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## **GENIE Status**

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Simulations for Neutrinos

April 27, 2015



# GENIE Release Roadmap: 2.9/2.10

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- Latest production release is 2.8.6, released 14 November, 2014.
- New release candidate 2.9.0 (beta) will graduate into production release 2.10.0
- Model introduction release.
  - Effective Spectral Functions from A. Bodek, E. Christy, B. Coopersmith (EPJC (2014) 74:3091). (B. Coopersmith and A. Bodek)
  - Very-High Energy extension (5 TeV, working toward PeV) (K. Hoshina)
  - Inclusive Eta production. (J. Liu)
  - New Berger-Sehgal resonant pion production model, tuned with MiniBooNE data by J. Nowak. Berger, Sehgal Phys. Rev. D76, 113004 (2007) & Kuzmin, Lyubushkin, Naumov Mod. Phys. Lett. A19 (2004) 2815 (J. Nowak and S. Dytman)
  - Improved hA FSI model. (S. Dytman and N. Geary)
  - Single Kaon production model by Alam, Simo, Athar, and Vacas (PRD 82, 033001 (2010)). (C. Marshall and M. Nirkko)



# GENIE Release Roadmap: 2.9/2.10

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- Updates to the flux driver (R. Hatcher):
  - Implemented a GFluxDriverFactory, where flux drivers can self-register and be returned by name.
  - Introduced two common flux interfaces GFluxExposureI and GFluxFileConfigI, allowing GNuMIFlux, GSimpleNtpFlux and the external GDk2NuFlux to be used interchangeably. Other flux drivers can start to incorporate these.
  - Renamed gevgen\_numu executable to gevgen\_fnal to emphasize its use in FNAL experiments at other beam lines (e.g. Dune); executable will dynamically pick up GDk2NuFlux flux driver if available (i.e. no longer a build dependence). Flux entries from the input driver will be copied to a branch along side the GHepRecord; flux metadata from all ntuple files will be copied to the output file.
- Event records:
  - Reinstated method in GHepRecord to return the stored KinePhaseSpace\_t, allowing records to be fully recreated from elements stored in non-genie formats (R. Hatcher)
  - Note: We updated the XclsTag object with a new field for strange production for one of the new models in 2.9.0. This could impact the ability of some users to read older GENIE files.

# GENIE Release Roadmap: 2.9/2.10

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- Other changes in 2.9 / 2.10
  - Numerous updates to the validation packages for new models and some improvements to the old ones.
  - Changed the numerical integration routines to use GSL (GNU Scientific Library).
    - Necessary for several new, higher-dimensional models:
      - Single kaon in 2.9/2.10
      - Alvarez-Ruso et al, Coherent Pion in 2.12
  - Results in many small changes in the total cross section splines.
    - One or two wiggles are puzzling, but most are "arbitrary" and well within uncertainties (examples to follow).



# GENIE Release Roadmap: 2.12

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- GENIE 2.12.0 - likely this Summer/Fall
  - QEL Lambda production (J. Poage and H. Gallagher)
  - Berger-Sehgal coherent pion production (PRD 79, 053003 (2009)) (G. Perdue, H. Gallagher, D. Cherdack)
  - Local Fermi Gas & Nieves et al CCQE with RPA (J. Johnston and S. Dytman)
  - Valencia Model Meson Exchange-Currents (J. Schwer and R. Gran)
  - Alvarez-Ruso et al microscopic coherent pion production (PRC 75, 055501 (2007) and PRC 76, 068501 (2007)) (S. Dennis and S. Boyd)
  - Oset FSI model (T. Golan)
  - Kaon FSI (F. de Maria Blaszczyk, S. Dytman)
  - Z expansion of QEL form factor (Hill et al, PRD 84, 073006) (A. Meyer)
  - Benhar Spectral Functions (C. Mariani, M. Jen, and A. Furmanski)
    - Ambitious to get it all... (and I may have forgotten something)
- GENIE 3.0 - likely early 2016
  - New default physics tune incorporating all of these models and recent neutrino-nucleus cross section data, plus many tuning and data comparison tools.



# Tuning GENIE 3.0

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- Our long-term goal is just getting underway.
  - Simultaneous fits to multiple neutrino data sets (e.g., pion production in MiniBooNE, MINERvA, and T2K).
  - These sorts of fits always have lots of trouble! Understanding correlated uncertainties between measurements, Peele's Pertinent Puzzle (normalization uncertainties common to an entire data set can cause bias in a least-squares minimization), etc.
  - Significant software and production infrastructure required (we would like to run large pieces of the required simulation on the Open Science Grid).
    - Substantial recent progress here, especially in the fitting framework, but we're also finally beginning to make some inroads into automated validation on the Grid.

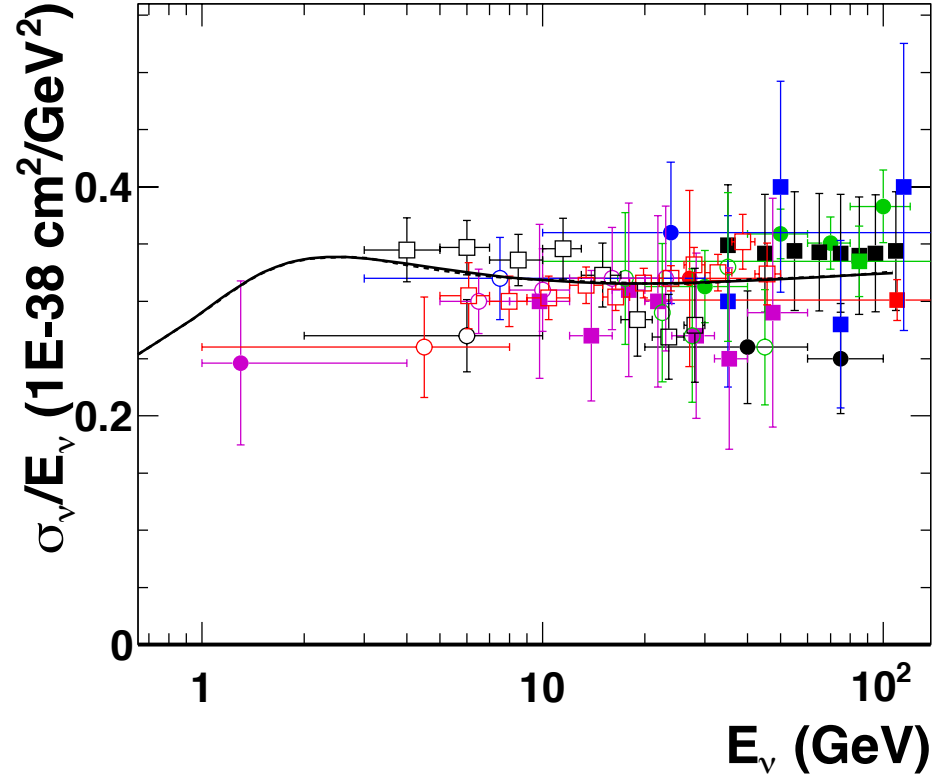
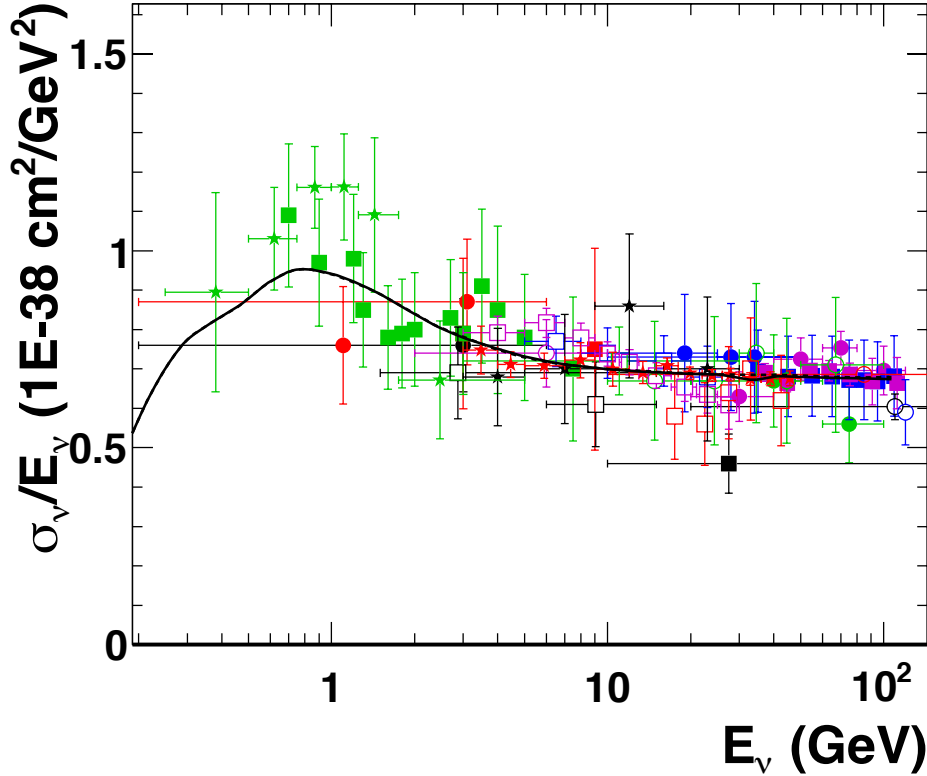
## Other Updates

