

Reactor Experiments and Underground Capabilities

Karsten Heeger
Yale University

Snowmass on the Mississippi, July 30, 2013

Reactor Neutrinos

55 years of liquid scintillator detectors with varying baselines.

2012 - Measurement of θ_{13} with Reactor Neutrinos

2008 - Precision measurement of Δm_{12}^2 . Evidence for oscillation

2003 - First observation of reactor antineutrino disappearance



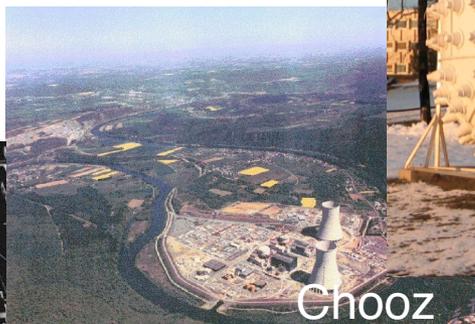
1995 - Nobel Prize to Fred Reines at UC Irvine

1980s & 1990s - Reactor neutrino flux measurements in U.S. and Europe

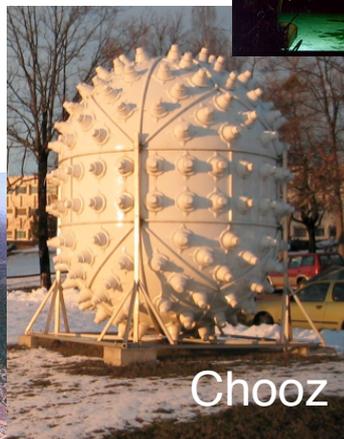
1956 - First observation of (anti)neutrinos



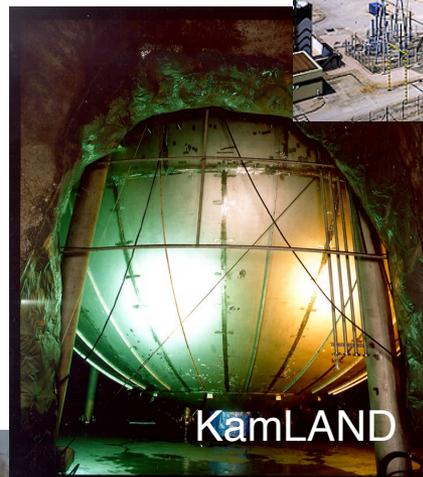
Savannah River



Chooz



Chooz



KamLAND



Daya Bay, Double Chooz, RENO

Past Reactor Experiments

- Hanford
- Savannah River
- ILL, France
- Bugey, France
- Rovno, Russia
- Goesgen, Switzerland
- Krasnoyarsk, Russia
- Palo Verde
- Chooz, France

Finding the right underground location near reactors....

Reactor Antineutrinos

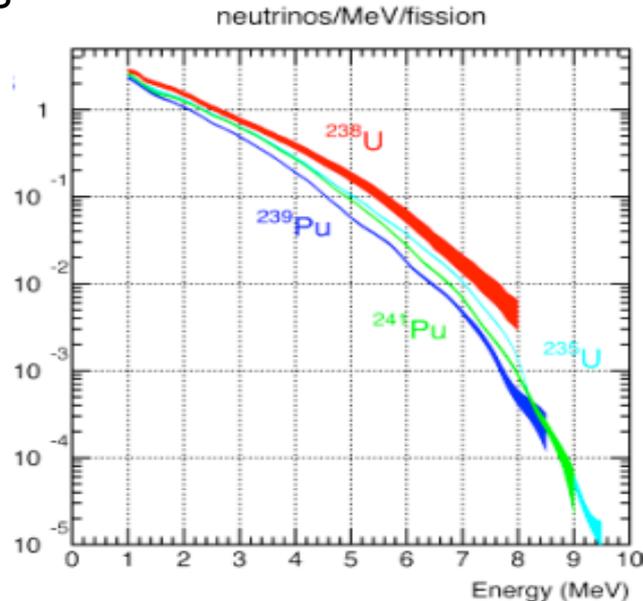
Source

$\bar{\nu}_e$ from β -decays
of n-rich fission products



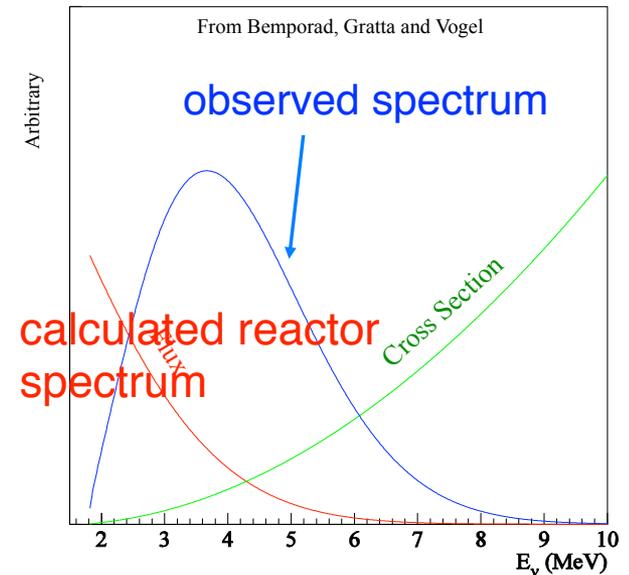
pure $\bar{\nu}_e$ source

> 99.9% of $\bar{\nu}_e$ are produced by fissions in ^{235}U ,
 ^{238}U , ^{239}Pu , ^{241}Pu



Detection

inverse beta decay
 $\bar{\nu}_e + p \rightarrow e^+ + n$

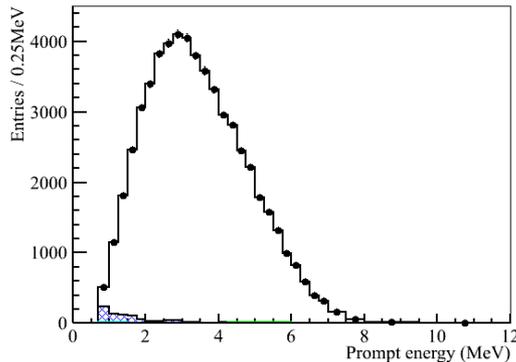
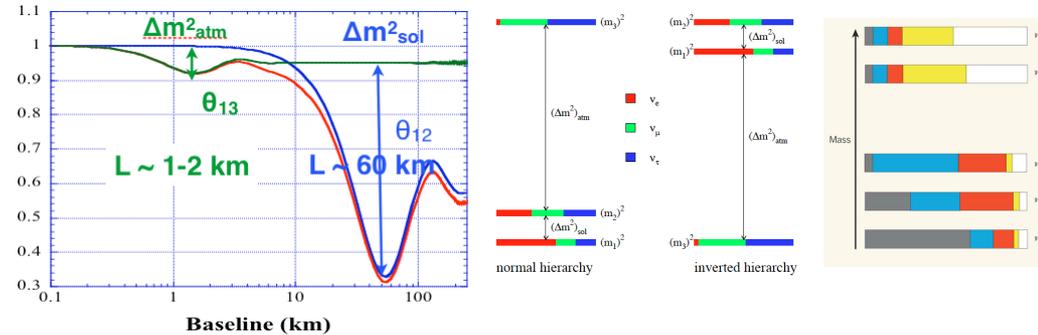


mean energy of $\bar{\nu}_e$: 3.6 MeV
only disappearance
experiments possible

Scientific Accomplishments and Goals

Neutrino Oscillation Experiments

- precision measurement of oscillation parameters Δm^2_{21} , Δm^2_{ee} , $\sin^2 2\theta_{13}$
- search for short-baseline oscillations
- probing the mass hierarchy



