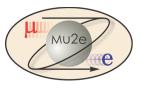
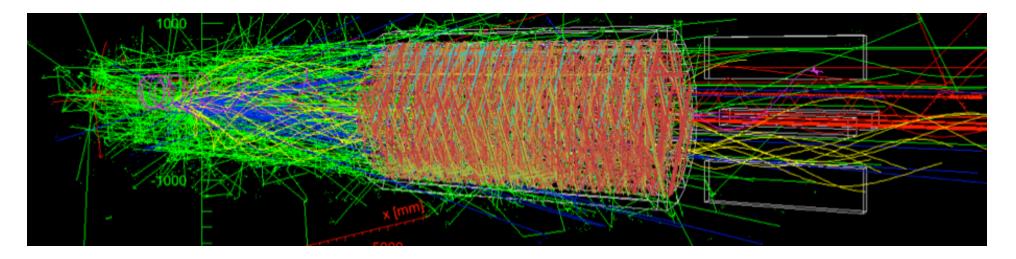


#### The Mu2e Experiment





Ron Ray Mu2e Project Manager









#### Mu2e Collaboration



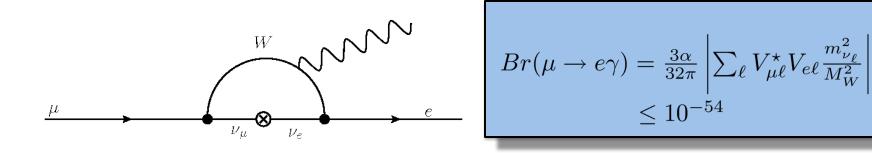
~160 Collaborators, 26 Institutions, 3 Countries

#### Introduction

- Mu2e is a search for Charged Lepton Flavor Violation (CLFV) via the coherent conversion of  $\mu^- N \rightarrow e^- N$
- Target sensitivity has great discovery potential
  - Goal: Single-event-sensitivity of 2 x 10<sup>-17</sup> (relative to ordinary  $\mu$  capture)
  - Goal: <0.5 events background
    - $\circ$  Yields Discovery Sensitivity for all rates > few 10<sup>-16</sup>
- Most new physics models so far postulated provide new sources of flavor phenomena
- Quark flavor is violated. Neutrino flavor is violated.
  - Both implied something profound about the underlying physics
  - Both garnered Nobel Prizes
- What about charged lepton flavor?

# Neutrino Oscillations and CLFV

- Neutrinos oscillate, so lepton flavor is not conserved.
- Charged leptons *must* mix through neutrino loops.
  - But the mixing is so small, it's effectively forbidden.



• No Standard Model pollution! Observation is unambiguous evidence for new physics.

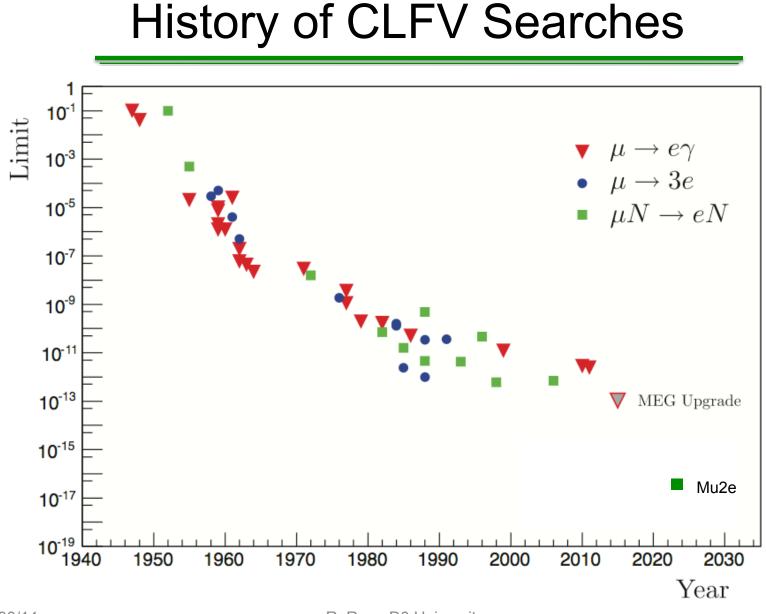
# **Charged Lepton Flavor Violation**

- Significant worldwide interest from Energy and Intensity Frontiers and the theoretical community
  - Rare CLFV decays of  $\mu$ ,  $\tau$ , K, B-mesons
- Rates negligible in vSM but wide array of new physics models predict rates that are <u>measurable in next generation experiments</u>.
  - Sensitive to new physics well above the TeV scale  $(10^3 10^4 \text{ TeV})$ .
- Rates of CLFV processes are model dependent and vary widely depending on the underlying physics.
  - CLFV processes are powerful discriminators.
- The most stringent limits on CLFV come from muons because of the relative "ease" of producing an intense source.
- Muon-to-electron conversion offers excellent discovery potential across a breadth of models and will explore impressive mass scales.

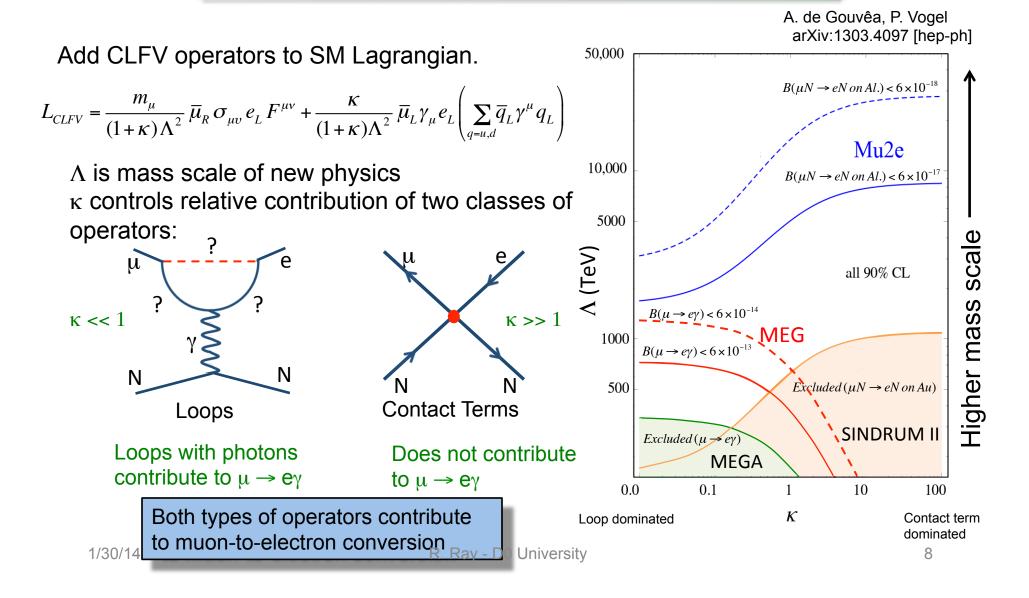
#### Some CLFV Processes

Process	Current Limit	Next Generation exp
τ> μη	BR < 6.5 E-8	
τ> μγ	BR < 6.8 E-8	10 <sup>-9</sup> - 10 <sup>-10</sup> (Belle II)
τ> μμμ	BR < 3.2 E-8	
$\tau \longrightarrow$ eee	BR < 3.6 E-8	
K <sub>L</sub> > eμ	BR < 4.7 E-12	
K+> π+e <sup>-</sup> μ+	BR < 1.3 E-11	NA62
B <sup>0</sup> > eμ	BR < 7.8 E-8	Belle II, LHCb
B+> Κ+eμ	BR < 9.1 E-8	
μ+> <b>e</b> +γ	BR < 5.7 E-13	10 <sup>-14</sup> (MEG)
μ+> e+e+e-	BR < 1.0 E-12	10 <sup>-16</sup> (PSI)
μN> eN	R <sub>μe</sub> < 7.0 Ε-13	10 <sup>-17</sup> (Mu2e, COMET)

