



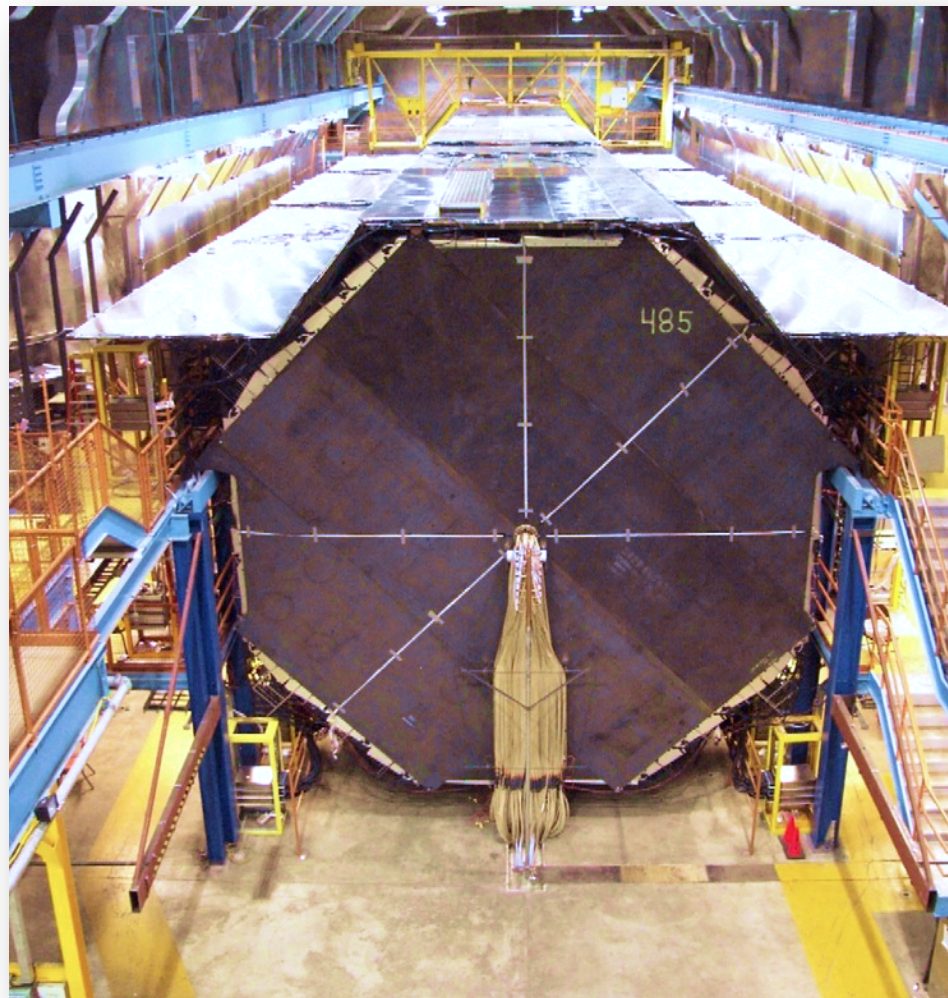
New Results from MINOS

Jeff Nelson



William & Mary

Fermilab User's Meeting
June, 12th 2012

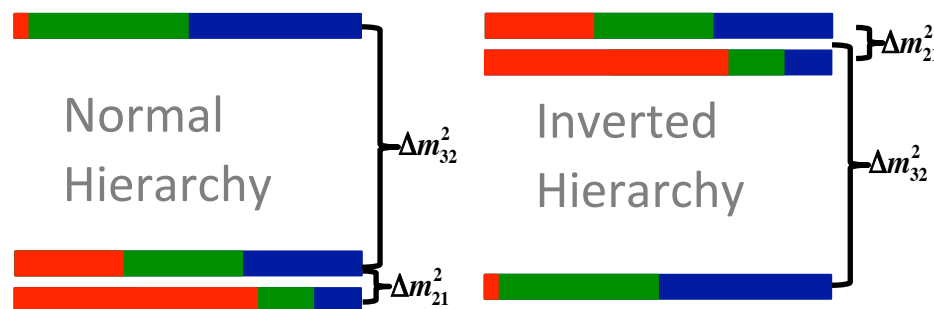




MINOS physics goals

- Test the $\nu_\mu \rightarrow \nu_\tau$ oscillation hypothesis for atmospheric neutrino disappearance
 - > Measure precisely $|\Delta m_{32}^2|$ and $\sin^2 2\vartheta_{23}$
- Test exotic models
 - Decay, decoherence, sterile neutrinos, velocity
- Search for $\nu_\mu \rightarrow \nu_e$ oscillations
- Compare ν to $\bar{\nu}$ oscillations
- Atmospheric neutrinos
- Neutrino interaction physics
- Cosmic ray physics

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



$$\Delta m_{32}^2 = m_3^2 - m_2^2$$

Useful Approximations

ν_μ disappearance (2 flavors):

$$P(\nu_\mu \rightarrow \nu_x) = 1 - \sin^2 2\vartheta_{23} \sin^2(1.27 \Delta m_{32}^2 L/E)$$

ν_e appearance (first order):

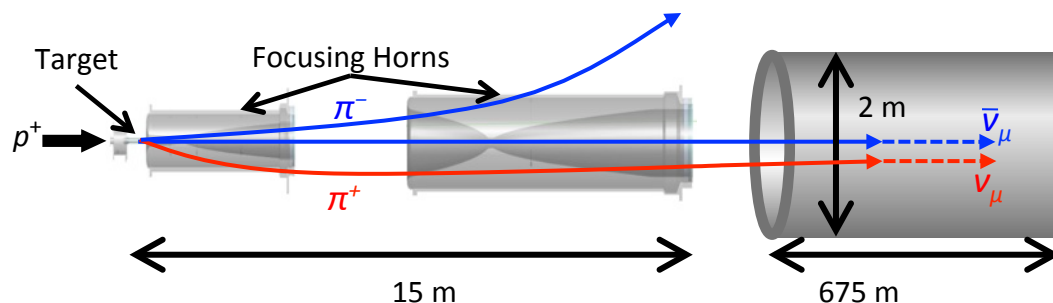
$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 \vartheta_{23} \sin^2 2\vartheta_{13} \sin^2(1.27 \Delta m_{31}^2 L/E)$$

where L , E are experimental parameters &

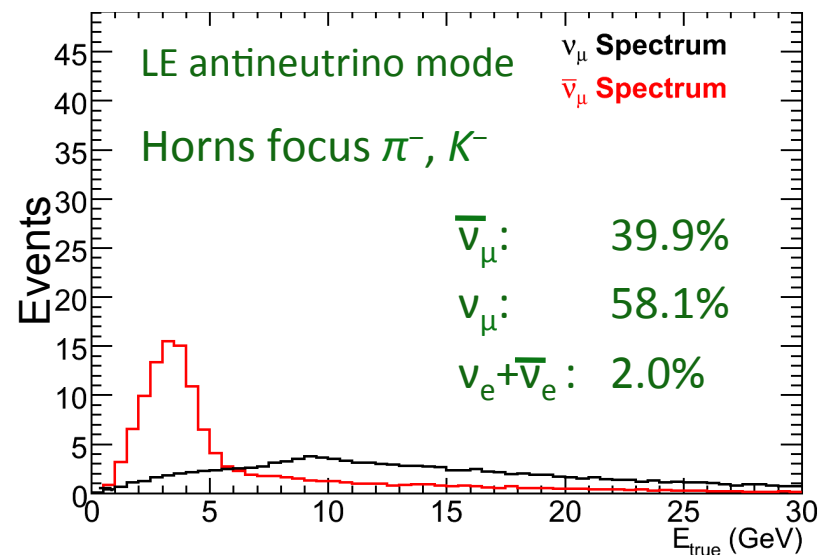
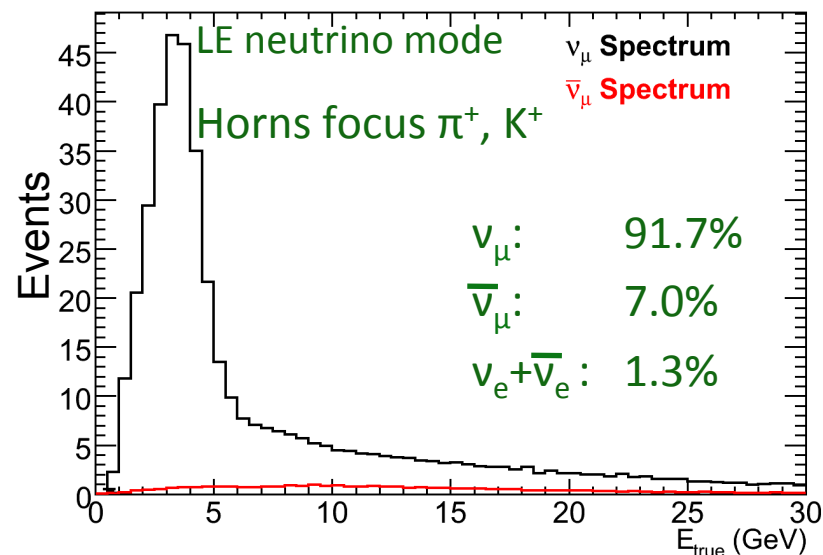
ϑ_{23} , ϑ_{13} , Δm_{32}^2 are to be determined