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- New project incubators page on HepForge:
  - <https://genie.hepforge.org/load.php?include=incubator>
  - Some links on the page are not publicly accessible.
  - Main point of the page is to provide a place the community can clearly see what we're working on.
- Plan to release (at least) a citable arXiv e-print for future production releases.
  - We won't be seeking peer-review for these e-prints.
  - The author list will be comprised of all contributors to the release.
  - Find a bug and you can be a co-author on the 2.10.0 eprint!

$\nu_\mu$  CC inclusive

$\bar{\nu}_\mu$  CC inclusive



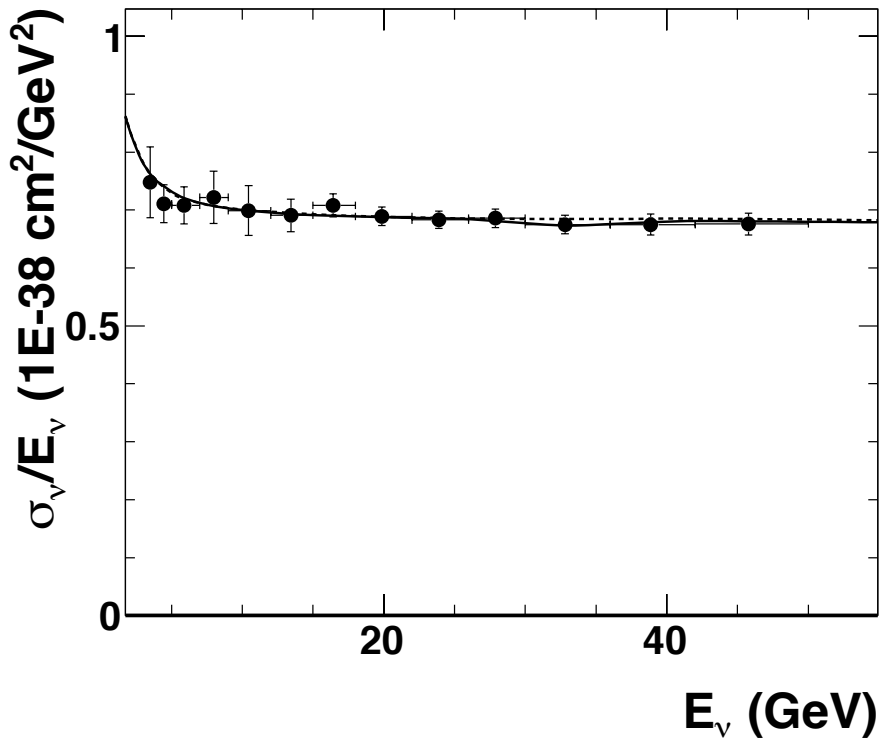
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- CCFR,C,0 [Bosetti et al., Phys.Lett.B70:295 (1977)]
- CCFR,C,1 [Colley et al., Zeit.Phys.C2:187 (1979)]
- CCFR,C,2 [Bosetti et al., Phys.Lett.B110:167 (1982)]
- CCFR,C,3 [Parker et al., Nucl.Phys.B232:1 (1984)]
- CHARM\_7E,0 [Baltay et al., Phys.Rev.Lett.44:916 (1980)]
- CHARM\_7E,4 [Baker et al., Phys.Rev.D25:617 (1982)]
- CHARM\_7E,5 [Seligman et al., Nevis Report 292 (1982)]
- CCFR,F,0 [MacFarlane et al., Zeit.Phys.C26:1 (1984)]
- CHARM\_9 [Jonker et al., Phys.Lett.B99:265 (1981)]
- CHARM\_9 [Allaby et al., Zeit.Phys.C38:403 (1983)]
- FNAL\_15FT,1 [Kitagaki et al., Phys.Rev.Lett.49:38 (1982)]
- FNAL\_15FT,2 [Baker et al., Phys.Rev.Lett.51:735 (1983)]
- Gargamelle,0 [Eichten et al., Phys.Lett.B46:274 (1973)]
- Gargamelle,1 [Campanella et al., Phys.Lett.B84:181 (1979)]
- Gargamelle,12 [Morfin et al., Phys.Lett.B104:235 (1981)]
- Gargamelle,13 [Morfin et al., Phys.Lett.B104:235 (1981)]
- IHEP\_ITEP,0 [Asratyan et al., Phys.Lett.B76:239 (1978)]
- IHEP\_ITEP,1 [Asratyan et al., Phys.Lett.B76:239 (1978)]
- IHEP\_ITEP,3 [Vovenko et al., Sov.J.Nucl.Phys.30:528 (1979)]
- IHEP\_JINR,1 [Anikeev et al., Zeit.Phys.C70:39 (1986)]
- MINOS,0 [Adamson et al., Phys.Rev.D81:37 (2010)]
- MINOS,1 [Adamson et al., Phys.Rev.D81:072002 (2010)]
- SciBooNE,0 [Nakajima et al., Phys.Rev.D83:012005 (2011)]
- v290
- v286

