

Lepton Flavor Violation and SUSY Dark Matter




UNIVERSITY OF MINNESOTA

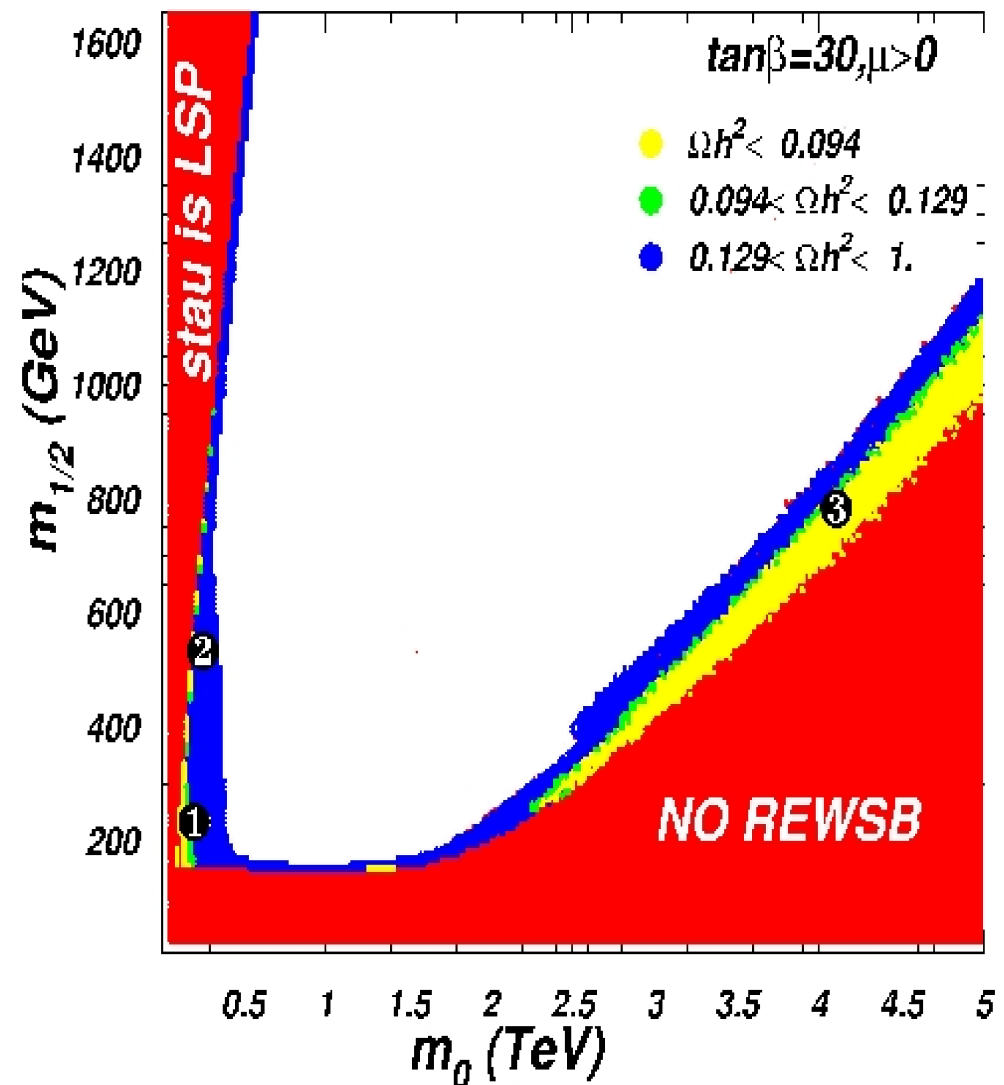
Azar Mustafayev

in collaboration with V.Barger and D.Marfatia PLB'08 and PRD'09

Dark Matter and mSUGRA


- WMAP measured relic density

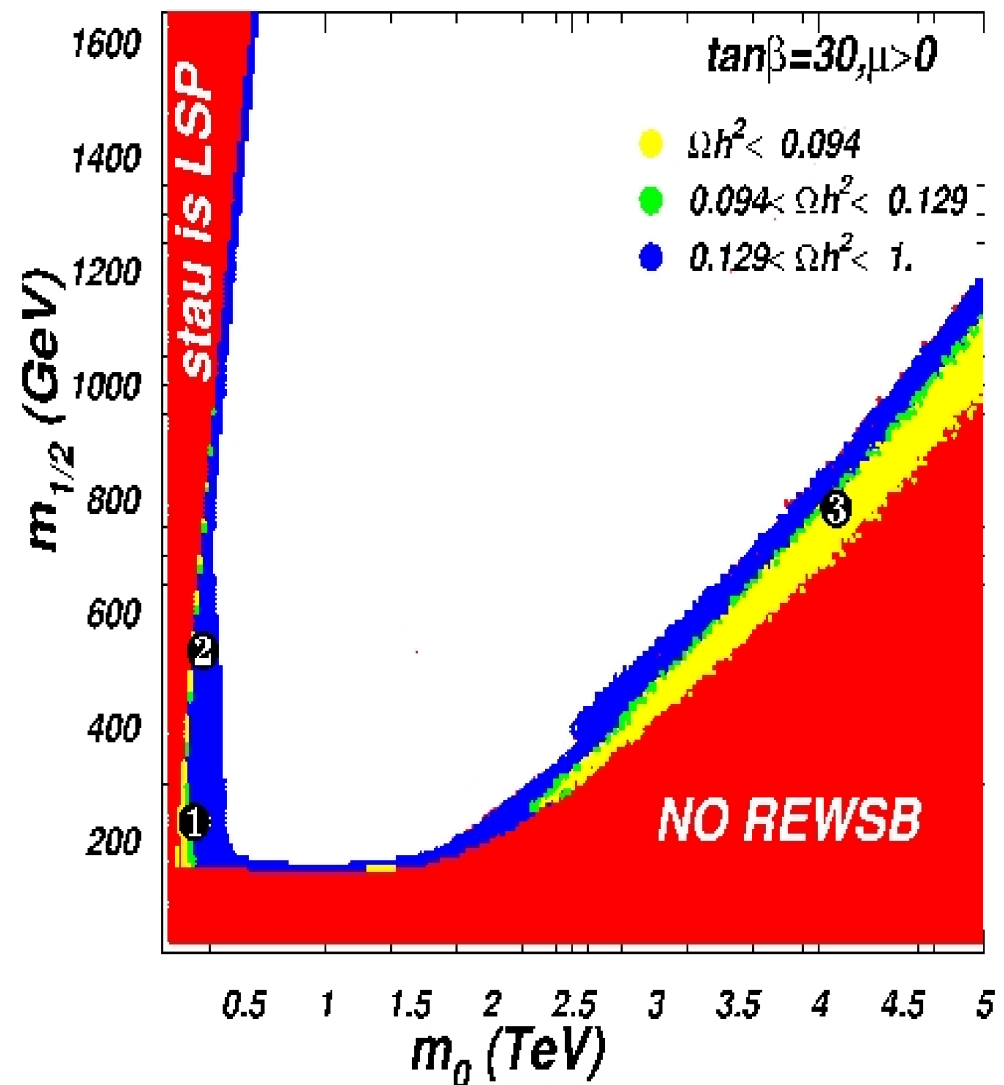
$$\Omega_{CDM}h^2 = 0.1120^{+0.0074}_{-0.0076}$$
- Lightest neutralino \tilde{Z}_1 is good DM candidate
- Most of para space ruled out due to too high RD $\Omega_{\tilde{Z}_1}h^2 \gg 0.1$
- mSUGRA para space:
 $m_0, m_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$
- Several narrow RD-allowed regions in mSUGRA
 SPS benchmark points





Dark Matter and mSUGRA

- WMAP measured relic density

$$\Omega_{CDM}h^2 = 0.1120^{+0.0074}_{-0.0076}$$
- Lightest neutralino \tilde{Z}_1 is good DM candidate
- Most of para space ruled out due to too high RD $\Omega_{\tilde{Z}_1}h^2 \gg 0.1$
- mSUGRA para space:
 $m_0, m_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$
- Several narrow RD-allowed regions in mSUGRA
