

# Detector performances and cosmic-ray reconstruction efficiency in MicroBooNE

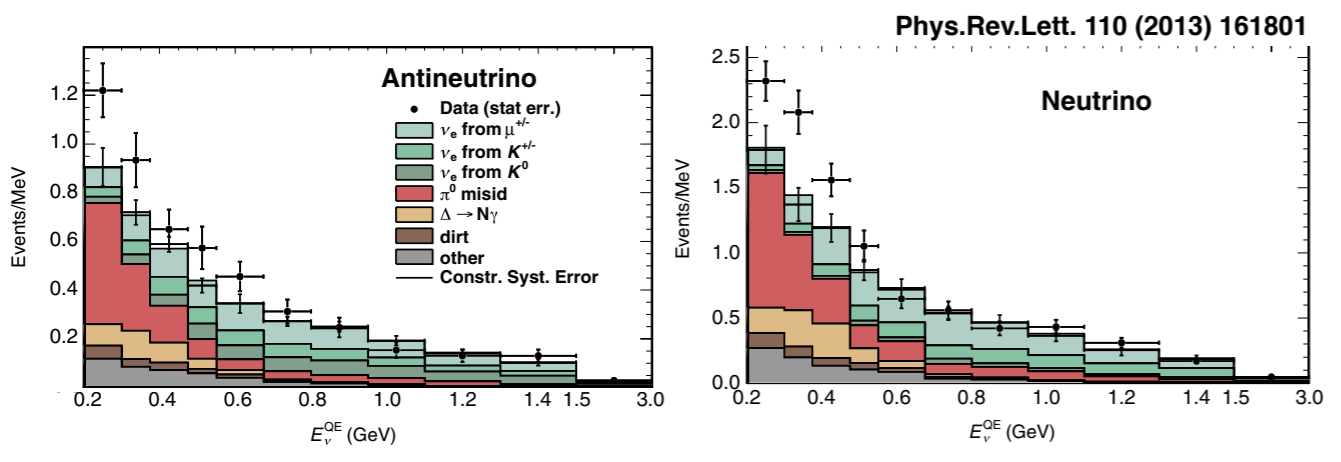
**Stefano Roberto Soletti**

WIN 2017, University of California - Irvine, 23<sup>rd</sup> June 2017



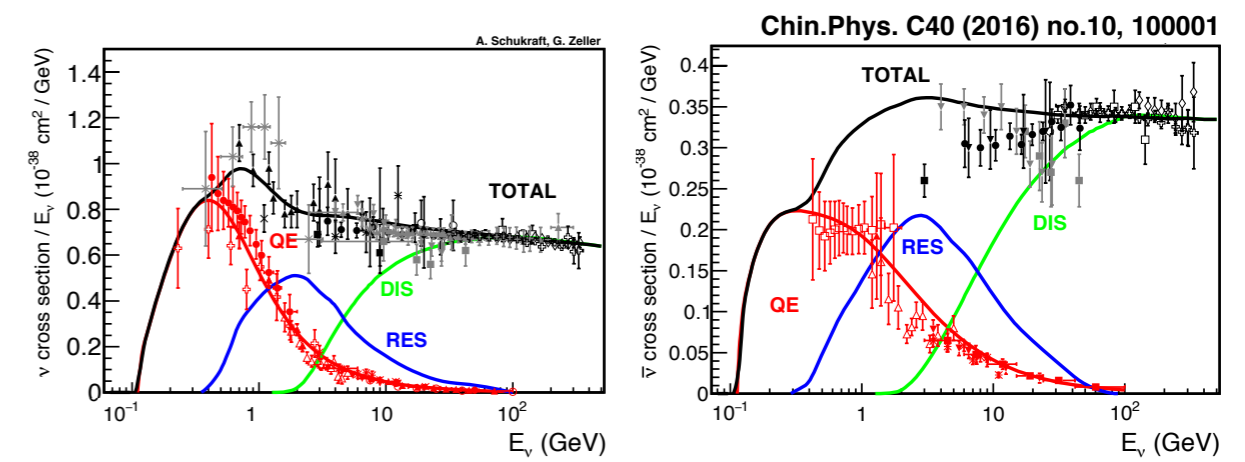
## Investigate short-baseline neutrino anomalies

- LSND excess.
- MiniBooNE low energy excess.
- Reactor neutrino anomalies.



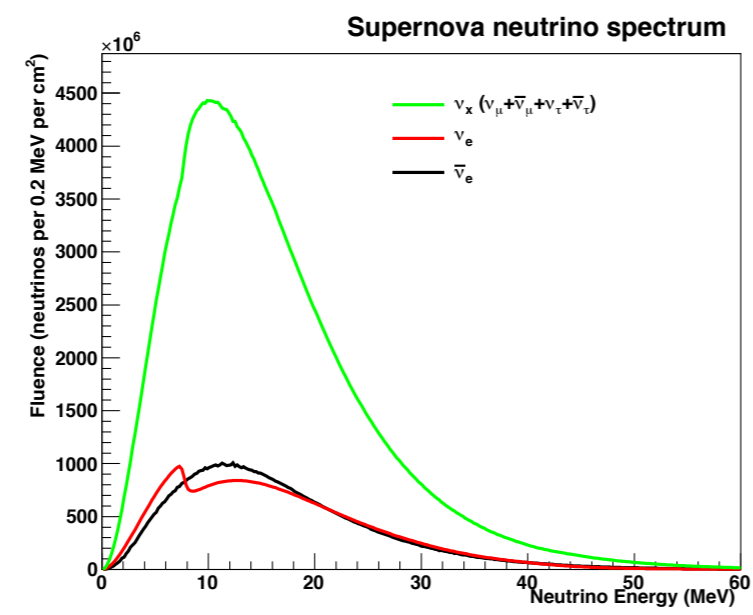
## Cross-section measurements

- Precise measurements of  $\nu$ -Ar cross section for future LAr experiments (DUNE).
- Probe different theories of nuclear effects in  $\nu$ -Ar scattering.



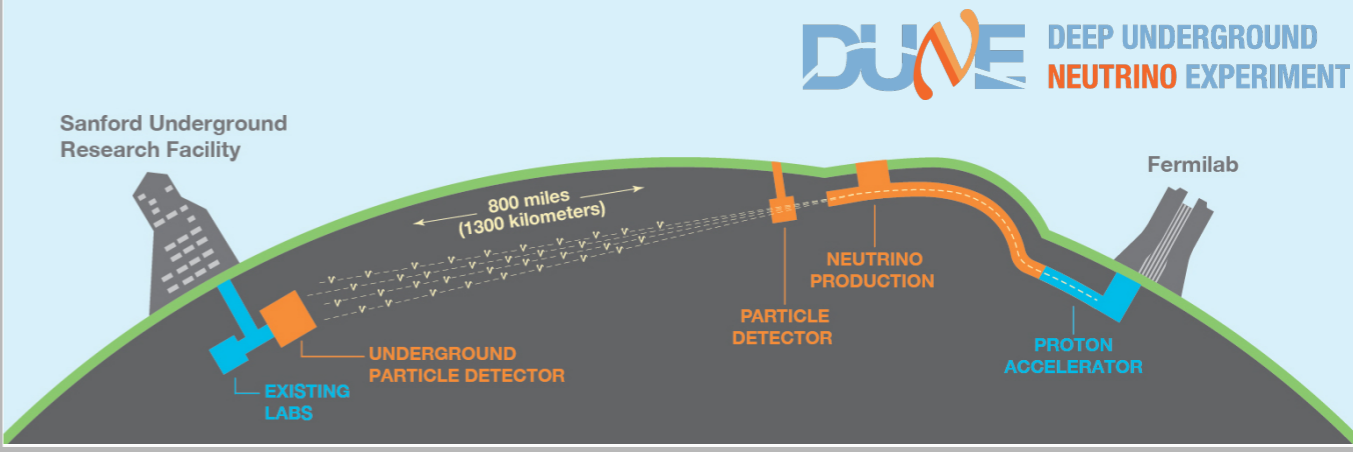
## Supernova neutrino and exotic physics

- Supernova neutrinos (~10 MeV) using the SNEWS alert system
- Proton decay backgrounds study:  $K_L^0 + p \rightarrow K^+ + n$



## LArTPC detector R&D

- Detector effects to further develop LArTPC technology
  - Space charge effect, wire response, noise studies, electron lifetime...
- Essential for future experiments.

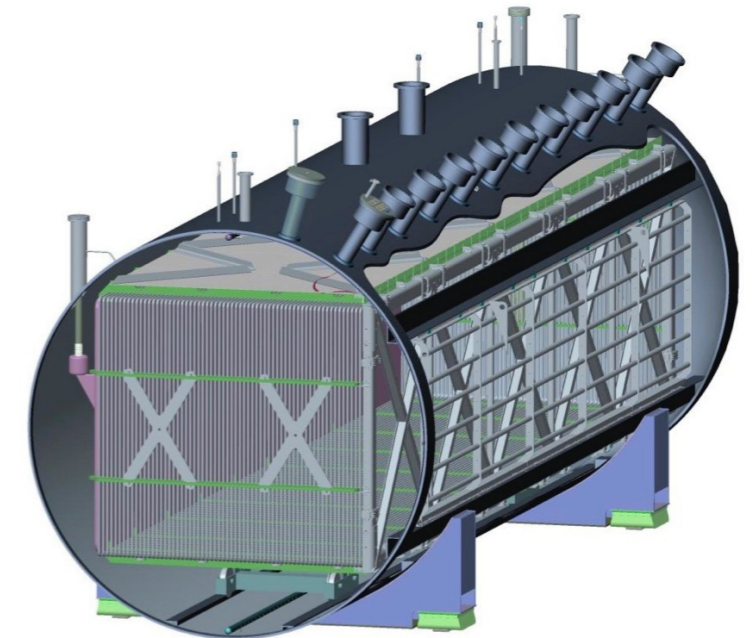




## What is it?

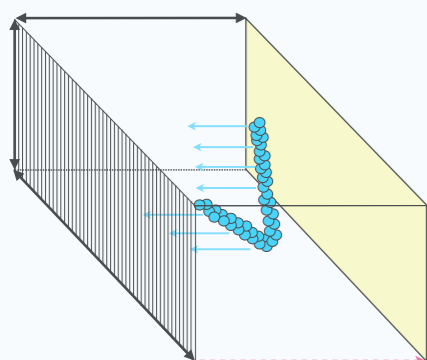
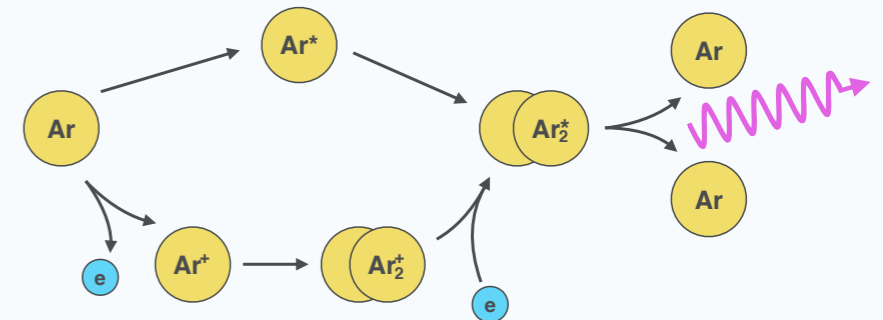
MicroBooNE is a 85 active mass ton **Liquid Argon Time Projection Chamber** at Fermilab, located:

- on the axis of the Booster Neutrino Beam (BNB);
- off the axis of the Neutrinos at the Main Injector (NuMI) beam.

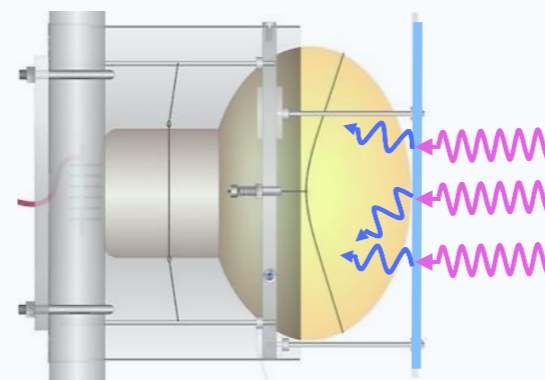


## How does it work?

As a charged particle passes through the liquid argon, the argon atom experiences one of two processes, it becomes ionized, or excited. Both paths end by de-excitation via **scintillation**.

