

Results from EXO-200

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Excellence Cluster ‘Universe’

for the EXO-200 collaboration

ALSO SEE:

**EXO-200 and nEXO Posters
(# 93 - 100) in Session II**

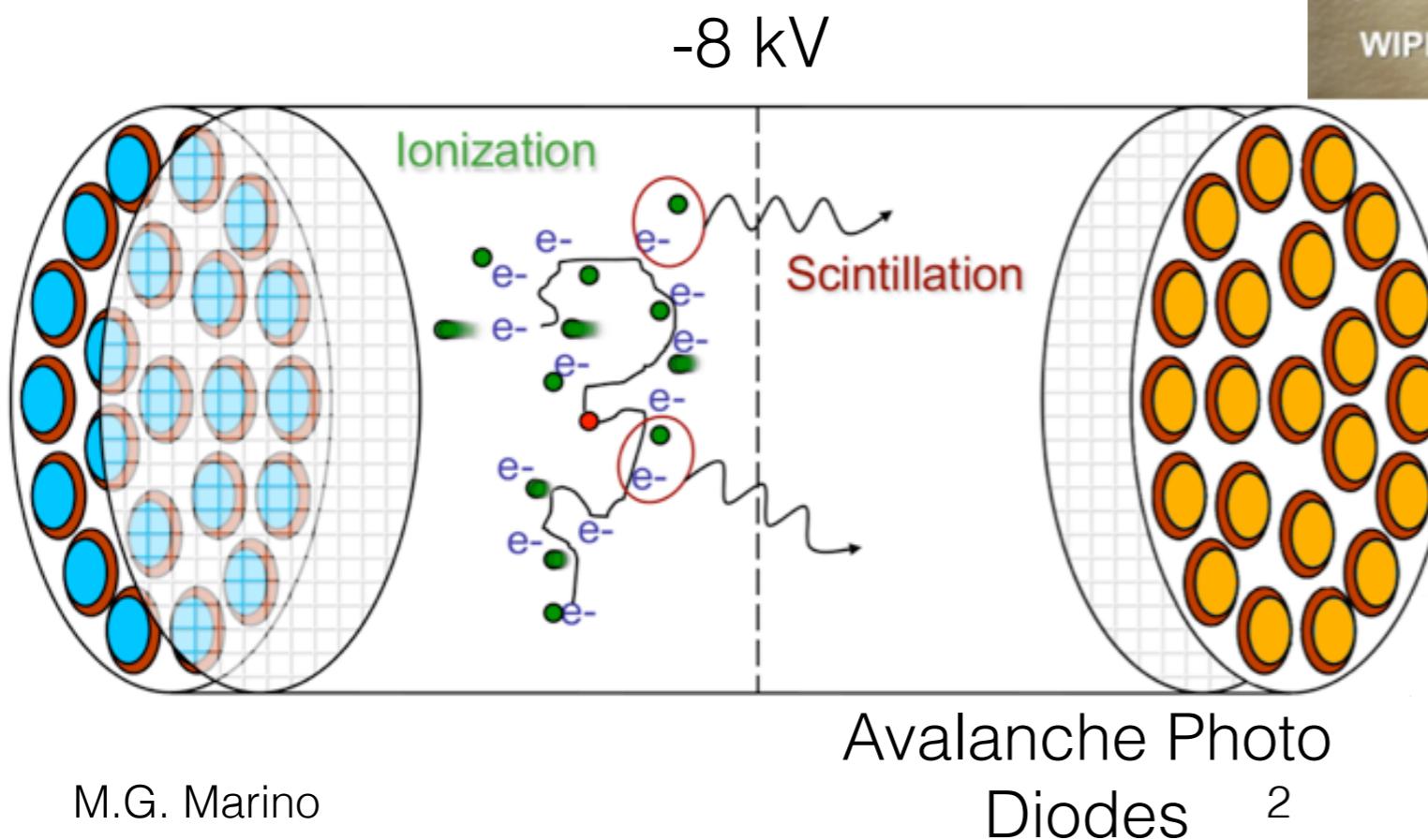
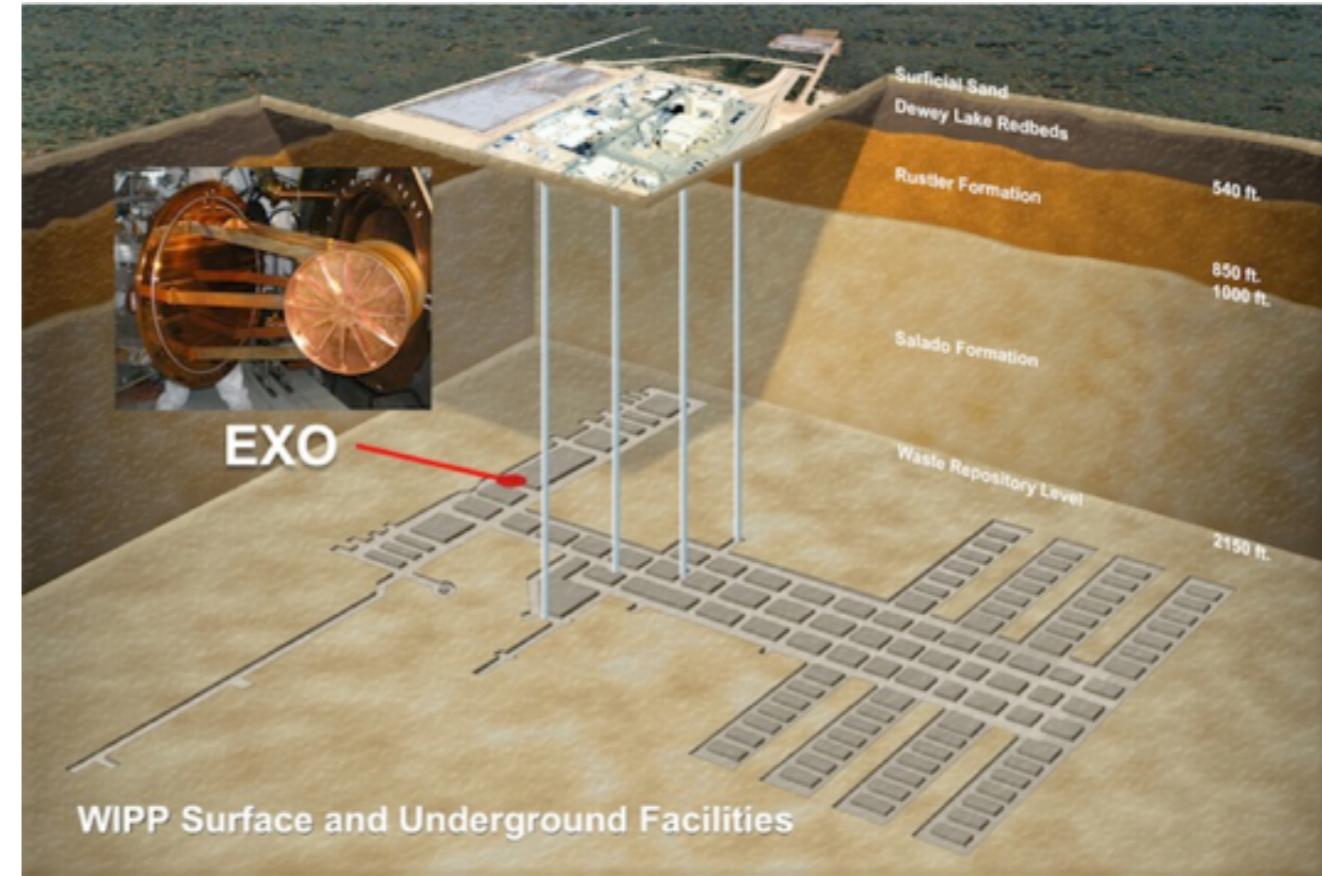
Neutrino 2014, Boston, USA
6 June 2014

EXO

Enriched
Xenon
Observatory

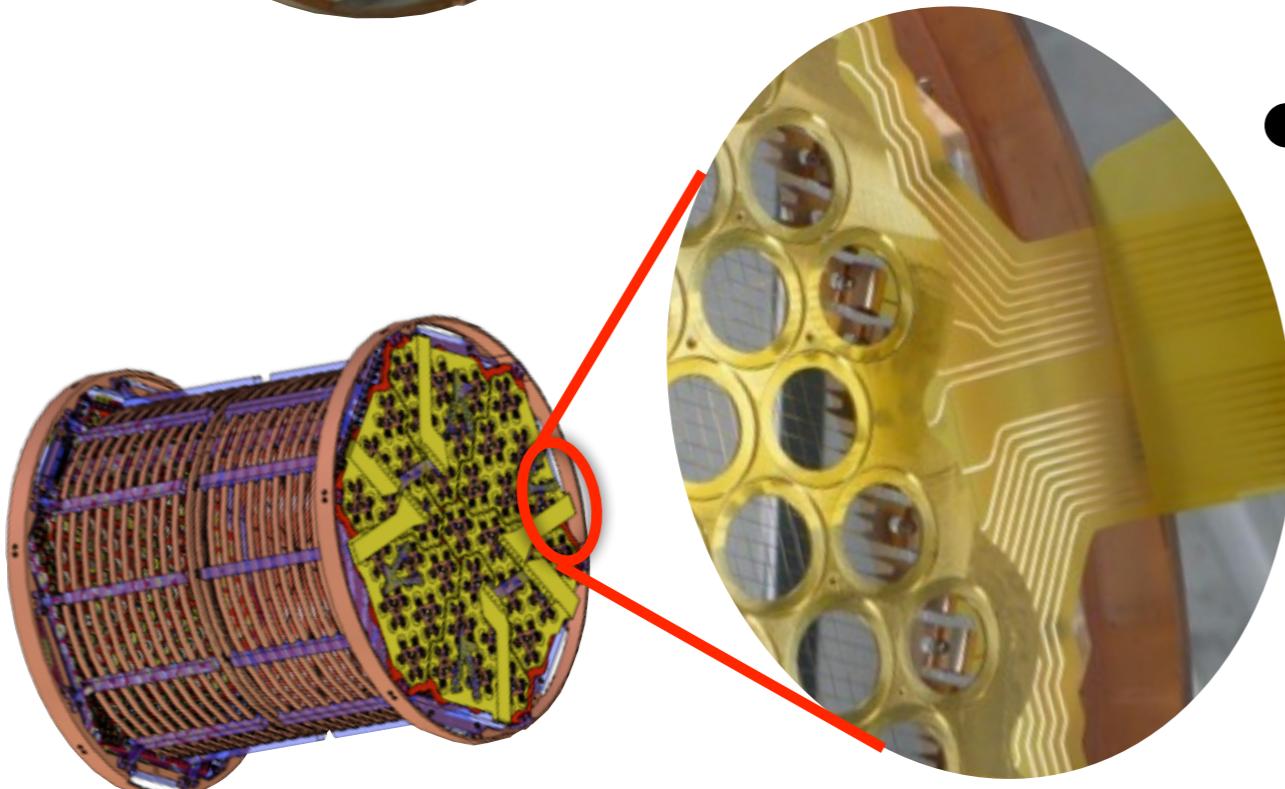
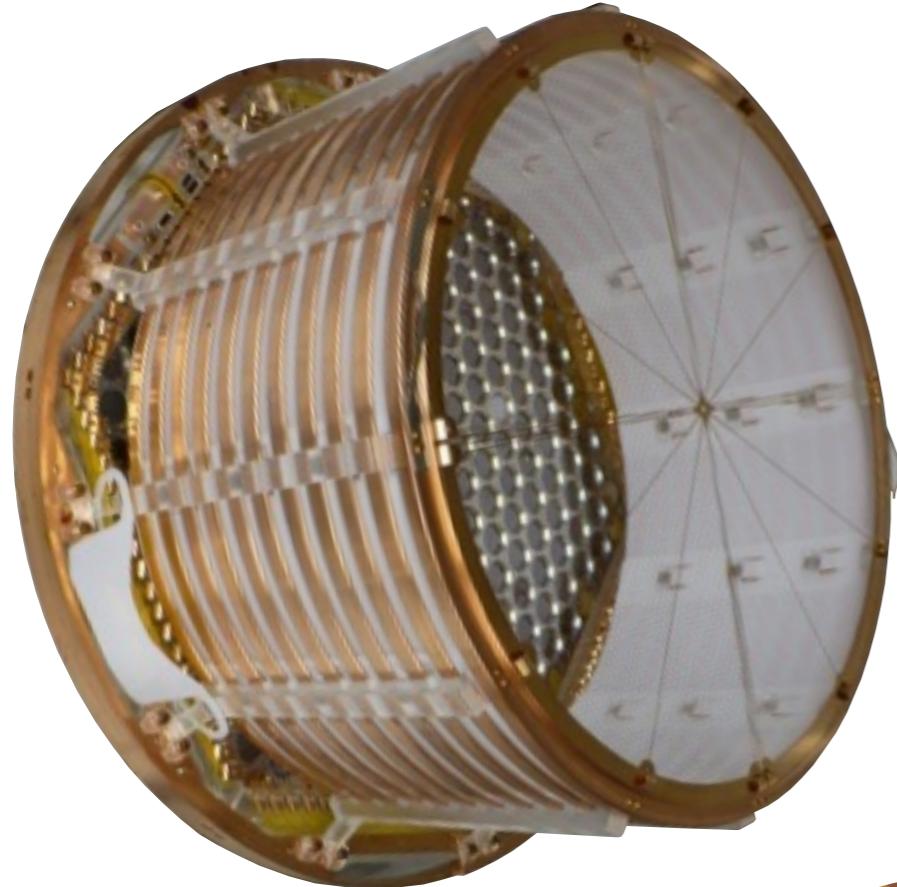


- Liquid Xe Time Projection Chamber (TPC)
- Enriched ^{136}Xe to 80.6%
- Q-value 2458 keV



