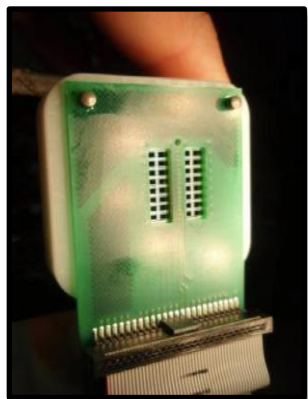
14-kton, fine-grained,
low-Z, highly-active
tracking calorimeter
→ 344,000 channels

Near detector:

0.3-kton version of
the same
→ 18,000 channels

32-pixel APD

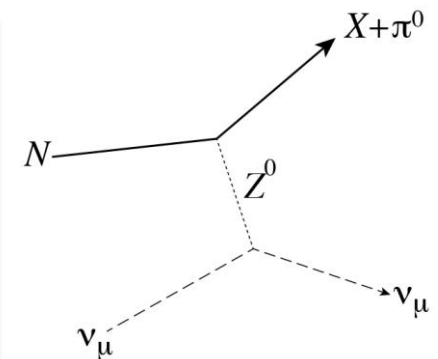
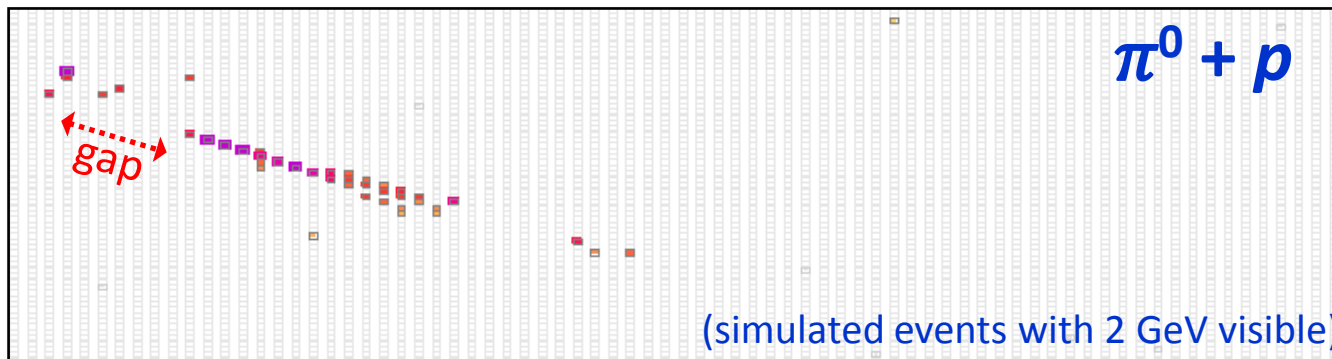
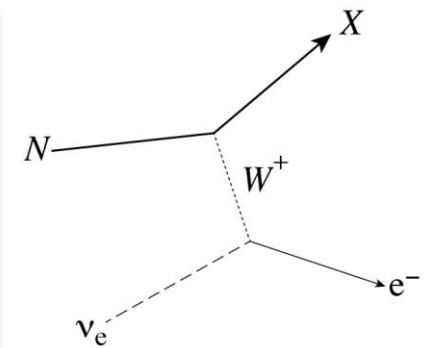
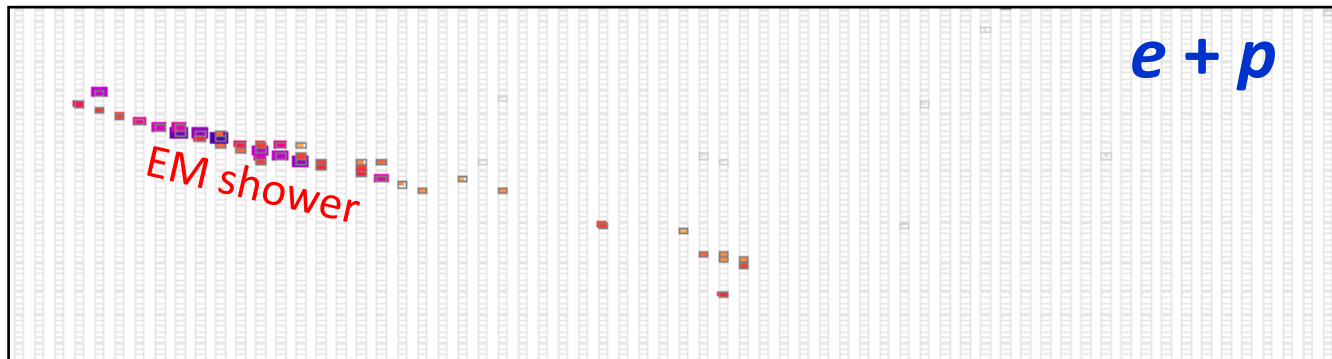
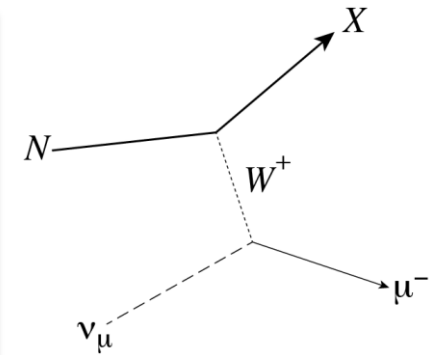
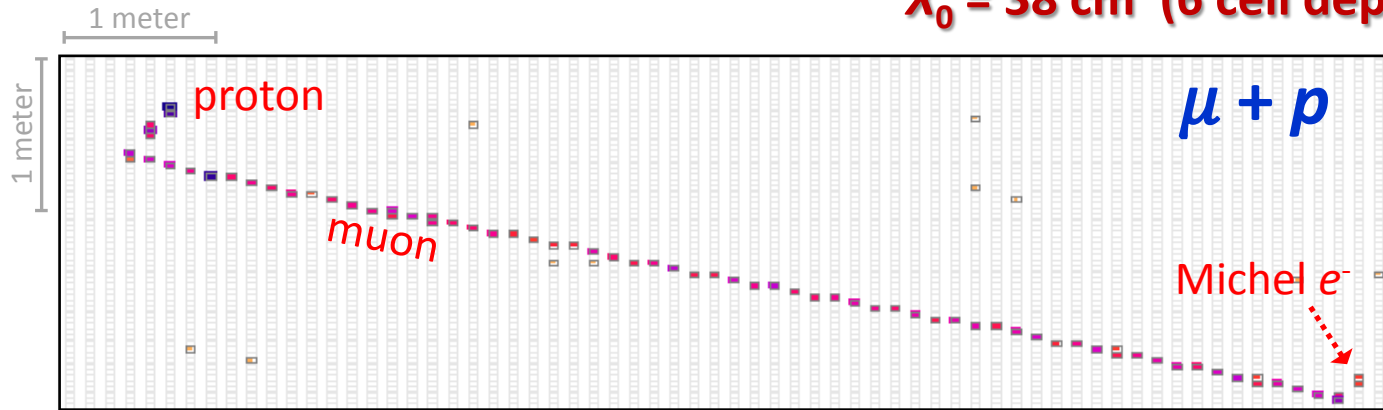
Fiber pairs
from 32 cells

