

A rectilinear channel for muon cooling towards micron scale emittances

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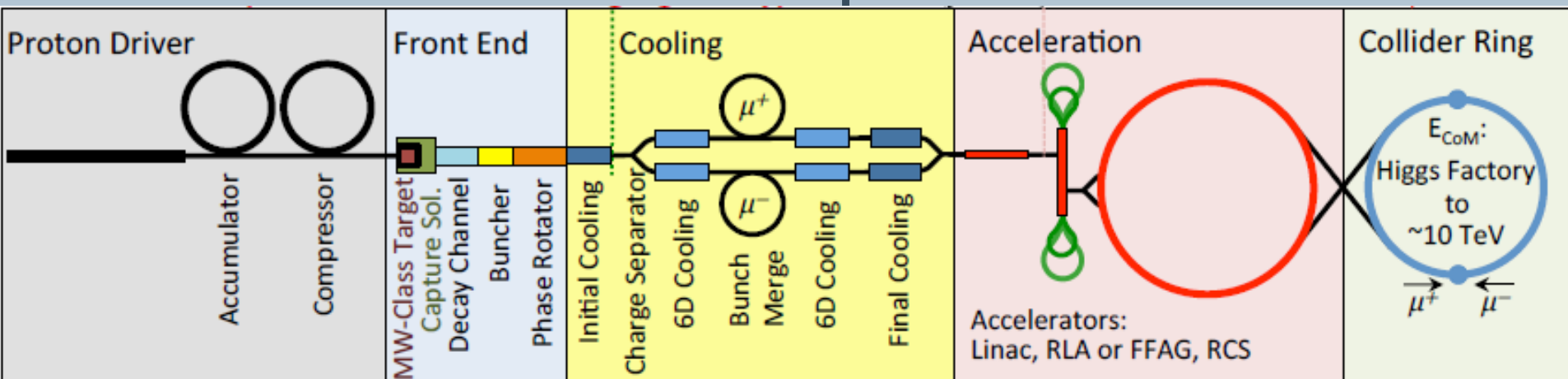
MAP Collaboration Meeting, Fermilab, Batavia, IL, USA

May 20, 2015

Outline

- Introduction to muon accelerators
- Challenges
- Cooling solution for a muon accelerator
 - Introduction to ionization cooling
 - Lattice design features
 - End-to-End simulation
 - Challenges & feasibility issues
- Summary

Muon Collider parameters

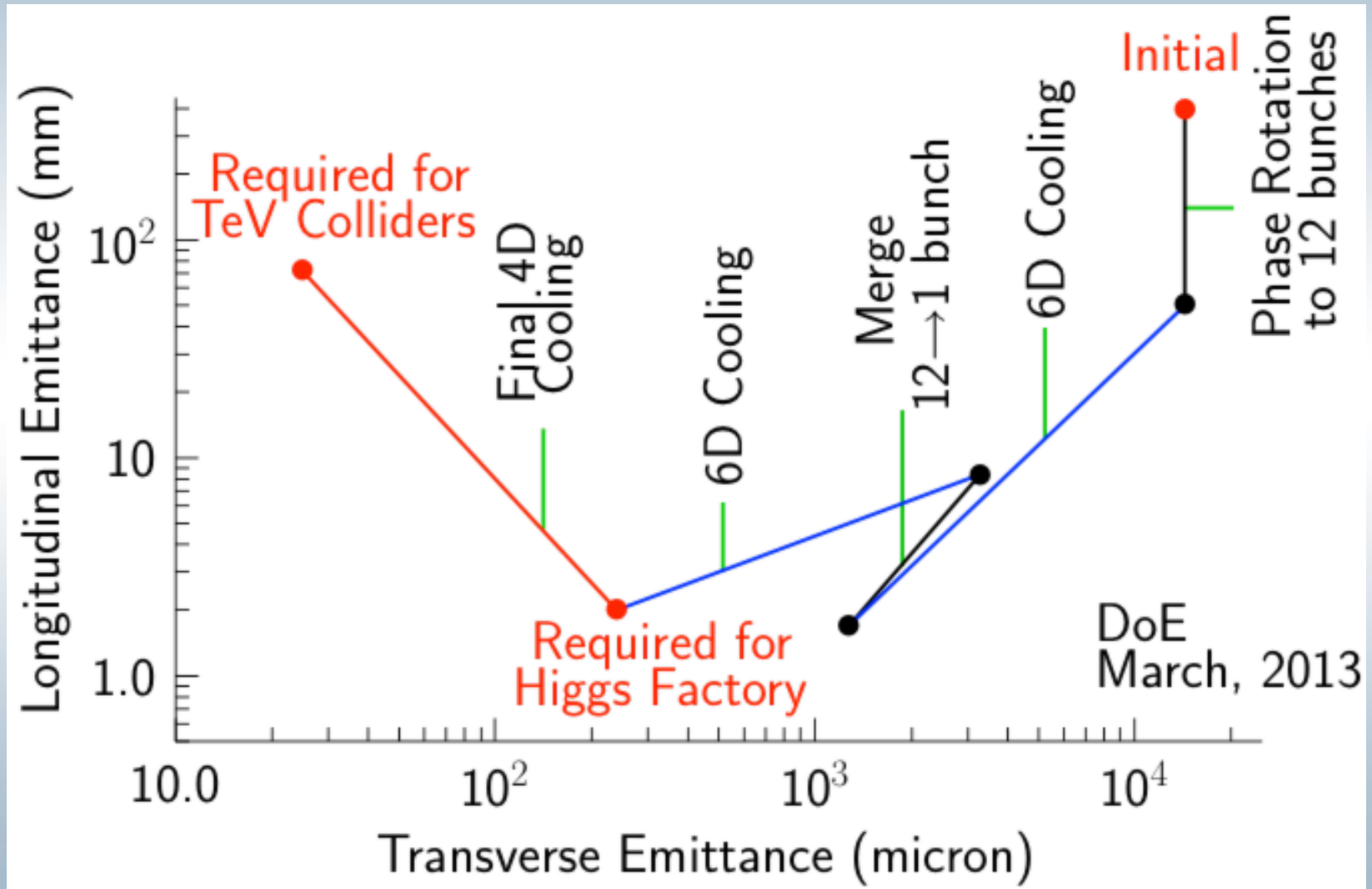


Muon Collider Parameters								
Parameter	Units	Higgs Factory		Top Threshold Options		Multi-TeV Baselines		Accounts for Site Radiation Mitigation
		Startup Operation	Production Operation	High Resolution	High Luminosity			
CoM Energy	TeV	0.126	0.126	0.35	0.35	1.5	3.0	6.0
Avg. Luminosity	$10^{34} \text{cm}^{-2} \text{s}^{-1}$	0.0017	0.008	0.07	0.6	1.25	4.4	12
Beam Energy Spread	%	0.003	0.004	0.01	0.1	0.1	0.1	0.1
Higgs* or Top* Production/ 10^7sec		3,500*	13,500*	7,000*	60,000*	37,500*	200,000*	820,000*
Circumference	km	0.3	0.3	0.7	0.7	2.5	4.5	6
No. of IPs		1	1	1	1	2	2	2
Repetition Rate	Hz	30	15	15	15	15	12	6
β^*	cm	3.3	1.7	1.5	0.5	1 (0.5-2)	0.5 (0.3-3)	0.25
No. muons/bunch	10^{12}	2	4	4	3	2	2	2
No. bunches/beam		1	1	1	1	1	1	1
Norm. Trans. Emittance, ϵ_{TN}	$\pi \text{ mm-rad}$	0.4	0.2	0.2	0.05	0.025	0.025	0.025
Norm. Long. Emittance, ϵ_{LN}	$\pi \text{ mm-rad}$	1	1.5	1.5	10	70	70	70
Bunch Length, σ_s	cm	5.6	6.3	0.9	0.5	1	0.5	0.2
Proton Driver Power	MW	4 [#]	4	4	4	4	4	1.6