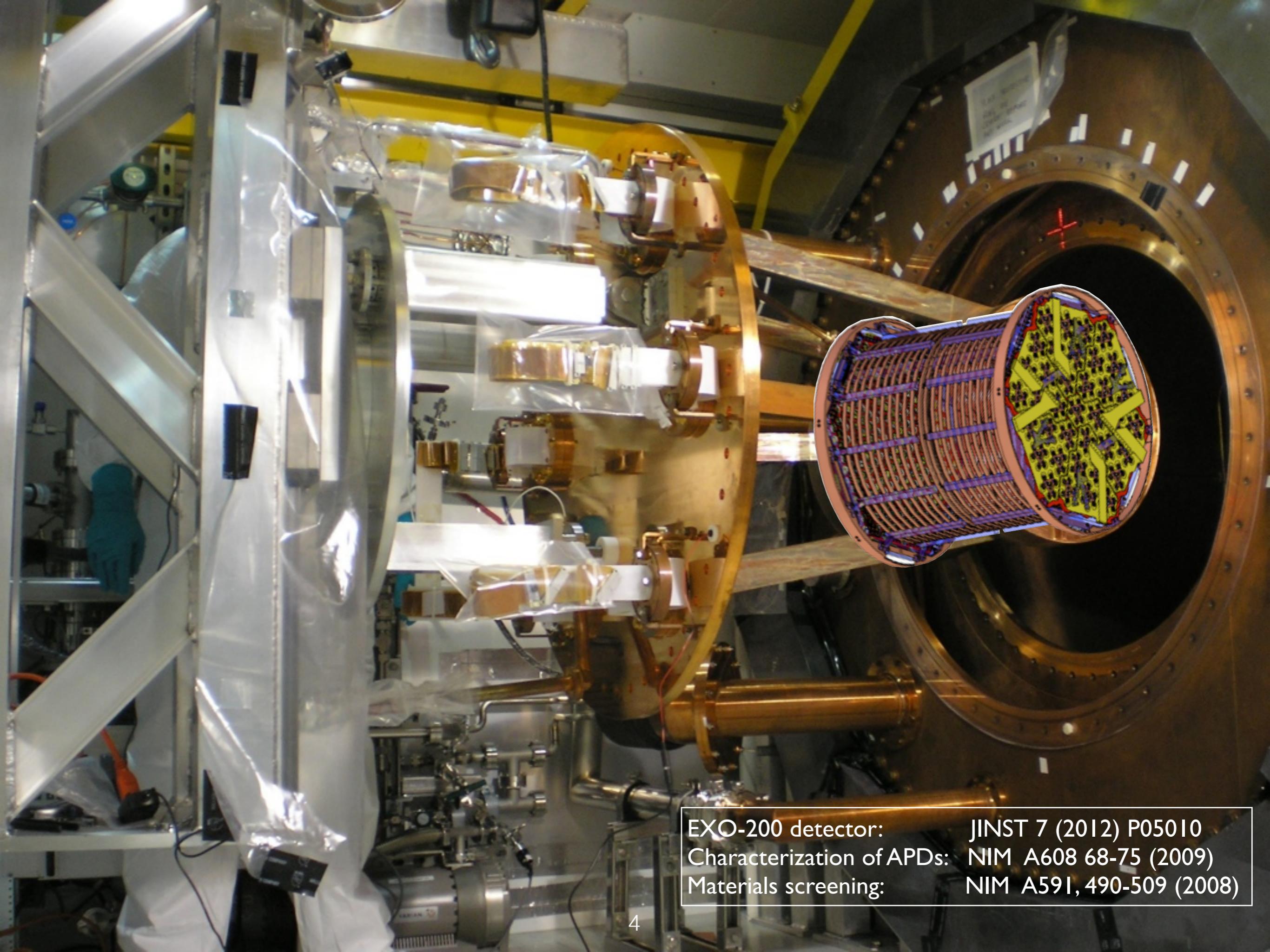
- Located at Waste Isolation Pilot Plant (WIPP) in Carlsbad, NM, USA
- 1585 meters water equivalent

The EXO-200 TPC

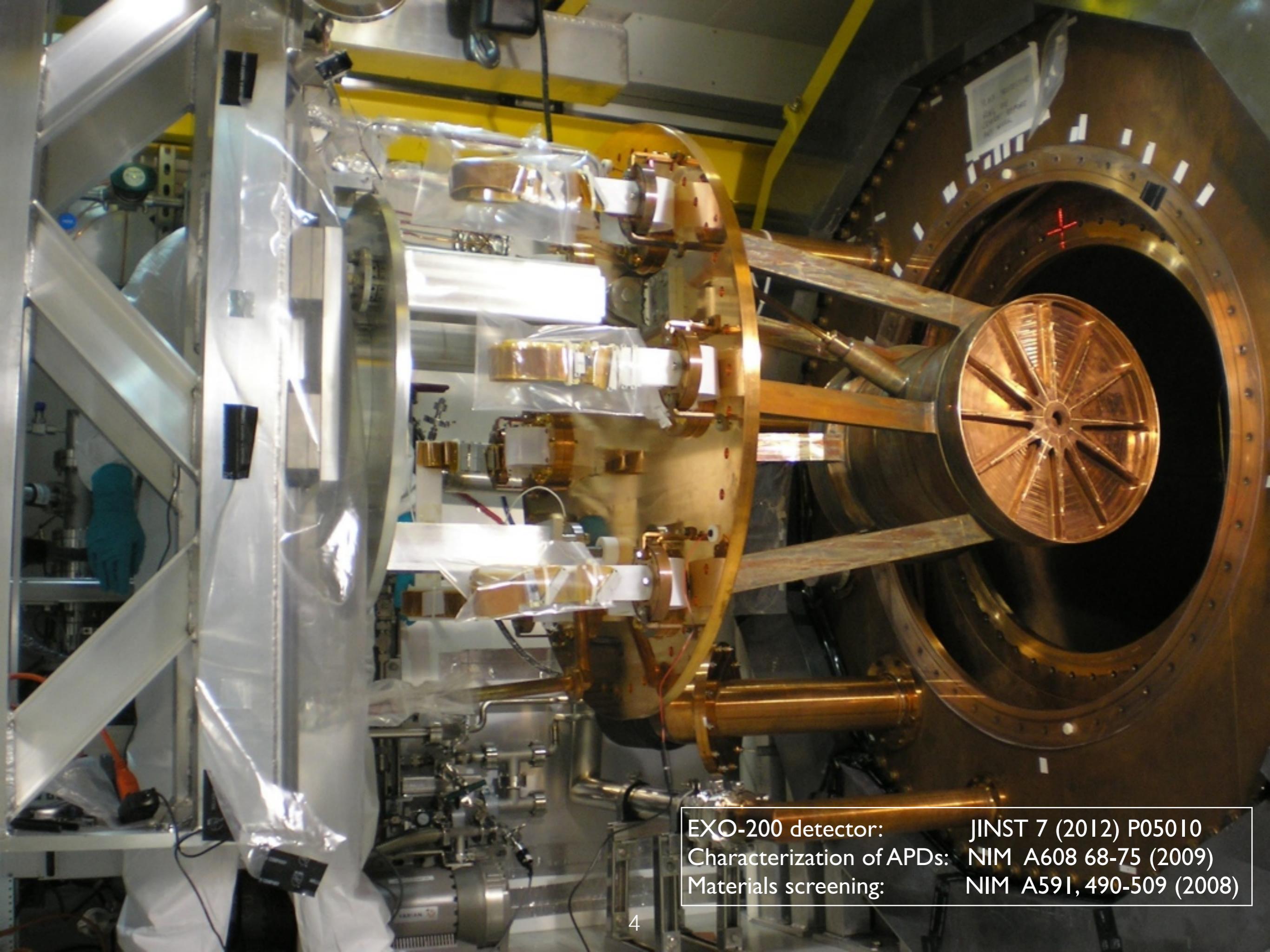


Two almost identical halves reading
ionization and 178 nm **scintillation**, each
with:

- 38 U triplet wire channels (charge)
- 38 V triplet wire channels, crossed at 60° (induction)
- 234 large area avalanche photodiodes (APDs, light in groups of 7)
- All signals digitized at 1 MS/s, ±1024 µs around trigger (2 ms total)
- Drift field 376 V/cm

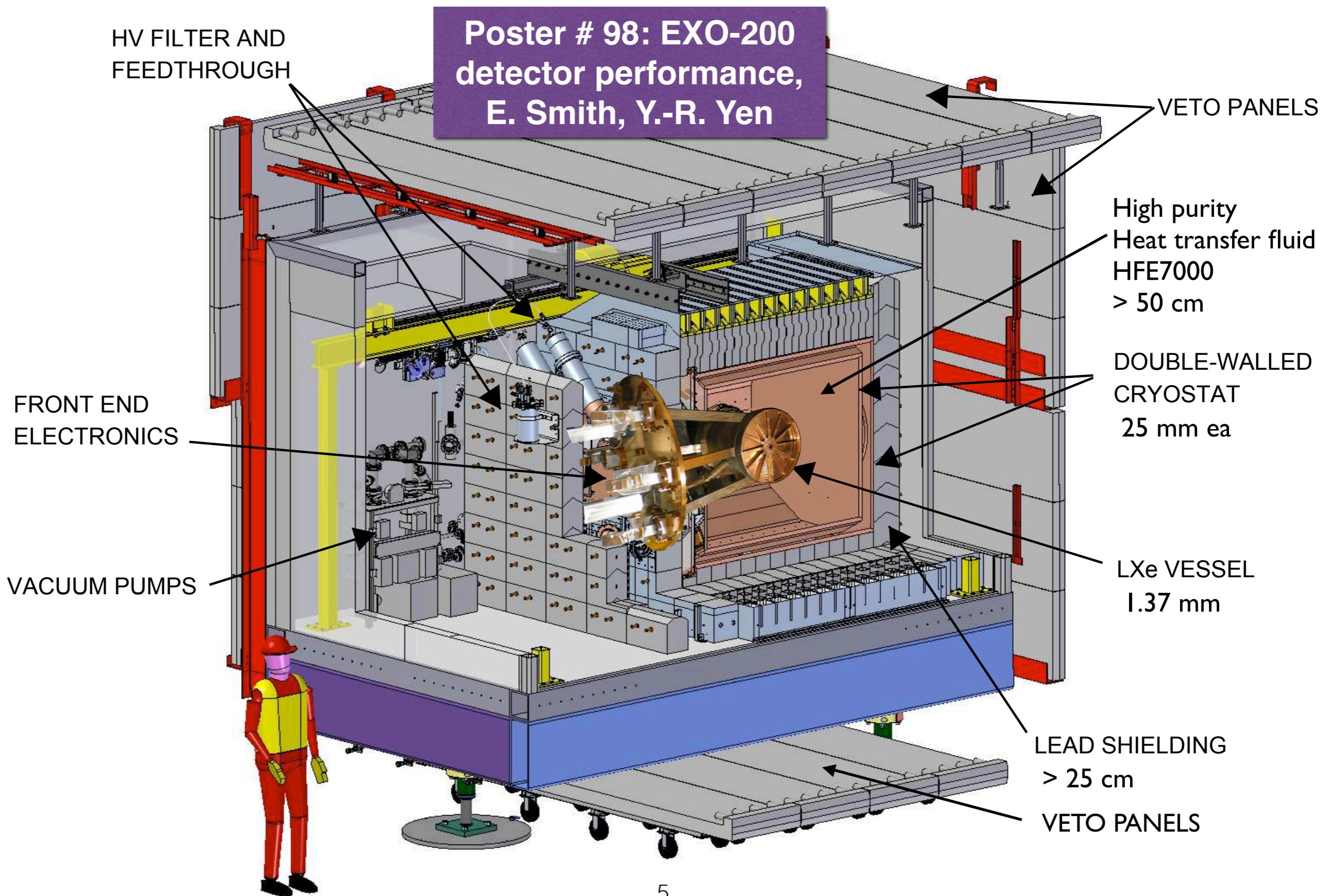


EXO-200 detector: JINST 7 (2012) P05010
Characterization of APDs: NIM A608 68-75 (2009)
Materials screening: NIM A591, 490-509 (2008)