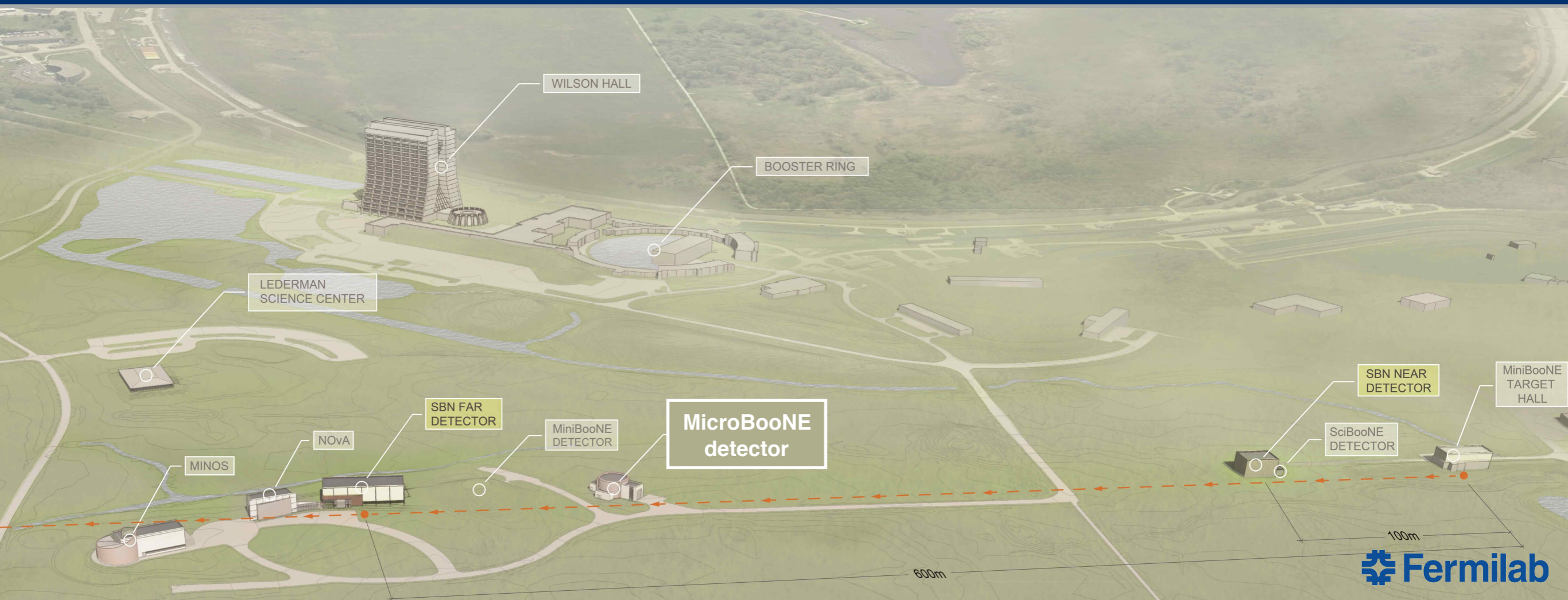


Ionized **electrons** travel through the liquid argon in a constant electric field and are **collected by the TPC**.



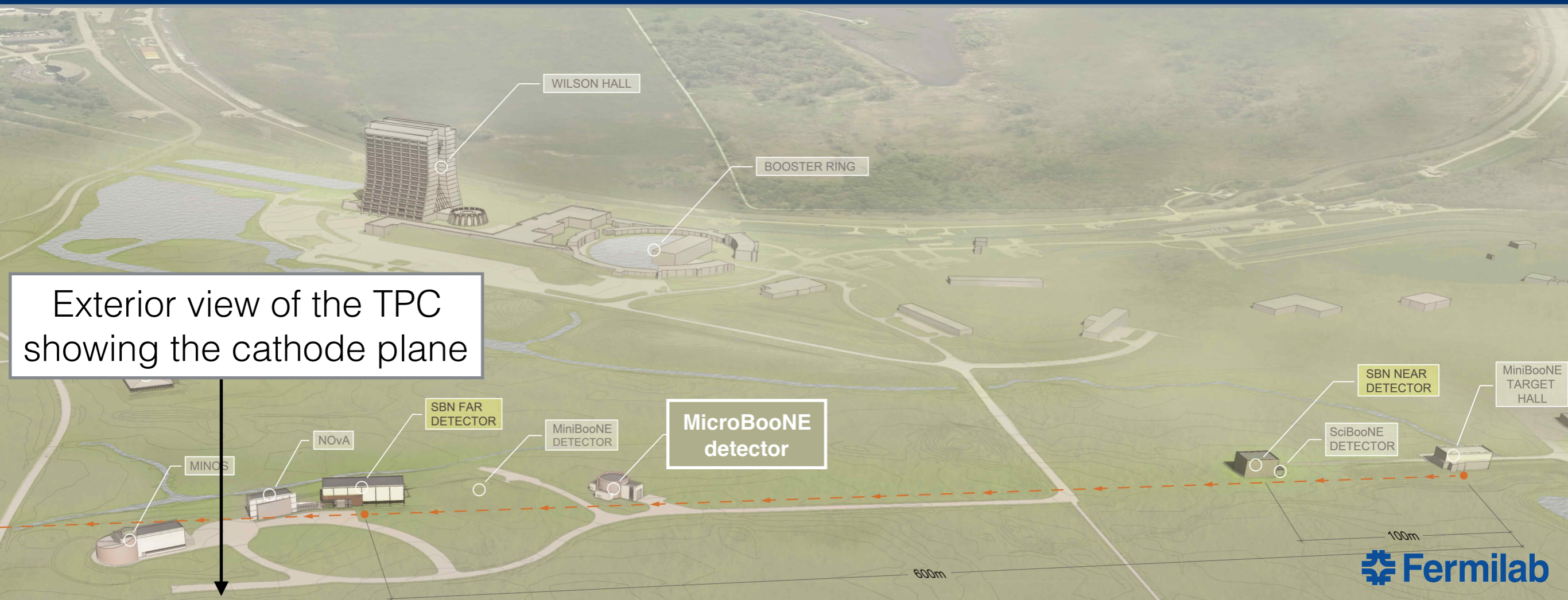
The **photons** are collected by one of the 32 8-inch **photomultipliers (PMTs)**,