Reactor Antineutrino Spectra, Fuel, Cores

- precision measurement of reactor antineutrino spectrum
- studying reactor core evolution
- absolute measurement of reactor antineutrino flux

Detector development

- large, homogeneous low-background scintillator detectors underground
- R&D towards segmented detectors near surface



US researchers have played leading roles in recent and past experiments

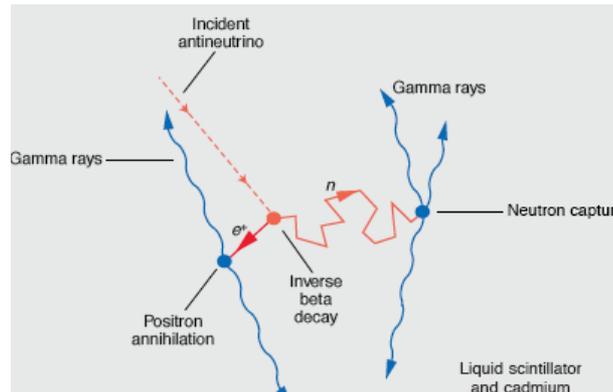
Detecting Antineutrinos

inverse beta decay



signal: delayed coincidence
between positron and neutron
capture

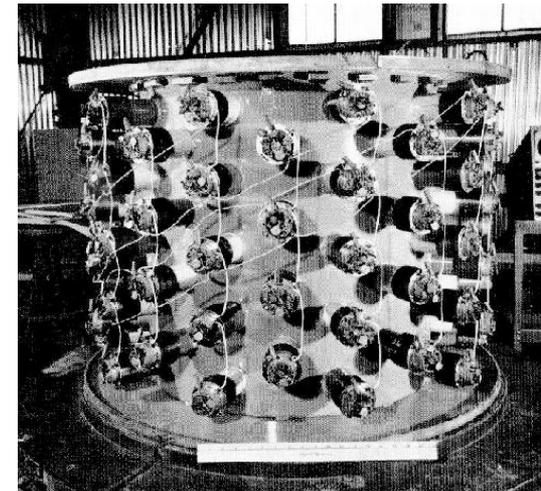
Shielding: 4 ft of soaked sawdust



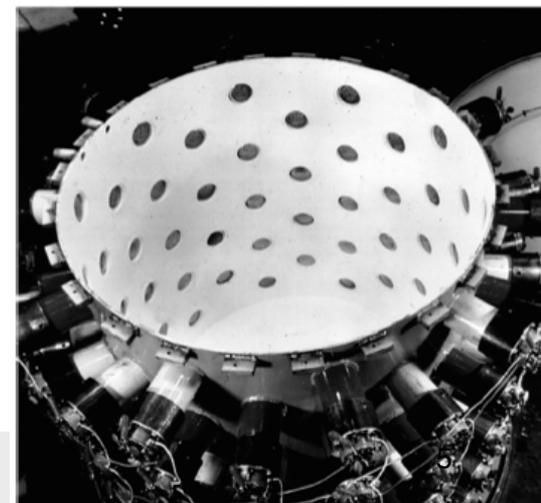
pioneering efforts by Reines
and Cowan in the US



Reines, Cowan



*early experiments founds that shielding and background
reduction are essential*



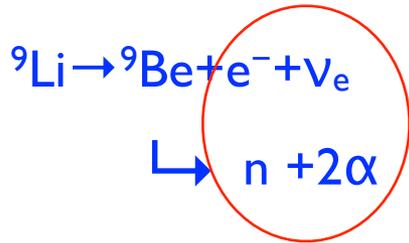
Backgrounds in Reactor Experiments

accidentals

cosmogenic-induced backgrounds

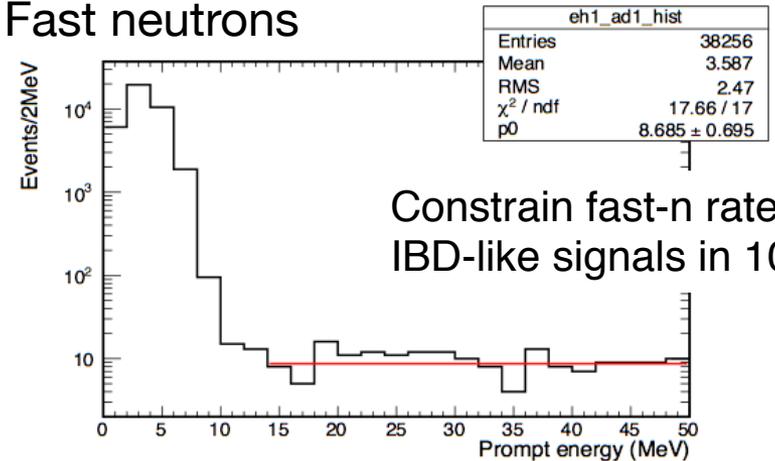
correlated events can mimic antineutrino signal

- prompt: β -decay
- delayed: neutron capture



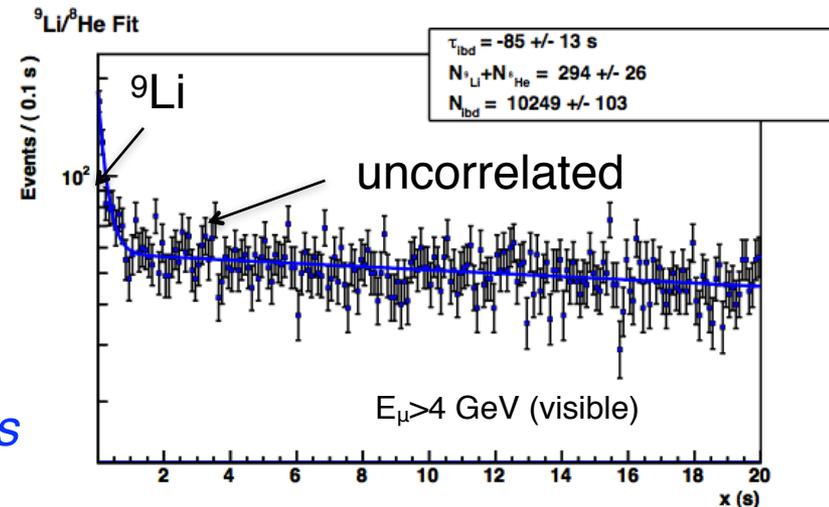
background considerations require radiopure detectors in underground locations