The most sensitive CLFV probes use muons



#### Model Independent Evaluation



#### Mu2e Physics Reach

		AC	RVV2	AKM	$\delta LL$	FBMSSM	LHT	RS	
★★★ = Discovery Sensitivity	$D^0 - \bar{D}^0$	***	*	*	*	*	***	?	
	$\epsilon_K$	*	***	***	*	*	**	***	
	$S_{\psi\phi}$	***	***	***	*	*	***	***	
	$S_{\phi K_S}$	***	**	*	***	***	*	?	
	$A_{\rm CP}\left(B  ightarrow X_s \gamma ight)$	*	*	*	***	***	*	?	
	$A_{7,8}(B ightarrow K^{\star}\mu^+\mu^-)$	*	*	*	***	***	**	?	
	$A_9(B\to K^*\mu^+\mu^-)$	*	*	*	*	*	*	?	
	$B \to K^{(\star)} \nu \bar{\nu}$	*	*	*	*	*	*	*	
	$B_s  ightarrow \mu^+ \mu^-$	***	***	***	***	***	*	*	a
	$K^+ \to \pi^+ \nu \bar{\nu}$	*	*	*	*	*	***	***	Xiv
	$K_L  ightarrow \pi^0  u ar u$	*	*	*	*	*	***	***	:09
	$\mu \to e \gamma$	***	***	*	***	***	***	***	<u> </u>
	$\tau \to \mu \gamma$	***	***	*	***	***	***	***	133
	$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***	3[h
	$d_n$	***	***	***	**	***	*	***	arXiv:0909.1333[hep-ph]
	$d_e$	***	***	**	*	***	*	***	[dc
	$(g-2)_{\mu}$	***	***	**	***	***	*	?	

W. Altmannshofer, A.J.Buras, S.Gori, P.Paradisi, D.M.Straub

Table 8: "DNA" of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models  $\bigstar \bigstar \bigstar$  signals large effects,  $\bigstar \bigstar$  visible but small effects and  $\bigstar$  implies that the given model does not predict sizable effects in that observable.

R. Ray - D0 University

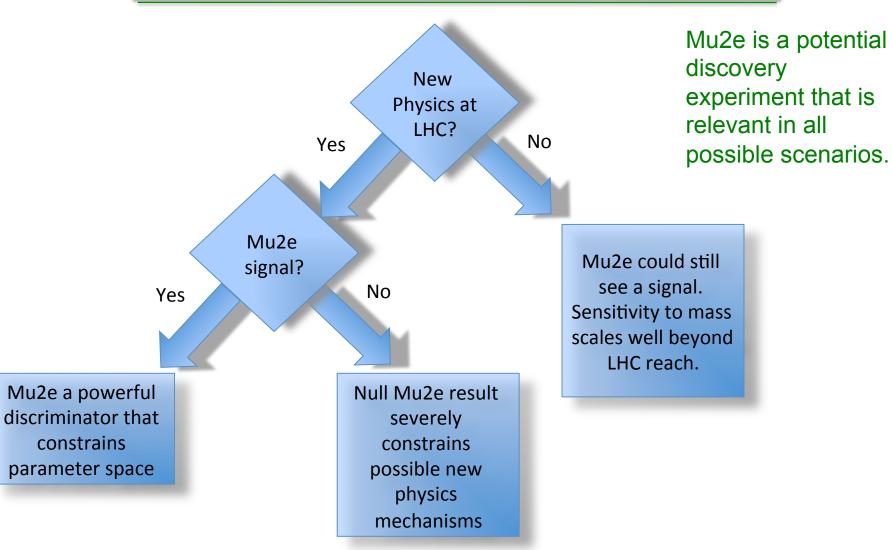
#### Mu2e Physics Reach

	AC	RVV2	AKM	$\delta$ LL	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	***	*	*	*	*	***	?
$\epsilon_K$	*	***	***	*	*	**	***
$S_{\psi\phi}$	***	***	***	*	*	***	***
$S_{\phi K_S}$	***	**	*	***	***	*	?
$A_{\rm CP}\left(B  ightarrow X_s \gamma ight)$	*	*	*	***	***	*	?
$A_{7,8}(B\to K^*\mu^+\mu^-)$	*	*	*	***	***	**	?
$A_9(B\to K^*\mu^+\mu^-)$	*	*	*	*	*	*	?
$B \to K^{(\star)} \nu \bar{\nu}$	*	*	*	*	*	*	*
$B_s \to \mu^+ \mu^-$	***	***	***	***	***	*	*
$K^+ \to \pi^+ \nu \bar{\nu}$	*	*	*	*	*	***	***
$K_L  ightarrow \pi^0  u ar{ u}$	*	*	*	*	*	***	***
$\mu \to e \gamma$	***	***	*	***	***	***	***
$\tau \to \mu \gamma$	***	***	*	***	***	***	***
$\mu + N \rightarrow e + N$	***	***	***	***	***	***	***
$d_n$	***	***	***	**	***	*	***
$d_e$	***	***	**	*	***	*	***
$(g-2)_{\mu}$	***	***	**	***	***	*	?

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#### Mu2e has discovery sensitivity across the board

#### Mu2e and the LHC



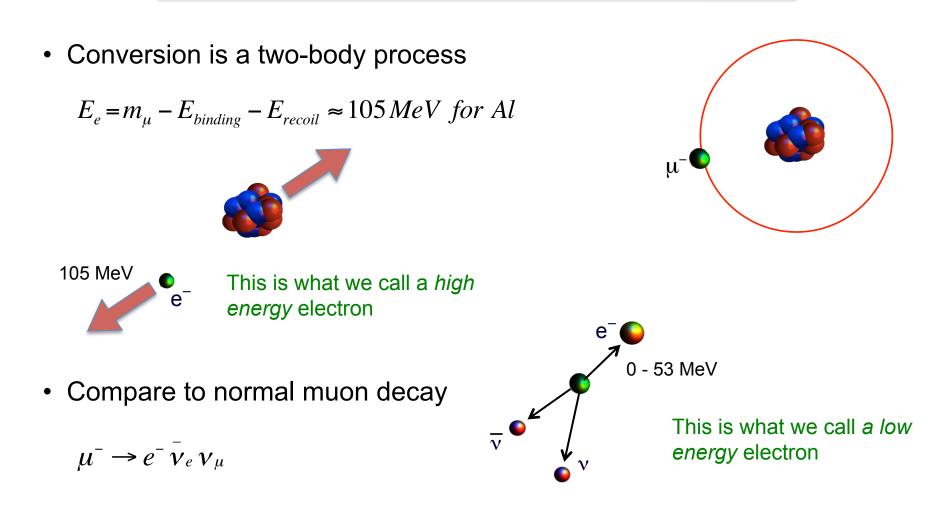
#### How does Mu2e work?

# Mu2e Concept

- Generate an intense pulsed beam of low momentum  $\mu^{\scriptscriptstyle -}$ 
  - Use 8 GeV protons from Booster
- Stop muons in orbit around a nucleus
  - $\tau_{\mu}^{Al} = 864 \ ns$
- Wait for prompt backgrounds to go away
  - Pulsed beam
- Measure delayed electron spectrum
  - Signal is a monoenergetic 105 MeV electron

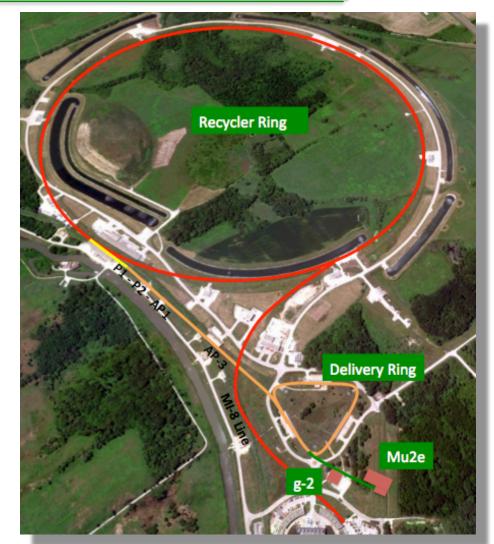
# Sounds easy. But achieving 10<sup>-17</sup> sensitivity is very challenging