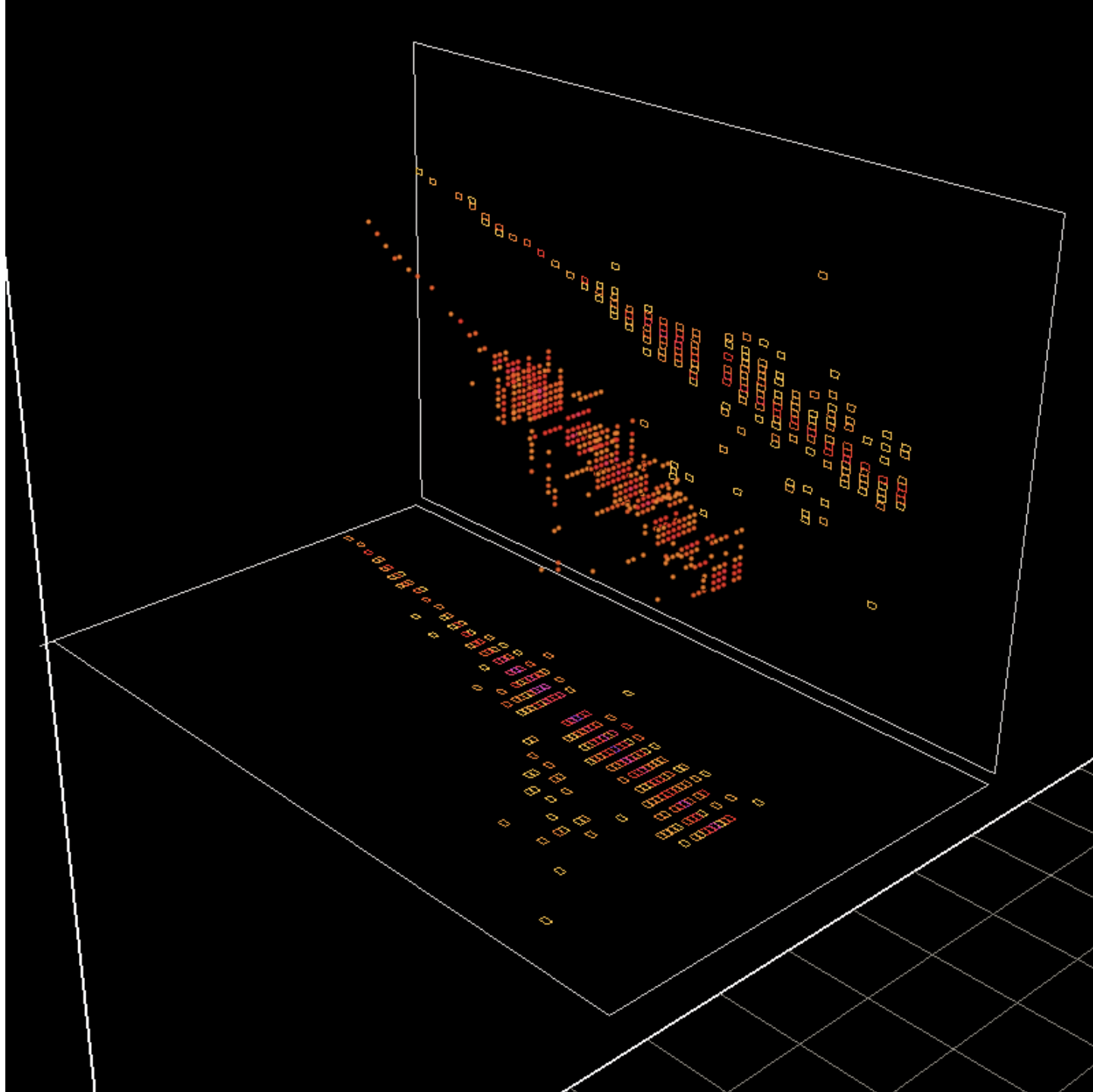


Events in NOvA

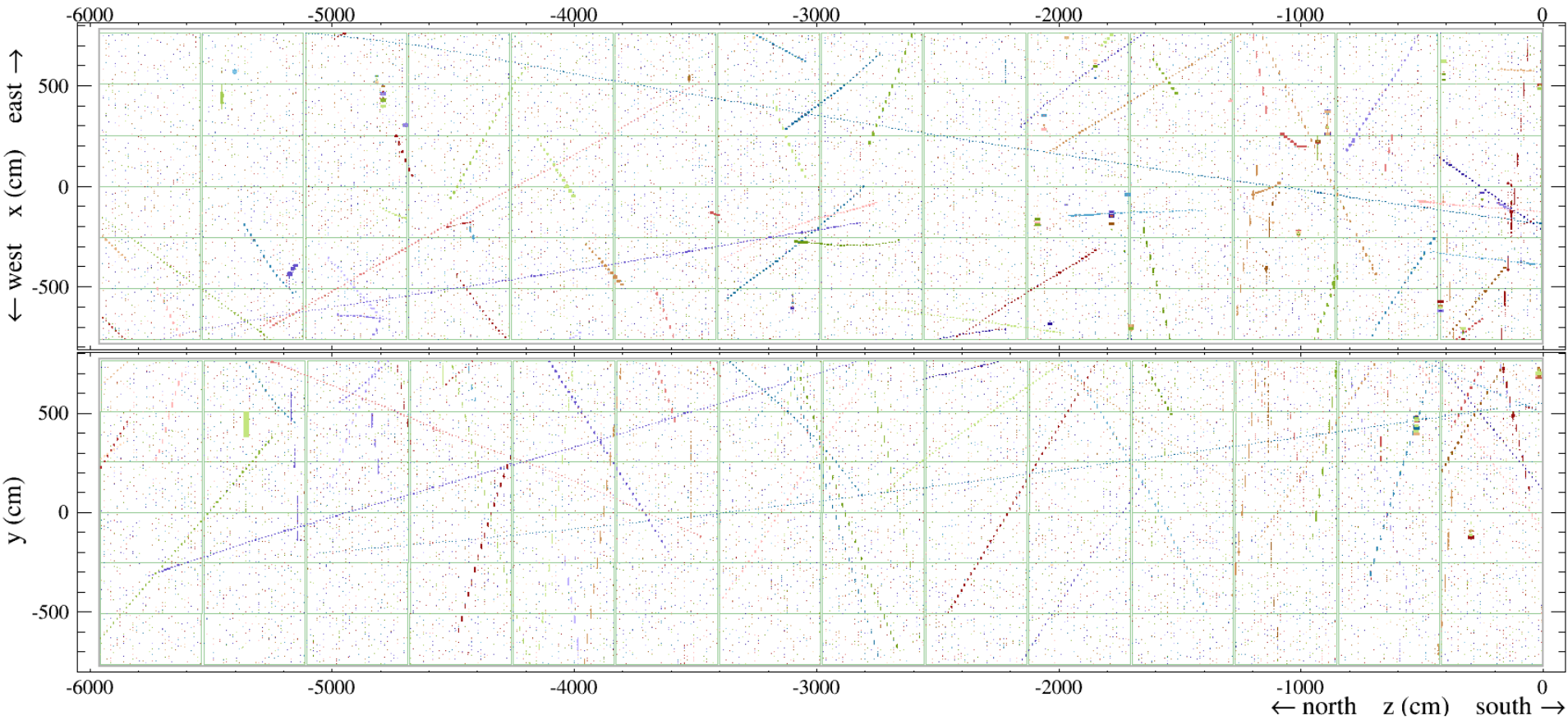
Superb spatial granularity for a detector of this scale

$X_0 = 38 \text{ cm}$ (6 cell depths, 10 cell widths)