- BEBC,1 [Bosetti et al., Phys.Lett.B70:273 (1977)]
- BEBC,3 [Colley et al., Zeit.Phys.C2:187 (1979)]
- BEBC,5 [Bosetti et al., Phys.Lett.B110:167 (1982)]
- BEBC,7 [Parker et al., Nucl.Phys.B232:1 (1984)]
- BNL\_7FT,1 [Fanourakis et al., Phys.Rev.D21:562 (1980)]
- CCFR,3 [Seligman et al., Nevis Report 292 (1982)]
- CHARM\_1 [Jonker et al., Phys.Lett.B99:265 (1981)]
- CHARM\_5 [Allaby et al., Zeit.Phys.C38:403 (1983)]
- FNAL\_15FT,4 [Taylor et al., Phys.Rev.Lett.51:739 (1983)]
- FNAL\_15FT,5 [Asratyan et al., Phys.Lett.B137:122 (1984)]
- Gargamelle,1 [Eichten et al., Phys.Lett.B46:274 (1973)]
- Gargamelle,11 [Erriquez et al., Phys.Lett.B80:309 (1979)]
- Gargamelle,13 [Morfin et al., Phys.Lett.B104:235 (1981)]
- IHEP\_ITEP,1 [Asratyan et al., Phys.Lett.B76:239 (1978)]
- IHEP\_ITEP,3 [Vovenko et al., Sov.J.Nucl.Phys.30:528 (1979)]
- IHEP\_JINR,1 [Anikeev et al., Zeit.Phys.C70:39 (1986)]
- MINOS,1 [Adamson et al., Phys.Rev.D81:072002 (2010)]
- v290
- v286



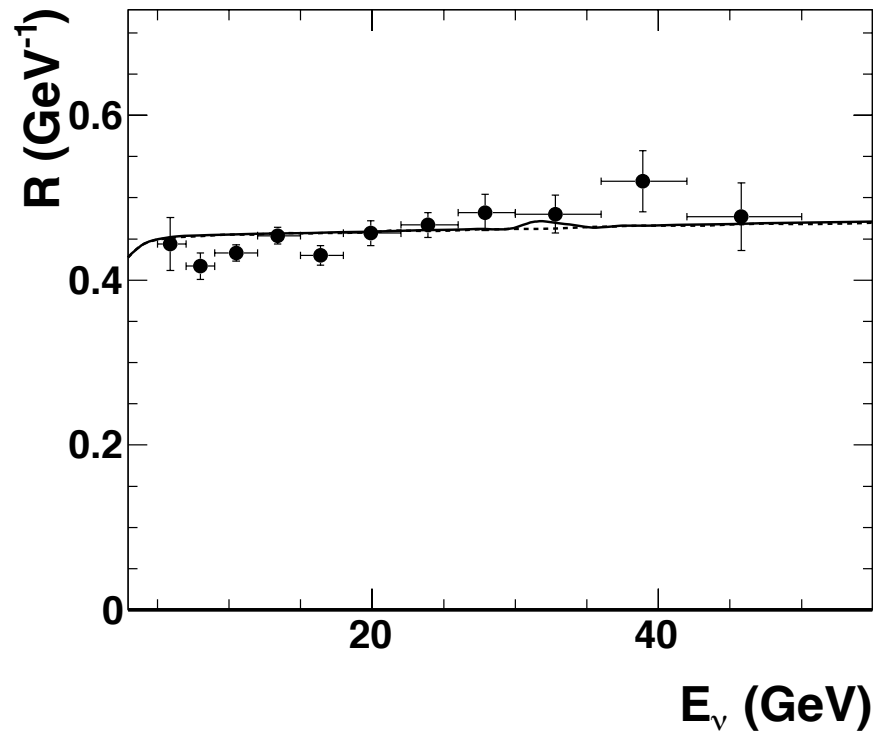
# 2.9.0 vs. 2.8.6

$\nu_\mu$  CC inclusive, MINOS data only



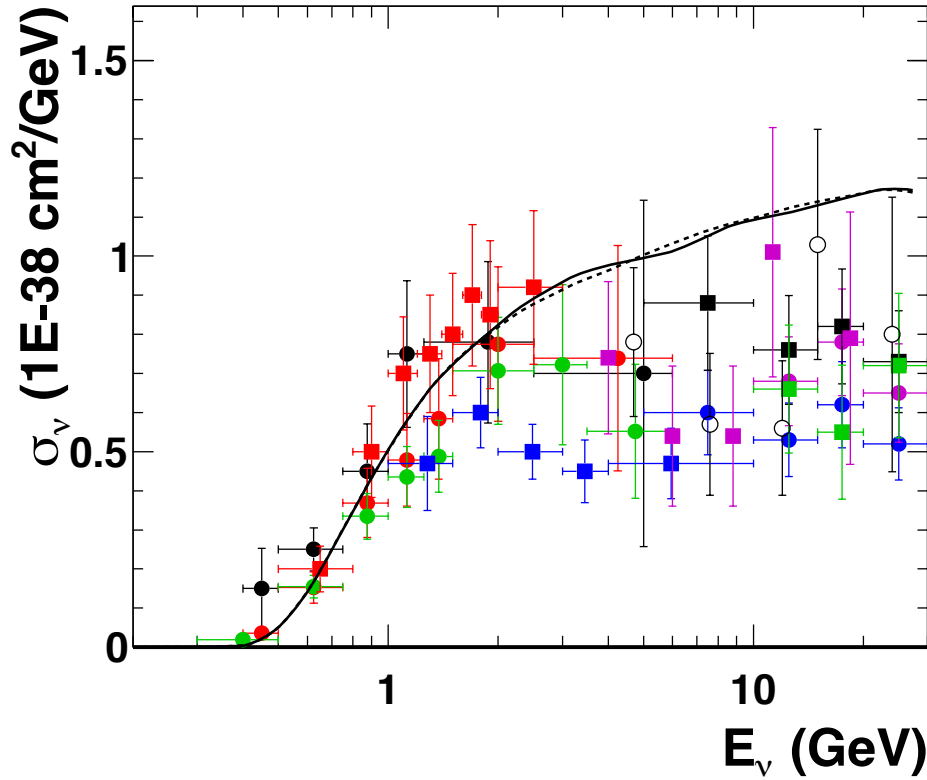
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 — v290  
 - - - v286

