

Managed by Fermi Research Alliance, LLC for the U.S. Department of Energy Office of Science

Quantum Sensors - Fundamental Experiments and Infrastructure

William Wester
Physics Advisory Meeting
8 Jul 2017



Managed by Fermi Research Alliance, LLC for the U.S. Department of Energy Office of Science

Quantum Sensors - Fundamental Experiments and Infrastructure

William Wester
Physics Advisory Meeting
8 Jul 2017

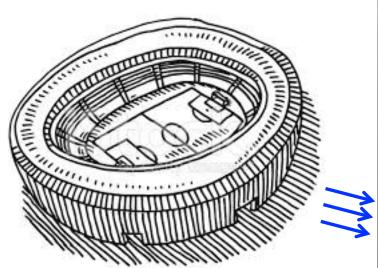
Thanks to A. Chou, A. Sonnenschein, D. Bowring, A. Moretti, A. Tollestrup, A. Grasselino, A. Romanenko, B. Benson, A. Anderson, S. Rahlin, J. Tiffenberg, J. Estrada, D. Kubic, S. Chattopadhyay, plus all the associated University colleagues and others not at Fermilab.

Axions first! ADMX: axion dark matter experiment

- For the first time, an experiment has sensitivity to the QCD axion that solves the strong-CP problem and could account for the dark matter in the universe.
- Over 20 years of R&D, but now a 1) high magnetic field surrounds a 2) high-Q and tunable copper cavity with 3) superconducting quantum interference device (SQUID) readout at 4) sub-kelvin temperatures. This is ADMX in the context of the DOE gen2 dark matter program.
- Sensitivity to possible 10⁻²³ W of power from dark matter axions converting into detectable single microwave photons.



ADMX – direct detection of axionic dark matter





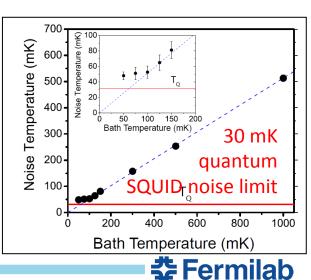
Tune the μ -wave cavity to the frequency corresponding to the kinetic energy of the axion.

Scan in frequency for sensitivity as a function of the mass of the axion.

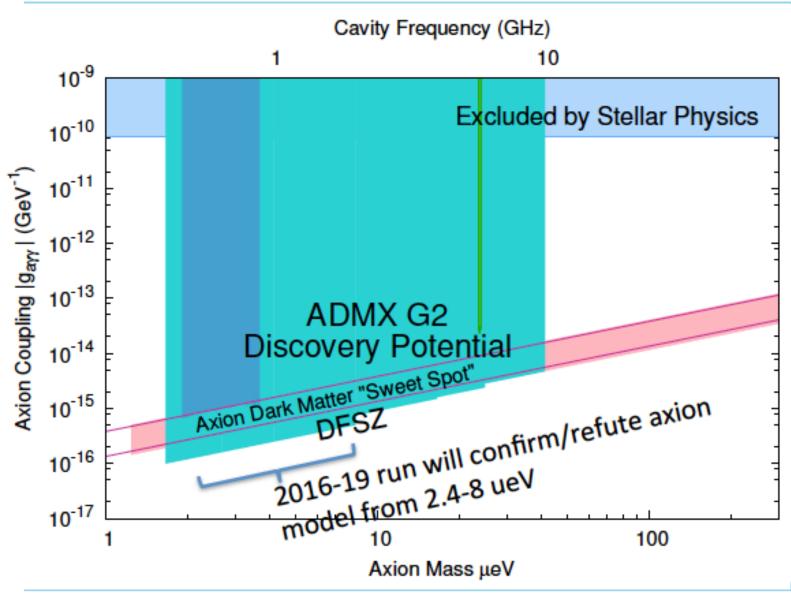
 $\lambda = h/p \sim 100m$ coherent cloud of an axion field making up the dark matter halo, passing through the ADMX apparatus.

$$\frac{df}{dt} \approx 1.68 \,\text{GHz/year} \, \left(\frac{g_{\gamma}}{0.36}\right)^4 \left(\frac{f}{1 \,\text{GHz}}\right)^2 \left(\frac{\rho_0}{0.45 \,\text{GeV/cc}}\right)^2.$$