#### Introduction

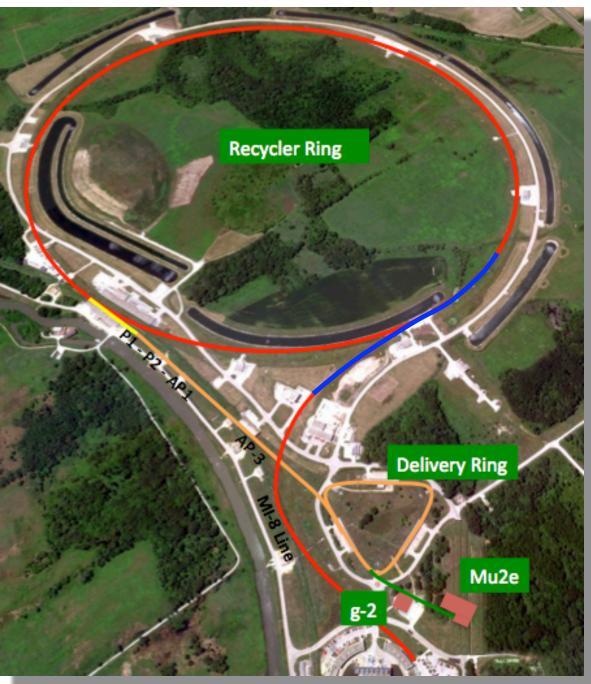


#### Fermilab Accelerator Complex

- Fermilab has a powerful and flexible accelerator complex.
- Multiple machines and multiple rings affords great flexibility and the ability to operate multiple programs simultaneously.
- End of Tevatron program freed up the pbar source and the Recycler Ring for other programs.
- Mu2e can reuse existing rings to package and deliver protons to create the world's most intense muon beam with no impact on the 120 GeV neutrino program (NOvA).



R. Ray - D0 University

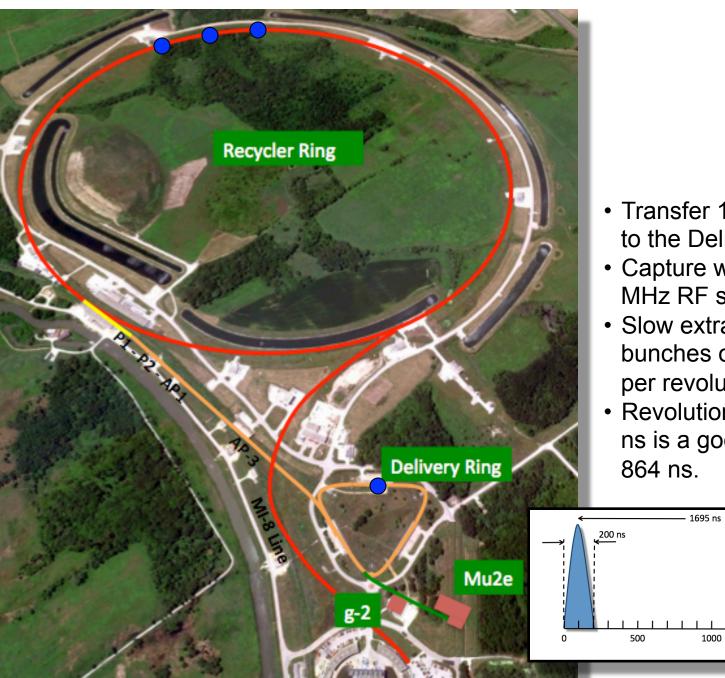


Transfer 2 Booster batches of 8 GeV protons to Recycler ring during Main Injector ramp.

R. Ray - D0 University



Rebunch into 4 bunches using new 2.5 MHz RF system.



- Transfer 1 bunch at a time to the Delivery Ring.
- Capture with a new 2.4 MHz RF system
- Slow extract microbunches of ~ 10<sup>7</sup> protons per revolution to Mu2e.
- Revolution time of 1695 ns is a good match to  $\tau^{AI}$  = 864 ns.



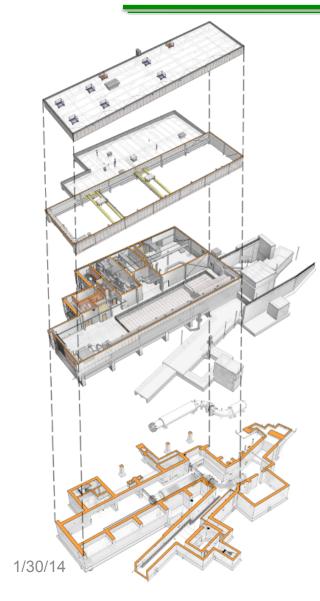
2000 Time (ns)

1500

1/30/14

R. Ray - D0 University

### **Experimental Hall**





- 100% design from A&E in hand
- Solicit bids in early 2014.
- Bids in hand for CD-2
- Begin site prep work this summer
- Begin building construction on October 1 using FY14 construction funds.

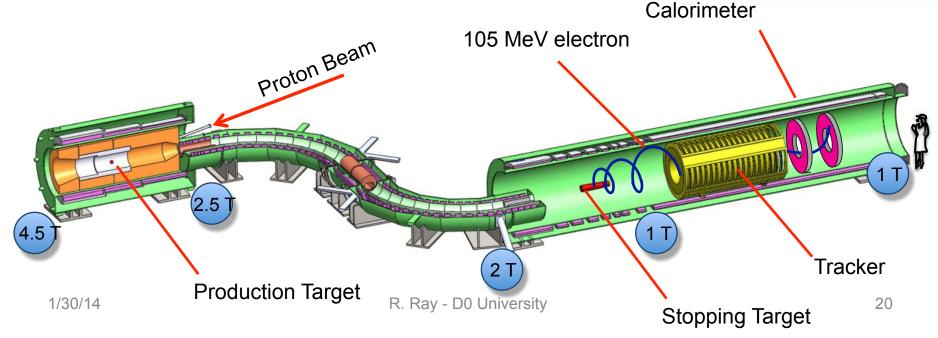
R. Ray - D0 University

• Solenoids capture pions, form secondary muon beam, preserve timing structure, provide magnetic field for momentum analysis and help to reject backgrounds

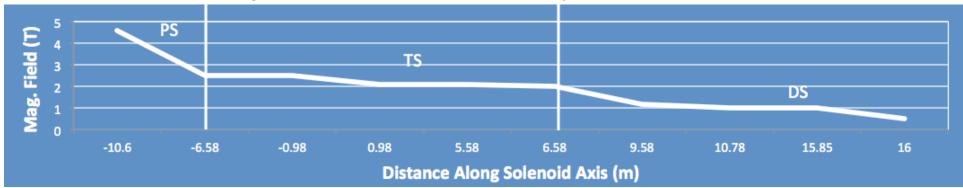
**Cosmic Ray Veto and Stopping** 

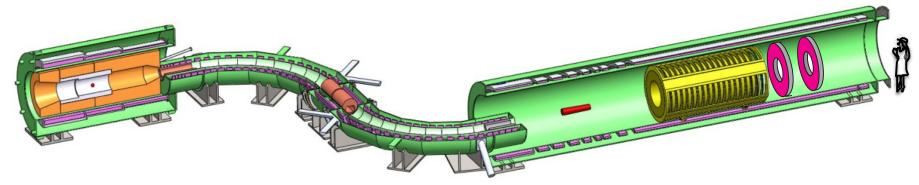
**Target Monitor not shown** 

- Most efficient way of producing an intense, low energy muon beam
- 2 targets
- Tracker Straw tubes
- Calorimeter BaF2 crystals
- Cosmic Ray Veto Scintillator, WLS fibers, SiPMs
- Stopping Target Monitor Crystal
- Warm bore of solenoids evacuated to 10<sup>-4</sup> to 10<sup>-5</sup> Torr.