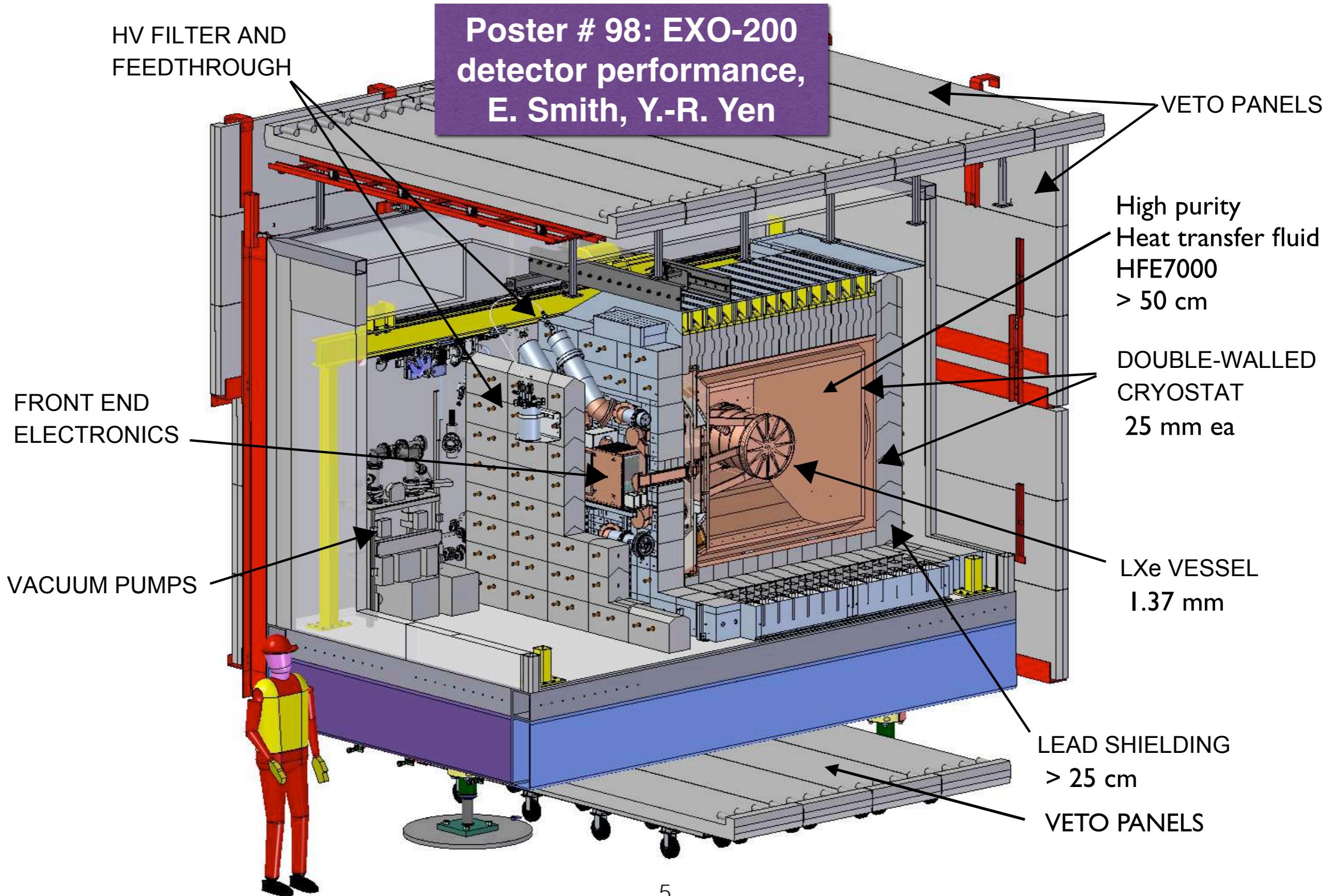


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The EXO-200 Detector

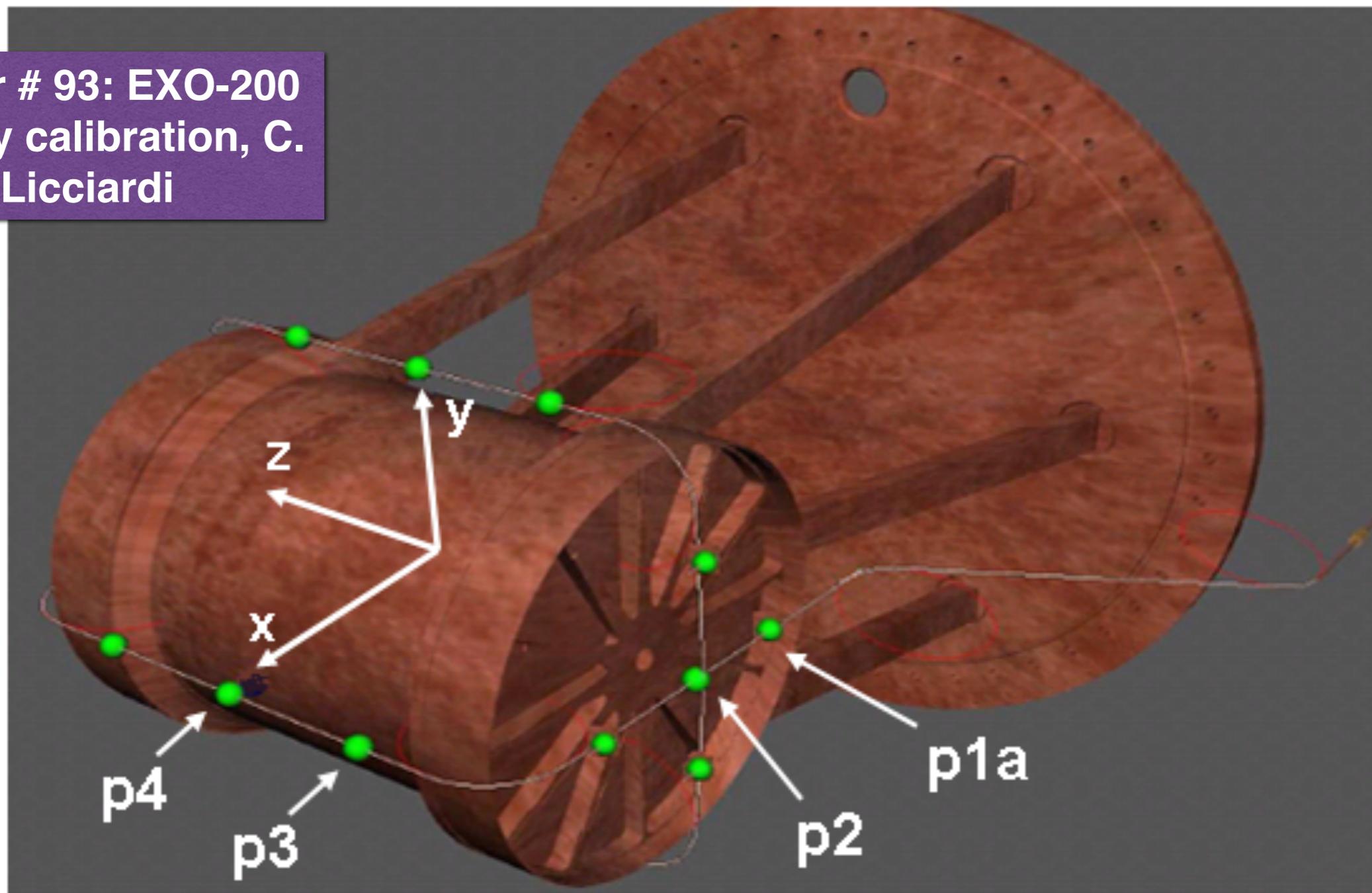


The EXO-200 Detector



Calibration

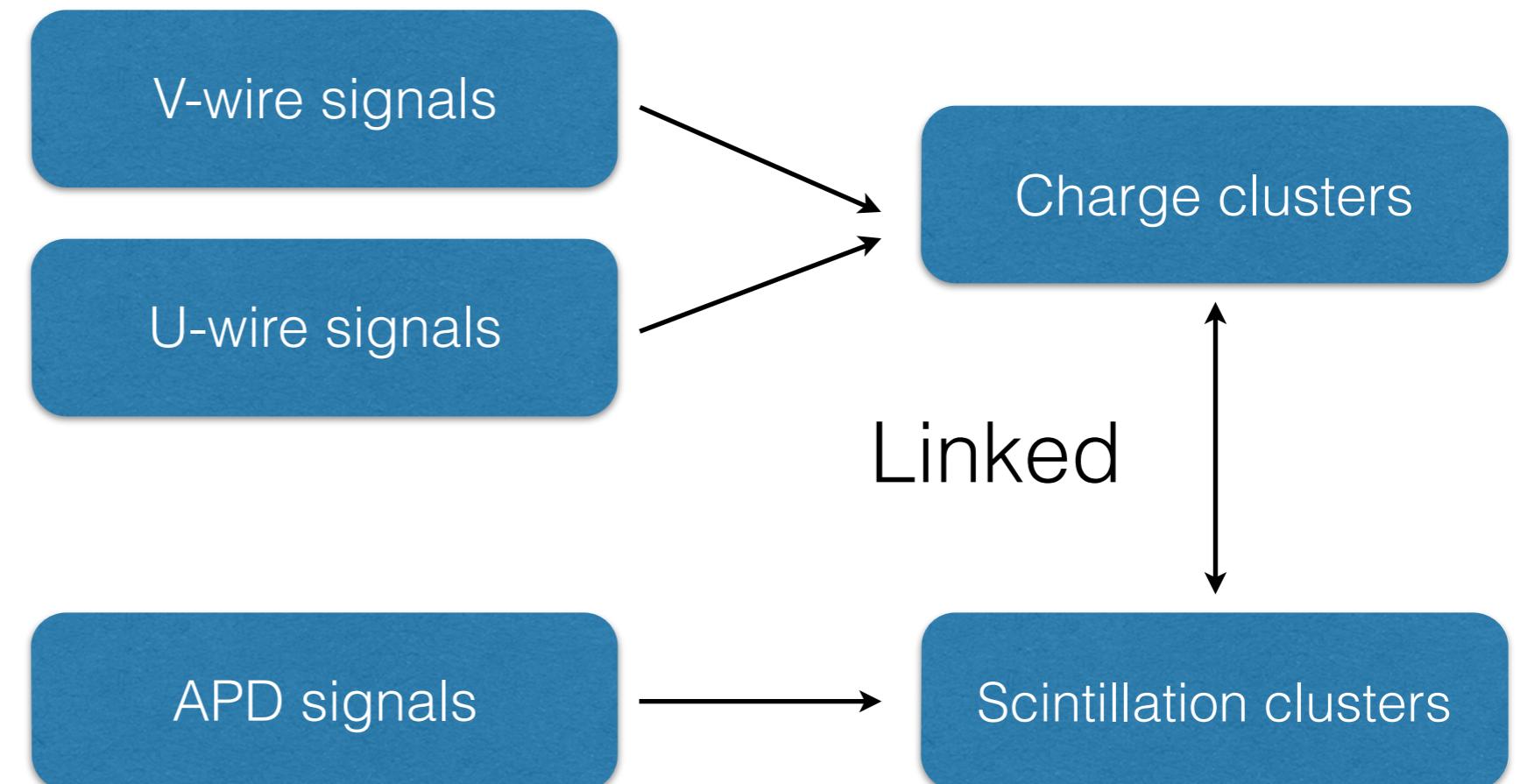
Poster # 93: EXO-200
Energy calibration, C.
Licciardi



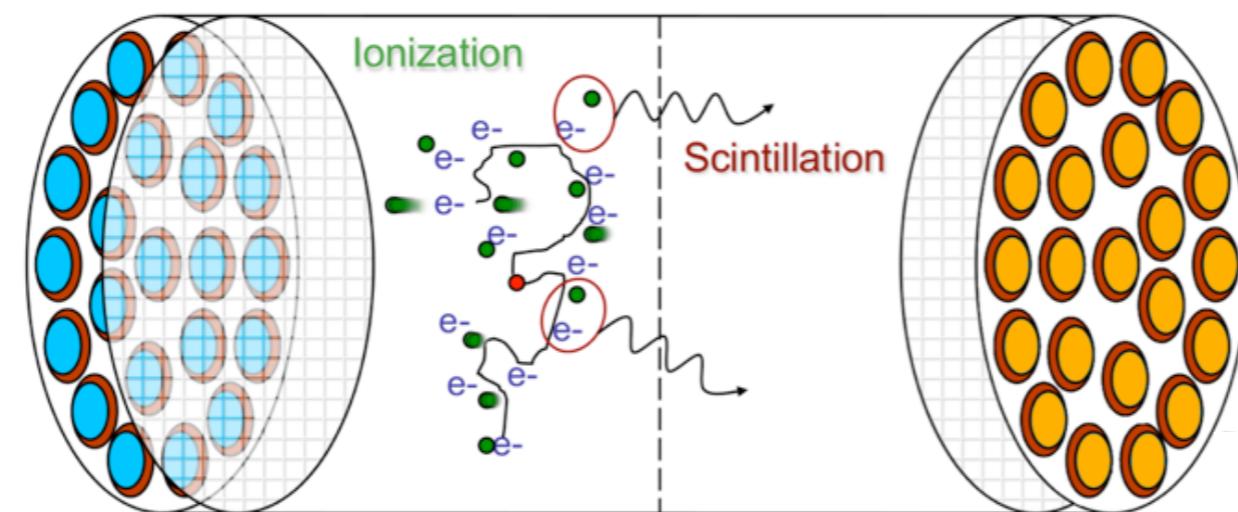
- Primary calibration - 2615 keV line from ^{228}Th source (deployed at least 2 times weekly)
- Periodic (every ~3 months) campaigns with ^{228}Th , ^{60}Co , ^{226}Ra , and ^{137}Cs , used also for systematics studies

Event reconstruction

1. Event position
2. Event multiplicity
3. Energy measurement

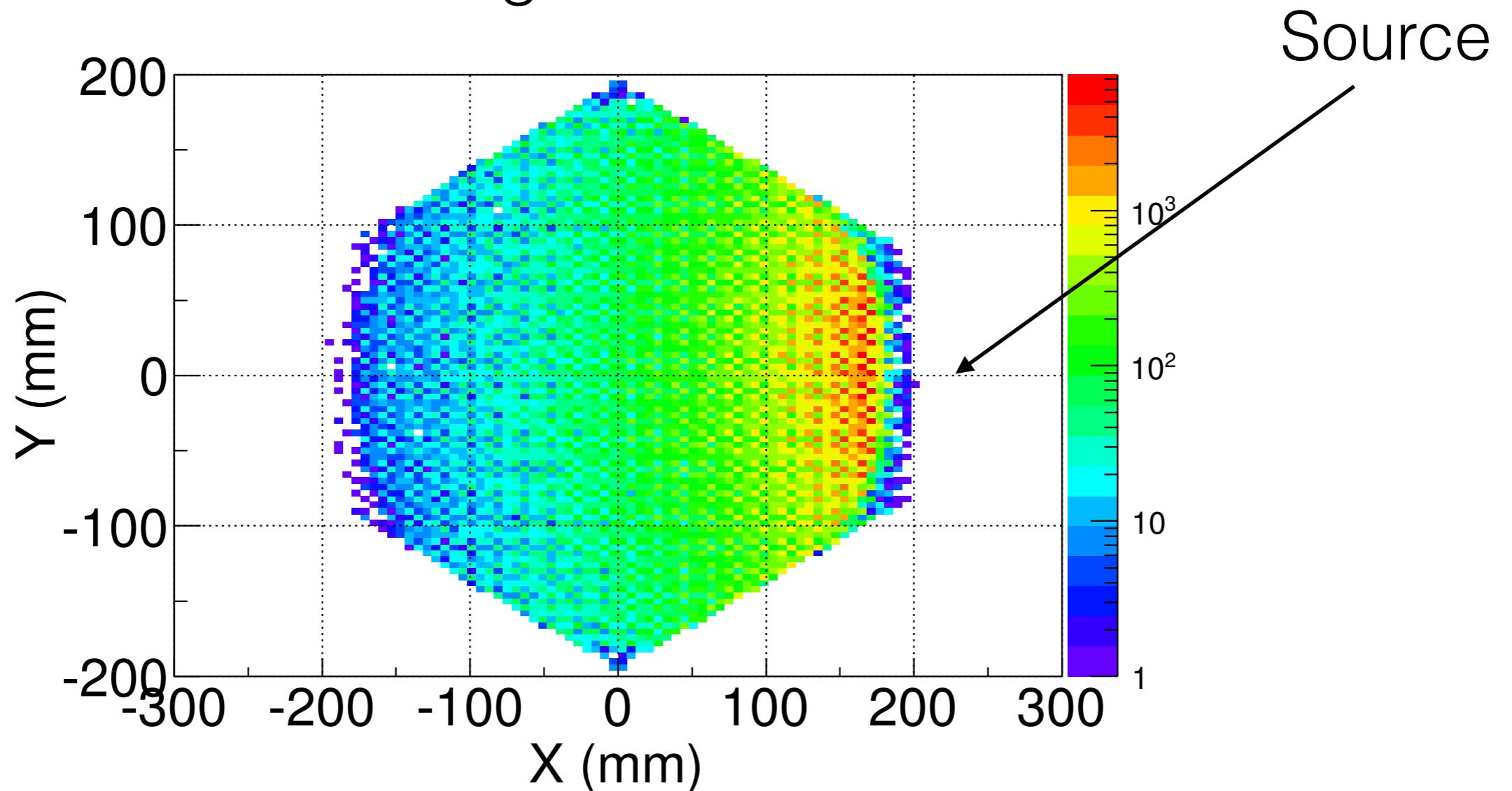


**Poster # 100: EXO-200
Event reconstruction,
R. MacLellan**



Position/multiplicity reconstruction

Combine signals into “clusters”

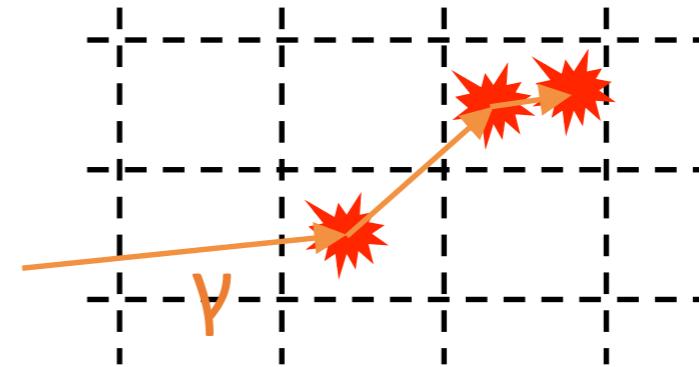
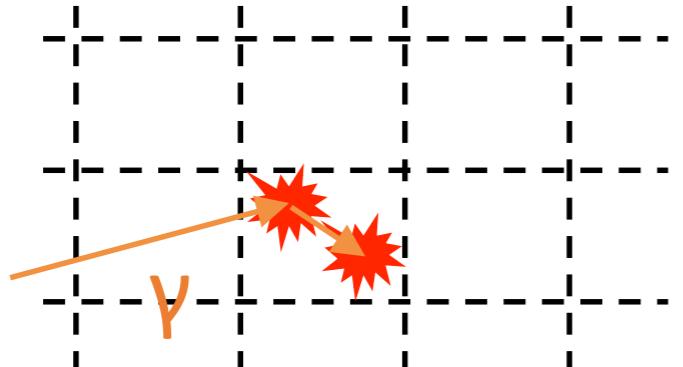


Uncertainty, 2.4 (1.2) mm U (V) + 1.5 mm shift (taken as systematic error), Z (0.5 mm), measured using internal decays on the cathode

