

**Repercussions of a geometric theory of  
flavor for future directions  
in High Energy Physics**

**Amarjit Soni, HET, BNL**

**Snowmass @ Univ of Minnesota 7/31/13**

- **Talk is primarily based on arXiv: 1303.5056**  
**see also Davoudiasl, McElmurry and A.S.**  
**arXiv:1206.4062**
- **See also talks at**
- **FPCP 2012, Hefei China, May 2012**
- **ICHEP 2012 Melbourne (July 2012)**
- **Solvay workshop on “Facing the Scalar Sector” , Brussels, May 29-31, 2013**

# Main point

- Expt results on Higgs –SM like as well as flavor suggest scale of new physics is not  $\sim 1$  TeV

So what is it ....?

What are its consequences?

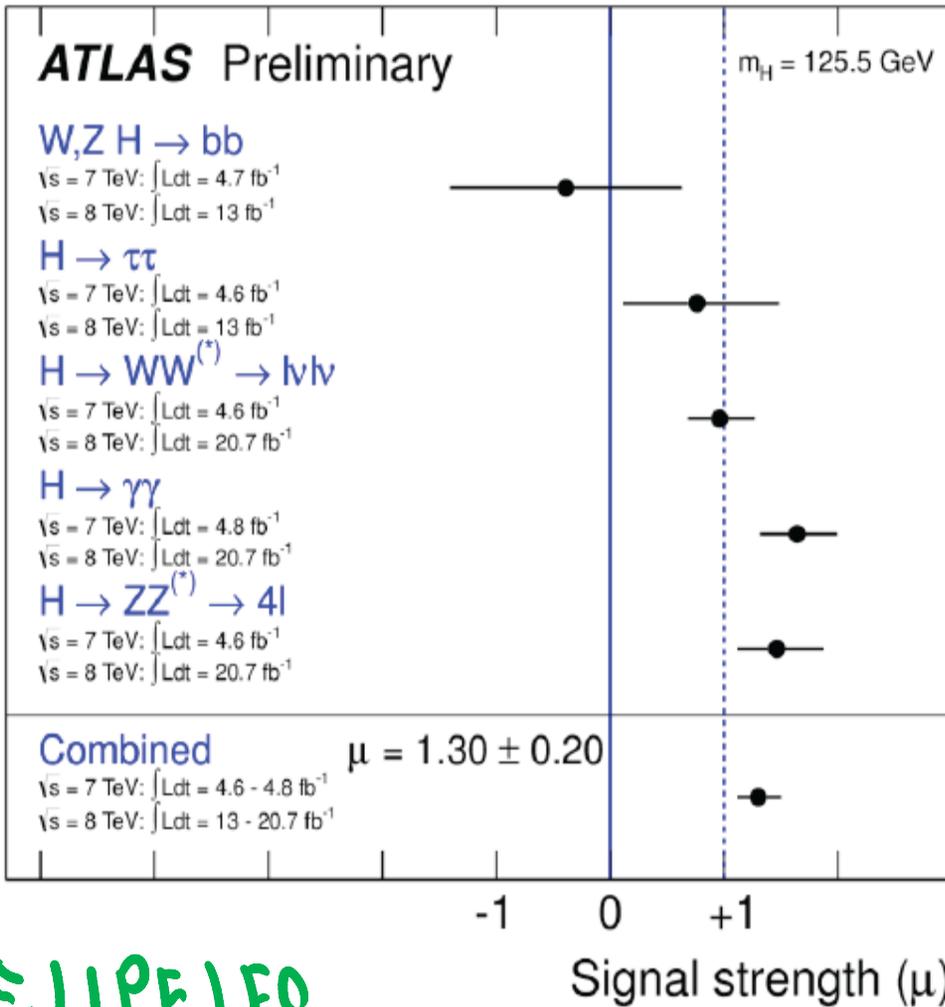
# Outline

- Works but .....
- Scale of NP=> future directions
- From the vantage point of a geometric theory of flavor scale was not expected to be less than around 10 TeV \*
- summary

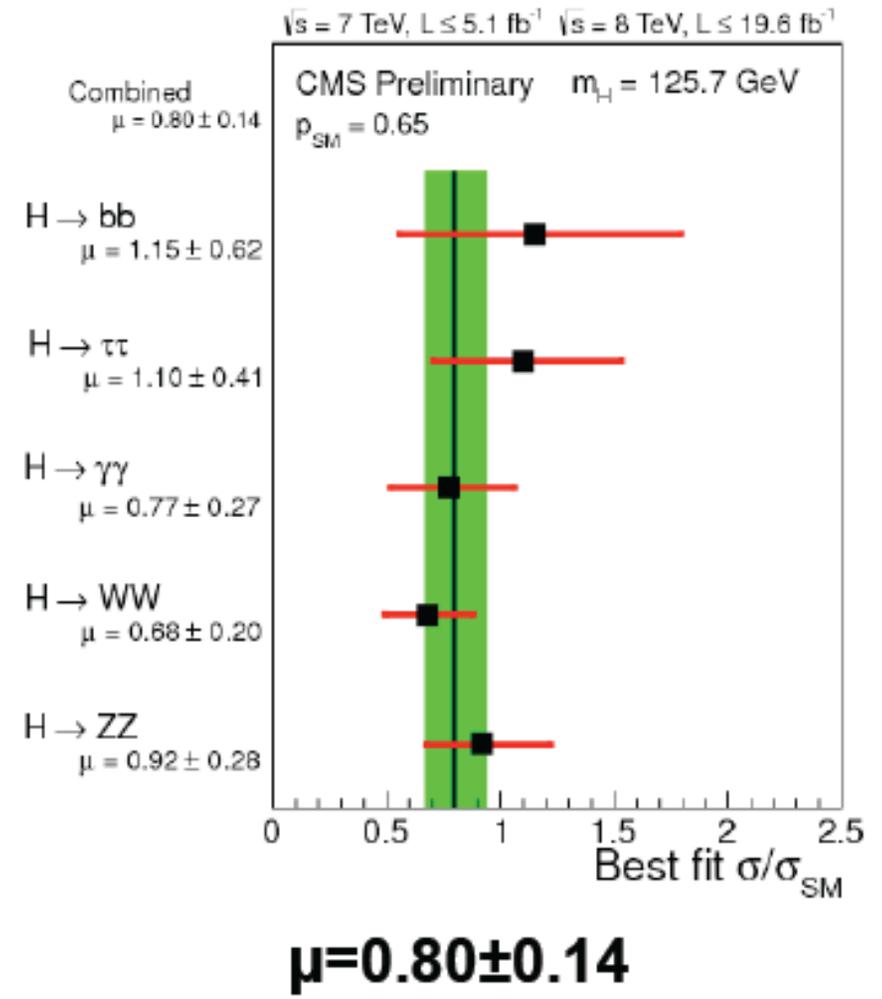
\* Flavor is not a footprint and, in particular flavor-alignment is a serious issue

