Electrophobic Scalar Boson and Muonic Problems

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Muonic Problem

- Proton Radius Puzzle
- Lamb Shift in $2S_{1/2} 2P_{3/2}$ in μ H [1,2] $r_p = 0.84087(39)$ fm
- CODATA 2014 from L.S. in H and e -p scattering [3] $r_p = 0.8751(61)$ fm
- More than 5σ away
- *e*–*p* scattering? Nuclear theory? New physics?

[1] R. Pohl *et al.*, Nature (London) **466**, 213 (2010).

- [2] A. Antognini *et al.*, Science **339**, 417 (2013).
- [3] P. J. Mohr, D. B. Newell, and B. N. Taylor, arXiv:1507.07956.

Muonic Problem

• Muon Anomalous Magnetic Moment

$$a_{\mu} = \frac{(g-2)_{\mu}}{2}$$

- The measurement at BNL [1] differs from SM prediction [2,3] by more than 3σ away.
- $\Delta a_{\mu} = 287(80) \times 10^{-11}$

[1] T. Blum, A. Denig, I. Logashenko, E. de Rafael, B. Lee Roberts, T. Teubner, and G. Venanzoni, arXiv:1311.2198.
[2] M. Davier, A. Hoecker, B. Malaescu, and Z. Zhang, Eur. Phys. J. C **71**, 1515 (2011); **72**, 1874(E) (2012).
[3] K. Hagiwara, R. Liao, A. D. Martin, D. Nomura, and T. Teubner, J. Phys. G **38**, 085003 (2011).

New Physics

- A Scalar Boson, ϕ , can solve both problems [1,2].
- Assume Yukawa Interaction $\mathcal{L} \supset e\epsilon_f \phi \overline{\psi}_f \psi_f$, where $\epsilon_f = \frac{g_f}{e}$, and $f = e, \mu, p, n$, etc.
- The classical potential between f_1 and f_2 by exchanging a scalar boson is

$$V_{\phi}(r) = -\epsilon_1 \epsilon_2 \alpha \frac{e^{-m_{\phi}r}}{r}.$$

D. Tucker-Smith and I. Yavin, Phys. Rev. D 83, 101702 (2011).
 E. Izaguirre, G. Krnjaic, and M. Pospelov, Phys. Lett. B 740, 61 (2015).

New Physics

- Make no other *a priori* assumptions. ($\epsilon_{\mu} = \epsilon_{p}$ [1], $\epsilon_{n} = 0$ [1,2], mass-weighted [2], etc.)
- $\epsilon_{\mu}\epsilon_{p} > 0$ (using Lamb Shift of μ H)
- Without lose of generality, we set $\epsilon_{\mu} > 0$ and $\epsilon_{p} > 0$.
- ϵ_e and ϵ_n are allowed to have either sign.

D. Tucker-Smith and I. Yavin, Phys. Rev. D 83, 101702 (2011).
 E. Izaguirre, G. Krnjaic, and M. Pospelov, Phys. Lett. B 740, 61 (2015).

Muon and Nucleon Observables

- The Lamb shift using laser spectroscopy in muonic hydrogen, deuterium, and helium-4
- $(g 2)_{\mu}$
- Low energy scattering of neutron on ²⁰⁸Pb
- The NN scattering length
- The binding energy per nucleon in nuclear matter
- The binding energy difference of ${}^{3}\text{He}$ and ${}^{3}\text{H}$









Electron Observables

- Electron Beam dump experiments, E137, E141, and Orsay.
- $(g-2)_e$
- Resonance in Bhabha Scattering
- The Lamb shift using laser spectroscopy in electronic hydrogen



Coupling to Electron ϵ_{ρ}

- A: APEX, HPS, DarkLight, VEPP-3, MAMI, MESA, etc.
- B: de-excitation of ${}^{16}O^*$ [1].
- A1 at MAMI and BaBar, Astronomical observables, Kaon decay, etc.
- mass-weighted $\epsilon_l \propto m_l^n$
- electrophobic



 ϵ_{e}



 m_{ϕ} (MeV)





μH Lamb Shift

• $\delta E_L^{lN} = -\frac{\alpha}{2a_l} \epsilon_l \epsilon_N f(a_{lN} m_{\phi})$ (for lepton-nucleus) $f(x) = \frac{x^2}{(1+x)^4}$ a_{lN} is the Bohr radius • $\delta E_L^{\mu H} = -0.307(56)$ meV [1-4]

[1] R. Pohl *et al.*, Nature (London) **466**, 213 (2010).
 [2] A. Antognini *et al.*, Science **339**, 417 (2013).
 [3] P. J. Mohr, D. B. Newell, and B. N. Taylor, arXiv:1507.07956.
 [4] A. Antognini et al., EPJ Web Conf. **113**, 01006 (2016).