Magnetic Field Profile – Driven by the science requirements



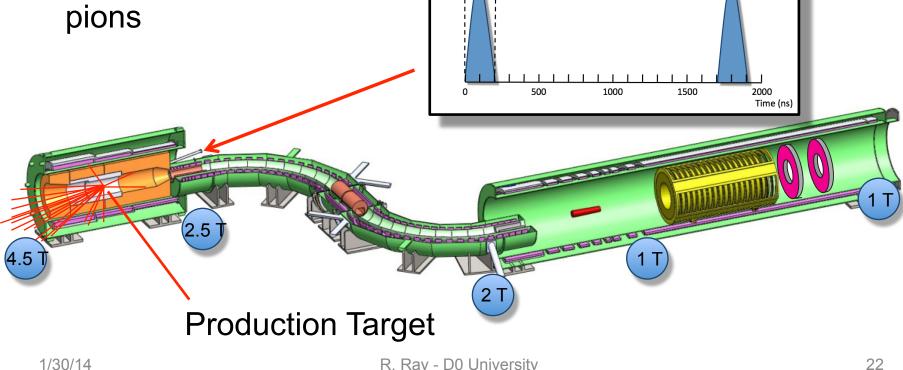


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#### **Production Solenoid**

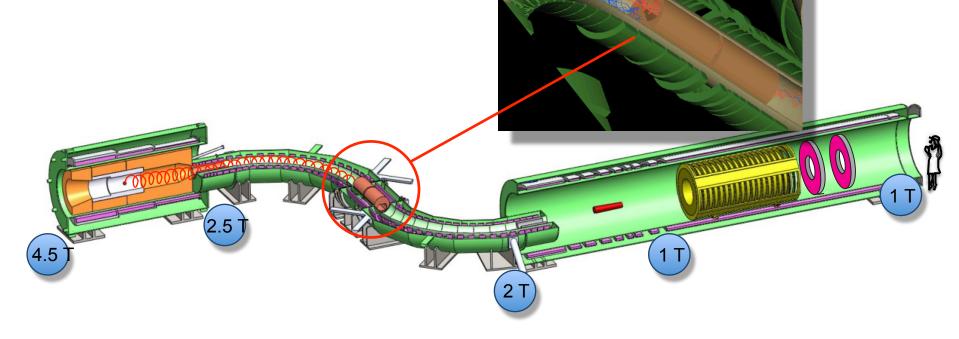
- Production target
- Graded field
- Captures secondary pions





#### **Transport Solenoid**

 Collimation system selects muon charge and momentum range



**Transport Solenoid** 

- Collimation system selects muon charge and momentum range
- Pbar window in middle of central collimator
- Directs 10<sup>10</sup> Hz of  $\mu^-$  to stopping target

4.5

40 MeV/c μ<sup>-</sup>

All Stopped muon

20 30 40 50 60 70 80 90 100 110 120

P (MeV/c)

Stopping Target

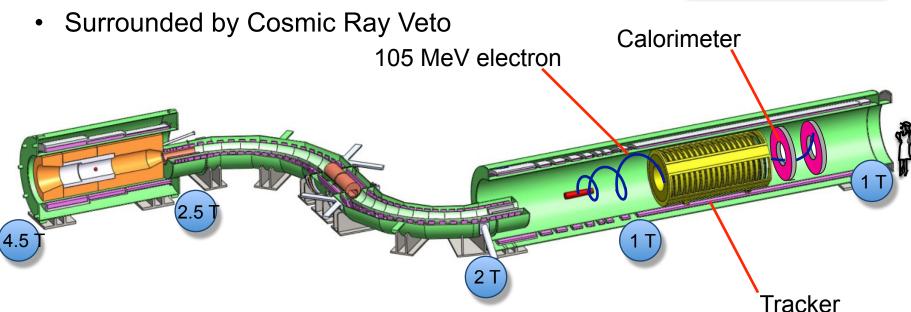
25000

Events/1.7 MeV/c 120000 10000

5000

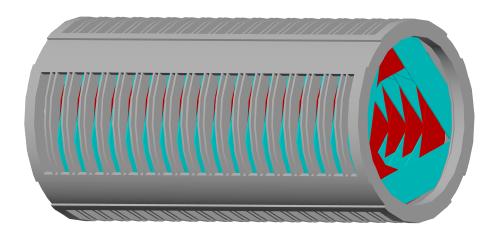
**Detector Solenoid** 

- Graded field upstream for acceptance and background suppression
- Uniform field downstream for momentum analysis
- Muon stopping target
- Tracker
- Calorimeter

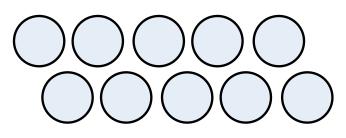


Cosmic Ray Veto not shown

#### Tracker

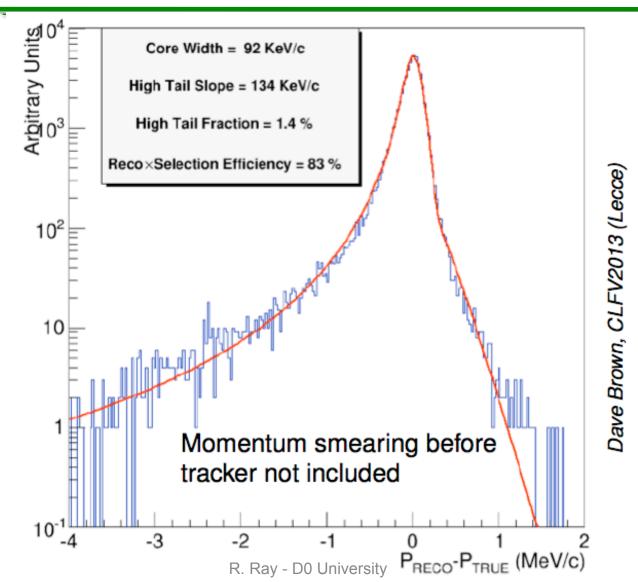


- ~ 3 m long
- 1 T "uniform" B field
- Operates in 10<sup>-4</sup> T vacuum
- Good tracks make 2 3 turns
- TDC readout at both ends of wire
  - $\Delta t$  gives position along the wire
- ADC readout for particle ID



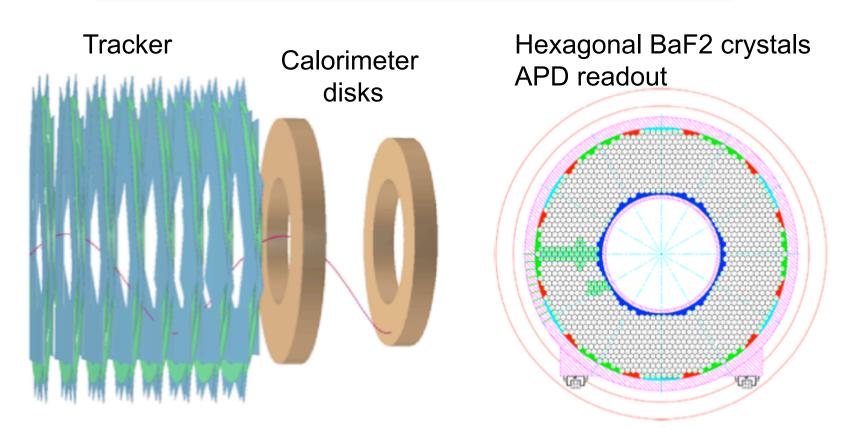
- ~ 25,000 straw tubes mounted transverse to solenoid axis
- 5 mm diameter
- 15 μm wall thickness
- 25 µm wire
- Good resolution required to control background from muon decay in orbit
- Resolution dominated by multiple scattering and straggling in stopping target.

#### **Tracker Momentum Resolution**



1/30/14

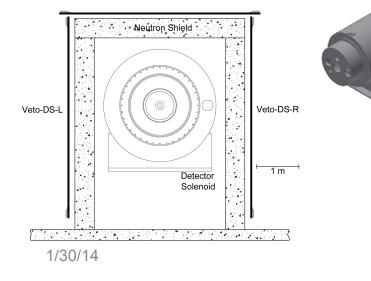
## Calorimeter



Provides good timing, PID, track seed, trigger. Corse confirmation of tracker signal events.