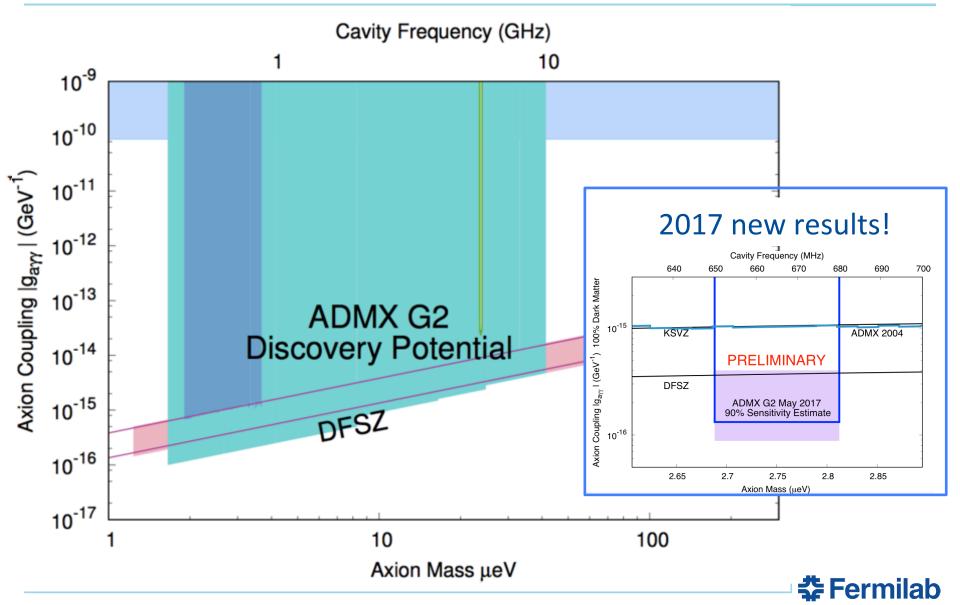
$$\left(\frac{5}{SNR}\right)^2 \left(\frac{B_0}{8 \,\text{T}}\right)^4 \left(\frac{V}{100 \text{l}}\right)^2 \left(\frac{Q_L}{10^5}\right) \left(\frac{C_{010}}{0.5}\right)^2 \left(\frac{0.2 \, K}{T_{sys}}\right)^2$$



