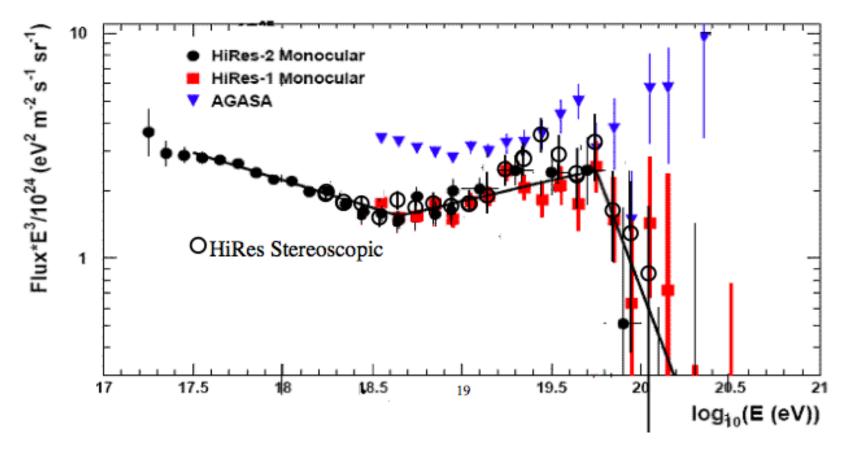
The Telescope Array Low-Energy Extension (TALE)



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Features in the UHECR spectrum: The minority straight-forward View

- CMBR photons interact with cosmic ray protons:
 - Pion production makes the *GZK suppression*: $E < 6x10^{19}$ eV if cosmic rays travel > 50 Mpc.
 - $-e^+e^-$ production: threshold ~ $4x10^{17}$ eV, excavates the **ankle**.
- Pair production pileup + galactic/extragalactic transition: the second knee.
- Three spectral features in the UHE regime.
- But NO single experiment has measured all three features: the exact energies (or even the ratios) are not known; i.e., basic information is in doubt.
- The field needs an experiment with WIDE energy coverage!! good resolution!! and good systematics!!

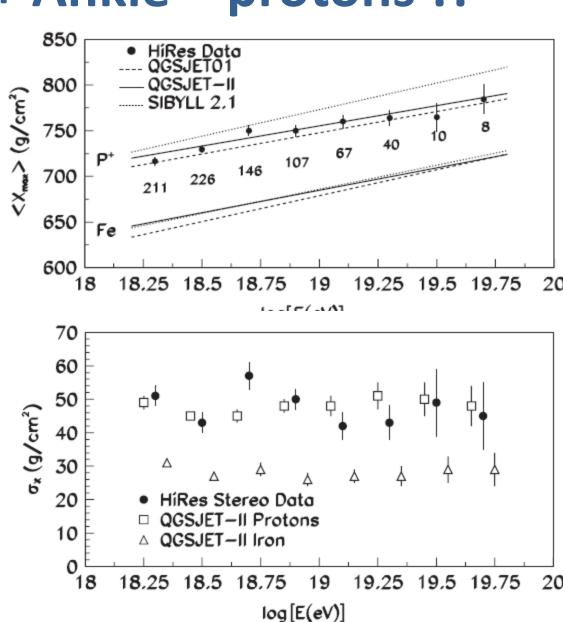


- HiRes was first to observe the GZK suppression
 - Physical Review Letters 100 (10): 101101 (5σ significance)
 - Confirmed in stereo spectrum (4σ significance)
- Ankle ("dip") was also seen by HiRes

HiRes: GZK + Ankle = protons!?