$\bar{\nu}_\mu$  CC inclusive /  $\nu_\mu$  CC inclusive, MINOS data only

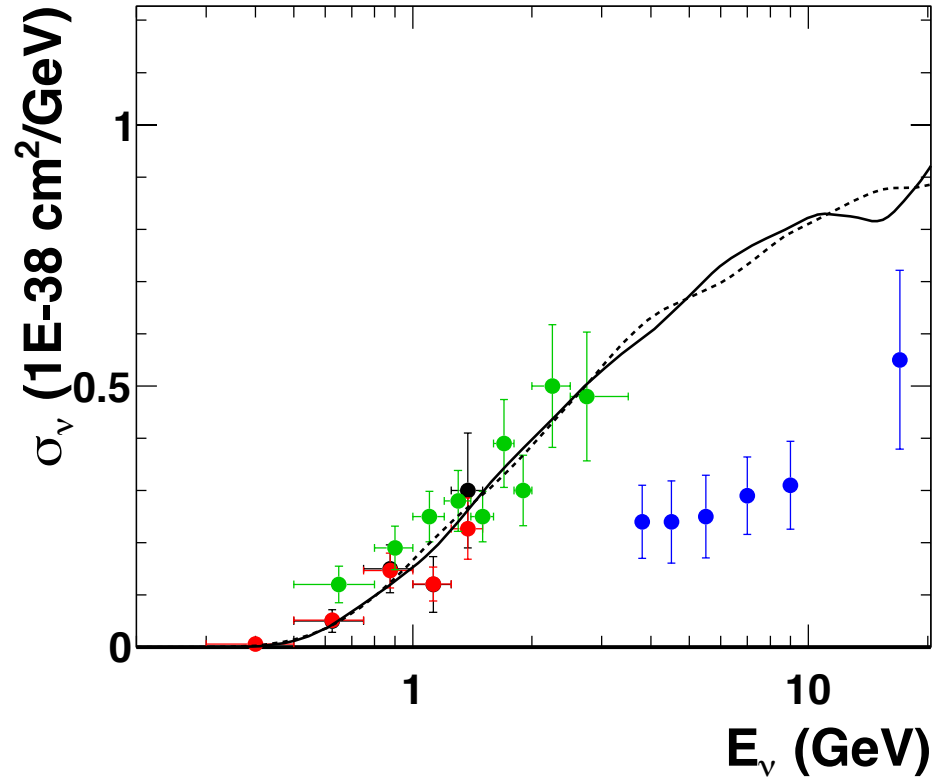


● MINOS,2 [Adamson et al., Phys.Rev.D81:072002 (2010)]  
 — v290  
 - - - v286

$\nu_\mu$  CC1 $\pi^+$  ( $\nu_\mu$  p  $\rightarrow$   $\mu^-$  p  $\pi^+$ )



$\nu_\mu$  CC1 $\pi^0$  ( $\nu_\mu$  n  $\rightarrow$   $\mu^-$  p  $\pi^0$ )



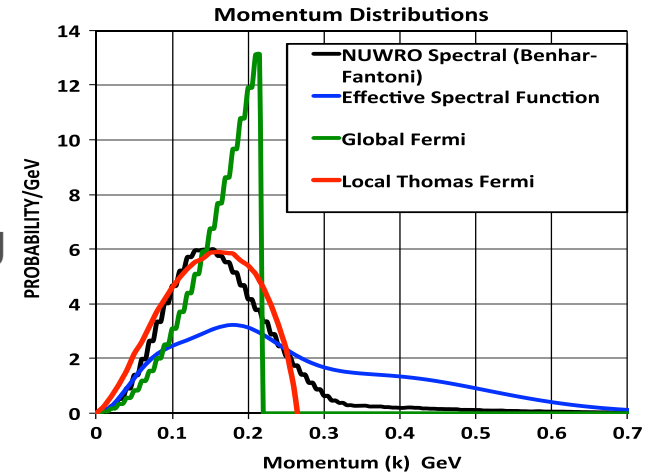
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- BEBC,4 [Allen et al., Nucl.Phys.B176:269 (1980) ]
- BEBC,9 [Allen et al., Nucl.Phys.B264:221 (1986) ]
- BEBC,13 [Allasia et al., Nucl.Phys.B343:285 (1990) ]
- BNL\_7FT,5 [Kitagaki et al., Phys.Rev.D34:2554 (1986)]
- FNAL\_15FT,0 [Bell et al., Phys.Rev.Lett.41:1008 (1978)]
- Gargamelle,4 [Lerche et al., Phys.Lett.B78:510 (1978) ]
- SKAT,4 [Ammosov et al., Sov.J.Nucl.Phys.50:67 (1988)]
- SKAT,5 [Grabosch et al., Zeit.Phys.C41:527 (1988)]
- v290
- - - v286

- ANL\_12FT,6 [Barish et al., Phys.Rev.D19:2521 (1979)]
- ANL\_12FT,9 [Radecky et al., Phys.Rev.D25:1161 (1982)]
- BNL\_7FT,6 [Kitagaki et al., Phys.Rev.D34:2554 (1986)]
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- v290
- - - v286

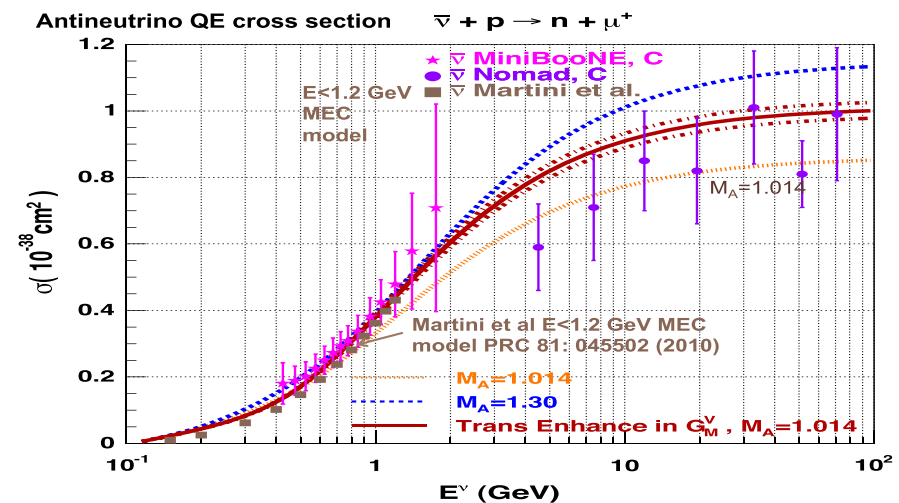
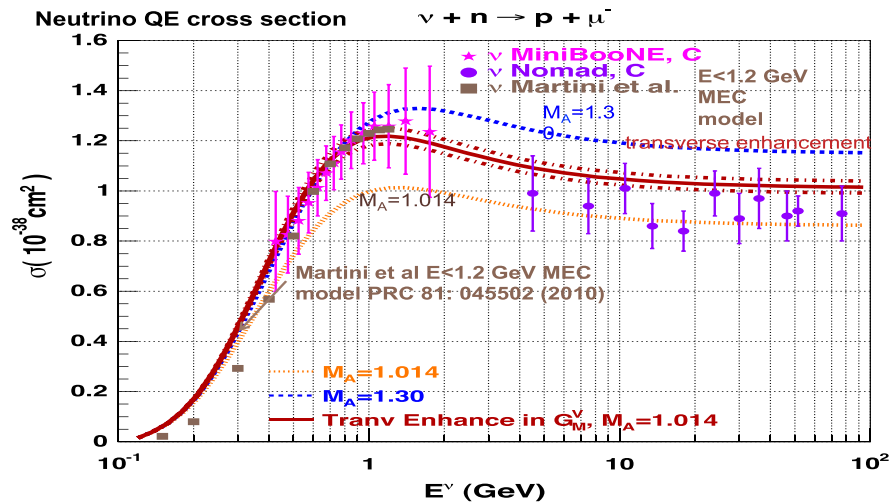


