

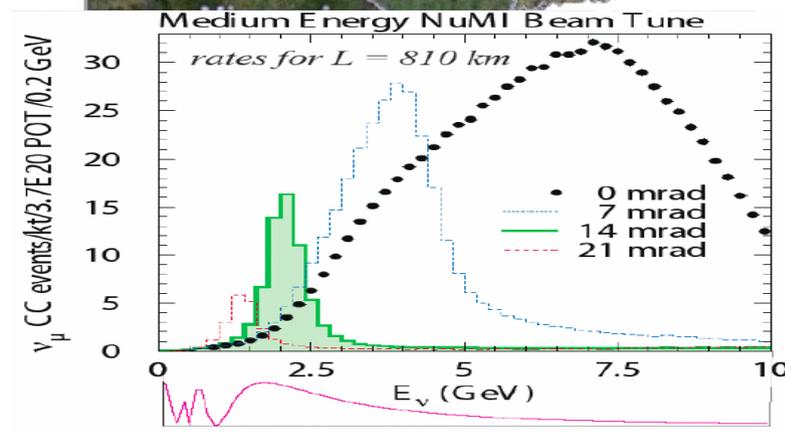


Cosmic Ray Background Rejection for First ν_e Appearance Analysis in NOvA

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NOvA Experiment

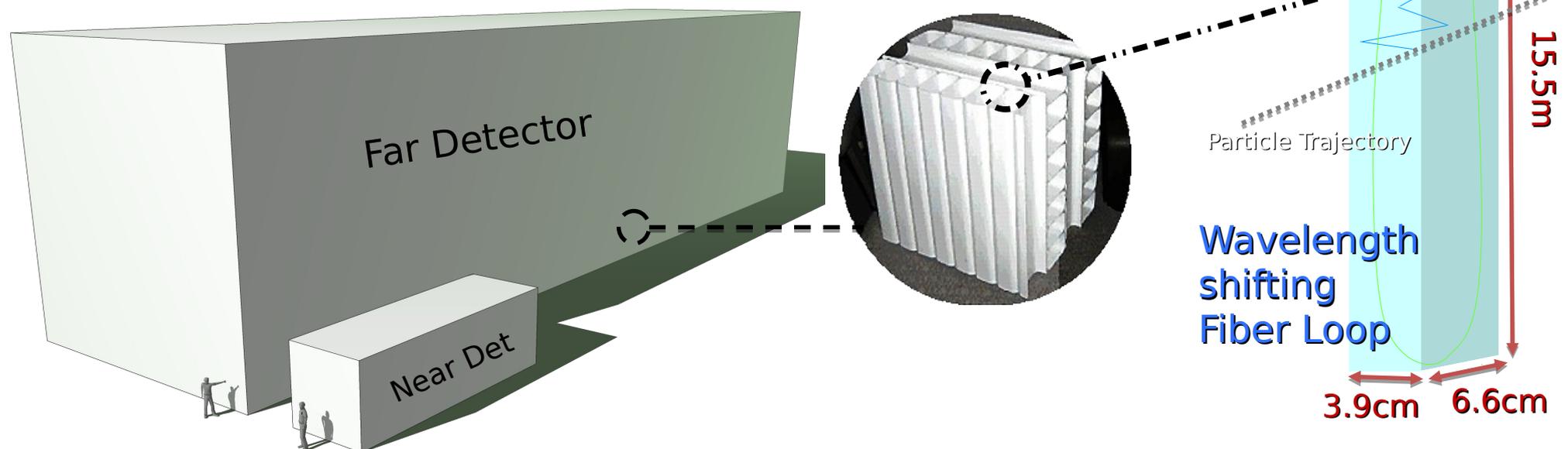
- NOvA is a **long-baseline accelerator neutrino experiment**, using neutrinos from **NUMI** muon neutrino beam at Fermilab.
- NOvA has two functionally identical detectors, along the neutrino beam direction **810 km** apart.
- The detectors are **14 mrad off beam axis**, where the muon neutrino beam narrowly peaks at 2 GeV, at which the majority of the beam events match the oscillation maximum.