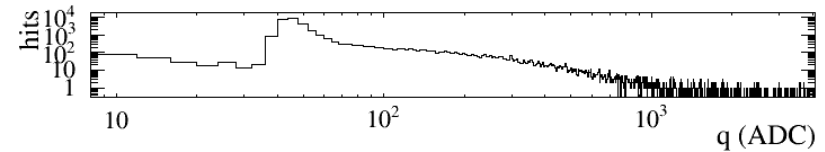
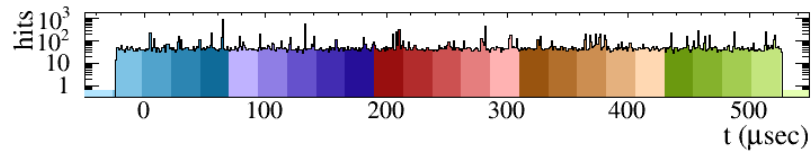


Raw data from 500 μs of Far Detector activity

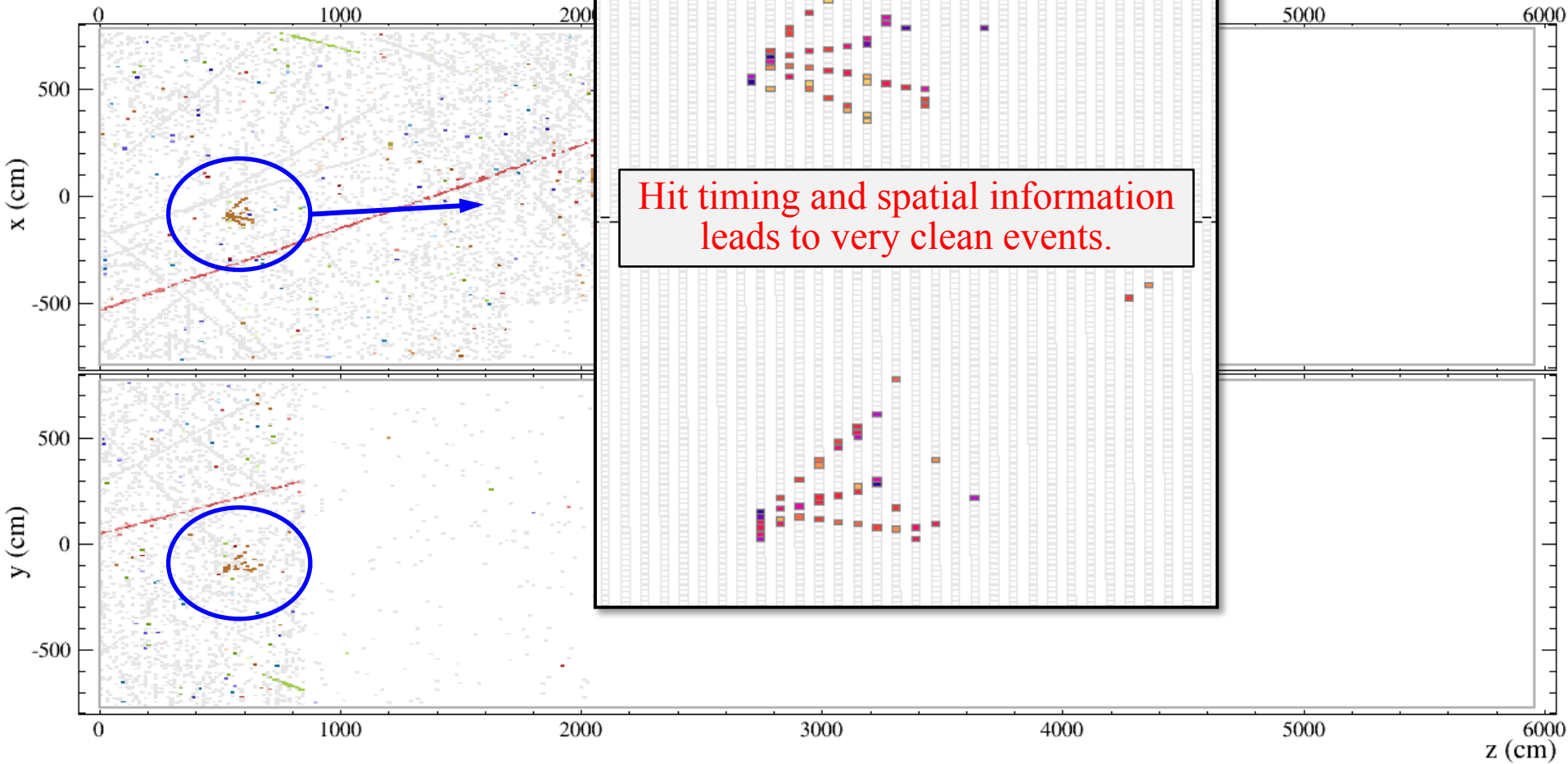


NOvA - FNAL E929

Run: 16674 / 15
Event: 68101 / NuMI
UTC Fri Aug 8, 2014
19:12:22.145823424