# **FITS LIKE A GLOVE!**

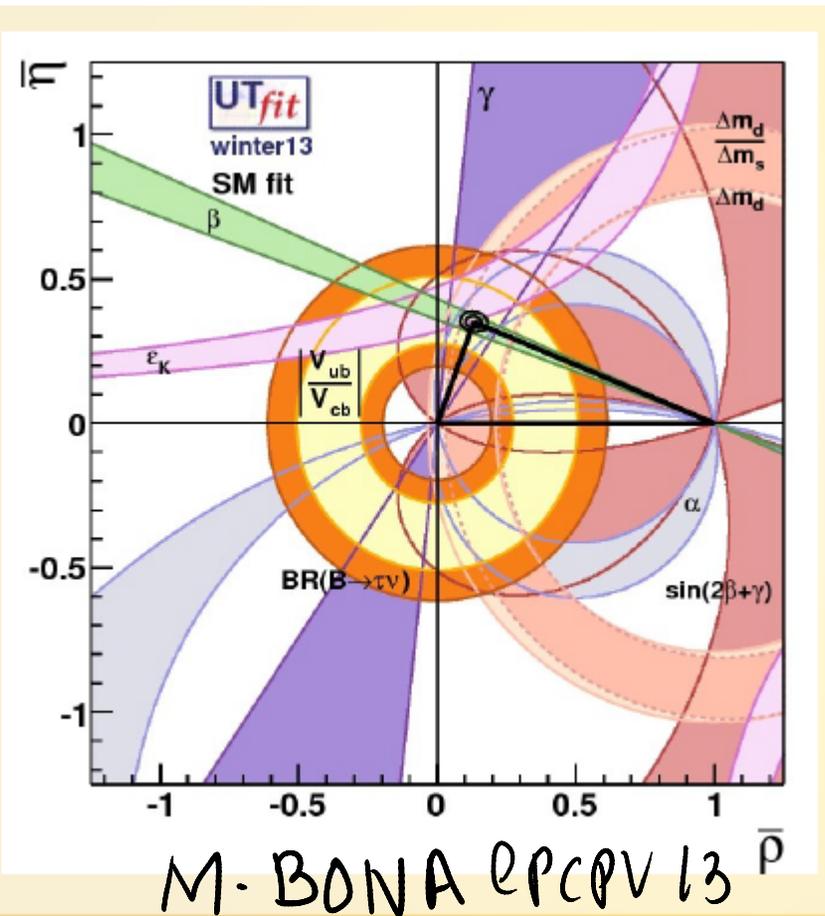
## **[OR DOES IT?]**



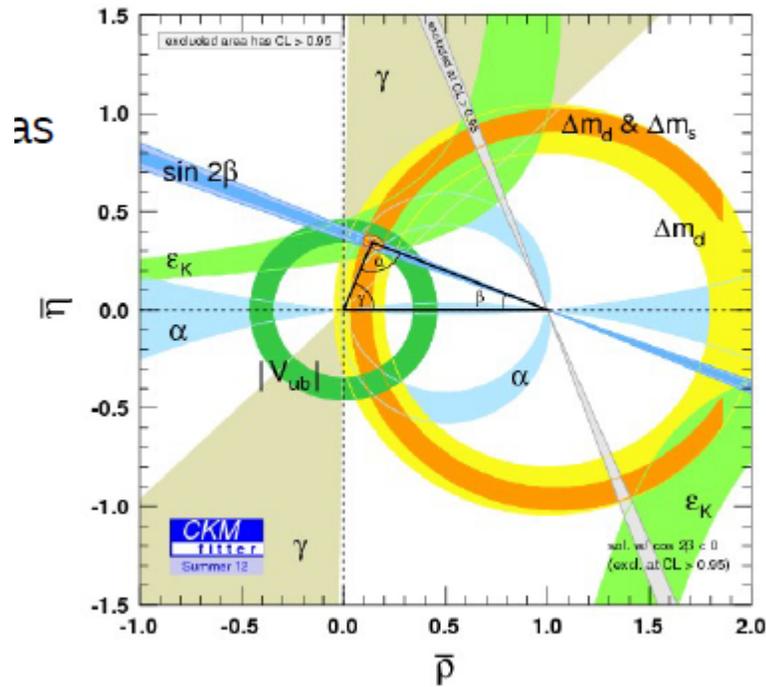
E. LIPELES  
 PBF 2013



S. BOSE PBF 2013



<http://ckmfitter.in2p3.fr>  
see also <http://www.utfit.org>



T. GERSHON  
@BF2013

SM-CKM paradigm works rather well.  
No glaring discrepancy  
OTOH tests only  $\sim 10-15\%$  accuracy

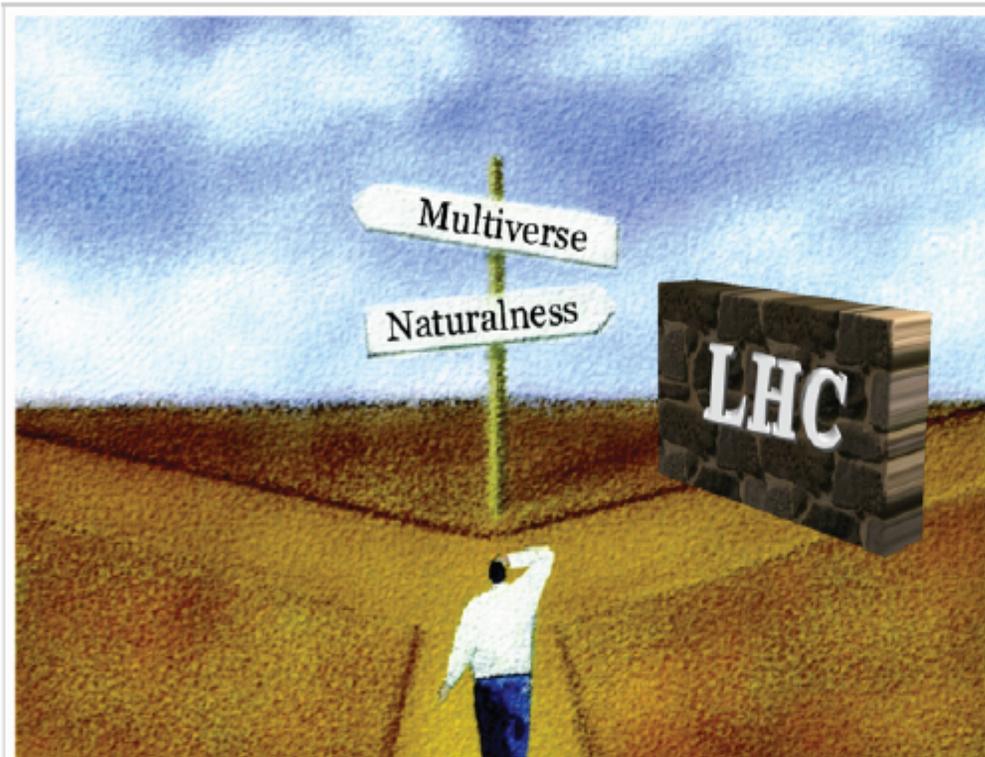
## Is Nature Unnatural?

Decades of confounding experiments have physicists considering a startling possibility: The universe might not make sense.

by: [Natalie Wolchover](#)

May 24, 2013

[email](#) [print](#)



Is the universe natural or do we live in an atypical bubble in a multiverse? Recent results at the Large Hadron Collider have forced many physicists to confront the latter possibility. (Illustration: Giovanni Villadoro)

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Science Lives

## A lesson from history (I)

"A special search at Dubna was carried out by E. Okonov and his group. They did not find a single  $K_L \rightarrow \pi^+ \pi^-$  event among **600 decays** into charged particles [12] (Anikira et al., JETP 1962). At that stage the search was terminated by the administration of the Lab. The group was unlucky."

-**Lev Okun**, "The Vacuum as Seen from Moscow"

1964:  $BF = 2 \times 10^{-3}$

A failure of imagination ? Lack of patience ?

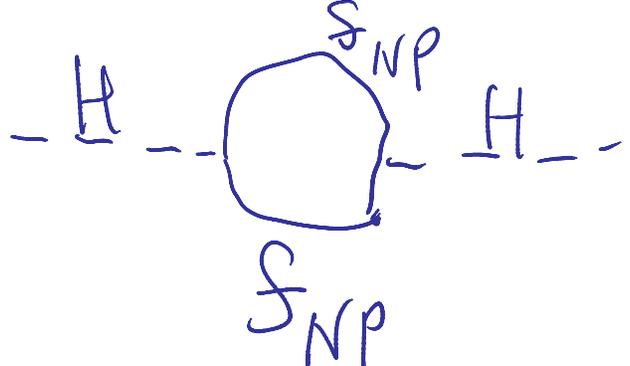
Christenson,  
Cronin, FITCH  
& TURLAY  
BNL 1964

DRAWING STRONG CONCLUSIONS  
BASED ON 20% tests is  
too risky.

**SHOULD WE BE SHOCKED TO FIND THAT  
THE SCALE OF NEW PHYSICS IS NOT  $\sim 1$   
TEV & APPEARS TO BE BIGGER?**

# Outstanding Th.puzzles of our times

- Hierarchy puzzle**



A Feynman diagram showing a Higgs boson (H) interacting with a loop of new particles (NP). The loop is represented by a circle with two vertices, each labeled  $S_{NP}$ . Dashed lines represent the Higgs boson entering and exiting the loop.

$$-H - \text{loop} - H \sim \frac{g_{NP}^2}{16\pi^2} \Lambda_{NP}^2 \Rightarrow \Lambda_{NP} \lesssim \text{TeV}$$

to avoid fine tuning  $m_H$

- Flavor puzzle**

$\Delta flavor = 2$  e.g.



A Feynman diagram showing a flavor-changing neutral current process. A quark line starts as a strange quark (s) and ends as a down quark (d). A wavy line labeled  $G_{NP}$  connects two vertices. At the first vertex, a down quark (d) and a strange quark (s) meet. At the second vertex, a down quark (d) and a strange quark (s) meet.

$$\sim \frac{g_{NP}^2}{\Lambda_{NP}^2} \Rightarrow \Lambda_{NP} \gtrsim 10^3 \text{ TeV}$$

to avoid constraint from  $\Delta C M_K$

# The Randall-Sundrum (RS) idea

**Island Universes in Warped Space-Time**

According to string theory, our universe might consist of a three-dimensional "brane," embedded in higher dimensions. In the model developed by Lisa Randall and Raman Sundrum, gravity is much weaker on our brane than on another brane, separated from us by a fifth dimension. (Time is the unseen fourth dimension.)

**GRAVITY BRANE**  
(where gravity is concentrated)

**Fifth dimension**  
Space is warped by energy throughout five-dimensional space-time. As a result, gravity is much weaker on our brane.

**Gravitons,**  
which transmit gravity, are closed strings, which are not confined to either brane.

**Warped space-time**  
Because space-time is warped, things are exponentially bigger and lighter closer to our brane.

**BRANE**  
(our universe)

The ends of **open strings**, whose oscillations are particles and forces other than gravity, are stuck to our brane.