## Cosmic Ray Veto

- Cosmic Ray Veto covers all of Detector Solenoid
   and half of Transport Solenoid
- Nearly hermetic veto on top and sides with 0.9999
   overall efficiency
- Shielding of neutrons from stopping target and collimators required.
- 4 layers of extruded scintillator
- SiPM readout



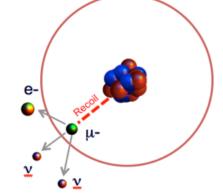
#### Backgrounds

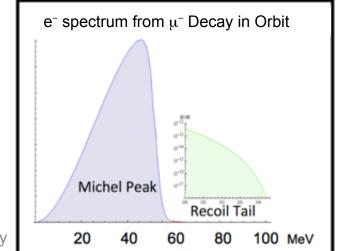
Backgrounds fall into 4 categories:

- Muon decay in orbit
- Prompt processes where an electron is detected shortly after the arrival of a particle at the stopping target
- Cosmic rays
- Antiprotons

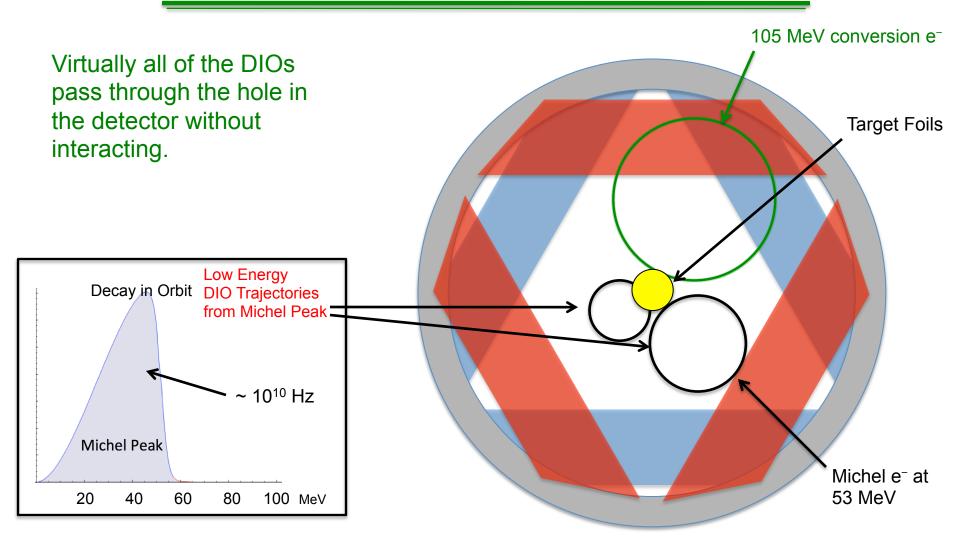
# Muon Decay in Orbit Background

- μ at rest can produce an electron whose energy is at most half of the μ rest mass, the decay of a bound μ can result in an electron with energy approaching that of a conversion electron.
- In kinematic limit where the v's carry away no energy the electron recoils against the nucleus simulating the two-body conversion process.
- DIO spectrum is a distorted Michel peak and a small recoil tail extending to the conversion energy, falling as  $(E_{endpoint} - E_e)^5$ near the endpoint.
  - ~3 × 10<sup>-13</sup> of the DIO electrons within 3 MeV of endpoint.
  - Sets the scale for the required energy resolution of the experiment.



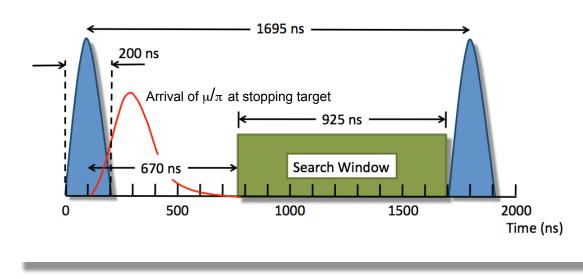


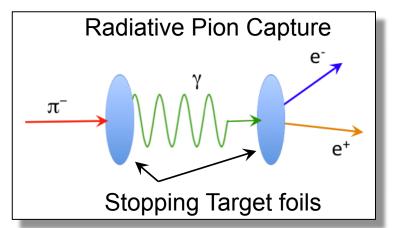
#### Rate from DIOs



#### **Prompt Backgrounds**

*Prompt background:* Processes where the detected background electron is nearly coincident in time with the arrival of a beam particle at the muon stopping target.

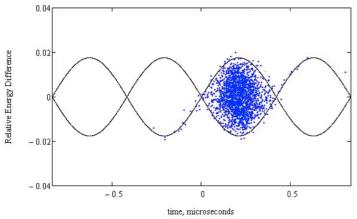




Pulsed beam combined with long  $\tau^{Al}$ , delayed search window and extinction of beam between pulses reduces prompt backgrounds like Radiative Pion Capture.

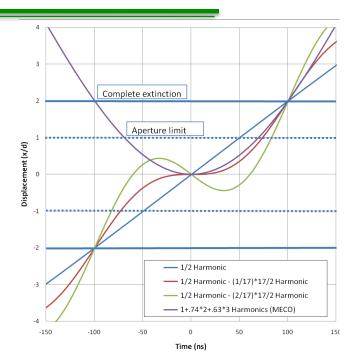
## Extinction

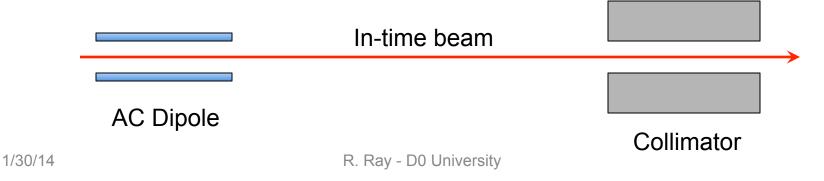
- Extinction requirement of 10<sup>-10</sup> is not trivial, but can be accomplished through a combination of internal and external extinction.
- Internal
  - We use RF to form and hold tight bunches in both the Recycler and Delivery Rings.
  - <10<sup>-5</sup> in Recycler Ring
  - Extinction in Delivery Ring < 10<sup>-4</sup>
     Degrades during slow extraction



#### Extinction

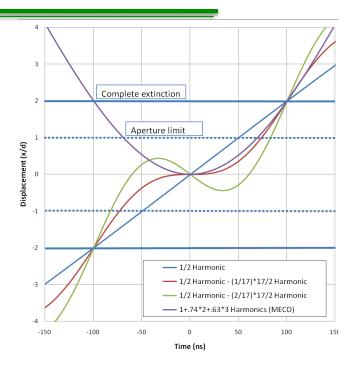
- External extinction is accomplished with a high frequency AC Dipole operating at 300 kHz with an admixture of a 5.1 MHz 17th harmonic.
- Unwanted beam swept into collimators.
- MARS simulations indicate this is adequate, but not a trivial magnet to build.
  - Extinction from AC dipole < 10<sup>-7</sup>





#### Extinction

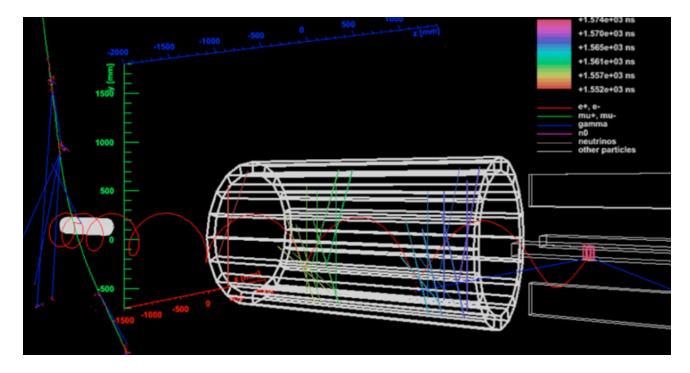
- External extinction is accomplished with a high frequency AC Dipole operating at 300 kHz with an admixture of a 5.1 MHz 17th harmonic.
- Unwanted beam swept into collimators.
- MARS simulations indicate this is adequate, but not a trivial magnet to build.
  - Extinction from AC dipole < 10<sup>-7</sup>
- Bunches at 600 kHz beyond the state-ofthe-art for traditional kickers.