Producing the NuMI neutrino beam



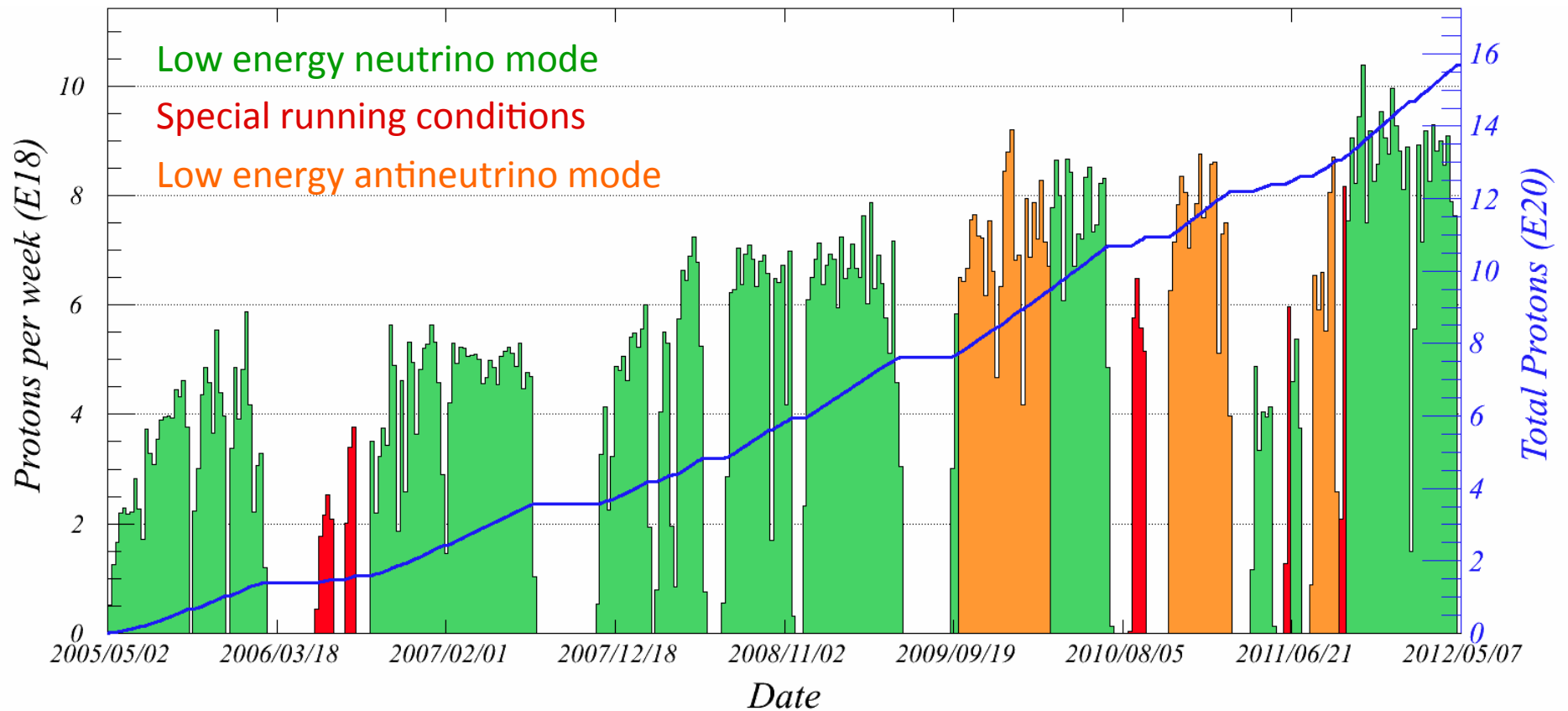
- 120 GeV protons strike target
 - > 10 μ s pulse every 2.2 s
 - > Typical running at 340 kW
 - > 3.5×10^{13} protons/pulse
- 2 magnetic horns focus secondary π/K
 - > Moveable target & horn provides variable beam energy
 - > Majority of running in low-energy tune
 - > Can focus positives (neutrino mode) or negatives (antineutrino mode)





The MINOS LE exposure (2005-2012)

Thanks for all the protons!

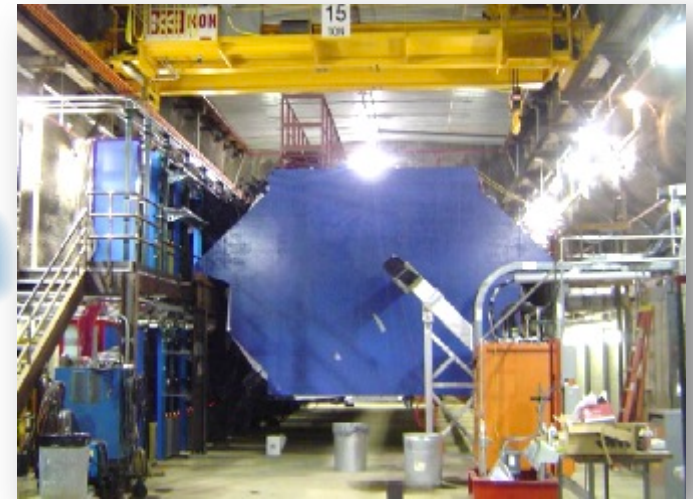


15.6×10^{20} protons on target [POT]



MINOS: 2 magnetized iron-scintillator tracking calorimeters

NIMA 596 190 (2008)

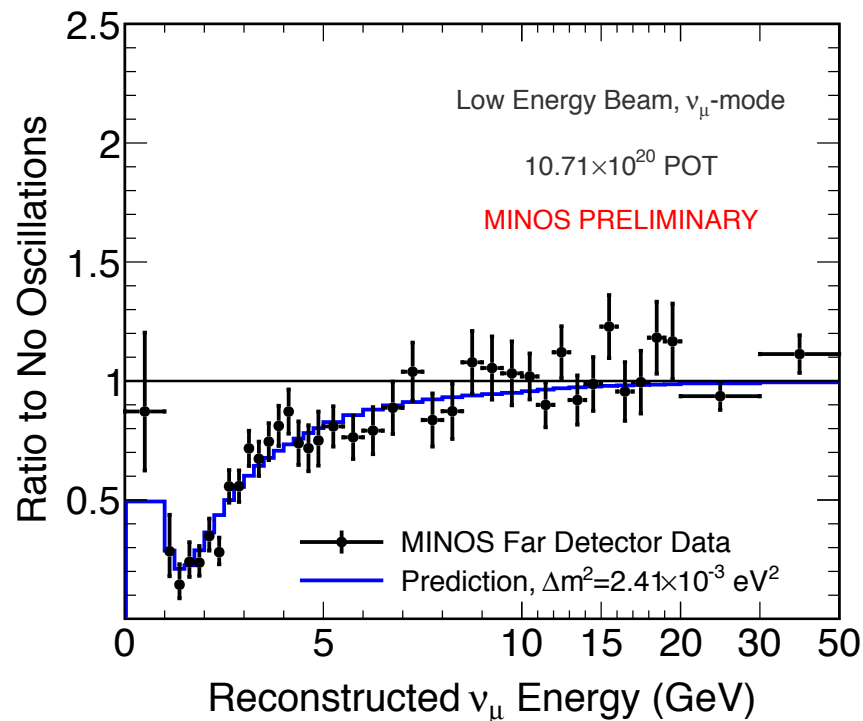
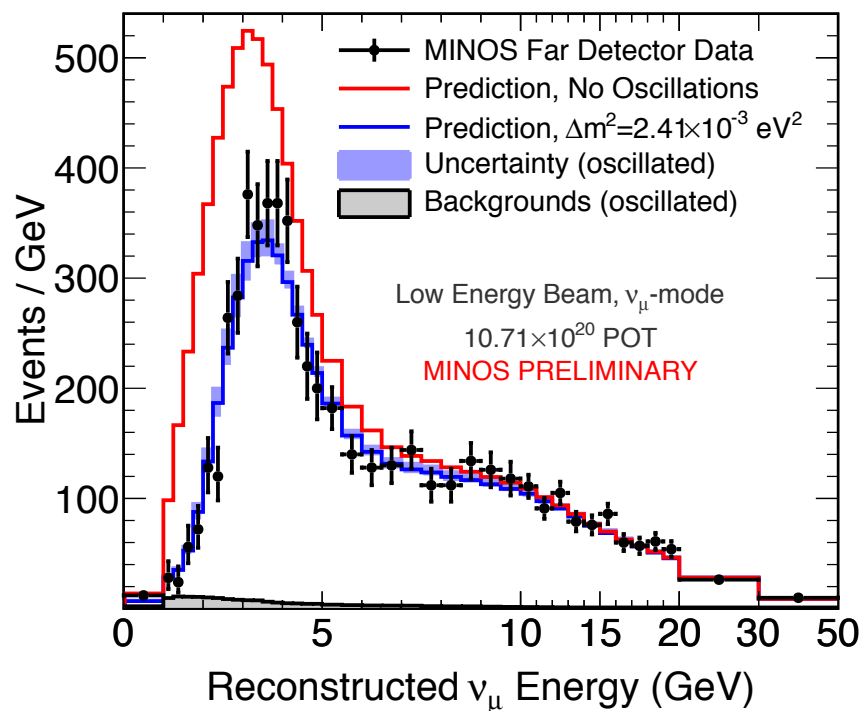