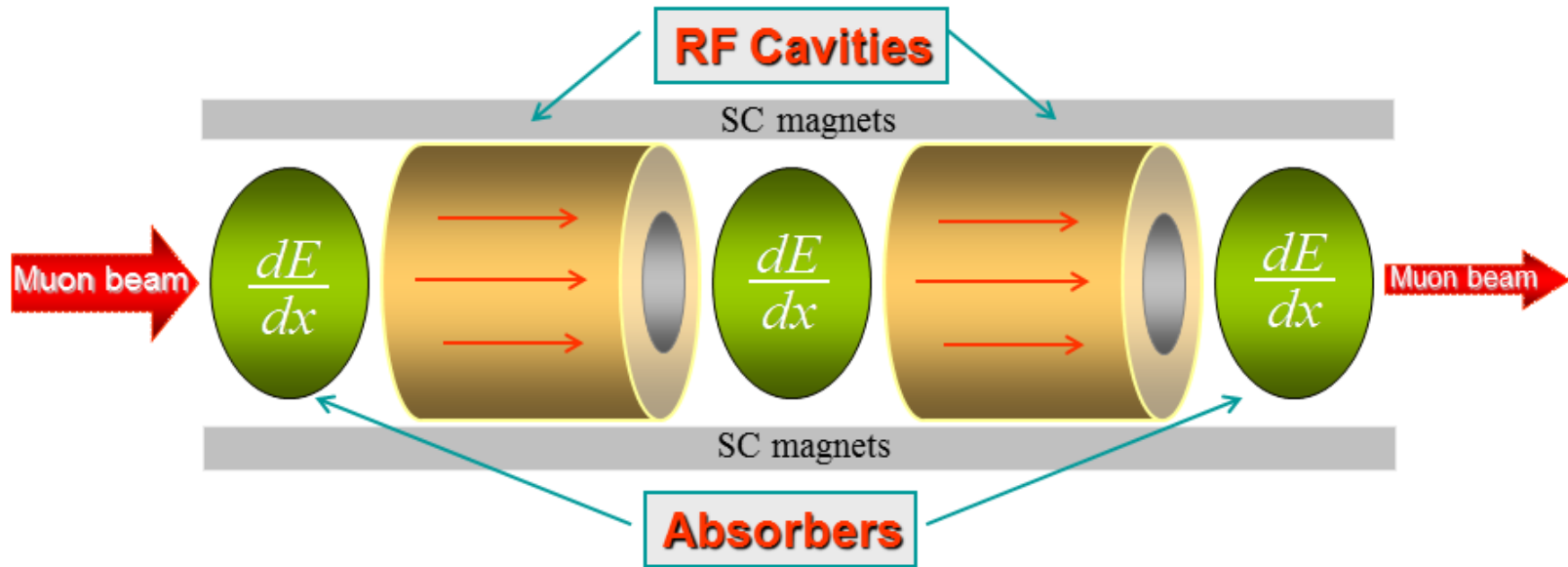
Dealing with muons: Challenges

- Muon beams are born as tertiary beams
 - Protons $\rightarrow$ pions $\rightarrow$ muons
 - Challenge: How to capture & transport muons efficiently?
- Muons are born within a large phase-space
 - To obtain luminosities $O(10^{34}) \text{ cm}^{-2}\text{s}^{-1}$ need to reduce the initial phase-space by 6 orders of magnitude
 - Challenge: How to cool the muon beam? This talk
- Muons decay fast ($\sim 2 \mu\text{s}$ at rest)
 - At the same time, deal with high momentum protons, electrons...
 - Challenge: How to accelerate fast the beam? Collider ring design?

The cooling challenge

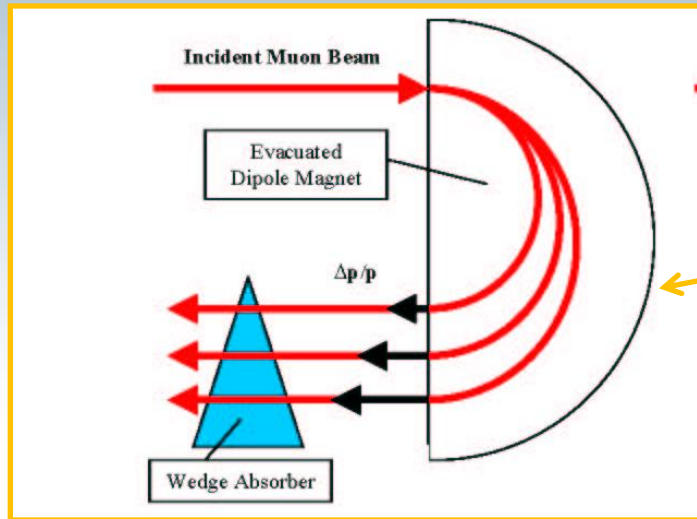


Ionization cooling concept



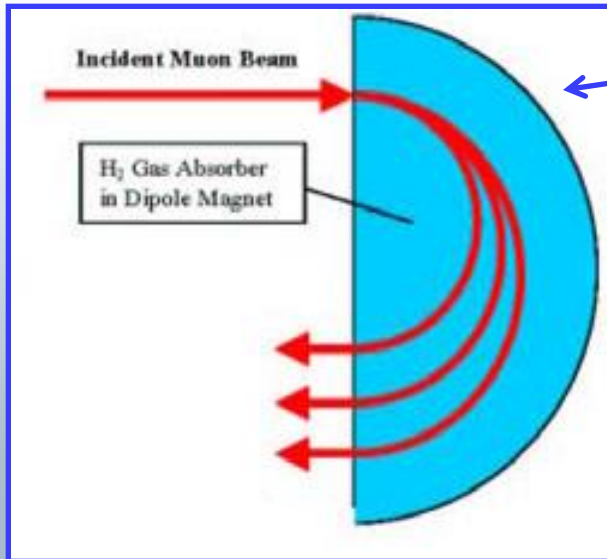
- Given the short muon lifetime, ionization cooling is the only practical method that can be realized
- Energy loss in discrete absorbers
- rf cavities to compensate for lost longitudinal energy
- Multi-tesla magnetic field to confine muon beams

Emittance exchange for 6D cooling



Concept 1: Generate dispersion and cool via emittance exchange in a wedge absorber

Concept 2: Energy loss dependence on path length in a continuous absorber



- Two concepts, same principle
- Dispersion is introduced to spatially separate muons of different momenta
- This study focuses on channels with discrete absorbers only!

