# Measuring $v_{\mu}$ and $\overline{v}_{\mu}$ oscillation parameters with MINOS

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### Introduction

#### Neutrino oscillations

- Two mass splittings
- Three mixing angles

MINOS can make precision measurements with neutrinos

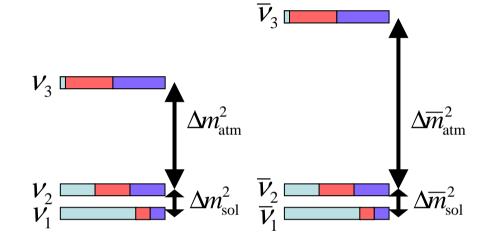
- Largest mass splitting
- $\triangleright$  Mixing angle  $\theta_{23}$

Corresponding antineutrino parameters are much less precisely known

- No direct precision measurements exist
- MINOS will be the first

A difference between the two would be very interesting

- Non-standard interactions
- CPT violation

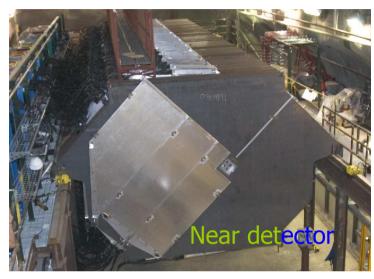


$$\begin{pmatrix} v_e \\ v_{\mu} \\ v_{\tau} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$





### The MINOS experiment



Beam neutrinos pass through two detectors

Detectors are magnetized to 1.3 T

- Allows the measurement of particle charge sign
- Also allows measurement of particle momenta

#### Two detectors to mitigate systematics

- e.g. neutrino flux or cross section mismodelings
- Use measured near detector data to predict what should be observed at the far detector before oscillations
- $\triangleright$  An observed  $v_{\mu}$  deficit at the far detector tells us about the oscillation parameters







### Beam composition

Charged current interactions in the near detector

- > 91.7%  $v_{\mu}$
- $\rightarrow$  7.1%  $\overline{v_{\mu}}$
- $\triangleright$  1.3%  $v_e$ +  $\overline{v}_e$

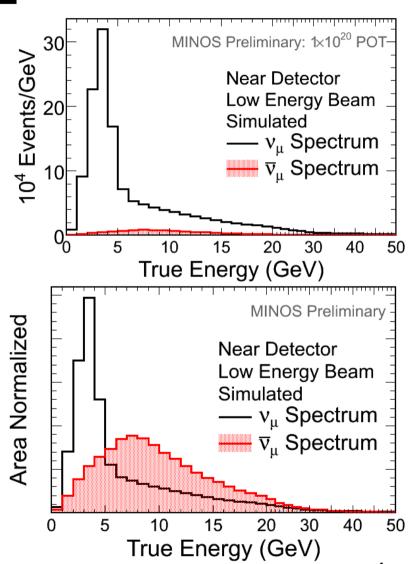
 $v_{\mu}$  and  $\overline{v}_{\mu}$  energy spectra are significantly different

 $v_{\mu}$  spectrum peaks at 3 GeV

Near the region of most oscillations

 $\overline{v}_{\!\scriptscriptstyle \prime\prime}$  spectrum peaks at 8 GeV

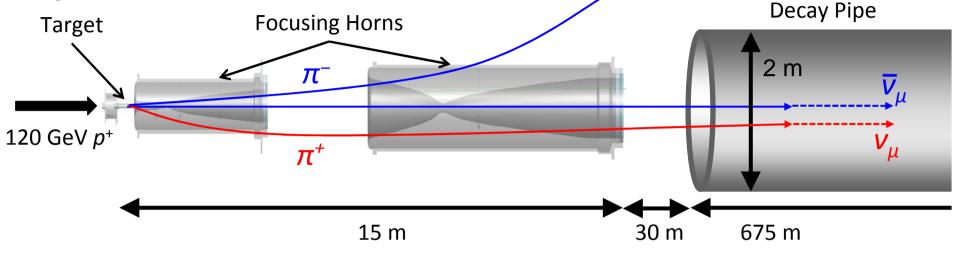
- > Away from the oscillation region
- Reducing the sensitivity to oscillations