Functionally equivalent

- Magnetized steel planes $\langle B \rangle = 1.3 \text{ T}$
- $1 \times 4.1 \text{ cm}^2$ scintillator strips
- 2.54 cm steel sheets
- Moliere radius = 3.7 cm
- Sampling = 1.4 radiation lengths



Muon neutrino oscillation results



Observed: 2894

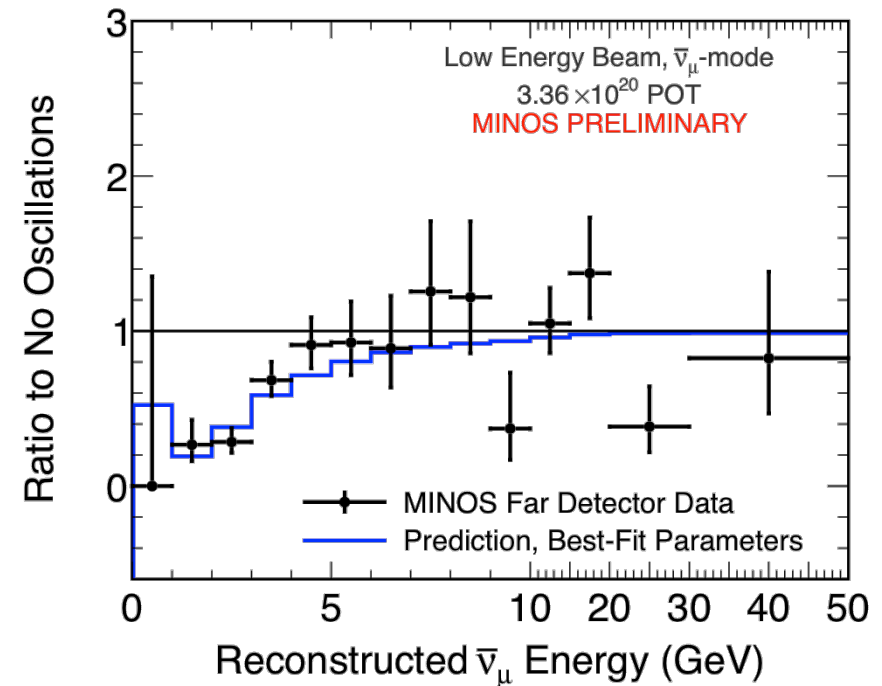
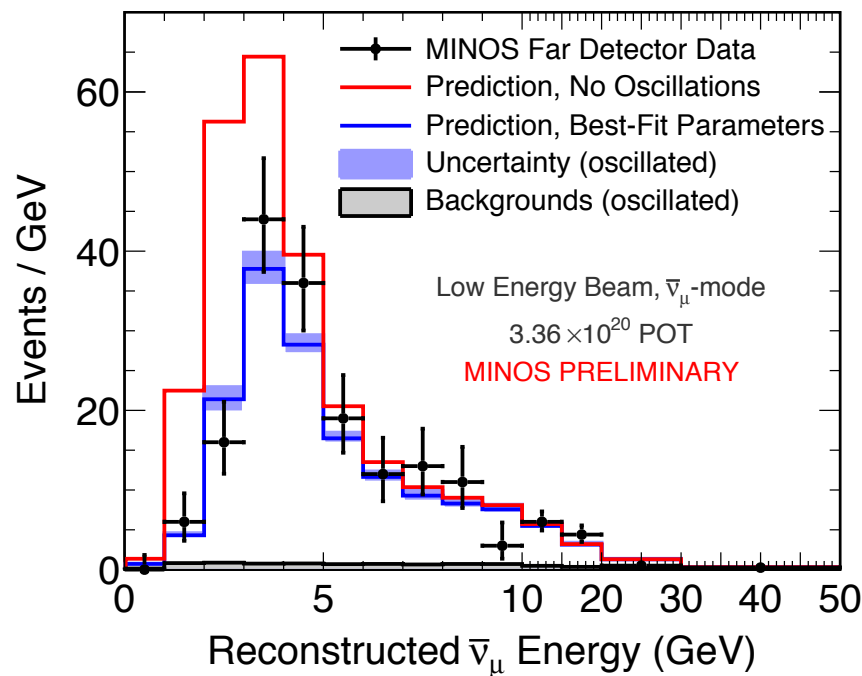
Predicted (no oscillations): 3564

$$|\Delta m^2| = 2.41^{+0.11}_{-0.10} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) = 0.937^{+0.044}_{-0.047}$$



Muon antineutrino oscillation results



Observed: 226

Predicted (no oscillations): 312

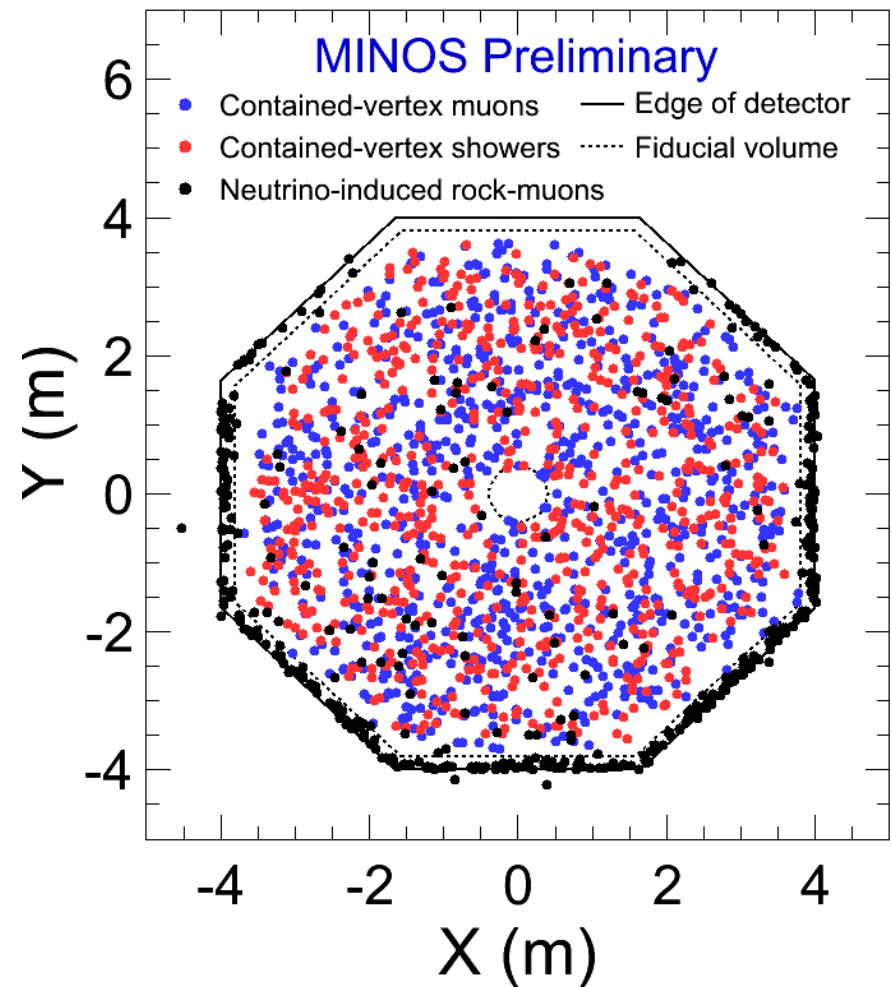
$$|\Delta \bar{m}^2| = 2.64^{+0.28}_{-0.27} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) > 0.78 \text{ (90\% C.L.)}$$