Neutrino interaction in (partial) Far Detector



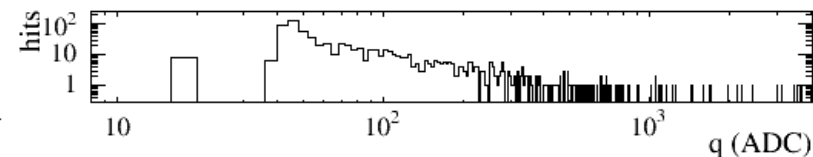
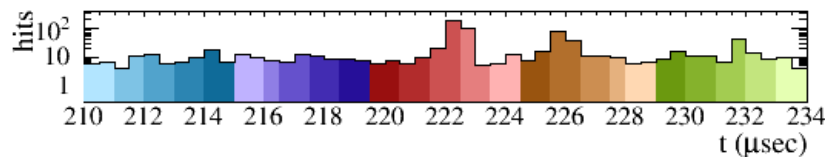
NOvA - FNAL E929

Run: 11988 / 48

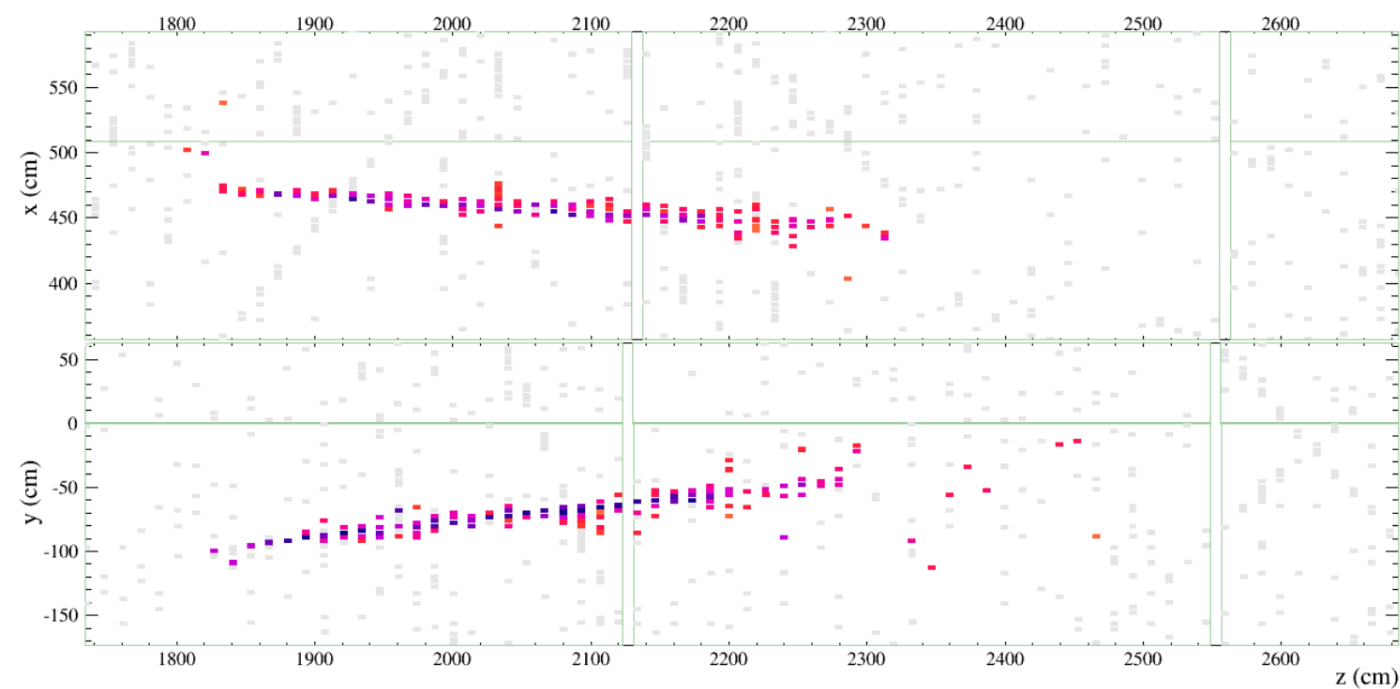
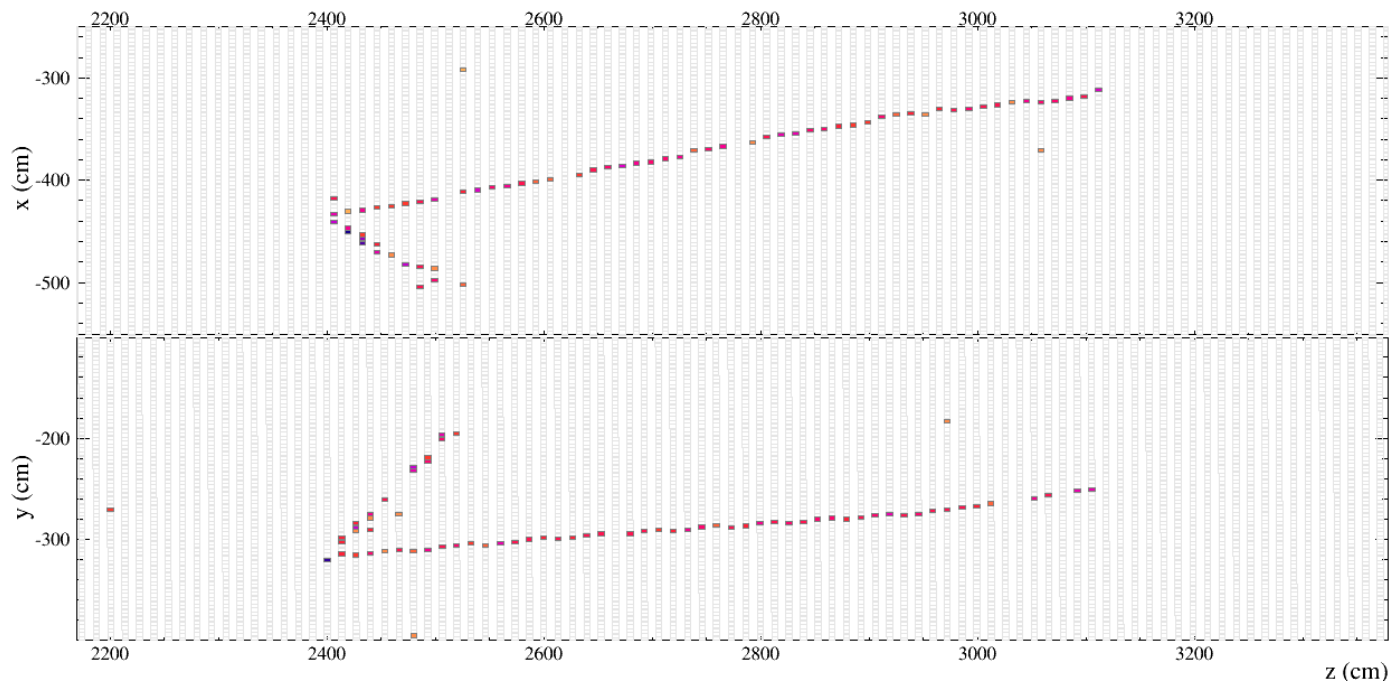
Event: 187563 / NuMI