## 2.9/2.10: Effective Spectral Functions

- The Effective Spectral Function model combines a superscaling formalism together with hadronic energy sharing prescription to form a complete QE model.
  - An eight parameter spectral function is fit to the superscaling function extracted from electron scattering data (plus two parameters for binding energy and 2p2h fraction).
- Implemented by B. Coopersmith (also implemented Transverse Enhancement Model)

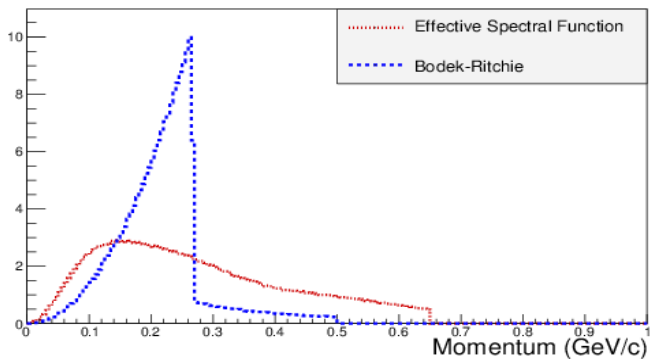


Bodek, Christy, Coopersmith EPJ C (2014) 74:3091

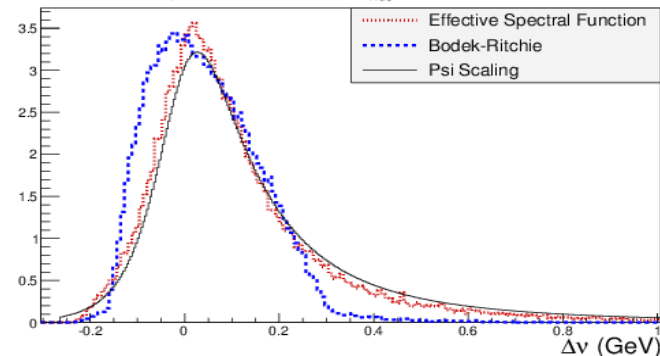


# Ar-40 Results

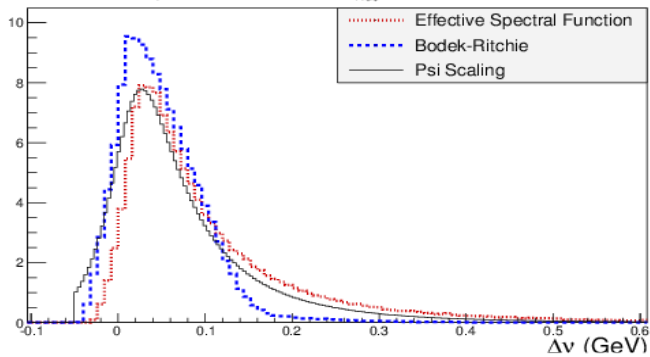
GENIE- Initial Nucleon Momentum Distributions, Ar-40



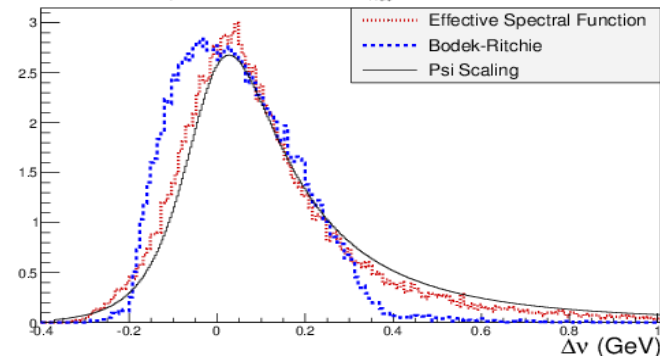
$\nu\text{-}Q^2/2M_p$  for  $E_\nu^{\text{True}} = 10.0$  GeV,  $Q_{\text{True}}^2 = 0.50 \pm .05$  GeV<sup>2</sup>, Ar-40



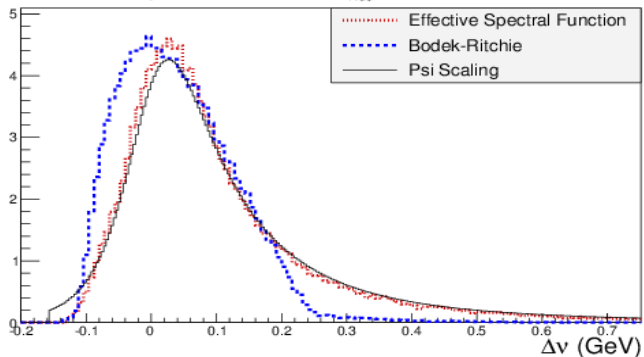
$\nu\text{-}Q^2/2M_p$  for  $E_\nu^{\text{True}} = 10.0$  GeV,  $Q_{\text{True}}^2 = 0.10 \pm .05$  GeV<sup>2</sup>, Ar-40



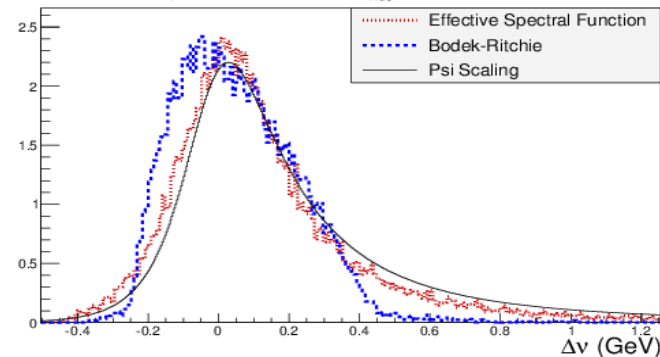
$\nu\text{-}Q^2/2M_p$  for  $E_\nu^{\text{True}} = 10.0$  GeV,  $Q_{\text{True}}^2 = 0.70 \pm .05$  GeV<sup>2</sup>, Ar-40



$\nu\text{-}Q^2/2M_p$  for  $E_\nu^{\text{True}} = 10.0$  GeV,  $Q_{\text{True}}^2 = 0.30 \pm .05$  GeV<sup>2</sup>, Ar-40



$\nu\text{-}Q^2/2M_p$  for  $E_\nu^{\text{True}} = 10.0$  GeV,  $Q_{\text{True}}^2 = 1.00 \pm .05$  GeV<sup>2</sup>, Ar-40