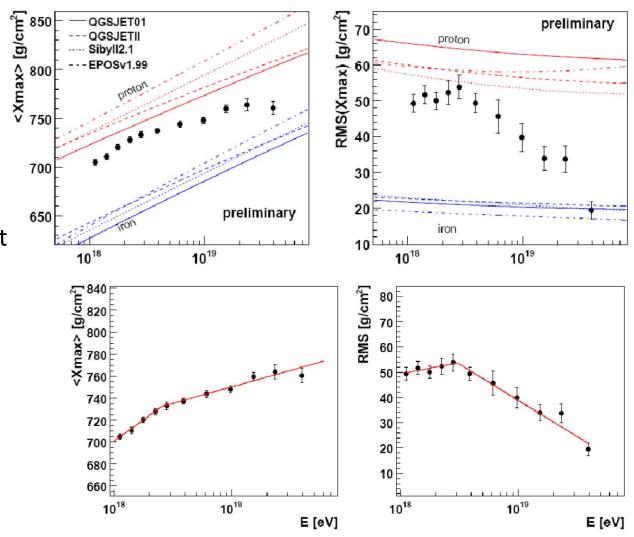
- The observation of the GZK and ankle features are consistent with proton dominated composition (e.g. Berezinsky & Gazizov Phys Rev. D 74, 043005 2006)
- Comparison of HiRes
 <Xmax> vs. energy plot
 againt hadronic models
 show a unchanging (from
 the slope), light
 composition

HiRes Result: Phys Rev Lett. 104, 161101 (2010)

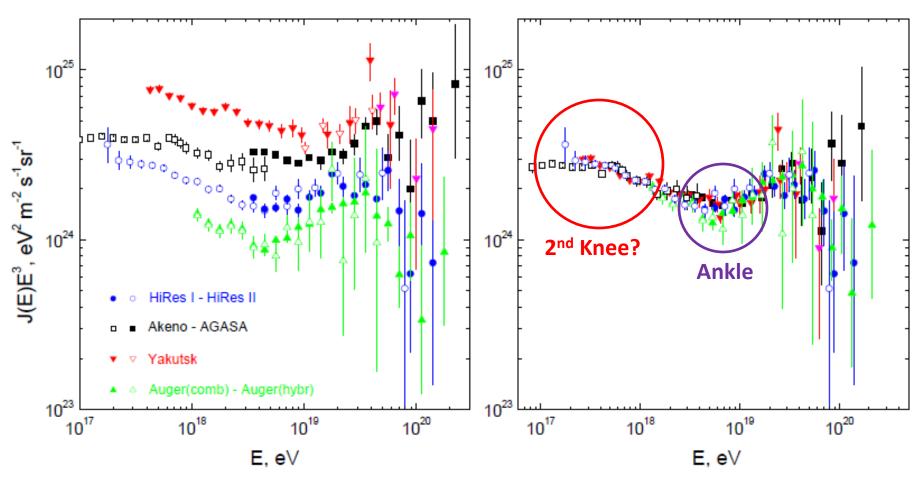


- >60% of the UHECR field disagrees with proton dominance
- The popular interpretation seems to be that the breaks in these plots suggests transition to iron starting below the ankle
- The two-component interpretation of these plots are of course problematic: a 50%-50% mix of iron and proton would have RMS > ~70 g/cm²

Gallup Poll Headline: Proton Dominance Approval Rating Down



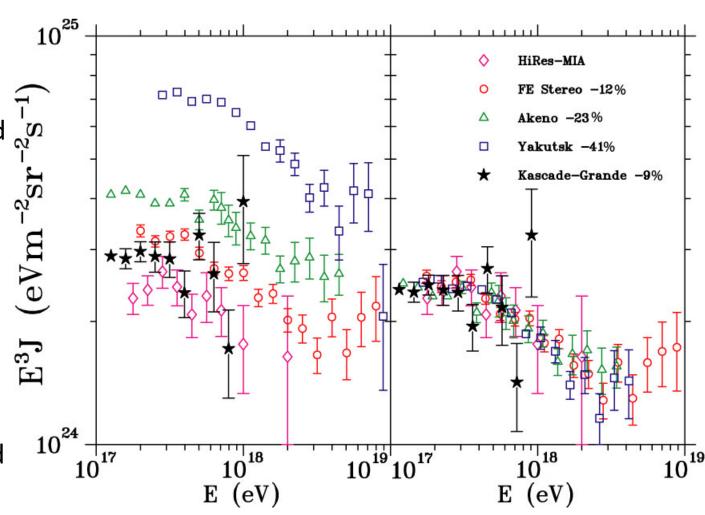
UHECR spectra consistency



 Nevertheless, the spectra from the four previous experiments can be made to agree both in shape and normalization by energy rescaling alone The 2nd knee feature has been seen by 4 previous experiments