Fast neutrons



Constrain fast-n rate using IBD-like signals in 10-50

Correlated β -n decay

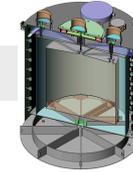


Reactor Neutrino Oscillation Experiments



$\bar{\nu}_e$

$\bar{\nu}_{e,x}$



far

$\bar{\nu}_{e,x}$

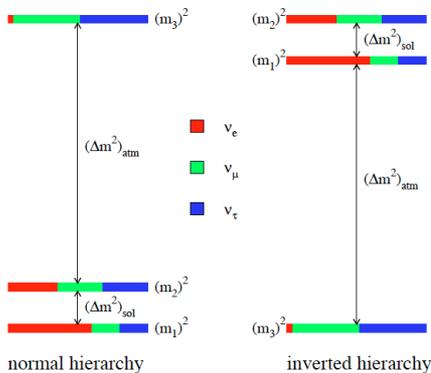
Measure (non)- $1/r^2$ behavior of ν_e interaction rate

$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_\nu} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E_\nu} \right)$$

for 3 active ν , two different oscillation length scales: $\Delta m_{12}^2, \Delta m_{23}^2$

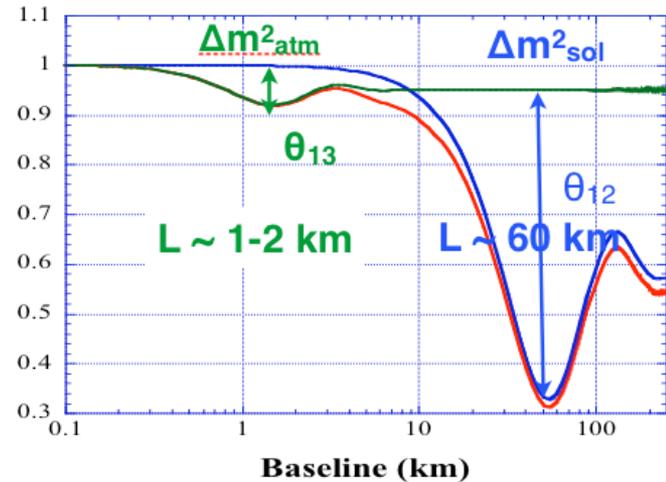
$L/E \rightarrow \Delta m^2$

amplitude of oscillation $\rightarrow \theta$



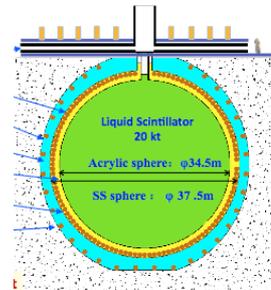
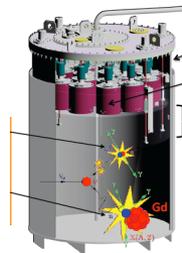
$$\Delta m_{12}^2 \sim 7.6 \times 10^{-5} \text{ eV}^2$$