# Cosmic Ray Background

- Scales with run time, not POT. Modest improvement from previous experiments all that is required.
- ~1 event/day without CRV



### Antiprotons

#### A source of background

- Negatively charged, so they can be transported to stopping target.
- Do not decay
- Heavy and slow, so they arrive during signal window, even with a perfect extinction system.
- 2 GeV shower

#### Mitigation

- Thin window in the central collimator
- Reduces muon yield by 15% but effectively annihilates antiprotons

# Summary of Backgrounds

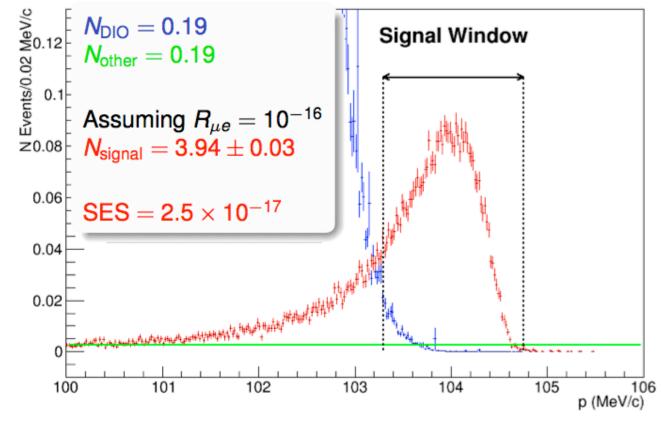
3 years at 1.2 × 10<sup>20</sup> protons/year (8 kW Beam power)

Background	Expected events
Muon decay in orbit	$0.20 \pm 0.06$
Antiproton induced	$0.10 \pm 0.06$
Radiative pion capture*	$0.04 \pm 0.02$
Cosmic rays**	0.050 ± 0.013
Beam electrons*	$0.001 \pm 0.001$
Muon decay in flight*	$0.010 \pm 0.005$
Total	$0.4 \pm 0.1$

#### \* Assumes 10<sup>-10</sup> beam extinction \*\* For 10<sup>-4</sup> cosmic ray veto inefficiency

## Signal

3 years at 1.2 × 10<sup>20</sup> protons/year (8 kW Beam power)



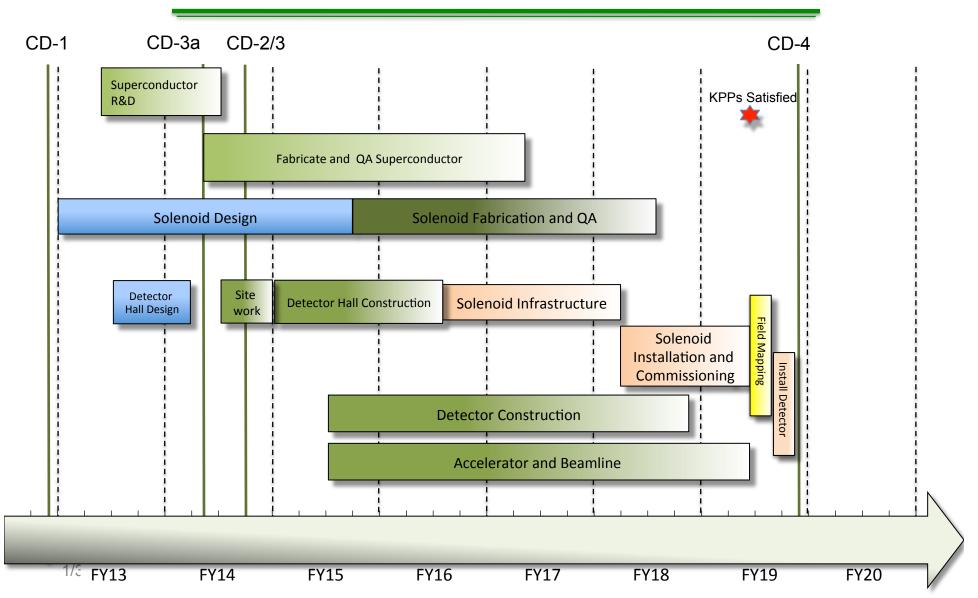
Full G4 and event reconstruction without truth inputs R. Ray - D0 University

1/30/14

### **Project Status**

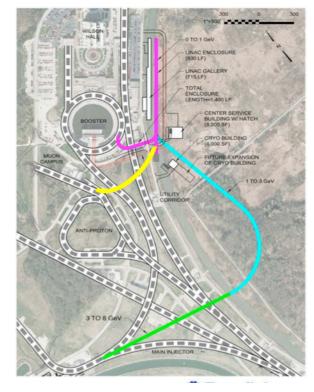
- Mu2e received CD-0 in December 2009
- CD-1 in July 2012
- CD-3a, to authorize long lead solenoid conductor procurement, scheduled for March 2014
- CD-2/3 scheduled for July 2014
- Total Project cost of ~ \$250M
  - \$40M spent thus far.
- Mu2e Conceptual Design Report can be found at
  - arXiv:1211.7019

#### Schedule

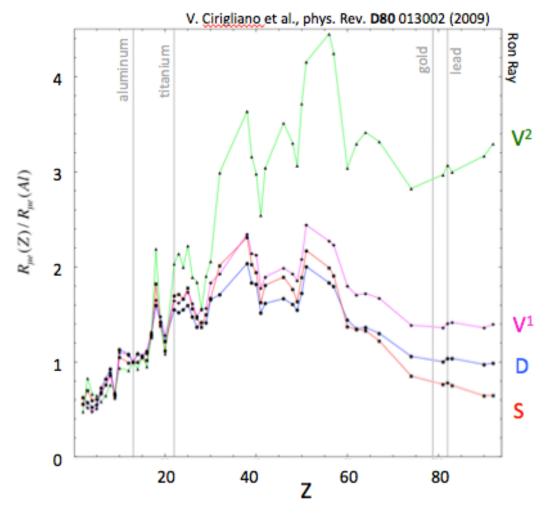


# Mu2e II

- Mu2e uses 8 kW proton beam from Booster
- PIP II provides an upgrade path to x10 more beam power
  - Narrower proton pulses
  - No pbar background
  - Flexible beam structure
  - Run simultaneously with g-2
  - Requires some modest upgrades to Mu2e apparatus to handle higher beam power.
  - Important physics goals regardless of results from first phase of Mu2e.
    - $\,\circ\,$  More beam power for a more sensitive search
    - Flexible time structure of PIP II beam allows access to different stopping target nuclei where model dependent effects vary significantly.



#### Mu2e II – If we see a signal



Greater discrimination at high Z, but shorter muon lifetimes for higher Z nuclei.

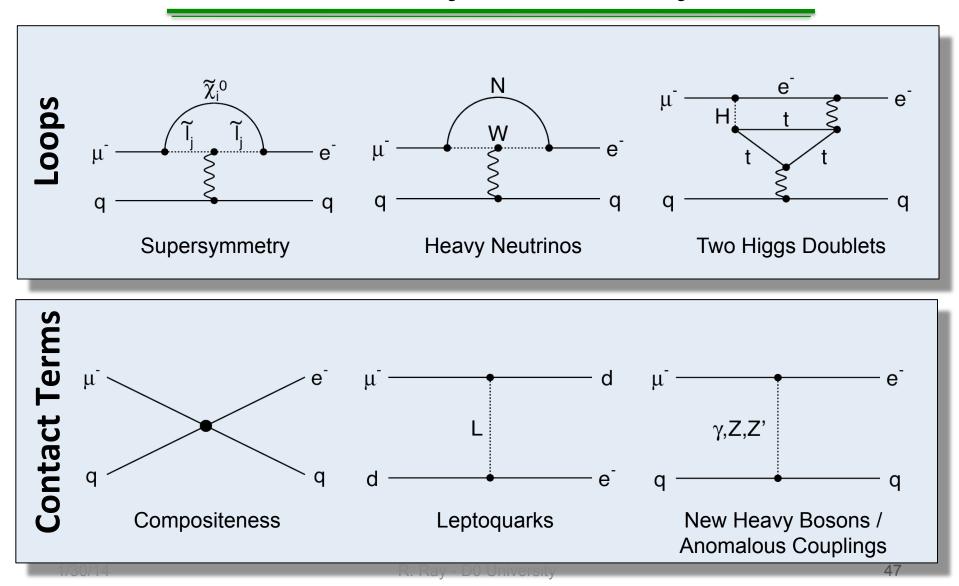
1/30/14

## Summary

- Mu2e will either:
  - Reduce the limit for R<sub>µe</sub> by ~ four orders of magnitude (R<sub>µe</sub> < 6 x 10<sup>-17</sup> @ 90% C.L.)
  - Discover unambiguous proof of Beyond Standard Model physics and
    - Provide important information either complementing LHC results or probing mass scales up to 10<sup>4</sup> TeV
- With upgrades, we could extend the limit by an order of magnitude or study the details of new physics
- There is plenty of interesting work to do and we are still looking for new collaborators...