Ionization cooling theory

- Transverse cooling:

$$\frac{d\varepsilon_{\perp}}{ds} = -\frac{\varepsilon_{\perp}}{\beta^2 E} \frac{dE}{ds} + \frac{\beta_{\perp}(13.6 \text{ MeV})^2}{2\beta^3 E_{\mu} X_0 E}$$

Cooling term

scattering term

- Minimum emittance:

$$\varepsilon_{\perp}^{eq} = \left(\frac{dE}{ds} \right)^{-1} \frac{\beta_{\perp}(13.6 \text{ MeV})^2}{2\beta E_{\mu} X_0}$$

- Emittance evolution:

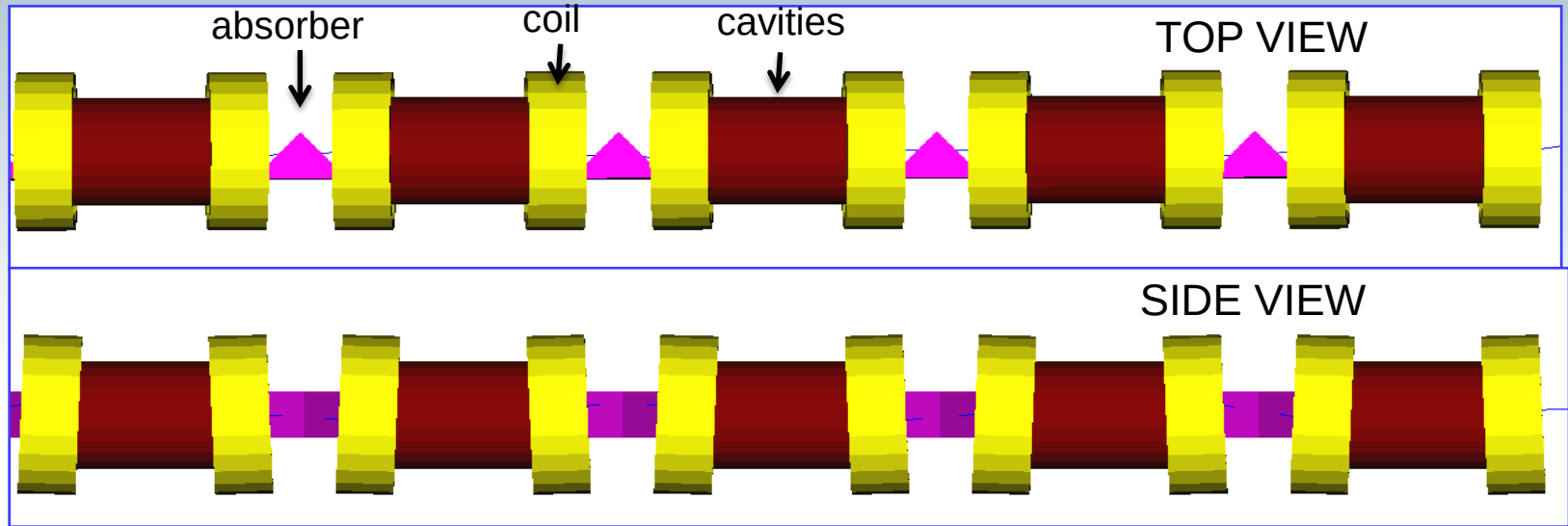
$$\varepsilon_{\perp}(s) = \varepsilon_{\perp}^{eq} + (\varepsilon_{\perp}^i - \varepsilon_{\perp}^{eq}) \exp(-s/s_{\perp}^c)$$

$$s_{\perp}^c = \beta^2 E / (dE/ds)$$

- Emittance can be controlled by:

- Material: Product of radiation length X_0 and energy loss rate dE/ds large. Best choices are LiH & H
- Magnet strength: Transverse beta function β_T must be small. Thus, we progressively taper the magnetic field towards higher values

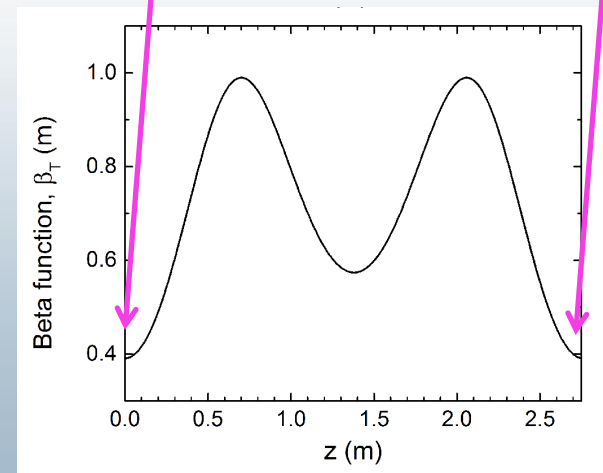
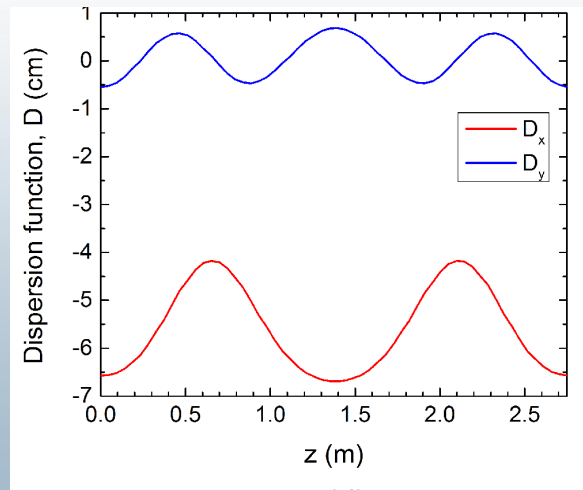
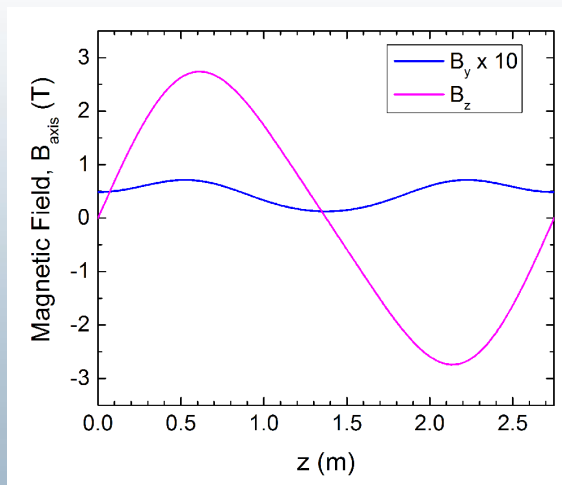
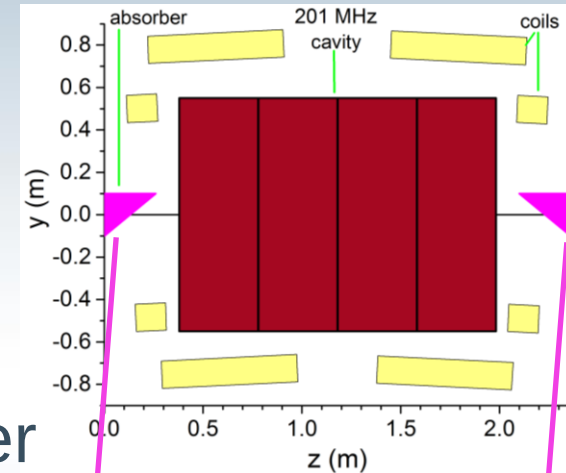
Cooling channel solution