UTC Sat Dec 14, 2013

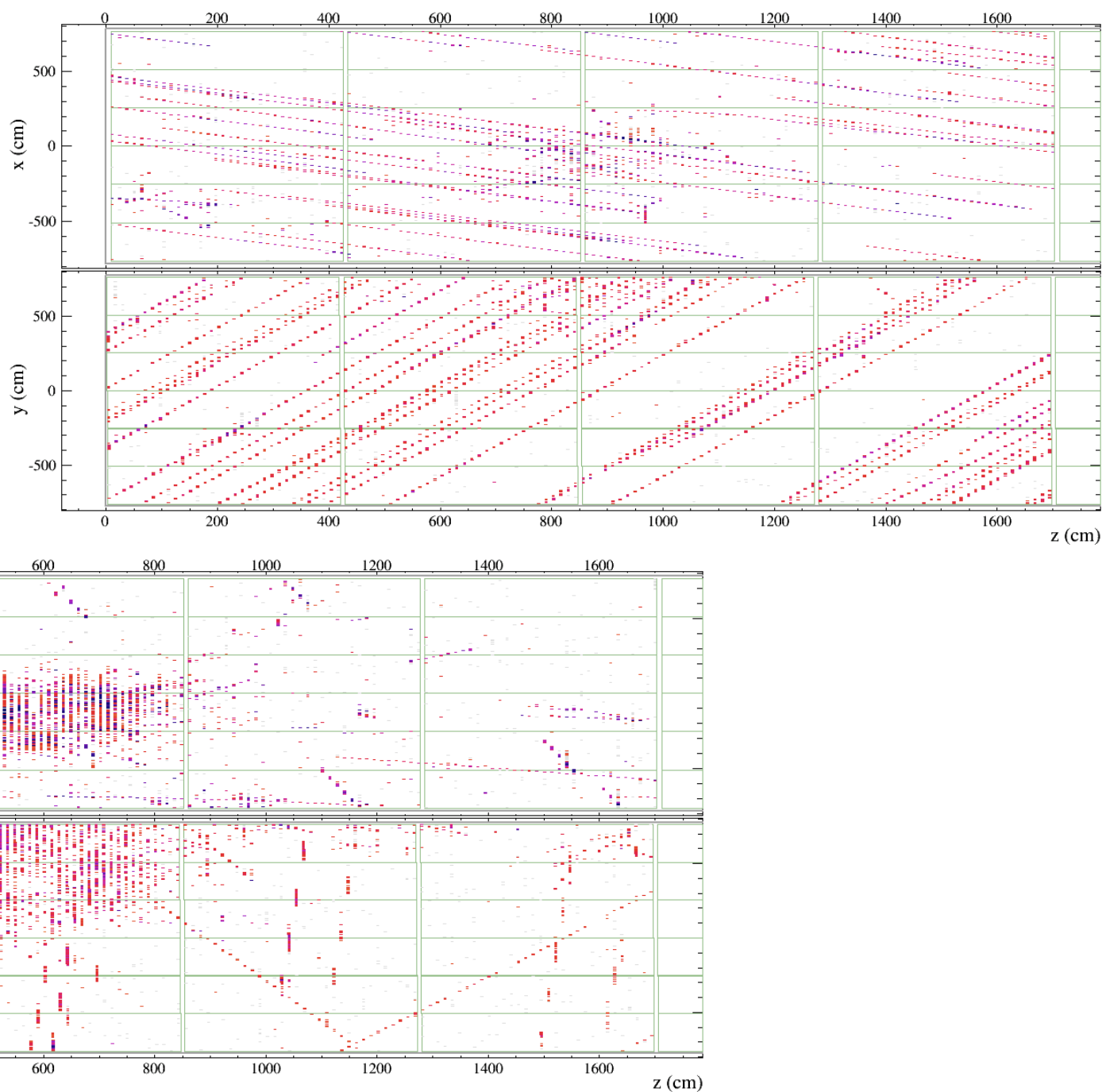
09:12:49.228821216



Charged-current candidates in the Far Detector



Air shower and high-energy cosmic event



$$\nu_{\mu} \rightarrow \nu_{\mu} \text{ and } \bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu}$$