(Wikipedia)

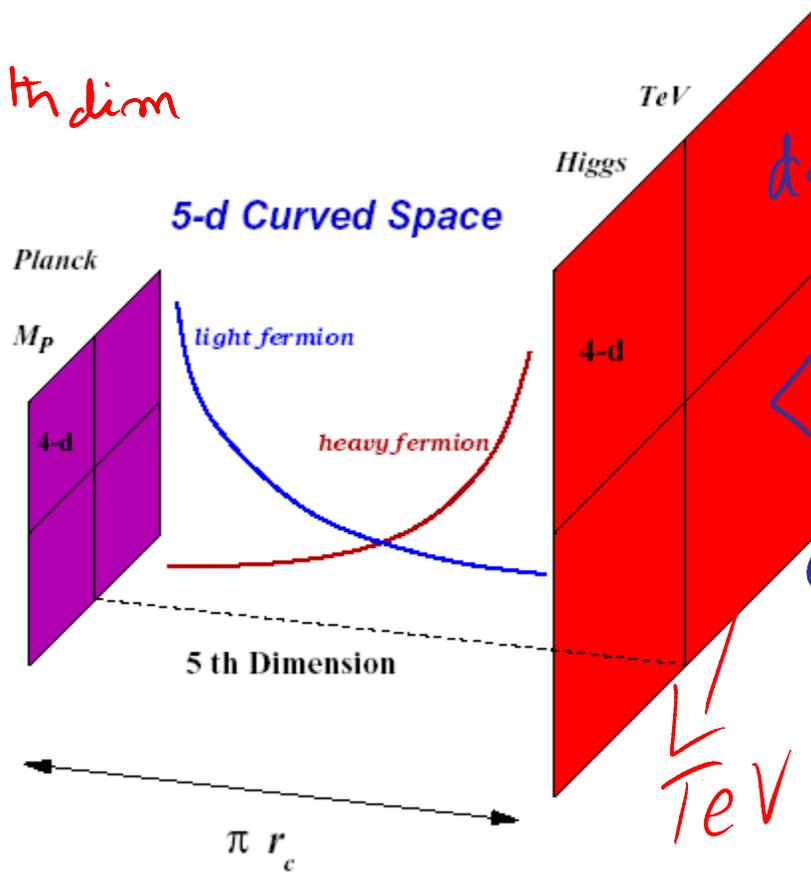
[Stolen from Newbert]

# **INSIGHTS FROM A GEOMETRIC THEORY OF FLAVOR**

# RANDALL+SUNDRUM '99

[FIG BY  
H DAVOUDI/ASL]

Points along 5<sup>th</sup> dim  
correspond to  
diff. eff.  
4d scale!



$$ds^2 = e^{-2\sigma} \eta_{\mu\nu} dx^\mu dx^\nu - r_c^2 d\psi^2$$

$$\langle H_4 \rangle = e^{-6\sigma} \langle H_5 \rangle$$

$$G = \frac{1}{2} r_c \pi \sim \frac{1}{12}$$

$\rightarrow$   
 $M_P$

Figure 1: Warped geometry with flavor from fermion localization. The Higgs field resides on the TeV-brane. The size of the extra dimension is  $\pi r_c \sim M_P^{-1}$ .

## Simultaneous resolution to hierarchy and flavor puzzles

*Fermion “geography” (localization) naturally explains:*

Grossman&Neubert; Gherghetta&Pomarol; Davoudiasl, Hewett & Rizzo

- Why they are light (or heavy)
  - FCNC for light quarks are severely suppressed automatically
  - RS-GIM MECHANISM (Agashe, Perez, AS'04) flavor changing transitions though at the *tree level* (resulting from rotation from interaction to mass basis) are suppressed roughly to the same level as the loop in SM  $\Rightarrow$  CKM hierarchy
  - **O(1) CP ubiquitous;.....nedm, in fact ALL DIR-CP [ $\epsilon'/\epsilon$ ,  $\gamma$ ,  $\Delta ACP(B \Rightarrow K\pi)$ ,  $\Delta(\sin 2\beta)$ ;  $S[B \Rightarrow K^* \rho\gamma]$ ;  $\Delta ACP(D)$ ..] are an exceedingly important path to BSM-phase and new physics**
  - Most flavor violations are driven by the top
- > ENHANCED  $t \rightarrow cZ(h)$  ....A VERY IMPORTANT “GENERIC” PREDICTION..Agashe, Perez, AS'06

$$\Delta m_K : 10^3 \text{ TeV} \Rightarrow R_{SFL} \sim 10 \text{ TeV}$$

EXTENSIVE RECENT STUDIES by BURAS et al and NEUBERT et al

## Localization parameters of the 3-families of quarks

$$\begin{array}{lll} c_{Q_1} = -0.579, & c_{Q_2} = -0.517, & c_{Q_3} = -0.473 \\ c_{u_1} = -0.742, & c_{u_2} = -0.558, & c_{u_3} = +0.339 \\ c_{d_1} = -0.711, & c_{d_2} = -0.666, & c_{d_3} = -0.553 \end{array}$$

Table from  
M. Neubert  
@Moriond09

⇒ masses of the 6 quarks in RS!

# REMINISCENT OF CHENG-SHER ANZATS

$$f_{ij} = \lambda_{ij} \sqrt{m_i m_j} / v$$

# Minimal tuning ansatz & naturalness

- $RS_{Fl}$  is the scale at which all flavor constraints are satisfied....
- This scale seems to be set by Kaon mixings and is around 10 TeV. This is high enough that EWPC are automatically satisfied.
- While scale of new physics could be bigger than this, its likely that it is close to  $RS_{Fl}$  so as to minimize the degree of tuning needed....:

$$v^2 / [RS_{Fl}]^2 \sim 10^{-3} \gg \gg [v/M_{pl}]^2$$

Naturalness is not at stake at least not yet!

## B-Factory Signals for a Warped Extra Dimension

Kaustubh Agashe,<sup>1,\*</sup> Gilad Perez,<sup>2,†</sup> and Amarjit Soni<sup>3,‡</sup>

<sup>1</sup>*Department of Physics and Astronomy, Johns Hopkins University, Baltimore, Maryland 21218-2686, USA*

<sup>2</sup>*Theoretical Physics Group, Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA*

<sup>3</sup>*Brookhaven National Laboratory, Upton, New York 11973, USA*

(Received 23 June 2004; published 10 November 2004)

*m<sub>KK</sub> ~ 3 TeV*

We study predictions for  $B$  physics in a class of warped extra dimension models recently introduced, where few ( $\sim 3$ ) TeV Kaluza-Klein masses are consistent with electroweak data due to custodial symmetry. As in the standard model (SM), flavor violations arise due to the heavy top quark leading to striking signals: (i) New physics contributions to  $\Delta F = 2$  transitions are comparable to the SM, so the success of the SM unitarity triangle fit is a “coincidence.” Thus, clean extractions of unitarity angles are likely to be affected, in addition to  $O(1)$  deviation from the SM prediction in  $B_s$  mixing. (ii)  $O(1)$  deviation from various SM predictions for  $B \rightarrow X_s l^+ l^-$ . (iii) Large mixing-induced  $CP$  asymmetry in radiative  $B$  decays. Also, the neutron electric dipole moment is roughly 20 times larger than the current bound so that this framework has a “ $CP$  problem.”

TABLE I. Contrasting signals from RS1 with the SM.

	$\Delta m_{B_s}$	$S_{B_s \rightarrow \psi\phi}$	$S_{B_d \rightarrow \phi K_s}$	$Br[b \rightarrow sl^+l^-]$	$S_{B_{d,s} \rightarrow K^*, \phi\gamma}$	$S_{B_{d,s} \rightarrow \rho, K^*\gamma}$
RS1	$\Delta m_{B_s}^{\text{SM}}[1 + O(1)]$	$O(1)$	$\sin 2\beta \pm O(0.2)$	$Br^{\text{SM}}[1 + O(1)]$	$O(1)$	$O(1)$
SM	$\Delta m_{B_s}^{\text{SM}}$	$\lambda_c^2$	$\sin 2\beta$	$Br^{\text{SM}}$	$\frac{m_s}{m_b} (\sin 2\beta, \lambda_c^2)$	$\frac{m_d}{m_b} (\lambda_c^2, \sin 2\beta)$

# 3 TeV scale was an unfortunate oversight

- FC KK-glu exchanges give rise to LR currents
- LR currents cause enhanced Kaon mixings.....
- [Beall, Bander, AS, PRL '82]....
- C also M.Bona et al [UTFit] arXiv:0707.0636
- Agashe, Papucci, Perez and Pirjol, hep-ph/0509117
- Gedalia, Isidori and Perez, arXiv: 0905.3264

$$\Rightarrow m_{KK} \sim 10 \text{ TeV}$$

# EWPC/Little Hierarchy

- Imposing custodial symmetry, by enlarging  $SU_2 \Rightarrow SU_2 \times SU_2$  and by adding additional fermions,  $m_{KK} > \sim 3 \text{ TeV}$  [see Agashe, Delgado, May & Sundrum'03]
- Once  $RS_{FI} > 10 \text{ TeV}$  is imposed, EWPC likely automatically satisfied; there is now **“Little Hierarchy”** i.e. tuning  $O(10^{-3})$  is somewhat worse but the setup is more economical & simpler; perhaps the experimental indications to date are that this simplicity is preferred. However, not the last word yet.

# Higgs is SM-like =>

- Light SM-like Higgs strengthens case for  $m_{KK} > \sim 10$  TeV in warped framework

See Azatov, Toharia, Zhu, arXiv 1006.5939

M. Carena et al, 1204.0008

Davoudiasl, McElmurry, A. S. 1206.4062

1. With  $m_{KK} \sim 10$  TeV resulting set up is simpler and economical as then may NOT need to enforce custodial symmetry  
(which requires introducing more dofs....see Agashe et al...)
2. With  $m_{KK} \sim 10$  TeV, tuning is somewhat worse but has the advantage of a more economical theoretical framework and may well be preferable.
3. At LHC at best only radion may be accessible.