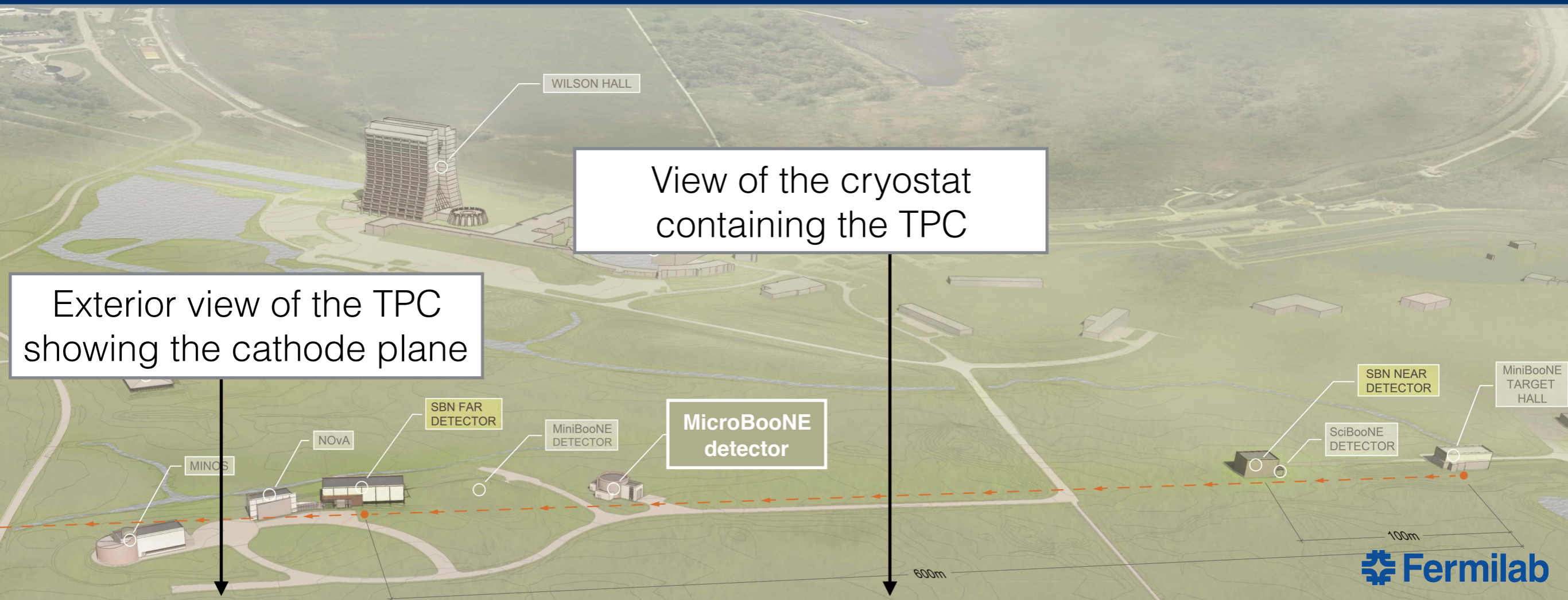
arXiv:1612.05824





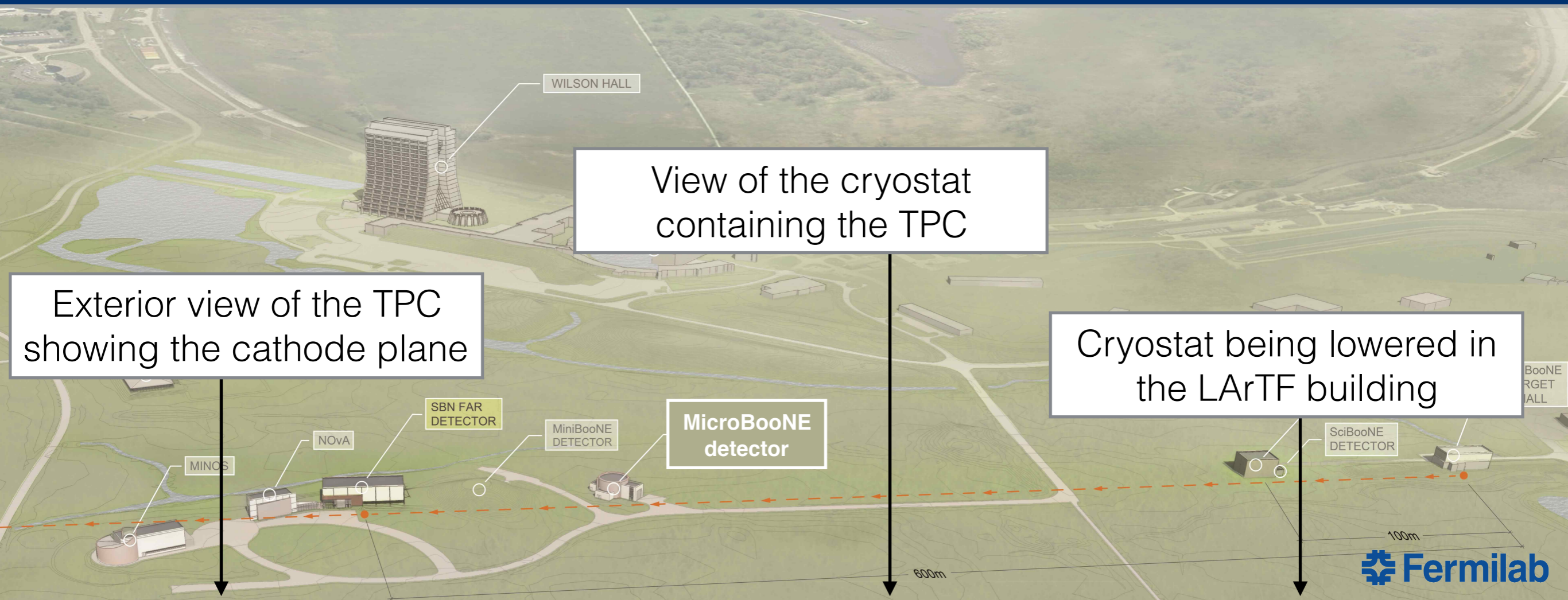
arXiv:1612.05824





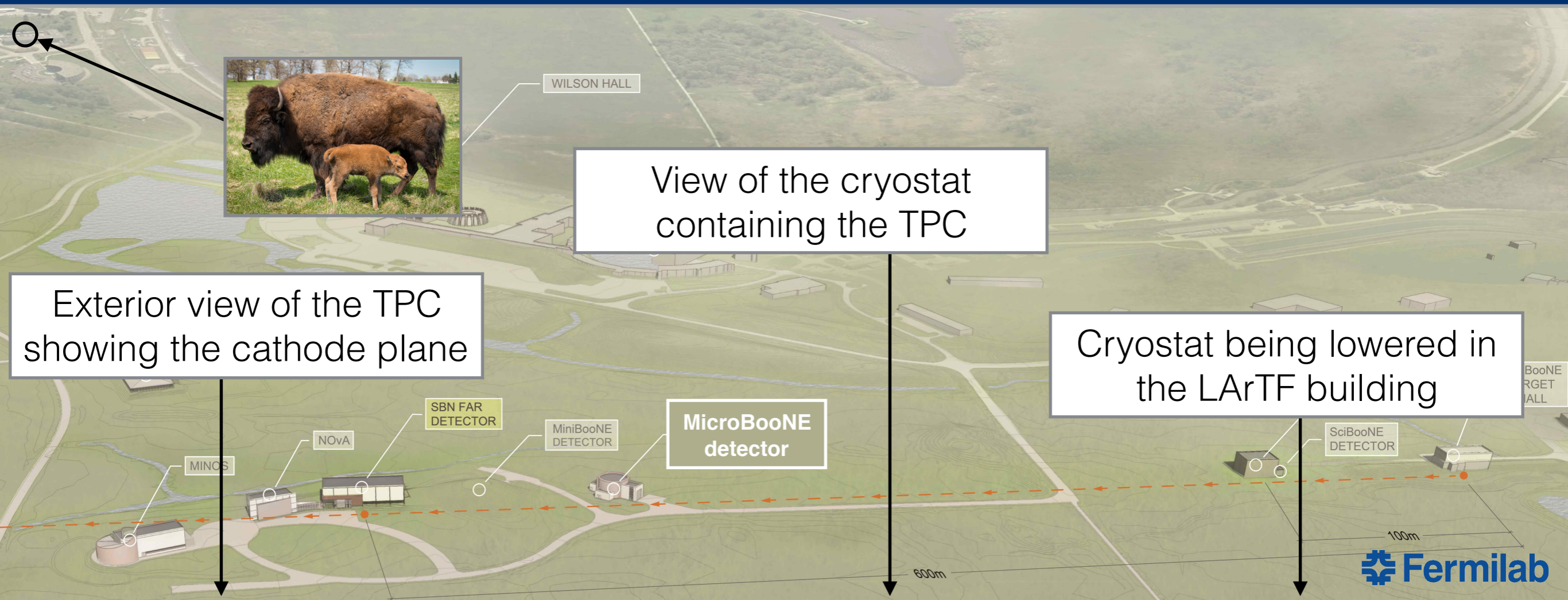
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**$\mu$ BooNE**

Collection plane

Proton track  
 $\nu$  beam

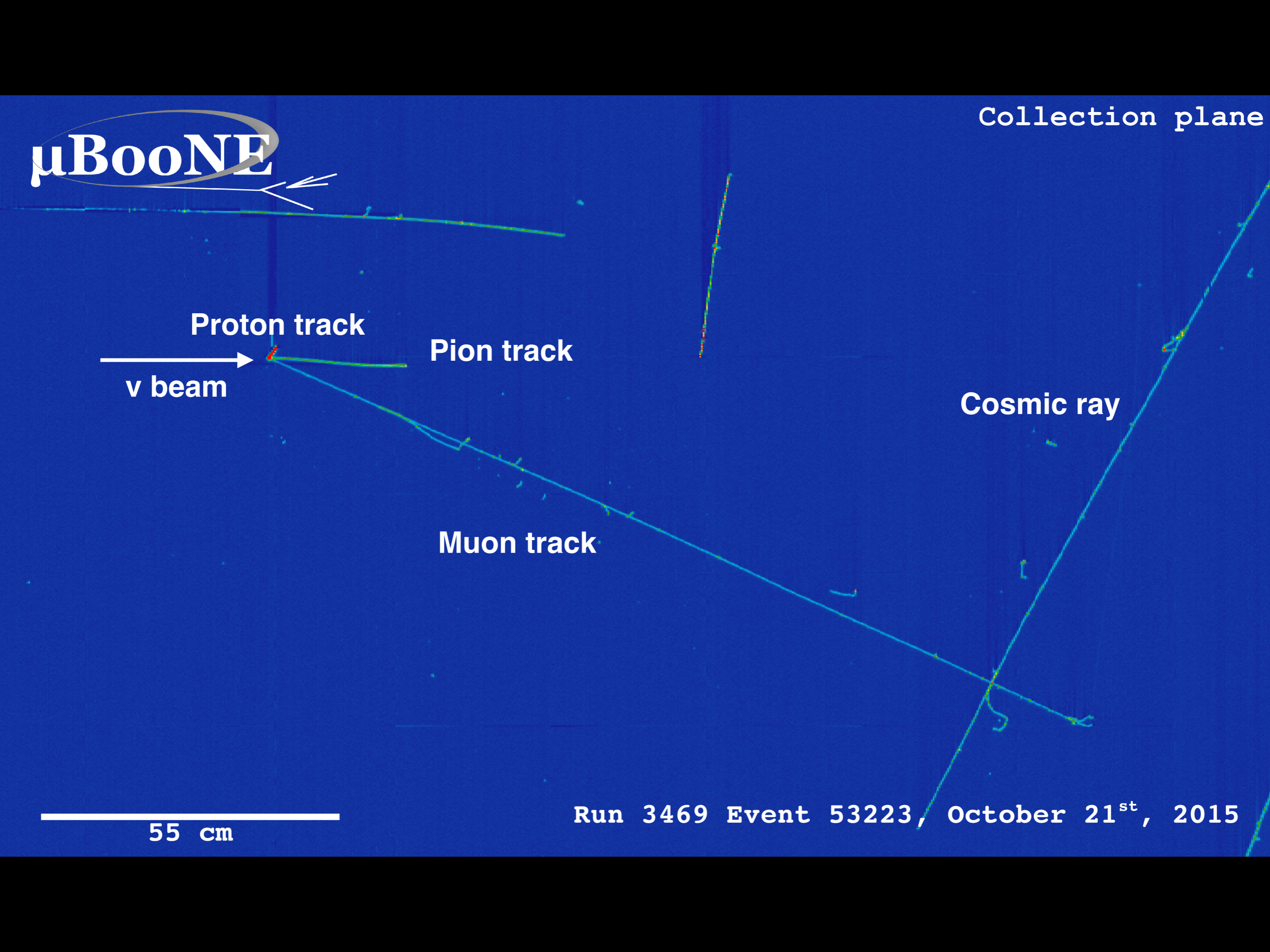
Pion track

Muon track

Cosmic ray

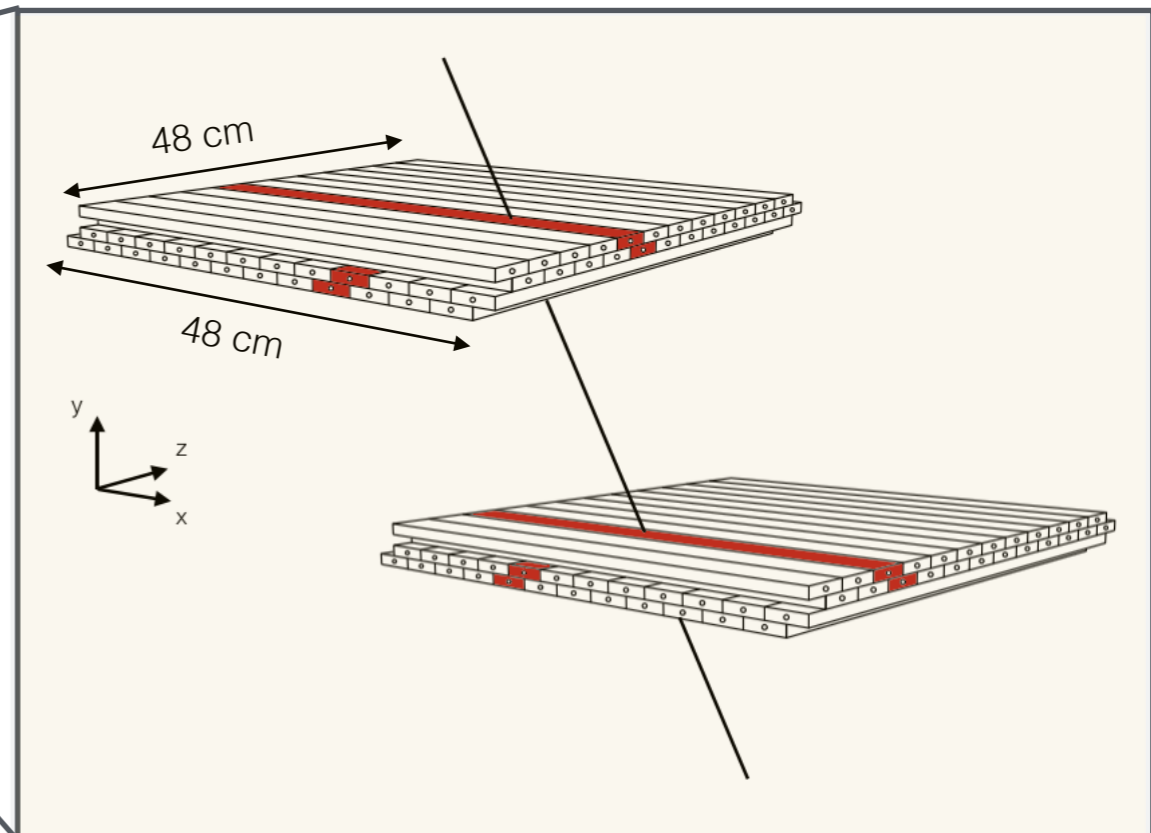
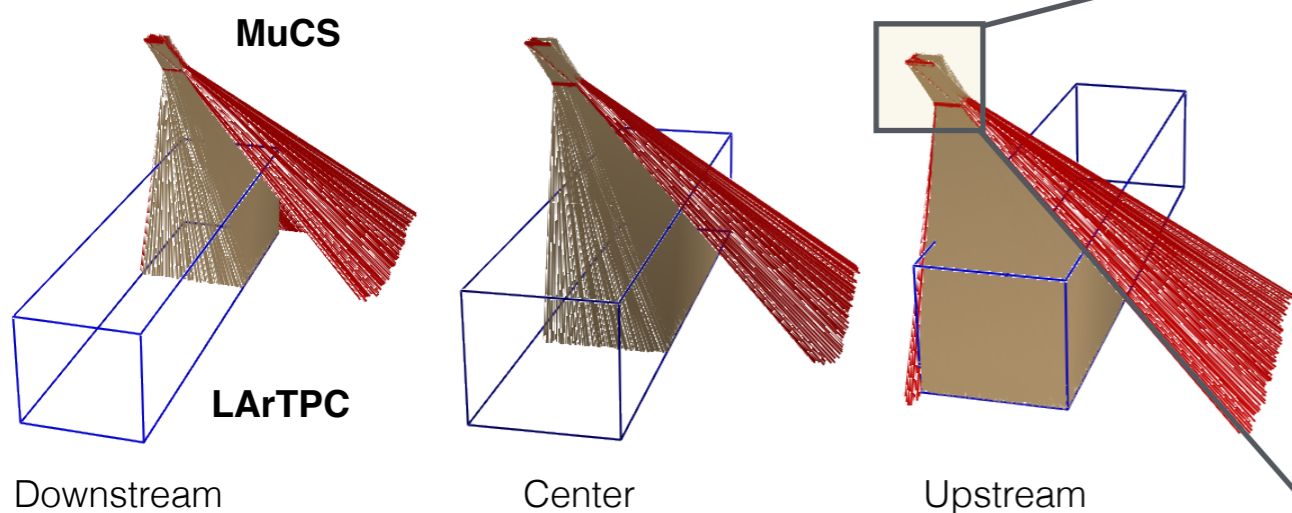
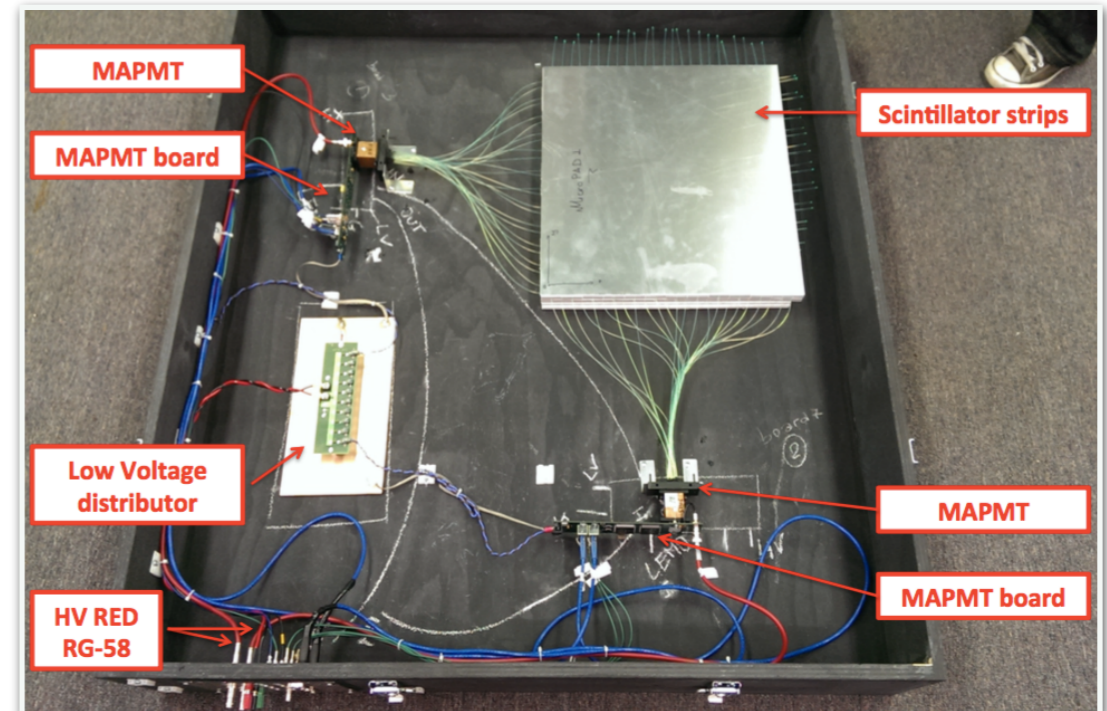
55 cm