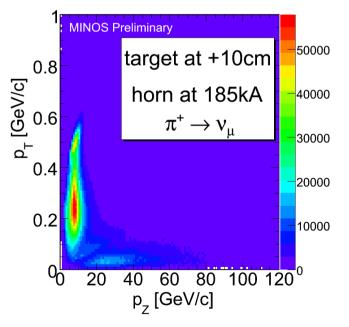


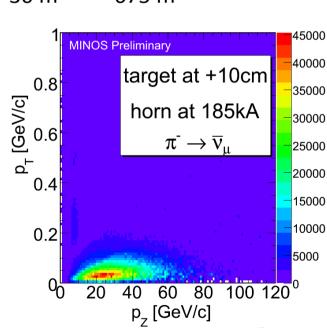


Why are the spectra different?



- $\nu_{\mu}$  spectrum is dominated by focused high-p<sub>T</sub> pions
- Majority of  $\overline{v}_{\mu}$  come from low-p<sub>T</sub> pions which travel down the centre of the horns where there is no magnetic field



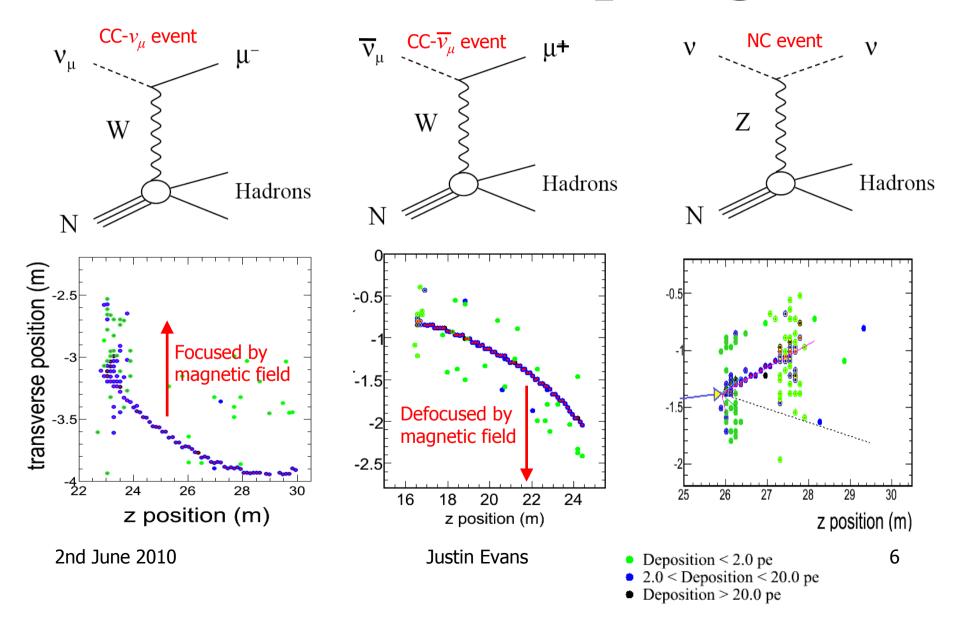


2nd June 2010



### \*UCL

### MINOS event topologies



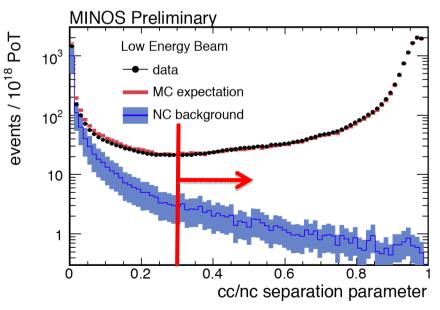


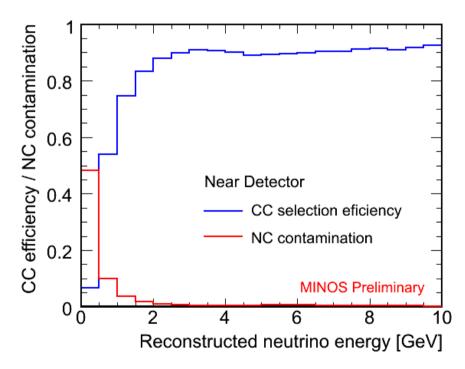


# CC $v_{\mu}$ event selection

Aim to separate charged and neutral current  $v_{\mu}$  interactions Four variables combined using a k-nearest-neighbour algorithm