MINOS atmospheric neutrinos & antineutrinos

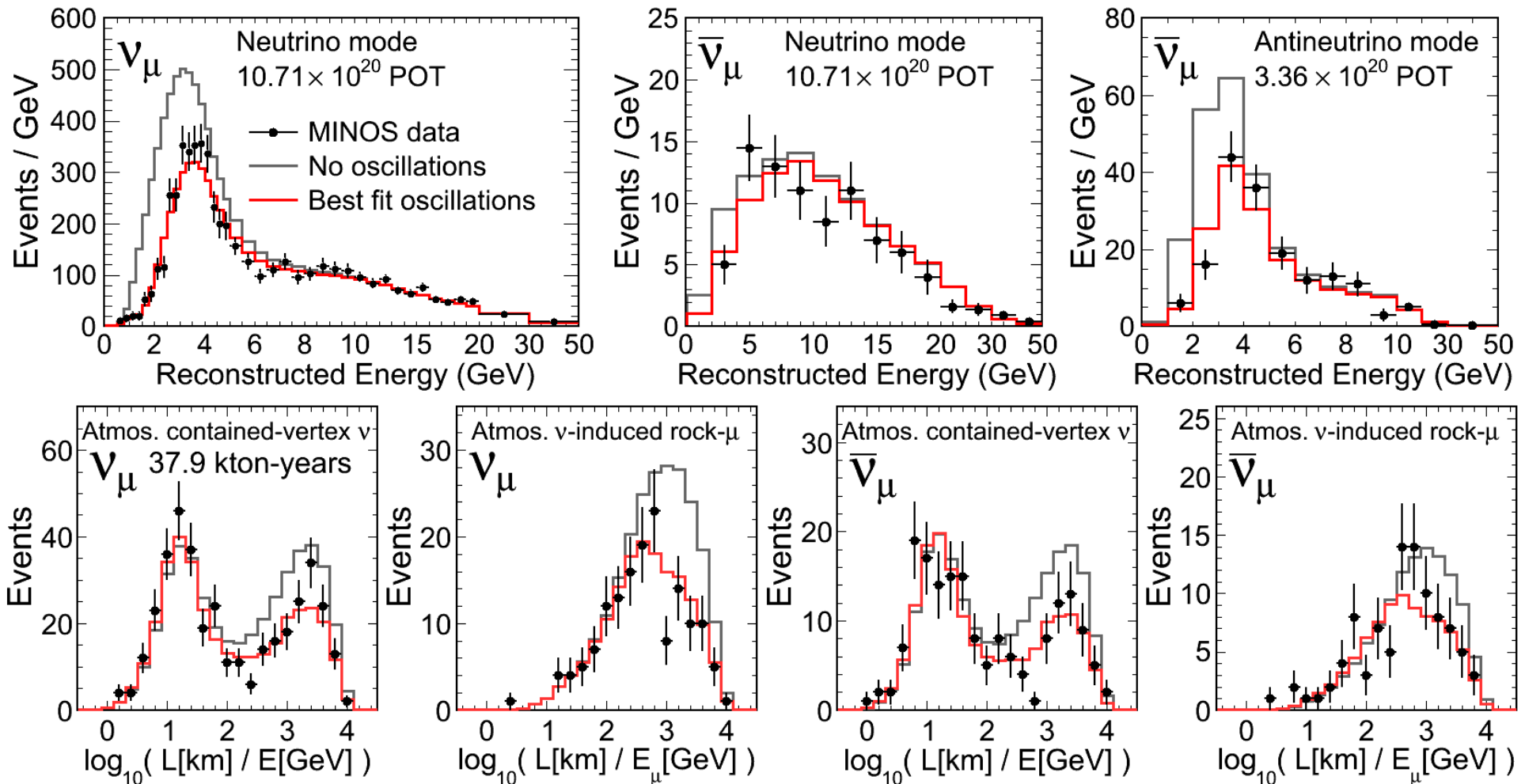
- Magnetized far detector with scintillator veto shield on top and sides
 - > Located 705m underground (2070 mwe)
 - > Event-by-event muon charge identification
- 2553 live-days or 37.9 kt-yrs
- Observed events
 - > 905 contained-vertex muons
 - > 466 neutrino-induced rock-muons
 - Charged-current muon neutrino interactions
 - > 701 contained-vertex showers
 - Mainly charged-current electron neutrino & neutral-current interactions





MINOS beam & atmospheric neutrinos & antineutrinos combined in a fit

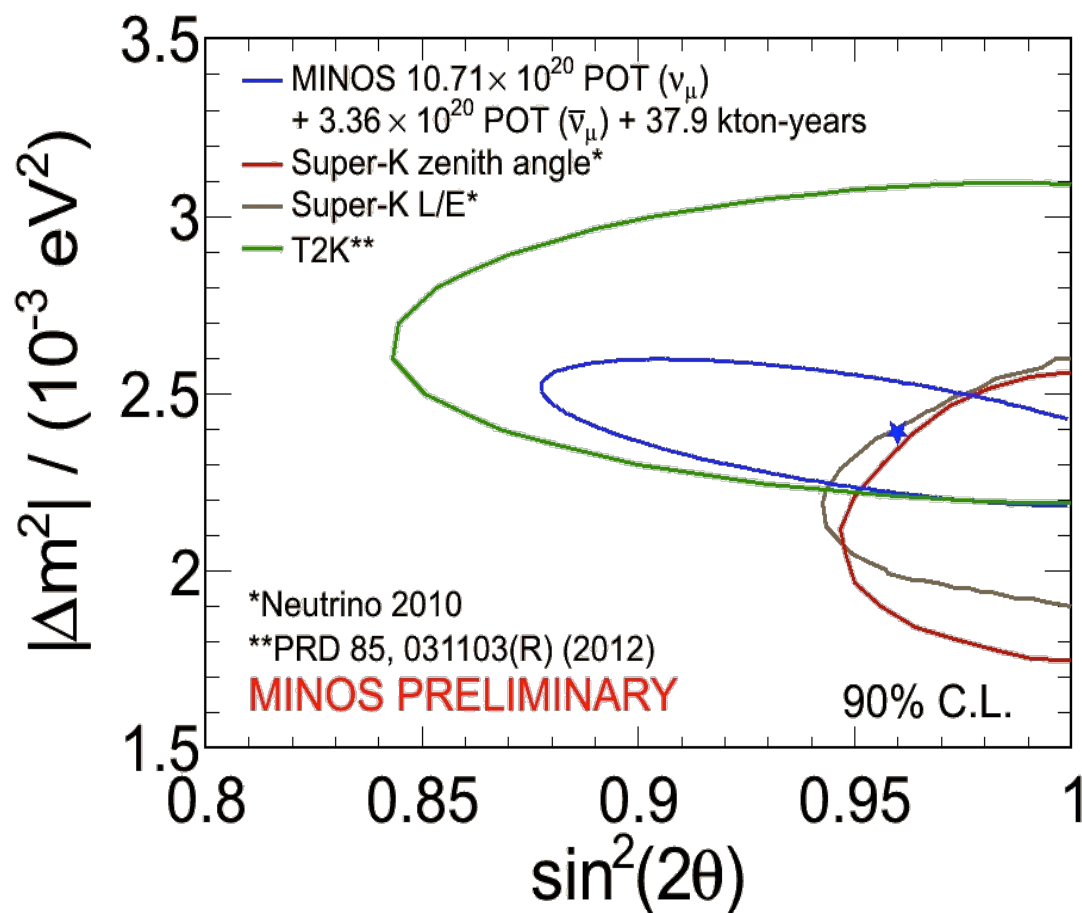
MINOS PRELIMINARY





MINOS full-exposure results vs the world's results

2-flavour oscillation parameters Δm^2 and $\sin^2 2\theta$ are determined by applying a maximum likelihood fit to the MINOS beam and atmospheric neutrino data

