

# The physics goals and reach of ELBNF28

22 January 2015  
ELBNF Proto-Collaboration Meeting  
Fermilab

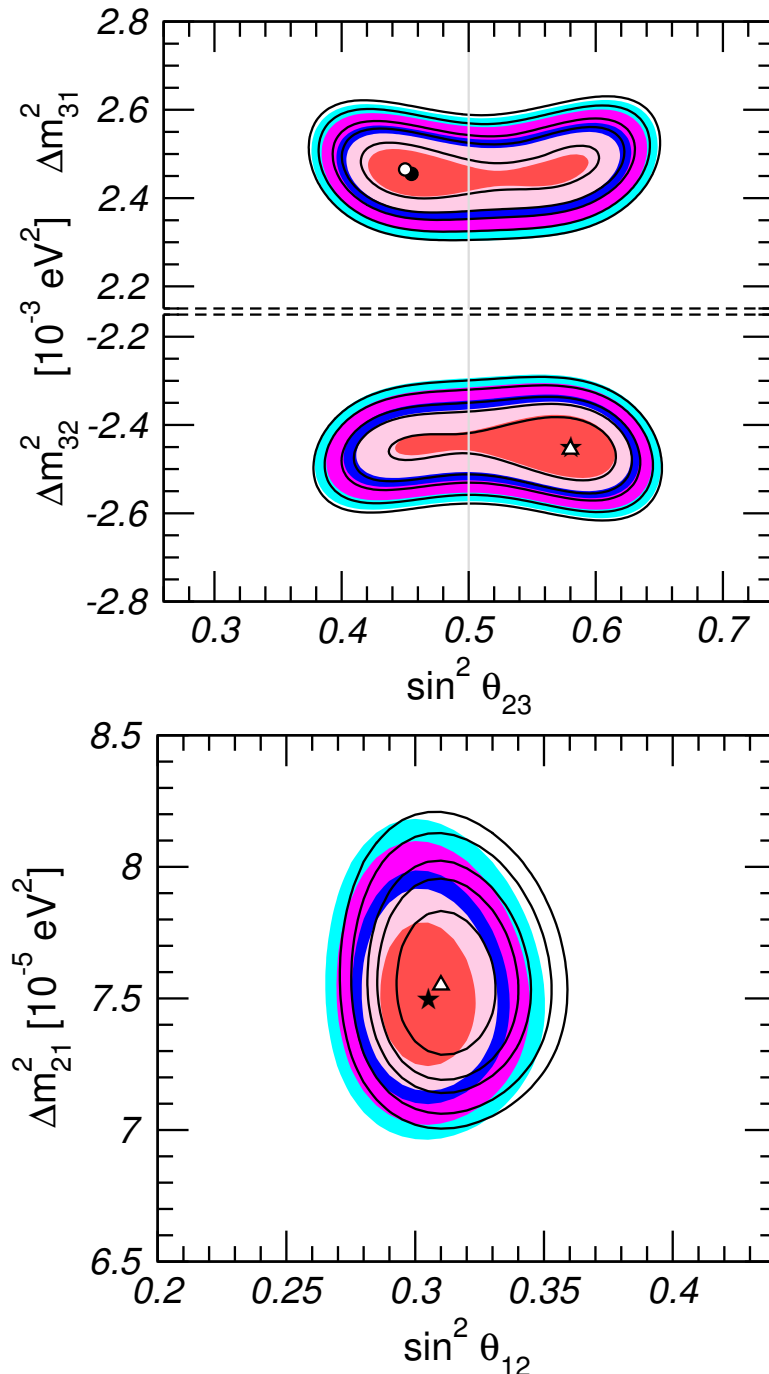
**Silvia Pascoli**  
IPPP – Durham University



# Outline

- 1. Present status of neutrino parameters**
- 2. LBL oscillation experiments physics goals:**
  - Mass ordering**
  - Leptonic CP-violation**
  - Precision measurement of parameters**
  - Testing the 3-neutrino scenario**
- 3. ELBNF28 physics reach**
- 4. Synergies and complementarities with other experiments**
- 4. Conclusions**

# Summary of current neutrino parameters



	Normal Ordering ( $\Delta\chi^2 = 0.97$ )	
	bfp $\pm 1\sigma$	$3\sigma$ range
$\sin^2 \theta_{12}$	$0.304^{+0.013}_{-0.012}$	$0.270 \rightarrow 0.344$
$\theta_{12}/^\circ$	$33.48^{+0.78}_{-0.75}$	$31.29 \rightarrow 35.91$
$\sin^2 \theta_{23}$	$0.452^{+0.052}_{-0.028}$	$0.382 \rightarrow 0.643$
$\theta_{23}/^\circ$	$42.3^{+3.0}_{-1.6}$	$38.2 \rightarrow 53.3$
$\sin^2 \theta_{13}$	$0.0218^{+0.0010}_{-0.0010}$	$0.0186 \rightarrow 0.0250$
$\theta_{13}/^\circ$	$8.50^{+0.20}_{-0.21}$	$7.85 \rightarrow 9.10$
$\delta_{\text{CP}}/^\circ$	$306^{+39}_{-70}$	$0 \rightarrow 360$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.50^{+0.19}_{-0.17}$	$7.02 \rightarrow 8.09$
$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.457^{+0.047}_{-0.047}$	$+2.317 \rightarrow +2.607$

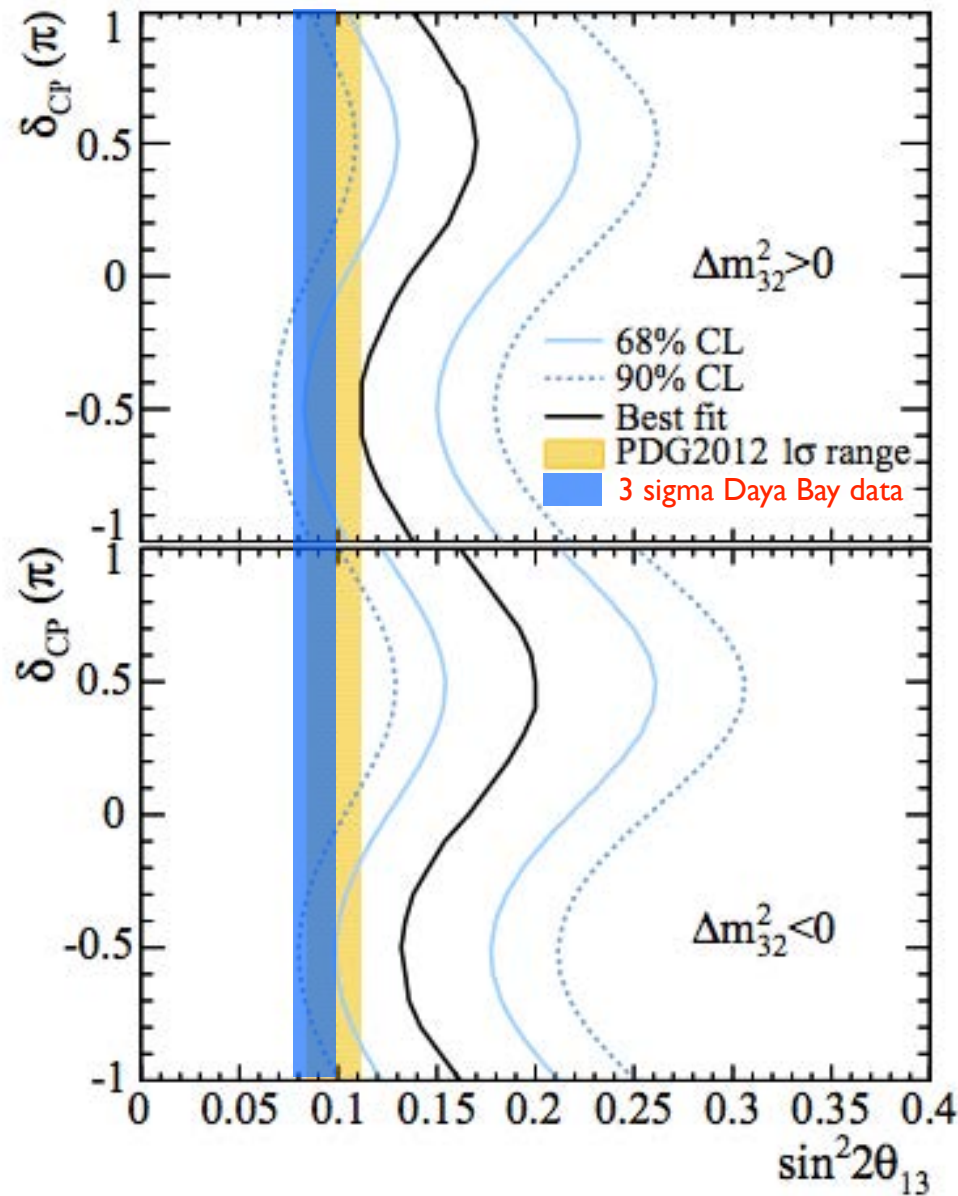
M. C. Gonzalez-  
Garcia et al.,  
NuFit,  
1409.5439



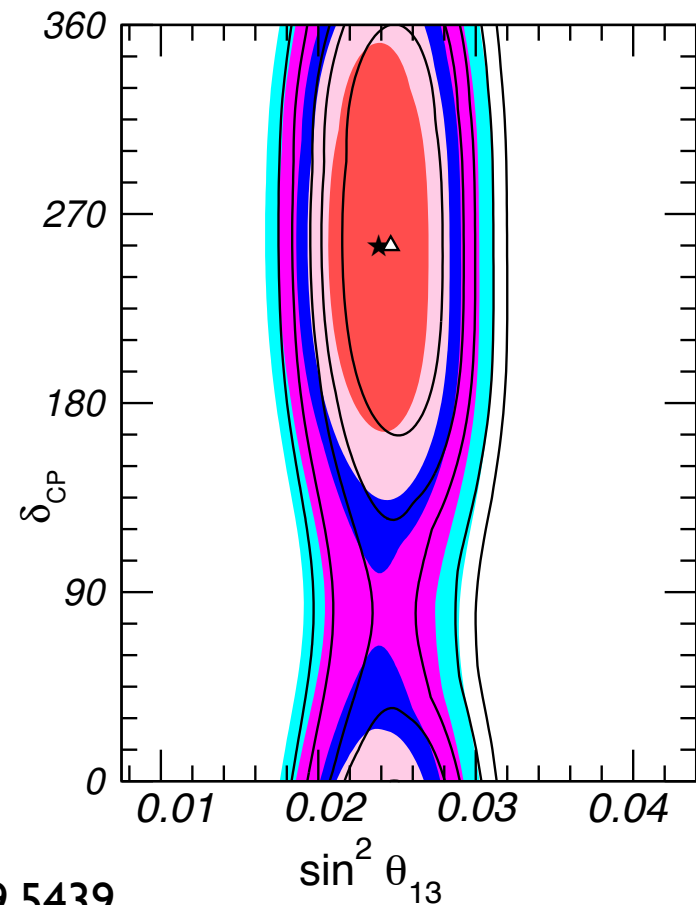
[www.invisibles.eu](http://www.invisibles.eu)

## Hints for CP violation?

There is a slight preference for CP-violation, which is mainly due to the combination of T2K and reactor neutrino data.