Run 3469 Event 53223, October 21<sup>st</sup>, 2015



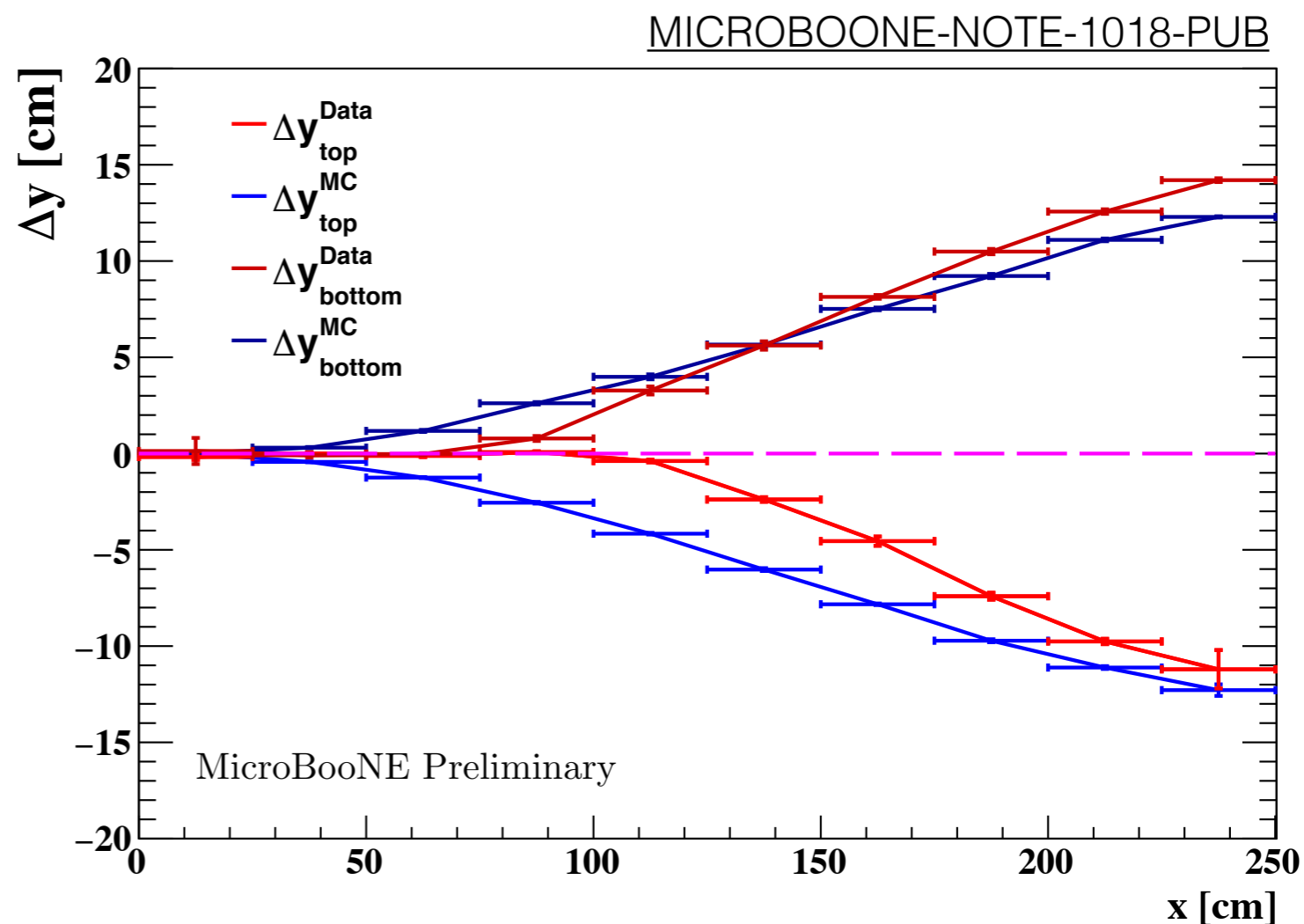
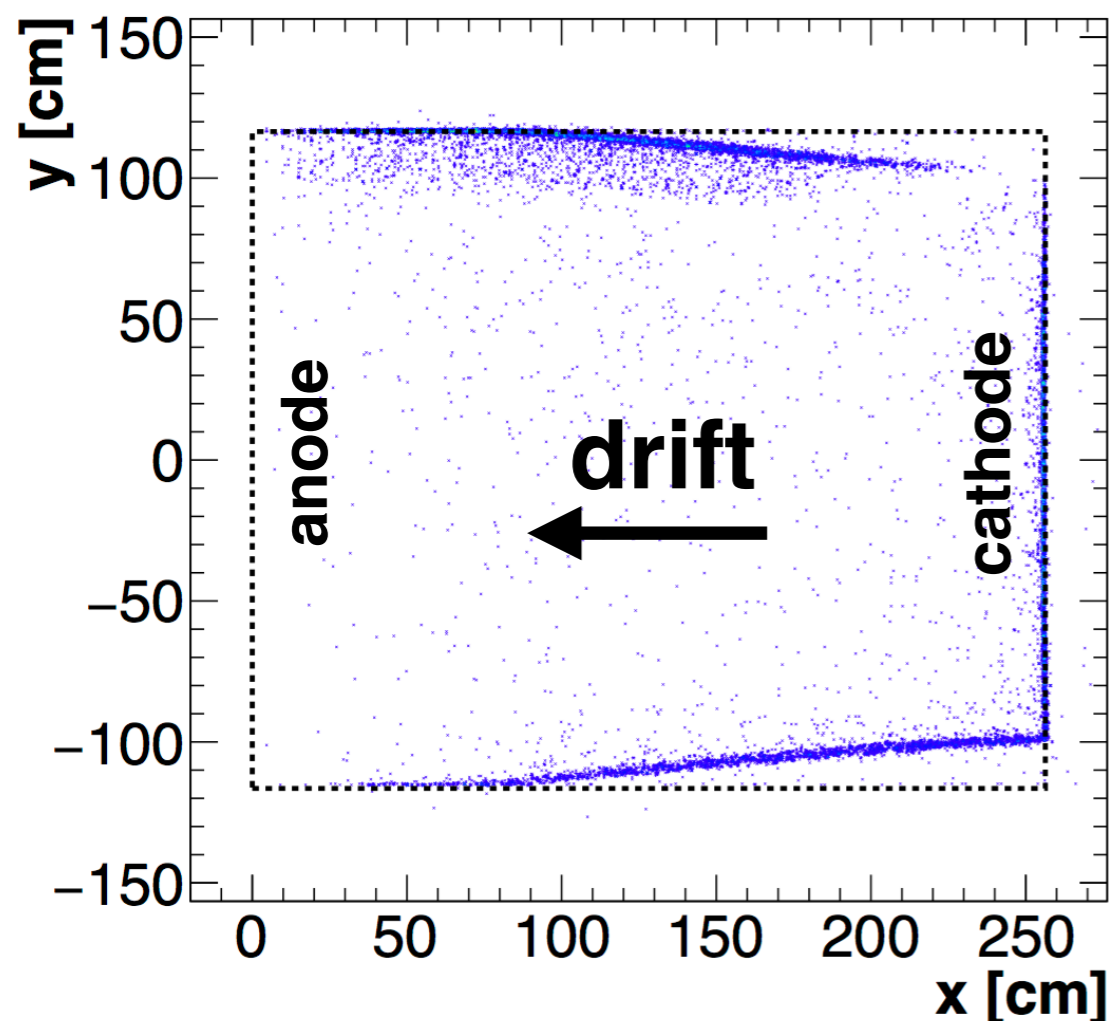


- Being MicroBooNE located near the surface, **cosmic muons can be a source of backgrounds to many analyses** (~10 cosmic muons per 2.2 ms drift time).
- A small muon counter stack (MuCS) has been installed on top of the TPC to help with several studies:
  - **Data reconstruction efficiency**
  - Optical system - TPC matching efficiency
  - Trigger efficiency
  - **Detector performances** (space-charge effect, collected charge, collected light...)