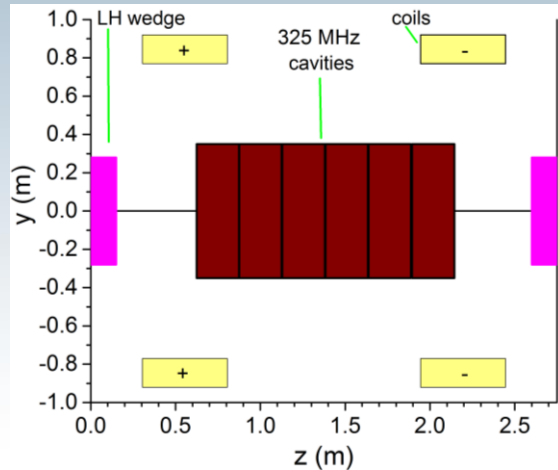
- We considered a rectilinear cooling scheme for a MC. Idea first proposed by V. Balbekov (Fermilab).
- Its simple geometry avoids several engineering challenges of previously considered schemes (rings or helix)

Cooling channel: How it works

- Coils are slightly tilted to generate a B_y field component
- This leads to dispersion, primarily in x.
- 6D cooling on wedge absorber
- Better, if beta is minimum at the absorber

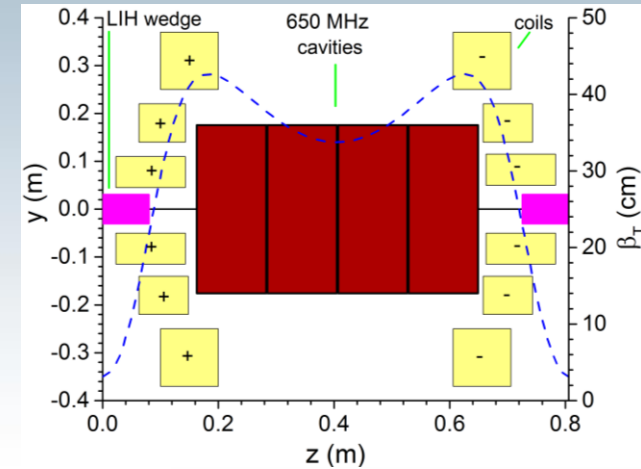


Tapered lattice design



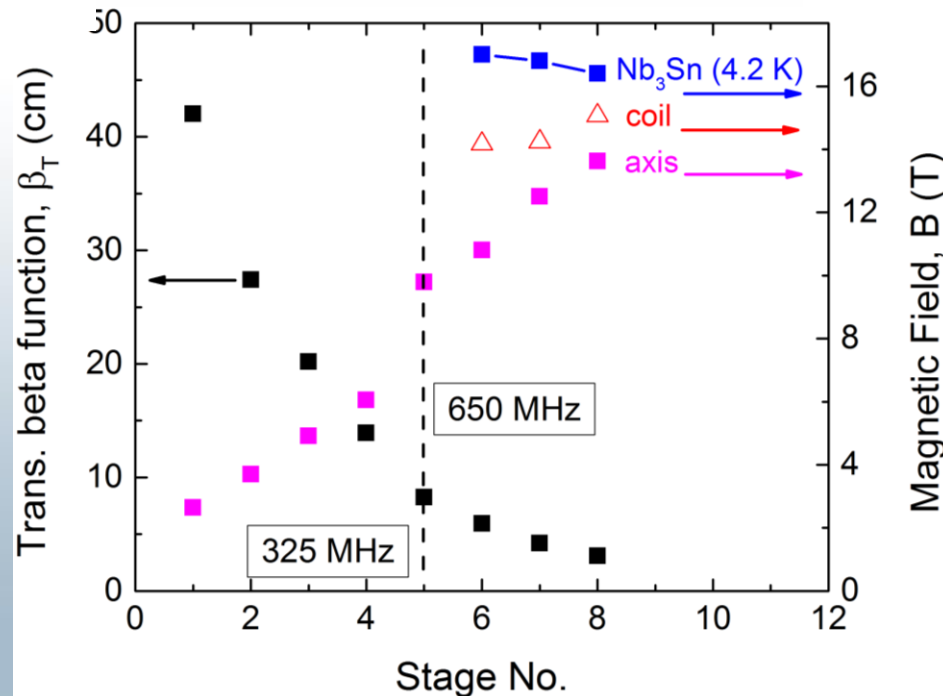
EARLY STAGE COOLING

- 275 cm long
- Coils far
- 325 MHz
- Axial B $\sim$ 3 T
- Beta $\sim$ 40 cm



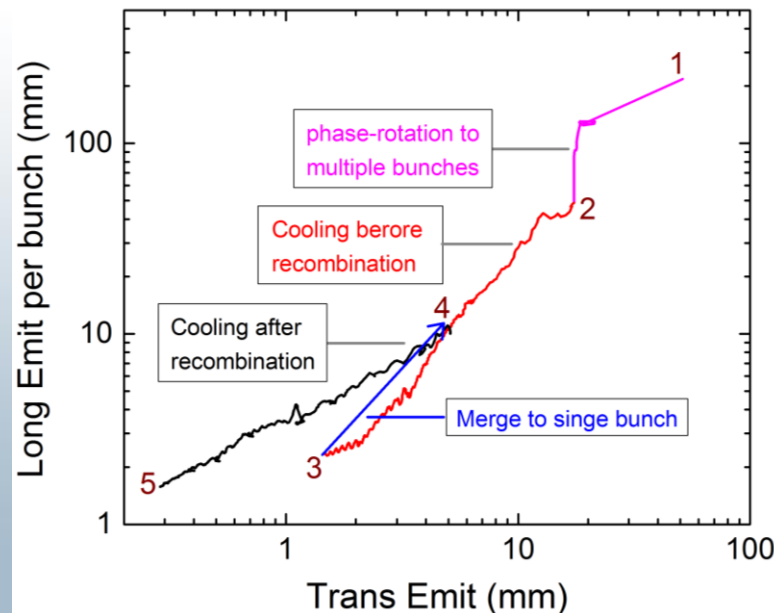
LATE STAGE COOLING

- 80 cm long
- Coils near axis
- 650 MHz
- Axial B $\sim$ 12 T
- Beta $\sim$ 3 cm