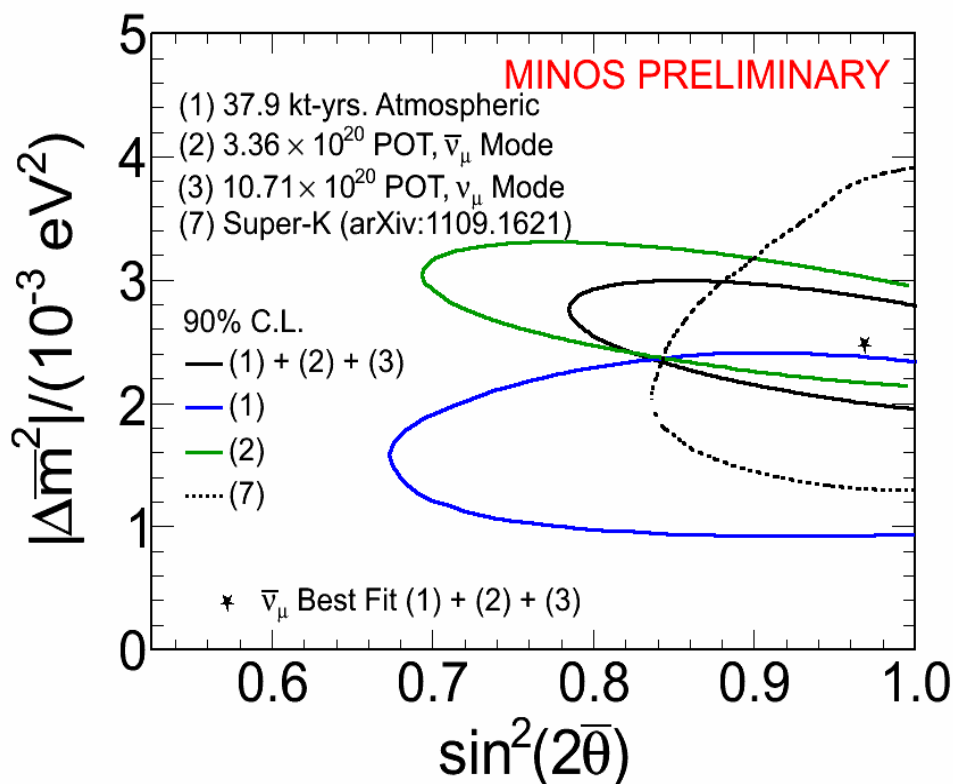


$$|\Delta m^2| = 2.39^{+0.09}_{-0.10} \times 10^{-3} \text{ eV}^2$$
$$\sin^2(2\theta) = 0.957^{+0.035}_{-0.036}$$
$$\sin^2(2\theta) > 0.90 \text{ (90\% C.L.)}$$



MINOS full-exposure results neutrinos vs antineutrinos

Extended fit allows different oscillation parameters for neutrinos and antineutrinos, and confidence limits are placed on the antineutrino oscillation parameters



$$|\Delta \bar{m}^2| = 2.48^{+0.22}_{-0.27} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) > 0.83 \text{ (90\% C.L.)}$$

$$|\Delta m^2| = 2.39^{+0.09}_{-0.10} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) = 0.957^{+0.035}_{-0.036}$$

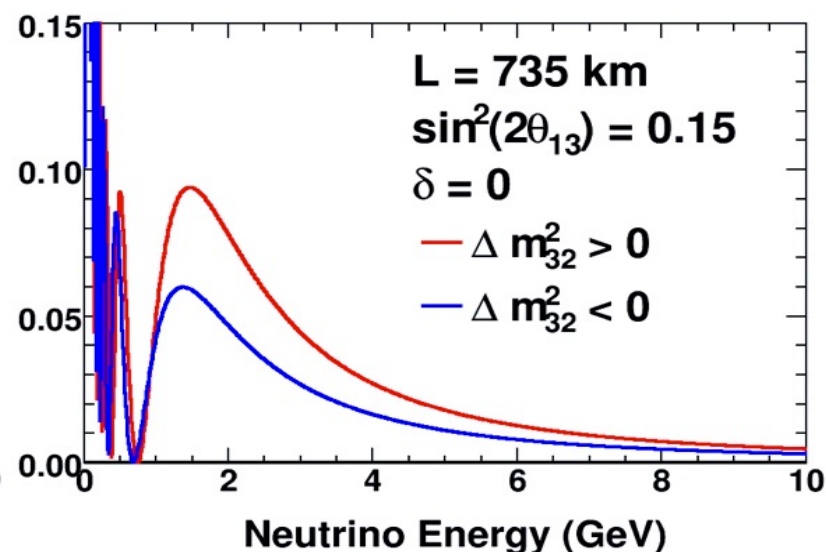
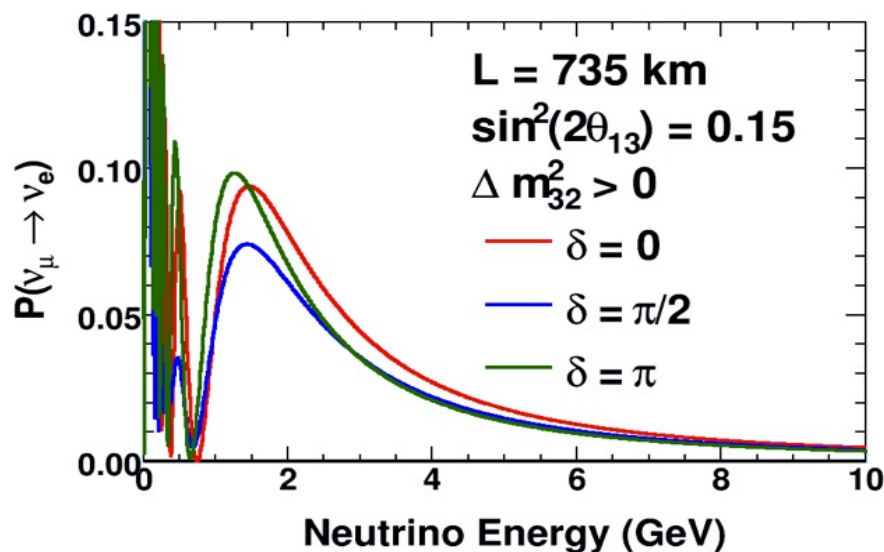
$$\sin^2(2\theta) > 0.90 \text{ (90\% C.L.)}$$

The new data resolves the tension between neutrino & antineutrino parameters



3-flavor oscillation phenomenology

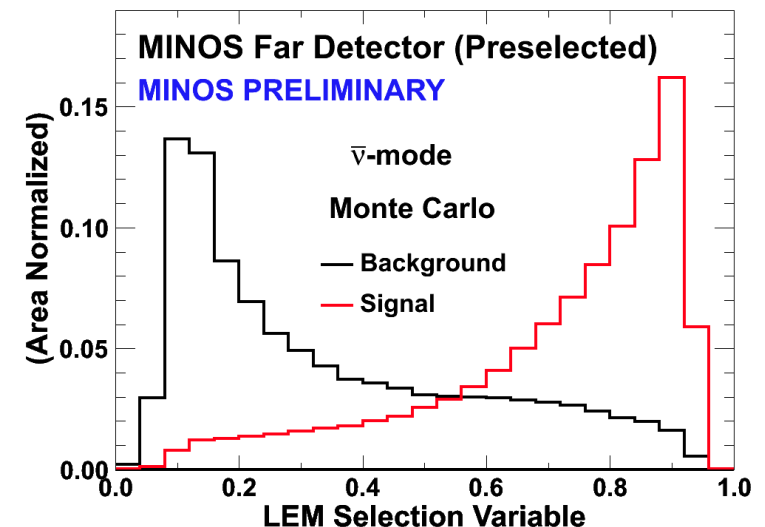
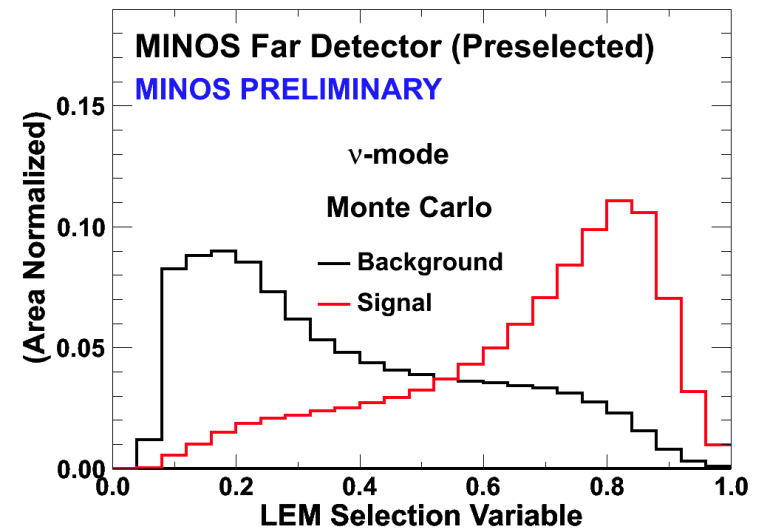
- Super-K / MINOS tell us $|\Delta m^2_{32}| \approx 3 \times 10^{-3} \text{ eV}^2$ $\theta_{23} \approx 45^\circ$
- SNO / Kamland tell us $\Delta m^2_{21} \approx 8 \times 10^{-5} \text{ eV}^2$ $\theta_{12} \approx 34^\circ$
- DC+RENO+DB tell us $\theta_{13} \approx 9^\circ$
- Examples of the effects of the higher order terms for MINOS