## Updated expectations (some based on APS'04)

- $\text{nedm}$  within factors of  $O(\text{few})$  close to Expt bound  $< 6 \times 10^{-26}$  e-cm
- Null tests extremely important ; Gershon & A. S. '07
- Time dependent CP Bd $\Rightarrow$   $K(\pi)\pi \gamma$  ; Bs  $\Rightarrow \phi \gamma \sim O(10\%)$  ...Atwood, Gronau, AS, '97; Atwood,GHS'04; Grinstein,Grossman,Ligeti,Pirjol'05...
- $\Delta \sin 2\beta(\text{penguins}) \sim O(\text{few } \%)$  i.e. comparable to QCD uncertainties.....Grossman & Worah; London & AS .....may need lattice
- $\Delta \gamma \sim O(2 \times 10^{-3})$  GLW; ADS, GGSZ; comp to theo uncertainties
- Charm CP esp. modes where SM predicts 0...e,g D $\Rightarrow$ KKX,  $\phi \pi^+$  ,  $\pi^+ \pi^0$ ..... See Brod,Grossman,Kagan,Zupan'12; Feldmann, Nandi , AS'12; Atwood, A.S. '13..
- $\varepsilon'/\varepsilon$  see also Gedalia, Isidori, Perez'09...RBC-UKQCD lattice
- $K \Rightarrow \pi^+ \nu \nu$  ;  $KL \Rightarrow \pi^0 \nu \nu$
- $t \Rightarrow c Z, ch$  Br  $O(10^{-7})$  ;  $t \Rightarrow c g$   $O(10^{-5})$  ;  $t \Rightarrow c \gamma$   $O(10^{-6})$  ; APS'06
- $ee \Rightarrow tc$  [Atwood, Reina, AS'95]; APS'07;  $R_{tc} \sim 10^{-6} - 10^{-5}$
- $t \text{edm} \sim O(10^{-20})$  Atwood et al Phys Rpt
- Triple correlation in  $ee \Rightarrow tth$  ; Bar-Shalom et al'95; Atwood et al Phys Rpt
- $\Delta$  SM in  $h \Rightarrow bb$

# Null Tests of SM-CKM

**Table 2.** Illustrative sample of *approximate null tests (ANTs)*, with rough SM expectations and theoretical errors, current experimental uncertainties and estimates of numbers of  $B$  mesons needed for a Super  $B$  Factory to approach the SM uncertainty. More details for each mode can be found in the text.

Observable	SM expectation	Current experimental uncertainty ( $B\bar{B}$ pairs used)	$B\bar{B}$ pairs needed
$\Delta S[\eta' K^0, \phi K^0, K^0 \bar{K}^0 K^0, \dots]$	$\sim(0 \pm 2)\%$	20% ( $6 \times 10^8$ )	$5 \times 10^{10}$
$\mathcal{A}_{CP}^{s+d}[M^0 X_{s+d}]$	$\lesssim 0.1\%$	–	$> 10^{12}$
$\mathcal{A}_{CP}[X_s \gamma]$	$(0.5 \pm 0.2)\%$	4% ( $2.4 \times 10^8$ )	$10^{11}$
$\mathcal{A}_{CP}[X_d \gamma]$	$(-10 \pm 5)\%$	–	$10^{11}$
$\mathcal{A}_{CP}[X_{s+d} \gamma]$	$(0.000 \pm 0.001)\%$	12% ( $10^8$ )	$> 10^{12}$
$\mathcal{A}_{CP}[X_s l^+ l^-]$	$(-0.2 \pm 0.2)\%$	26% ( $10^8$ )	$10^{12}$
$\mathcal{A}_{CP}[X_d l^+ l^-]$	$(4 \pm 4)\%$	–	$10^{12}$
$\mathcal{A}_{CP}[X_{(s,d)} l^+ l^-]$	–	–	$> 10^{12}$
$\Sigma(\mathcal{A}_{CP}(\pi K))$	$(0 \pm 1)\%$	15% ( $6 \times 10^8$ )	$> 10^{11}$
$\mathcal{A}_{CP}(\pi^+ \pi^0)$	$\lesssim 1\%$	6% ( $6 \times 10^8$ )	$10^{10}$
$S[K_S \pi^0 \gamma, \dots]$	$\sim(0 \pm 5)\%$	28% ( $6 \times 10^8$ )	$> 10^{10}$
$\langle p_t^\tau \rangle (D(X_c) \tau \nu_\tau)$	0	–	$> 10^{12}$

T. Gershon + AS JHEP'07

**Expected deviations tend  
to be very small, strongly  
suggesting we need to  
strengthen both our  
computational AND  
measurement  
infrastructure**

## CP VIOLATION IN TOP PHYSICS

David ATWOOD<sup>a</sup>, Shaouly BAR-SHALOM<sup>b</sup>, Gad EILAM<sup>c</sup>, Amarjit SONI<sup>d</sup>

Containing studies of many reactions  
for  $e^+e^-$  &  $pp$  e.g.  
 $t\bar{e}dm$ ;  $e^+e^- \rightarrow tc$ ;  $t\bar{t}h$  ...

Table 2: Real and Imaginary parts of  $d_t^\gamma$  and  $d_t^Z$  (in units of  $10^{-19}$  e-cm) for  $m_h = 100, 200$  and  $300$  GeV and for  $\sqrt{s} = 500$  GeV and  $\sqrt{s} = 1$  TeV (in parenthesis).  $\tan \beta = 0.3$  and Set I,II,III means  $\{\alpha_1, \alpha_2\} = \{\pi/4, \pi/2\}, \{\pi/4, 3\pi/4\}, \{\pi/4, \pi/4\}$ , respectively.

Type of moment ( $10^{-19}$ e - cm) $\Downarrow$	$m_h$ (GeV) $\Downarrow$	The different Sets of $\{\alpha_1, \alpha_2\}$ , $\tan \beta = 0.3$		
		Set I	Set II	Set III
$\Re(d_t^\gamma)$	100	1.97(3.77)	1.40(2.66)	1.40(2.66)
	200	-3.36(2.26)	-2.38(1.60)	-2.38(1.60)
	300	-4.75(1.27)	-3.36(0.90)	-3.36(0.90)
$\Im(d_t^\gamma)$	100	-23.89(-5.44)	-16.88(-3.84)	-16.88(-3.84)
	200	-16.56(-4.91)	-11.70(-3.47)	-11.70(-3.47)
	300	-11.34(-4.33)	-8.02(-3.06)	-8.02(-3.06)
$\Re(d_t^Z)$	100	0.62(1.25)	0.36(0.83)	0.52(0.93)
	200	-1.17(0.74)	-0.87(0.47)	-0.78(0.57)
	300	-1.57(0.40)	-1.04(0.24)	-1.18(0.33)
$\Im(d_t^Z)$	100	-7.96(-1.81)	-5.41(-1.21)	-5.85(-1.34)
	200	-5.45(-1.62)	-3.58(-1.08)	-4.12(-1.22)
	300	-3.64(-1.42)	-2.22(-0.93)	-2.91(-1.08)

See also  
Gunion,  
Grzadkowski  
& He, PRL 96

$e^+e^- \rightarrow t\bar{t}h$   
Triple  
Correlations

Atwood et al  
PRD 96

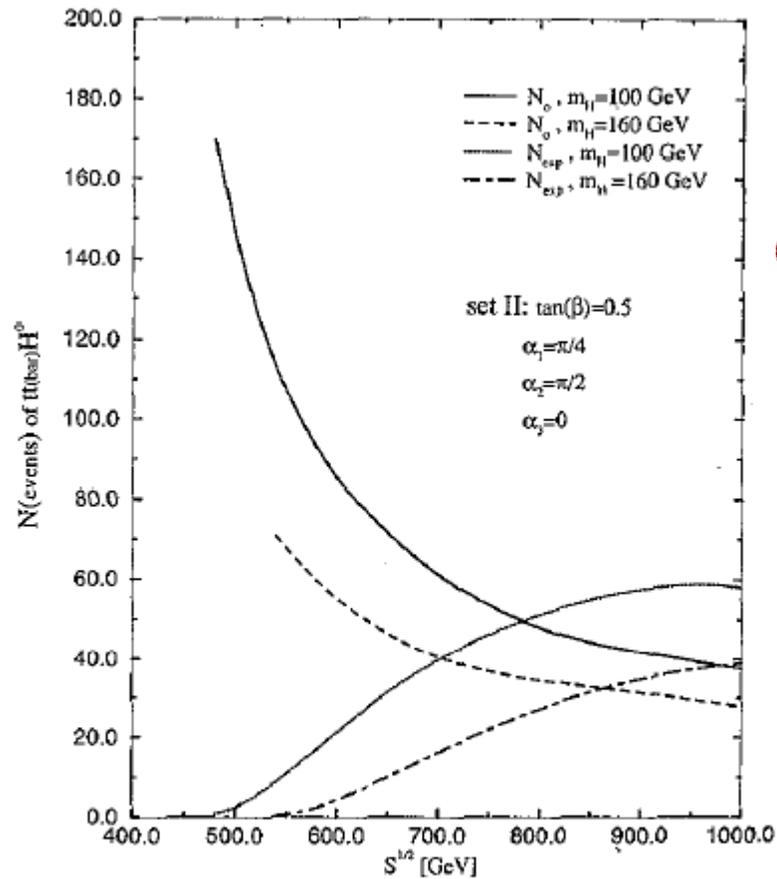


FIG. 3. Number of events,  $N_O$  ( $N_{exp}$ ) required (expected yearly), as a function of total beam energy for set II of the parameters and for  $m_{H^0} = 100$  and 160 GeV with unpolarized electron and positron beams.