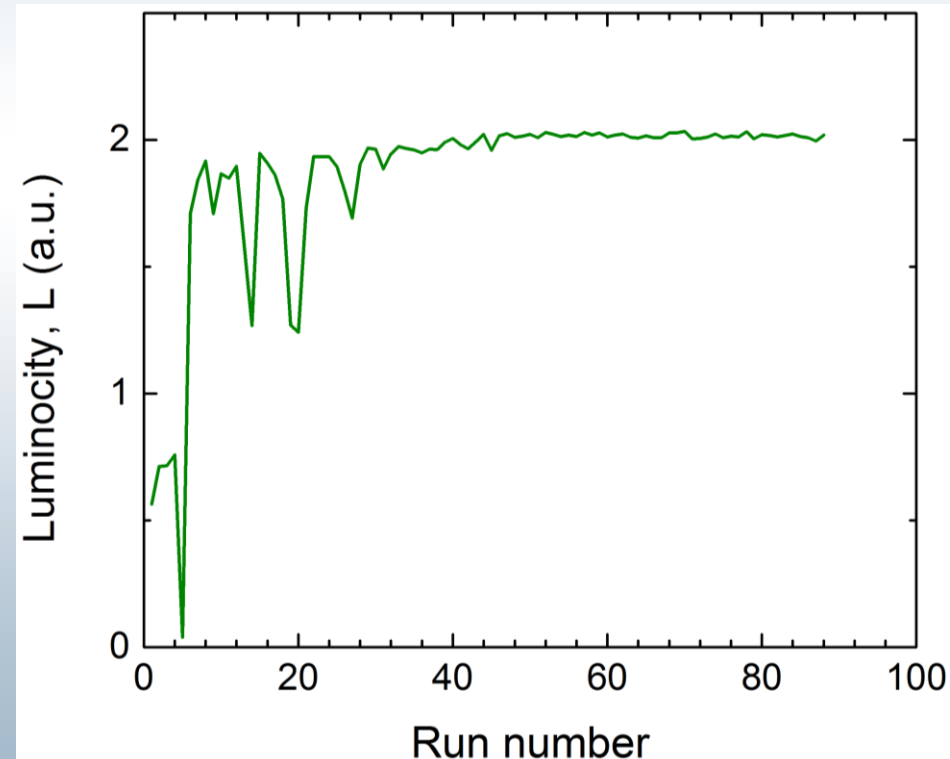
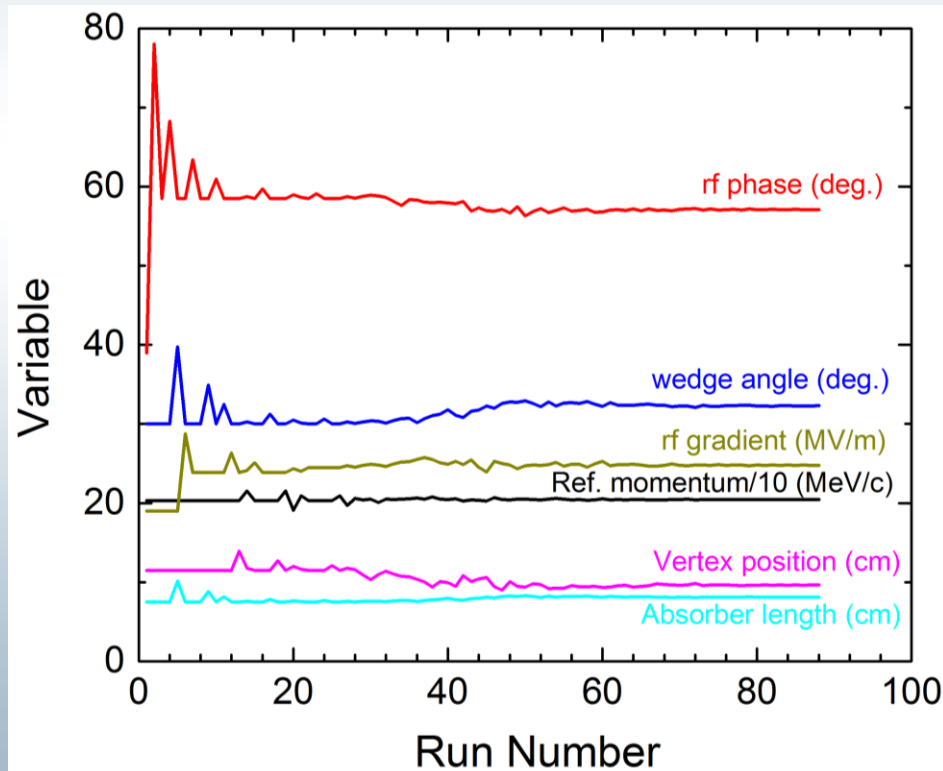
Numerical study with ICOOL

- End-to-end simulation starting from the muon beam birth place (point 1)
- Notable reduction of 6D-emittance a factor of 10^5
- Desired emittances for Higgs factory delivered!



Multivariable optimization

- 6 beam & lattice parameters were scanned with the goal to maximize the luminosity using the Nelder-Mead algorithm

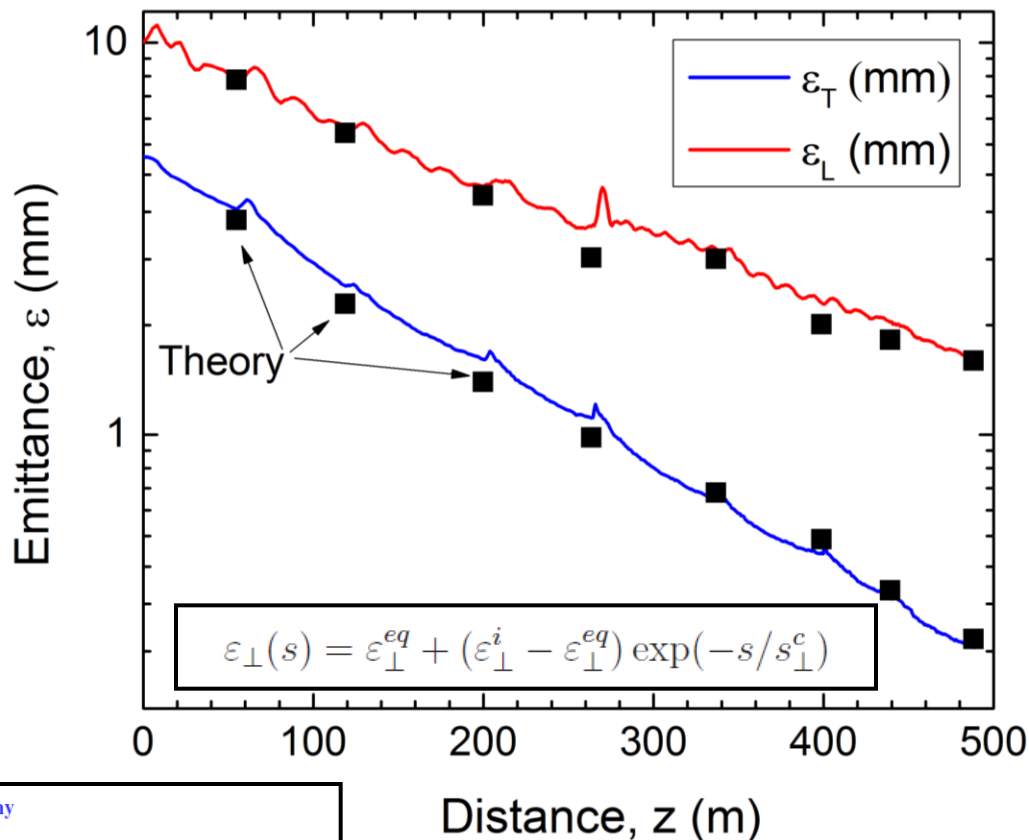


Theory and simulation

- Found good agreement between theory and simulation

$$\varepsilon_{\perp}^{eq} = \left(\frac{dE}{ds} \right)^{-1} \frac{\beta_{\perp} (13.6 \text{ MeV})^2}{2\beta E_{\mu} X_0}$$