$$\Delta m_{23}^2 \sim 2.4 \times 10^{-3} \text{ eV}^2$$



Reactor Experiments at Different Scales

Past&Present



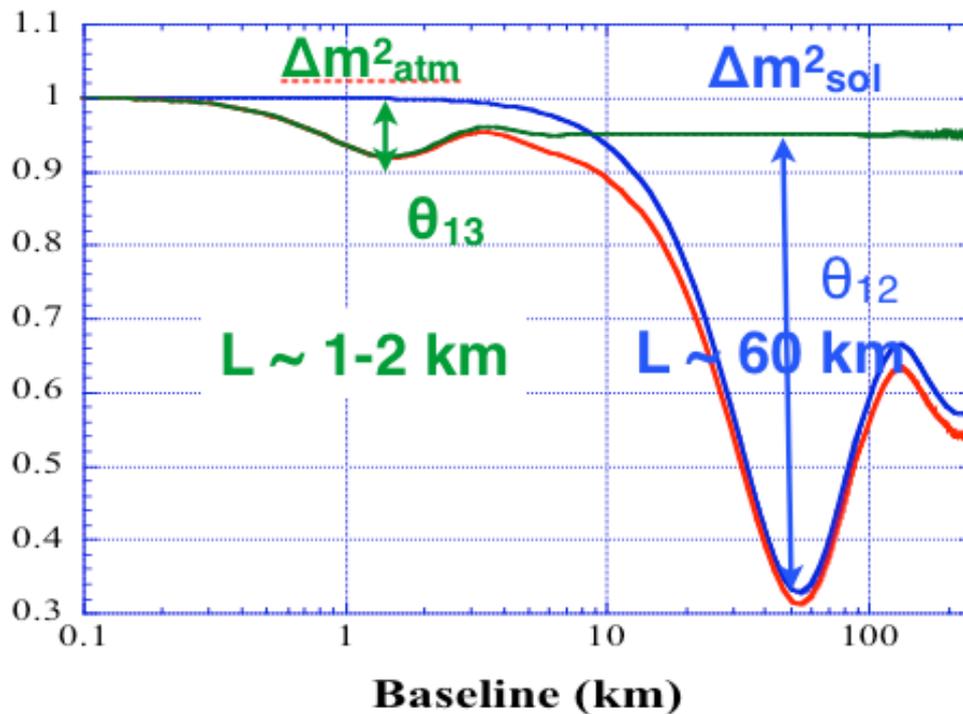
Future

1 ton/10m

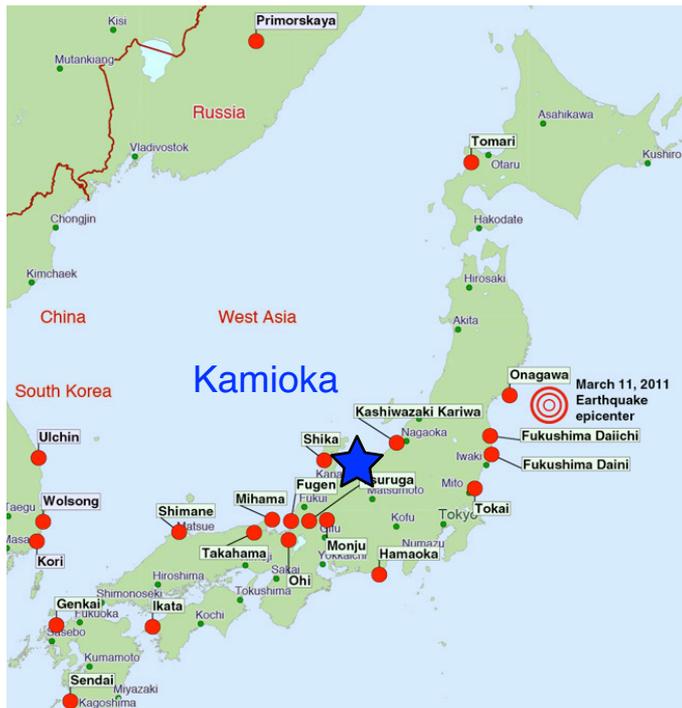
20 ton/1km

20 kton/50km

1 kton/180km



KamLAND at ~180km Baseline

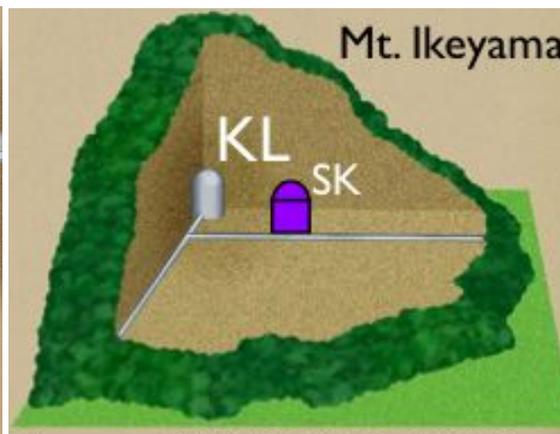
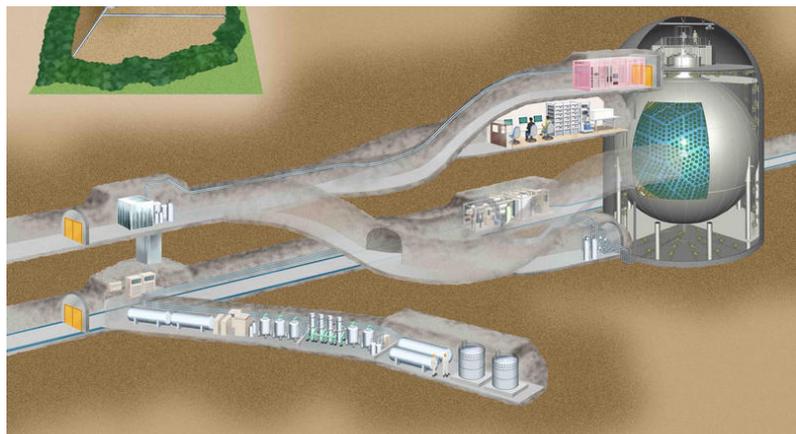
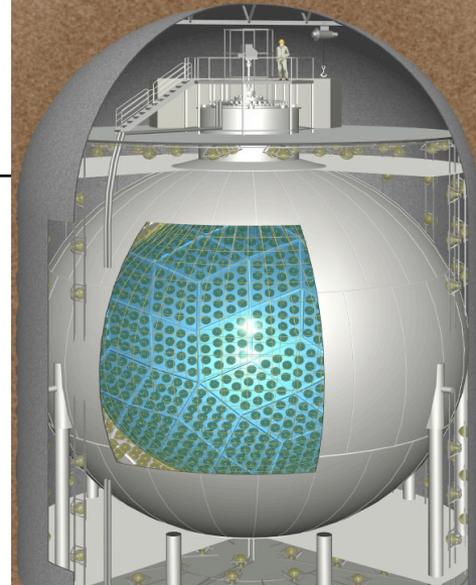


measures antineutrinos from 55 reactors

mean, flux-weighted reactor distance ~ 180km

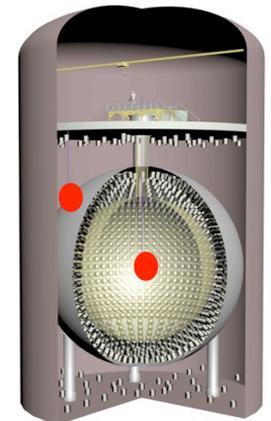
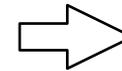
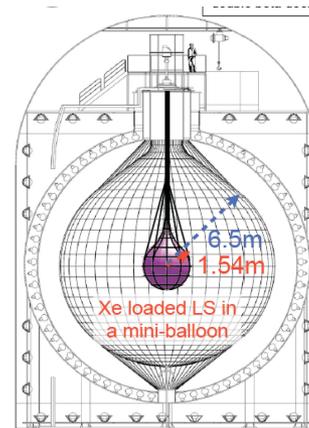
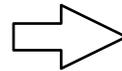
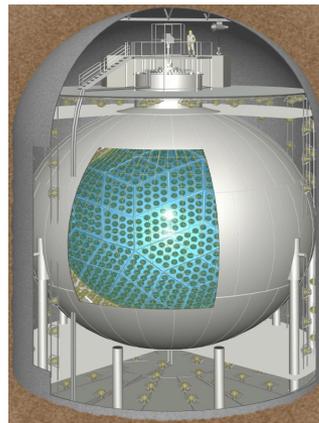
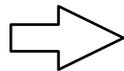
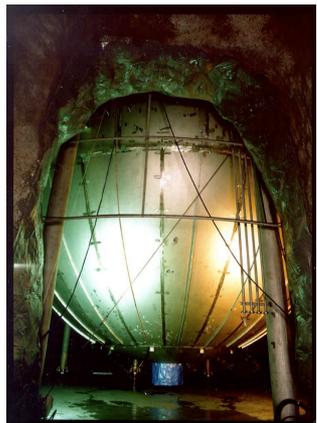
2700 mwe overburden

1kt liquid scintillator detector in dedicated underground facility



Kamiokande → KamLAND → KamLAND-Zen

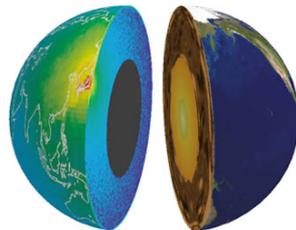
Underground excavation at Kamioka has enabled a long-term program of underground science



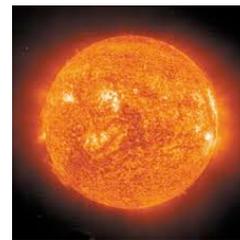
reactor ν



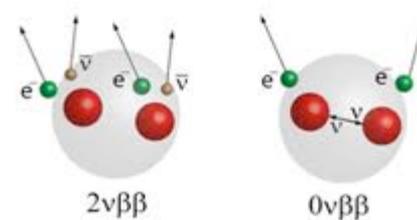
geo ν



solar ν



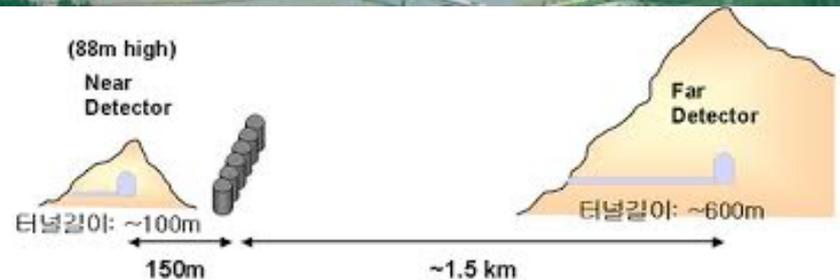
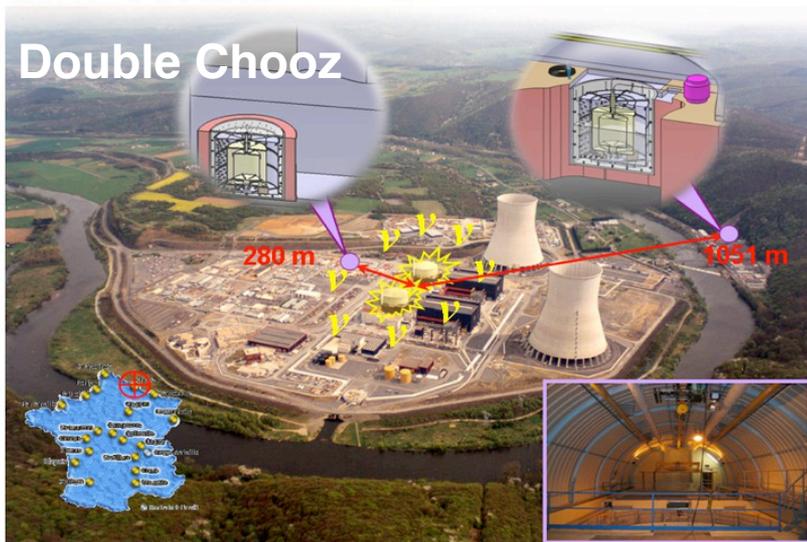
$0\nu\beta\beta$