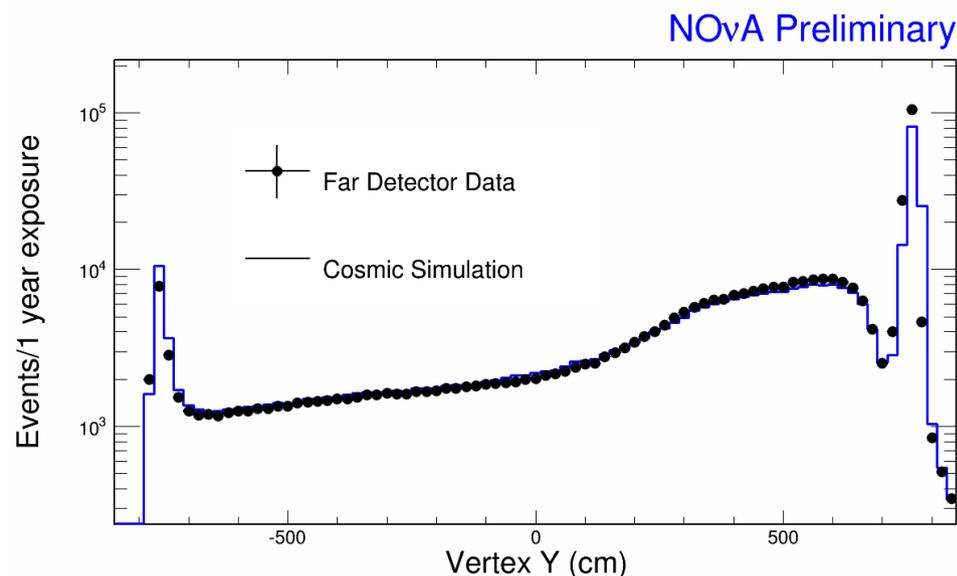
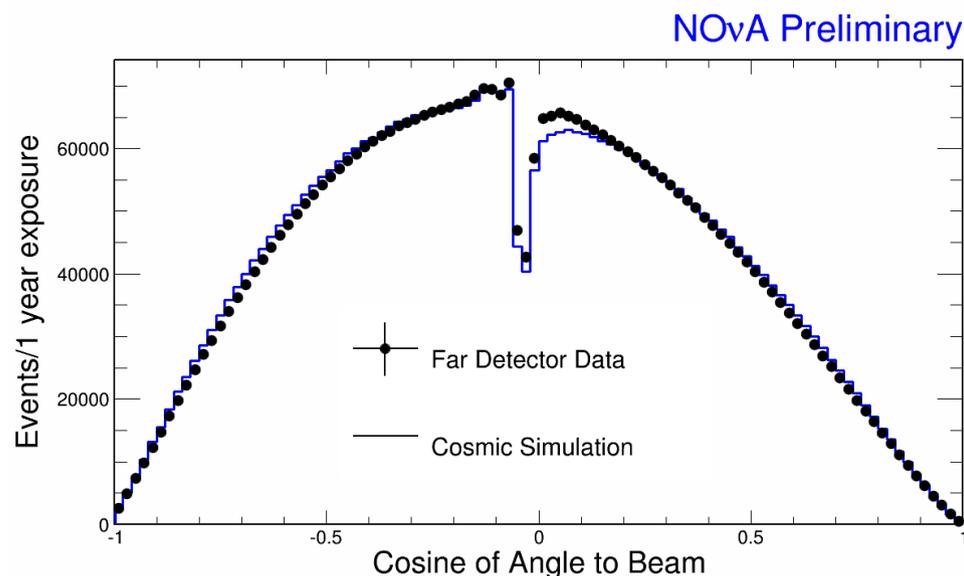
NOvA Detectors

- The NOvA experiment has two functionally identical, finely **segmented liquid scintillator** detectors.

	Location	Size
Near Detector	In front of NUMI beam. Underground	0.3 kton
Far Detector	810 km from the near detector, on the surface	14 kton