Total error in fiducial volume due to position reconstruction: **1.73%**

Position/multiplicity reconstruction

Background measurement/reduction



Events with **> 1** charge cluster: multi-site (MS) events

Event with **1** charge cluster: single-site (SS) events

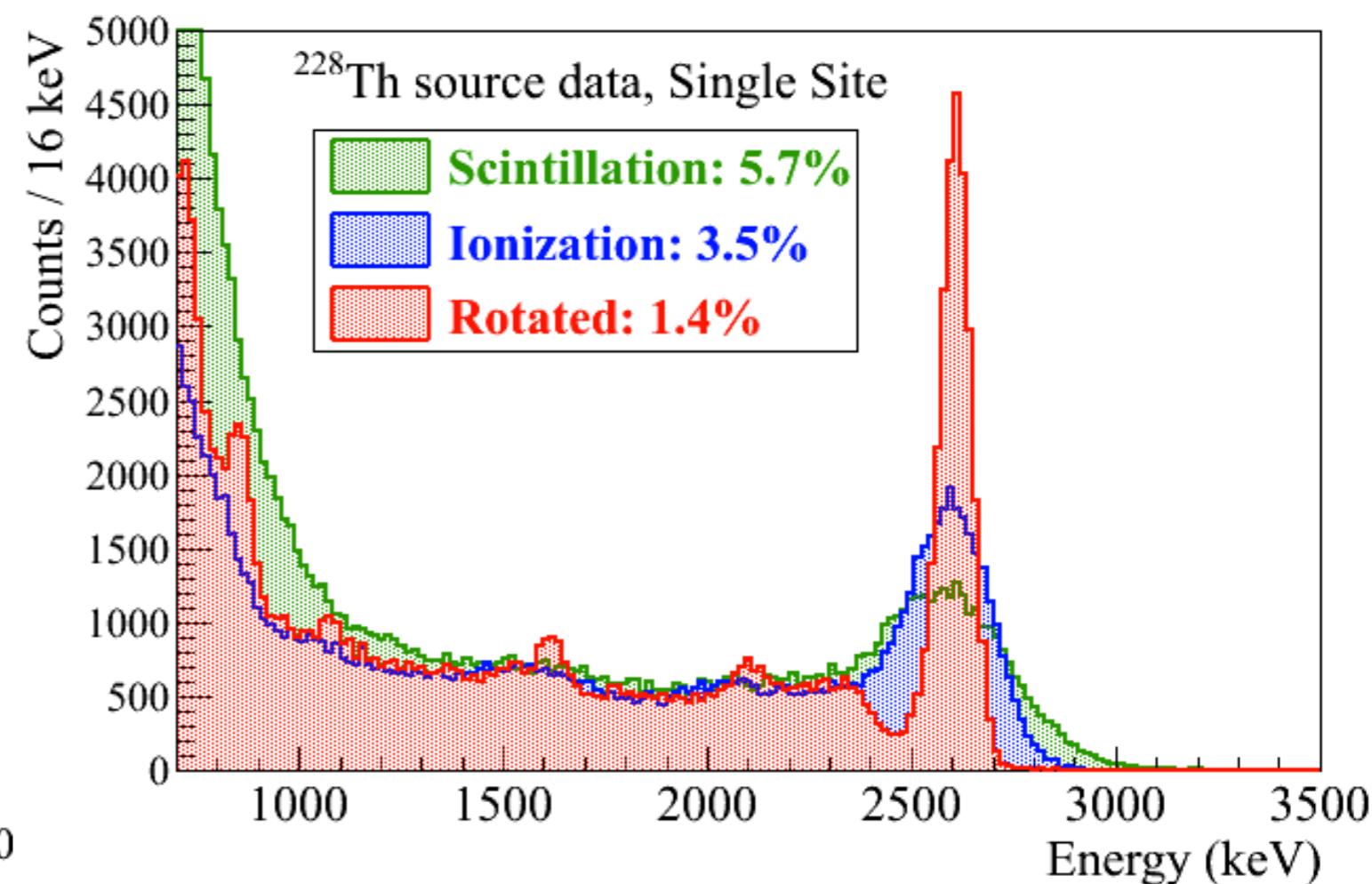
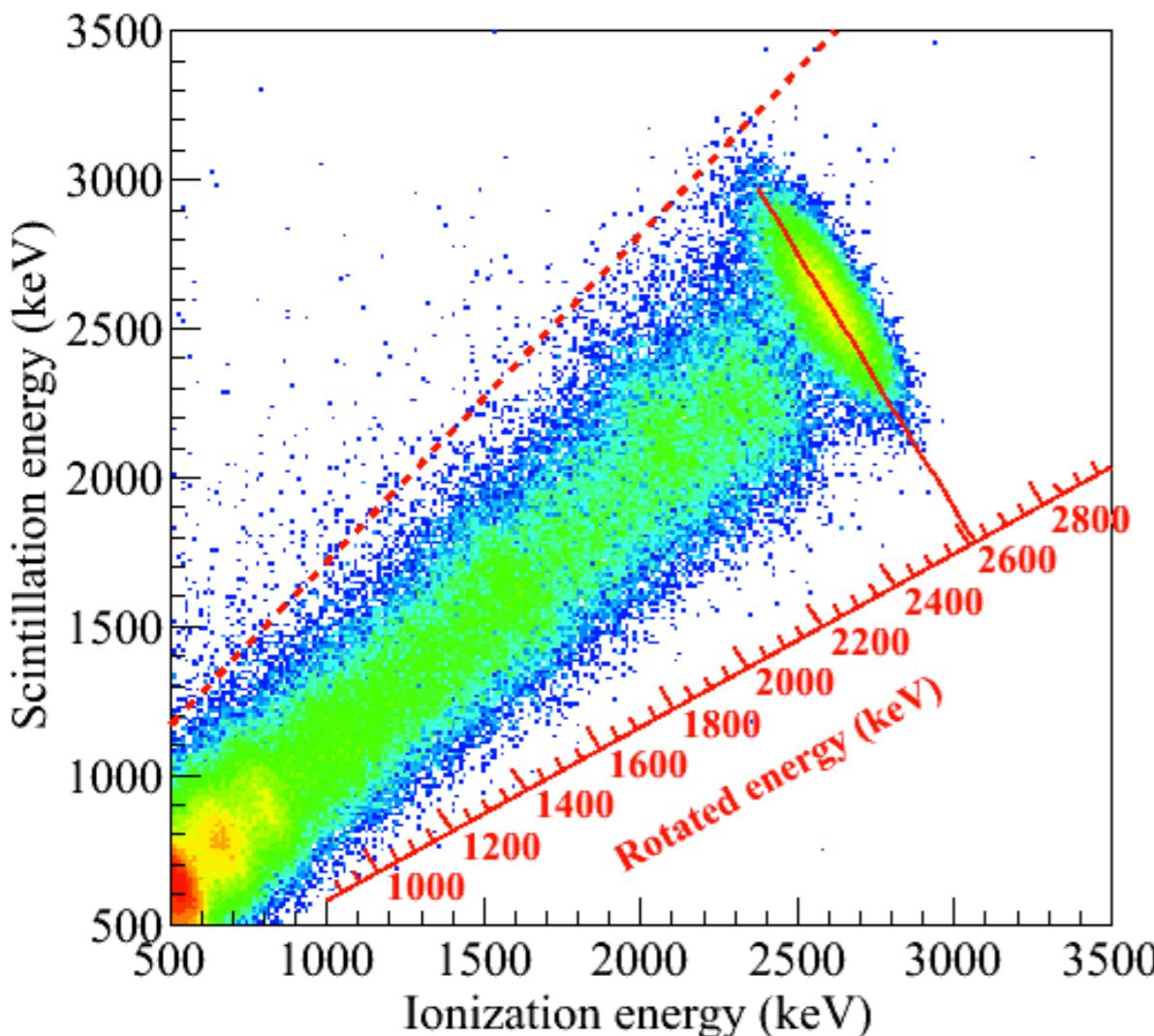
0νββ: ~90% SS

γs: ~30% SS at 0νββ Q-value

But we don't throw MS events away! Use them in the fit to help predict background...

Energy measurement

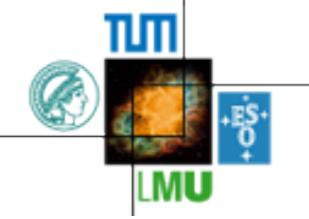
Combination of charge and light



'Rotation angle' determined weekly using ^{228}Th source data, defined as angle which gives best 'rotated' resolution

Energy resolution is dominated by APD noise (more on this later)

Previous results (Run 2a)



Most precise measurement of the $2\nu\beta\beta$ half-life

$$T_{1/2}^{2\nu\beta\beta} = 2.165 \pm 0.016(\text{stat}) \pm 0.059(\text{sys}) \times 10^{21} \text{ yr}$$

PRC 89, 015502 (2014)

$2\nu\beta\beta$ Sig-to-Bgd ratio:
11: 1

Inner 40% Fiducial Volume:
19: 1

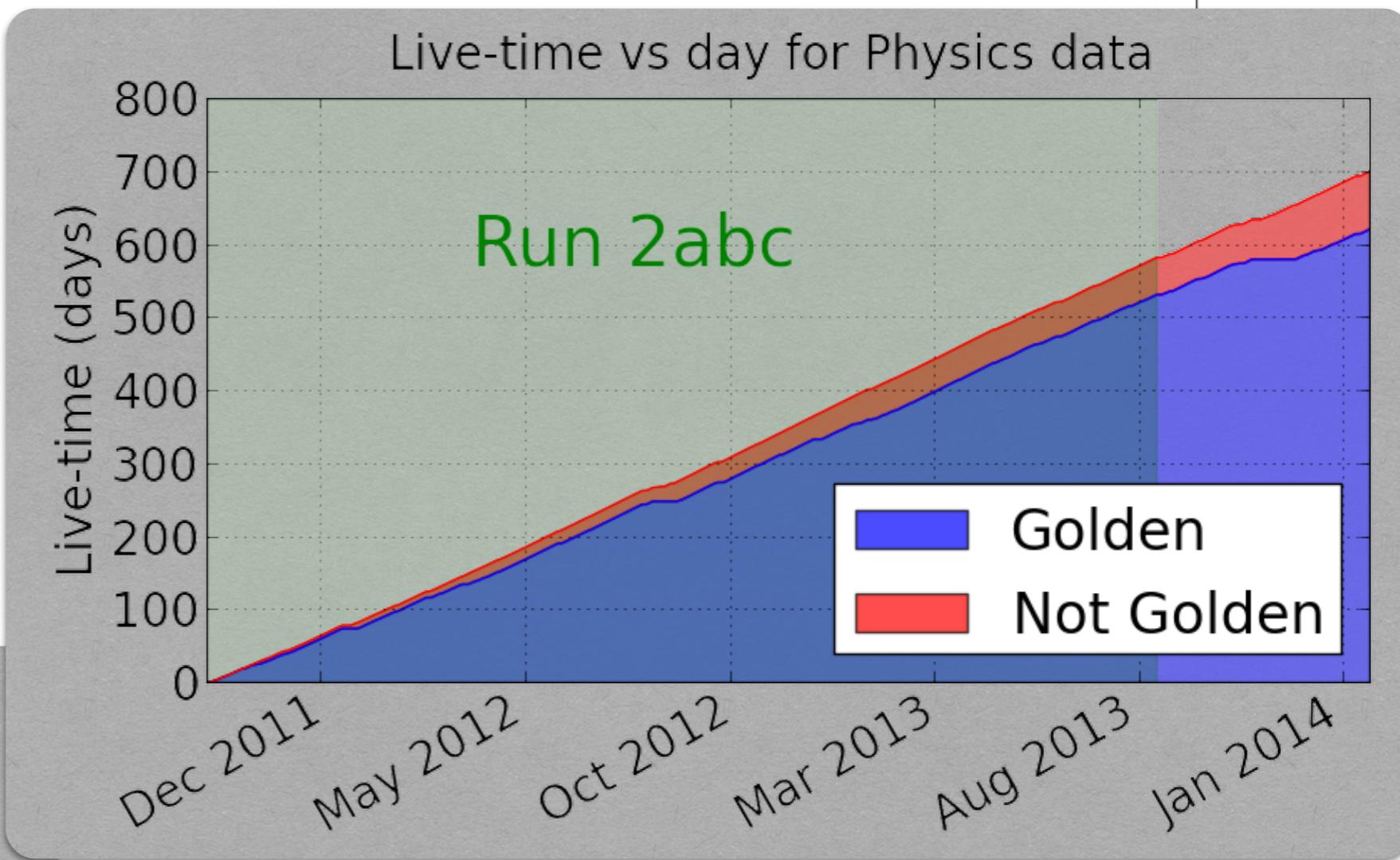
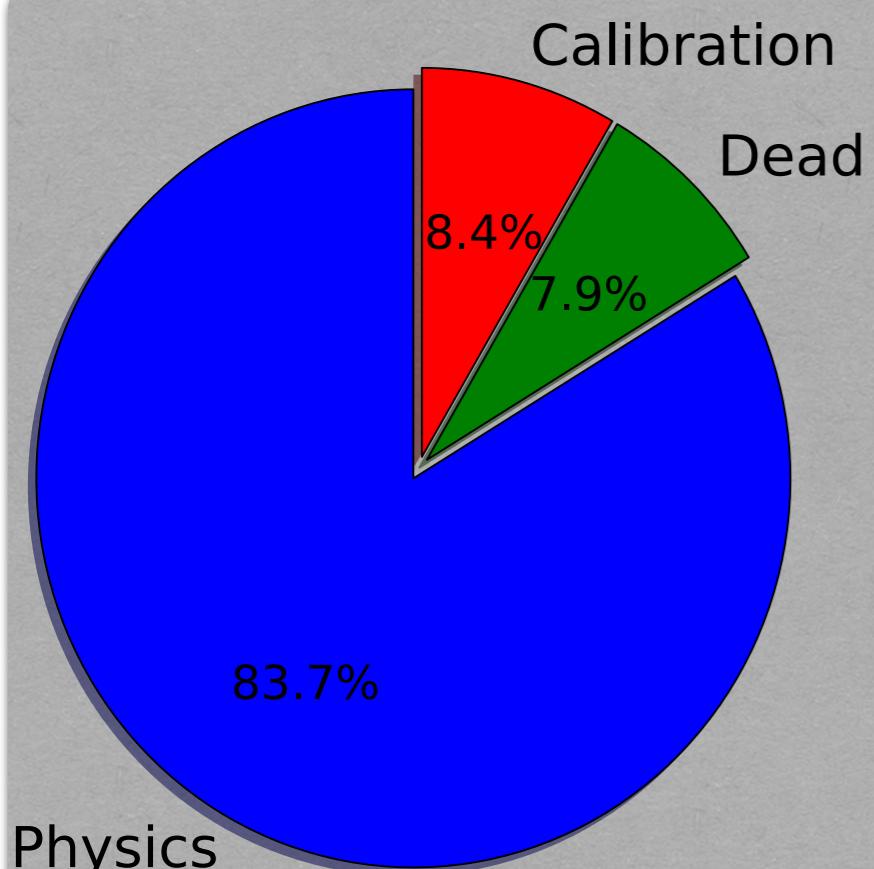
Above reference includes analysis details