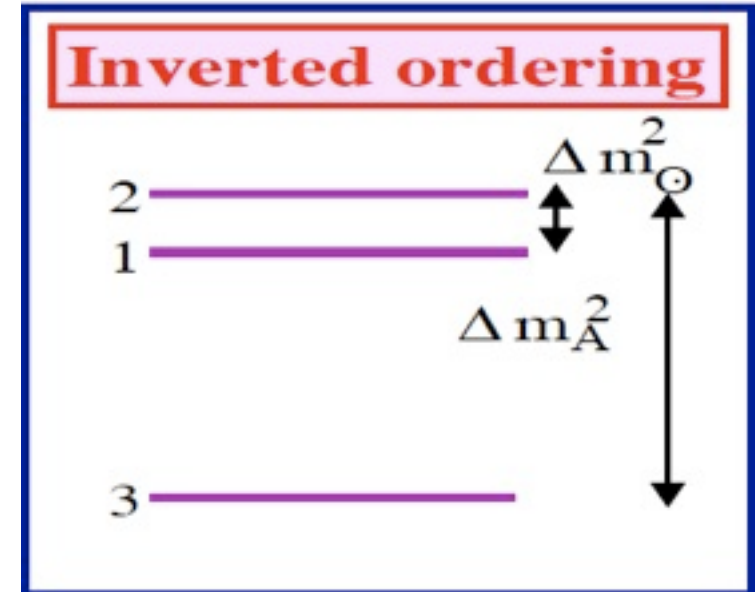
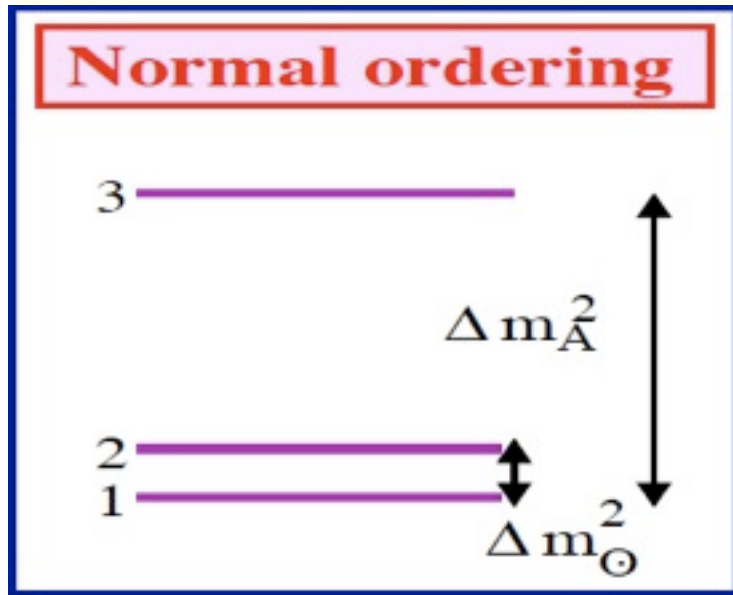


T2K Coll. PRL 112, 061802 (2014)





$\Delta m_s^2 \ll \Delta m_A^2$  implies at least 3 massive neutrinos.



$$\begin{aligned}
 m_1 &= m_{\min} \\
 m_2 &= \sqrt{m_{\min}^2 + \Delta m_{\text{sol}}^2} \\
 m_3 &= \sqrt{m_{\min}^2 + \Delta m_A^2}
 \end{aligned}$$

$$\begin{aligned}
 m_3 &= m_{\min} \\
 m_1 &= \sqrt{m_{\min}^2 + \Delta m_A^2 - \Delta m_{\text{sol}}^2} \\
 m_2 &= \sqrt{m_{\min}^2 + \Delta m_A^2}
 \end{aligned}$$

Measuring the masses requires:

- the mass scale:  $m_{\min}$
- the mass ordering.

The mixing matrix has 3 angles and 1(3) CPV phases.

# Open Phenomenology questions

- **1. What is the nature of neutrinos?**
- **2. What are the values of the masses?** Absolute scale (KATRIN, ...?) and the mass ordering (MO).
- **3. Is there CP-violation?** Its discovery in the next generation of LBL depends on the value of  $\delta$ .
- **4. What are the precise values of mixing angles?** Do they suggest an underlying pattern?
- **5. Is the standard picture correct?** Are there NSI? Sterile neutrinos? Other effects?

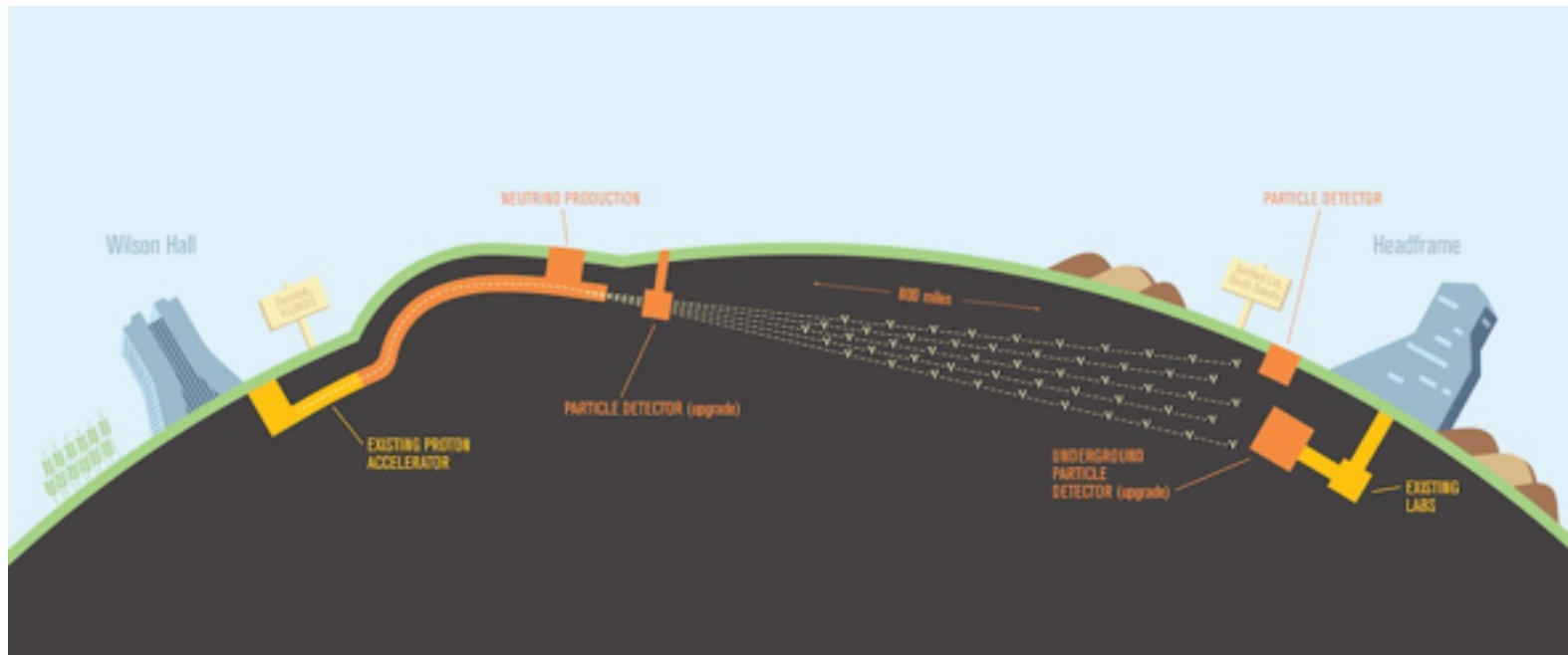
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**How can we search for the mass ordering and leptonic CP-violation in long baseline neutrino oscillation experiments?**

# Long-baseline neutrino oscillations and the ordering

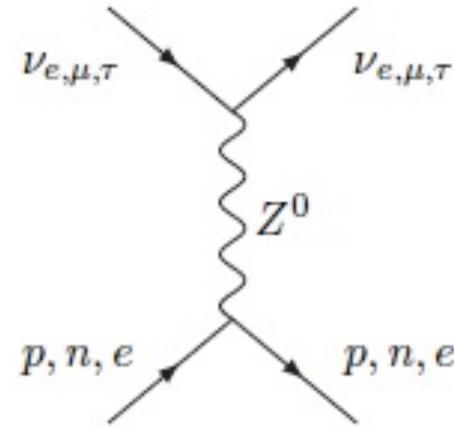
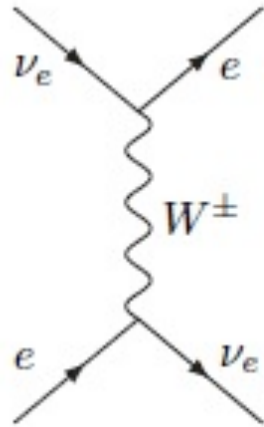
- When neutrinos travel through a medium, they interact with the background of e, p and n.