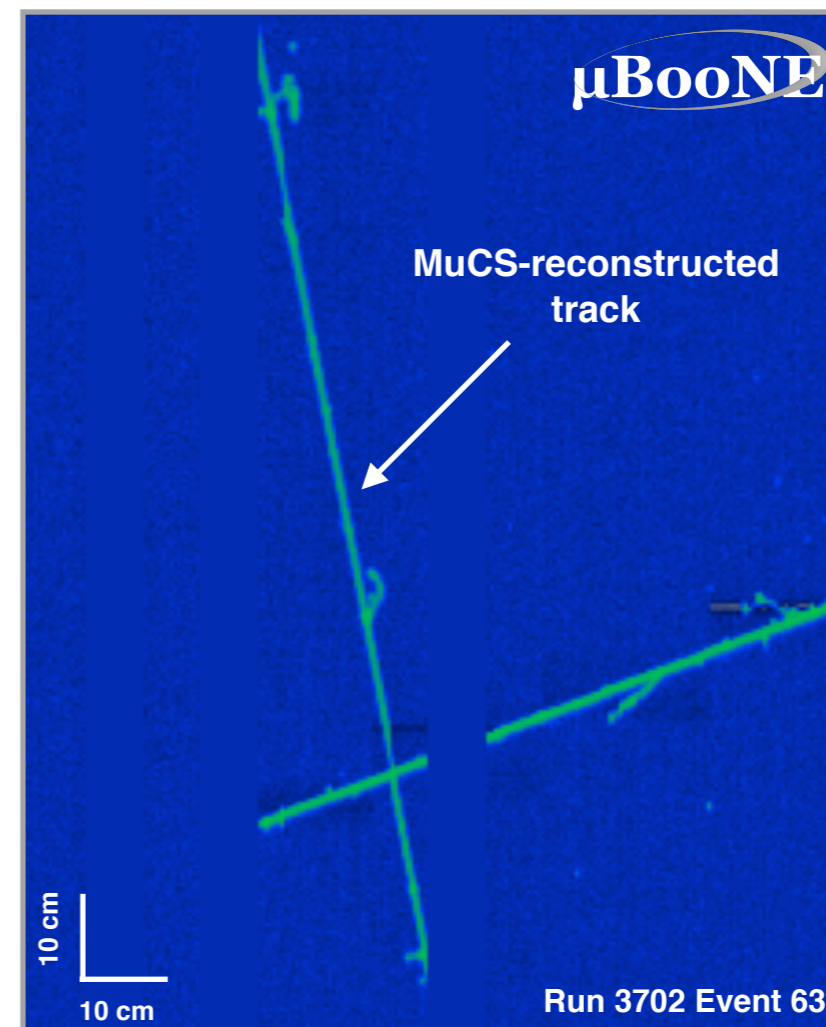
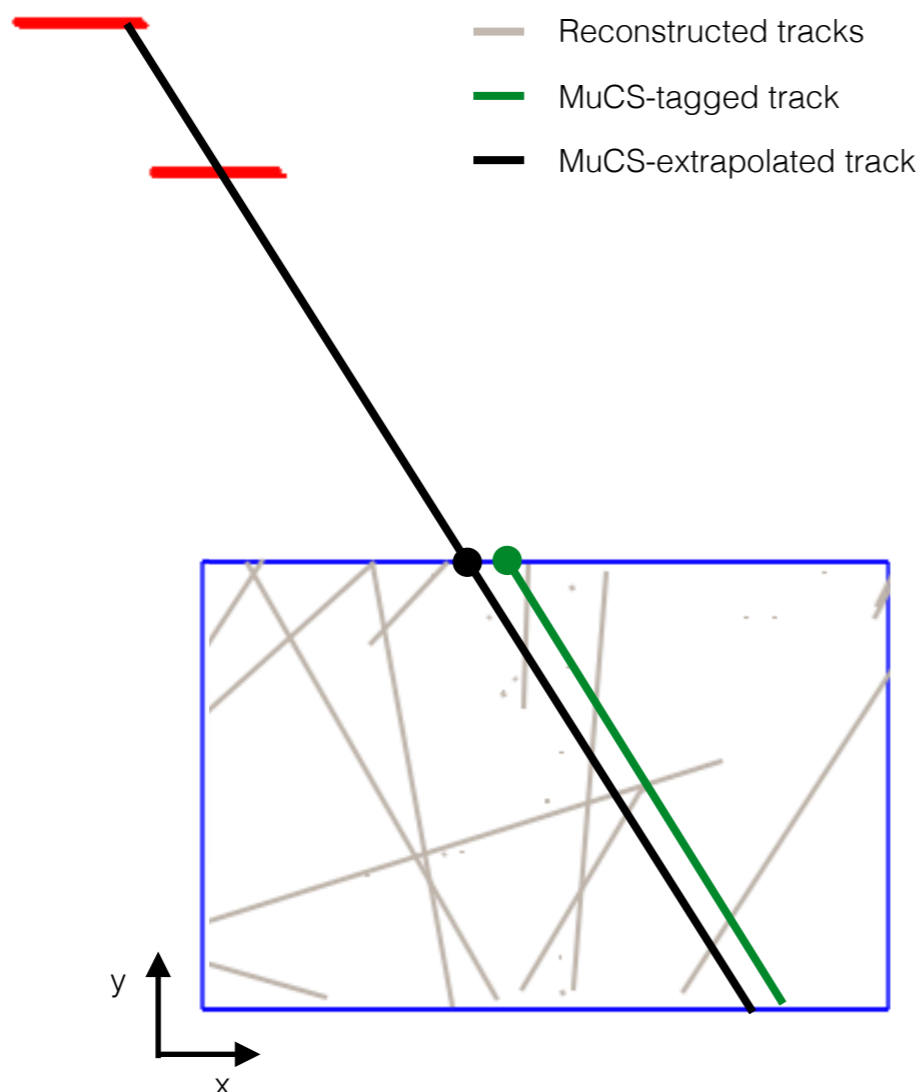


- The positive argon ions can cause a **distortion in the electrical field** of the TPC.
- We can use the MuCS dataset to quantify this **space-charge effect** (SCE).
- SCE simulation qualitatively reproduces effect.
  - Agrees in normalization and base shape features.
  - Offset near anode probably caused by liquid argon flow.
- Can impact track/shower reconstruction and calorimetry.





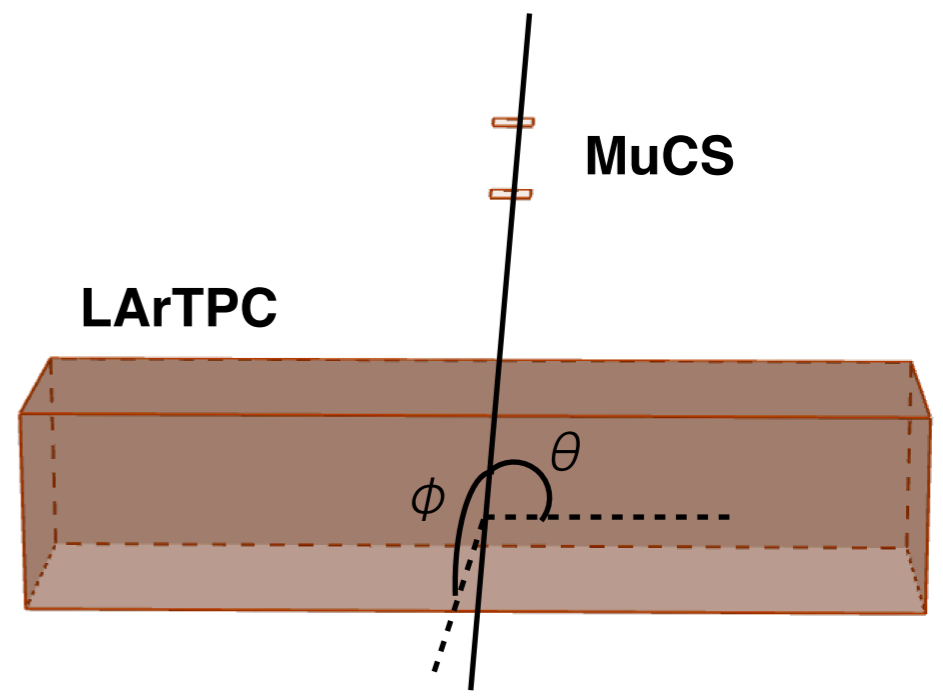
- Each MuCS event is **triggered by a cosmic ray going through the MuCS panels**.
- Each MuCS event will contain more than one reconstructed cosmic-ray track (we have  $\sim 10$  cosmic rays per 2.2 ms drift window).
- We find the **reconstructed track with the starting points closest to the intersection** between the extrapolated MuCS trajectory (**MuCS-extrapolated track**) and the TPC, within a maximum distance (**MuCS-tagged track**).
- Number of MuCS-tagged tracks is corrected by the purity and the acceptance of the cut on the maximum distance.





Using the  $(\mathbf{x}_{\text{top}}, \mathbf{y}_{\text{top}}, \mathbf{z}_{\text{top}})$  and  $(\mathbf{x}_{\text{bottom}}, \mathbf{y}_{\text{bottom}}, \mathbf{z}_{\text{bottom}})$  points given by the MuCS panels, it is possible to measure the reconstruction efficiency as a function of  $\theta$ ,  $\phi$  and extrapolated length in the TPC  $L$ .

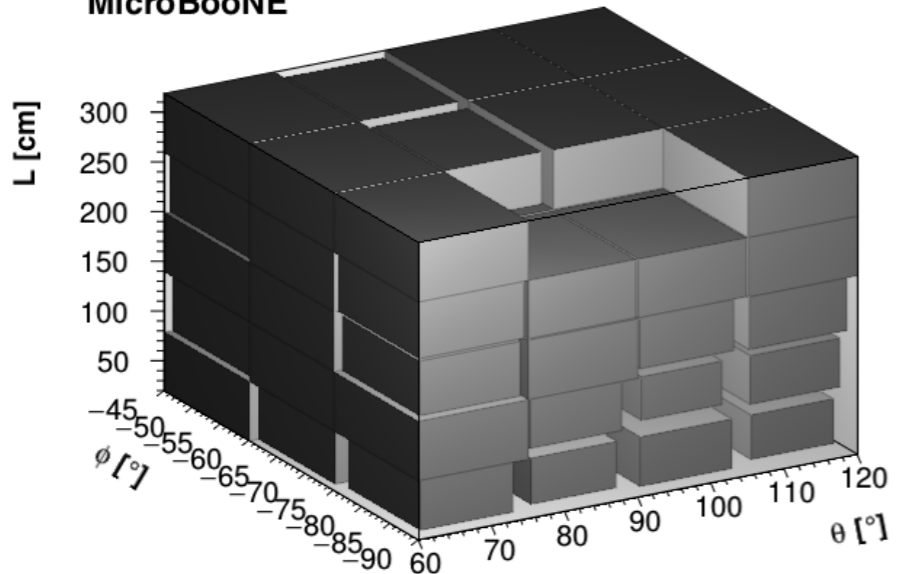
$$\phi = \text{atan} \left( \frac{y_{\text{top}} - y_{\text{bottom}}}{x_{\text{top}} - x_{\text{bottom}}} \right), \quad \theta = \text{acos} \left( \frac{z_{\text{top}} - z_{\text{bottom}}}{r} \right)$$



## Monte Carlo

$$\epsilon_{\text{MC}} = \frac{\text{N. of reconstructed cosmic rays}}{\text{N. of generated cosmic rays}} = 97.3 \pm 0.1 \%$$

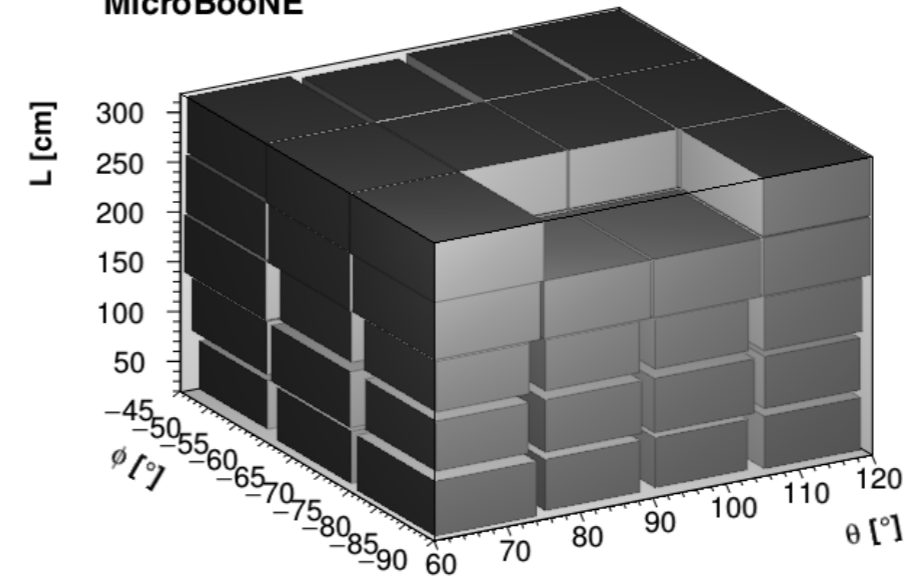
MicroBooNE



## Data

$$\epsilon_{\text{data}} = \frac{\text{N. of reco. MuCS cosmic-ray events}}{\text{N. of MuCS triggered events}} = 97.1 \pm 0.1 \text{ (stat)} \pm 1.4 \text{ (sys)} \%$$