$$s_{\perp}^c = \beta^2 E / (dE/ds)$$



TUPME021

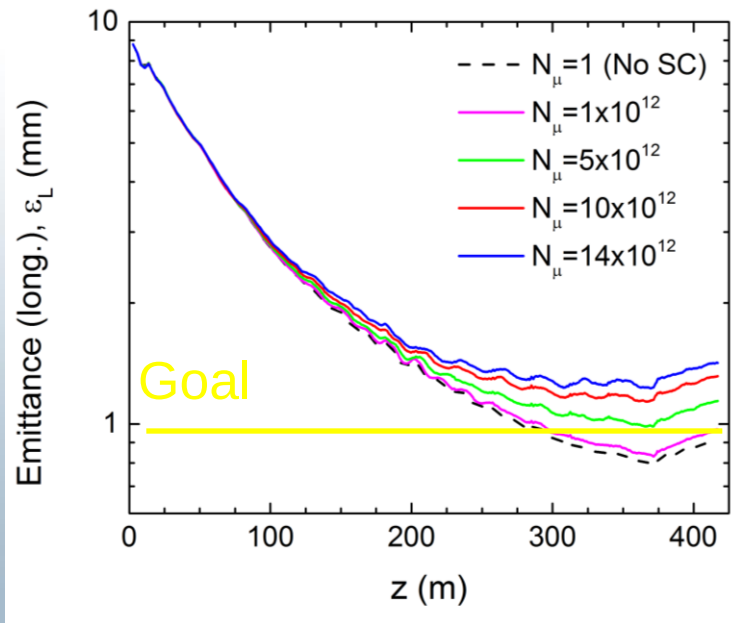
Proceedings of IPAC2014, Dresden, Germany

THEORETICAL FRAMEWORK TO PREDICT EFFICIENCY OF IONIZATION COOLING LATTICES*

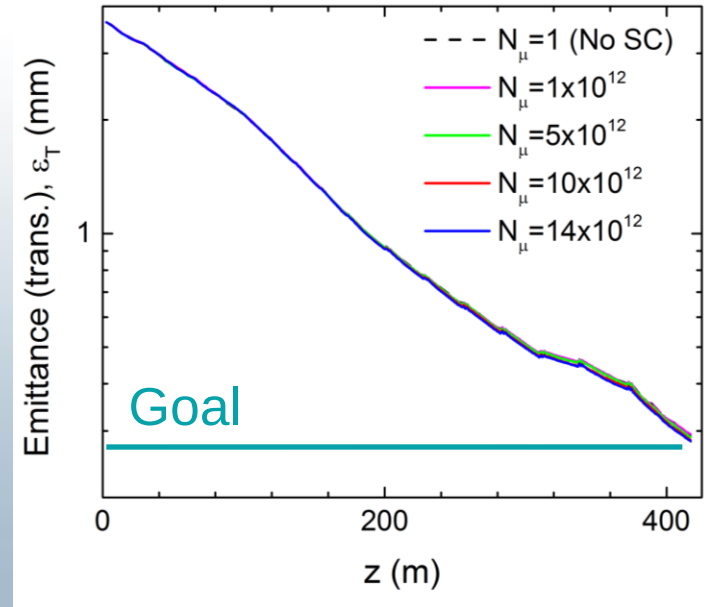
Diktys Stratakis[#], Brookhaven National Laboratory, Upton, NY, USA
David Neuffer, Fermi National Accelerator Laboratory, Batavia, IL, USA

Influence of space-charge

- At the end of the cooling, the bunch is 2 cm long, has 5×10^{12} muons with a peak current of 5 kA at 200 MeV/c!
- We examined the influence of SC fields on cooling.
- SC causes particle loss and longitudinal emittance growth



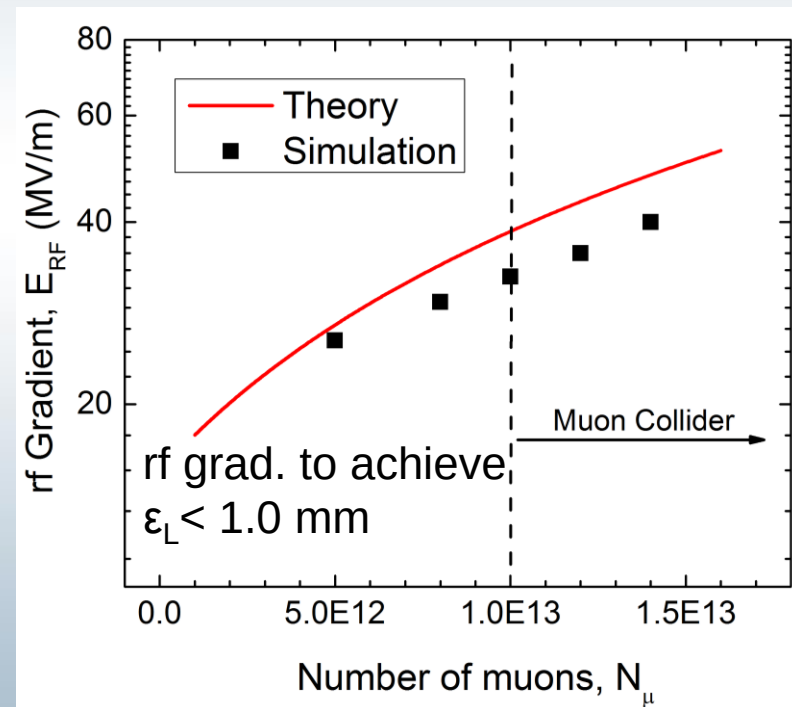
WARP
Simulation



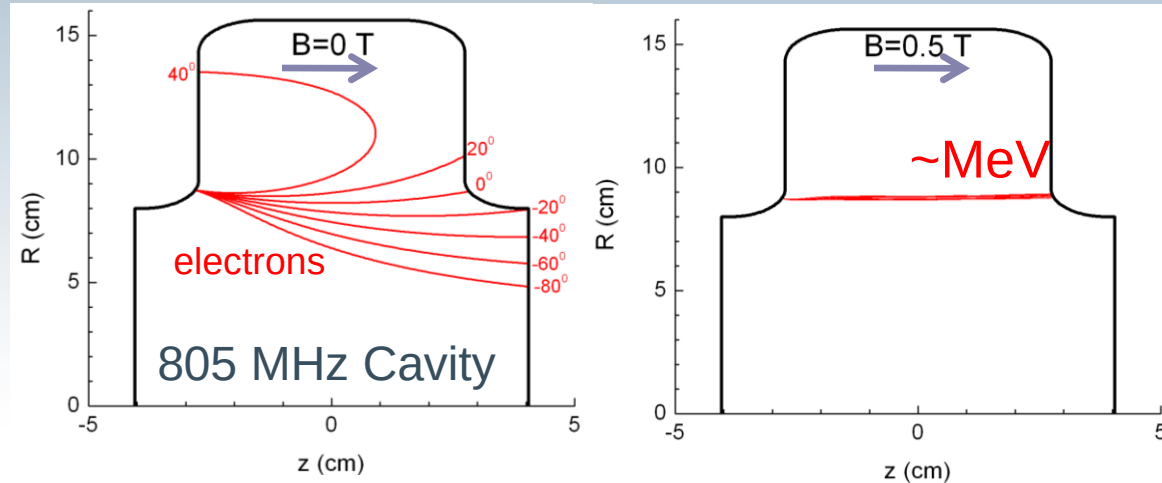
Space-charge compensation (2)