**Constraining the FC  $h_{tc}(u)$  vertex  
in the context of RS [i.e. (partial)  
compositeness] is of special  
significance**

*$h_{tc}(u)$  vertex should be beaten to death!*

# Constraining FCNC: Exploit (PARTIAL) COMPOSITENESS

Agashe et al'06

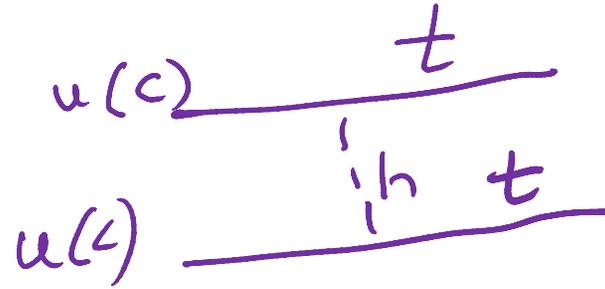
**CSAKI ET AL JHEP'08**

**AGASHE ET AL PRD'07**

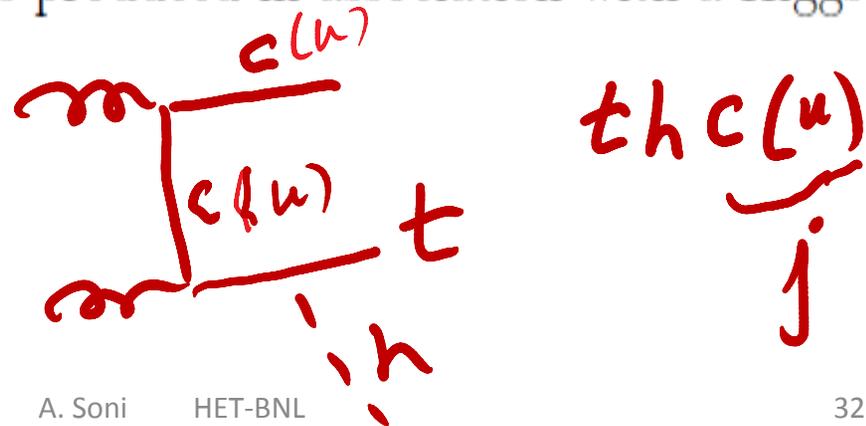
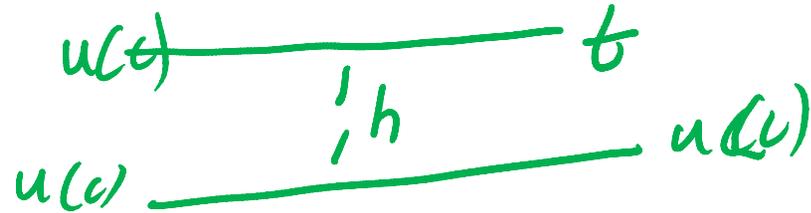
**AGASHE, CONTINO, PRD'09**

**AZATOV ET AL PRD'09**

**KEREN-ZUR ET AL, NPB '13**



- (a) the same-sign top-pair, e.g.  $pp \rightarrow tt(\bar{t}\bar{t})$ ,
- (b) processes where a top-quark is produced in association with a light jet, e.g.  $pp \rightarrow t\bar{j}_u(\bar{t}j_u)$ , where  $j_u = u, c$ , and,
- (c) processes where the top-quark is produced in association with a Higgs and a light jet, e.g.  $pp \rightarrow t\bar{j}_u h(\bar{t}j_u h)$ .



## Effective tree-level

$$\sqrt{\xi_{tc}^2 + \xi_{tu}^2} \lesssim \begin{cases} 0.25 \text{ (0.2)} & \text{for process (a)} \\ 0.9 \text{ (0.9)} & \text{for process (b)} \\ 0.26 \text{ (0.1)} \times 10^{-4} & \text{for process (c)} \end{cases}$$

$$\sqrt{s} = 7 \text{ TeV,} \\ 5/f_6$$

$$\sqrt{s} = 8 \text{ TeV, } 22/f_6$$

## Production level constraints

$$[RS \text{ bias } \{t_c\} \gg \{t_u\}]$$

Harnik, Kopp, Zupan  
arXiv:1209.1397

Technique	Coupling	Constraint
$D^0$ oscillations [48]	$ Y_{uc} ^2,  Y_{cu} ^2$	$< 5.0 \times 10^{-9}$
	$ Y_{uc}Y_{cu} $	$< 7.5 \times 10^{-10}$
$B_d^0$ oscillations [48]	$ Y_{db} ^2,  Y_{bd} ^2$	$< 2.3 \times 10^{-8}$
	$ Y_{db}Y_{bd} $	$< 3.3 \times 10^{-9}$
$B_s^0$ oscillations [48]	$ Y_{sb} ^2,  Y_{bs} ^2$	$< 1.8 \times 10^{-6}$
	$ Y_{sb}Y_{bs} $	$< 2.5 \times 10^{-7}$
$K^0$ oscillations [48]	$\text{Re}(Y_{ds}^2), \text{Re}(Y_{sd}^2)$	$[-5.9 \dots 5.6] \times 10^{-10}$
	$\text{Im}(Y_{ds}^2), \text{Im}(Y_{sd}^2)$	$[-2.9 \dots 1.6] \times 10^{-12}$
	$\text{Re}(Y_{ds}^* Y_{sd})$	$[-5.6 \dots 5.6] \times 10^{-11}$
	$\text{Im}(Y_{ds}^* Y_{sd})$	$[-1.4 \dots 2.8] \times 10^{-13}$
single-top production [49]	$\sqrt{ Y_{tc}^2  +  Y_{ct} ^2}$	$< 3.7$
	$\sqrt{ Y_{tu}^2  +  Y_{ut} ^2}$	$< 1.6$
$t \rightarrow hj$ [50]	$\sqrt{ Y_{tc}^2  +  Y_{ct} ^2}$	$< 0.34$
	$\sqrt{ Y_{tu}^2  +  Y_{ut} ^2}$	$< 0.34$
$D^0$ oscillations [48]	$ Y_{ut}Y_{ct} ,  Y_{tu}Y_{tc} $	$< 7.6 \times 10^{-3}$
	$ Y_{tu}Y_{ct} ,  Y_{ut}Y_{tc} $	$< 2.2 \times 10^{-3}$
	$ Y_{ut}Y_{tu}Y_{ct}Y_{tc} ^{1/2}$	$< 0.9 \times 10^{-3}$
neutron EDM [37]	$\text{Im}(Y_{ut}Y_{tu})$	$< 4.4 \times 10^{-8}$

)  $pp \rightarrow t h j$  gives much tighter constraints.

**Table 2.** Constraints on flavor violating Higgs couplings to quarks. We have assumed a Higgs mass  $m_h = 125$  GeV, and we have taken the diagonal Yukawa couplings at their SM values.

$$\sqrt{\xi_{Stc}^2 + \xi_{Stu}^2} \lesssim \begin{cases} 0.47 \text{ (0.36)} & \text{for } pp \rightarrow l^\pm l^\pm + 2b - jets + \cancel{E}_T \\ 0.9 \text{ (0.9)} & \text{for } pp \rightarrow l^\pm + j + b - jet + \cancel{E}_T \\ 4.5 \text{ (1.8)} \times 10^{-4} & \text{for } pp \rightarrow l^\pm + j + 3b - jets + \cancel{E}_T \\ 52.5 \text{ (19.8)} \times 10^{-4} & \text{for } pp \rightarrow l^\pm + j + b - jets + 2\gamma + \cancel{E}_T \end{cases} \quad (5.2)$$