Search for electron appearance in MINOS

- Due to the coarse sampling we cannot identify electrons in an event-by-event basis
- Particle ID discriminator
 - > Library Event Matching (LEM)
 - > An ANN informed by properties of similar MC library events as defined by a charge overlap algorithm
- Define a signal region for events with $LEM > 0.6$ for a binned fit
 - > Optimal counting analysis is $LEM > 0.7$





Appearance of electrons

10.6×10^{20} POT (ν mode) + 3.3×10^{20} ($\bar{\nu}$ mode)

ν mode

Observe 88 events *

Expected

69.1 (background)

26.0 (signal**)

$\bar{\nu}$ mode

Observe 12 events *

Expected

10.5 (background)

3.1 (signal**)

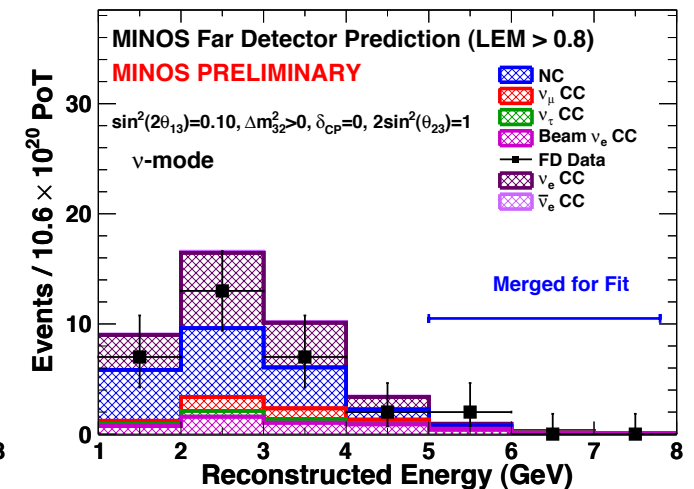
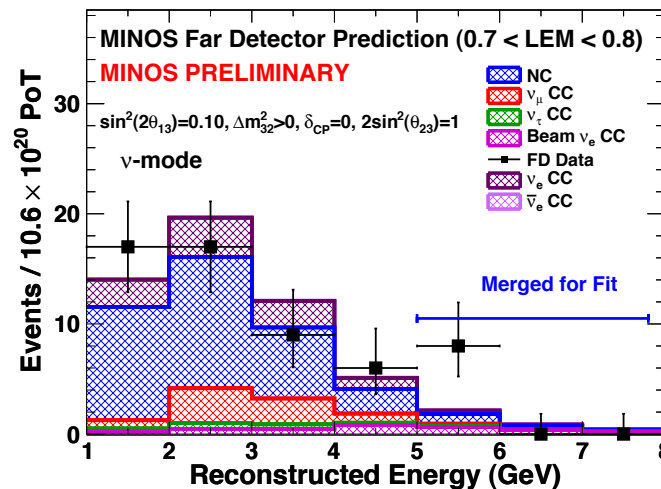
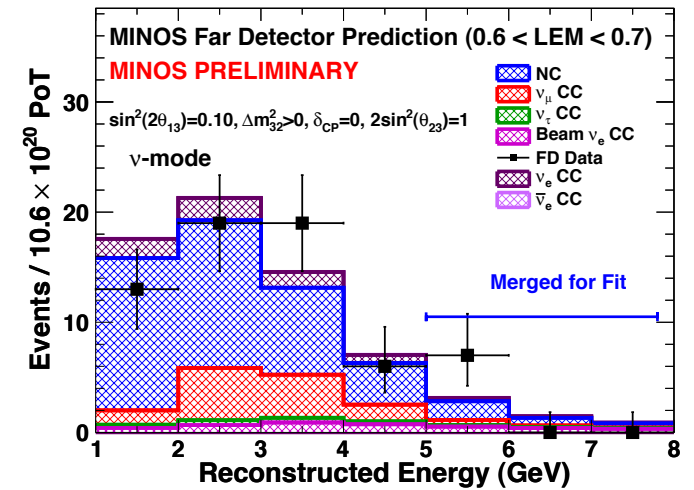
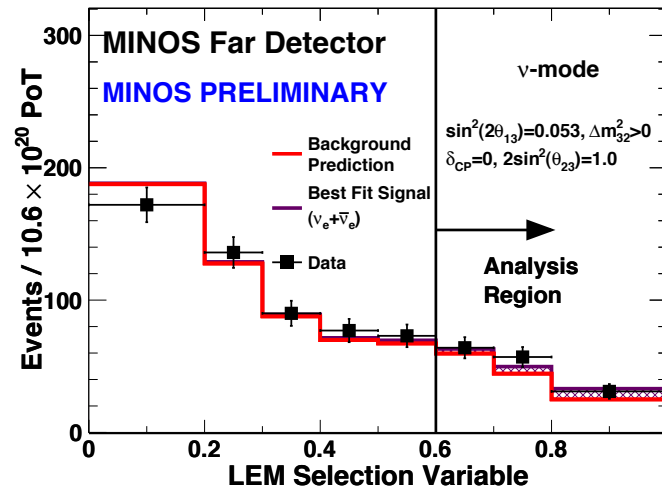
Fit to 15 LEM/energy bins