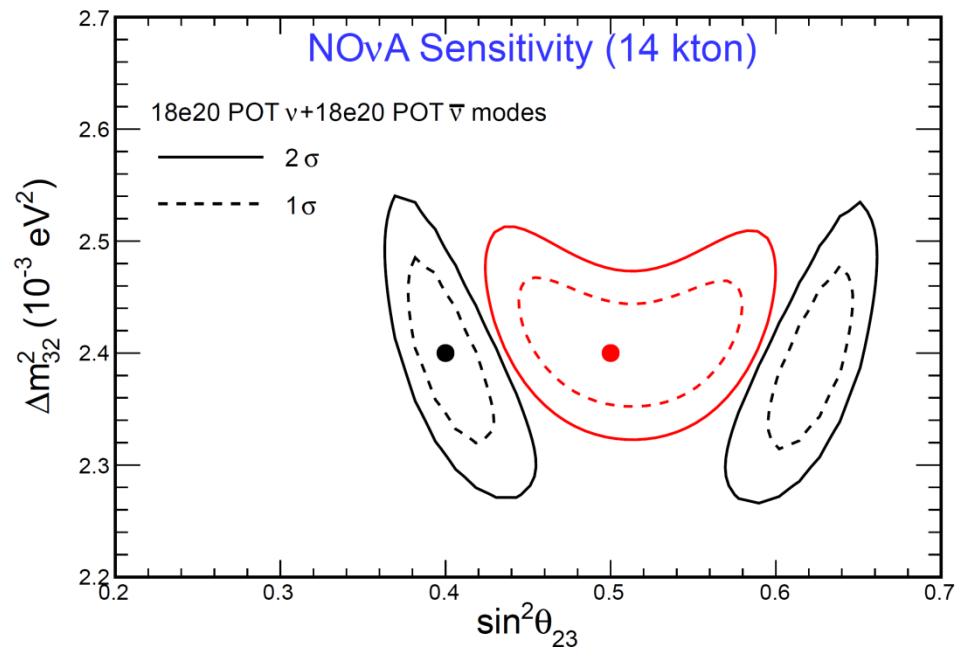
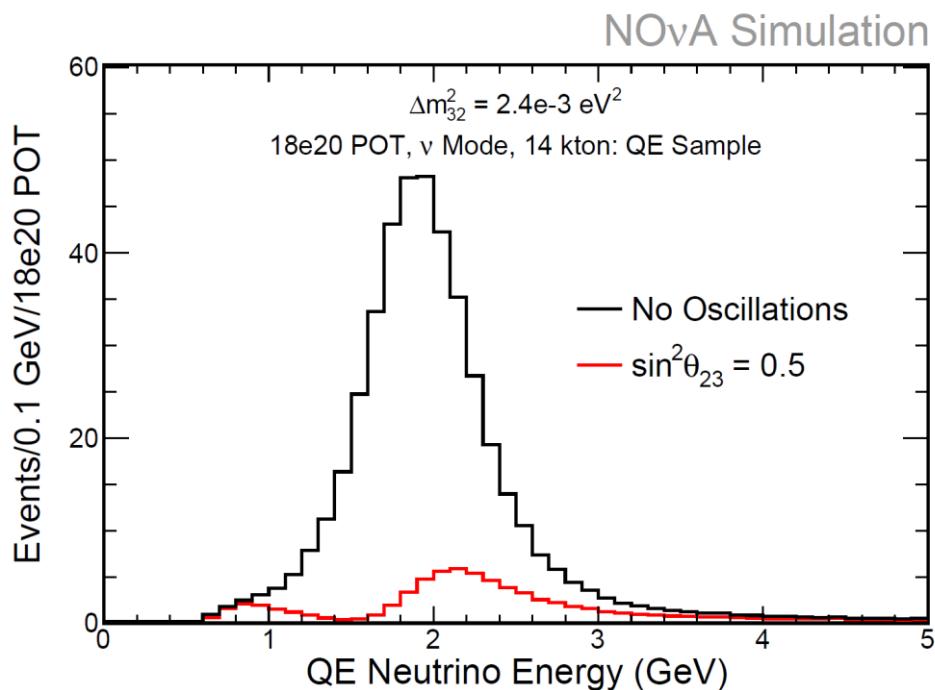
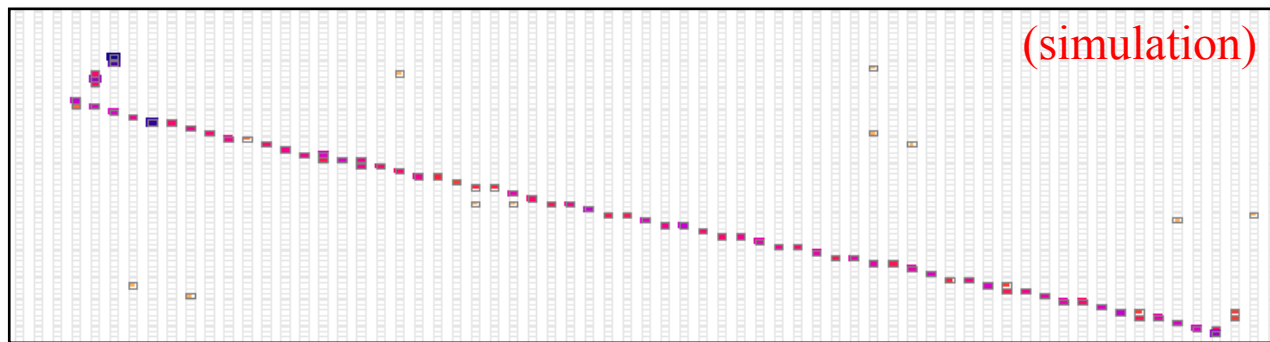
- **Below:** Oscillated and non-oscillated spectra (reconstructed energy) for ν_{μ} CC QE candidates