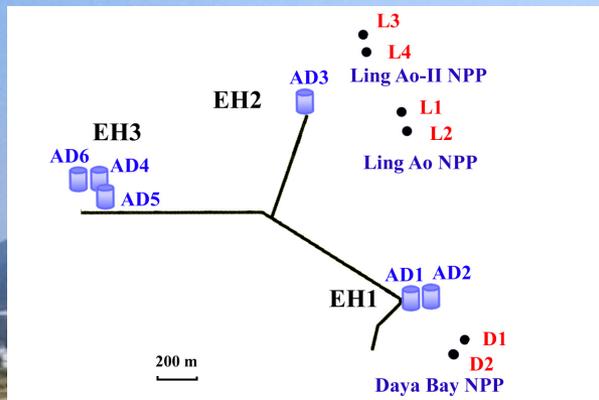
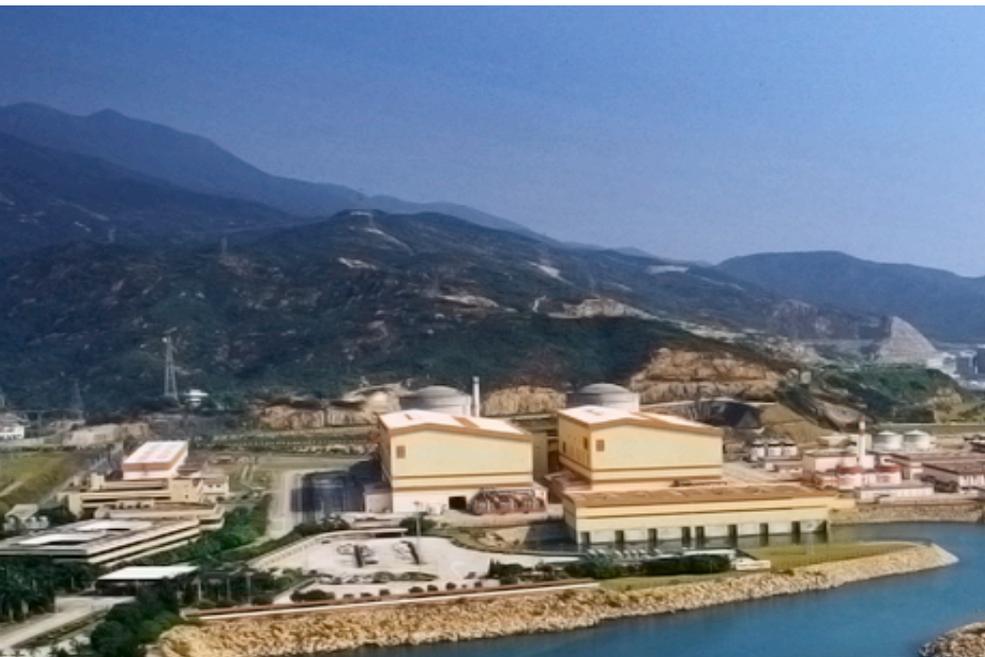
sterile ν ?

Reactor Experiments at ~1km Baseline

Facilities Designed for an Optimized Oscillation Measurement



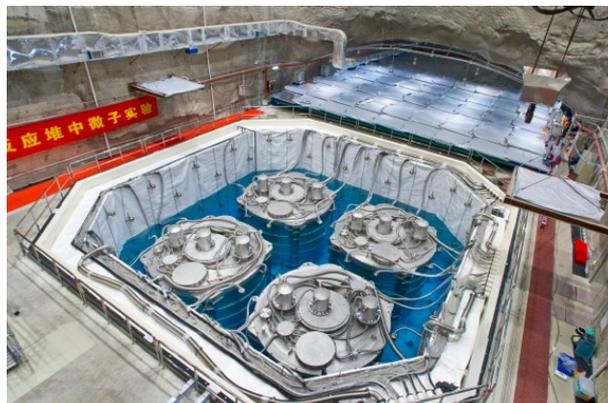
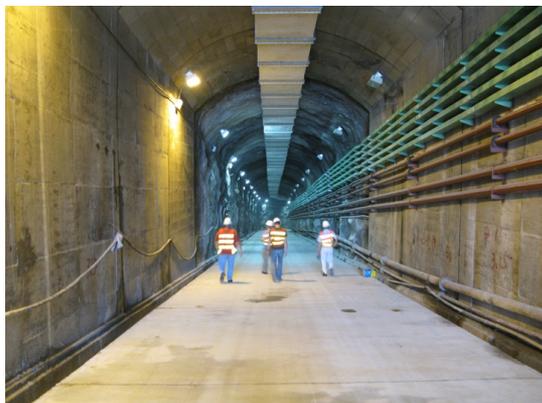
Reactor Experiments at ~1km Baseline