$\sqrt{s} = 7 \text{ TeV}, 5/\text{fb}$

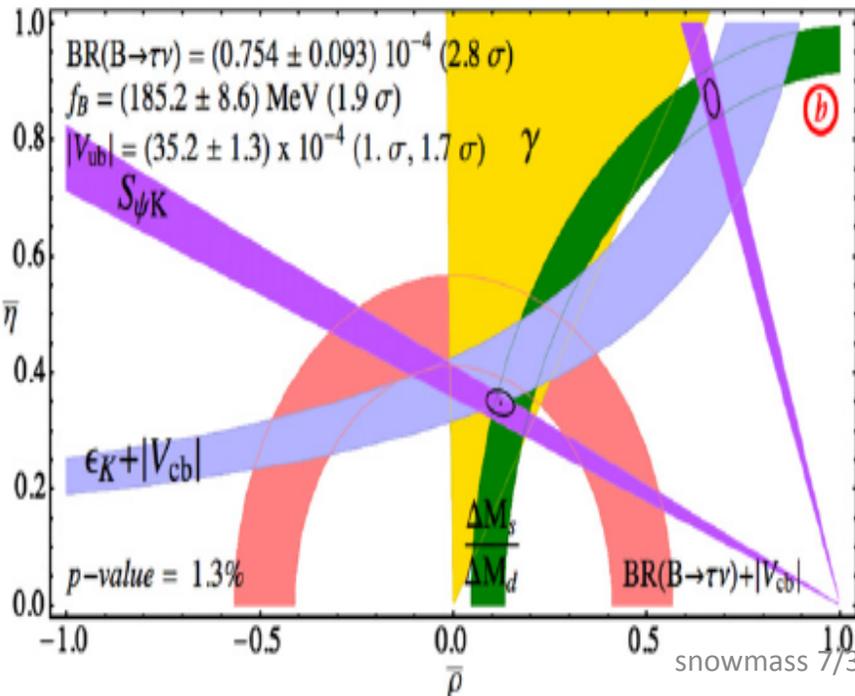
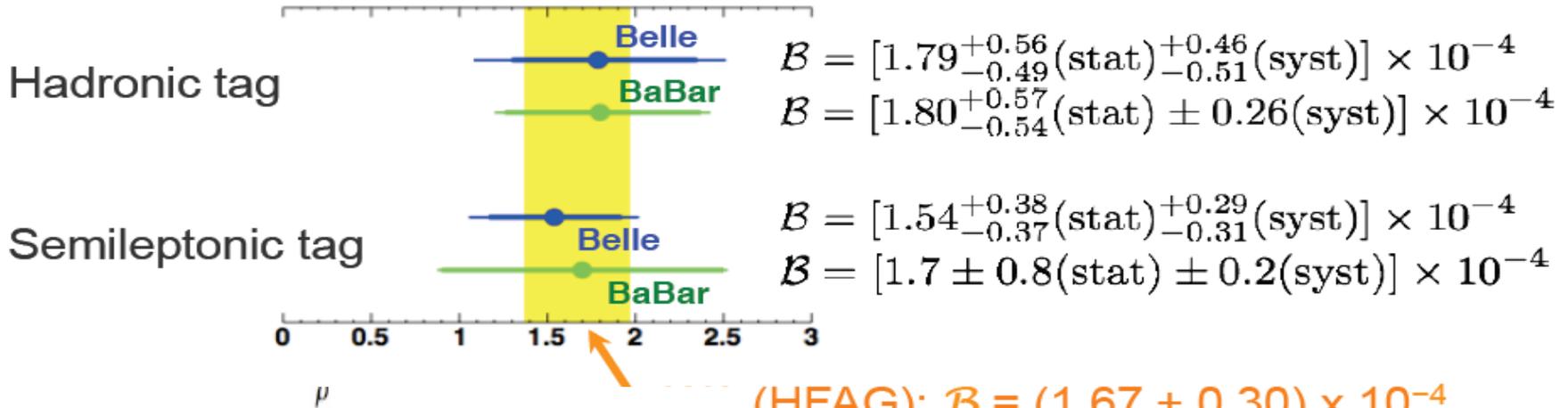
$(\sqrt{s} = 8 \text{ TeV}, 22/\text{fb})$

Observable	Value	Experiment
$\Gamma_t$	$2 \pm 0.7$ GeV	Tevatron [17]
$m_h$	$125.5 \pm 0.2$ (stat) $_{-0.6}^{+0.5}$ (sys) GeV $125.8 \pm 0.5$ (stat) $\pm 0.2$ (sys) GeV $125.7 \pm 0.4$ GeV	ATLAS [3] CMS [4] Combined
$\mathcal{R}_{gg\gamma\gamma}$	$1.65_{-0.30}^{+0.34}$ $1.11_{-0.30}^{+0.32}$ $1.36 \pm 0.23$	ATLAS [24, 25] CMS [26, 27] Combined
$\mathcal{R}_{gg2l2\nu}$	$1.01 \pm 0.31$ $0.76_{-0.21}^{+0.21}$ $0.84 \pm 0.17$	ATLAS [28, 29] CMS [30, 31] Combined
$\mathcal{R}_{gg4l}$	$1.7_{-0.4}^{+0.5}$ $0.91_{-0.24}^{+0.30}$ $1.12 \pm 0.26$	ATLAS [24, 32] CMS [30, 33] Combined
$D0$ -oscillations	$ \xi_{tu}\xi_{tc}  < 0.9 \times 10^{-3}$	UTfit Collaboration [34]

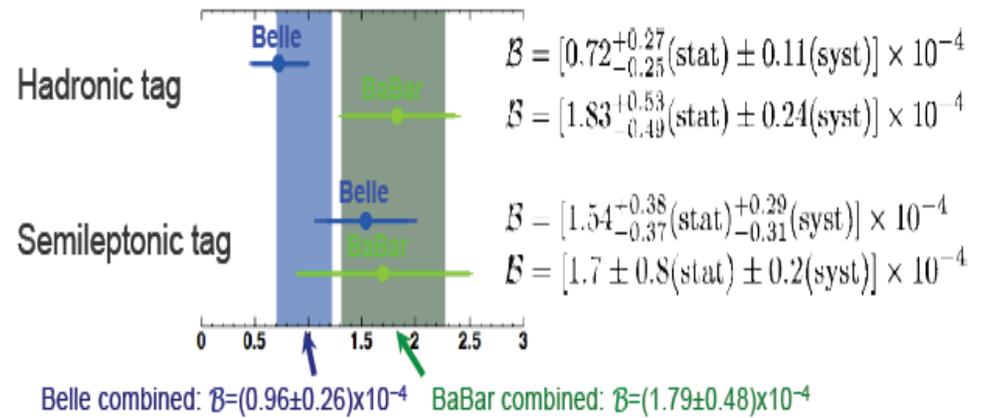
Table 3: Measured values of various observables used in our analysis; combined here means weighted average of ATLAS and CMS values for a given observable.

# ANOMALIES (A SAMPLE)

# Status for $B \rightarrow \tau \nu$ before ICHEP 2012



# Status for $B \rightarrow \tau \nu$ after ICHEP 2012



A naive world average:  $B = (1.15 \pm 0.23) \times 10^{-4}$

LUNCHEI + APS PLB '11

## $B \rightarrow D^{(*)} \tau \nu$ from BaBar and SM

$$\begin{array}{ccc}
 \mathcal{R}(D)_{\text{exp}} = 0.440 \pm 0.072 & \mathcal{R}(D^*)_{\text{exp}} = 0.332 \pm 0.030 & \\
 \updownarrow 2.0\sigma & \updownarrow 2.7\sigma & \\
 \mathcal{R}(D)_{\text{SM}} = 0.297 \pm 0.017 & \mathcal{R}(D^*)_{\text{SM}} = 0.252 \pm 0.003 & 
 \end{array}$$

SM expectations in S. Fajfer, J. Kamenik, I. Nisandzic, PRD 85, 094025 (2012).

## $B \rightarrow D^{(*)} \tau \nu$ from Belle

A. Bozek's averages (KEK-FF 2013):

(naive averages for inclusive and exclusive hadronic tags)

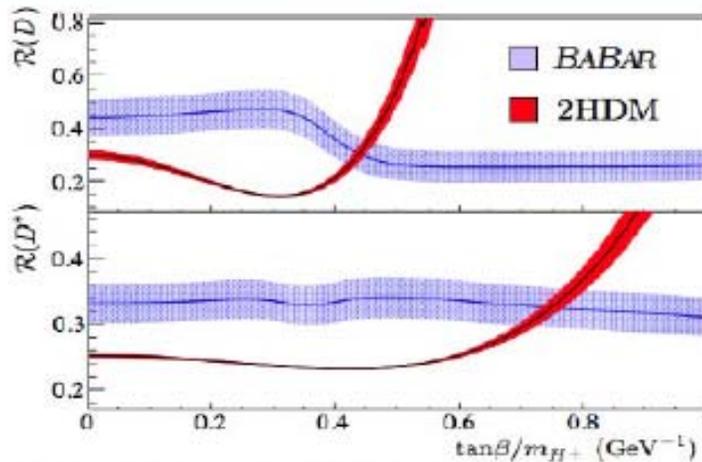
	Deviation from SM
$R(D) = 0.430 \pm 0.091$	$1.4 \sigma$
$R(D^*) = 0.405 \pm 0.047$	$3.0 \sigma$
Combined	$3.3 \sigma$

Correlation btw  $R(D)$  and  $R(D^*)$  neglected conservatively.

[B BHUYAN]



# MSSM: 2HDM-Type II



PRD 78, 015006 (2008)  
PRD 85, 094025 (2012)

$\tan\beta/m_{H^+} =$   
 $0.44 \pm 0.02 \text{ GeV}^{-1}$

$\tan\beta/m_{H^+} =$   
 $0.75 \pm 0.04 \text{ GeV}^{-1}$

R(D) and R(D\*) do not both agree with the predictions for any single value of  $\tan\beta/m_{H^+}$

2HDM-Type-II excluded over the full parameter space\* at 99.8% CL.