 SPS benchmark points
- **Neutrinos are massless!**



SUSY seesaw

- Observed neutrino oscillations  massive neutrinos
- Mass scale $\sum m_\nu \lesssim 1 \text{ eV}$
- Seesaw mechanism (type-I):
$$\mathcal{M}_\nu = -\mathbf{f}_\nu M_N^{-1} \mathbf{f}_\nu^T v_u^2$$
- Right-handed neutrino masses $M_N \sim (10^{10} - 10^{16}) \text{ GeV}$
- Neutrino Yukawas from $SO(10)$
 - ▶ $\mathbf{f}_\nu = \mathbf{f}_u$ if higgses in **10**
 - ▶ $\mathbf{f}_\nu = 3\mathbf{f}_u$ if higgses in **126**
- Common approach:
SPS points + RHNs  LFV predictions in RD-allowed regions

SUSY seesaw

- Observed neutrino oscillations  massive neutrinos

- Mass scale $\sum m_\nu \lesssim 1 \text{ eV}$

- Seesaw mechanism (type-I):

$$\mathcal{M}_\nu = -\mathbf{f}_\nu M_N^{-1} \mathbf{f}_\nu^T v_u^2$$

- Right-handed neutrino masses $M_N \sim (10^{10} - 10^{16}) \text{ GeV}$

- Neutrino Yukawas from $SO(10)$

▶ $\mathbf{f}_\nu = \mathbf{f}_u$ if higgses in **10**

▶ $\mathbf{f}_\nu = 3\mathbf{f}_u$ if higgses in **126**

- Common approach:

SPS points + RHNs  LFV predictions in RD-allowed regions 

RGE analysis

- Neutrino Yukawas appear in RGEs above seesaw scale

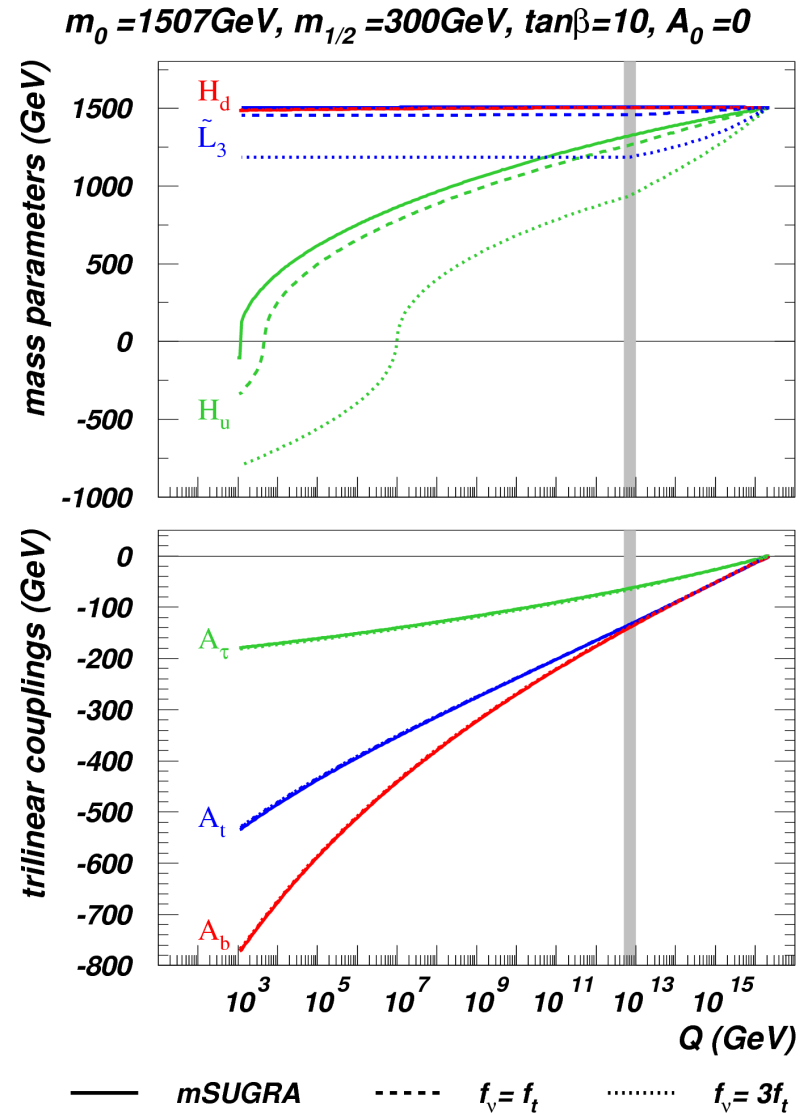
$$\frac{dm_{L_3}^2}{dt} \propto f_\tau^2 X_\tau + f_\nu^2 X_\nu$$