Credit:  
Symmetry  
magazine

- The background is CP and CPT violating, e.g. the Earth contains only particle and not antiparticles, and the resulting oscillations are CP and CPT violating.

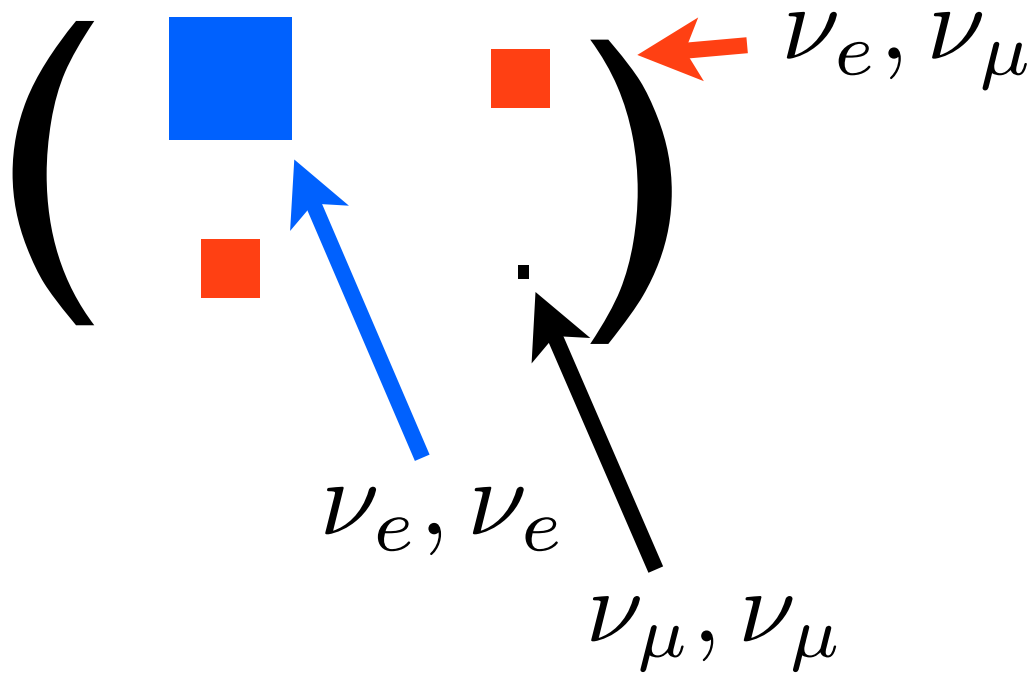
- Neutrinos undergo forward elastic scattering via CC and NC interactions.



- Matter effects are described by a potential  $V$  in the effective Hamiltonian which determines the time evolution.

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} -\frac{\Delta m^2}{4E} \cos(2\theta) + \sqrt{2} G_F N_e & \frac{\Delta m^2}{4E} \sin(2\theta) \\ \frac{\Delta m^2}{4E} \sin(2\theta) & \frac{\Delta m^2}{4E} \cos(2\theta) \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}$$

# Effective Hamiltonian in the flavour basis





## Effective Hamiltonian

$$\begin{pmatrix} \text{blue square} & \text{red square} \\ \text{red square} & \cdot \end{pmatrix}$$

$$\begin{pmatrix} \text{green square} & \text{red square} \\ \text{red square} & \cdot \end{pmatrix}$$

$$\begin{pmatrix} \text{blue square} & \text{red square} \\ \text{red square} & \cdot \end{pmatrix}$$

## Mixing angle

vacuum

$$\tan 2\theta \sim \frac{2 \text{ red square}}{\text{blue square}}$$

matter suppression (Sun, SN)

$$\tan 2\theta^M \sim \frac{2 \text{ red square}}{\text{blue square} + \text{green square}} \ll \tan 2\theta$$

MSW resonance (Sun, SN)

$$\tan 2\theta^M \sim \frac{2 \text{ red square}}{\text{blue square} - \text{gray square}} \sim \infty$$

# In long baseline experiments

$$\text{Blue square} - \frac{\Delta m^2}{2E} \cos(2\theta) \quad \boxed{\nu} + \sqrt{2}G_F N_e \quad \boxed{\bar{\nu}} - \sqrt{2}G_F N_e$$

For neutrinos

$$\begin{pmatrix} \boxed{\text{Blue square}} & \text{Red square} \\ \text{Red square} & \cdot \end{pmatrix}$$

$\Delta m^2 > 0$  enhancement

$$\tan 2\theta^M \sim \frac{2 \text{Red square}}{\text{Blue square} - + \boxed{+}}$$

For antineutrinos

$$\begin{pmatrix} \text{Blue square} & \text{Yellow square} & \text{Red square} \\ \text{Red square} & \cdot \end{pmatrix}$$

$\Delta m^2 > 0$  suppression

$$\tan 2\theta^M \sim \frac{2 \text{Red square}}{\text{Blue square} - + \text{Yellow square} -}$$

Matter effects modify the oscillation probability in LBL experiments.

$$P_{\nu_{\mu} \rightarrow \nu_e} = \sin^2 \theta_{23} \sin^2 2\theta_{13}^m \sin^2 \frac{\Delta m_{13}^2 L}{2}$$

The probability enhancement happens for

- neutrinos if  $\Delta m^2 > 0$
- antineutrinos if  $\Delta m^2 < 0$

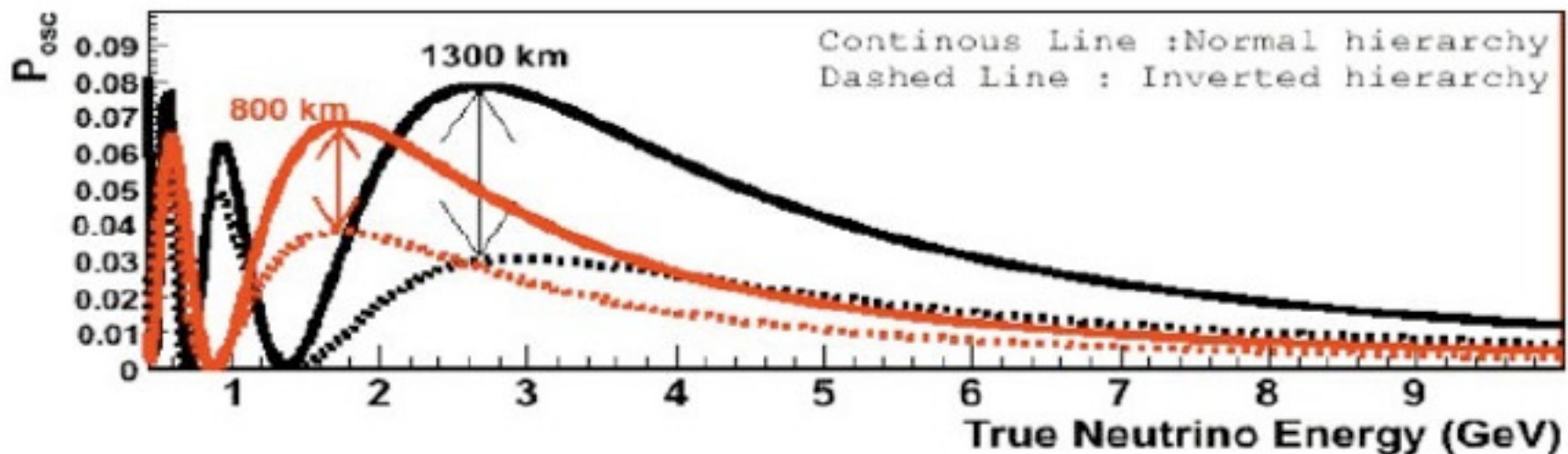
The impact of matter effects is stronger at higher energies and at longer baselines.

# The 3 neutrino probability can be approximated as

$$\begin{aligned}
 P_{\mu e} \simeq & 4c_{23}^2 s_{13}^2 \frac{1}{(1 - r_A)^2} \sin^2 \frac{(1 - r_A)\Delta_{31}L}{4E} \\
 & + \sin 2\theta_{12} \sin 2\theta_{23} s_{13} \frac{\Delta_{21}L}{2E} \sin \frac{(1 - r_A)\Delta_{31}L}{4E} \cos \left( \delta - \frac{\Delta_{31}L}{4E} \right) \\
 & + s_{23}^2 \sin^2 2\theta_{12} \frac{\Delta_{21}^2 L^2}{16E^2} - 4c_{23}^2 s_{13}^4 \sin^2 \frac{(1 - r_A)\Delta_{31}L}{4E}
 \end{aligned}$$

A. Cervera et al., hep-ph/0002108;  
K. Asano, H. Minakata, 1103.4387;  
S. K. Agarwalla et al., 1302.6773...

with  $r_A \equiv \frac{2E}{\Delta m_{31}^2} \sqrt{2} G_F N_e$



# ***CP-violation in LBL experiments***

CP-violation will manifest itself in neutrino oscillations, due to the delta phase. The CP-asymmetry:

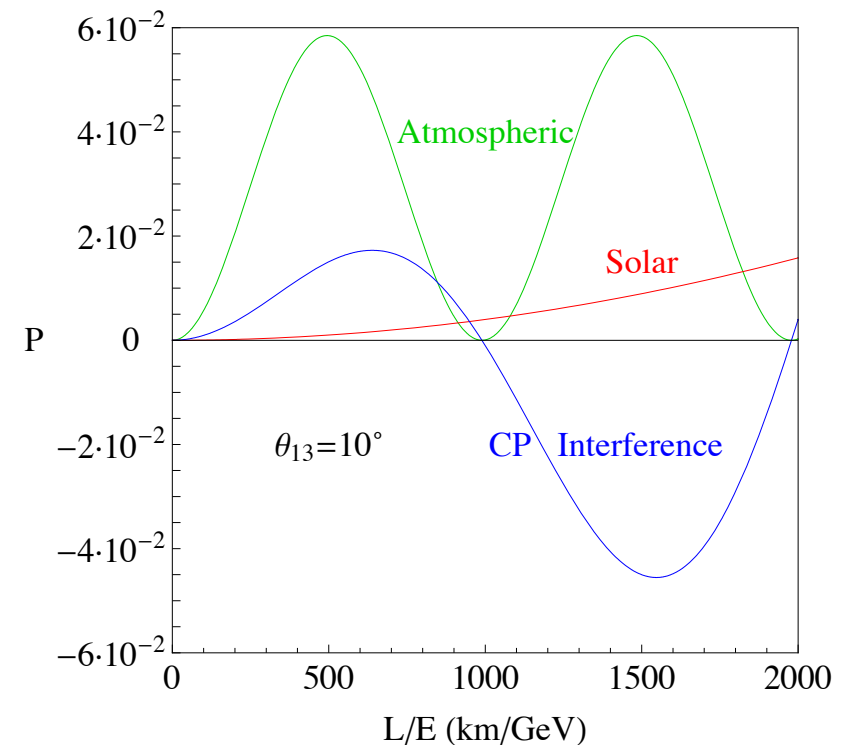
$$P(\nu_\mu \rightarrow \nu_e; t) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e; t) =$$
$$= 4s_{12}c_{12}s_{13}c_{13}^2s_{23}c_{23}\sin\delta \left[ \sin\left(\frac{\Delta m_{21}^2 L}{2E}\right) + \sin\left(\frac{\Delta m_{23}^2 L}{2E}\right) + \sin\left(\frac{\Delta m_{31}^2 L}{2E}\right) \right]$$

- CP-violation requires all angles to be nonzero.
  - It is proportional to the sine of the delta phase.
  - Effective 2-neutrino probabilities are CP-symmetric.
- CPV needs to be searched for in **LBL experiments** which have access to 3-neutrino oscillations.

$$\begin{aligned}
P_{\mu e} \simeq & 4c_{23}^2 s_{13}^2 \frac{1}{(1-r_A)^2} \sin^2 \frac{(1-r_A)\Delta_{31}L}{4E} \\
& + \sin 2\theta_{12} \sin 2\theta_{23} s_{13} \frac{\Delta_{21}L}{2E} \sin \frac{(1-r_A)\Delta_{31}L}{4E} \cos \left( \delta - \frac{\Delta_{31}L}{4E} \right) \\
& + s_{23}^2 \sin^2 2\theta_{12} \frac{\Delta_{21}^2 L^2}{16E^2} - 4c_{23}^2 s_{13}^4 \sin^2 \frac{(1-r_A)\Delta_{31}L}{4E}
\end{aligned}$$

A. Cervera et al., hep-ph/0002108;  
K. Asano, H. Minakata, I 103.4387;  
S. K. Agarwalla et al., I 302.6773...

- The CP asymmetry peaks for  $\sin^2 2\theta_{13} \sim 0.001$ . Large  $\theta_{13}$  makes its searches possible but not ideal.
- Degeneracies with the mass hierarchy and  $\theta_{23}$ .
- CPV effects are more pronounced at low energy.



P. Coloma, E. Fernandez-Martinez, JHEP1204

The **precision measurement of the oscillation parameters** will become very important in the future.

- The values of the mixing angles seem to indicate an underlying symmetry:  $\theta_{23} \sim 45^\circ$ ,  $\theta_{13}$  not too far from 0.
- Predictions for the CPV phase delta and relations among parameters in flavour models (e.g. sum rules).

**Crucial information in order to discriminate between different flavour models.**

**Tests of the standard 3-neutrino paradigm**

- **Sterile neutrinos** (as suggested or not by current hints). Synergy with SBN.
- **New interactions**: NSI, light mediators...
- **Decoherence, Lorentz violation...**

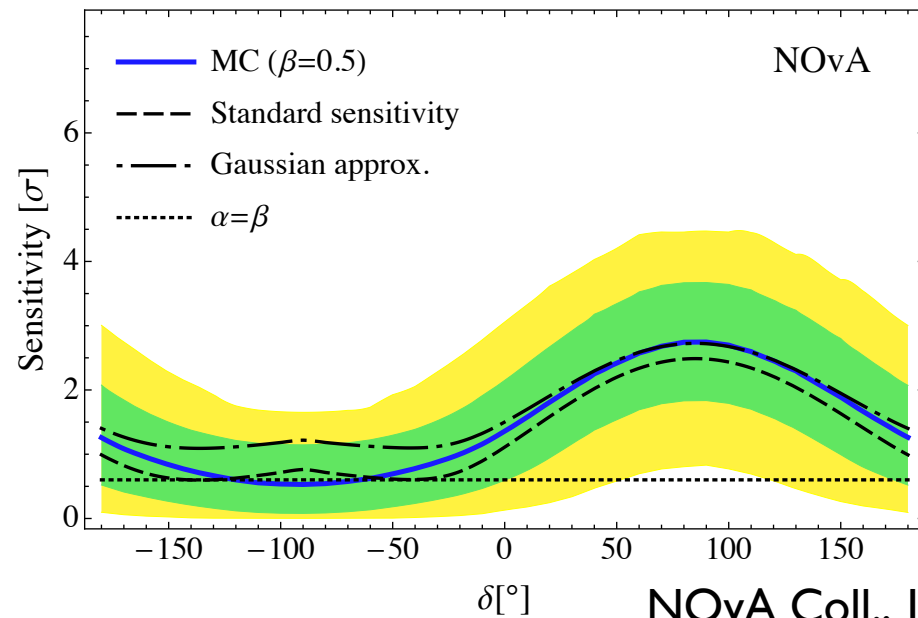
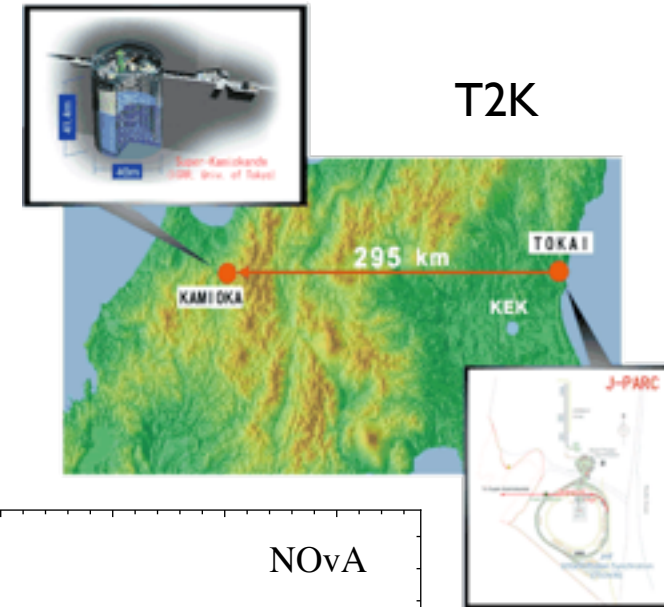
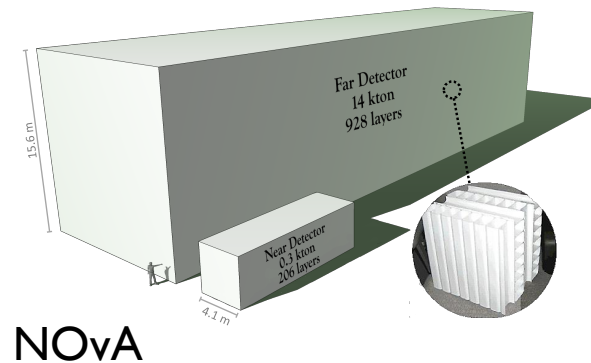
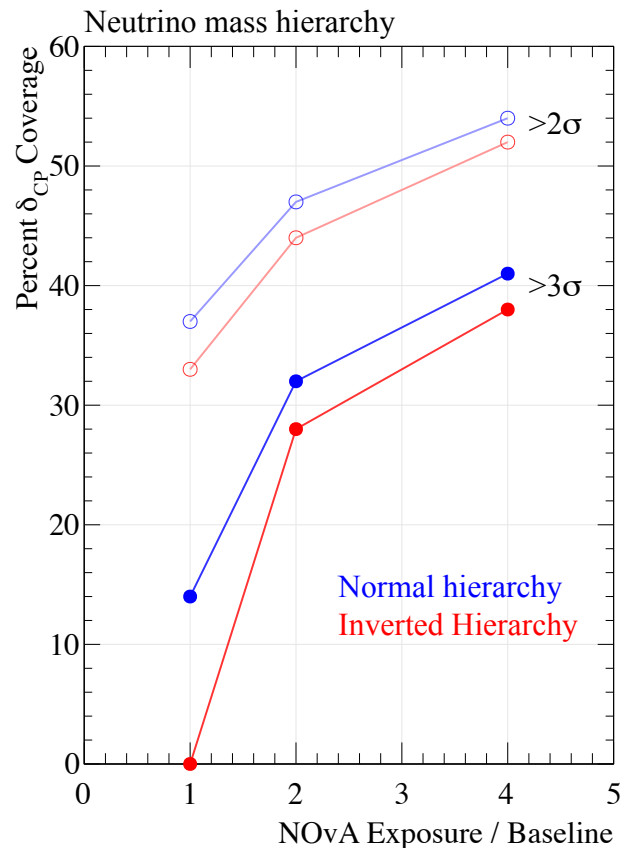
**A deviation from the standard picture would have a groundbreaking impact.**