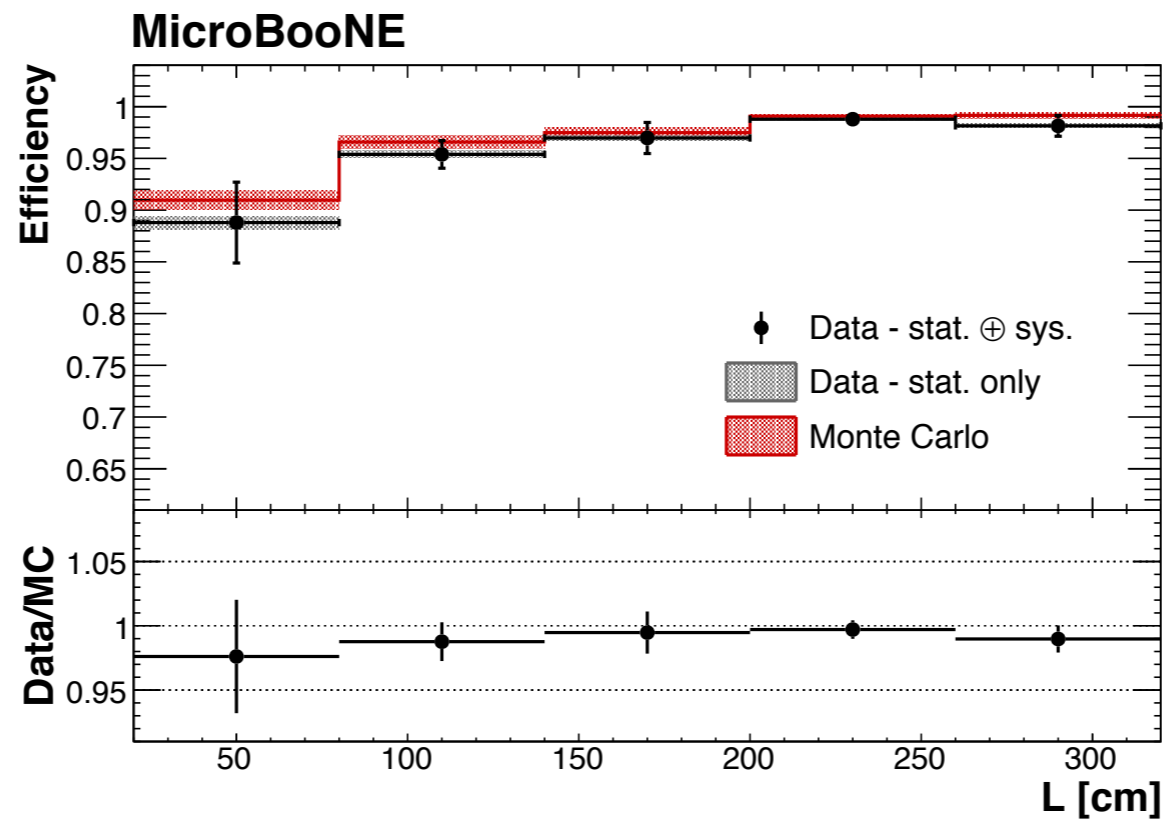
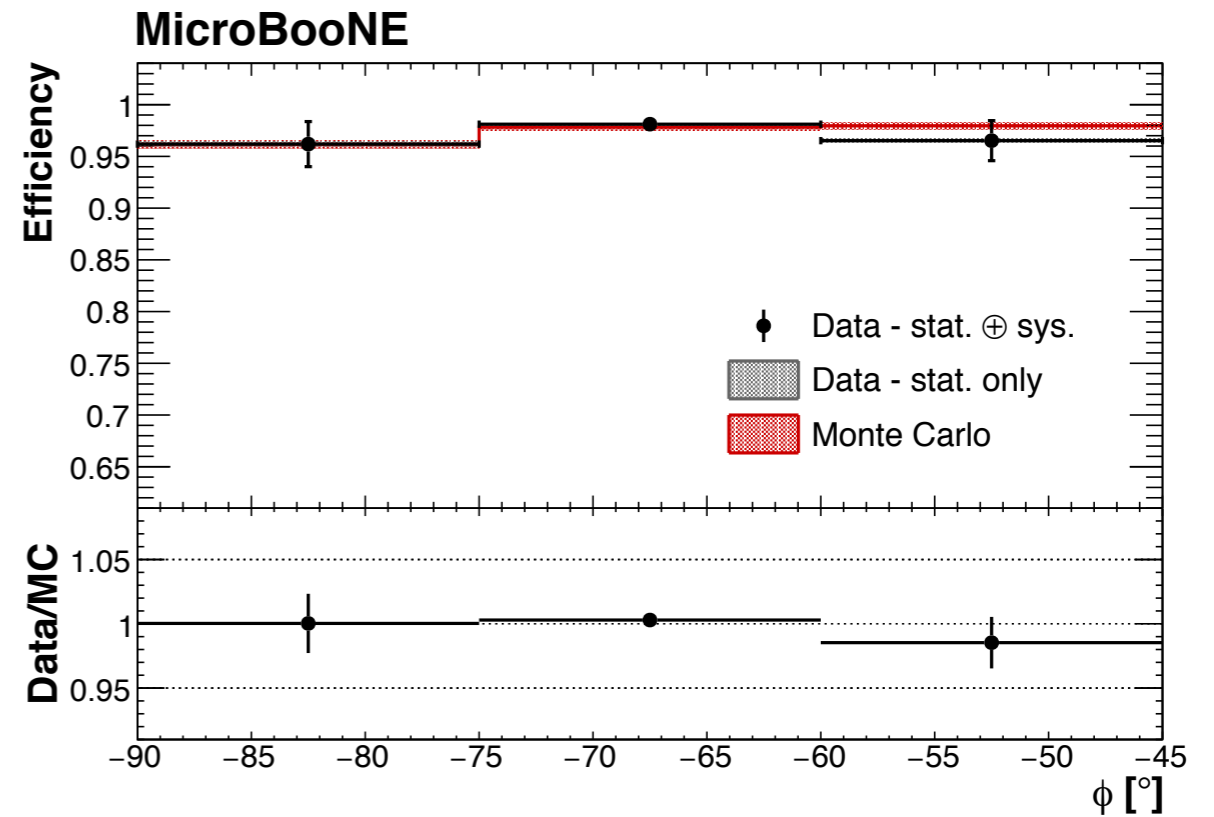
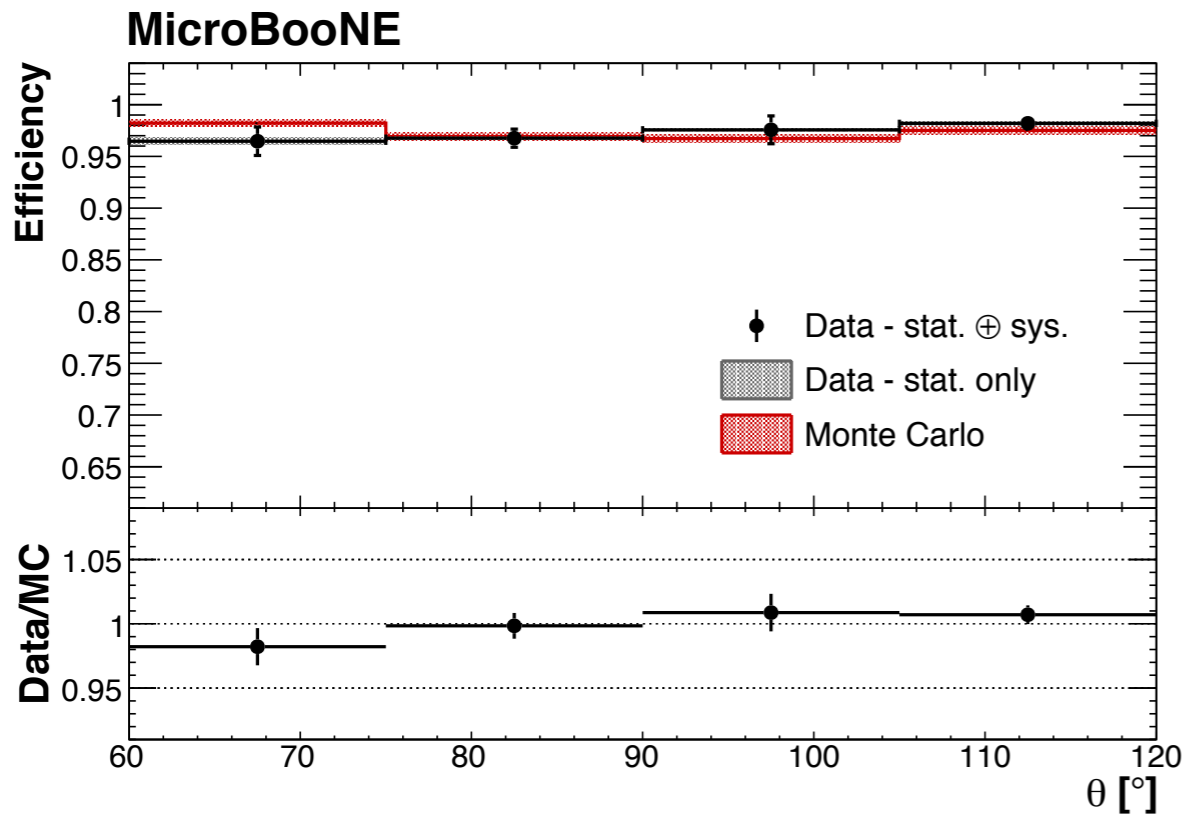
MicroBooNE