$$\frac{dm_{H_u}^2}{dt} \propto 3f_t^2 X_t + f_\nu^2 X_\nu$$

$$\frac{dA_\tau}{dt} \propto 3f_b^2 A_b + f_\nu^2 A_\nu$$

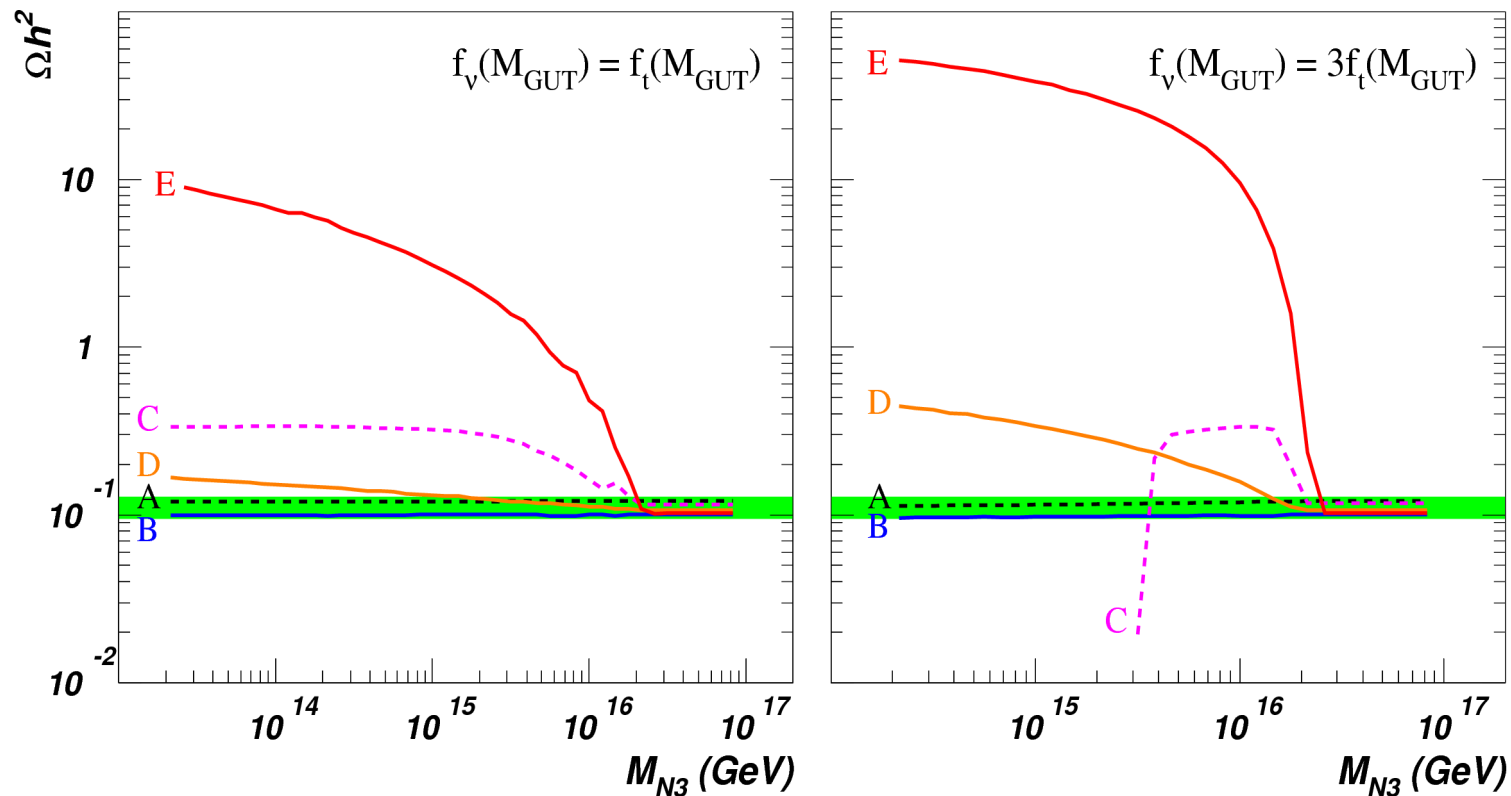
$$\frac{dA_t}{dt} \propto 3f_t^2 A_t + f_\nu^2 A_\nu$$

- EWSB: $\mu^2 \simeq -m_{H_u}^2$
- CP-odd Higgs mass $m_A^2 \simeq m_{H_d}^2 - m_{H_u}^2$
- $m_{L_3}^2 \rightarrow$ stau masses
- A-terms \rightarrow L-R sfermion mixing