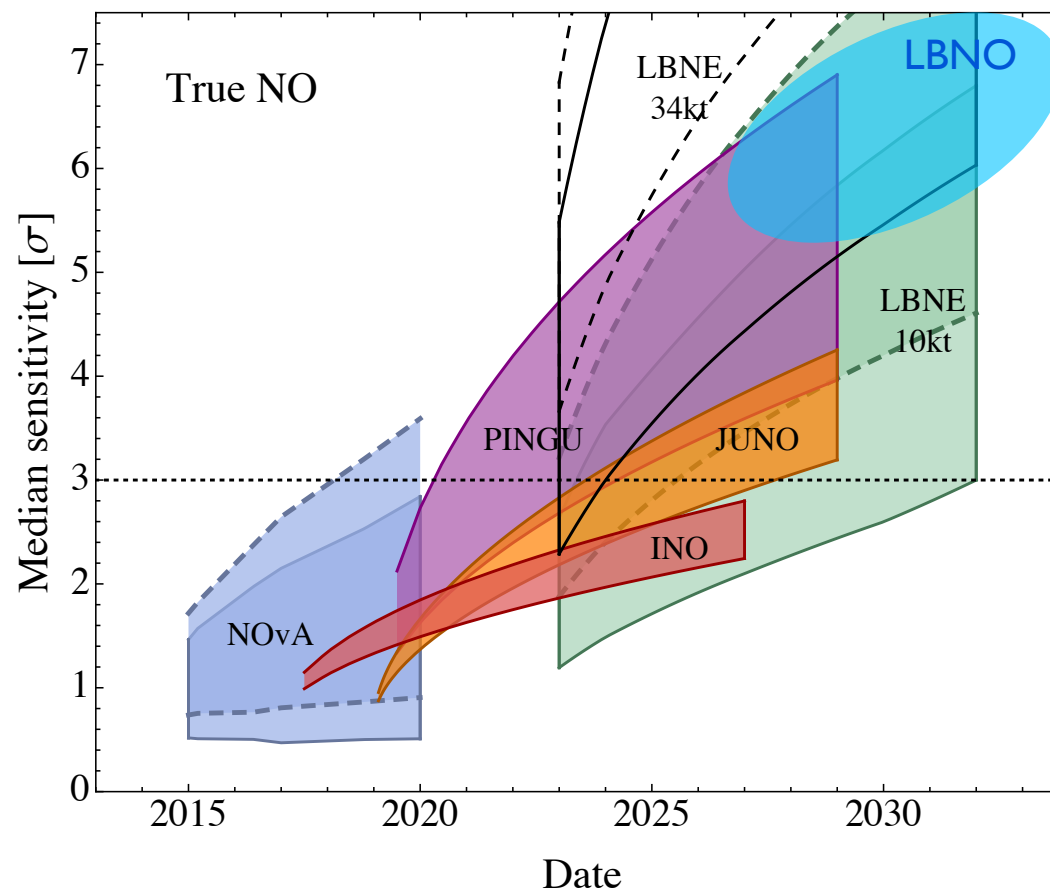


# What we will now by early 2020s?

T2K, NOvA will have collected several years of data, INO is expected to be online as PINGU and JUNO (RENO50) will.

## “NOvAplus”



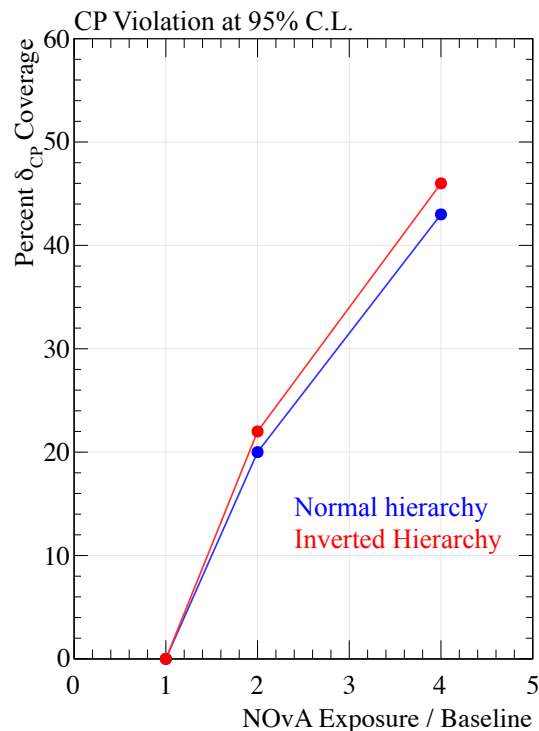


Blennow, Coloma,  
Huber, Schwetz,  
1311.1822

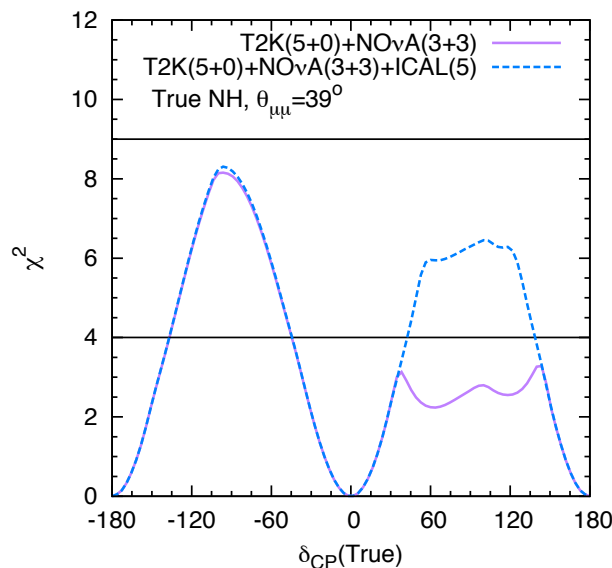
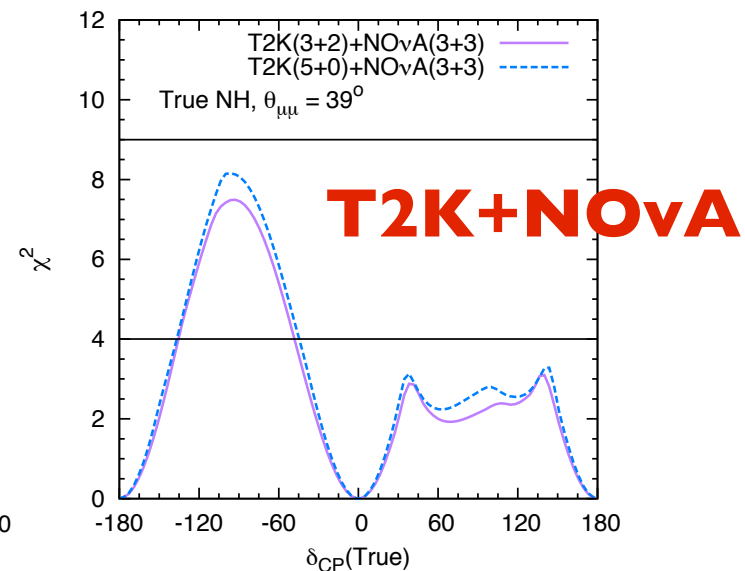
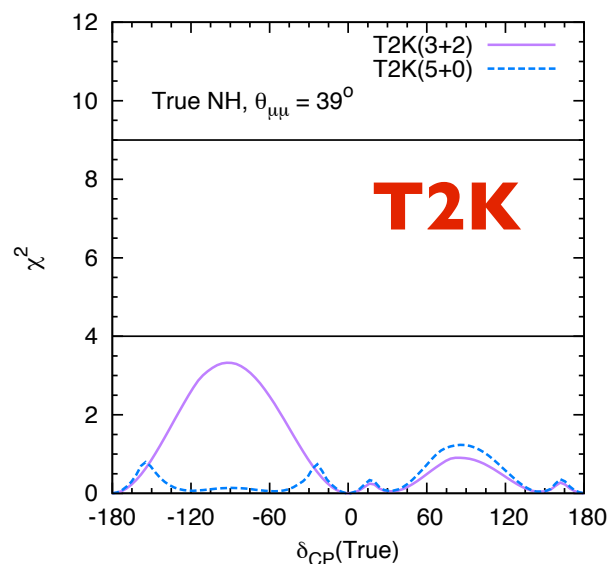
Experiment	Physics effect	Challenges
ELBNF	Matter effects in crust	delta, theta23
Atm nus (INO, PINGU, ORCA)	Matter effects (mantle, mantle+core)	theta23, energy and angular resolution
Reactor exp (JUNO, RENO50)	Vacuum oscillations	energy resolution and reconstruction

# Sensitivity to CPV

## “NOvAplus”



NOvA Coll., 1308.0106



M. Gosh et al., 1401.7243; see also Machado et al.; Huber et al.; For first studies of synergy between T2K and NOvA, see Mena, Nunokawa, Parke, hep-ph/0609011

If delta is in the right region with the right hierarchy, a sensitivity  $>2$  sigma could be achieved.