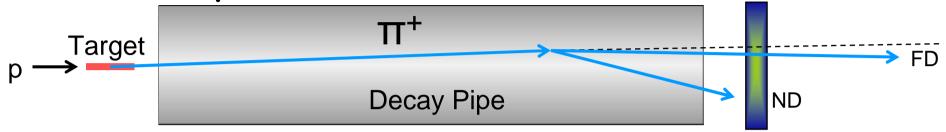
- > Track length
- Mean signal in track planes
- > Transverse track profile
- Signal fluctuation along the track







CC  $v_{\mu}$  beam extrapolation



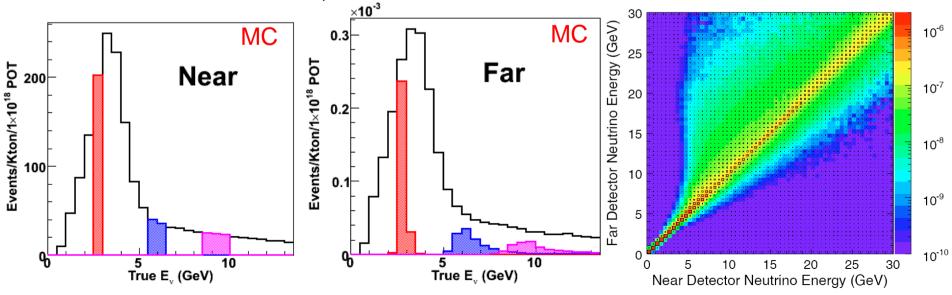
Use the measured ND energy spectrum to predict the FD spectrum:

Spread of pion decay directions smears neutrino energies

Different energy spectra at the two detectors

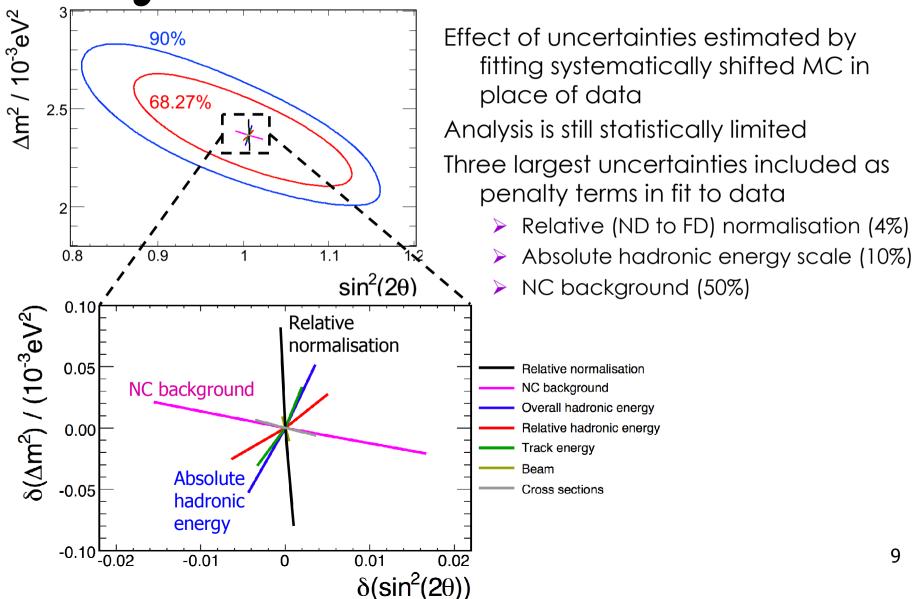
Encode the pion decay kinematics into a beam transfer matrix