Non-QE candidates fitted separately.

6% vs. 4.5% energy resolution using current techniques.)

- **Right:** example contours for a nominal 6-year run at two test points

Requirements of cosmic rejection now included (small change)



*Long baseline \rightarrow hierarchy sensitivity,
along with rest of oscillation reach*

Example point in ν parameter space

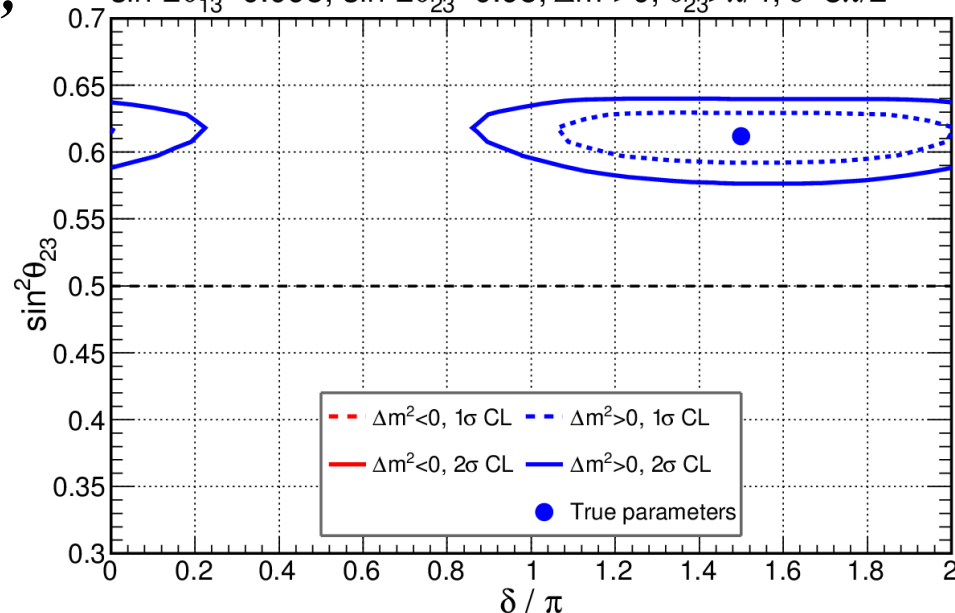
Simultaneously break ν_3 **flavor degeneracy** (θ_{23} octant), determine **mass hierarchy**, and constrain **CP phase δ** .

And a “degenerate” point...

Hierarchy and δ information now correlated. **Octant preference** still established.

Example NOvA contours, 3+3 yr

$\sin^2 2\theta_{13}=0.095$, $\sin^2 2\theta_{23}=0.95$, $\Delta m^2 > 0$, $\theta_{23} > \pi/4$, $\delta = 3\pi/2$



Example NOvA contours, 3+3 yr

$\sin^2 2\theta_{13}=0.095$, $\sin^2 2\theta_{23}=0.95$, $\Delta m^2 > 0$, $\theta_{23} > \pi/4$, $\delta = \pi/2$