μD and $\mu^4 He^+$ Lamb Shift

•
$$\delta E_L^{lN} = -\frac{\alpha}{2a_l} \epsilon_l \epsilon_N f(a_{lN} m_{\phi})$$
 (for lepton-nucleus
 $f(x) = \frac{x^2}{(1+x)^4}$, a_{lN} is the Bohr radius
• $\delta E_L^{\mu D} = -0.409(69)$ meV [1,2]

•
$$\delta E_L^{\mu^4 \text{He}^+} = -1.4(1.5) \text{ meV} [3,4]$$

[2] J. J. Krauth, M. Diepold, B. Franke, A. Antognini, F. Kottmann, and R. Pohl, Ann. Phys. (Amsterdam) 366, 168 (2016).

- [3] A. Antognini et al., EPJ Web Conf. **113**, 01006 (2016).
- [4] I. Sick, Phys. Rev. C 90, 064002 (2014).

^[1] R. Pohl et al., Science 353, 669 (2016).

Low Energy Neutron Experiments

• Measuring total cross section of Scattering of neutron on ²⁰⁸Pb

• In [1], it is assumed
$$g_p = g_n = g_N$$
,
$$\frac{g_N^2}{4\pi} \lesssim 8 \times 10^{-23} \left(\frac{m_\phi}{\mathrm{eV}}\right)^2.$$

• Using the replacement

$$\frac{g_N^2}{4\pi} \to \alpha \epsilon_n \left(\frac{A-Z}{A} \epsilon_n + \frac{Z}{A} \epsilon_p \right)$$

[1] H. Leeb and J. Schmiedmayer, Phys. Rev. Lett. 68, 1472 (1992).

NN scattering length

•
$$\Delta a = \overline{a} - a_{np}, \overline{a} = \frac{a_{pp} + a_{nn}}{2}$$

• $\Delta a_{exp} = 5.64(60) \text{ fm [1]}, \Delta a_{th} = 5.6(5) \text{ fm [2]}$
• $\Delta a_{\phi} = \overline{a}a_{np}M \int_{0}^{\infty} \Delta V \overline{u}u_{np}dr$
 $M = \frac{m_p + m_n}{2}; \Delta V = -\frac{1}{2}(\epsilon_p - \epsilon_n)^2 e^{-m_{\phi}r};$
 $u(r) \text{ is the zero energy } {}^{1}S_0 \text{ wave function},$
 $u(r) \rightarrow (1 - r/a) \text{ as } a \rightarrow \infty.$
• $\Delta a_{\phi} < 1.6 \text{ fm}$

[1] R. Machleidt and I. Slaus, J. Phys. G 27, R69 (2001).
[2] T. E. O. Ericson and G. A. Miller, Phys. Lett. 132B, 32 (1983).

Nucleon B. E. in Nuclear Matter

- The volume term in the semiempirical mass formula
- Nucleon self-energy correction in N = Z nuclear matter using Hartree approximation $\delta B_{N} = \frac{1}{2m_{\phi}^{2}} g_{N} (g_{p} + g_{n}) \rho + \cdots$ $\rho \sim 0.08 \text{ fm}^{-3}$ $\cdot \frac{\delta B_{p} + \delta B_{n}}{2} = \frac{1}{4m_{\phi}^{2}} (g_{p} + g_{n})^{2} \rho$ $= \left(+ \left(- \frac{1}{2m_{\phi}^{2}} + \frac{$ • $\frac{\delta B_p + \delta B_n}{1 \text{ MeV}} < 1 \text{ MeV}$



B. E. difference of ${}^{3}\text{He}$ and ${}^{3}\text{H}$

- ${}_{2}^{3}$ He- ${}_{1}^{3}$ H binding energy difference (763.76 keV)
- Coulomb interaction: 693 keV
- Charge asymmetry of p and n: 68 keV [1-5]
- Contribution from scalar boson potential

 $\frac{2\alpha}{\sqrt{3}\pi} \int_0^\infty q^2 dq \, \frac{\epsilon_p^2 - \epsilon_n^2}{q^2 + m_\phi^2} F(q^2) < 30 \text{ keV} [2, 6-8]$

[1] J. L. Friar, Nucl. Phys. A156, 43 (1970).