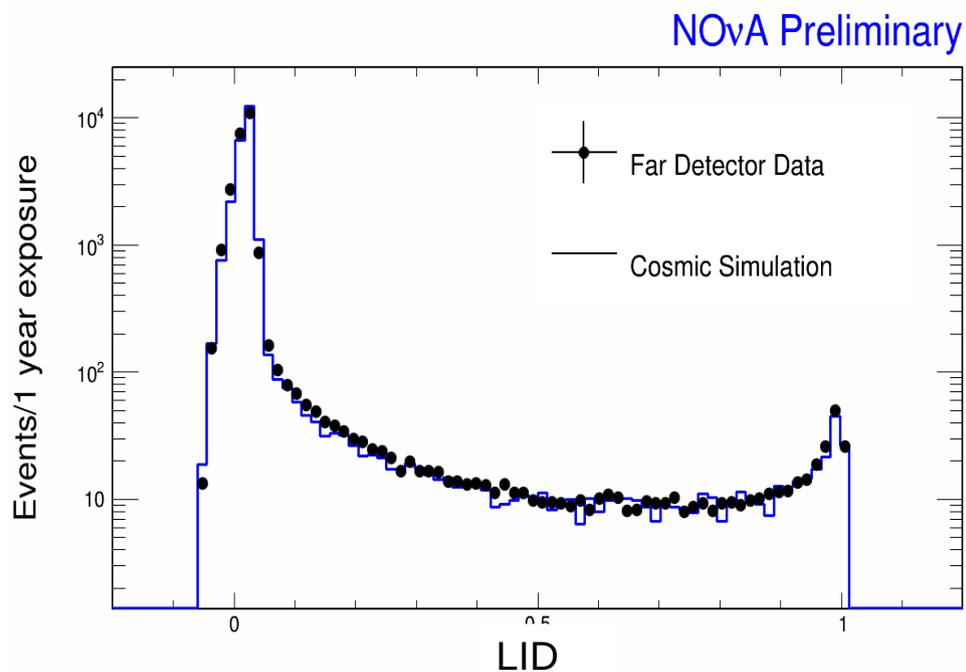
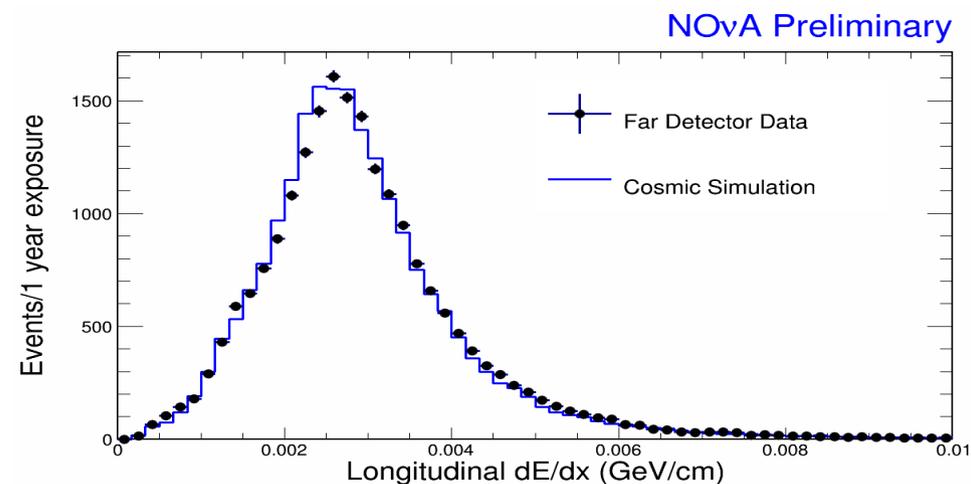
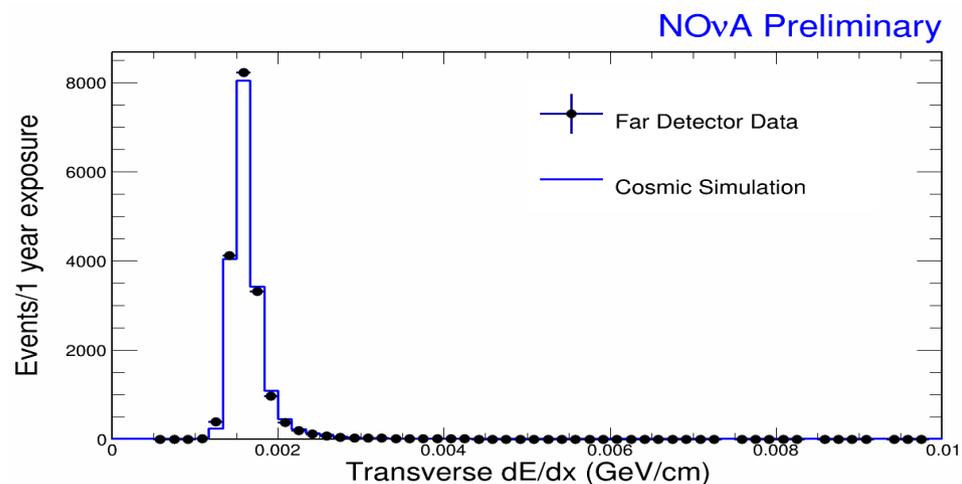


Data Versus MC Comparisons for Reconstruction Variables



- Data versus Monte Carlo comparisons specifically for two reconstruction variables,
 - Left: The cosine of the angle of reconstructed track with respect to the beam direction.
 - Right: The reconstructed vertex position in Y coordinate.
- The comparison of data to Monte Carlo simulations demonstrates that basic reconstructed variables are modeled well in our simulation.