- For a Muon Collider to obtain a longitudinal emittance < 1.0 mm the rf gradient of a 805 MHz cavity needs to surpass 32.5 MV/m
- Avoid longitudinal cooling to < 1.3 mm
- Required compensation gradient from theory:

$$E_{rf,c} = \frac{eN_0 g_0 c}{4\pi\epsilon_0 \sqrt{2\pi\gamma^2 \sigma_z^3} (2\pi f n \cos\phi)^2}$$



rf operation in B-fields



- Numerical simulations predict that the copper surfaces of a rf cavity may be damaged when $B > 1\text{ T}$

Contents lists available at ScienceDirect

Nuclear Instruments and Methods in Physics Research A

journal homepage: www.elsevier.com/locate/nima

Effects of external magnetic fields on the operation of high-gradient accelerating structures

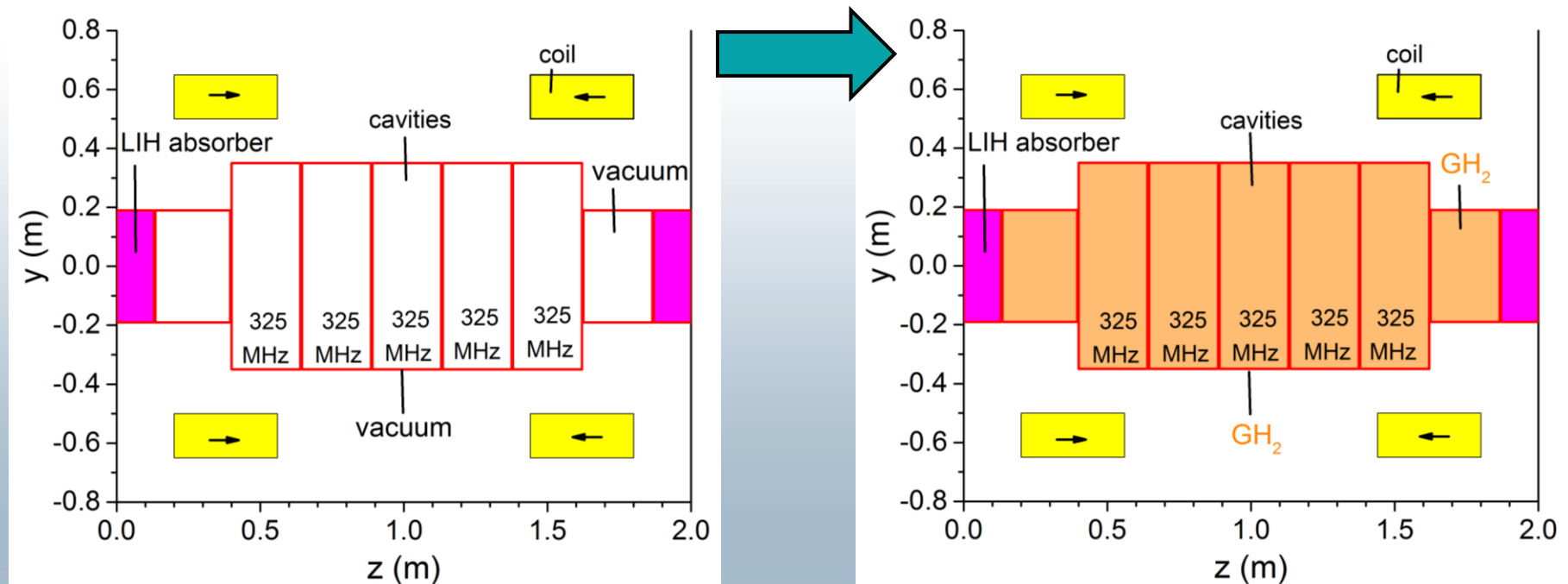
Diktys Stratakis*, Juan C. Gallardo, Robert B. Palmer

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NUCLEAR INSTRUMENTS & METHODS IN PHYSICS RESEARCH

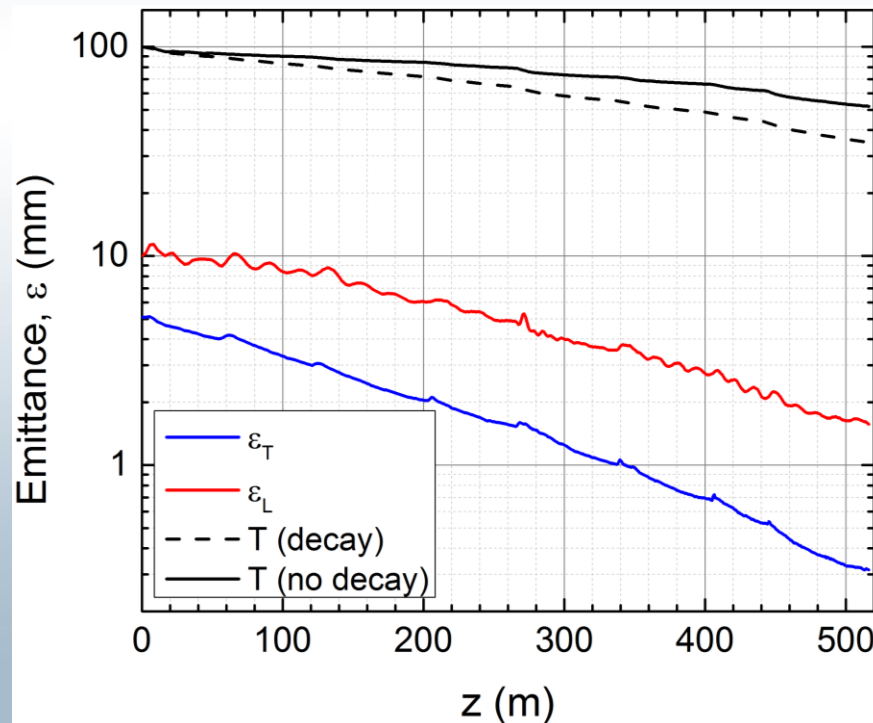
Hybrid solution

- Key Idea: Utilize gas filled cavities in a rectilinear channel
- Majority of cooling will be done in LiH and use gas only to protect the cavity from the high-field. Similar idea was used in the past for 4D cooling (Gallardo & Zisman, Nufact09)