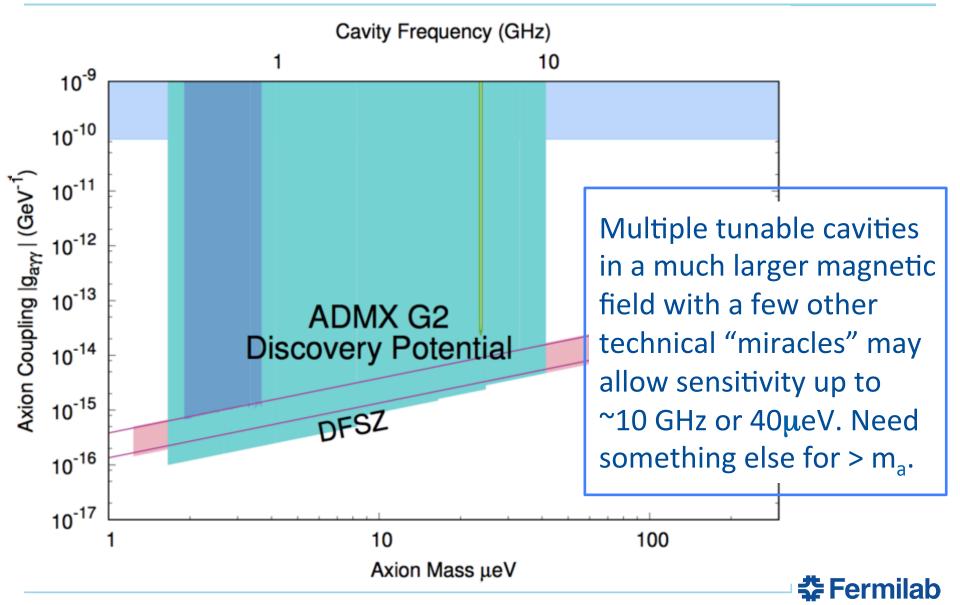
ADMX - gen2 discovery potential



ADMX - Initial results for gen2 project

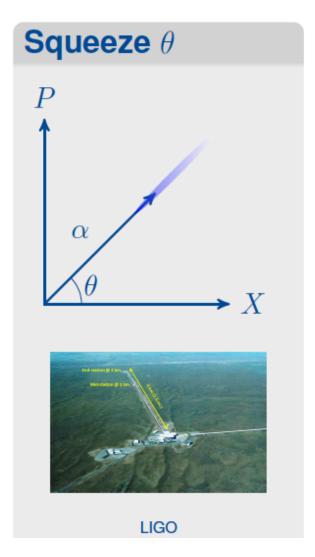


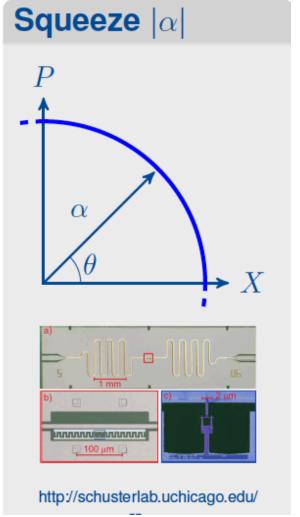
ADMX – possible to use multiple cavities to ~ 10 GHz



Quantum "Miracles"

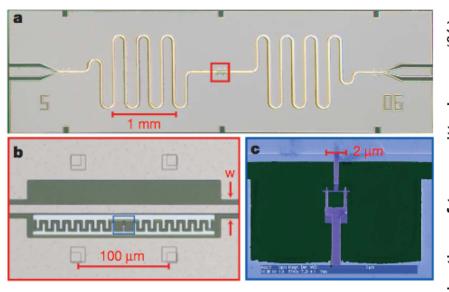
Coherent State α $\Delta P \Delta X \gtrsim \hbar/2$

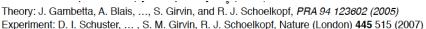


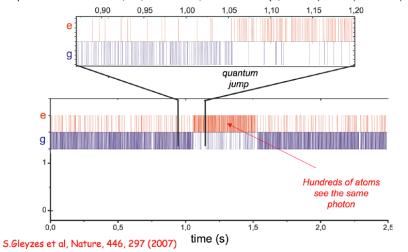


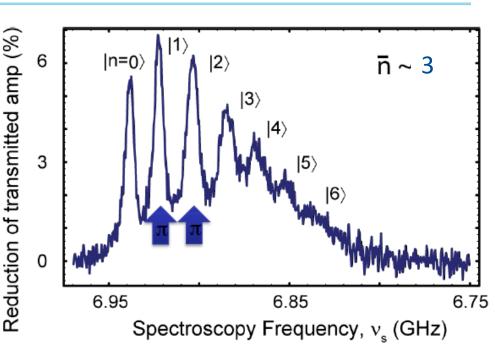


Qubits – non-destructive single microwave photon detector







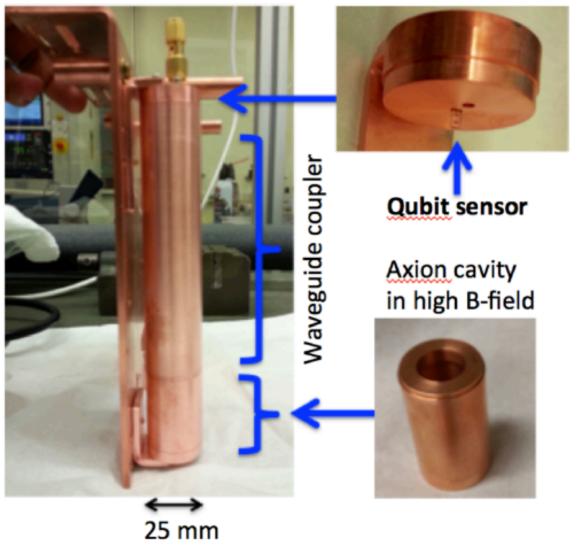


Aaron Chou's LDRD: detect single microwave photons in a high magnetic field using qubits with particular emphasis on improving the bit error rate -> enable high mass axion search.

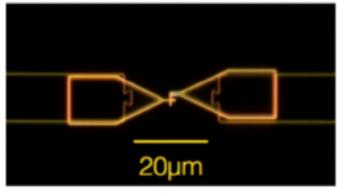


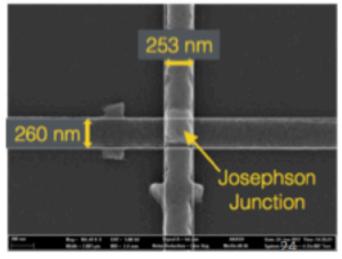
3D Qubit-based detector prototype

AC, Akash Dixit, D.Schuster (U.Chicago)







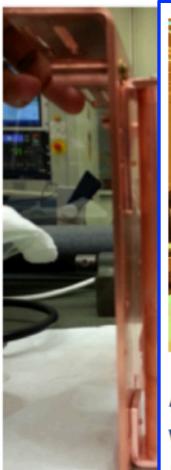


R&D supported by Heising-Simons Foundation.