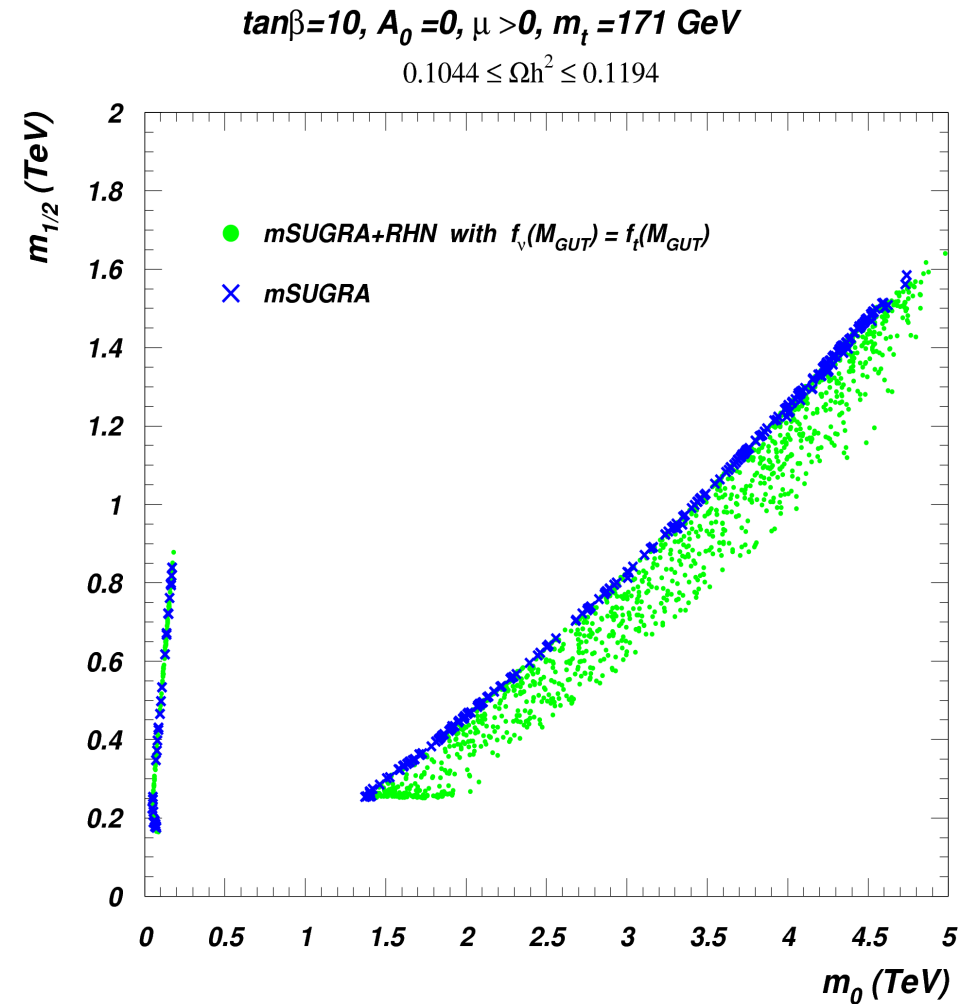
RHN effect on RD

- mSUGRA+RHN model parameters: m_0 , $m_{1/2}$, A_0 , $\tan \beta$, $\text{sign}(\mu)$, M_{N3}
- Benchmark points:
A=bulk, B=stau co-annih., C=stop co-annih., D=Higgs funnel, E=HB/FP
- New co-annihilation with $\tilde{\nu}_\tau$ in case C



RD in mSUGRA+RHN

- Random scan in $(m_0, m_{1/2}, M_{N3})$
- Neutrino Yukawa effects can be compensated by SSB parameters
→ shifts of RD-allowed regions
- Expect effects on DM rates and/or collider signatures and LFV rates



LFV in mSUGRA-seesaw

- Radiative lepton decay (in mass-insertion approx.)

$$BR(l_i \rightarrow l_j \gamma) \propto \frac{\alpha^3}{G_F^2 m_s^8} |(m_L^2)_{i \neq j}|^2 \tan^2 \beta$$

- Induced by neutrino Yukawa

$$(m_L^2)_{i \neq j} \simeq -\frac{1}{8\pi^2} (3m_0^2 + A_0^2) \sum_k (\mathbf{f}_\nu^T)_{ik} (\mathbf{f}_\nu^*)_{kj} \log \frac{M_{GUT}}{M_{N_k}}$$

- Sensitive to SSB values

$$m_s^8 \simeq 0.5 m_0^2 m_{1/2}^2 \left(m_0^2 + 0.6 m_{1/2}^2 \right)^2$$

- Similar expressions for $l_i \rightarrow 3l_j$ decays and $\mu \rightarrow e$ conversion in Nuclei
- Consider two extremal cases for each unification ($R_{\nu u} = 1, 3$)
 - Small mixing $\mathbf{f}_\nu = R_{\nu u} \mathbf{f}_u$
 - Large mixing $\mathbf{f}_\nu = R_{\nu u} \mathbf{V}_{TBM} \mathbf{f}_u^{diag}$