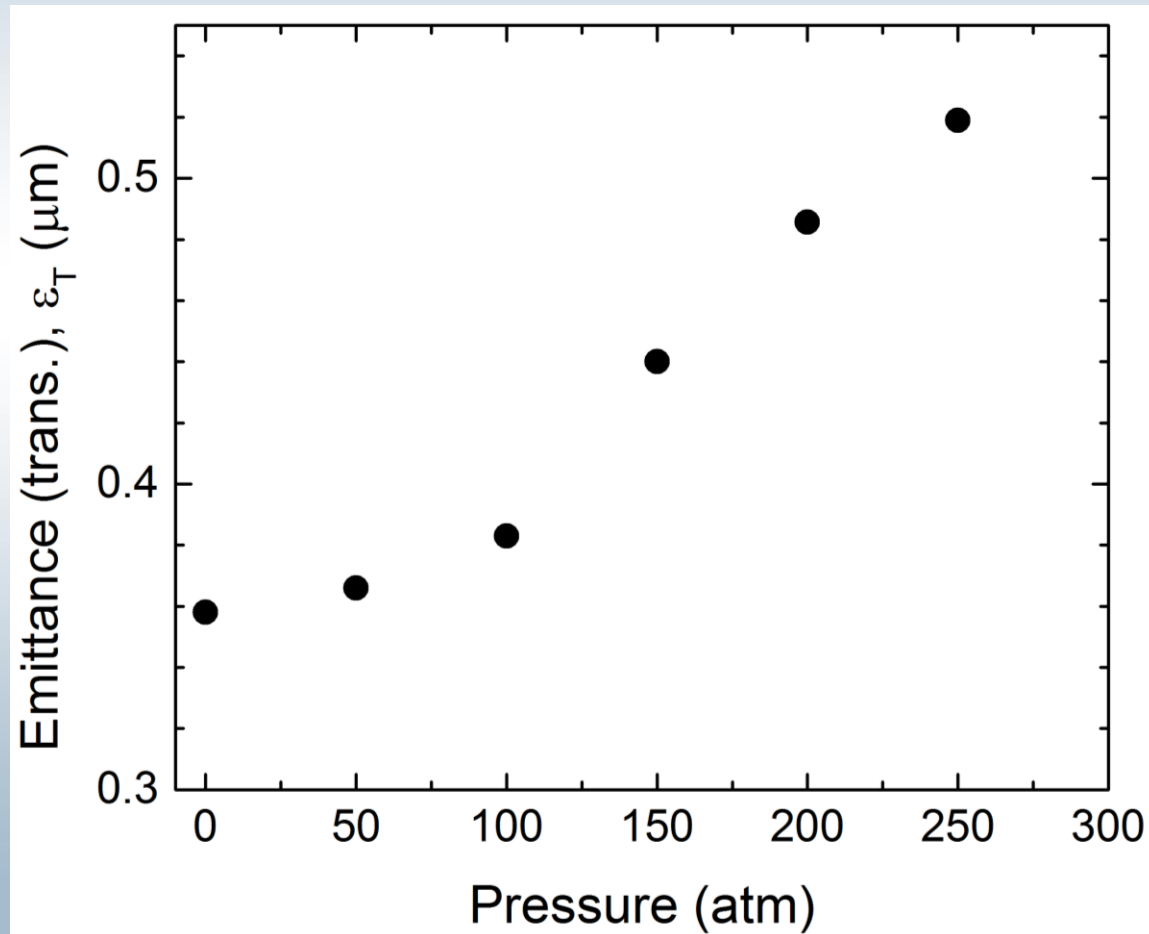
Lattice performance

- Essentially, the same performance as the conventional channel with vacuum cavities
- BUT there remains considerable work to do before a hybrid channel can be considered a validated cooling channel option.

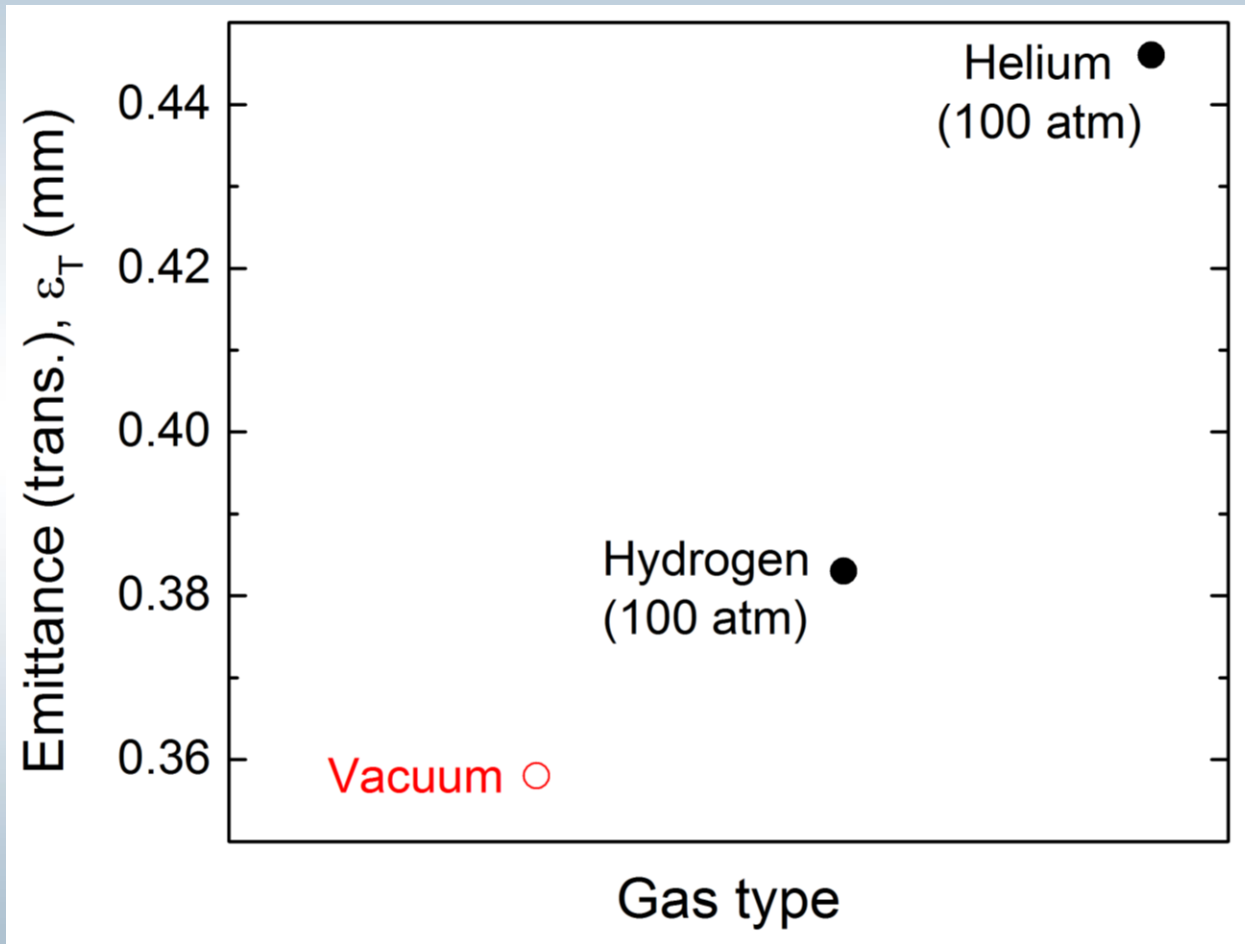


Sensitivity to pressure

- Final transverse emittance is correlated to gas pressure



Sensitivity to gas type



Summary

- We have presented a conceptual design of a rectilinear channel that in view its simple geometry may offer several technological advantages (compared to a helix or a ring)
- Showed reduction of 6D by at least 5 orders of magnitude.
- The influence of space-charge fields on the cooling process was thoroughly examined
- Nb₃Sn magnets can be used for late stage cooling
- A hybrid solution with gas filled cavities was presented