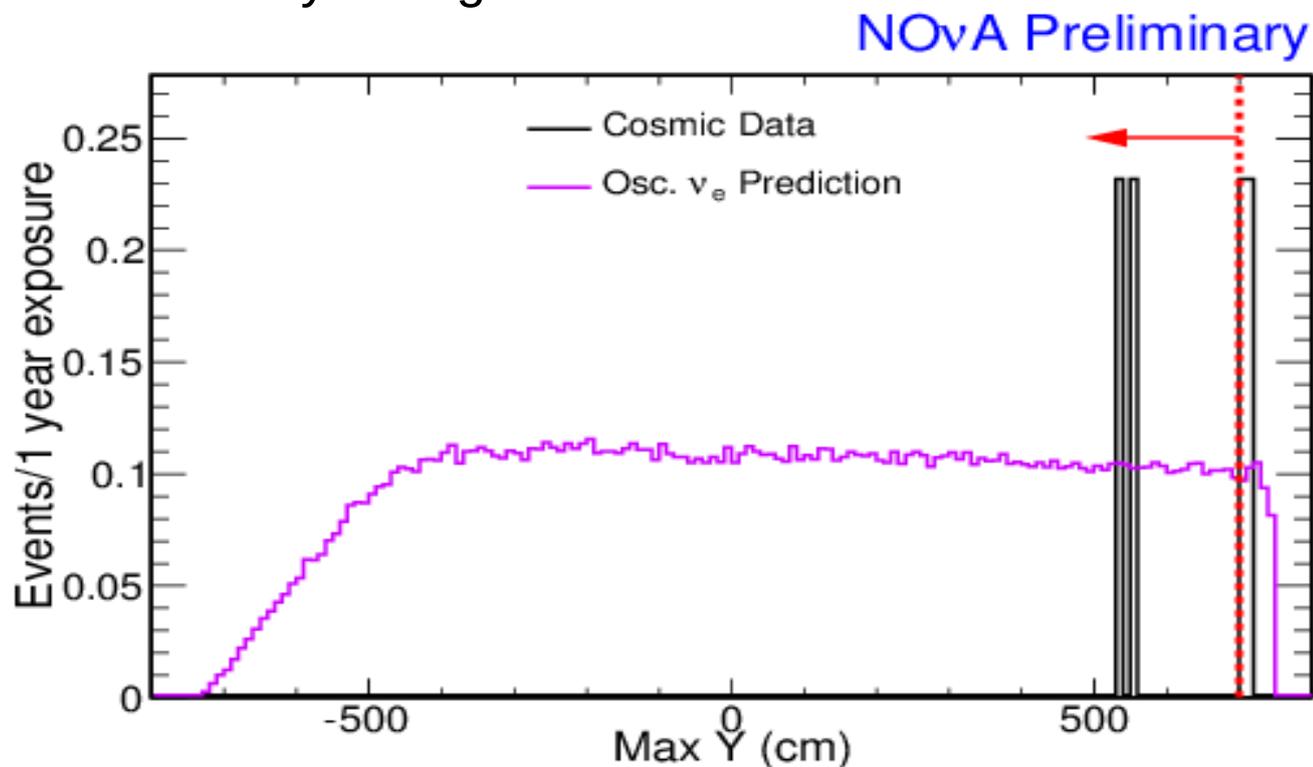
Data Versus MC Comparisons for Particle Identification Variables



- LID is one of the electron neutrino particle ID algorithms
 - It is based on shower shape and likelihood comparisons to various particle hypotheses.
- Cosmic ray data shows agreement with MC for these particle ID variables for events along the beam direction.