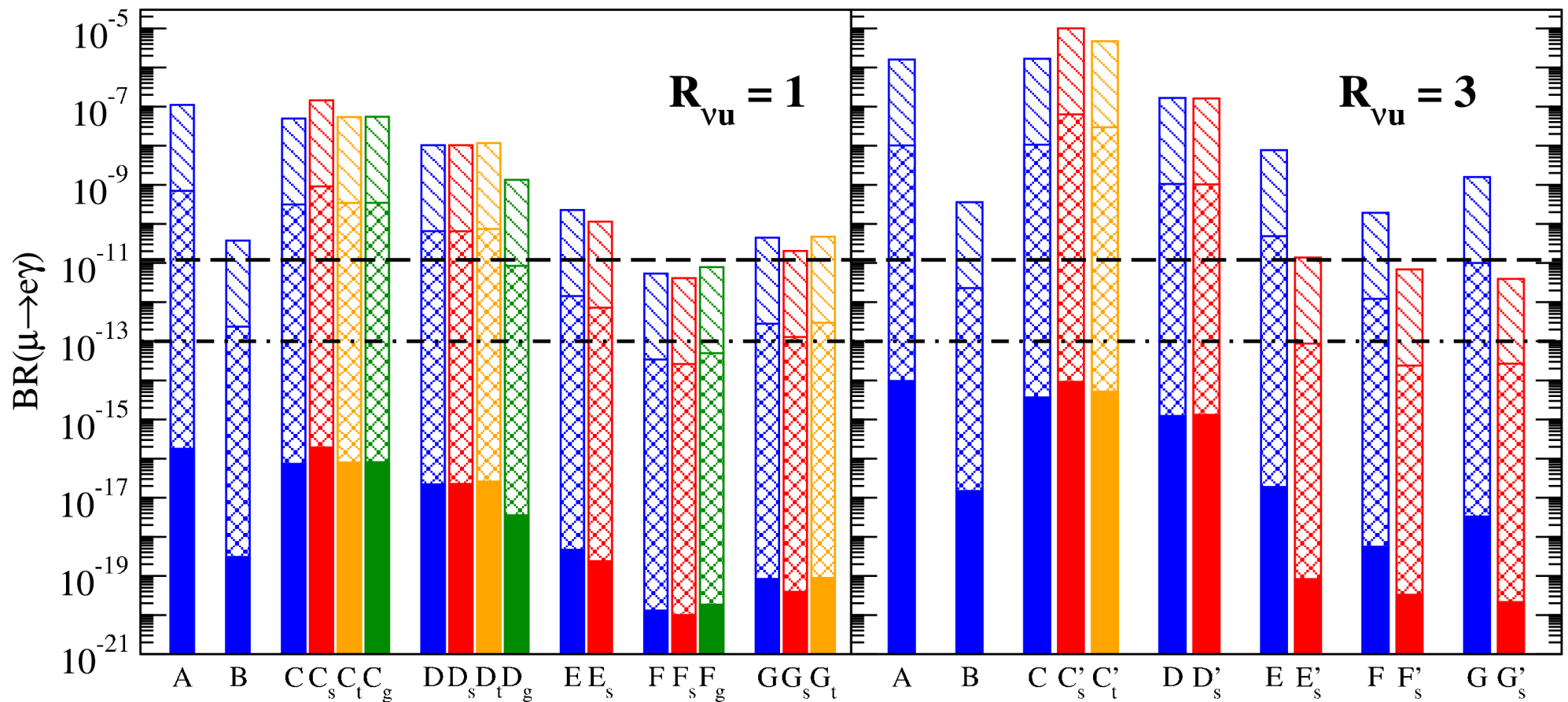
Large mixing case

Correct RD by adjusting one parameter:

● = m_0

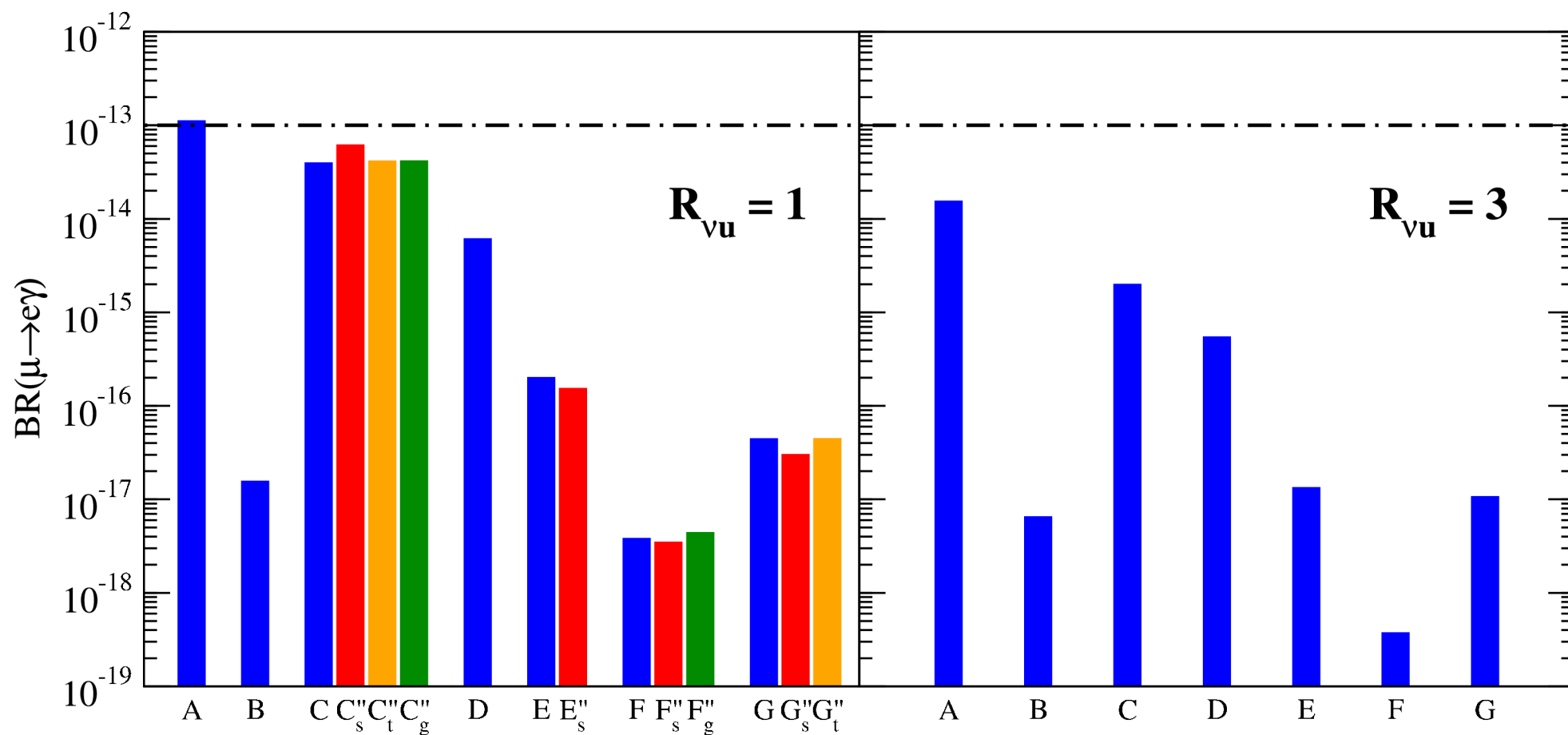
● = A_0

● = $m_{1/2}$



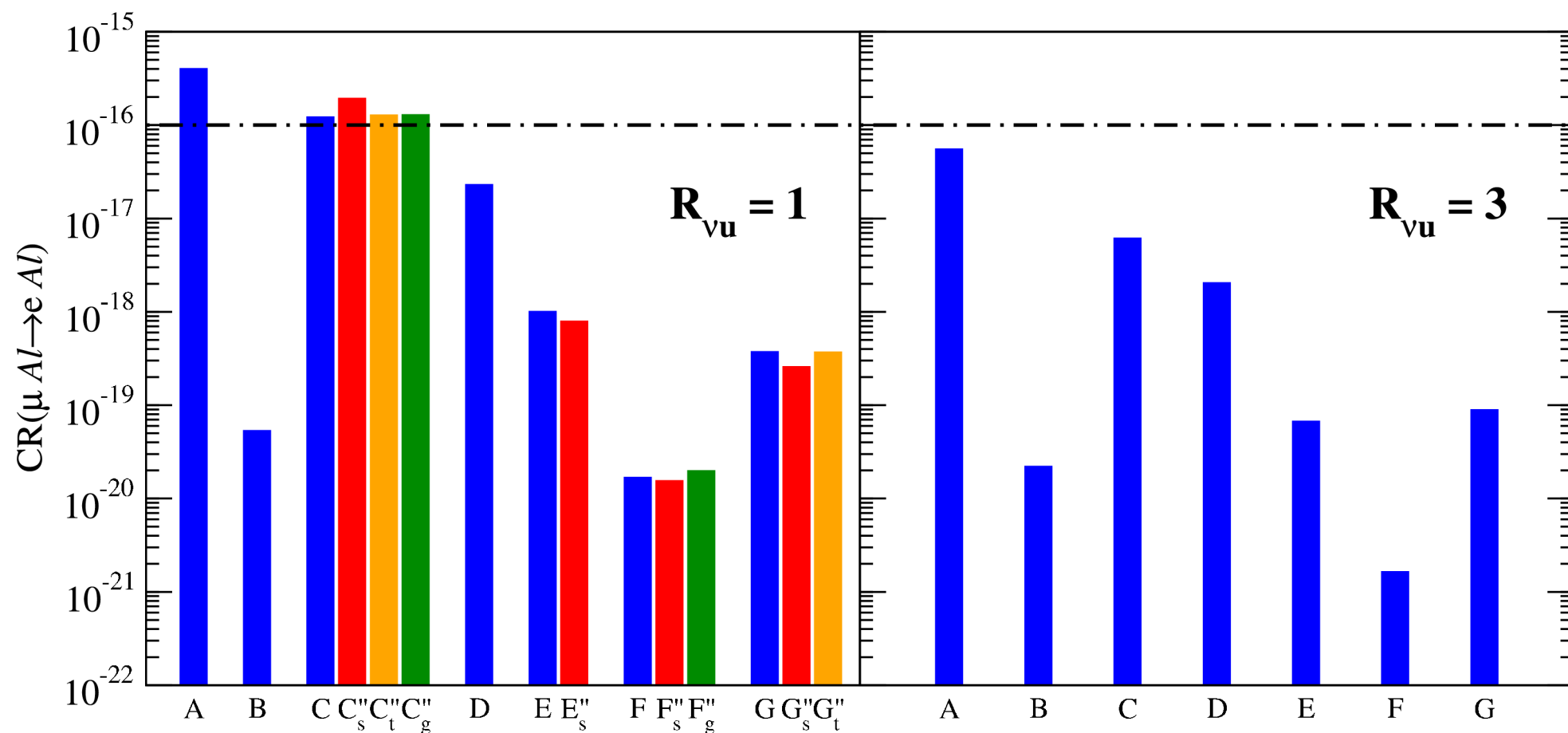
➤ Rates change from factor a few up to about order of magnitude

Small mixing case




- Completely untouched by current experimental bounds
- Larger yukawa produces order of magnitude *smaller* rates!

Small mixing case



- Completely untouched by current experimental bounds
- Larger yukawa produces order of magnitude *smaller* rates!
- Only Mu2e (or PRIME) will be able to test