- Can make the spectra agree in normalization and in location of feature by adjusting energy scales
- Recent results from Kascade-Grande seems to run out of statistical power at the feature and is inconclusive

The Second Knee



Galactic Sources

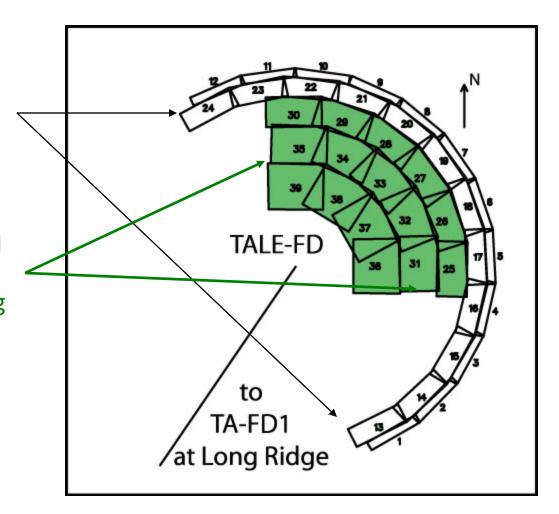
- While a minority of the UHECR field (~20%) might not agree that UHE cosmic rays are definitively correlated with AGN's, there appear to be consensus that there are Galactic sources of cosmic rays.
- Questions about Galactic sources:
 - What is the maximum energy they produce?
 - Is there anisotropy at 10^{18} eV?
- TA/TALE aim: attack these questions:
 - Measure spectrum and composition at lower energies where galactic contribution is larger.
 - Search for anisotropy along galactic plane, and just above the galactic center.
- X_{max} is the variable that discriminates between p and Fe primaries.
- Fluorescence gives direct observation
 best technique.
- Choose stereo and hybrid: each has x2 better X_{max} resolution than mono.
- Paradoxical indication by HiRes+(HiRes-MIA) and Fly's Eye stereo:
 - "early" transition; i.e., below the ankle (HiRes)
 - "late" transition; i.e., above supernova capability (Fly's Eye + AUGER?).
- The next experiment needs a WIDE energy range.

TALE Design Criteria

- Second Knee region:
 - Transition from Galactic to extra-galactic flux?
 - The energy scale of this feature is uncertain (somewhere in the 10¹⁷ eV decade)
 - Need long lever arm below
- → Hybrid measurements from 10^{16.5}-10¹⁸ eV
- Ankle region:
 - Need better, and more stereo data than HiRes (HiRes sites were too far apart at 12.6 km)
- →~6 km site separation stereo measurements

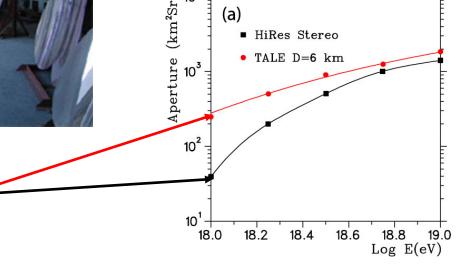
TALE fluorescence

- The Fluorescence detectors of TALE consist of two overlapping components:
 - 24 telescopes viewing up to 31° in elevation to provide stereo coverage in conjunction with existing TA fluorescence station
 - 15 "Tower" telescopes with 4m diameter mirrors (~3 fold increase in light collection area from HiRes) and viewing up to 73 in elevation
- Uses refurbished HiRes-2 FADC electronics and PMTs
- Tower: new, larger mirrors and scaled up PMT clusters with Winston Cones