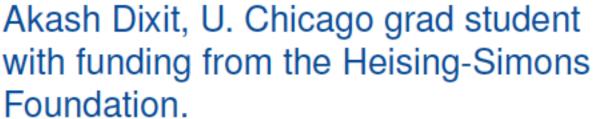
Aaron S. Chou, Aspen 3/21/17

3D Qubit-based detector prototype

AC, Akash Dixit, D.Schuster (U.Chicago)







Josephson

Junction

25 mm

R&D supported by Heising-Simons Foundation.

Aaron S. Chou, Aspen 3/21/17

Measurement in high field to be done with LDRD support

 Dilution refrigerator for sub-K operation with associated high field magnet. Plan to locate in the SiDet facility.



SiDet – technical area used for silicon particle trackers, CCDs, CMB sensors, etc.





Newly obtained test cryostats and lab space for LDRD project studying fine tuning of cavities using dielectrics with applied voltage.

(A. Sonnenschein)



Measurement in high field to be done with LDRD support

 Dilution refrigerator for sub-K operation with associated high field magnet. Plan to locate in the SiDet facility.



SiDet – technical area used for silicon particle trackers, CCDs, CMB sensors, etc.





2nd effort: Study qubit lifetime in high Q SRF cavities

We have begun R&D toward higher Q cavities for quantum computers, which

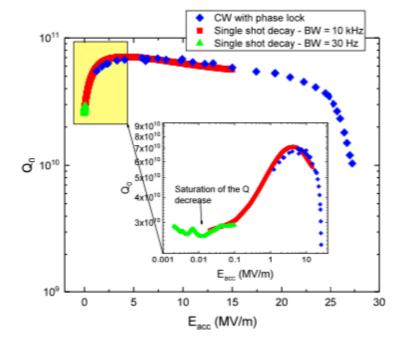
translates e.g. to longer coherence times

Supported by LDRD (A. Romanenko)

From PIP-2, LCLS-2 to Quantum Computing



Joe Lykken discussed this effort in his talk. Large volume dilution refrigerator leverages existing cavities



Nature of the Low Field Q Degradation in Superconducting Niobium Cavities

A. Romanenko[†] Fermi National Accelerator Laboratory, Batavia, IL 60510, USA

D. I. Schuster[‡]
The James Franck Institute and Department of Physics,
University of Chicago, Chicago, Illinois 60637, USA
(Dated: May 31, 2017)



Sub-kelvin research at Fermilab

| | What: | Target R&D | Unique Capability | Funding source |
|--|-----------------------------------|---|-----------------------------------|---|
| | ADR (50 mK) | MKIDs for optical cosmo. | SiDet focal plane packaging | KA25 |
| | ADR (50 mK), He-3 fridges (1K) | SPT, CMB-S4 | SiDet detector integration | LDRD |
| | Dilution fridge (10 mK) | Qubit-based photon detectors for axions | High field magnet* (>14 T) | LDRD |
| | Dilution fridge (10 mK) | Materials studies for high Q SRF cavities | SRF surface processing facility | LDRD |
| | Dilution fridge (10 mK) | Active veto for WIMP detectors, low threshold calorimetry | Underground** in NUMI (30 m.w.e.) | Northwestern, KA25 (install) KA23 (G3 DM) |



15

^{*}NHFML is good for quick sample testing, FNAL magnetic test stand is for device development.

^{**}Stanford underground test stand (Cabrera) no longer exists.

MKIDs: Microwave kinetic inductive detectors

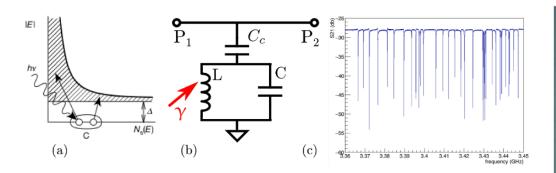
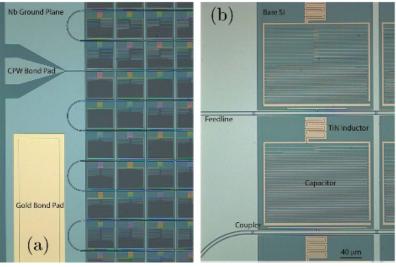


Figure 3 (a-b) When photons strike the inductor of an MKID, they break Cooper pairs and generate quasiparticles, changing the frequency of an LC circuit. (c) A measured microwave sweep across part of an existing optical MKID array, showing multiple frequency collisions and wide variation in the depth of resonances, both of which we propose to remedy.