* for LEM > 0.7

** assuming

$\sin^2(2\theta_{13})=0.1,$

$\delta=0, \Delta m^2 > 0$

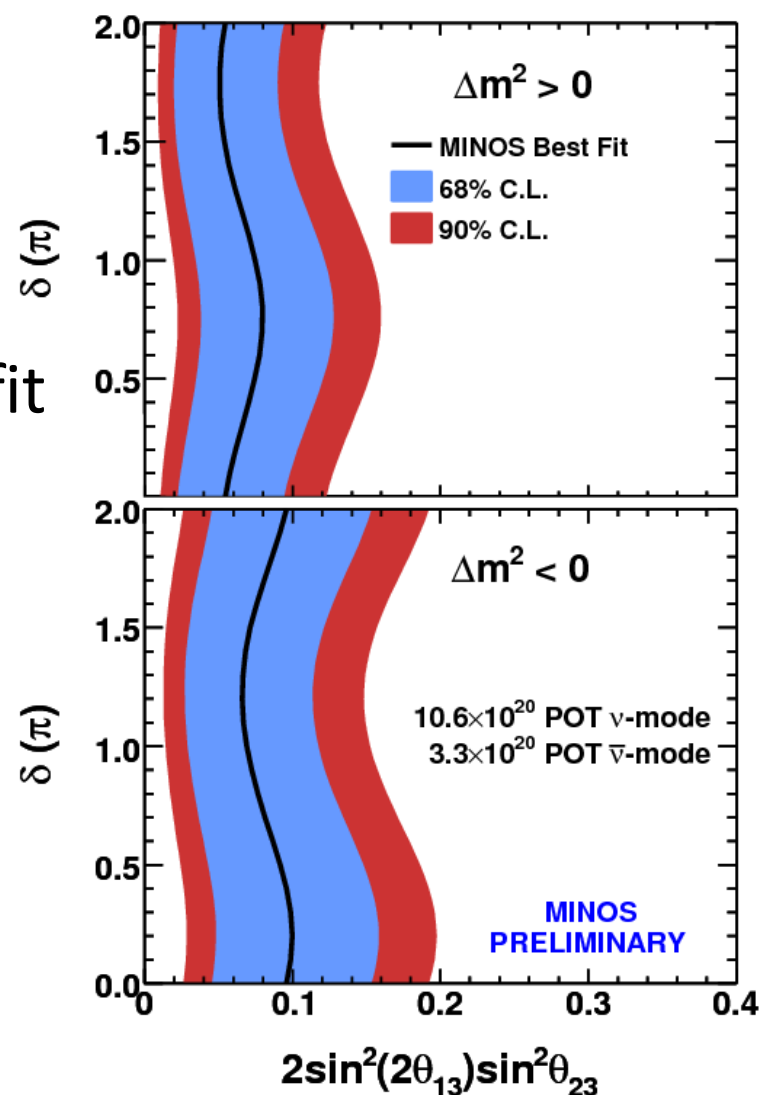




Results for appearance of electrons with both ν and $\bar{\nu}$ running combined

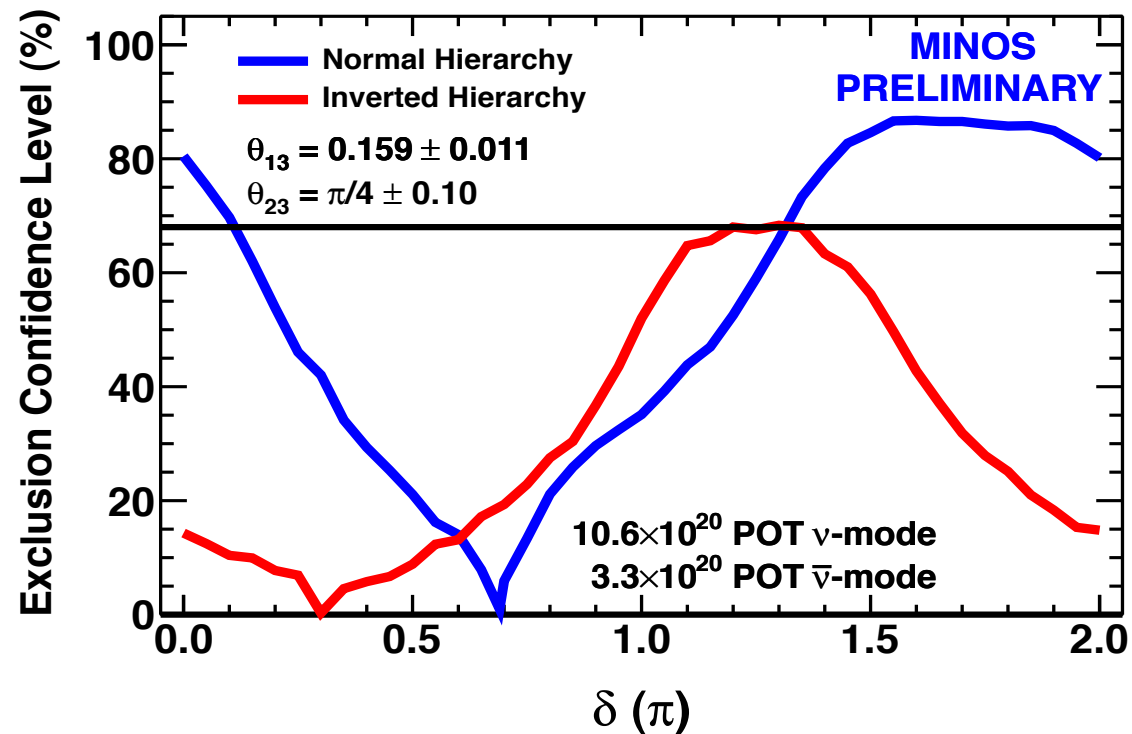
for $\delta_{CP} = 0$, $\sin^2(2\theta_{23}) = 1$,
with normal (inverted) hierarchy
 $\sin^2(2\theta_{23}) = 0.053$ (0.094) at best fit
 $0.01 < \sin^2(2\theta_{23}) < 0.12$ at 90% C.L.
($0.03 < \sin^2(2\theta_{23}) < 0.19$)

$\sin^2(2\theta_{23}) = 0$ excluded at 96% C.L.





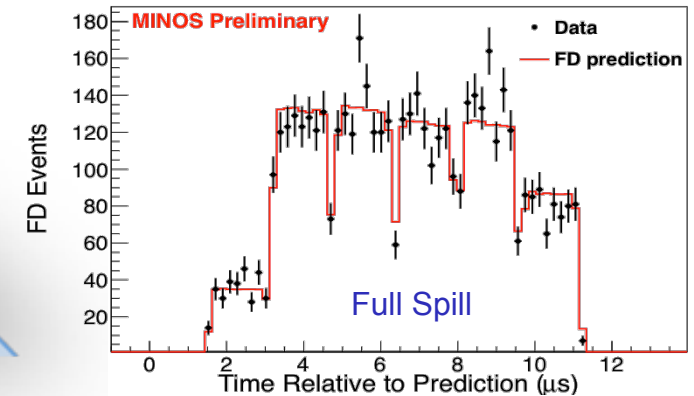
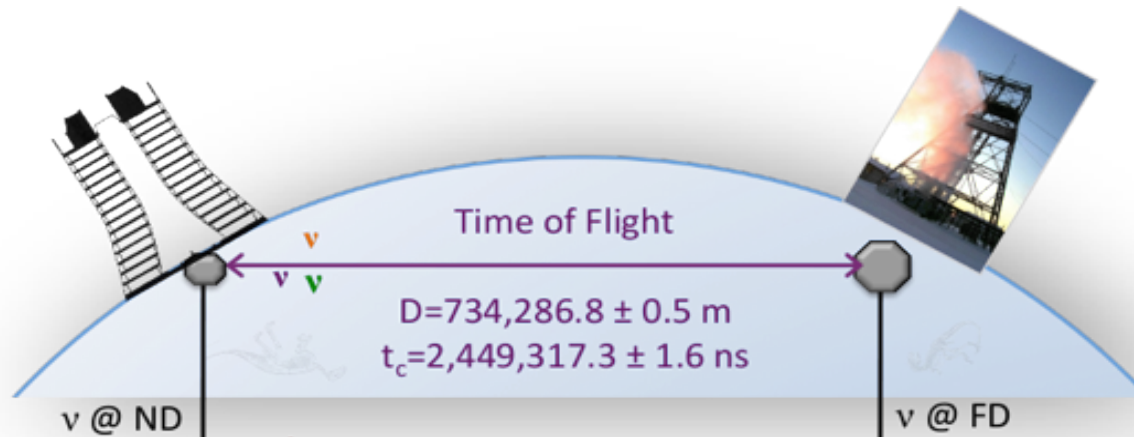
Delta and the hierarchy



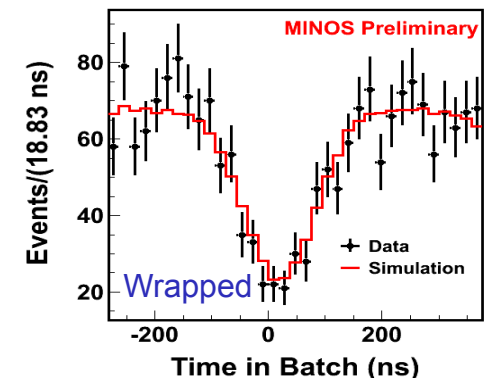
- For this plot we look at C.L. wrt the best fit in each hierarchy
(i.e. assume the hierarchy and see what is consistent with the data)
- Sensitivity to hierarchy: In pseudo experiments, we get the hierarchy right 60% of the time (compare to 50% by chance)
- With θ_{13} from reactor experiments, our data slightly favor inverted hierarchy $-2\Delta\ln L=0.2$



Measure the neutrino velocity by time of flight between the 2 detectors



- Initial result after first year of data indicated neutrinos arrived at FD earlier than expected
 $-126 \pm 32 \text{ (stat.)} \pm 64 \text{ (syst.) ns}^\dagger$
- We revisited this analysis with a factor of 8 more events & refined systematic errors
 $-15 \pm 11 \text{ (stat.)} \pm 29 \text{ (syst.) ns}$
- New timing system will greatly reduce the errors in the last 11 weeks of collected data (box to be opened) & during MINOS+ running





New TOF results: quoted in ns difference WRT speed of light over ~730km

Previously reported results

- MINOS¹ -126 ± 32 (stat.) ± 64 (syst.)
- OPERA² -57.8 ± 7.8 [+8.3, -5.9]

New results reported at Neutrino 2012

- MINOS -15 ± 31
 - OPERA $+1.6 \pm 1.1$ [+6.1, -3.7]
 - Borexino $+2.7 \pm 1.2 \pm 3$
 - ICARUS $+5.1 \pm 1.1 \pm 5.5$
 - LVD $+2.9 \pm 0.6 \pm 3$
 - All consistent with the speed of light
- } CNGS