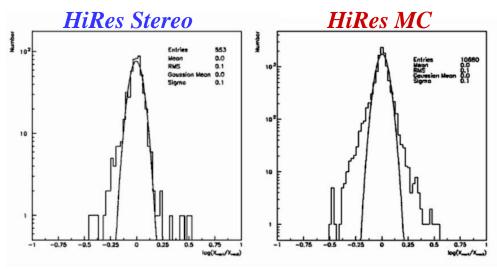




TALE Stereo

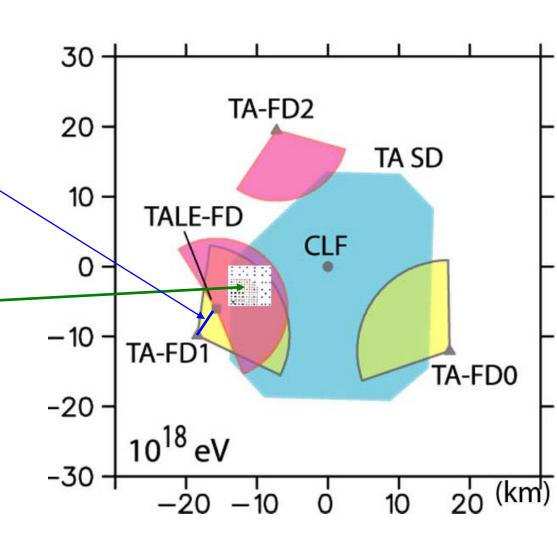


- Aperture is much flatter than the HiRes stereo aperture.
- Aperture at 10¹⁸ eV is ~6× that of HiRes stereo: comparable to hybrid aperture (is there a systematic effect between stereo and hybrid composition measurements?)
- Stereo: redundant measurement of shower properties (e.g. E and Xmax) which allows DIRECT validation of MC



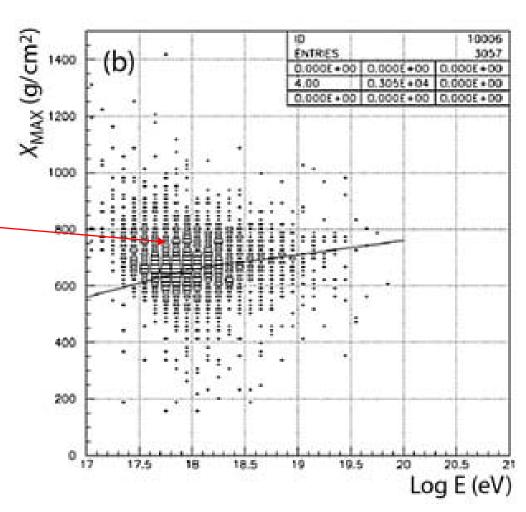
An example of TALE-FD placement

- Within the 4-7 km site separation window ideal for stereo fluorescence measurements in the 10¹⁸-10¹⁹ eV decade.
- Borders the periphery of the ground array where construction of infill arrays are possible (and needed)
- Availability of State
 Education trust Land
 (without BLM access and use constraints)



"31 Bias": Xmax below ~10¹⁸ eV

- X_{MAX} measurements below 10¹⁸ eV are beyond the scope of HiRes and Auger. TA is only a little better.
- Two-ring (<31° elevation) configuration introduces significant trigger bias toward low Xmax (heavy composition) showers
- TALE needs additional elements to cover this region, which contains the Second Knee Structure



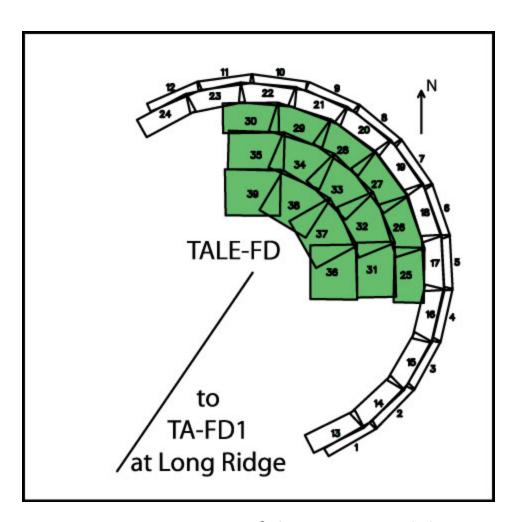
Data points show XMAX of triggered events: Line gives measured/corrected <XMAX>

13

The TALE Tower detector consists of 15 telescopes in its top three "rings":

- 6 (3) at 31-45°
- 5 (3) at 45-59°
- 4 (4) at 59-73°
- # in parenthesis shows the number of mirrors in the HiRes tower prototype at the same elevation
- The 6km telescopes also provide 16 telescopes directly below the top three rings compared to only 4 in the HiRes-prototype
- Stereo overlap with existing TA fluorescence Station validation of MC resolutions