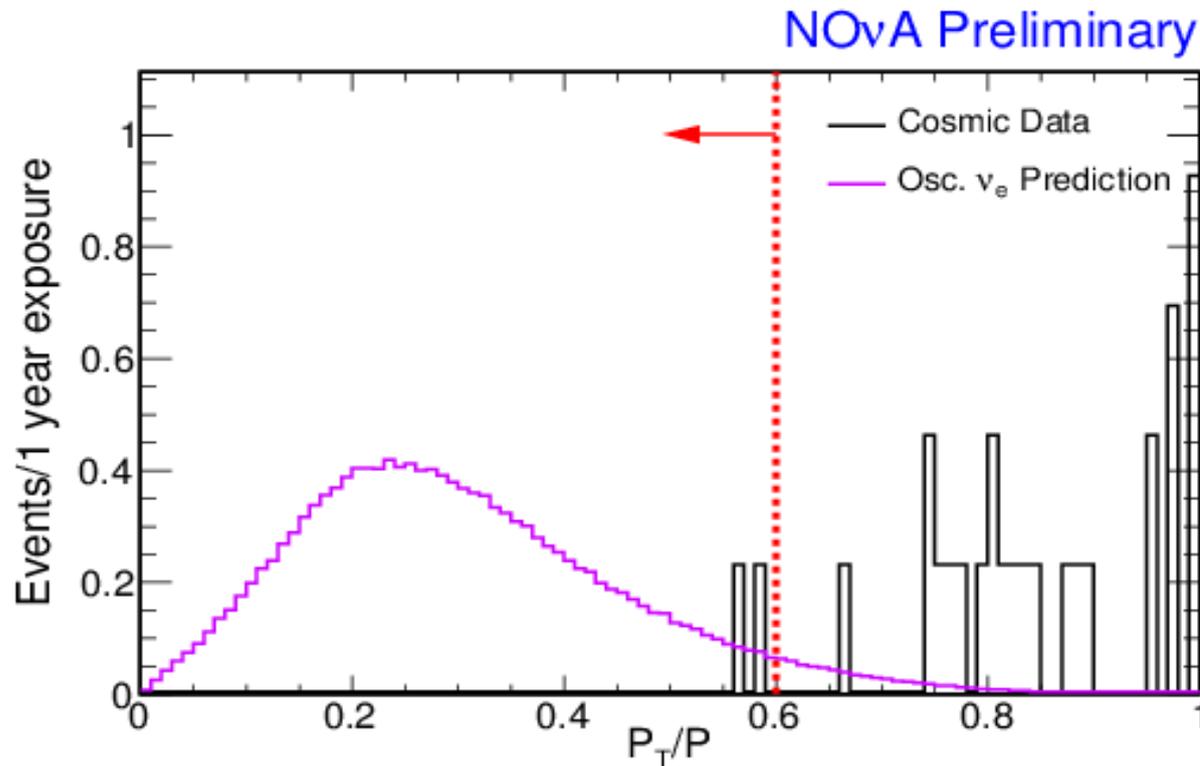
Cosmic Ray Background Rejection

- We have a set of selection cuts, including quality, containment, and particle ID cuts. After these cuts, we study a few additional variables to further separate signal from cosmic ray background events.



- One of the variables is Max Y, the larger value of either the start or stop Y position of the most energetic shower.
- Cut on this variable rejects cosmic background entering from top of the detector.