Summary

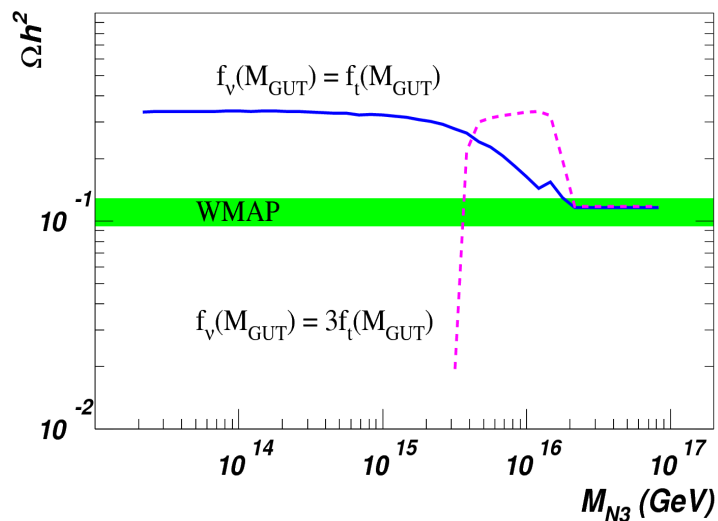
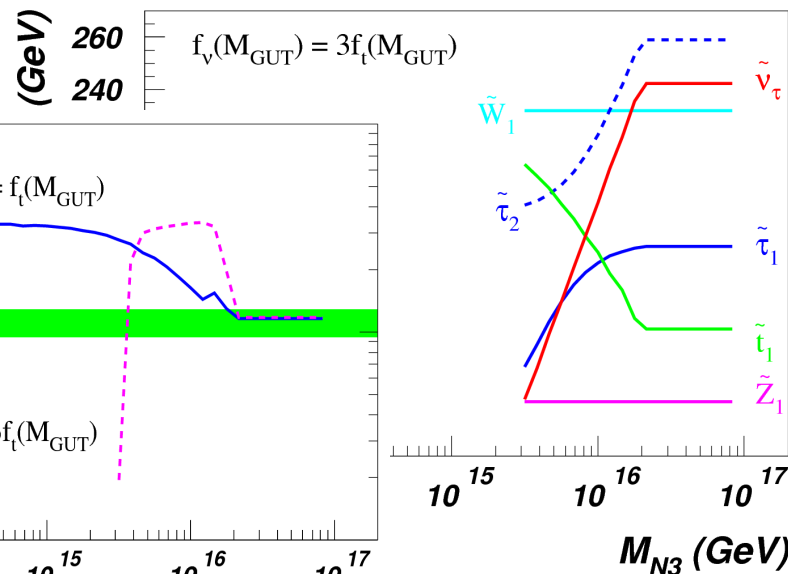
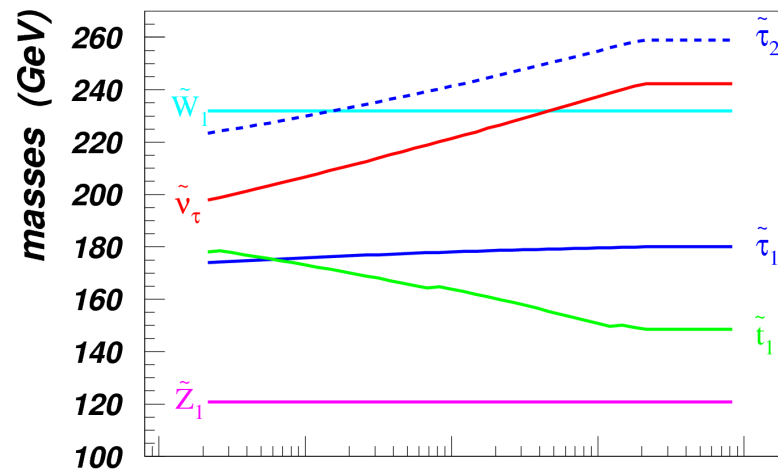
- Neutrino Yukawa coupling can **significantly affect** neutralino DM
- Most prominent when SSB slepton mass and/or trilinears large
 - Stop coannihilation, A-funnel and HB/FP regions of mSUGRA
- Effects in stau coannihilation and the bulk are small *only* because in mSUGRA $\tilde{\tau}_1 \simeq \tilde{\tau}_R$
- Neutrino-neutralino interplay can change LFV predictions for $R_{\nu u} = 1$ by factor ~ 5
- Large unification parameter $R_{\nu u} = 3$ is not excluded by LFVs and for small mixing produce order of magnitude *smaller* rates
 need Mu2e to probe

Backup slides

Stop coannihilation

- Small $m_0, m_{1/2}$ and large A_0
- Light $\tilde{t}_1 \implies$ rapid $\tilde{Z}_1 \tilde{t}_1 \rightarrow X_{SM}$
- RHN effects:
 - lighter $\tilde{\nu}_\tau, \tilde{\tau}_2 \simeq \tilde{\tau}_L$
 - smaller L-R mixing \longrightarrow heavier \tilde{t}_1
 - $\tilde{\tau}_1$ becomes $\tilde{\tau}_L$
- Coannih. with \tilde{t}_1 closes
- New coannih. channel with $\tilde{\nu}_\tau$

$m_0=150\text{GeV}, m_{1/2}=300\text{GeV}, \tan\beta=5, A_0=-1091\text{GeV}$



Stop coannihilation

- Small $m_0, m_{1/2}$ and large A_0
- Light $\tilde{t}_1 \implies$ rapid $\tilde{Z}_1 \tilde{t}_1 \rightarrow X_{SM}$
- RHN effects:
 - lighter $\tilde{\nu}_\tau, \tilde{\tau}_2 \simeq \tilde{\tau}_L$
 - smaller L-R mixing \longrightarrow heavier \tilde{t}_1
 - $\tilde{\tau}_1$ becomes $\tilde{\tau}_L$
- Coannih. with \tilde{t}_1 closes
- **New coannih. channel with $\tilde{\nu}_\tau$**
- Positron, muon and gamma rates are up by **100-150%** (**300-700%**)