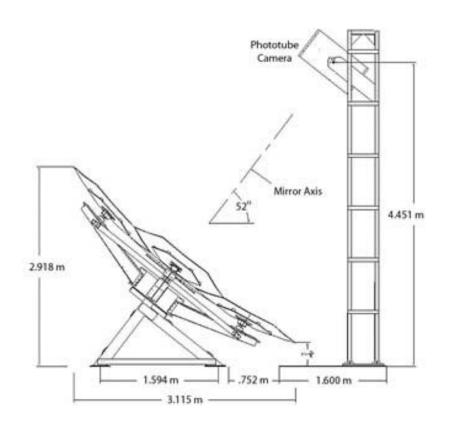
TALE Tower

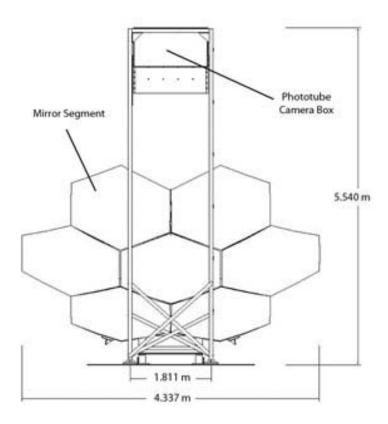


Top view projection of the viewing solid angles of the TALE telescopes

Tower Detector

- Use ~4 m diameter mirrors to triple the collection area over those of the redeployed HiRes mirrors in the 6km stereo detector.
- Eliminates trigger Xmax bias in the 10^{16.5}-10^{18.0} eV range
- Use scaled-up ~F1.1 optics identical to HiRes
- Re-use PMTs from HiRes telescopes
- Use Winston cones for light collection





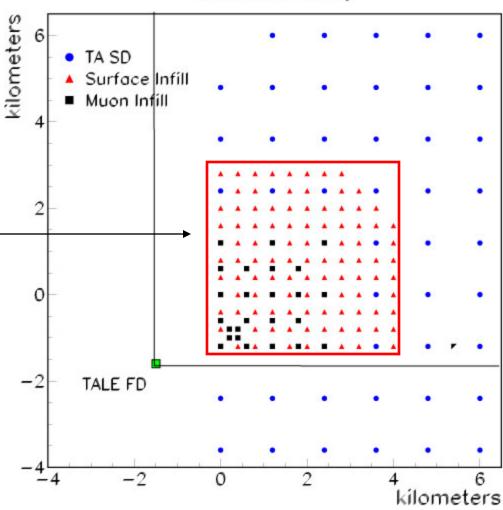
Hybrid with Infill Array (Japan)

TALE Infill Array

 Will place 111 additional surface array counters overlapping with main ground array: 4km x 4km

 For the Tower detector these counters provide hybrid trajectory reconstruction for the lowest energy events that fall outside of the stereo overlap

 25 muon detector array placed in the "inner corner" of the infill array, under 3m of packed soil

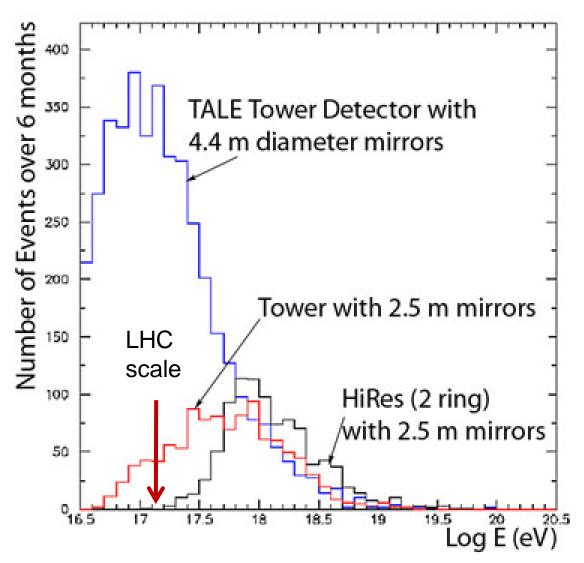


This 2.5km x 2.5km graded array is designed to work at 10^{16.5}-10¹⁸ eV

The increased mirror size will improve substantially the sensitivity of TALE in the 10^{16.5}-10^{17.5} eV energy decade