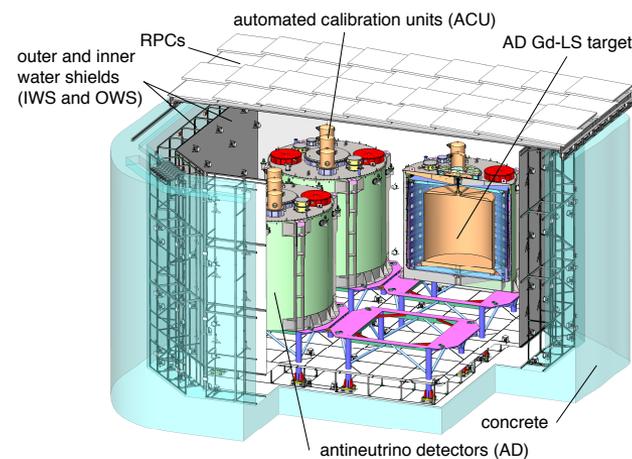
Example:
Daya Bay



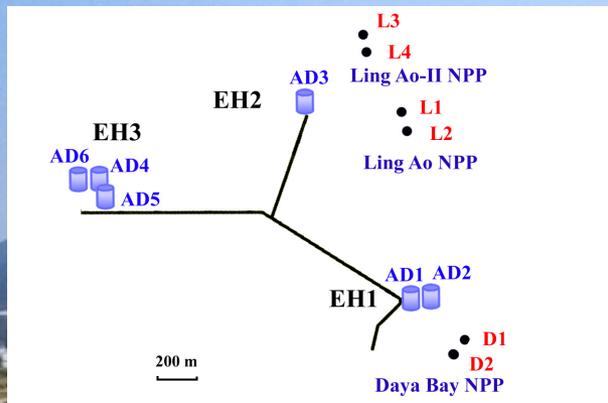
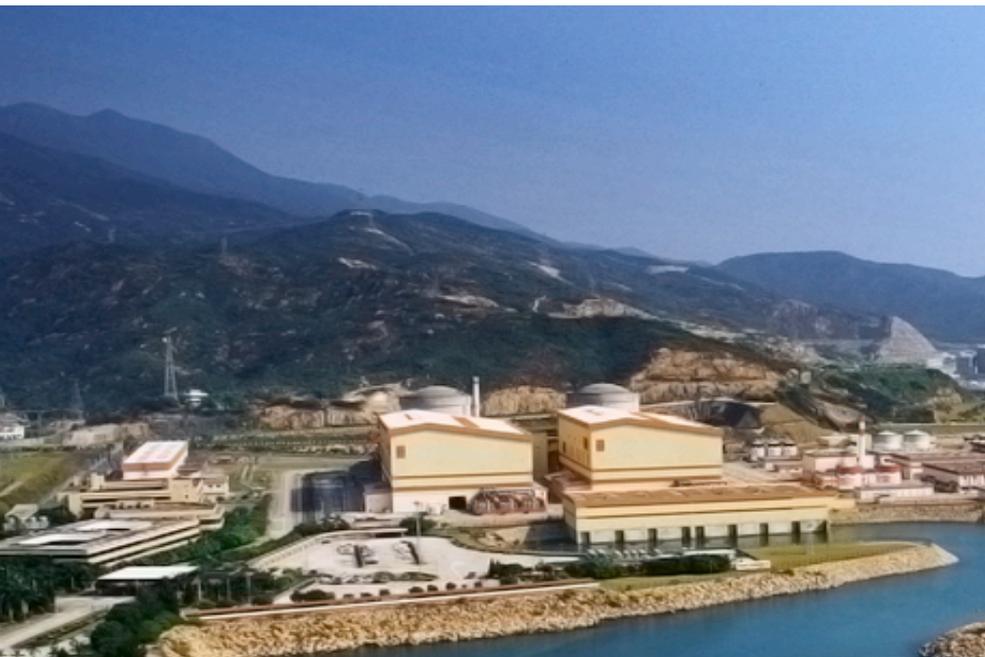
6 reactor cores
3 experimental halls
8 detectors



specialized underground infrastructure



Reactor Experiments at ~1km Baseline



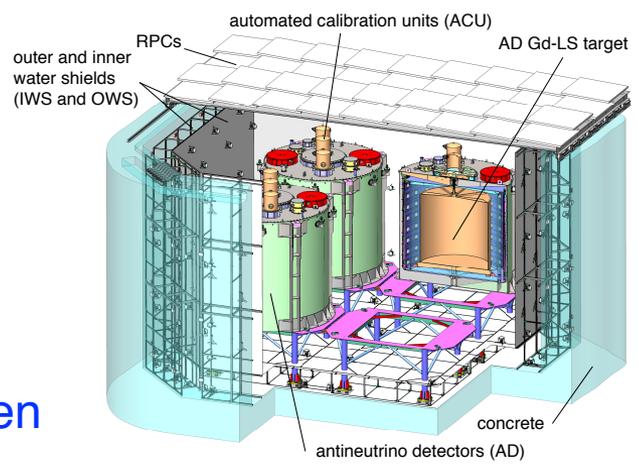
Example:
Daya Bay



6 reactor cores
3 experimental halls
8 detectors

	Overburden	R_μ	E_μ	D1,2	L1,2	L3,4
EH1	250	1.27	57	364	857	1307
EH2	265	0.95	58	1348	480	528
EH3	860	0.056	137	1912	1540	1548

experimental halls with several hundred mwe overburden

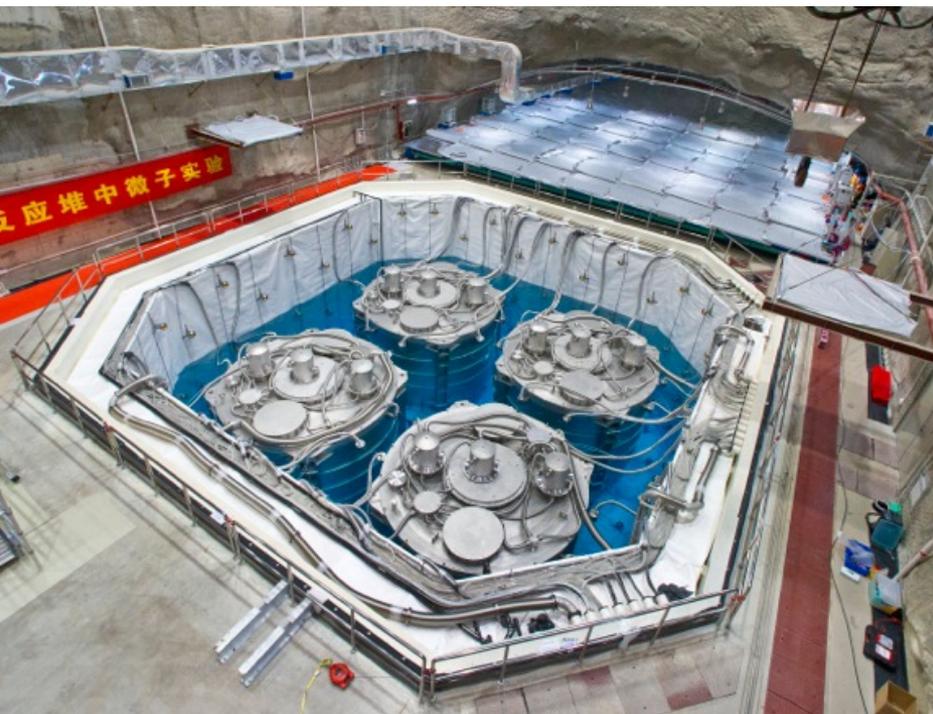


Reactor Experiments at ~1km Baseline

Experimental halls and underground infrastructure enable new R&D and auxiliary underground measurements

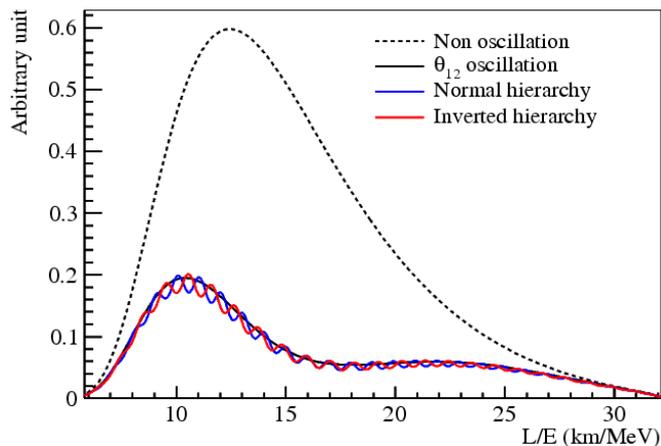
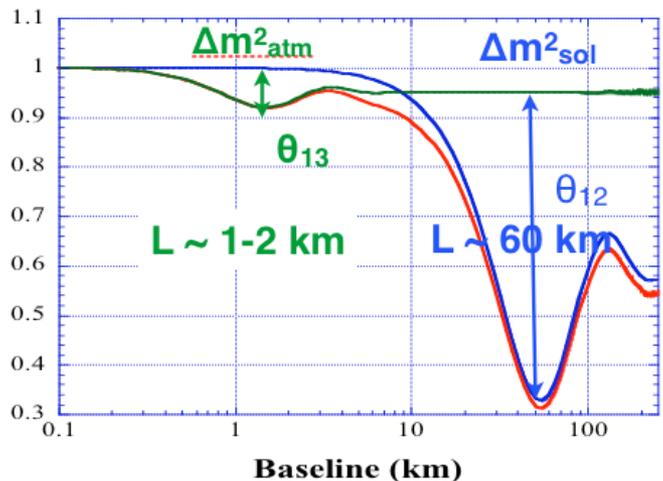
- R&D towards dark matter experiment in liquid storage pools
- muon and cosmogenic background studies