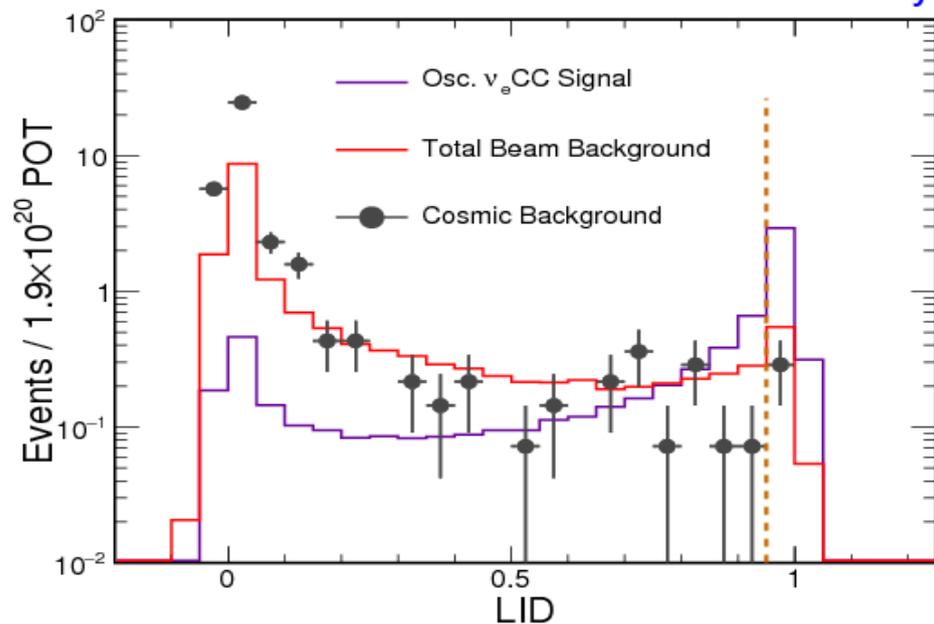
Cosmic Ray Background Rejection



- Another variable is P_T/P , the ratio of event transverse momentum (P_T) by event total momentum (P).
- The cut rejects predominately vertical cosmic rays with large P_T/P and retains the neutrinos that aligns with beam direction with small P_T/P .
- This cut also has a power in rejecting neutral current background, some of which has large P_T/P values.

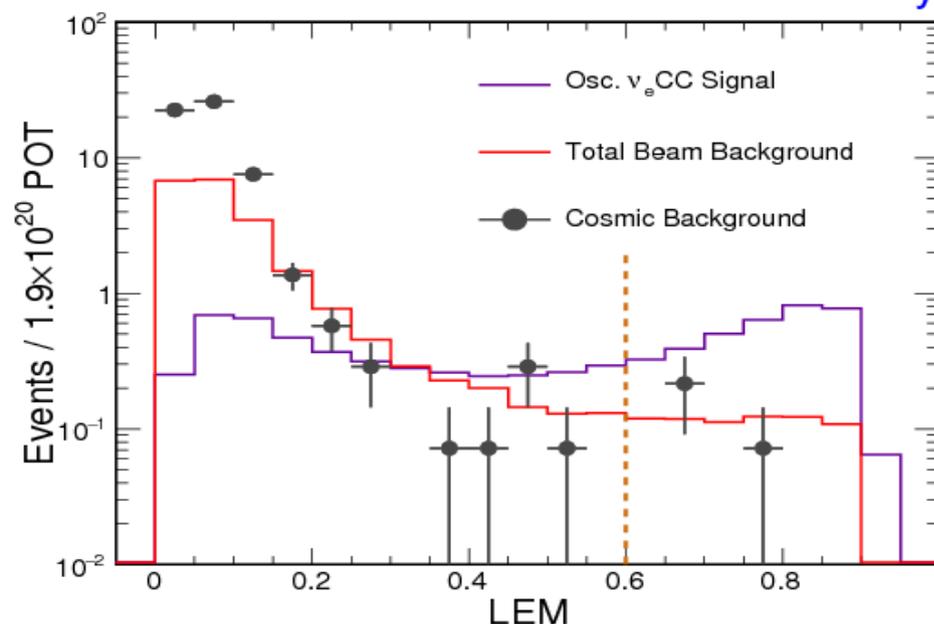
Electron Neutrino Particle Identifications

NOvA Preliminary



Based on shower shape likelihood assuming different particle hypotheses.

NOvA Preliminary



Matching events to the signal and beam background events in the library

- Two independent electron neutrino identification algorithms for the first ν_e appearance analysis to select electron neutrino from other event topologies, including cosmic ray background.

Cosmic Ray Background Prediction

	Cosmic background, LID	Cosmic background, LEM
No cut	1.49E+07	1.49E+07
Containment & quality	638325	967101
Cosmic rejection	5409.79	5791.31
Nue selection	0.29	0.29

- After all the quality, containment, cosmic rejection and ν_e selection cuts, we achieve **50 million to 1** cosmic background rejection.
- Based on the out of time numi data, we predict **0.29** cosmic background events in the two particle ID regions, LID and LEM, for 96 sec of livetime, the exposure for the first ν_e analysis sample.

Far Detector Event Prediction

	Osc. ν_e CC	Total bkg.	ν_μ CC	NC	Beam ν_e CC	Cosmic bkg
LID	3.25	1.02	0.05	0.32	0.33	0.29
LEM	3.48	1.14	0.05	0.41	0.36	0.29

- The neutrino signal and beam background event counts are extrapolated from the near detector data and scaled to $1.9e20$ POT, POT of the first analysis sample.
 - The oscillation weight is calculated without matter effect and with $\delta_{CP} = 0$ and $\sin^2 2\theta_{13} = 0.095$.
- **We have a method of effectively rejecting cosmic ray background to a level equivalent to other beam background components.**