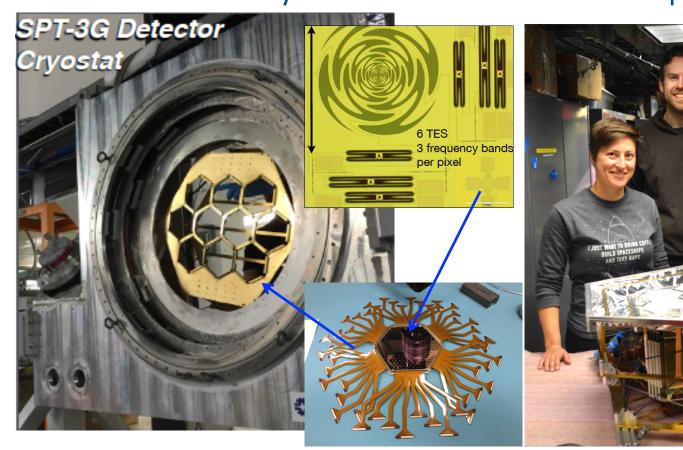


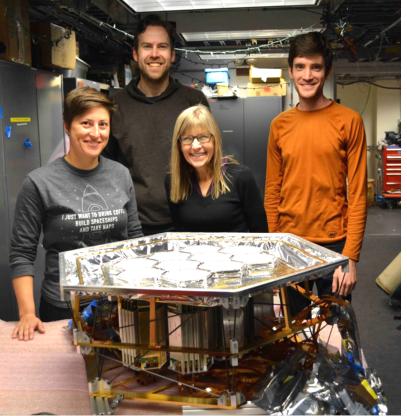
Single photon detection by MKIDs for optical / near-IR has been demonstrated, but continued R&D towards realizing a viable device -> goal towards using for cosmology where photons are collected and have an energy (wavelength) measured. Iteration on a next optimized devices supported through LDRD (J. Estrada).



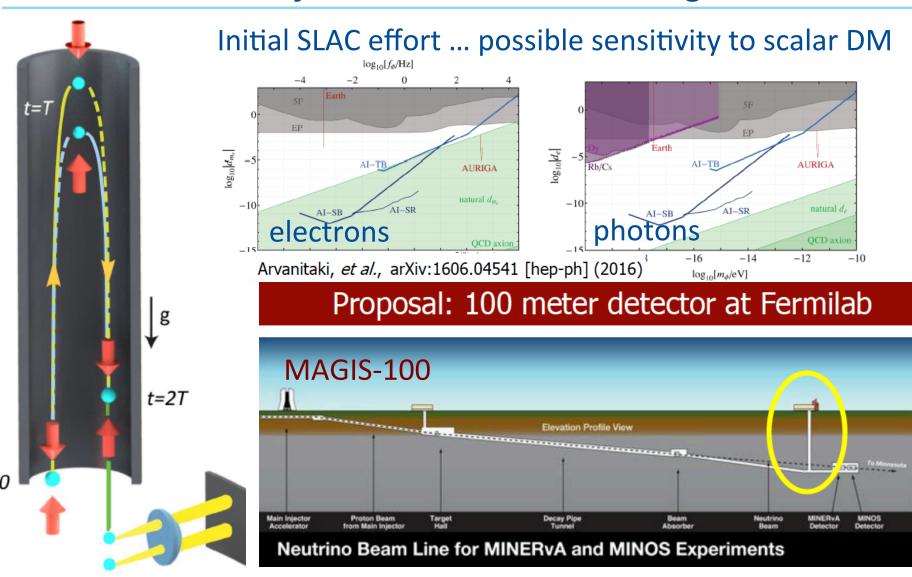
CMB R&D towards S4

For 3 years, B. Benson and his team have been working on R&D related to large scale focal plane assembly and testing for CMB. This effort lead to delivery of the S3 South Pole Telescope focal plane.