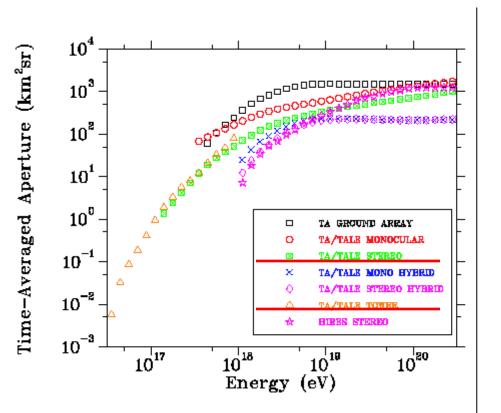
- Note the gain in sensitivity comes from the improvement in signal.
- Energy, angular and X_{max} resolutions comparable to HiRes stereo or TA hybrid

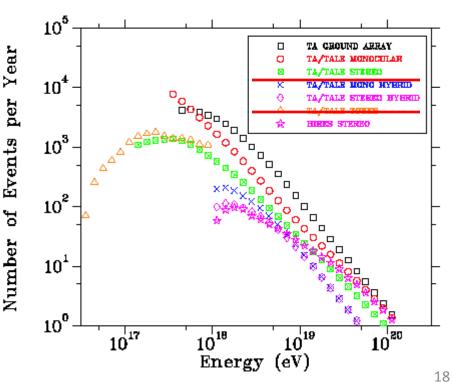
Improved Sensitivity



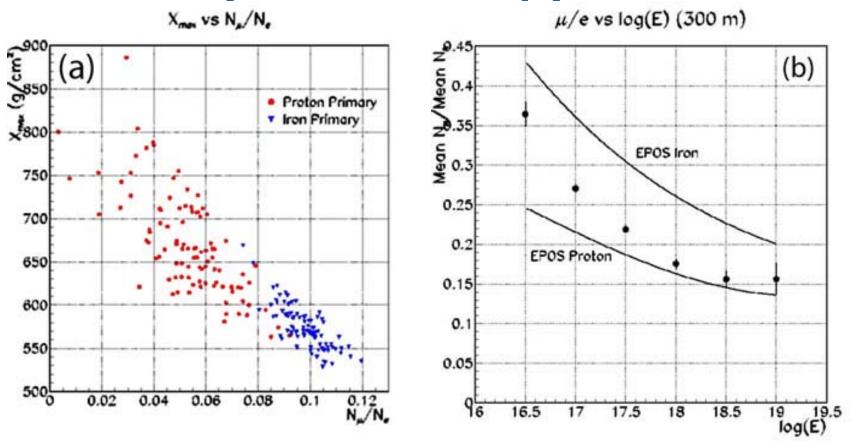
Aperture and Event Rate

- The TALE 6km stereo detector was designed for relatively flat aperture in the 10¹⁸-10¹⁹ eV decade
- The tower detector is specifically optimized to collect a maximum number of events in the 10¹⁷-10¹⁸ eV decade
- In both cases we will have long lever arms on either side of the spectral feature (ankle, 2nd knee) targeted





Composition via µ/e ratio



- Simulated 3-year TALE μ/e ratio ratio at 300m from core, at zenith angles of 25-35 degrees, for 10^{17} - 10^{18} eV decade: 50% protons at 10^{17} eV which increases to 80% protons at $10^{17.9}$ eV.
- The fit to the simulated data is shown with the curves for pure iron and proton CORSIKA w/ EPOS simulations.

TALE Summary

- In conjunction with TA, TALE provide overlapping (i.e. cross-calibrated) fluorescence (stereo or hybrid) coverage of cosmic rays from 10^{16.5}-10^{20.5} eV
 - Single energy scale for the energy spectrum over four decades
 - $-X_{\text{max}}$ composition measurement down to well below the "second knee" region
 - Anisotropy (with X_{max} tagging)
- A Non-Imaging CHerenkov Experiment (NICHE) is being proposed to extend energy coverage of the TA Observatory down to below the knee