Recent results (Run 2)



nature

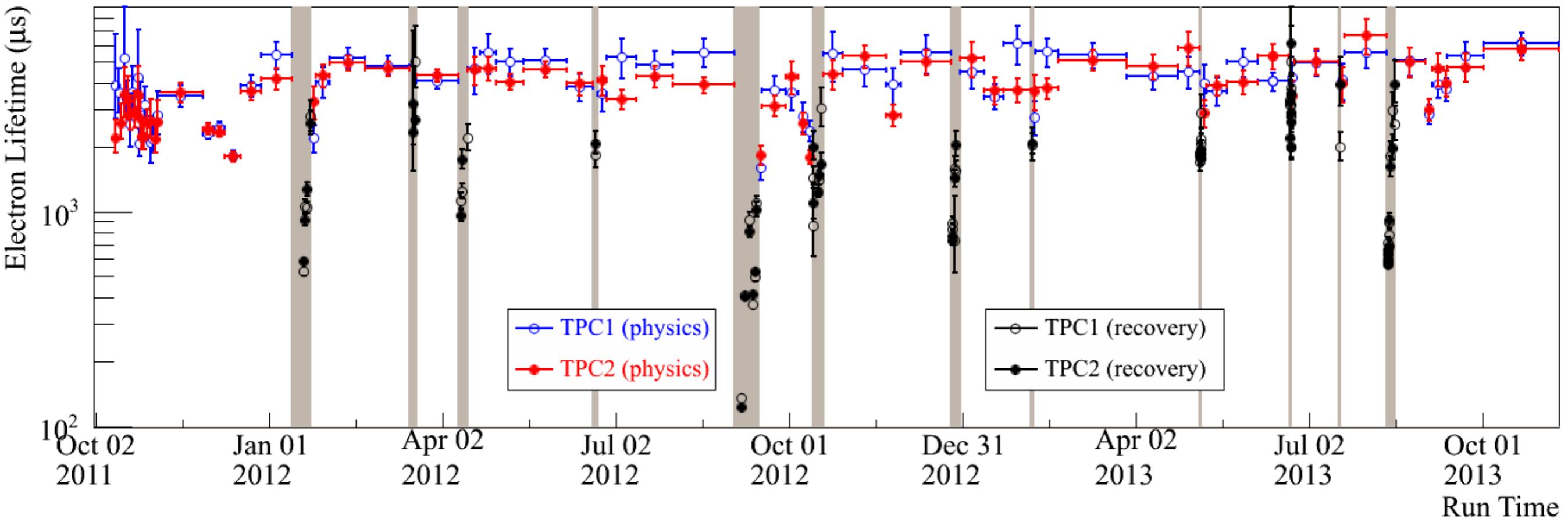
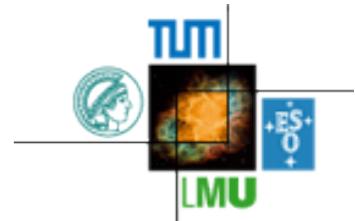
(12 June 2014, online 4 June)
doi:10.1038/nature13432



Accumulation of “Golden” data
 447.60 ± 0.01 days livetime
(100 kg·yr, 736 mol·yr ^{136}Xe
exposure)

(6 Oct 2011- 1 Sep 2013)

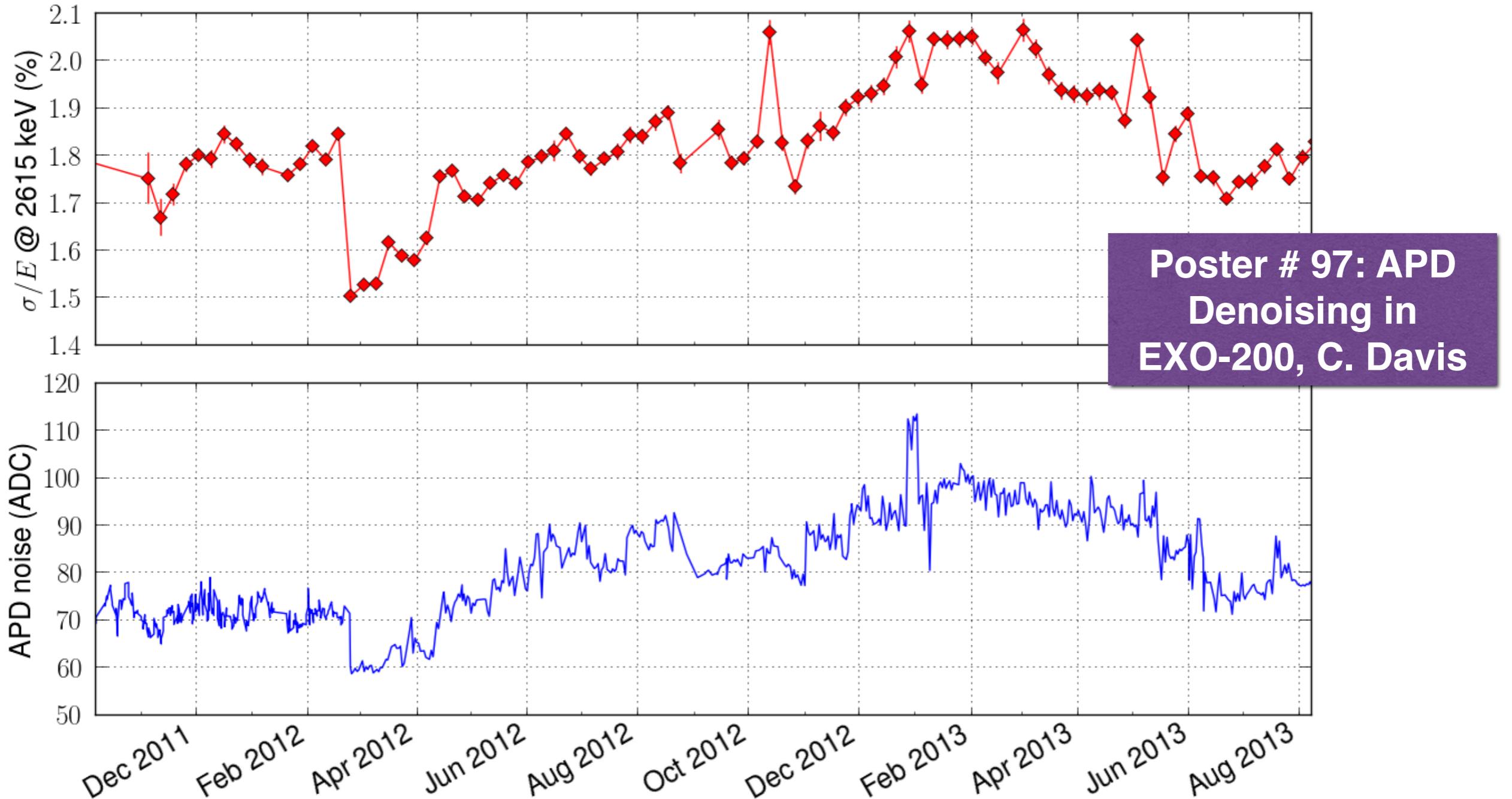
Xe Purity over Run 2



- Estimation based upon data from ^{228}Th source runs
- Purity strongly correlated with circulation pump speed
- At $\tau_e = 3 \text{ ms}$: drift time $< 110 \mu\text{s}$, loss of charge: 3.6% at full drift length

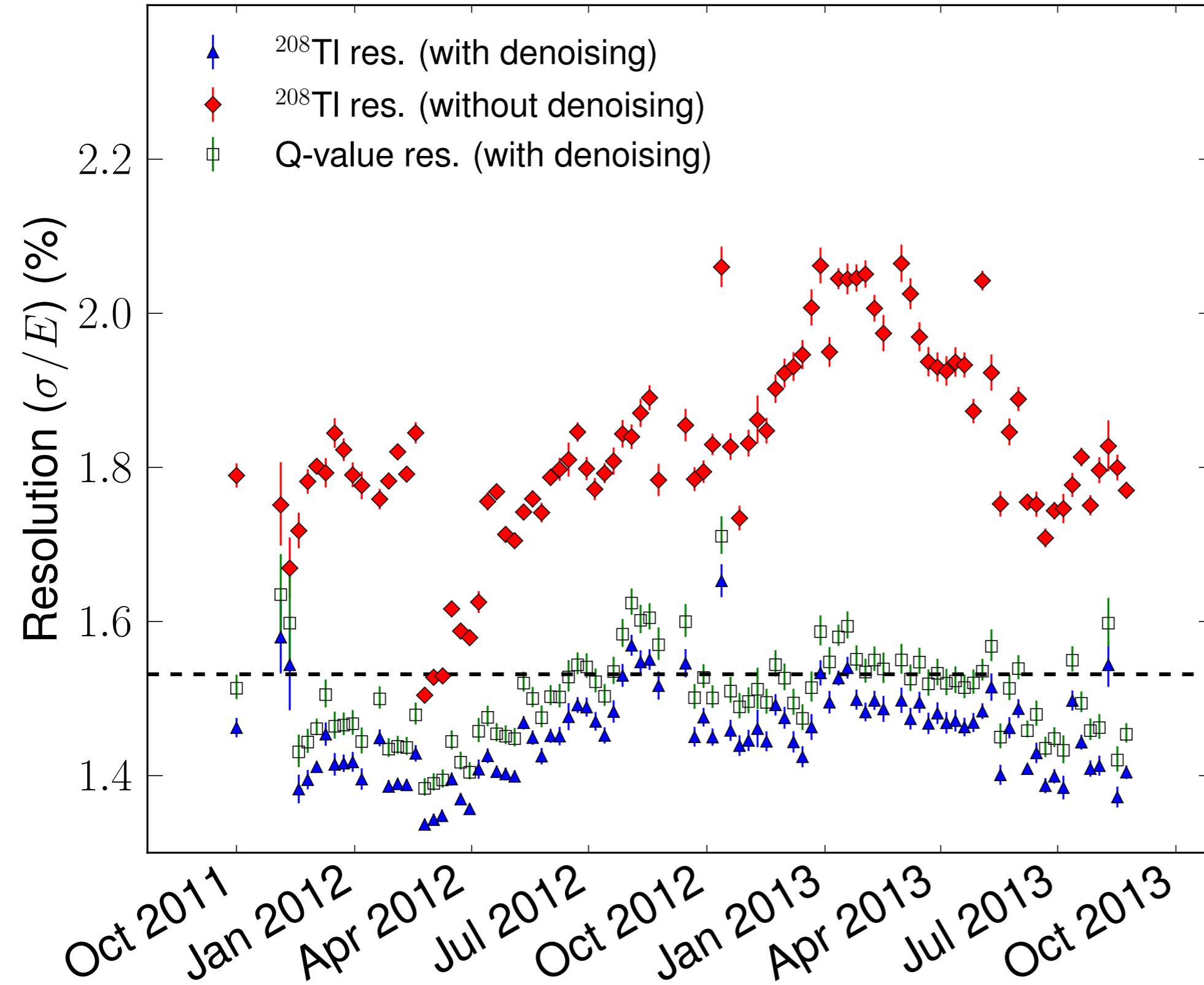
APD Denoising

Problem: Noise and resolution is dependent on the varying (correlated) noise of the APDs



(Current) solution: Find the optimum combination of APD signals *per event*, given position and noise