Example:
Daya Bay



Reactor Experiments at ~50km Baseline

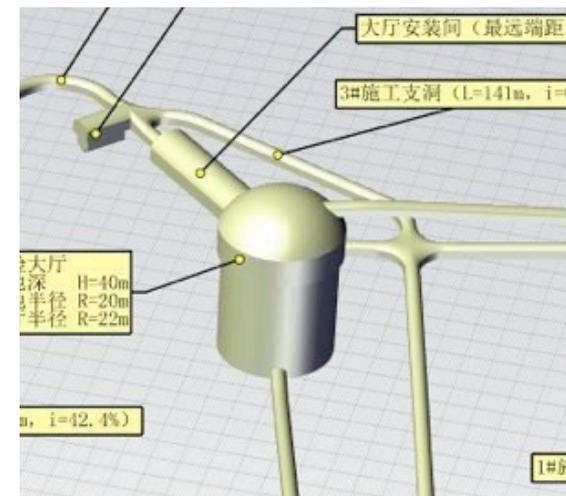
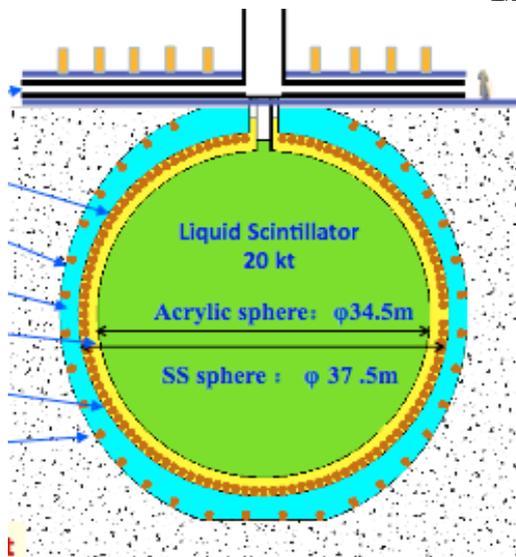
JUNO and RENO-50 (proposed)



aim to make precision measurement of reactor spectrum to determine mass hierarchy

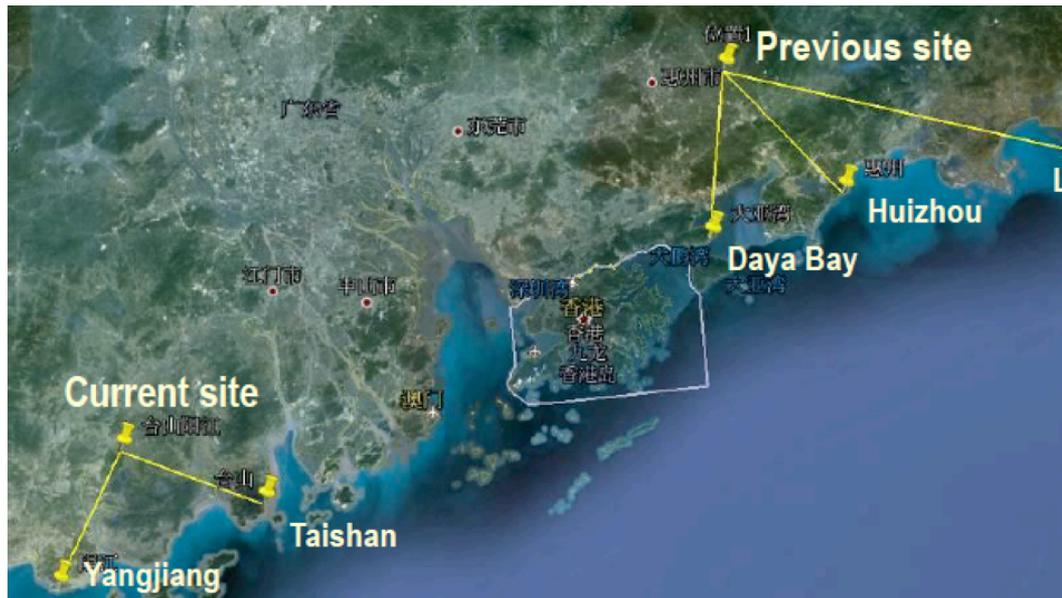
20x larger than KamLAND

baseline requires large precision detectors, and supporting underground facilities



Reactor Experiments at ~50km Baseline

JUNO (proposed)



Site selection

- ◆ Allowed region determined
- ◆ Experimental hall selected:
 - ⇒ In granite
 - ⇒ Mountain height: 270 m
- ◆ Preliminary geological survey completed:
 - ⇒ Review held on Dec. 17, 2012
 - ⇒ No show-stoppers
- ◆ Detailed geological survey started and first round data are available now
- ◆ Contacts with local government established, good support

From Y. Wang

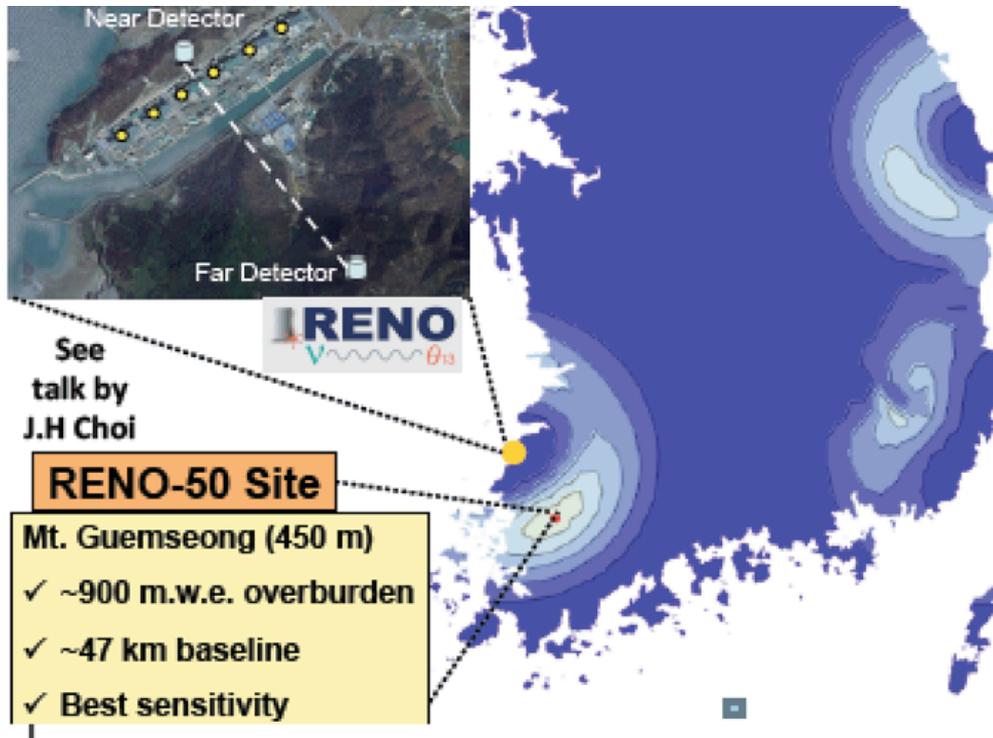
R&D from 2012-2015

Construction start ~ 2015/16?