Convert ND to FD spectrum





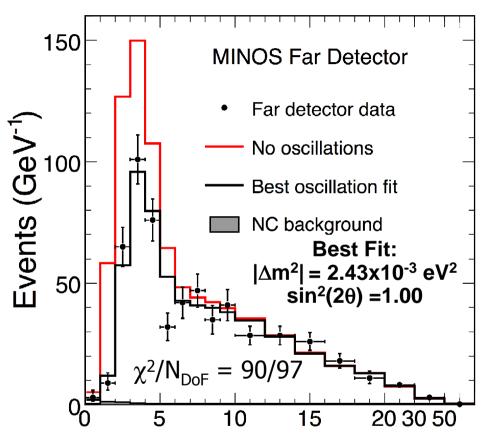
Systematic uncertainties







### Far detector data



FD data not looked at until the analysis was finalised

Expected 1065 ± 60 with no oscillations

Observed 848 events

Energy spectrum fit with the oscillation hypothesis

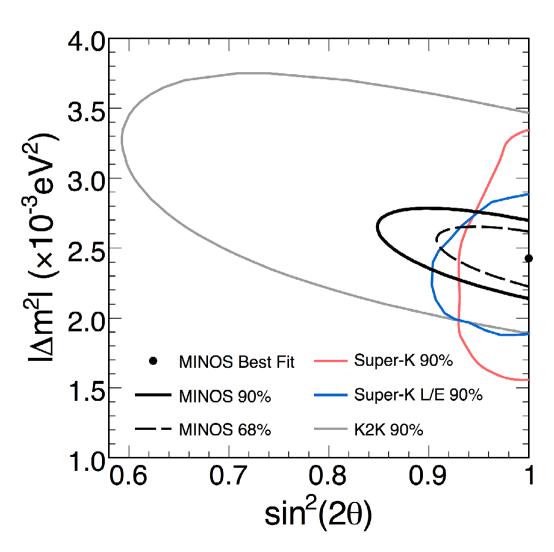
$$P(\nu_{\mu} \to \nu_{\tau}) = \sin^2(2\theta) \sin^2\left(\frac{1.27\Delta m^2 L}{E}\right)$$

Reconstructed neutrino energy (GeV)





### Allowed region



#### Constrained fit

- $\Delta m^2$  | =(2.43±0.13)x10<sup>-3</sup> eV<sup>2</sup> (68% C.L.)
- $\rightarrow$  sin<sup>2</sup>(2 $\theta_{23}$ ) > 0.90 (90% C.L.)
- $> \chi^2/N_{DOF} = 90/97$

#### Unconstrained fit

- $\rightarrow$   $|\Delta m^2| = 2.33 \times 10^{-3} \text{ eV}^2$
- $> \sin^2(2\theta_{23}) = 1.07$
- $\rightarrow$   $\Delta \chi^2 = -0.6$

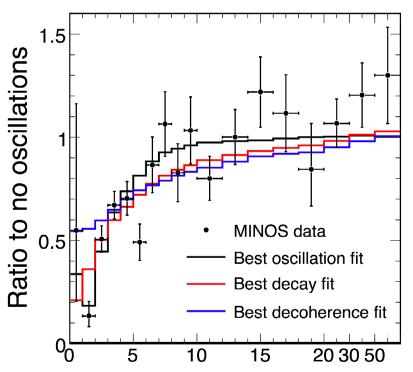




### Alternative models

#### Decay:

Two alternative disappearance models are disfavoured



Reconstructed neutrino energy (GeV)

$$P_{uu} = \left(\sin^2(\theta) + \cos^2(\theta) \exp(-\alpha L/2E)\right)^2$$

V. Barger et al., PRL82:2640(1999)  $\chi^2/\text{ndof} = 104/97$   $\Delta\chi^2 = 14$ 

disfavored at  $3.7\sigma$ 

#### **Decoherence:**

$$P_{\mu\mu} = 1 - \frac{\sin^2 2\theta}{2} \left( 1 - \exp\left(\frac{-\mu^2 L}{2E_v}\right) \right)$$

G.L. Fogli et al., PRD67:093006 (2003) 
$$\chi^2/\text{ndof} = 123/97$$
  $\Delta\chi^2 = 33$ 

disfavored at  $5.7\sigma$ 

### The future

#### MINOS has doubled its dataset

New results will be released this summer.