APD Denoising

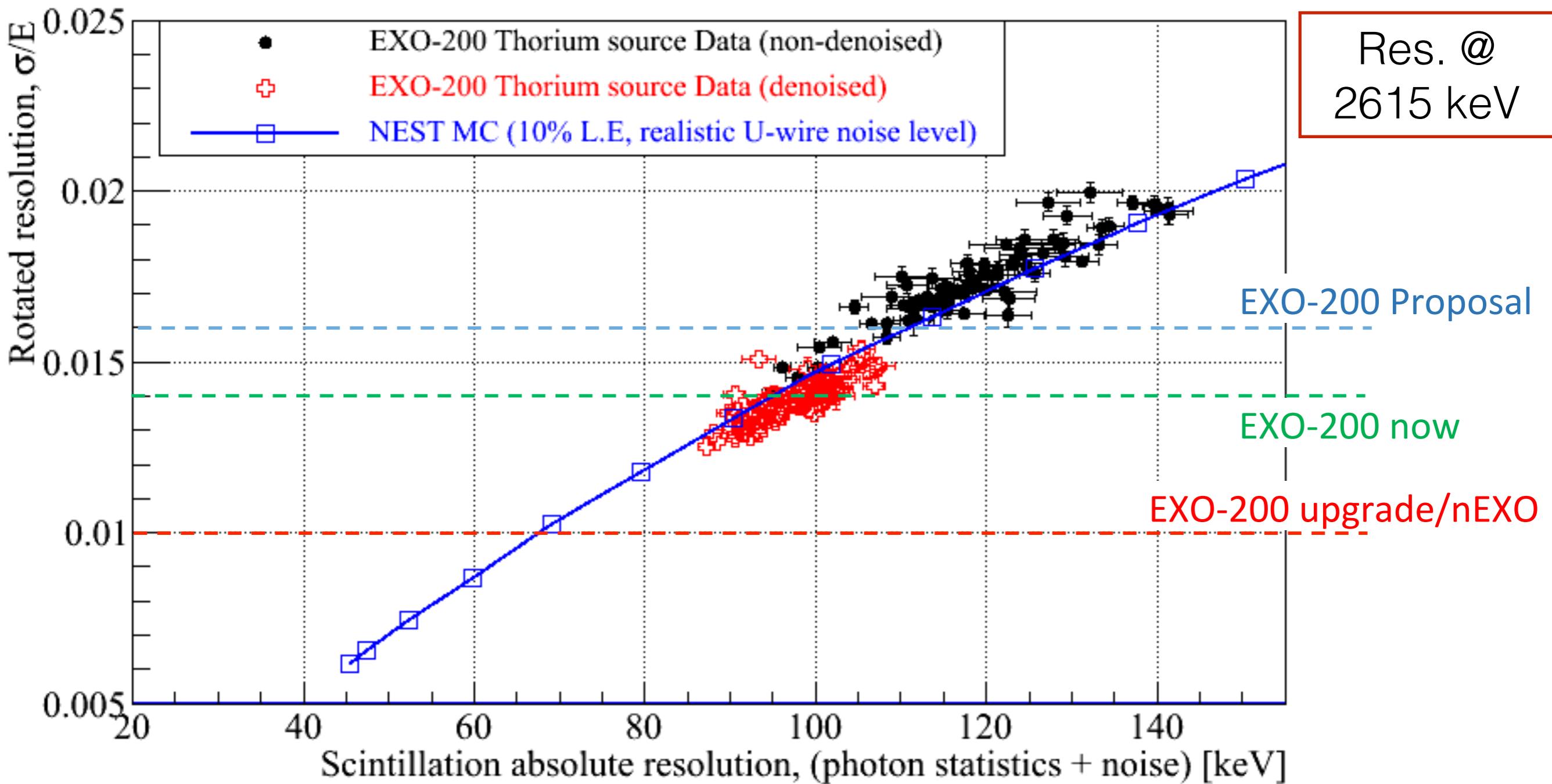


Poster # 97: APD
Denoising in
EXO-200, C. Davis

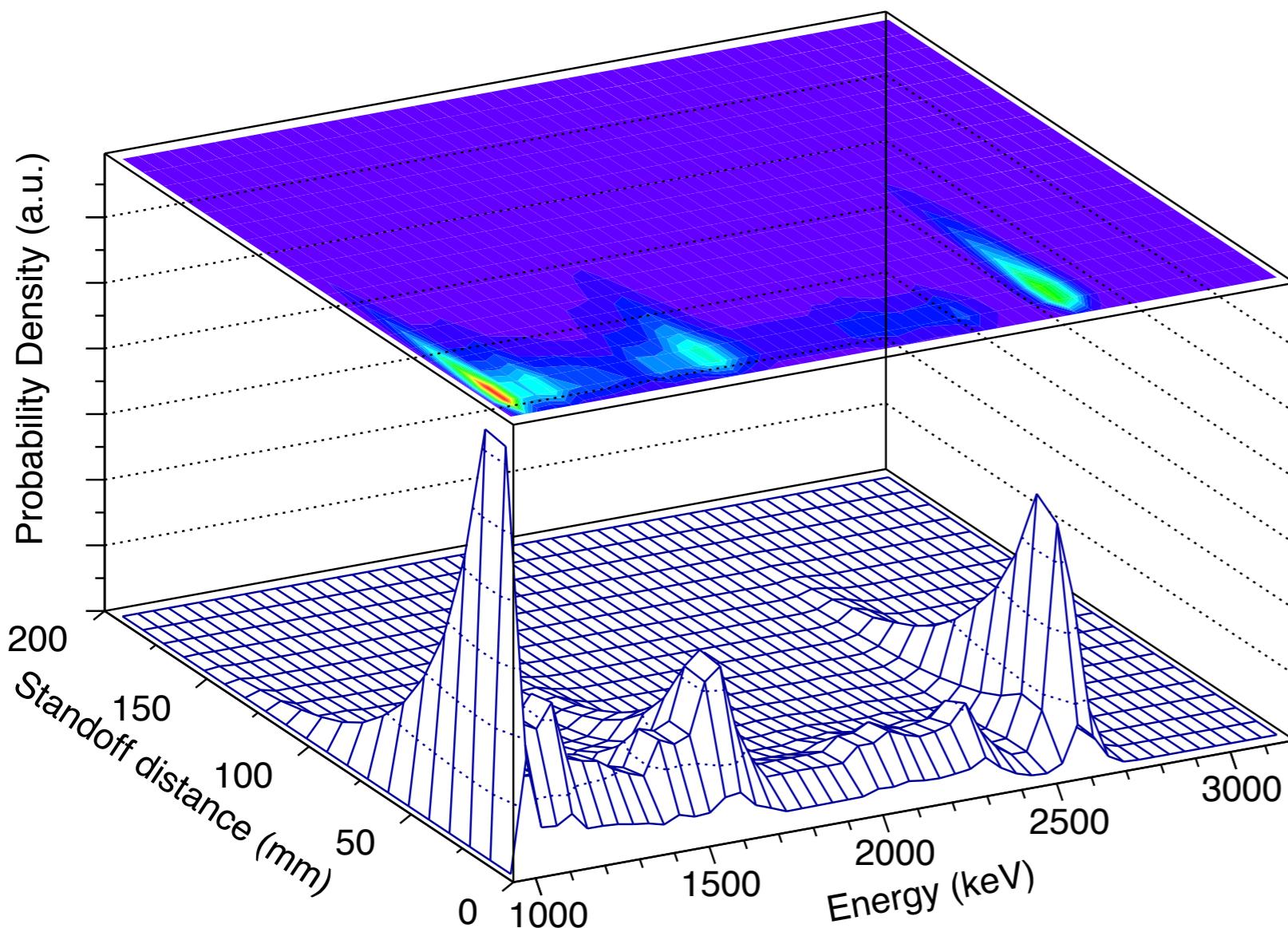
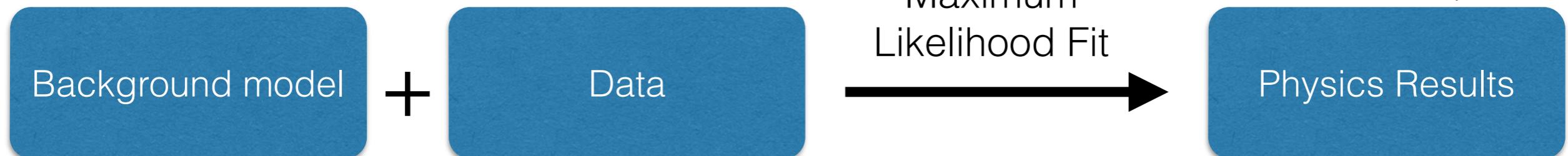
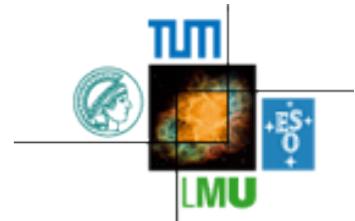
APD Denoising



Future: Upgrade of APD readout electronics to improve noise characteristics



Extracting physics results



- Variables
 - Energy
 - Position (standoff distance)
- Multiplicity
 - SS
 - MS

232Th PDF in Cu TPC Vessel (SS)

Systematic errors

- $0\nu\beta\beta$ detection efficiency:

Source:	Signal efficiency [%]:	Relative error [%]:
Summary from PRC 89 , 015502 (2014)	93.1	0.9
Partial reconstruction	90.9	7.8
Fiducial volume/rate agreement		3.4
Total:	84.6	8.6

- Region-of-interest (ROI) backgrounds:

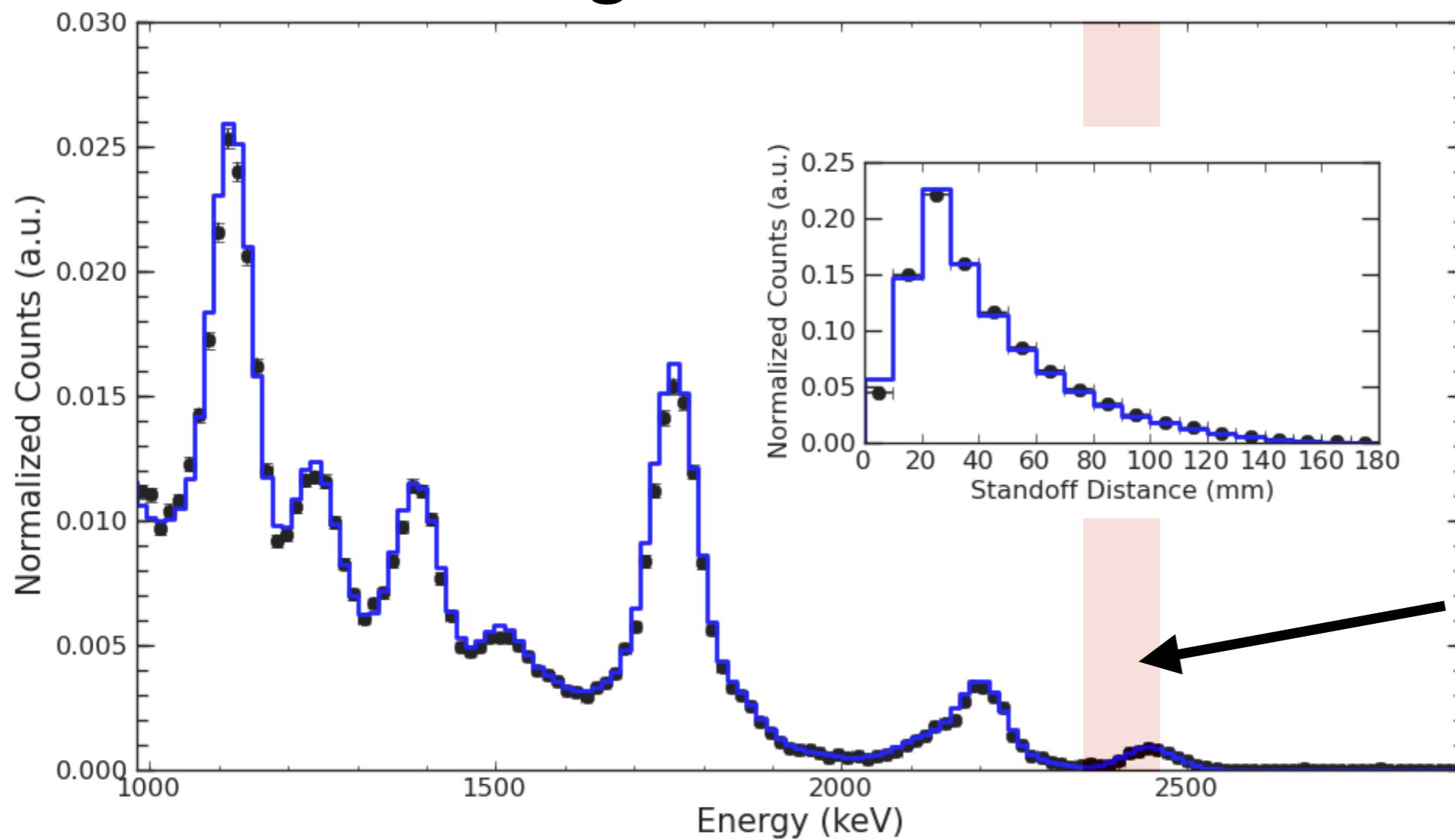
Source:	Relative error [%]:
Background shape distortion	9.2
Choice of background model components	5.7
Variation of energy resolution over time	1.5
Total:	10.9

- Location of $0\nu\beta\beta$ ROI:
 - Deviations between β and γ energy scale: $E_\beta = B \cdot E_\gamma \Rightarrow B = 0.999 \pm 0.002$
 - Single-site fraction error: **9.6%**

“ β -scale” allowed to float in fit

SS/MS fraction allowed to vary within this error in fit

Systematic errors agreement with simulation

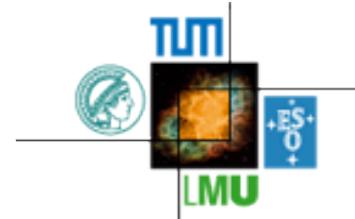


Determine how systematics issues affect background estimates in the ROI.
Allowed background in ROI (red region) to float within this error with respect to the rest of the energy spectrum.

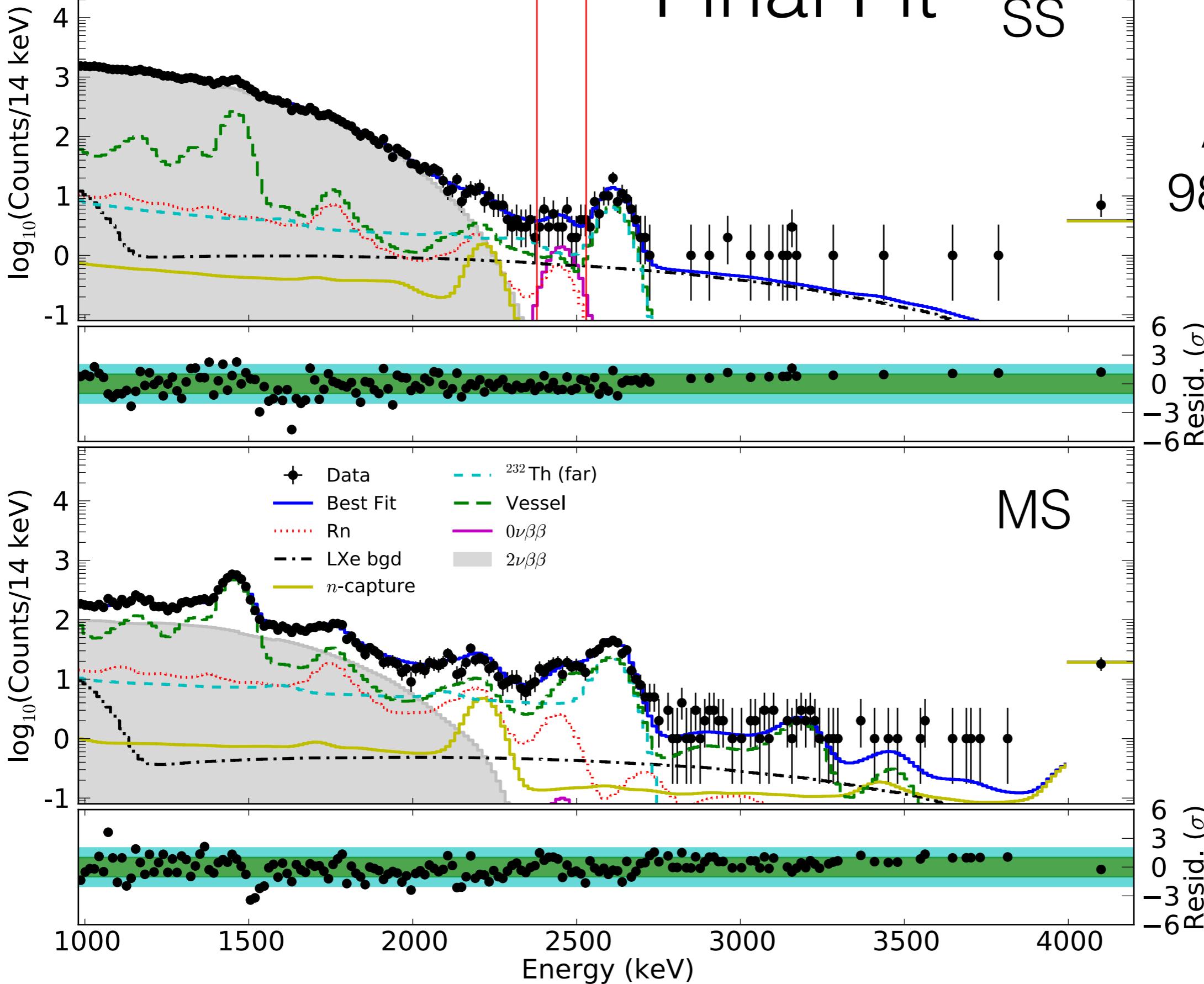
^{226}Ra source data, single-site: example distribution comparison used to estimate systematics (also used ^{228}Th , ^{60}Co , ^{137}Cs sources)