# Results

## One-dimensional projections



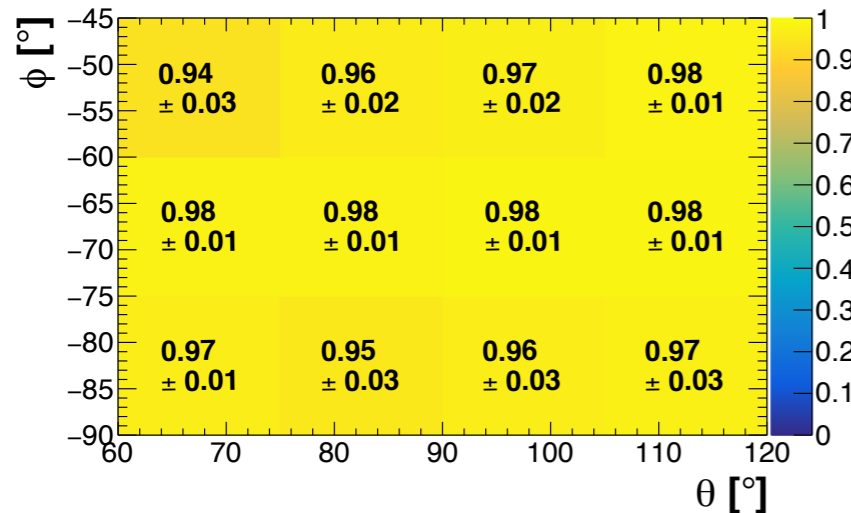


### Data

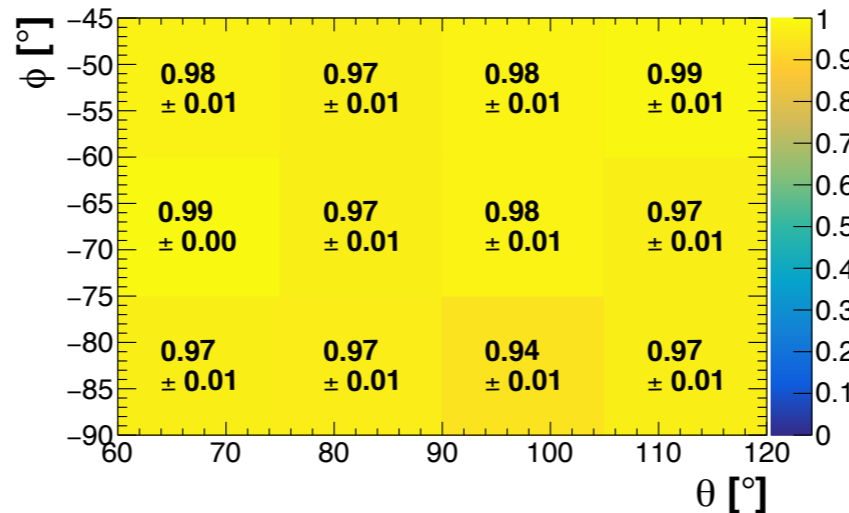
### Monte Carlo

### Data/Monte Carlo

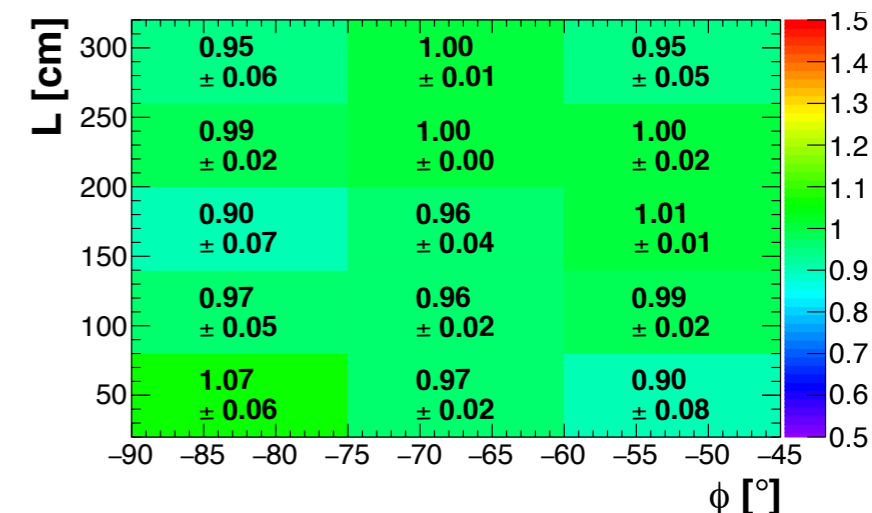
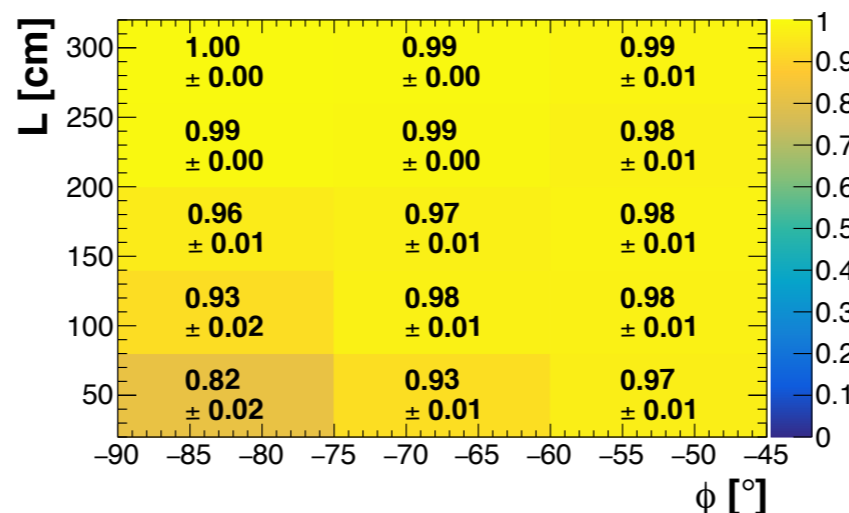
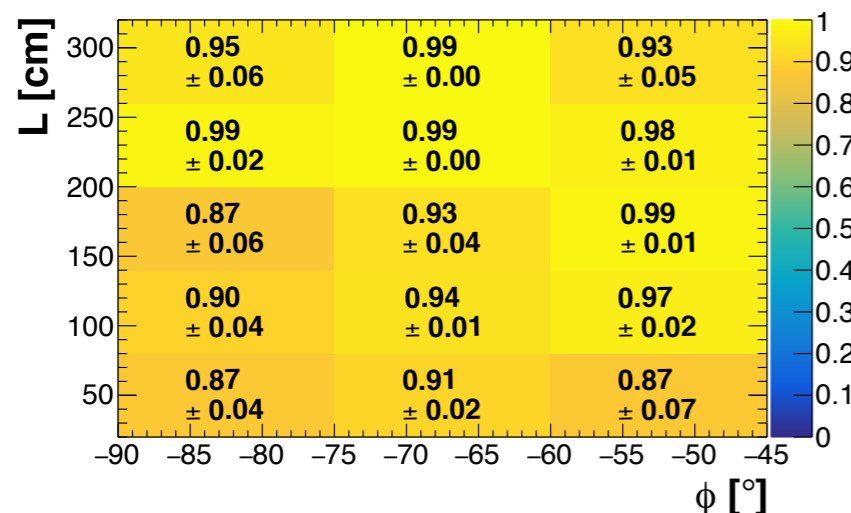
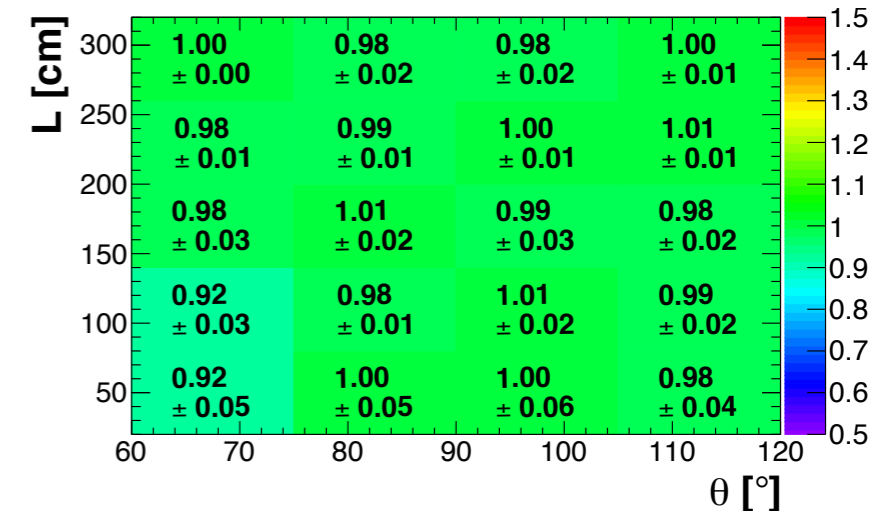
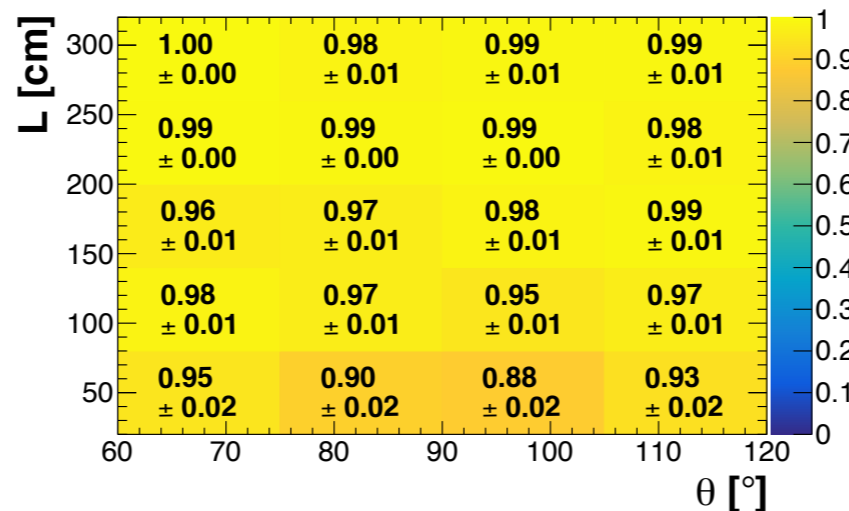
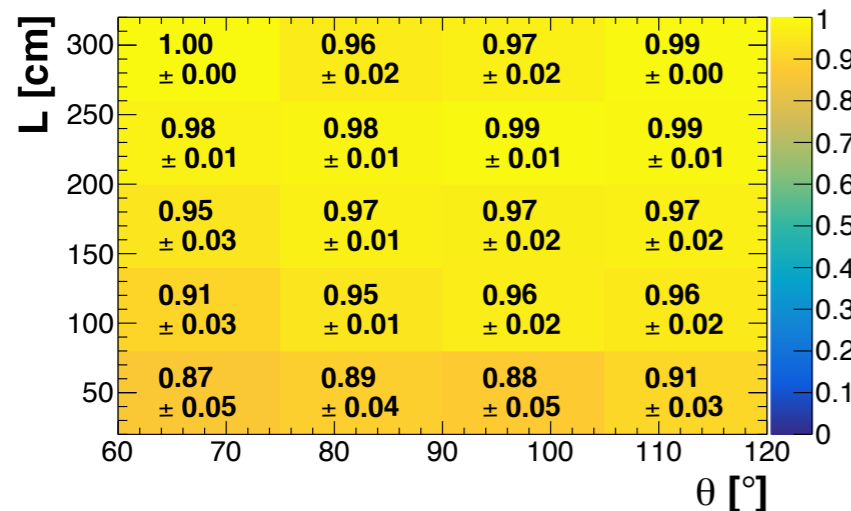
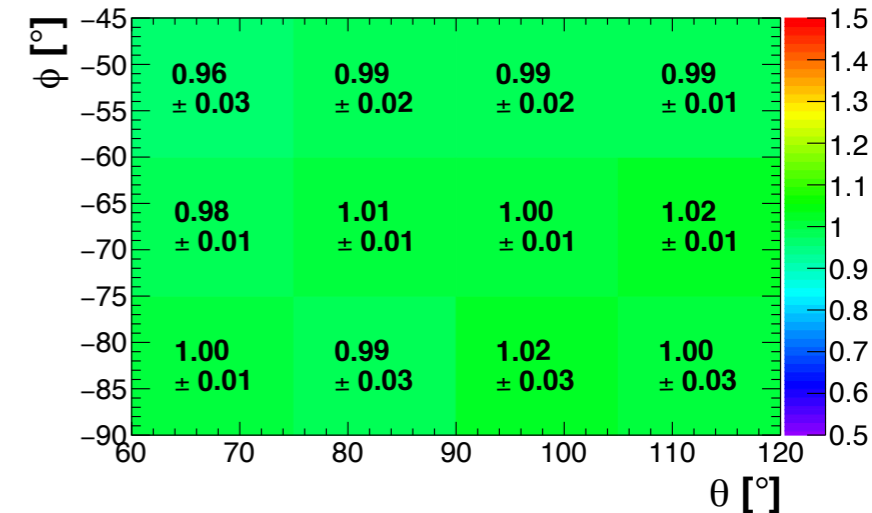
#### MicroBooNE