# ELBNF28

2021 Milestone: 10 kton LArTPC at SURF

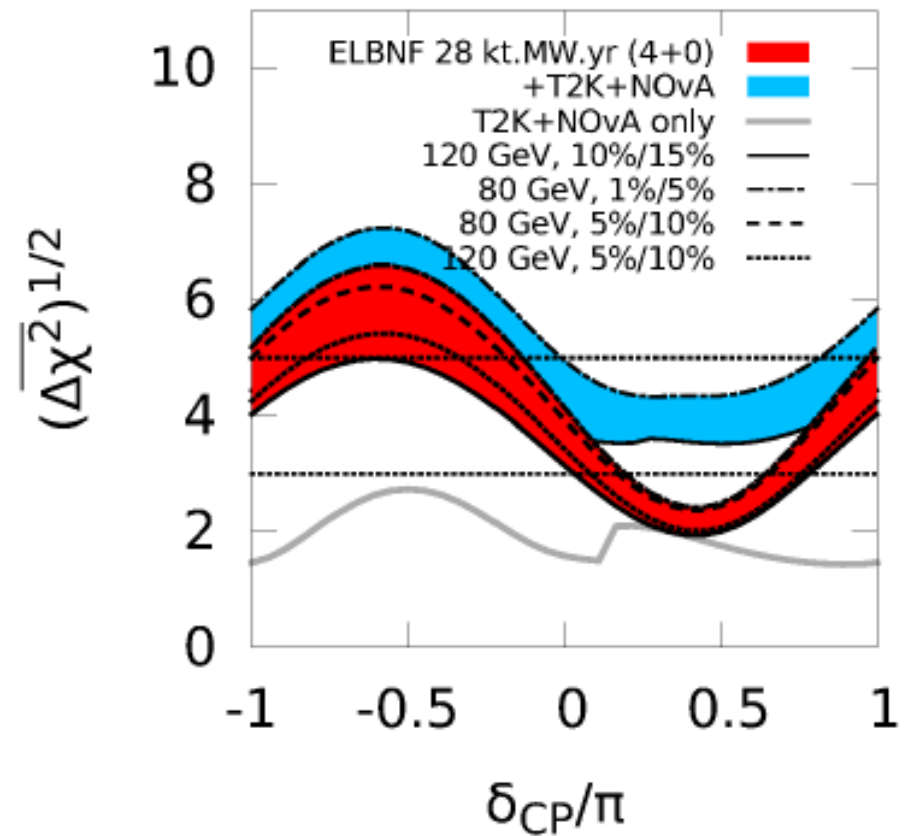
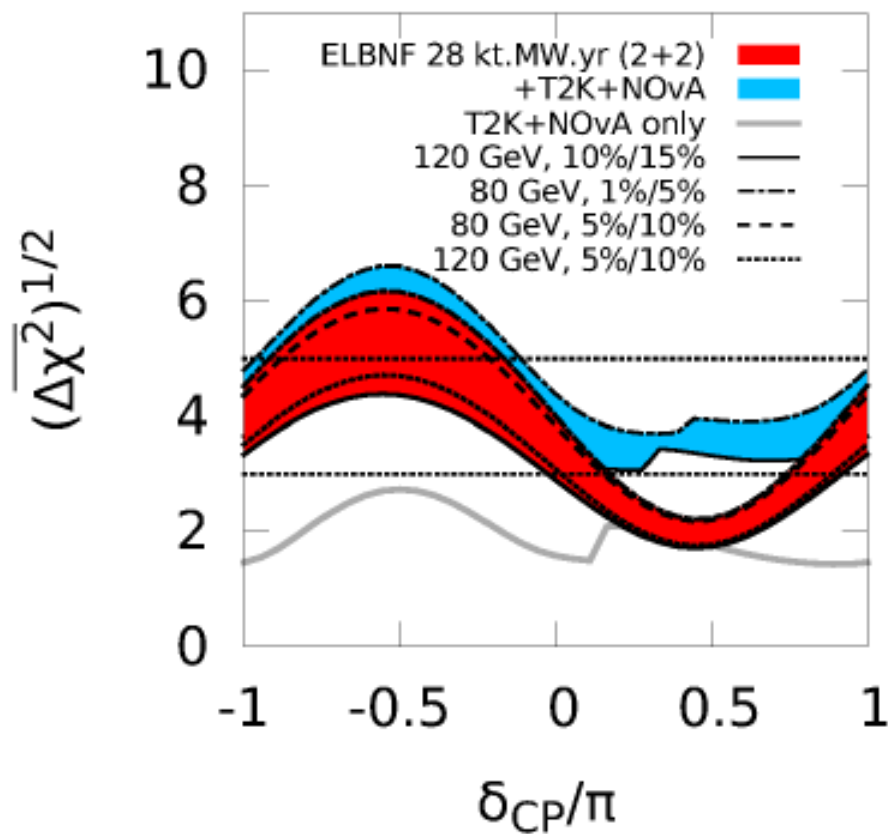
Beam assumptions for oscillation physics sensitivities:  
4 years (neutrino+antineutrino, or neutrinos only),  
700 kW with 80 GeV protons (28 MW-kton-yrs).

## Number of events

Run Mode	Signal Events			Background Events			
	$-\pi/2$	$\delta$ 0	$\pi/2$	$\nu_\mu$ NC	$\nu_\mu$ CC	$\nu_e$ Beam	$\nu_\tau$ CC
Neutrino	117	94	71	8	9	20	6
Antineutrino	18	23	25	4	5	12	4

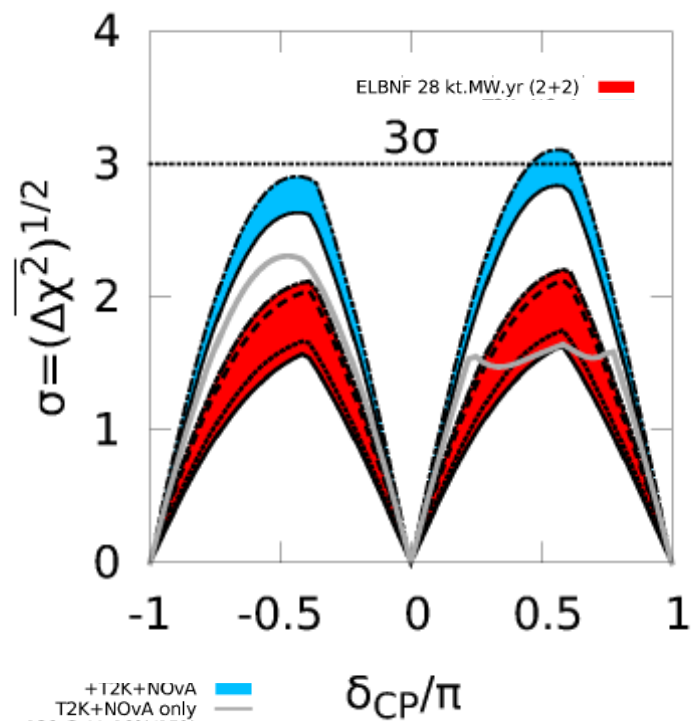
Strongly statistically limited!

# ELBNF28

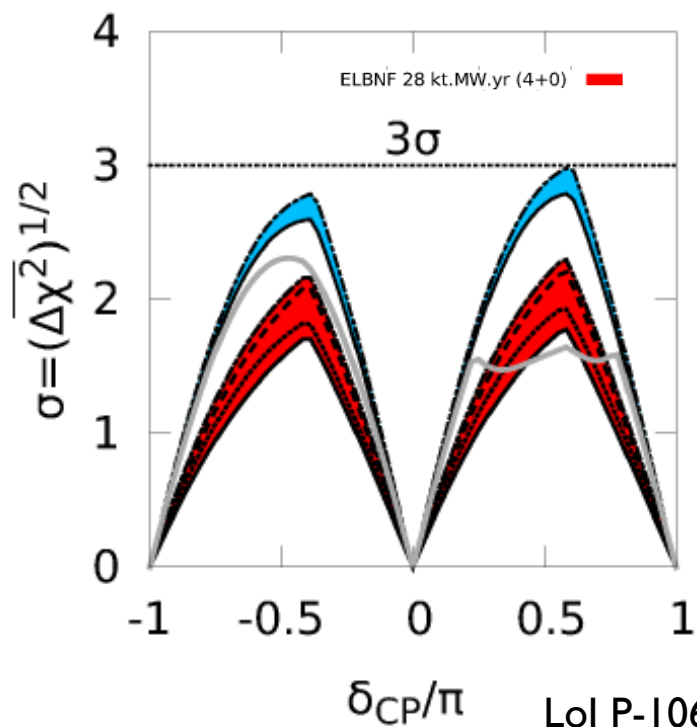


Lol P-1062, submitted to PAC Jan 15

Thanks to the strong matter effects, a sensitivity to MO of 2-6 sigma could be achieved, depending on the delta phase. For the favoured region (to be confirmed), better than 3 sigma determination of MO.