Early funding commitments in host country (China)

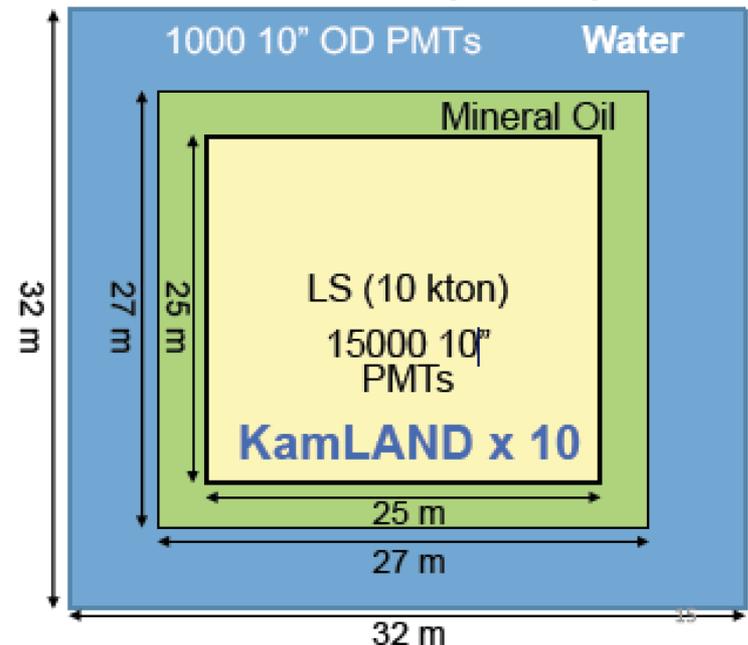
Reactor Experiments at ~50km Baseline

RENO-50 (proposed)



- Utilizing the current 6 RENO reactors
- Baseline ~47km
- Target mass 10kt
- Cylinder-shaped detector
- ➔ Simulation resolution is ~6% at 1MeV
- ➔ Need to improve photoelectrons

RENO-50 (default)



From W. Wang and RENO-50 workshop

Reactor Experiments at ~50km Baseline

Challenges of a 20kt LS Detector

- ◆ Large detector: >10 kt LS
- ◆ Energy resolution: $< 3\%/\sqrt{E}$ → 1200 p.e./MeV

	Daya Bay	BOREXINO	KamLAND	JUNO
LS mass	20t	~300t (100t F.V.)	~1 kt	20kt
Photocathode Coverage	~12%	~34%	~34%	?
Energy Resolution	~7.5%/√E	~5%/√E	~6%/√E	3%/√E
Light yield	~160 p.e./MeV	~500 p.e./MeV	~250 p.e./MeV	1200 p.e./MeV

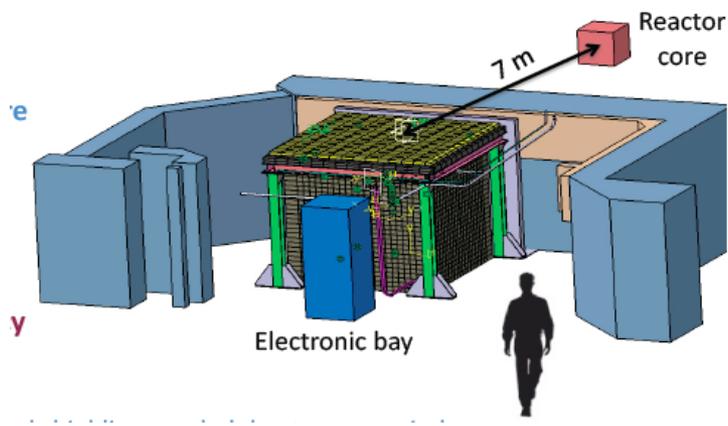
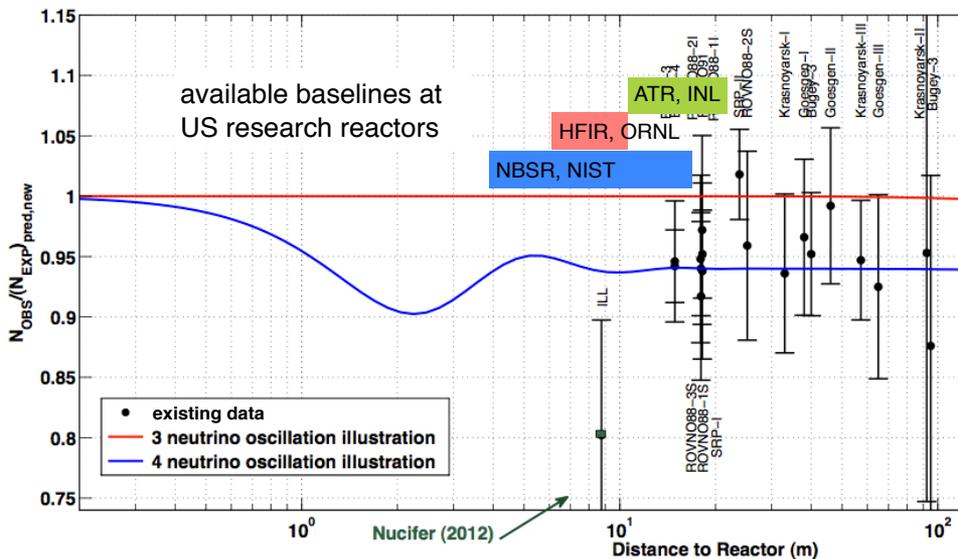
From Y. Wang

50-km baseline reactor experiments require large underground facilities and precision reactor antineutrino detectors

Reactor Experiments at Short Baseline (~10m)

Short-Baseline Experiments

- short-baseline requires much smaller detectors (ton-scale)
- several proposed experiments worldwide
- most of them near surface with little overburden
- considering segmented detector to mitigate backgrounds



Conclusions

- Reactor antineutrinos as probe of neutrino properties (oscillation parameters, mass hierarchy, short-baseline searches, magnetic moment)
- Detectors for reactor experiments at $> 100\text{m}$ baseline require medium-depth underground laboratories (several hundred mwe overburden).
- Strong US involvement in recent reactor experiments overseas (KamLAND, Daya Bay, Double Chooz)
- Reactor neutrino experiments and associated underground spaces have enabled R&D and development of new neutrino experimental initiatives overseas:
 - KamLAND \rightarrow KamLAND-Zen, CeLAND
 - Daya Bay \rightarrow JUNO, R&D for dark matter experiments in Daya Bay halls
 - RENO \rightarrow RENO-50
- Worldwide efforts towards future reactor experiments at medium baseline ($\sim 50\text{km}$) and short-baseline ($\sim 10\text{m}$). Funding commitments from host countries (RENO-50, JUNO, STEREO). May have US involvement. Opportunities for short-baseline searches in US.
- Future reactor experiments and underground space:
 - planned construction of dedicated underground space overseas for medium-baseline reactor experiments
 - multi-purpose use of underground space for neutrino and dark matter R&D in host countries
 - synergies with non-proliferation effort in US (see A. Bernstein's talk)