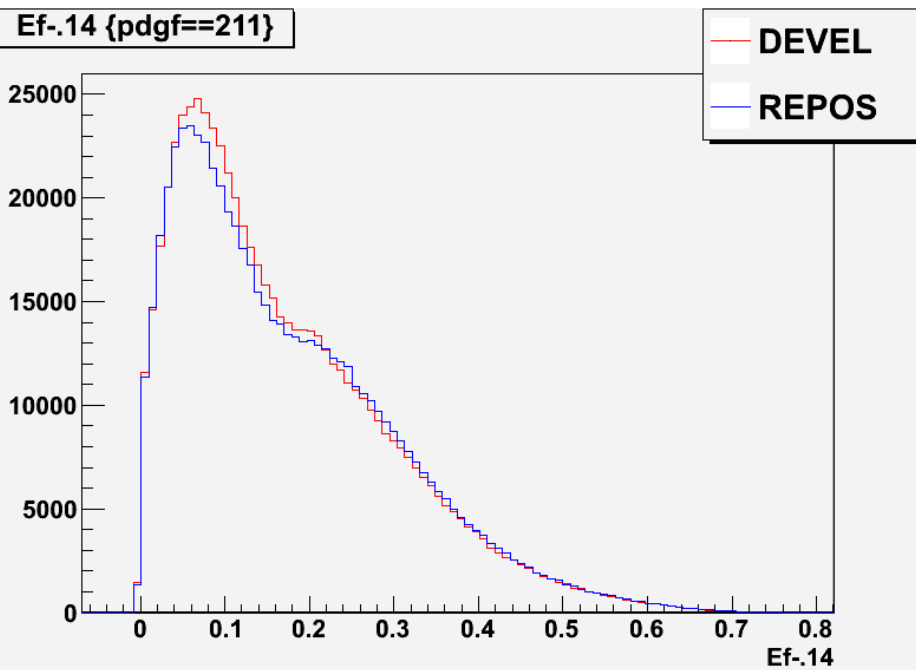
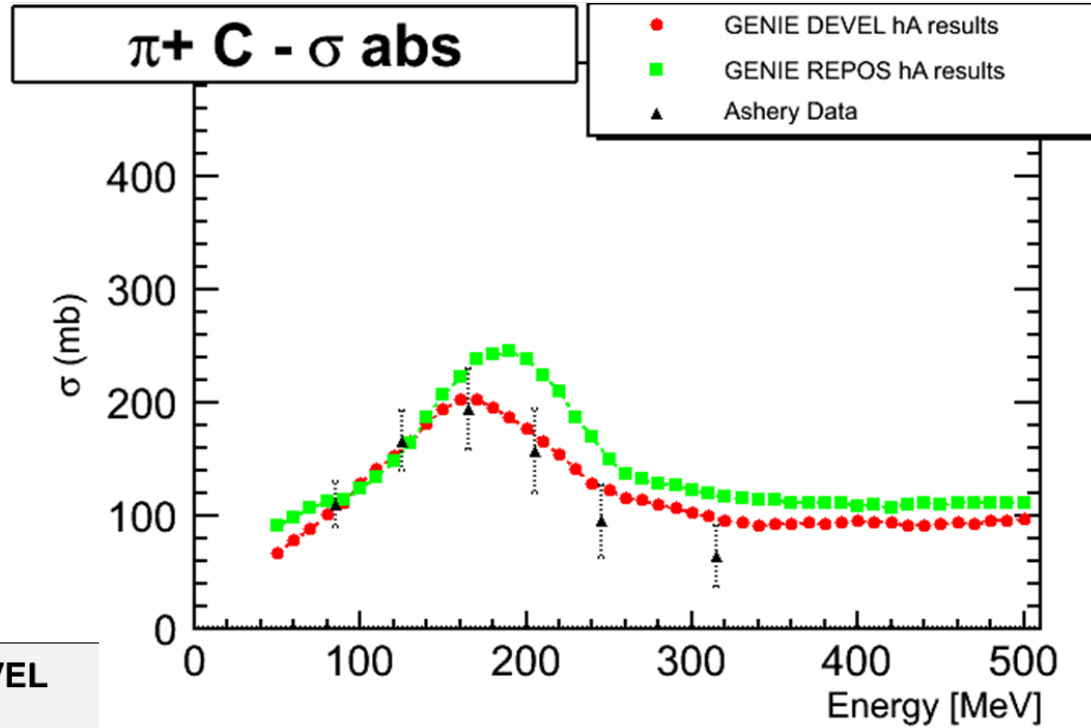
Brian Coopersmith- University of Rochester



## 2.9/2.10: Updated hA Model (FSI Model)

- Recall that in hA mode we parameterize a cascade with one effective interaction.
  - Easily re-weightable.
  - Good agreement with data.
- Previously we used data on iron and  $f(A)$  scaling.
  - Now including Li-7, C-12, Al-27, Fe-56, Nb-93, Bi-209
- Previously, all cross sections for different "fates" had an  $A^{2/3}$  dependence - but this doesn't agree with data.
  - Now absorption scales as  $A^{2/3+0.18}$ , charge exchange as  $A^{2/3}$ , elastic as  $A^{2/3 + 0.25}$ , inelastic like  $A^{2/3}$ , pion production as  $A^{2/3}$
  - Total cross section scales as  $A^{2/3}$  (used to convert a fate cross section to a fraction)
- Implemented by N. Geary and S. Dytman

Better A dependent performance.

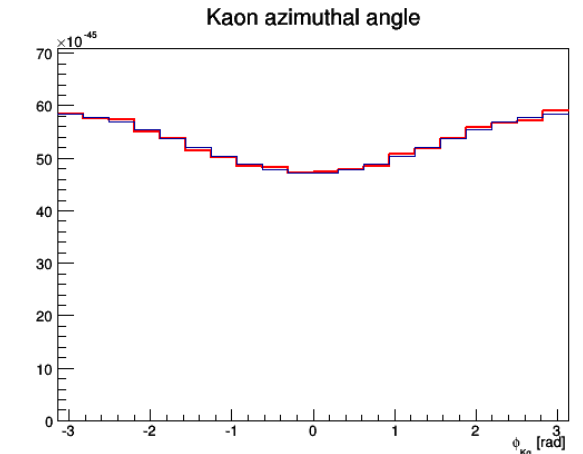
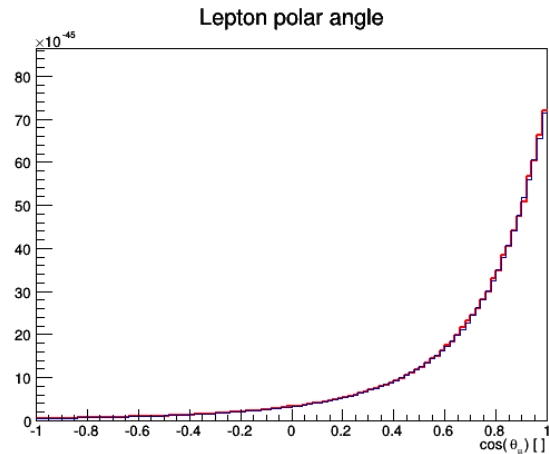
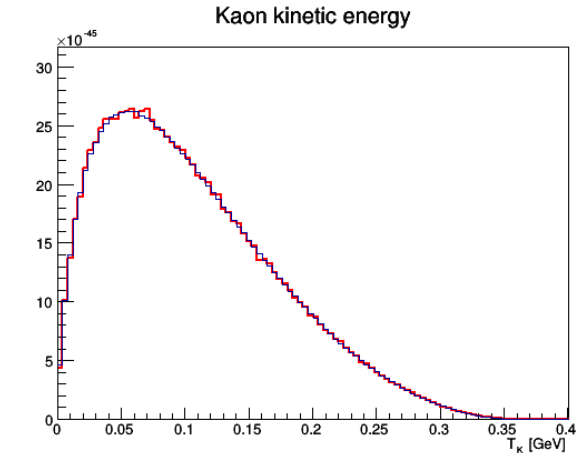
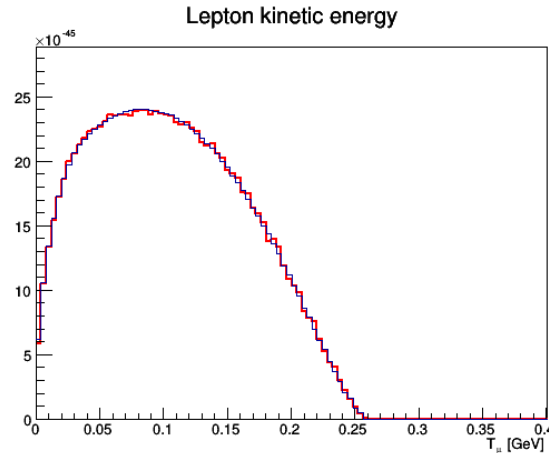


Shift to lower energy due to increased inelastics, more pions due to decreased (effectively) absorption.