#### We are also incorporating analysis improvements

Using events outside the fiducial volume

Interactions in the detector and surrounding rock

Adding in events with poorly measured charge

> Recovers around half the events at low energy

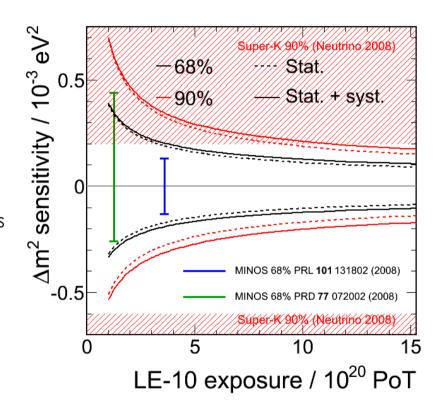
A new selection variable optimized for low energies

Grouping events according to their resolution

New hadronic shower energy estimator

- Based on a k-nearest-neighbour technique
- Significantly improves shower energy resolution

These improvements can increase our  $\Delta m^2$  sensitivity by 15%







# Selecting $\overline{v}_{\mu}$

Selecting events with a track reconstructed with positive charge

We must work harder to select a pure sample of  $\overline{v_u}$  events

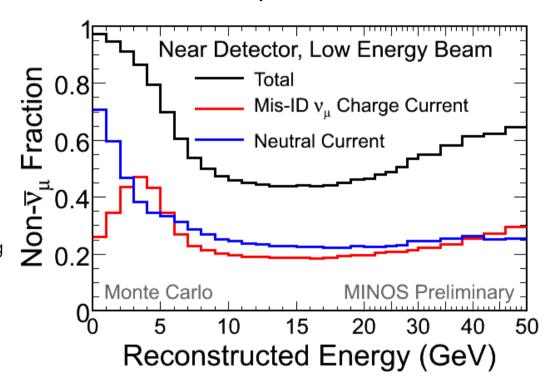
They are only 7% of all our CC events

Large backgrounds from other event-types

- ightharpoonup Mis-identified CC-  $v_{\mu}$  with wrong charge sign
- NC events where another particle fakes a track

Three additional selection criteria are used

All events with a positive curvature track







**MINOS Preliminary** 

# Selecting $\overline{v}_{\mu}$

Events

103

20

-- 2.9 ×10<sup>20</sup> POT

Near Detector

MC w/ flux error

MC background

NC component Mis-ID v., CC

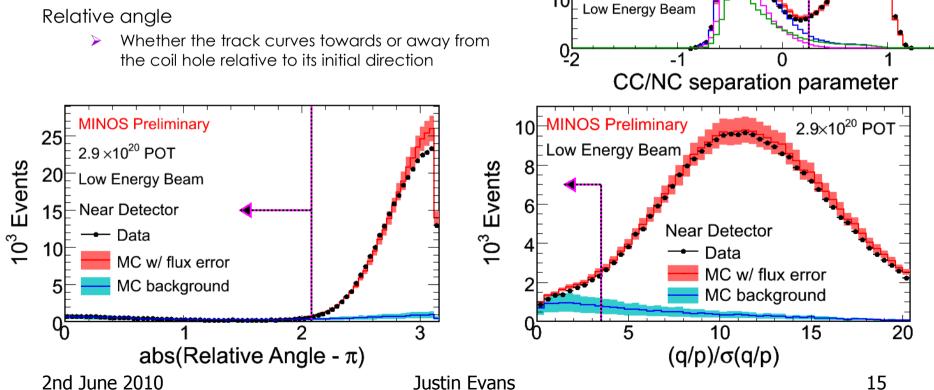
Data

#### CC/NC separator

- Likelihood-based with 3 probability density **functions**
- Removes both NC and mis-identified CC events

Track fit charge sign significance

#### Relative angle







### Selecting $\overline{v}_{\mu}$

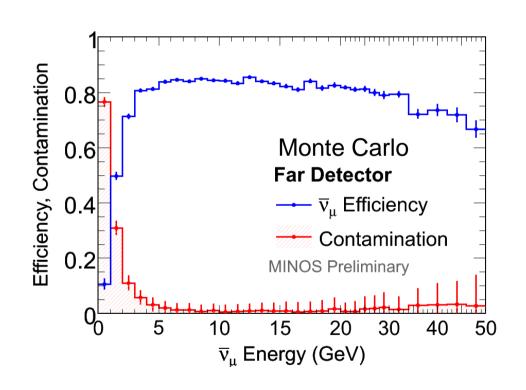
After all selection cuts:

Far detector

- > Efficiency of 82%
- Contamination of 3%

Optimized for physics sensitivity to oscillations with neutrino best fit parameters

Lower contamination is possible but at the cost of efficiency







### $\overline{v}_{\mu}$ far detector spectrum

Observe 42  $\overline{v}_{\mu}$ -CC events at the far detector

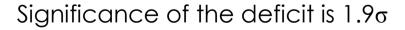
First direct observation of  $\overline{v}_{\mu}$  in an accelerator long-baseline experiment

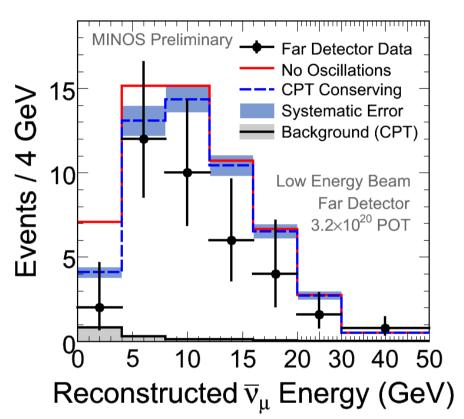
Predicted events with oscillations with the best-fit neutrino oscillation parameters

> 58.3 ± 7.6 (stat.) ± 3.6 (syst.)

Predicted with no oscillations

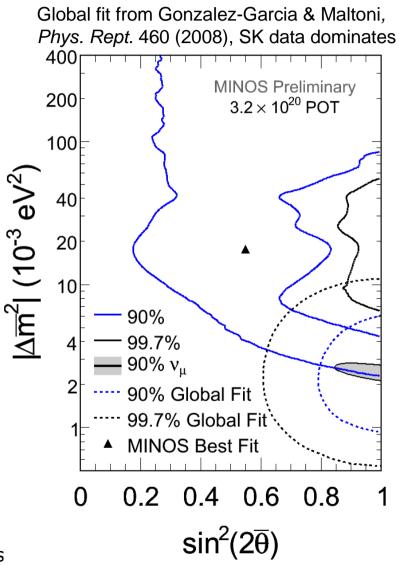
 $\rightarrow$  64.6 ± 8.0 (stat.) ± 3.9 (syst.)







- Feldman-Cousins contour, including systematics
- Null oscillation hypothesis excluded at 99%
- Previously allowed regions excluded at high confidence
- Allowed region from the  $v_{\mu}$  measurement is within the 90% contour





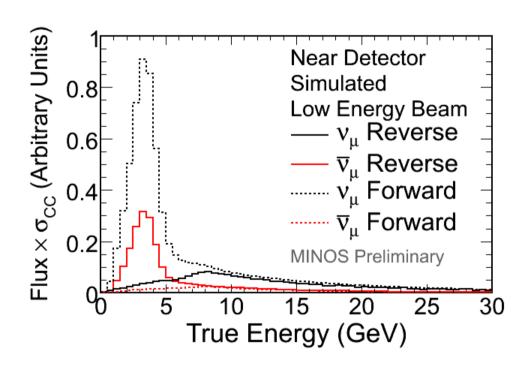


### Dedicated $\overline{v}_{\mu}$ running

MINOS has now taken data with a dedicated  $\overline{v}_{\mu}$  beam

- By reversing the current in the focusing horns
- From October 2009 to March 2010

Significantly enhances the  $\overline{v}_{\mu}$  flux in the oscillation signal region

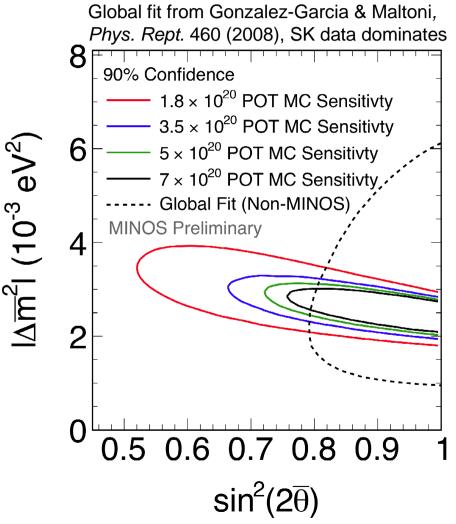






# Dedicated $\overline{v}_{\mu}$ running

- We will make the first ever precision measurement of the  $\overline{v}_{\mu}$  oscillation parameters
- Significantly reduce the uncertainty on  $\Delta \overline{m}^2$
- Results will be released this summer