### **Backup Slides**

#### **Discovery Sensitivity**



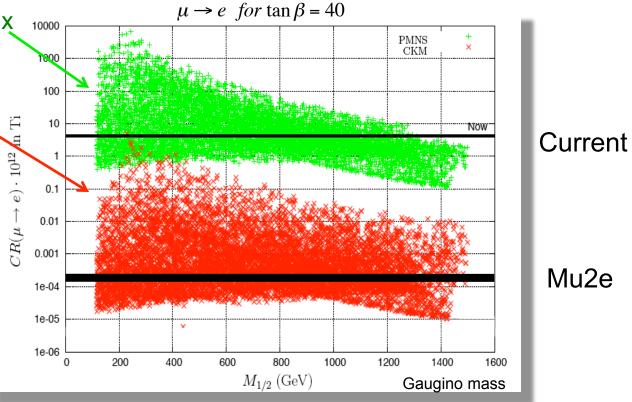
### SUSY and $\mu\text{e}$ Conversion

 $\mu \rightarrow e$  Conversion Rate vs.  $M_{1/2}$ Scan over LHC accessible SUSY-Gut parameter space.

Neutrino-like mixing matrix for Yukawa v's(PNMS)

CKM – like mixing matrix for Yukawa v's (Minimal  $\searrow$ Flavor Violation)

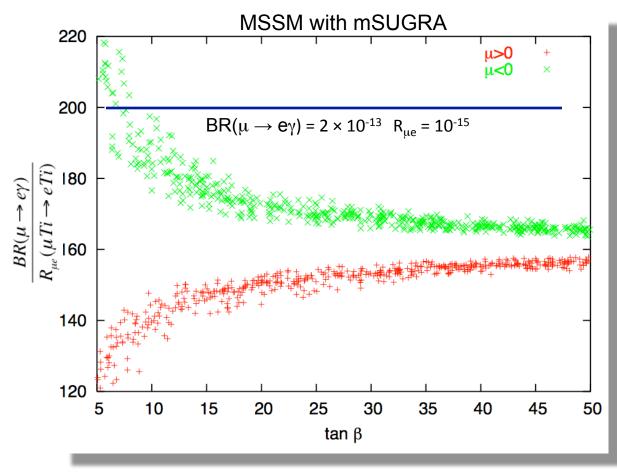
Mu2e can distinguish between PMNS and MFV and help us interpret discoveries at the LHC and provide input to neutrino physics at high mass scales.



Calibbi, Faccia, Masiero, Vempati Phys. Rev. D74:116002 (06)

R. Ray - D0 University

#### Pinning down SUSY with CLFV Measurements

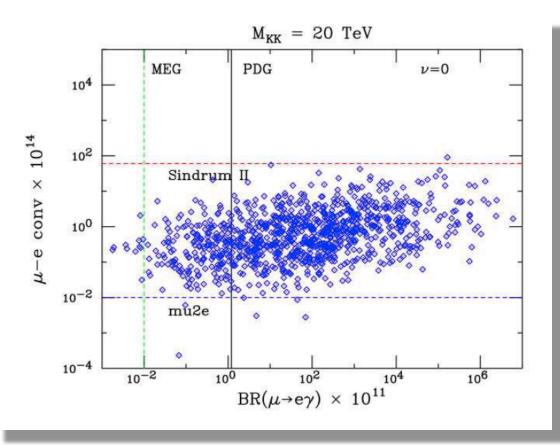


In the MSSM with msugra boundry conditions, and seesaw induced neutrino masses, BR( $\mu \rightarrow e\gamma$ ) and  $\mu$ -e conversion constrain msugra parameters.

C. E. Yaguna, hep-ph/0502014v2

#### Effect of Warped Extra Dimensions on Rates

Muon Conversion vs.  $\mu \rightarrow e\gamma$ 



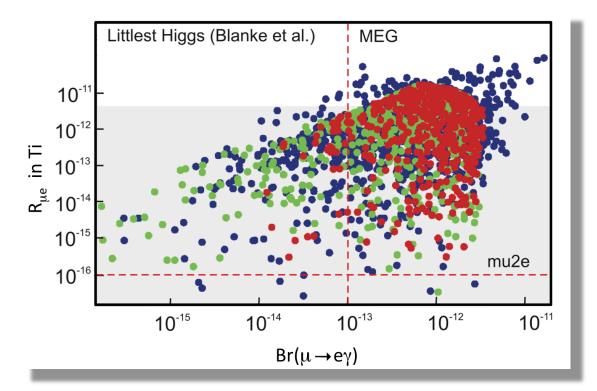
 $\mu$ -e conversion rate as a function of BR( $\mu \rightarrow e\gamma$ ) for RS model with one warped, compact extra dimension in the scenario where the Higgs is allowed to propagate in the bulk.

Mu2e spans nearly the entire parameter space.

K. Agashe, A. E. Blechman and F. Petriello, Phys. Rev. D 74, 053011 (2006).

## Effect of Littlest Higgs Model on Rates

Muon Conversion vs.  $\mu \to e \gamma$ 

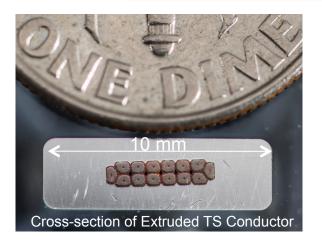


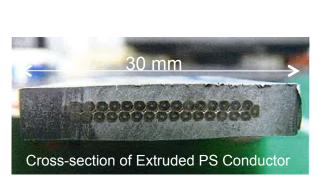
M. Blanke, A. J. Buras, B. Duling, A. Poschenrieder and C. Tarantino, JHEP 0705, 013 (2007).

Littlest Higgs Model. The different colored points refer to different choices for the structure of the mirror-lepton mixing matrix that gives rise to the CLFV effects.

Correlations between the two are distinctly different than for SUSY and could be used do differentiate between SUSY and mirror quarks and leptons at the LHC.

#### **Solenoid Status - Conductor**



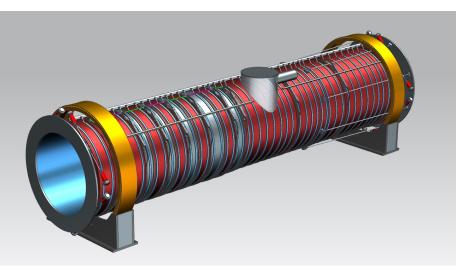


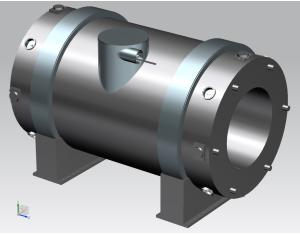


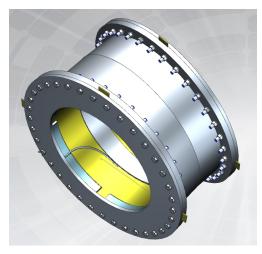
- Magnets are only as good as their conductor
- Solenoid conductor R&D is on the critical path and is nearly complete
- Successful conductor R&D validates decision to develop technically risky conductor in exchange for significantly simpler, cheaper, more robust solenoid designs.