## 2.9/2.10: Single Kaon Production

- Alam, Simo, Athar, and Vacas (PRD 82, 033001 (2010)).
- Implemented by C. Marshall and M. Nirkko

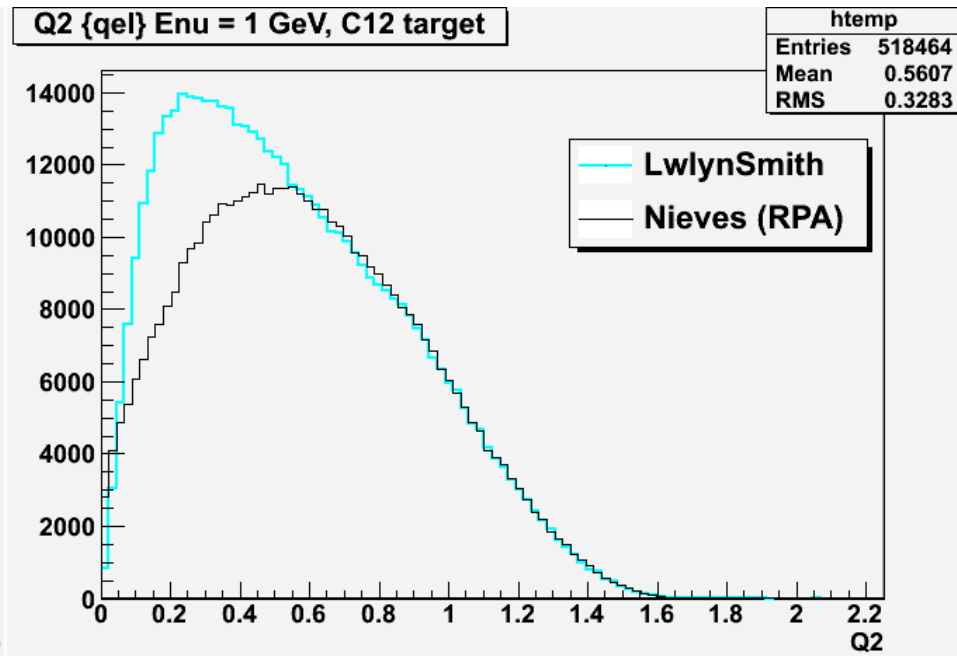
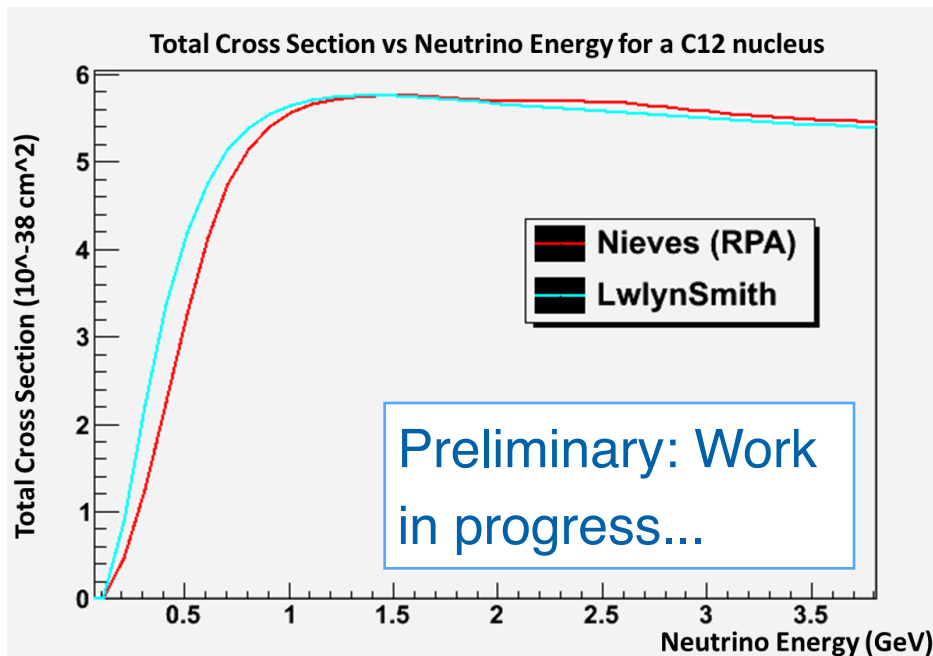


Blue histograms are from the a 4D integral based on the original paper and the red are from the GENIE implementation.



## 2.12: "Nieves CCQE with RPA"

- Potential replacement for current Llewellyn-Smith QEL
- Add RPA (long-range correlation)
- Use local Fermi Gas model for the nucleus (still under construction)
- Implemented by J. Johnston and S. Dytman