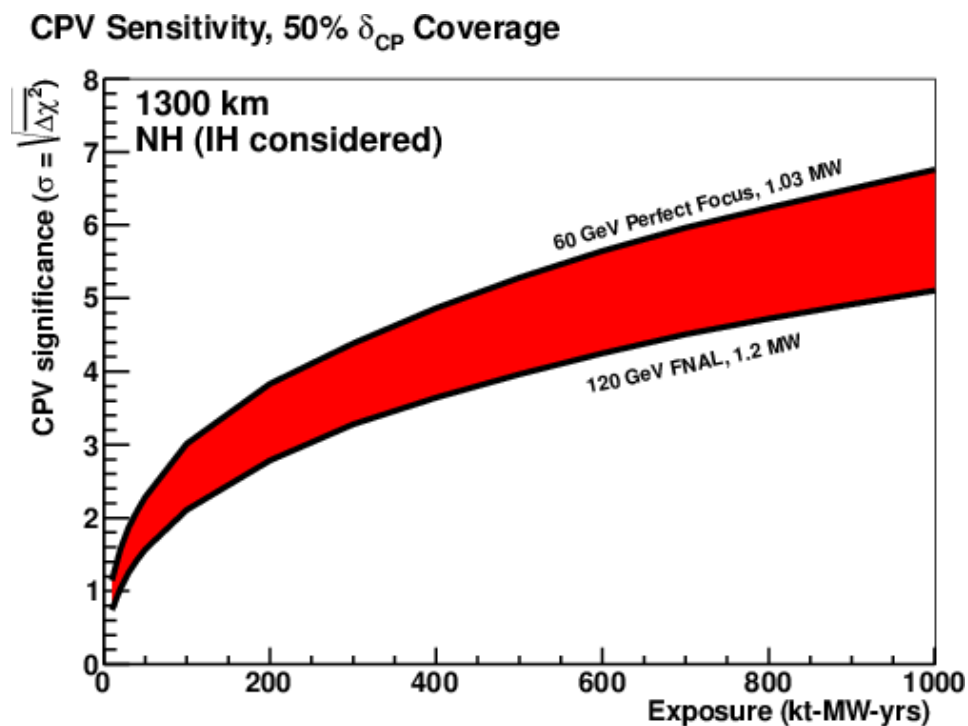
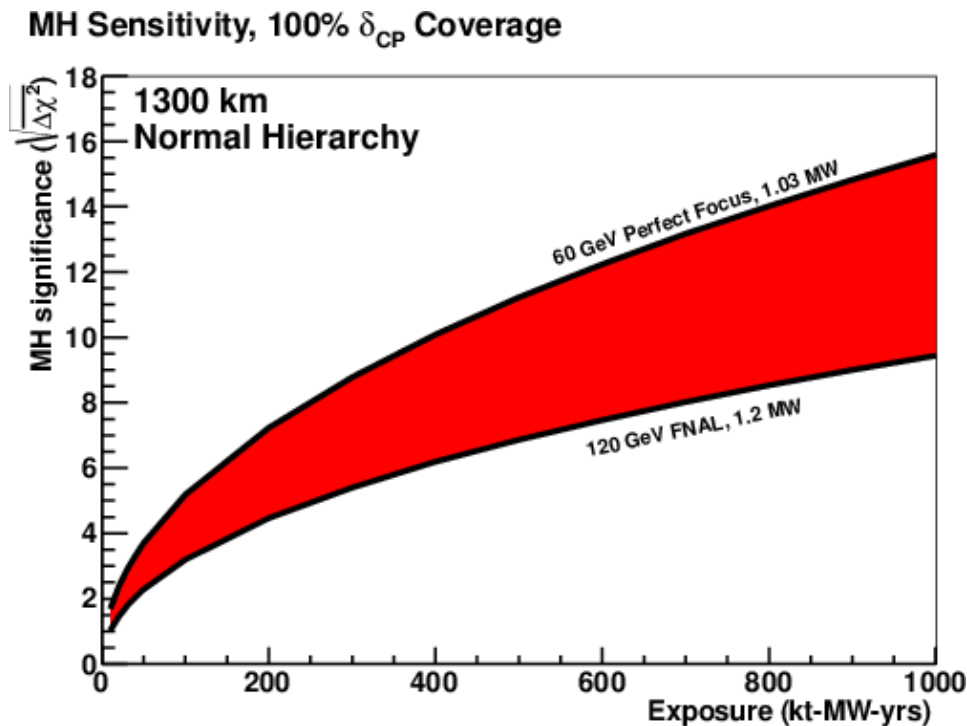


+T2K+NOvA  
 T2K+NOvA only  
 120 GeV, 10%/15%  
 80 GeV, 1%/5%  
 80 GeV, 5%/10%  
 120 GeV, 5%/10%



- This measurement is strongly limited by statistics.
- The addition of NOvA and T2K data makes a significant improvement in the reach.
- Running in neutrino and antineutrino mode is preferable.
- A 3 sigma CPV sensitivity could be within reach for maximal CPV.

# Effect of exposure



LOI P-1062

Predicted sensitivities should be read with some caution as they critically depend on:

- setup assumed: detector and its performance, beam...
- values of oscillation parameters and their errors;
- treatment of backgrounds and systematic errors.

Space for further optimizations.

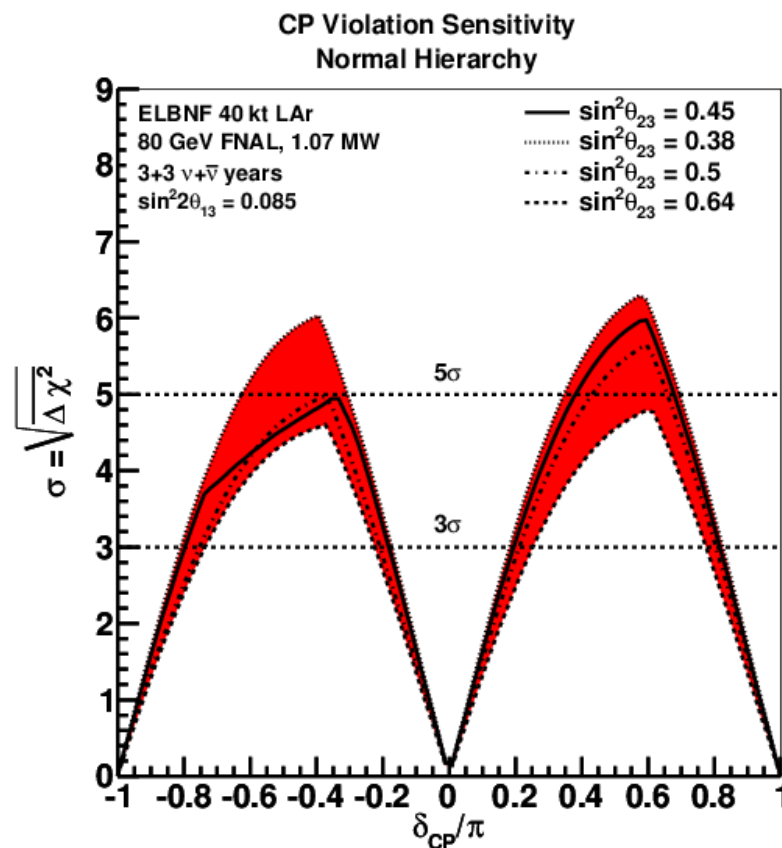
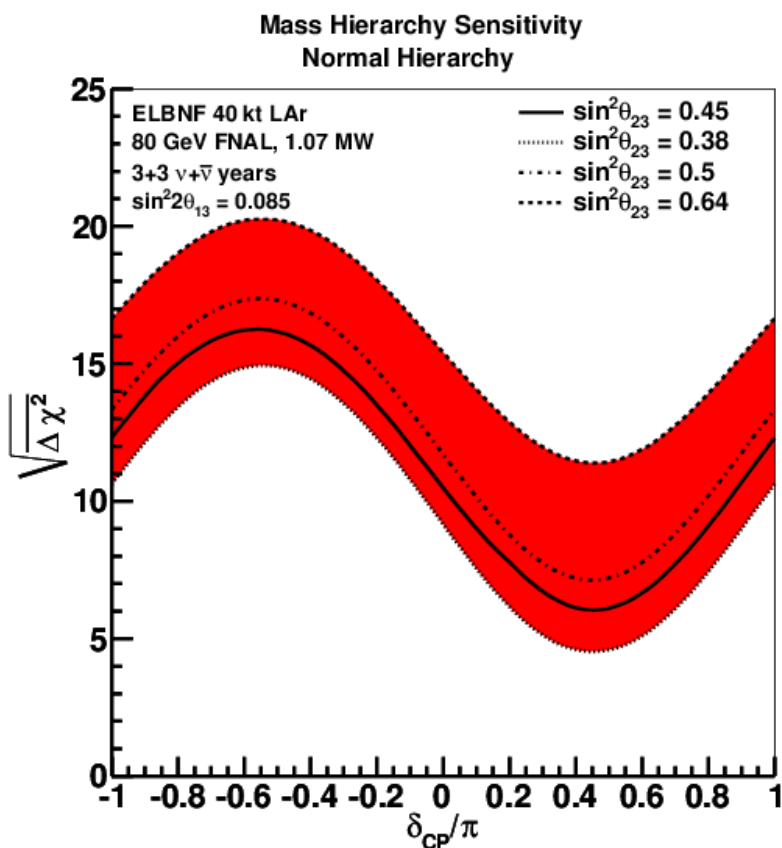