- [2] J. L. Friar and B. F. Gibson, Phys. Rev. C 18, 908 (1978).
- [3] S. A. Coon and R. C. Barrett, Phys. Rev. C 36, 2189 (1987).
- [4] G. A. Miller, B. M. K. Nefkens, and I. Slaus, Phys. Rep. 194, 1 (1990).
- [5] R. B. Wiringa, S. Pastore, S. C. Pieper, and G. A. Miller, Phys. Rev. C 88, 044333 (2013).
- [6] I. Sick, Prog. Part. Nucl. Phys. 47, 245 (2001).
- [7] F. P. Juster et al., Phys. Rev. Lett. 55, 2261 (1985).
- [8] J. S. Mccarthy, I. Sick, and R. R. Whitney, Phys. Rev. C 15, 1396 (1977).

Electron Beam Dump Experiments

- Search for new particle [1-3]
- E137 (20 GeV, 30C) [4]
- E141 (9 GeV, 0.32 mC) [5]
- Orsay (1.6GeV, 3.2 mC) [6]



FIG. 2. Layout of SLAC experiment E137.

- [1] J. D. Bjorken, R. Essig, P. Schuster, and N. Toro, Phys. Rev. D 80, 075018 (2009).
- [2] R. Essig et al., arXiv:1311.0029.
- [3] Y. S. Liu, D. McKeen and G. A. Miller, arXiv:1609.06781.
- [4] J. D. Bjorken, S. Ecklund, W. R. Nelson, A. Abashian, C. Church, B. Lu, L.W. Mo,
 - T. A. Nunamaker, and P. Rassmann, Phys. Rev. D 38, 3375 (1988).
- [5] E. M. Riordan et al., Phys. Rev. Lett. **59**, 755 (1987).
- [6] M. Davier and H. Nguyen Ngoc, Phys. Lett. B 229, 150 (1989).



- $\Delta a_l = \frac{\alpha}{2\pi} \epsilon_l^2 \xi\left(\frac{m_\phi}{m_l}\right), \xi(x) = \int_0^1 \frac{(1-z)^2(1+z)}{(1-z)^2+x^2z} dz$
- The most accurate method to measure α
- The shift of $(g 2)_e$ by ϕ results in the measured α by $\Delta \alpha = 2\pi \Delta a_e$ [1].
- A measurement of α , which is not sensitive to ϕ , using ⁸⁷Rb atom [2] gives 0.66 ppb uncertainty.
- $\Delta a_e < 1.5 \times 10^{-12}$ (taking 2 S.D.)

[1] M. Pospelov, Phys. Rev. D 80, 095002 (2009).

[2] R. Bouchendira, P. Clade, S. Guellati-Khelifa, F. Nez, and F. Biraben, Phys. Rev. Lett. 106, 080801 (2011).

Resonance in Bhabha Scattering

- GSI group
- No resonance were found at 97% C.L. within the experimental sensitivity [1]

$$\int d\sqrt{s} \left(\frac{d\sigma}{d\Omega}\right)_{c.m.} < 0.5 \text{ b eV/sr.}$$

[1] H. Tsertos, C. Kozhuharov, P. Armbruster, P. Kienle, B. Krusche, and K. Schreckenbach, Phys. Rev. D 40, 1397 (1989).

eH Lamb Shift

•
$$\delta E_L^{lN} = -\frac{\alpha}{2a_l} \epsilon_l \epsilon_N f(a_{lN} m_{\phi})$$
 (for lepton-nucleus)
 $f(x) = \frac{x^2}{(1+x)^4}$, a_{lN} is the Bohr radius

•
$$\delta E_L^{eH} < 14 \text{ kHz} [1] (5.8 \times 10^{-11} \text{ eV})$$

Supernova 1987A [1]

- Cooling: produce φ to escape SN electron-positron annihilation [2,3] Nucleon-nucleon bremsstrahlung [4]
- Trapping: scatter ($e\phi \rightarrow e\phi$), decay [5]
- Absorption: $e\phi \rightarrow e\gamma$, $N\phi \rightarrow N\gamma$ [5]

- [2] H. K. Dreiner, C. Hanhart, U. Langenfeld and D. R. Phillips, Phys. Rev. D 68, 055004 (2003).
- [3] H. K. Dreiner, J. F. Fortin, C. Hanhart and L. Ubaldi, Phys. Rev. D 89, no. 10, 105015 (2014).
- [4] E. Rrapaj and S. Reddy, arXiv:1511.09136.
- [5] G. G. Raffelt, *Stars as laboratories for fundamental physics : The astrophysics of neutrinos, axions, and other weakly interacting particles*, Chicago, USA: Univ. Pr. (1996)

^[1] A. Burrows and J. M. Lattimer, Astrophys. J. **307**, 178 (1986).