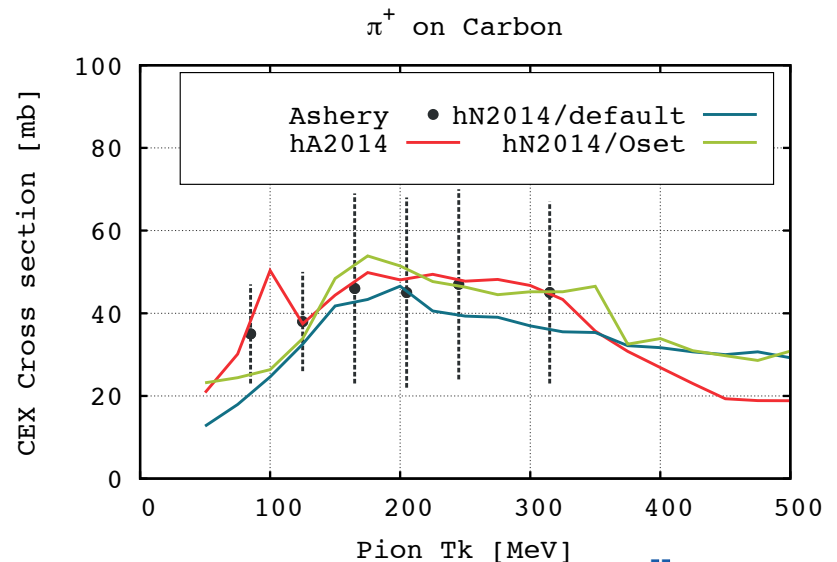
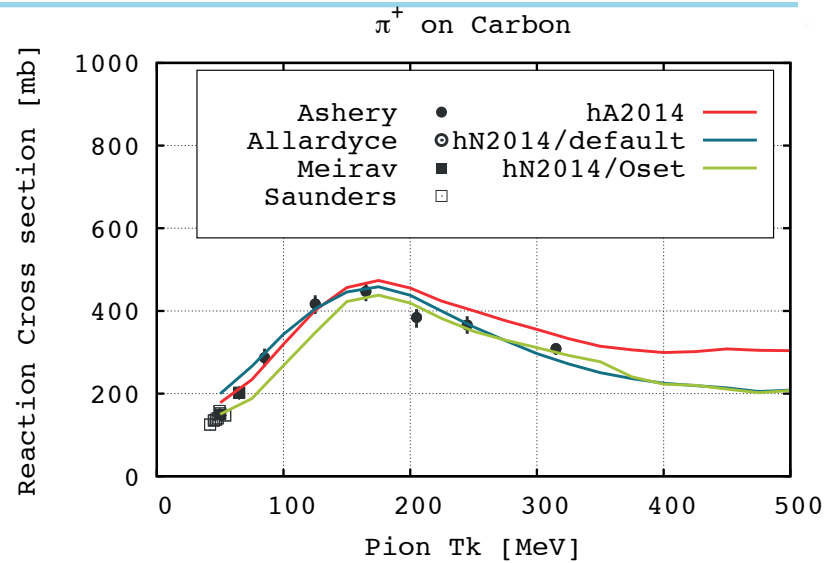
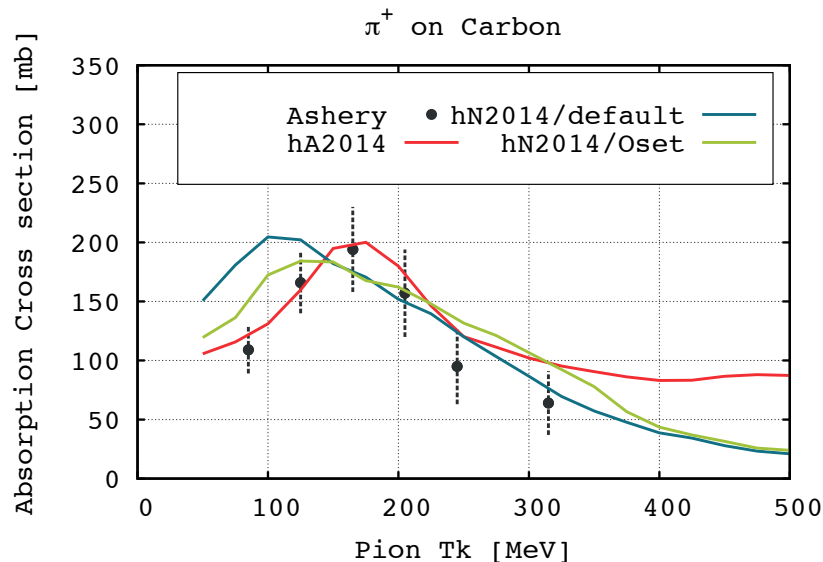






# 2.12: Oset FSI

- E. Oset et al, Nucl. Phys. A484 (1998) 557-592
- E. Oset et al, Nucl. Phys. A468 (1987) 631
- Nuclear effects are introduced as modifications of the  $\Delta$  width.
  - Plus many "hidden tricks" being worked through...
- GENIE hN pion model will be similar to NEUT and NuWro.
- Implemented by T. Golan.



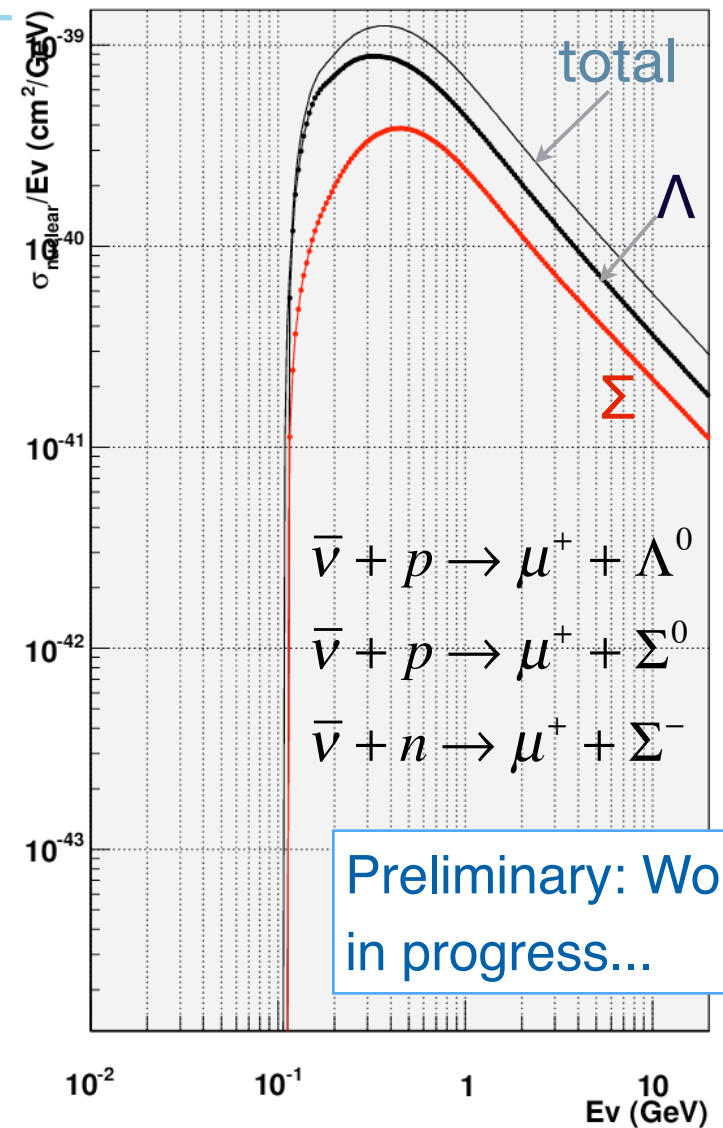


## 2.12: Quasielastic Hyperon Production

- Cabibo and Chilton, v136, N6B (1965)
- Pais, Ann. Phys. 63, 361 (1971)
- If we assume SU(3), we may write the  $p \rightarrow \Lambda$  transition in terms of the form factors for the  $n \rightarrow p$  transition.
- Implemented by J. Poage, E. Morrissey, H. Gallagher

$$\sigma_{\Delta S=1} \approx \tan^2 \theta_C \sigma_{QEL} = 0.05 \sigma_{QEL}$$

$$\sigma_{\Lambda} / \sigma_{QEL} (10 \text{ GeV}) = 0.038$$





# Conclusions

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- 2.9/2.10 is available.
- 2.12 is under construction with a loose "Summer" timetable.
- Work on 3.0 is also underway (building tuning and fitting infrastructure).
- We always welcome interested parties to collaborate / contribute!