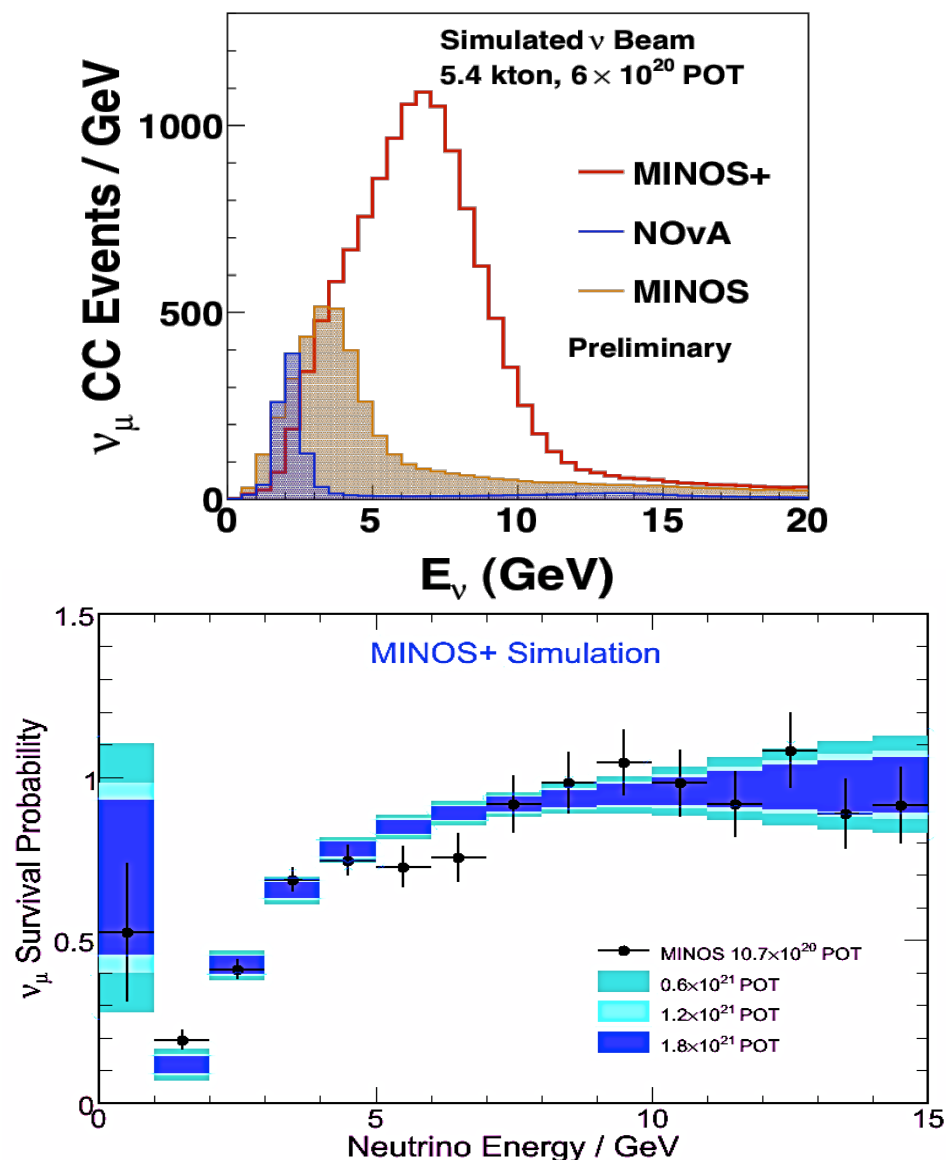
¹ Phys. Rev. D76 (2007) 072005

² arxiv1109.4897v2



MINOS+

medium-energy beam exposure

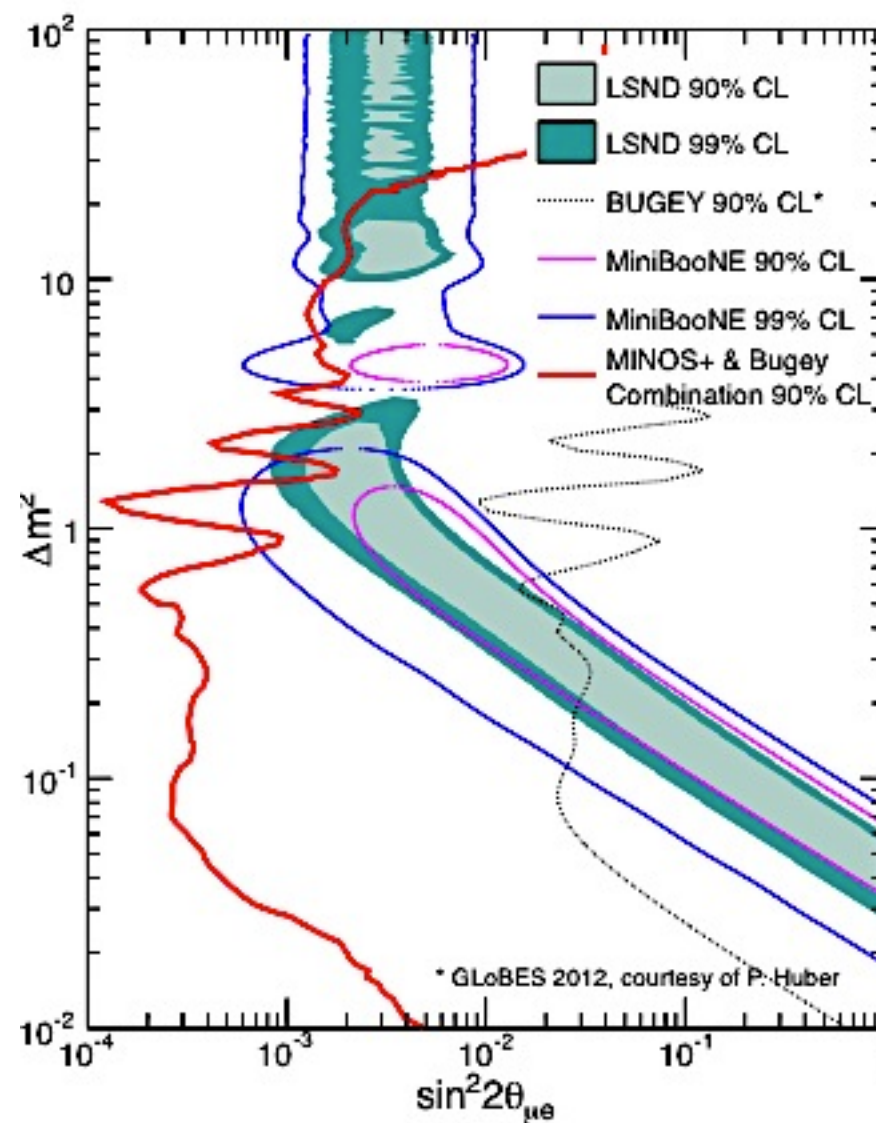
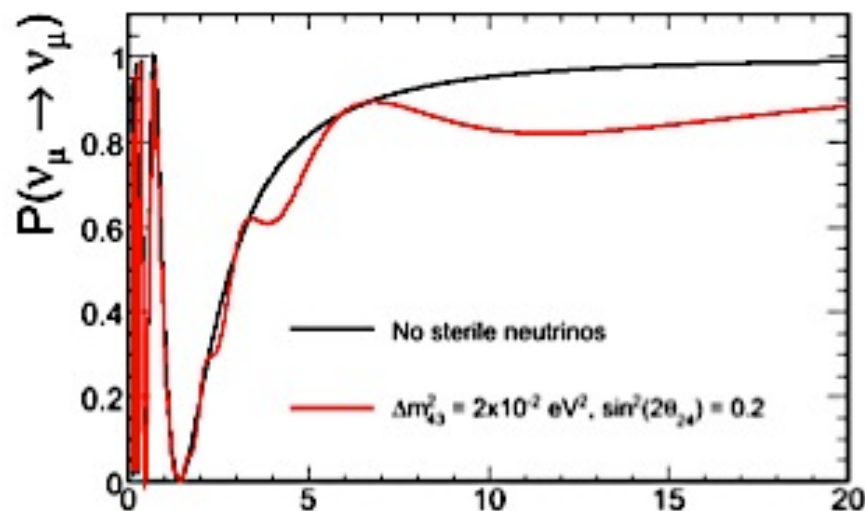


- MINOS will continue to run in the NOvA era
 - > On-axis ME beam peaks above the oscillation dip
- We get a flood of events
 - > Roughly 4000 far detector ν_μ CC events/year
- Unique test of oscillation paradigm
 - > Sensitivity to exotic signals



MINOS+ sensitivity

Using complementary information from Bugey and MINOS+ we can almost rule out all of the low mass LSND region





Summary of new MINOS results

- MINOS has completed its low-energy running after 1.56×10^{21} POT

muon neutrino disappearance

$$|\Delta m^2| = 2.39^{+0.09}_{-0.10} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) = 0.957^{+0.035}_{-0.036}$$

$$\sin^2(2\theta) > 0.90 \text{ (90\% C.L.)}$$

muon antineutrino disappearance

$$|\Delta \bar{m}^2| = 2.48^{+0.22}_{-0.27} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) > 0.83 \text{ (90\% C.L.)}$$

electron neutrino appearance

$$0.01 \text{ (0.03)} < \sin^2(2\theta_{13}) < 0.12 \text{ (0.19)} \text{ at 90\% C.L.}$$

$$\text{for } \delta_{CP} = 0, \sin^2(2\theta_{23}) = 1, \text{ normal (inverted) hierarchy}$$

$$\sin^2(2\theta_{13}) = 0 \text{ excluded at 96\%}$$

- Neutrino velocity is consistent with the speed of light
- MINOS+ has an exciting program based on NuMI medium-energy running