Final Fit

SS



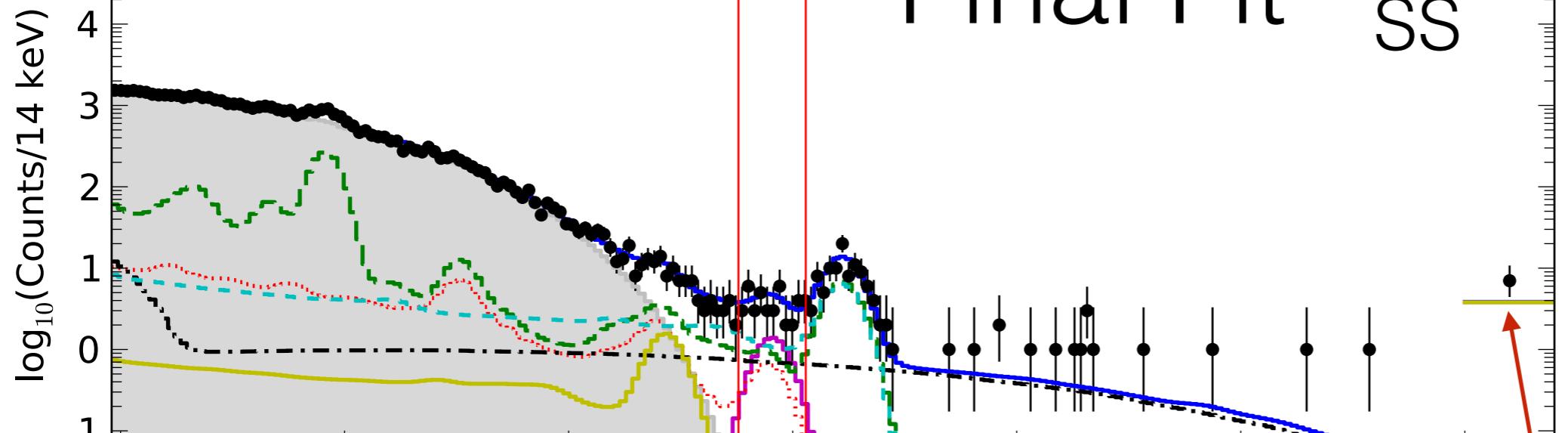
Analysis:
980 to 9800
keV



Final Fit

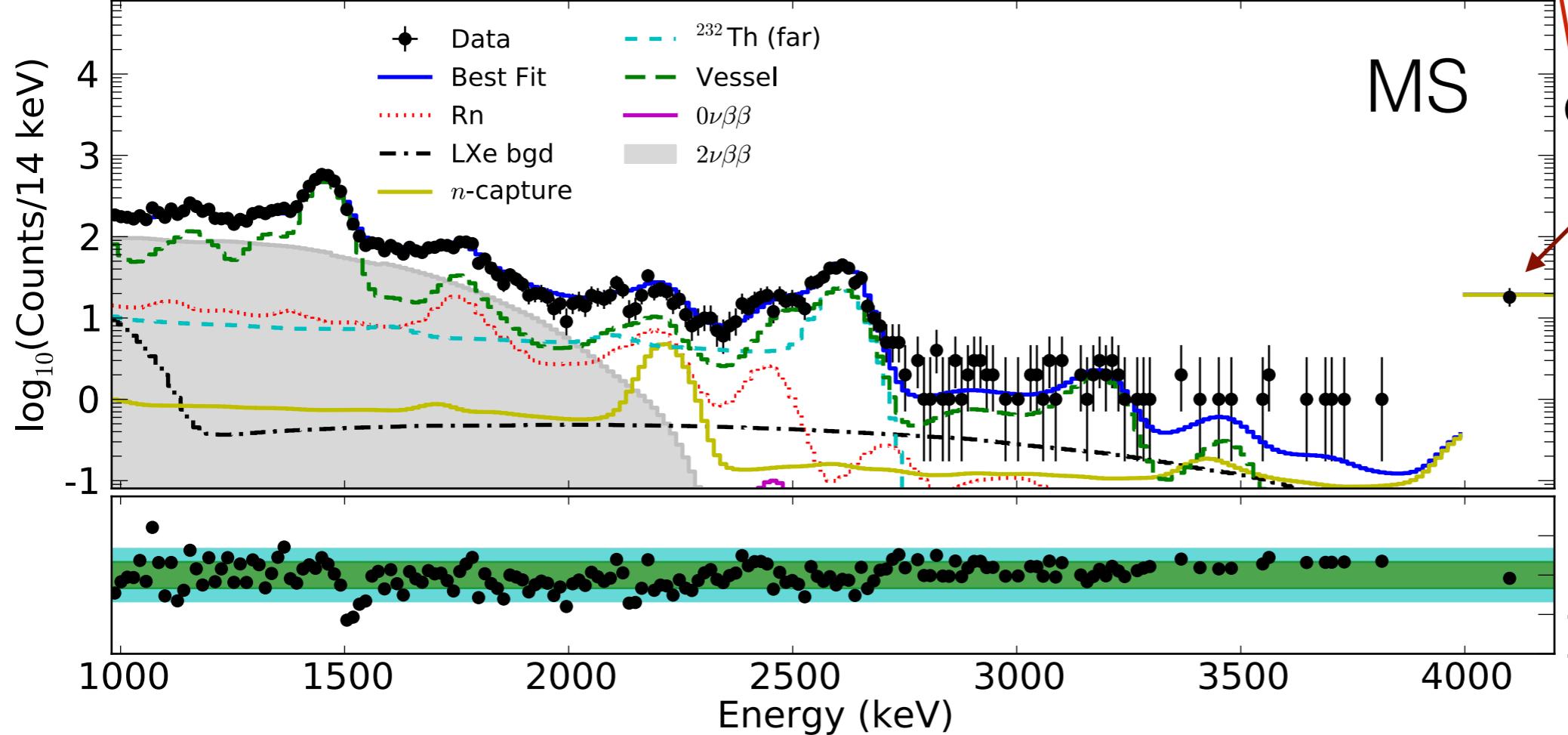
SS

Analysis:
980 to 9800
keV

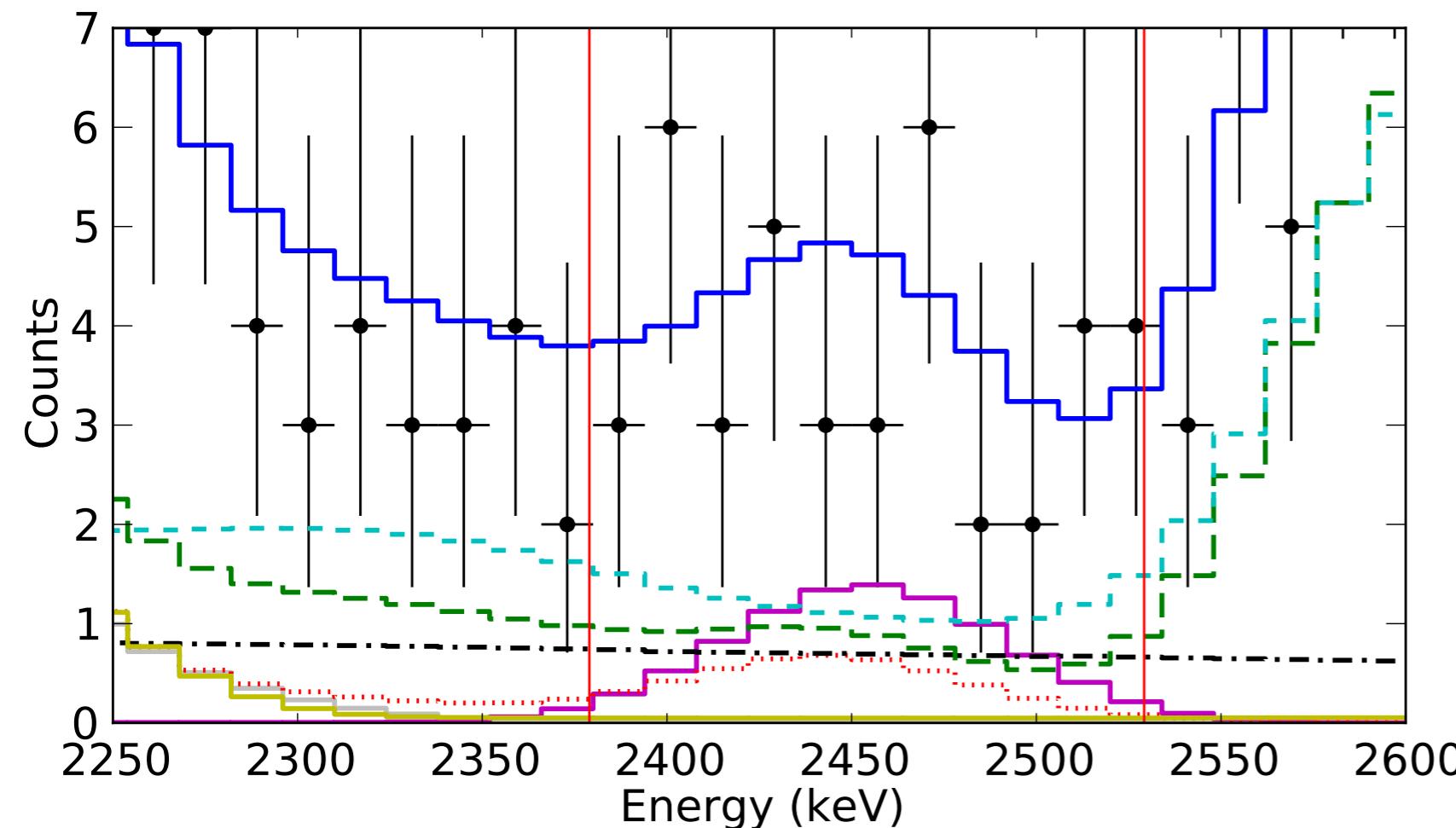


MS

High-energy
counts dominated
by μ -produced
neutrons



Looking for $0\nu\beta\beta$



Backgrounds in $\pm 2\sigma$ ROI

Th-228 chain	16.0
U-232 chain	8.1
Xe-137	7.0
Total	31.1 ± 3.8

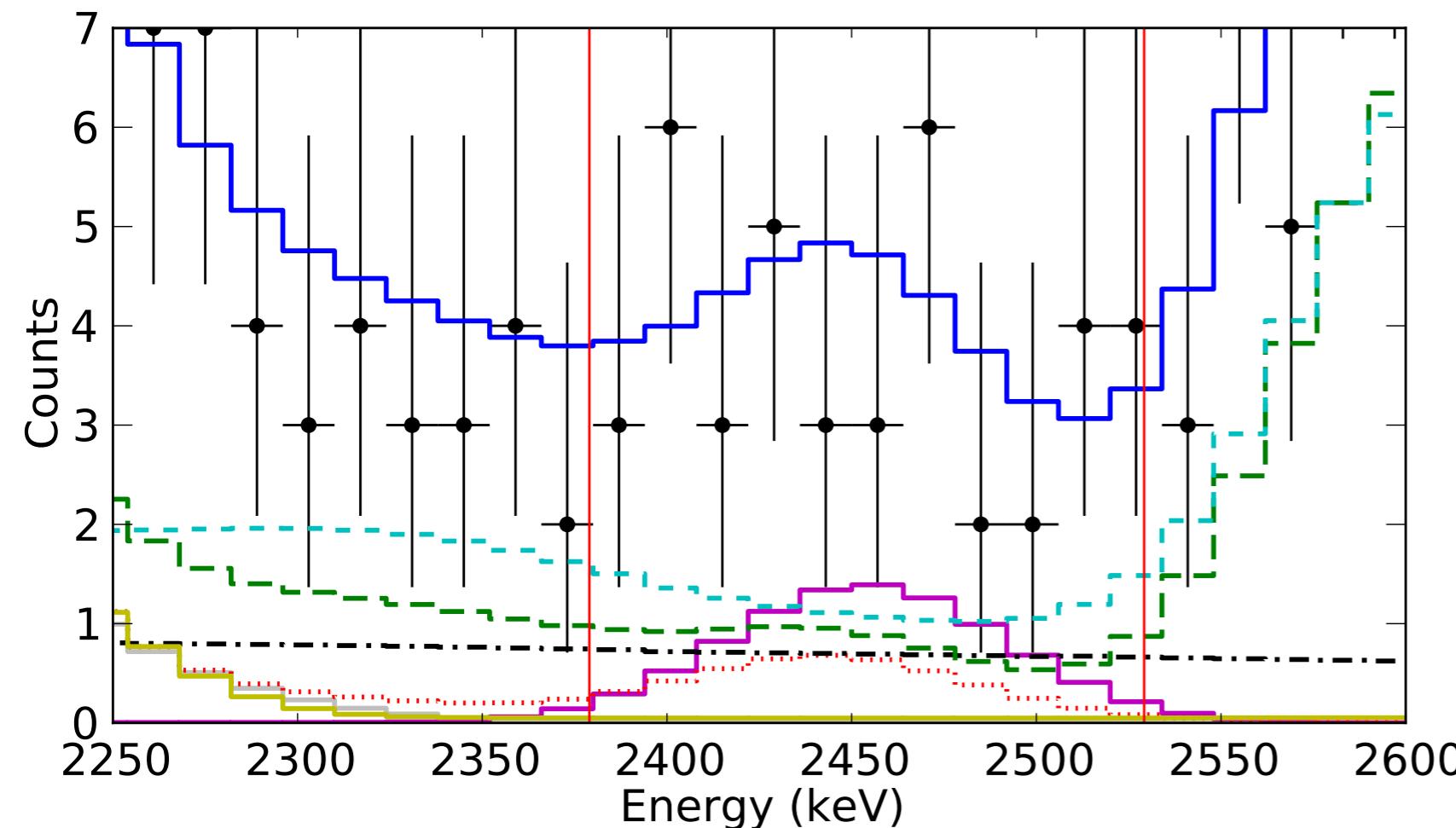
- Data
- Best Fit
- ... Rn
- - - LXe bgd
- n -capture
- ^{232}Th (far)
- Vessel
- $0\nu\beta\beta$
- $2\nu\beta\beta$

From profile likelihood:

$$\begin{aligned} T_{1/2}^{0\nu\beta\beta} &> 1.1 \cdot 10^{25} \text{ yr} \\ \langle m_{\beta\beta} \rangle &< 190 - 450 \text{ meV} \\ &\quad (90\% \text{ C.L.}) \end{aligned}$$

Nature (2014)
doi:10.1038/nature13432

Looking for $0\nu\beta\beta$



Backgrounds in $\pm 2\sigma$ ROI

Th-228 chain	16.0
U-232 chain	8.1
Xe-137	7.0
Total	31.1 ± 3.8

consistent with activation estimates
from muon-veto-tagged data

- Data
- Best Fit
- ... Rn
- - - LXe bgd
- n -capture
- · · · ^{232}Th (far)
- — — Vessel
- 0 $\nu\beta\beta$
- 2 $\nu\beta\beta$

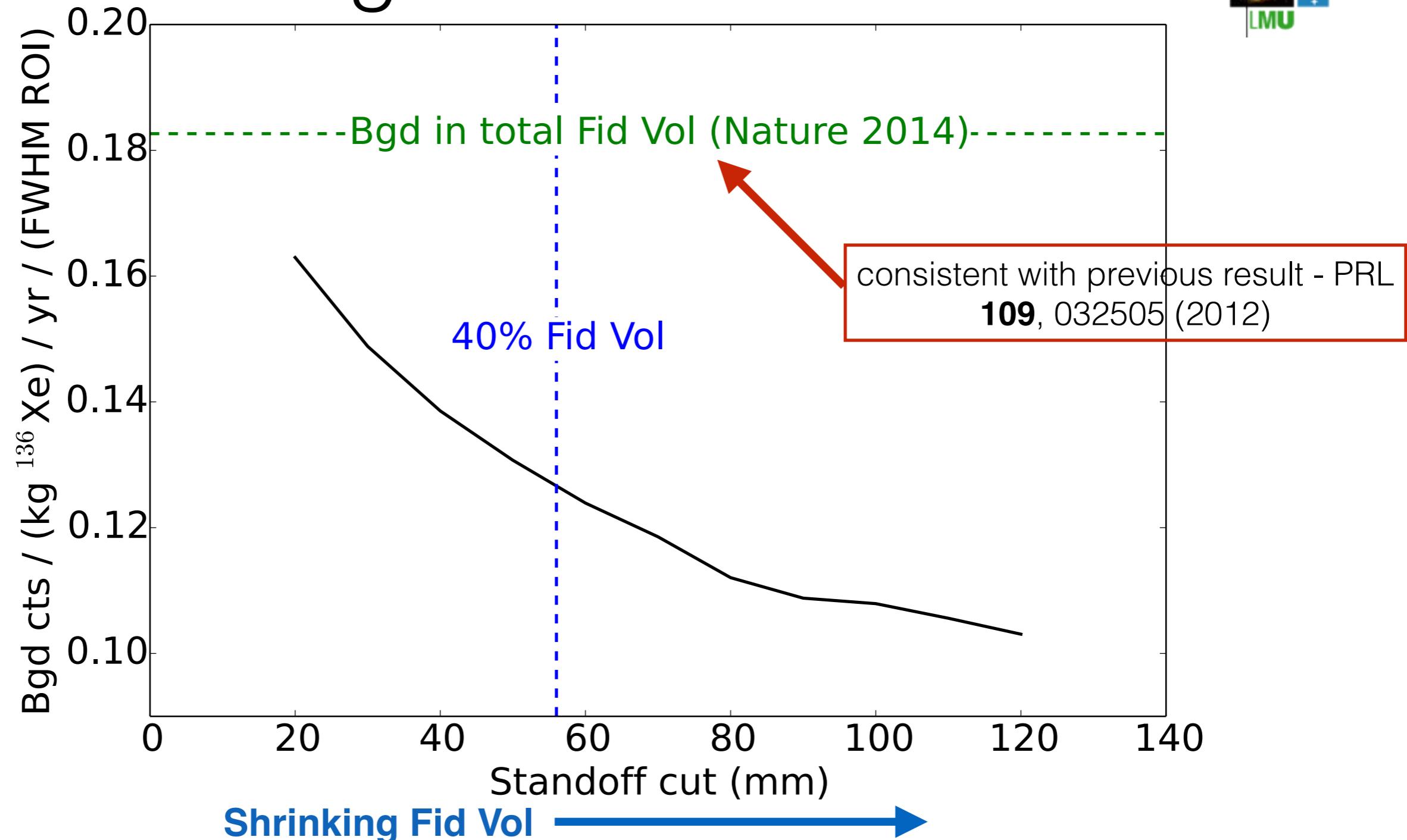
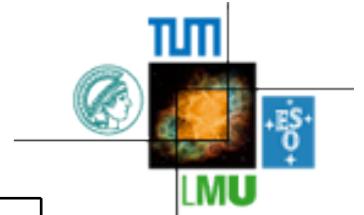
From poster # 96: Cosm.
neutron bgds for
 $0\nu\beta\beta$ search, J. Albert

$$T_{1/2}^{0\nu\beta\beta} > 1.7 \times 10^{25} \text{ yr}$$

$\langle m_{\beta\beta} \rangle < 190 - 450 \text{ meV}$
(90% C.L.)