# ELBNF-34kton

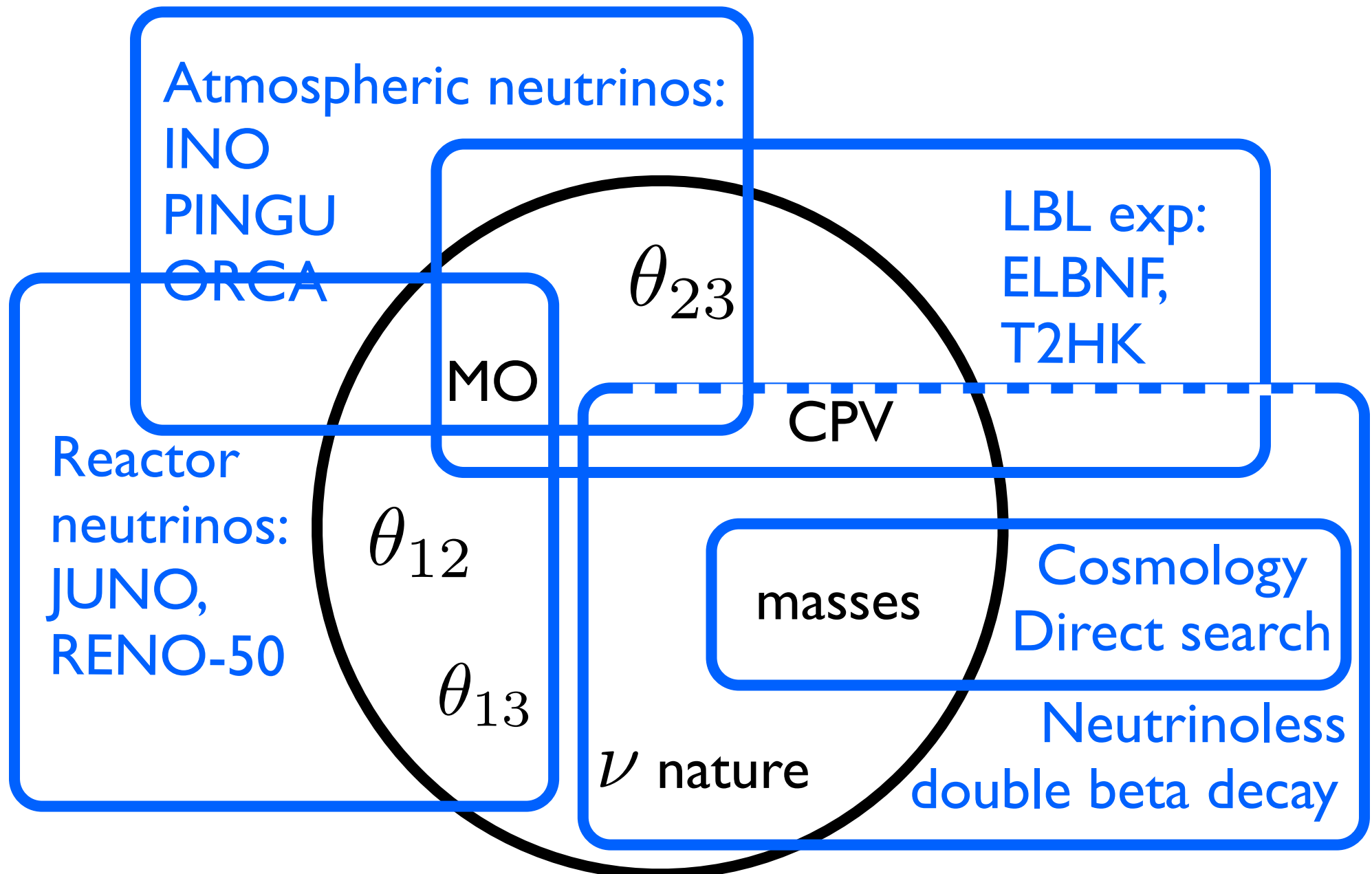
40 kton LArTPC, 80 GeV 1.07 MW beam, 3 yrs  
neutrinos + 3 yrs antineutrinos

LoI P-1062, submitted to PAC Jan 15

Run Mode	Signal Events			Background Events			
	$-\pi/2$	$0$	$\pi/2$	$\nu_\mu$ NC	$\nu_\mu$ CC	$\nu_e$ Beam	$\nu_\tau$ CC
Neutrino	1068	864	649	72	83	182	55
Antineutrino	166	213	231	41	42	107	33



# Complementarity with other searches



Also: Tests of standard neutrino paradigm

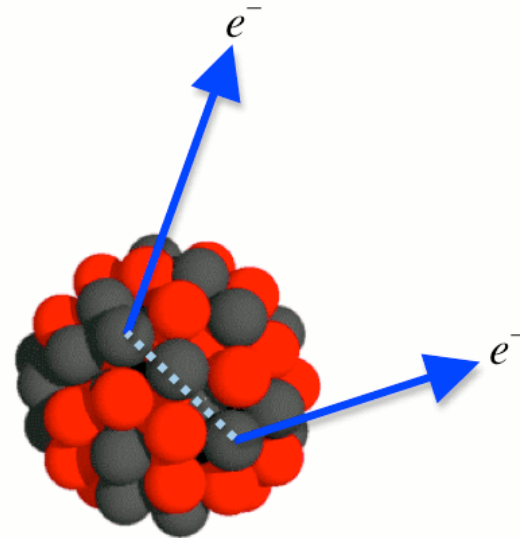
# Synergy

LBL

with

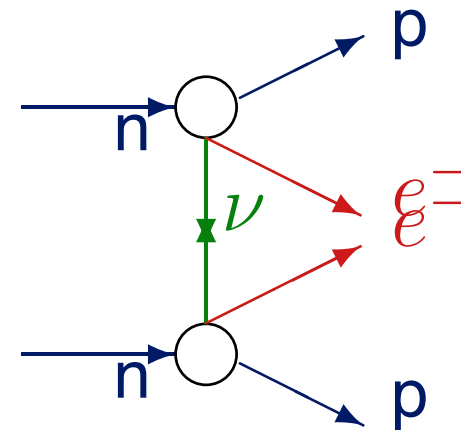
Neutrinoless  
double beta decay

Neutrinoless double beta decay,  $(A, Z) \rightarrow (A, Z+2) + 2 e^-$ , will test the nature of neutrinos.



Thanks to R. Saakyan,  
talk at NuPhys 2014

At the fundamental level, exchange of light Majorana neutrino (or other exotic mechanism).



The half-life time depends on neutrino properties

$$\left[ T_{0\nu}^{1/2}(0^+ \rightarrow 0^+) \right]^{-1} \propto |M_F - g_A^2 M_{GT}|^2 |\langle m \rangle|^2$$

- $|\langle m \rangle| = m_{ee}$  : the effective Majorana mass parameter

$$|\langle m \rangle| \simeq |m_1 \sin^2 \theta_{12} + m_2 \cos^2 \theta_{12} e^{i\alpha_{21}} + m_3 \sin^2 \theta_{13} e^{i\alpha_{31}}|$$

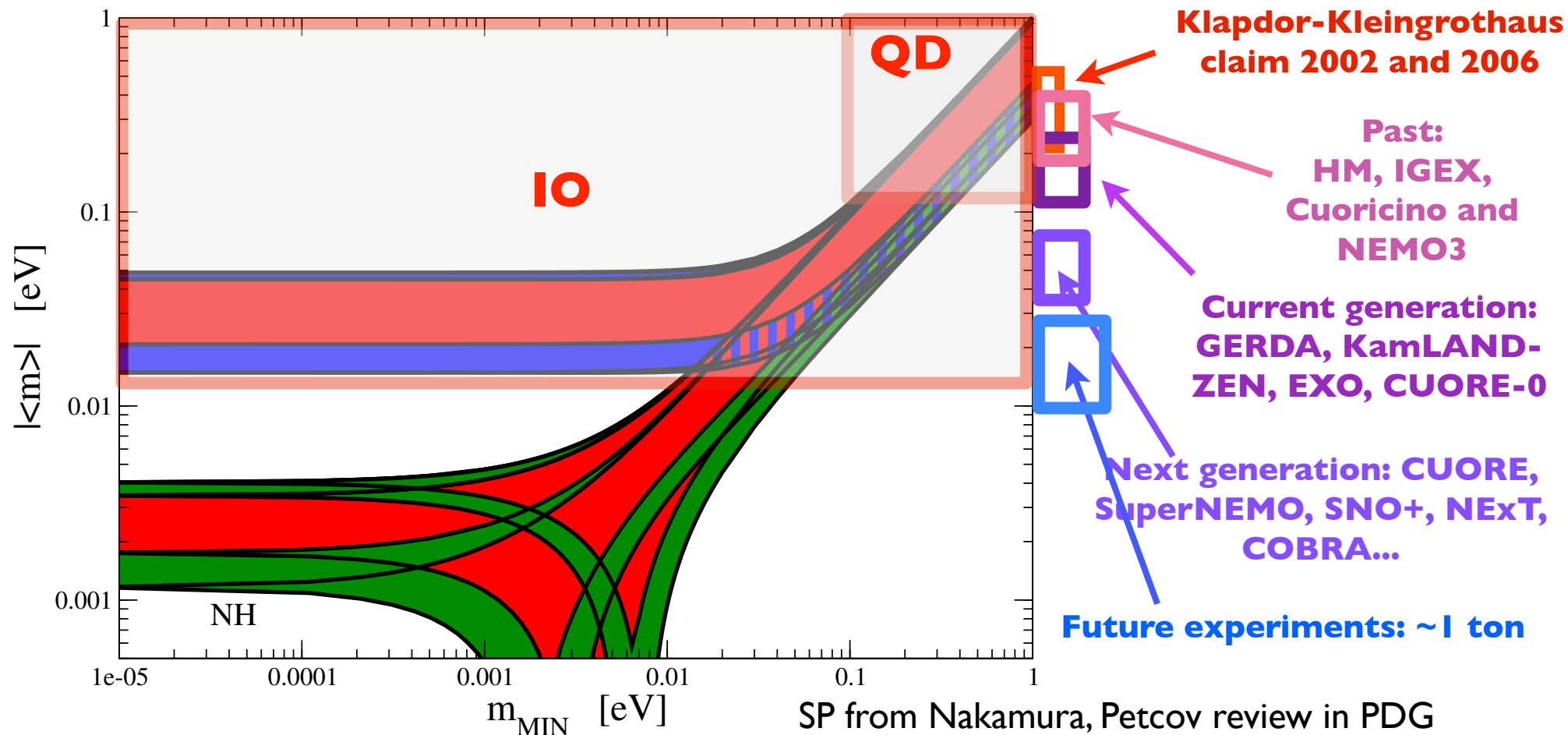
Mixing angles (known)

CPV phases (unknown)

Masses (partially known)

Example: **IH** ( $m_3 \ll m_1 < m_2$ ):  $10 \text{ meV} < |\langle m \rangle| < 50 \text{ meV}$

$$|\langle m \rangle| \simeq \sqrt{\Delta m_{31}^2} |\cos^2 \theta_{12} + \sin^2 \theta_{12} e^{i\alpha_{21}}|$$



If:

LBL, atm...  
finds IO

Nullless dbeta decay  
No signal down  
to  $m_{ee} \sim 10$  meV

Nus are Dirac particles or cancellations in double beta decay (e.g. low energy see-saw)

# Conclusions

- LBL experiments can search for the mass ordering, CPV, precise values of the oscillation parameters, tests of the 3-neutrino scenario.
- There is a strong complementarity and synergy with other searches (e.g. neutrinoless double beta decay, atmospheric, reactor and supernova neutrinos, cosmology,...).
- ELBNF28 is a milestone towards ELBNF with an improved physics reach compared to T2K and NOvA: sensitivity to MO and CPV (depending on  $\delta$ ).