#### Solenoid Status

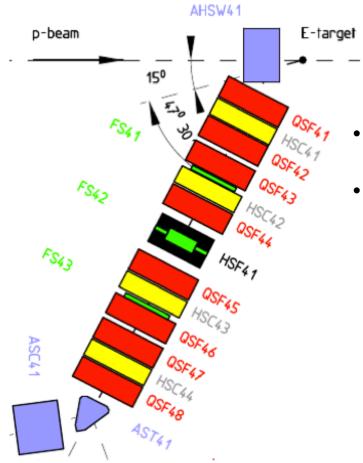
- Solicitation packages for PS/DS sent out this week
- Reference designs completed by Fermilab.
- Vendors complete designs while conductor being fabricated
- Begin solenoid construction in FY15







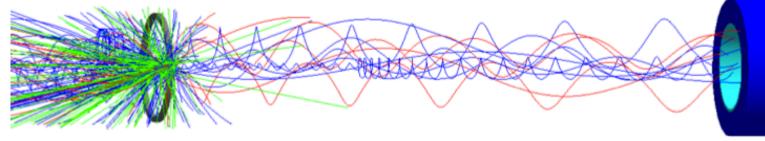
#### The usual way to collect muons



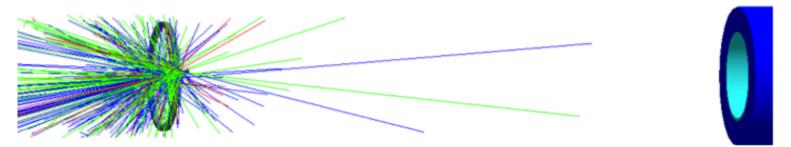
- PSI  $\pi$ E5 beamline used by SINDRUM II
  - 1.3 MW proton beam produces 10<sup>8</sup> μ<sup>+</sup>/s
- Mu2e produces  $10^{10} \mu^{-}$ /sec with an 8 kW proton beam.

#### Mu2e Muon Collection

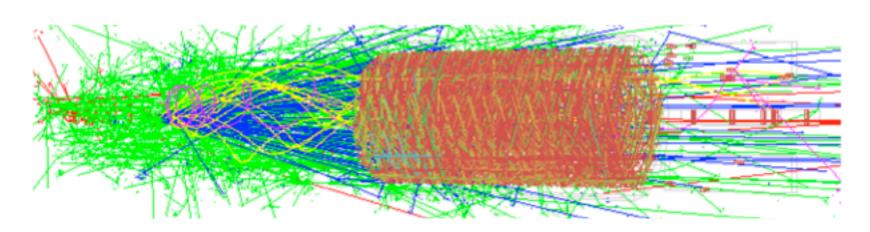
#### Proton target in solenoidal B field Soft pions confined. Collect most of resulting muons.



#### Compare to B = 0

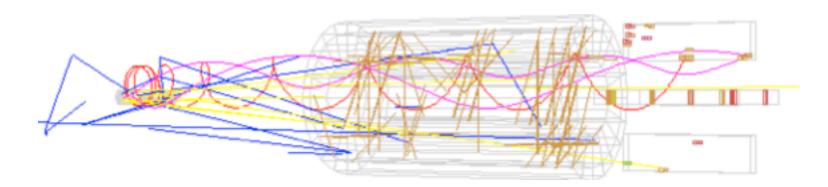


#### Pattern Recognition



Single proton pulse: particles and hits from 500 – 1694 ns

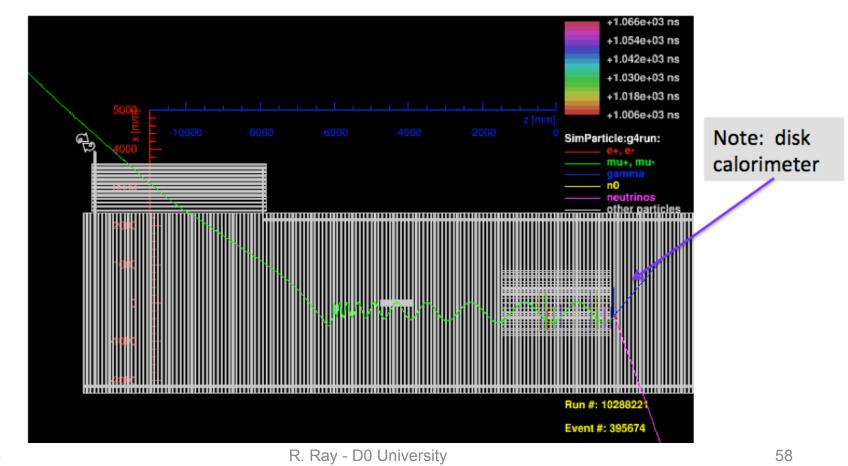
#### Pattern Recognition



Single proton pulse: Particles and hits in ±50 ns window around conversion

#### **Cosmic Ray Veto**

#### A cosmic ray muon vetoed by the calorimeter

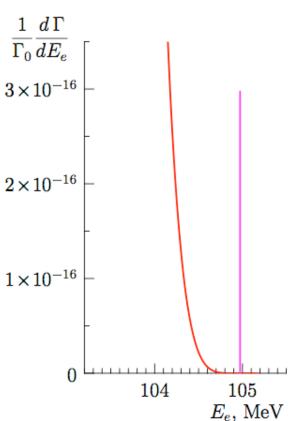


# DIO Endpoint

- Latest spectrum computation by Czarnecki, Tormo, Marciano (2011)
- End point expansion

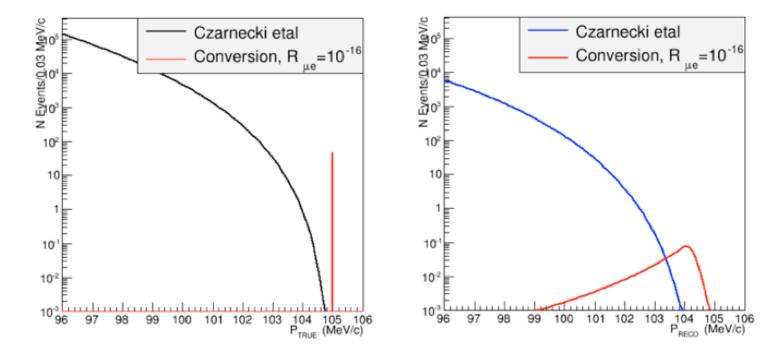
$$\frac{1}{\Gamma_0} \frac{d\Gamma}{dE_e} = B \left( E_\mu - E_e - \frac{E_e^2}{2m_N} \right)^5$$

- Small but steep tail
- DIO electron differs from signal only by its momentum



## DIOs

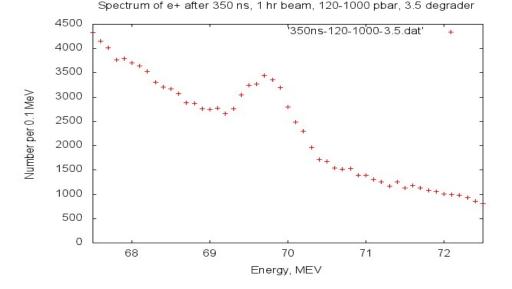
- Momentum resolution is important!
- High side tail pushes DIOs into the signal region

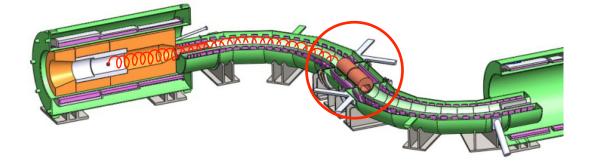


## **Understanding Resolution**

#### Calibration of this device not easy.

- Use stopped  $\pi^+ \rightarrow e^+ \nu$  decay for absolute calibration
  - Monochromatic line at ~70 MeV
- Requires lowering the field, lowering the beam intensity, and rotating the central collimator







#### **Resolution Function**

- Cosmic muons can spallate electrons from the calorimeter that can reflect in the gradient, allowing the particle to be measured twice
- Measuring the same particle twice allows us to extract the tracker momentum resolution function and understand the tails
- Can also measure energy loss, straggling in upstream material