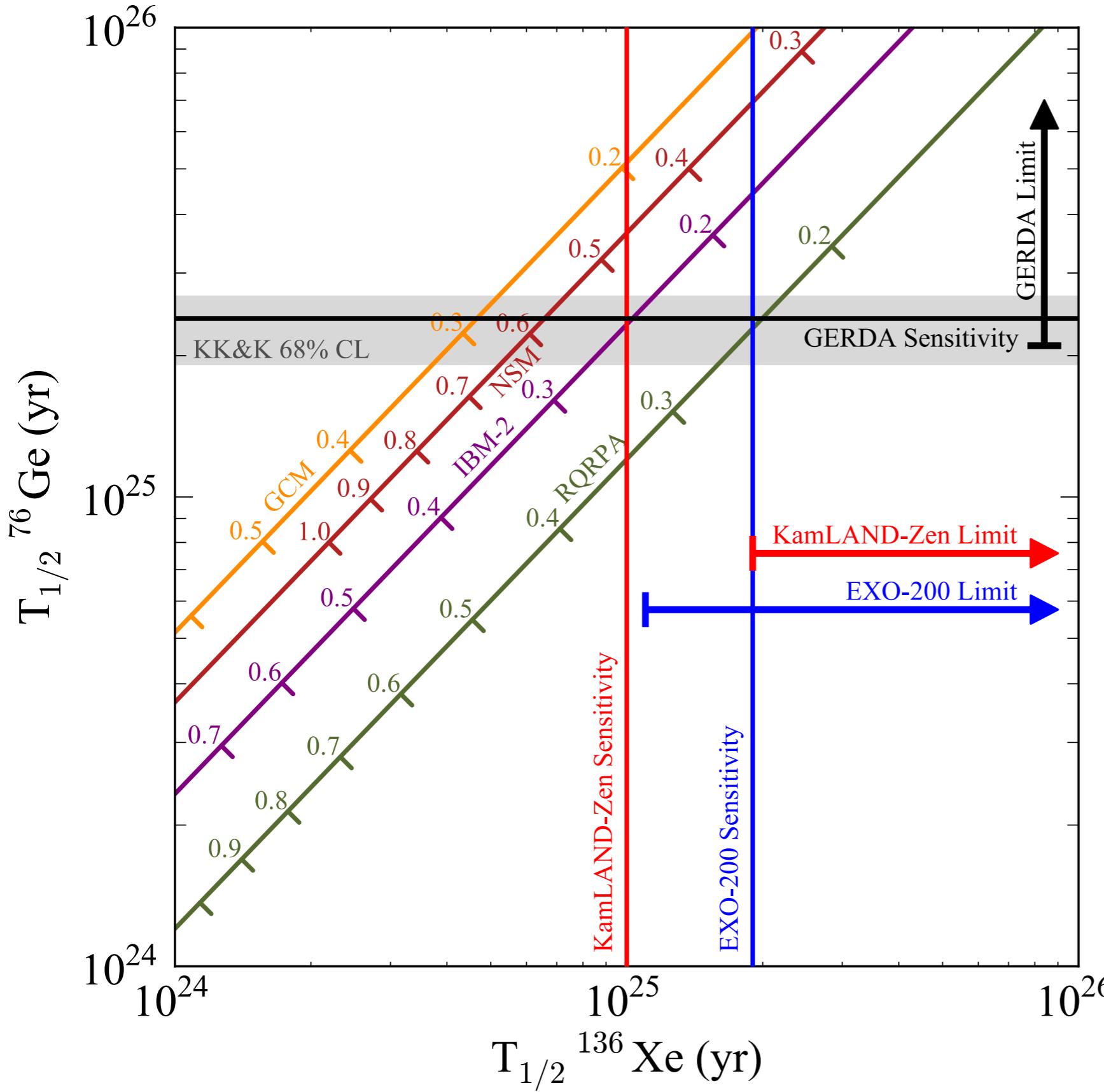
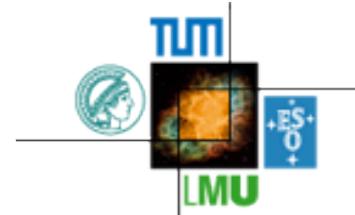
Nature (2014)
doi:10.1038/nature13432

Backgrounds in the ROI



30% reduction of background index in inner 40% fiducial volume (reduction limited by size of TPC)

$0\nu\beta\beta$ status comparison



EXO-200:
Nature (2014),
doi:10.1038/nature13432

GERDA Phase 1:
PRL 111 (2013) 122503

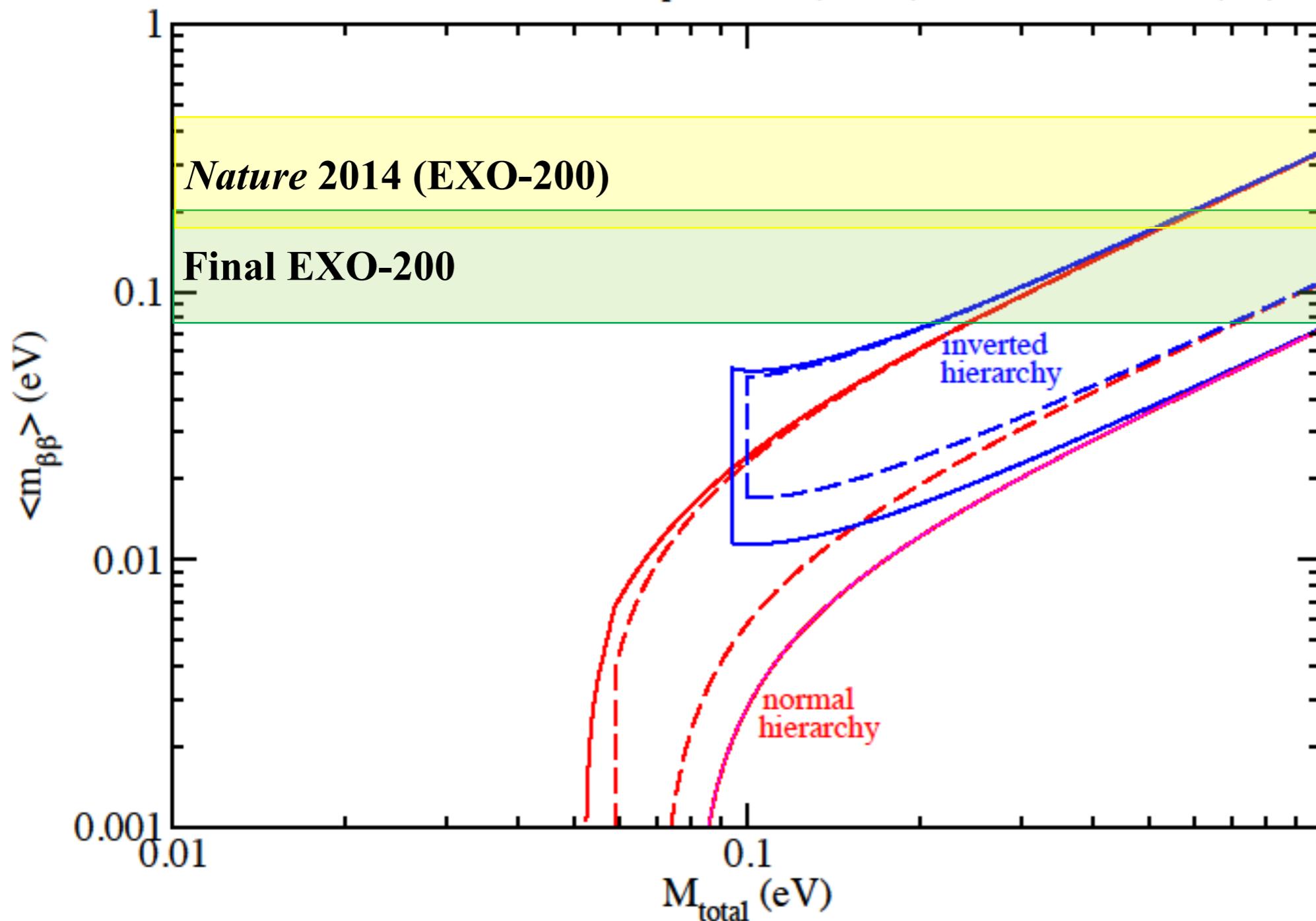
KamLAND-Zen:
PRL 110 (2013) 062502

KK&K Claim:
Mod. Phys. Lett., A21
(2006) 1547

Sensitivity outlook

Effective Majorana mass vs. M_{total}

For the mean values of oscillation parameters (dashed) and for the 3σ errors (full)



EXO-200 ultimate sensitivity (90%CL):
2 years additional livetime with Rn removal

Current EXO-200 status

- **WIPP events:**

- **5 Feb 2014** - Fire in WIPP underground
- **13-14 Feb 2014** - underground access, including with EXO-200 members: "... amount of soot in the project area ... **extensive but recoverable.**"
- **14 Feb 2014, ~23:00** - Airborne radiological event



DOE Accident Inv. Rep., Mar 2014

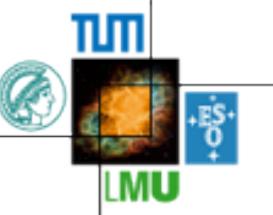
- **EXO-200 status:**

- Was taking (source) data **during above events**. Network/control access has continued.
- **In late February, once date of next access was no longer clear:** With remote system access, Xe was successfully recovered (as designed), TPC/Cryostat was warmed

- **Outlook**, "...reassure our commitment to make access to the unique underground environment at WIPP available to the scientific community." - DOE

- Must still be verified that northern air circuits free of contamination, hoist/ground control operations must return to normal, etc.
- "... **access in the Autumn time frame....**"

Summary



- EXO-200 is among the most sensitive $0\nu\beta\beta$ experiments, achieving $100 \text{ kg}\cdot\text{yr}$ of ^{136}Xe exposure. Substantial improvements in understanding systematics, improving energy resolution have been recently achieved.
- Data taking is currently halted due to WIPP closure, planned access in Fall 2014:
 - Planned deployment of new electronics, deradonator
- nEXO plans to take advantage of the scalability of LXe in a next-generation experiment.

See L. Winslow's talk and:

nEXO Posters:

94 The nEXO Experiment: L. Wen

95 Ba tagging for nEXO: B. Fairbank

98 EXO-200 detector performance, lessons
for nEXO: E. Smith, Y.-R. Yen

99 Radio assay for nEXO: D. Auty

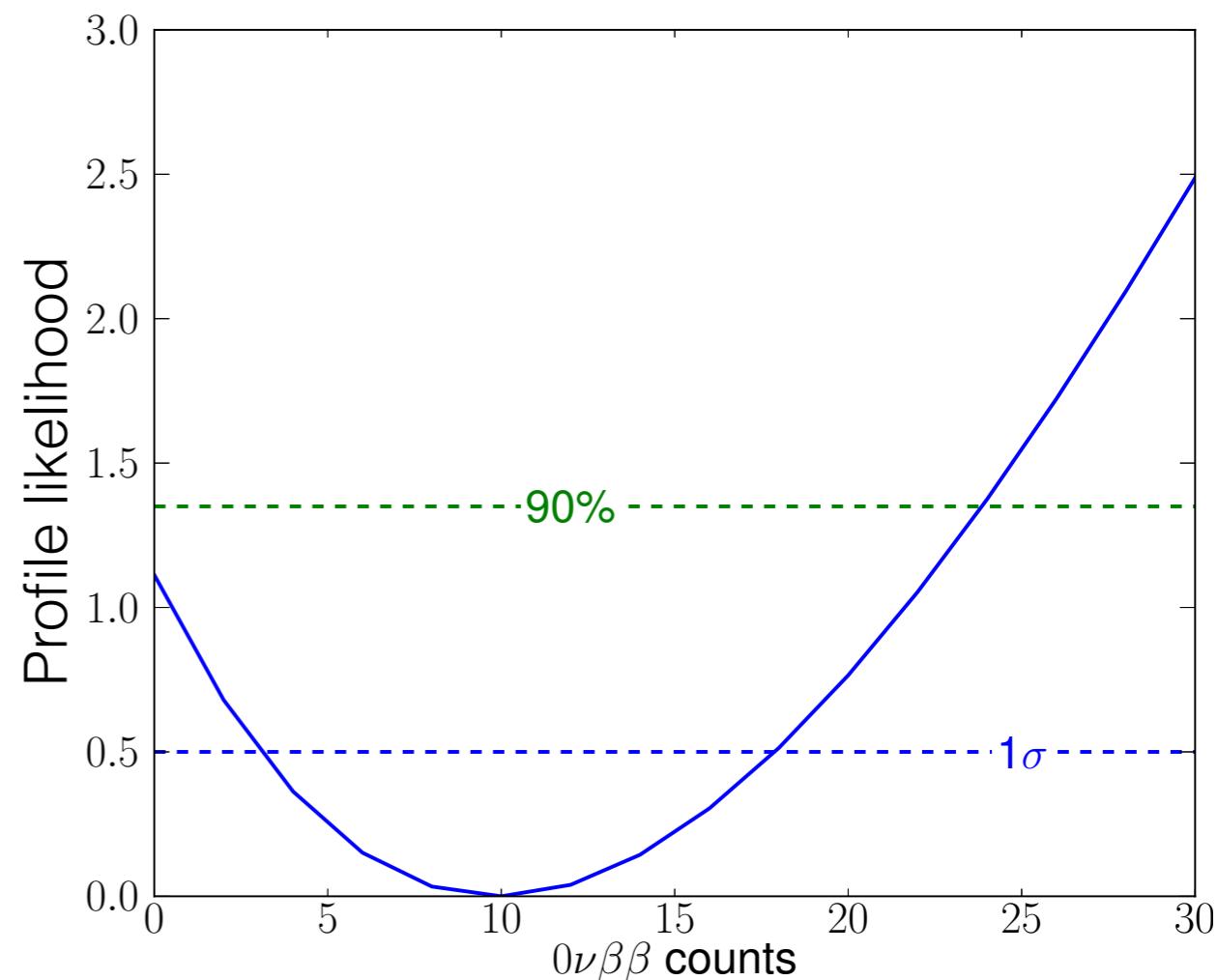
$T_{1/2}^{0\nu\beta\beta} > 1.1 \cdot 10^{25} \text{ yr}$

Sensitivity: $1.9 \cdot 10^{25} \text{ yr}$

$\langle m_{\beta\beta} \rangle < 190 - 450 \text{ meV (90\% C.L.)}$

Nature (2014)

doi:10.1038/nature13432



The EXO Collaboration



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Extra slides

Final Fit, Standoff Distance projection