\* relies on  $m_{H^+} < 15 \text{ GeV}$  being excluded based on  $b \rightarrow s\gamma$

Perhaps premature

If true another huge stab at SUSY

# FORWARD-BACKWARD TOP ASY

SM e NNLO<sup>+</sup>  
 $0.072^{+0.011}_{-0.007}$

N2

Observable	Values	Experiment
$A_{FB}^t$	$0.19 \pm 0.065$	DØ Collaboration [1]
	$0.158 \pm 0.074$	CDF Collaboration [2]
	$0.176 \pm 0.05$	Combined
$A_{FB}^{t,low}$	$0.078 \pm 0.048$	DØ Collaboration [1]
	$-0.022 \pm 0.043$	CDF Collaboration [2]
	$0.023 \pm 0.032$	Combined
$A_{FB}^{t,high}$	$0.115 \pm 0.060$	DØ Collaboration [1]
	$0.266 \pm 0.062$	CDF Collaboration [2]
	$0.188 \pm 0.043$	Combined
$\sigma_{t\bar{t}}^{Tevatron}$	$8.18^{+0.98}_{-0.87}$ pb	DØ Collaboration [8]
$\sigma_{l\pm l\pm}^{LHC}$	$< 1$ fb	ATLAS & CMS Collaborations [9]

Table 1: Measured values of various observables used in our analysis; combined here mean weighted averages.

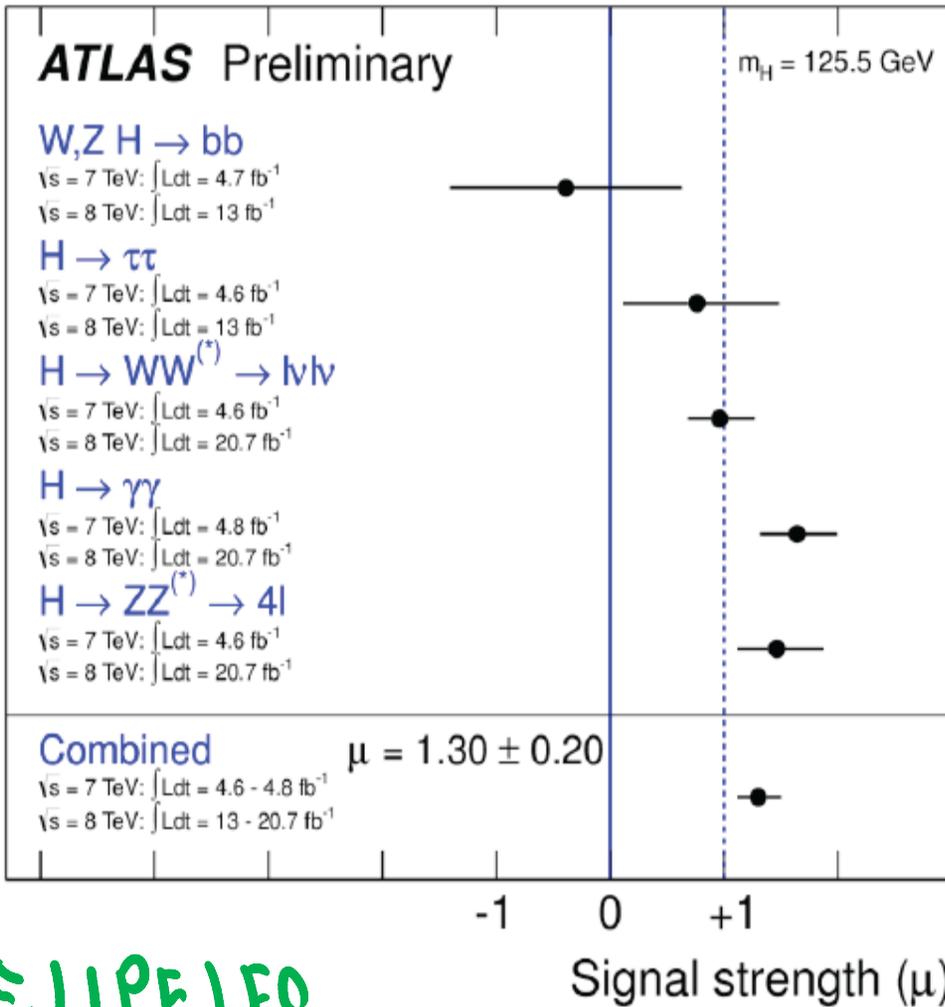
+

V. Ahrens, A. Ferroglia, M. Neubert, B. D. Pecjak and L. L. Yang, JHEP 1009, 097 (2010) [arxiv:1003.5827 [hep-ph]].  
 Atwood, Gupta, AS, arXiv1301.2250

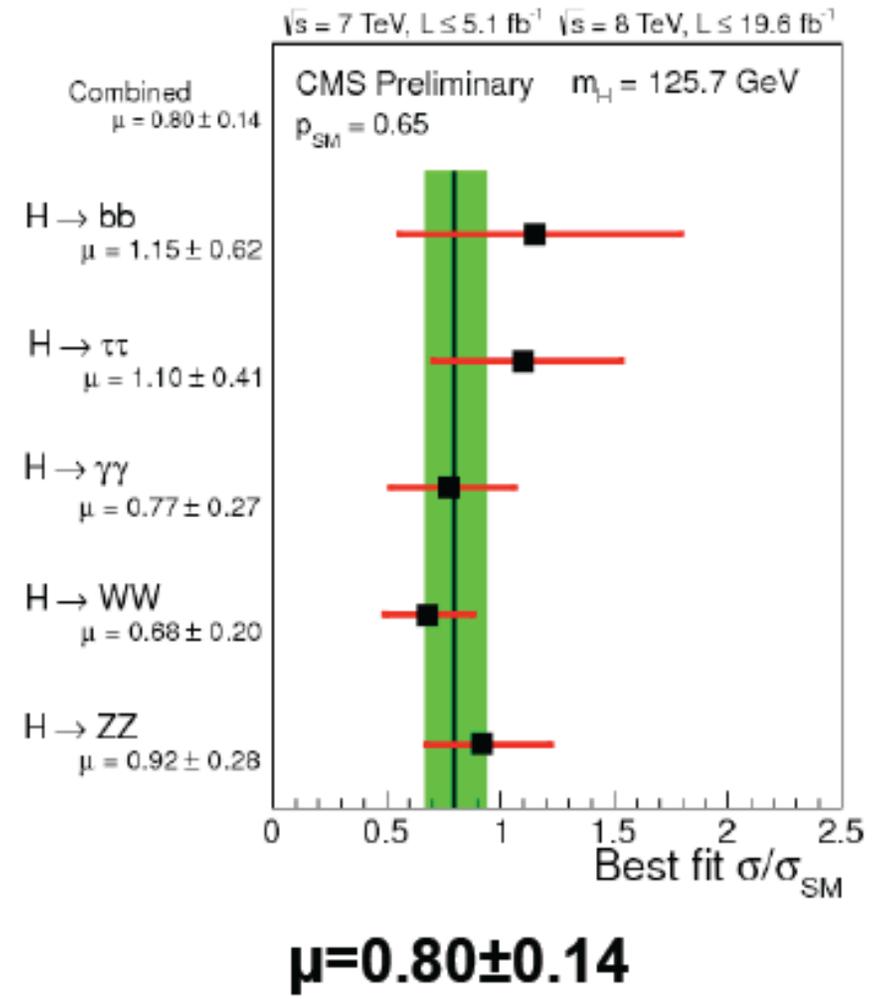
powerful discriminator SST

SMS-Flavor-CP A. Soni

Berger et al 1101.5625  
 Aguilar-Saavedra, Perez-Victoria, 1104.1385  
 Degrande et al 1104.1798



E. LIPELES  
 PBF 2013



S. BOSE PBF 2013

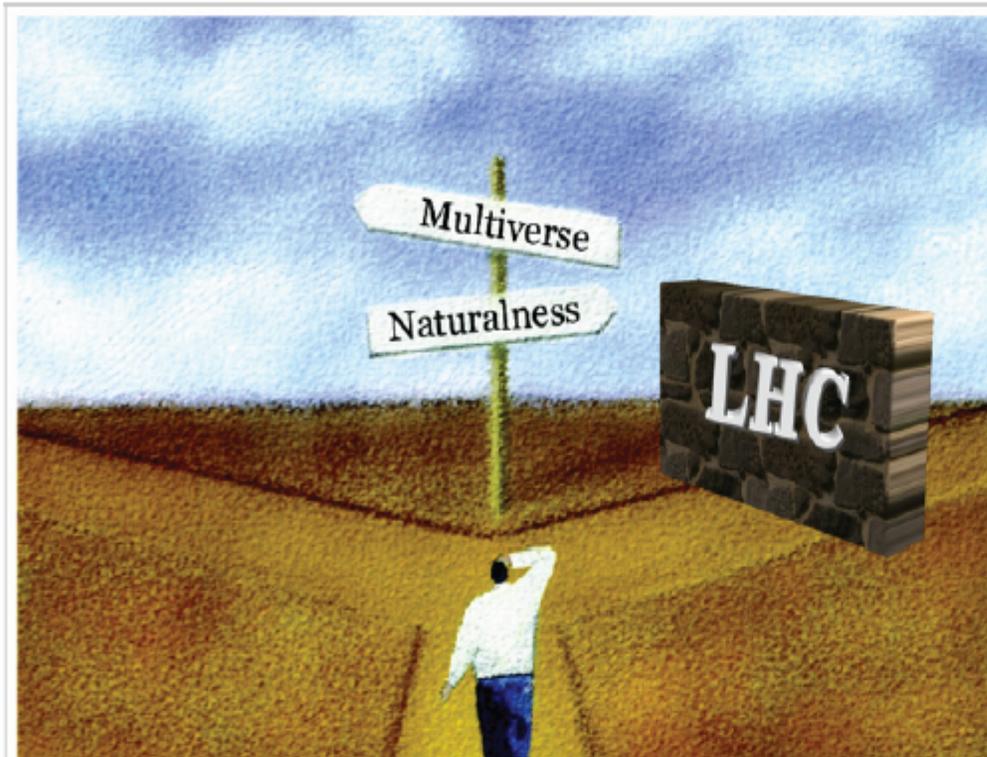
## Is Nature Unnatural?