#### MicroBooNE

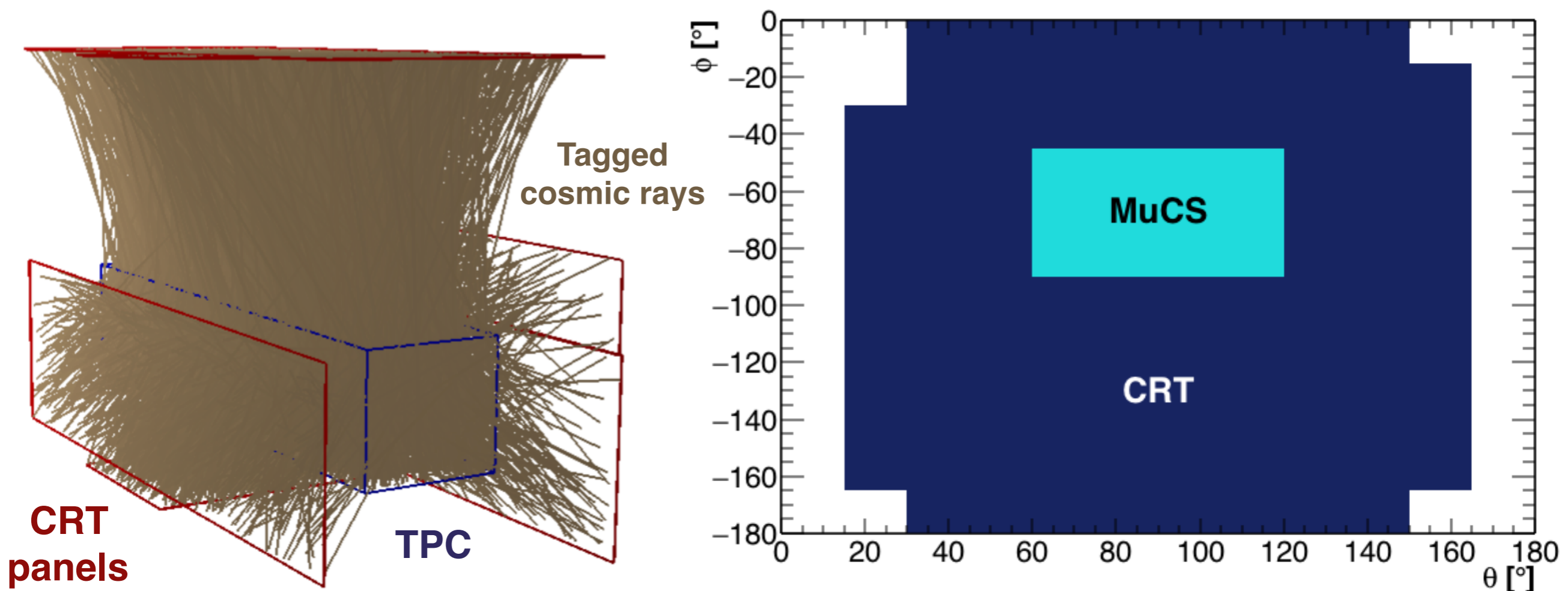


#### MicroBooNE





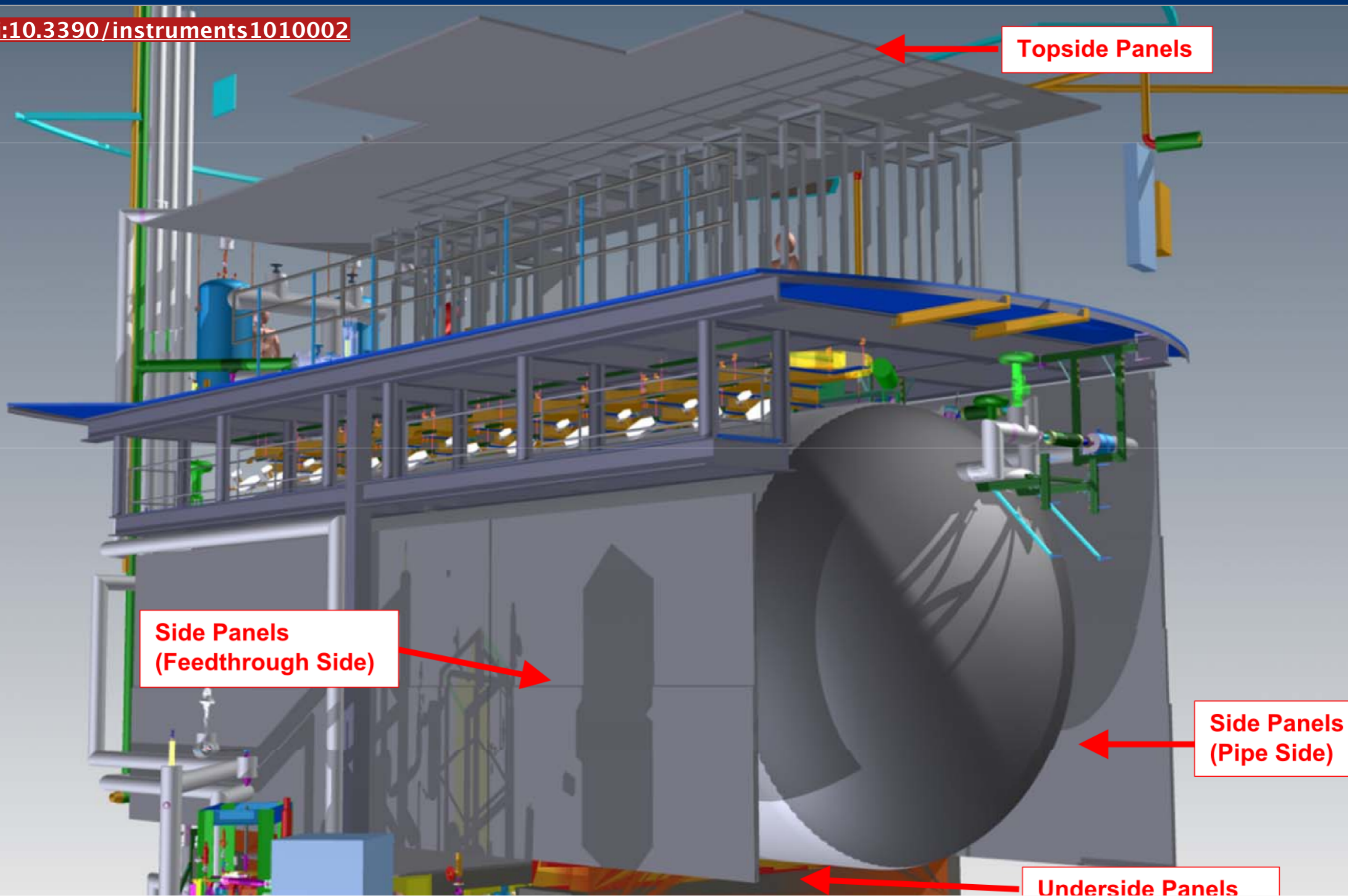
- This analysis represents a small-scale demonstration of the method that can be used with the data coming from the **Cosmic Ray Tagger**, a system of scintillation panels able to tag 85% of the cosmic-ray flux.
- The Cosmic Ray Tagger installation has been completed in January, 2017.
- The angular coverage provided by the CRT is much larger than the one of the MuCS and close to 100%. It will be possible to **measure efficiency-corrected quantities**, such as the cosmic-ray flux in MicroBooNE, and mitigate the cosmic-ray background.





# Cosmic Ray Tagger

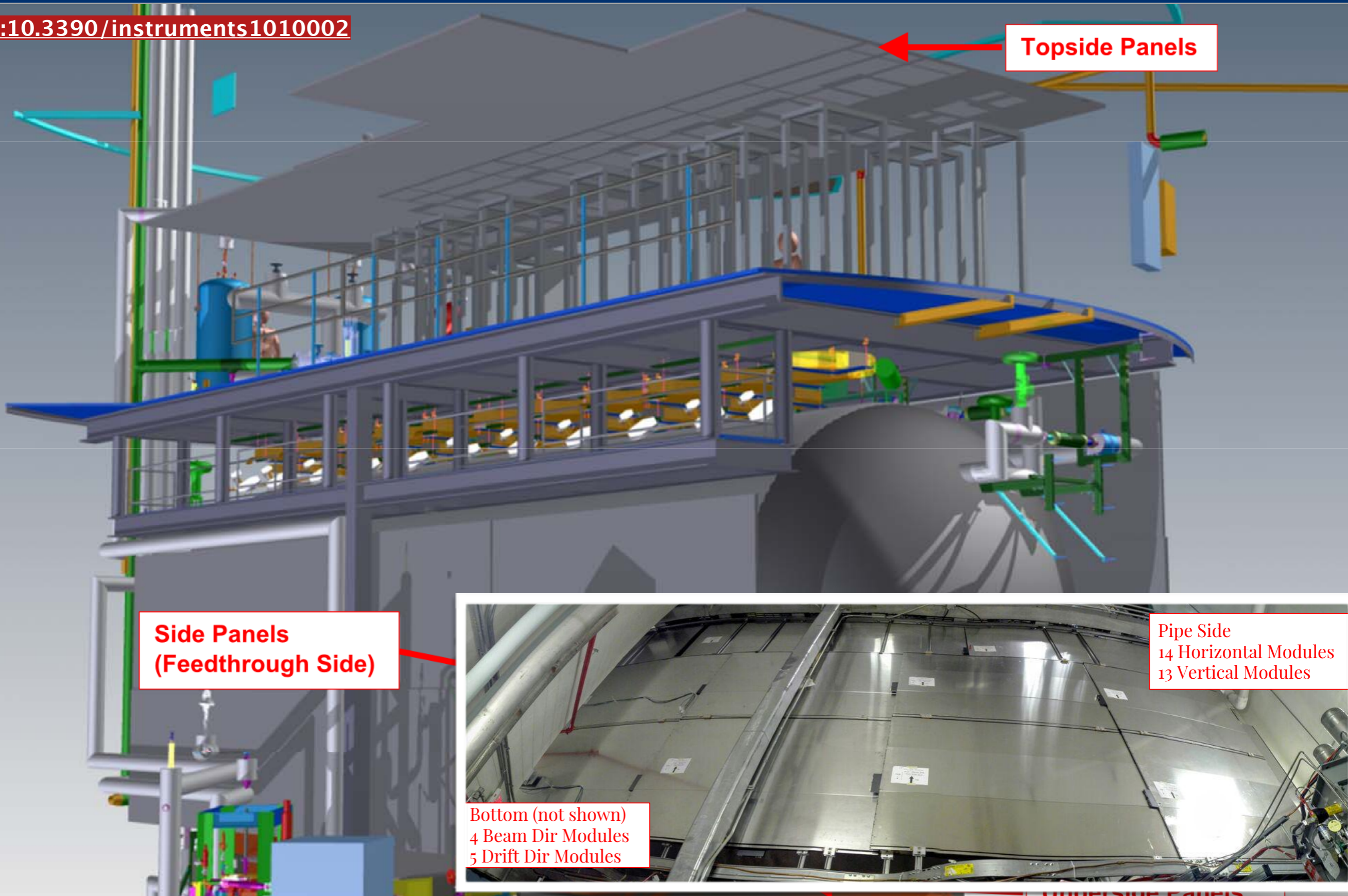
[doi:10.3390/instruments1010002](https://doi.org/10.3390/instruments1010002)





# Cosmic Ray Tagger

[doi:10.3390/instruments1010002](https://doi.org/10.3390/instruments1010002)



**Side Panels  
(Feedthrough Side)**

**Topside Panels**

**Pipe Side**  
14 Horizontal Modules  
13 Vertical Modules

**Bottom (not shown)**  
4 Beam Dir Modules  
5 Drift Dir Modules

**Underside Panels**



# Cosmic Ray Tagger

doi:10.3390/instruments1010002

← **Topside Panels**

**FT Side**  
6 Horizontal Modules  
7 Vertical Modules

**Side Panels  
(Feedthrough Side)**

**Pipe Side**  
14 Horizontal Modules  
13 Vertical Modules

**Bottom (not shown)**  
4 Beam Dir Modules  
5 Drift Dir Modules

→ **Underside Panels**



# Cosmic Ray Tagger

doi:10.3390/instruments1010002

← **Topside Panels**

me

**FT Side**  
6 Horizontal Modules



**Side Panels**  
(Feedthrough)

**Pipe Side**  
14 Horizontal Modules  
13 Vertical Modules

**Bottom (not shown)**  
4 Beam Dir Modules  
5 Drift Dir Modules

← **Underside Panels**



- The MicroBooNE experiment has a **broad physics program** and LArTPC R&D. The detector is up and running: 15 public notes and 5 papers already published.
- A small muon counter stack has the capabilities to assess **several performances of the LArTPC.**
- **Space-charge effect** must be taken into account when reconstructing LArTPC information, but it can be correctly simulated.
- The finding and **reconstruction efficiency of tracks in MicroBooNE is very high** (97.1%) and in good agreement with the simulation.
- The analysis has been included in a paper to be submitted to JINST.
- The **Cosmic Ray Tagger** provides increased coverage for efficiency studies and cosmic-ray background mitigation. Extremely important for future SBN program (ICARUS, SBND) and DUNE.





THANK YOU!