### Summary

MINOS has made the world's most accurate measurement of the atmospheric neutrino mass splitting

- $\rightarrow$   $|\Delta m^2| = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2 (68\% \text{ c.l.})$
- $\Rightarrow$  sin<sup>2</sup>(2 $\theta_{23}$ ) > 0.90 (90% c.l.)

Alternative models disfavoured

 $\triangleright$  Decay at 3.7 $\sigma$ , decoherence at 5.7 $\sigma$ 

First direct detection of  $\overline{v}_{\mu}$  in an accelerator long-baseline experiment

A new measurement of the  $v_{\mu}$  oscillation parameters will be released this summer

- Double the dataset
- Analysis improvements

MINOS has taken data with a dedicated antineutrino beam

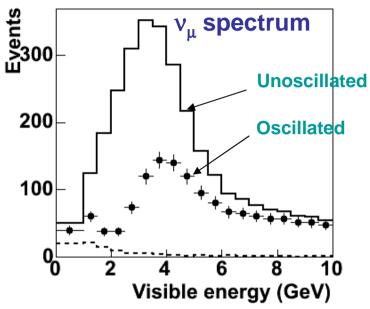
- The first precision measurement of the  $\overline{v_{\mu}}$  oscillation parameters
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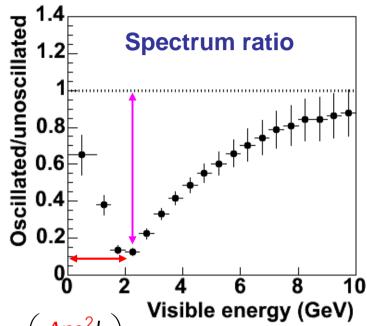
## Backup

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### Measuring oscillations





$$P(v_{\mu} \rightarrow v_{\mu}) = 1 - \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{E}\right)$$

Near detector measures the energy spectrum before oscillations occur At the FD we see an energy-dependent deficit due to oscillations

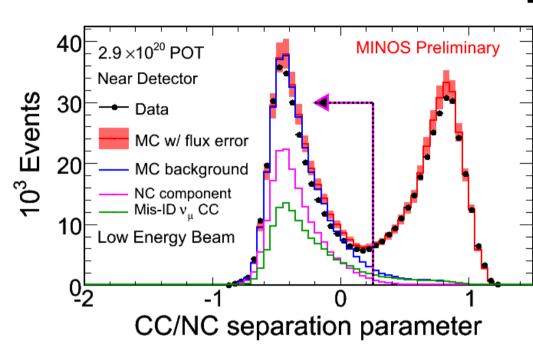
- ightharpoonup Position of the dip give  $\Delta m^2$
- $\triangleright$  Depth of dip gives  $\sin^2(2\theta)$

An alternative disappearance model (e.g. decay or decoherence) would give a different energy-dependence





### CC/NC separator



Likelihood-based with 3 Probability Density Functions:

- > Track length
- Pulse height fraction in track
- > Pulse height per plane

Use CC/NC separation parameter developed for previous analyses

Cut removes both NC and mis-identified v<sub>u</sub> CC events

Mis-ID  $v_{\mu}$  CC events tend to be inelastic events where the muon is obscured by a large hadronic shower

high-y events



Errors that are the same as the  $v_{\mu}$  analysis

- Normalization: ±4%
  Relative reconstruction eff., detector livetime and mass
- Muon energy: range ±2%, curvature ±4% (Error from curvature increased slightly due to exiting tracks)
- $\triangleright$  Beam extrapolation:  $1\sigma$  error band from beam fit
- > Relative shower energy: ±3%
- > Absolute shower energy: ±10%

Errors that are specific for this analysis

- Decay pipe production
- > Background: ±50%