Decades of confounding experiments have physicists considering a startling possibility: The universe might not make sense.

by: [Natalie Wolchover](#)

May 24, 2013

[email](#) [print](#)



Is the universe natural or do we live in an atypical bubble in a multiverse? Recent results at the Large Hadron Collider have forced many physicists to confront the latter possibility. (Illustration: Giovanni Villadoro)

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Science Lives

Gee, don't see no NP signals  
Flavor: Told you so!

# **KEY MESSAGES FROM A CANDIDATE THEORY OF FLAVOR**

## Key messages from a candidate theory of flavor

1. In a candidate theory, the gigantic tension between hierarchy and flavor puzzle gets dramatically ameliorated. *Thus remarkably RS-leads to lowering of  $\Lambda_{flavor}$  from  $\sim 1000$  to  $\sim 10$  TeV*

II. Due to flavor mis-alignment,  $O(1)$  BSM phases occur naturally;  $\Rightarrow$  direct CP is an extremely powerful probe of flavor alignment and holds the key to unlocking new physics. For this purpose, fortunately, there are many observables :  $N_{edm}$ ;  $\varepsilon'/\varepsilon$ ;  $\gamma$ ;  $\Delta \sin 2\beta$  from  $B_d \Rightarrow \eta' K_s$ ,  $\phi K_s$ ,  $3 K_s$ ....;  $ACP(B \Rightarrow K\pi)$ ,  $S[B \Rightarrow K^* (\rho)\gamma]$ ;  $\Delta ACP(D)$ ....but expected signals tend to be small.

**III. Top quark is very sensitive to flavor violation;  $t \Rightarrow c Z$ ;  $t \Rightarrow c h$ ,  $pp \Rightarrow t c h X$  etc need to be vigorously pursued**

*Agashe, Blechman, Petriello  
106*

**IV. Lepton flavor violation is a natural prediction  $\Rightarrow$  Searches for  $\tau = \mu \gamma$ ,  $3 \mu \dots$ ;  $B_s \Rightarrow \tau \mu \dots$  are very important.**

## V. Expected size of corrections to Higgs couplings

- Deviation from SM  $\sim O(1\%)$

[assuming  $m_{KK} > \sim 10 \text{ TeV}$  ]

- Small corrections are a concern.

- VI. For direct observation of KK-particles of mass  $> \sim 10 \text{ TeV}$  need a Gigantic International Hadron Collider (GIHC)  $\sim 100 \text{ TeV}$  cm energy

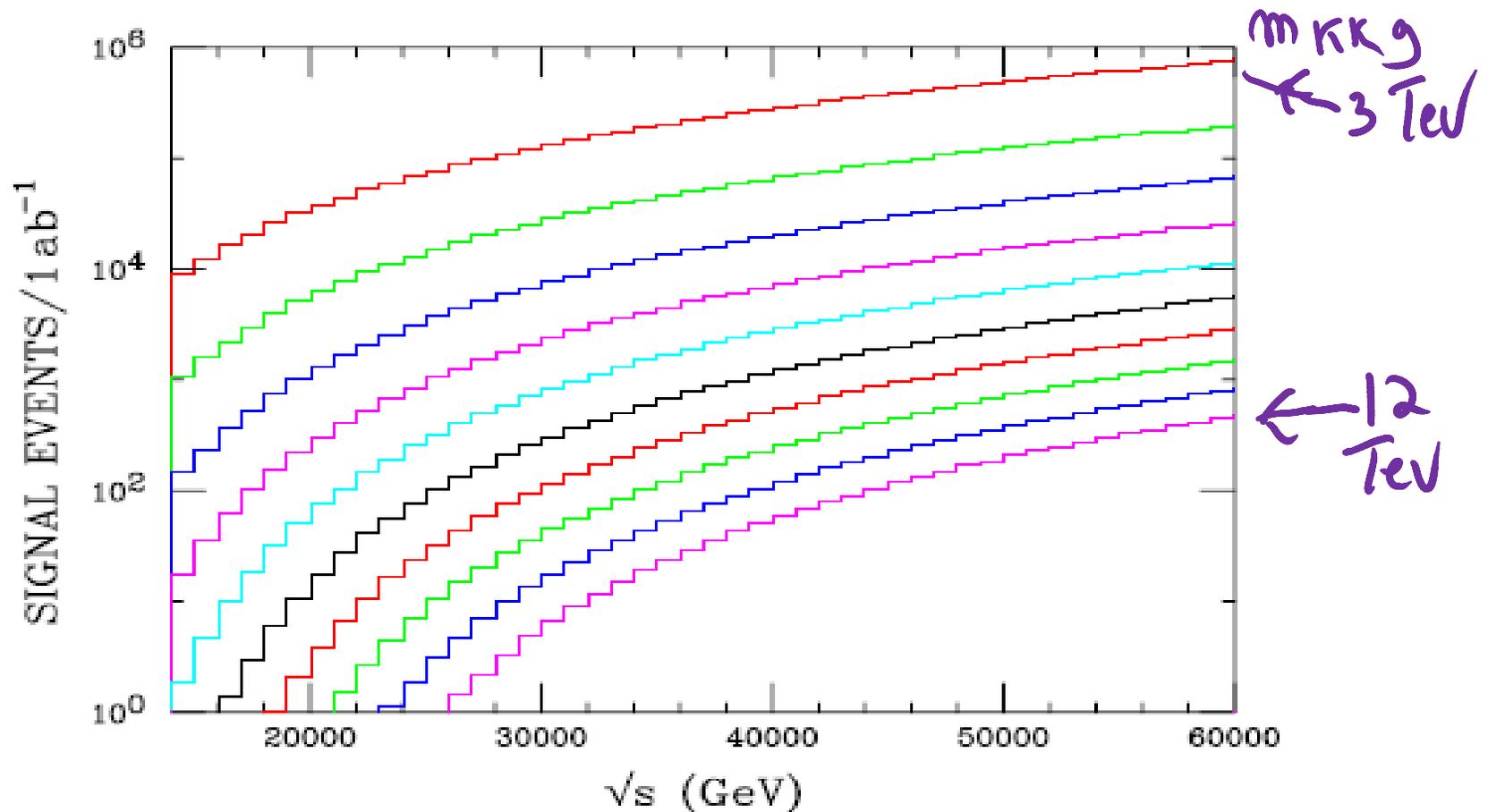


FIG. 10 (color online). Signal rate for a possible gluon KK resonance as a function of the collider energy employing the cuts described in the text. Branching fractions and efficiencies have been neglected. From top to bottom, the results are shown for gluon KK masses in the range from 3 to 12 TeV in steps of 1 TeV.

# Lesson learnt from $\nu$ 's

~ Circa 1983, after long and arduous efforts,  $\Delta m^2$  upper bound used to be around a few  $\text{eV}^2$  but efforts to search oscillations continued basically because there was no good theoretical reason for  $m_\nu$  to be zero.

- *Recall it took more than a decade beyond '83* and  $\Delta m^2$  had to be lowered by almost 4 orders of magnitude (!) before osc were discovered.
- **Moral: Physical “principles” [ i.e Naturalness in this instance ] shouldn't be abandoned easily .....We'll just have to work harder to get to it but this should have been anticipated if enough attention had been paid to flavor alignment**

Recall SSC  $\sim 40$  TeV 1990 technologically completely feasible.

We should be SERIOUSLY

THINKING of

GIGANTIC INTERNATIONAL  
HADRON COLLIDER [GIHC]

$\sim 100$  TeV CM

"GEEK"

# Summary & Outlook

- While naturalness is not tangible, [clearly  $10^{-2}$  OR  $10^{-4}$  are very different from  $10^{-34}$  ], flavor-alignment places specific constraints...has been telling us for long that scale of NP  $\gg 1$  TeV
- Specifically RS-flavor (which gives a nice geometric understanding of flavor & simultaneously of EW-Plank hierarchy ) strongly suggests scale is unlikely less than  $\sim 10$  TeV and the following deserve attention:
- Dir CP probes [e.g.  $\text{nedm}$ ,  $\varepsilon'/\varepsilon$ ,  $S[B \Rightarrow K \rho \gamma]$ ;  $\gamma$ ; Null Tests, D-CP...
- Top FV via e.g.  $t \Rightarrow c Z$ ;  $t \Rightarrow c h$ ;  $pp \Rightarrow t c h$ ;  $e^+ e^- \Rightarrow t c$
- $\tau$ FV:  $\tau \Rightarrow \mu \gamma$ ;  $3 \mu$ ;  $B_s \Rightarrow \tau \mu$
- Expected deviation to higgs couplings  $\sim O(1\%)$  need attention.
- Precise measurements & precise computations deserve high priority.
- It is essential to have high sensitivity flavor experiments AND we should be seriously thinking of a GHC as the next step in our adventure.

# XTRA