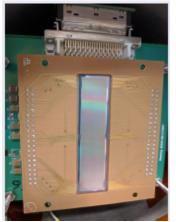


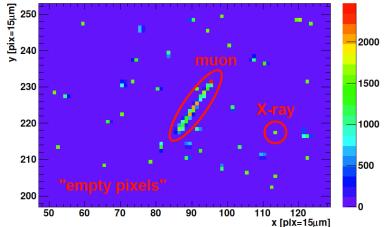
Atom Interferometry for new scalars – ultra light DM



SENSEI – Sub-Electron Noise SkipperCCD Exp. Instrument

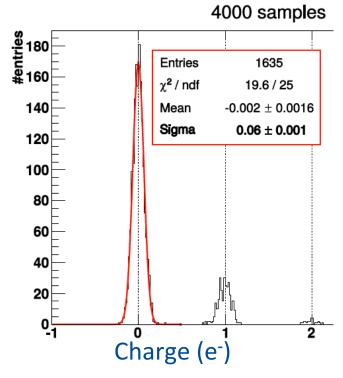
 New CCD device with novel architecture that allows multiple reads of the same pixel to reduce noise.
 (J. Tiffenberg)







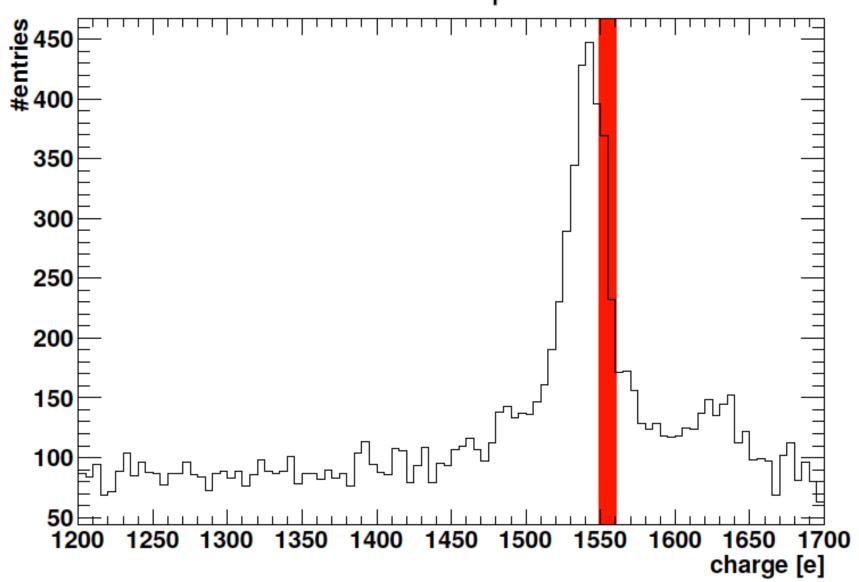
Skipper device and detector along with an exposure with 4000 reads/pixel



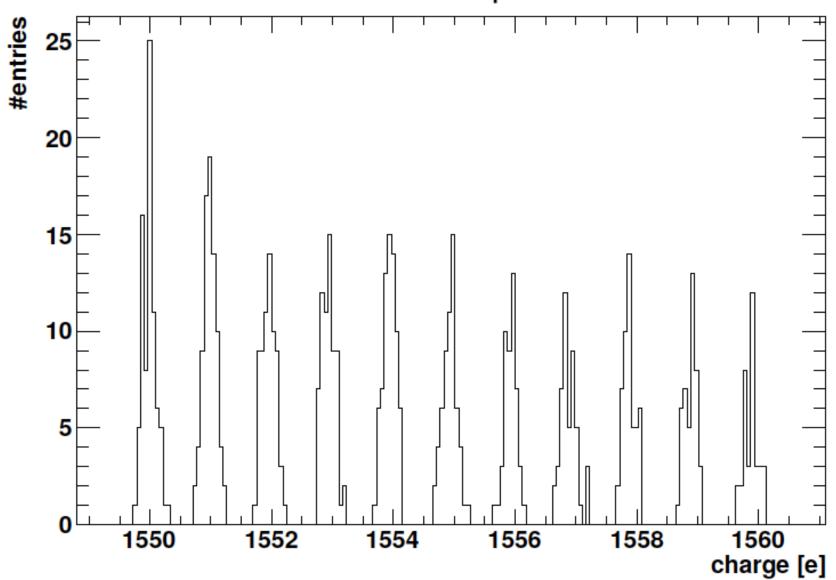
Sub-electron noise!



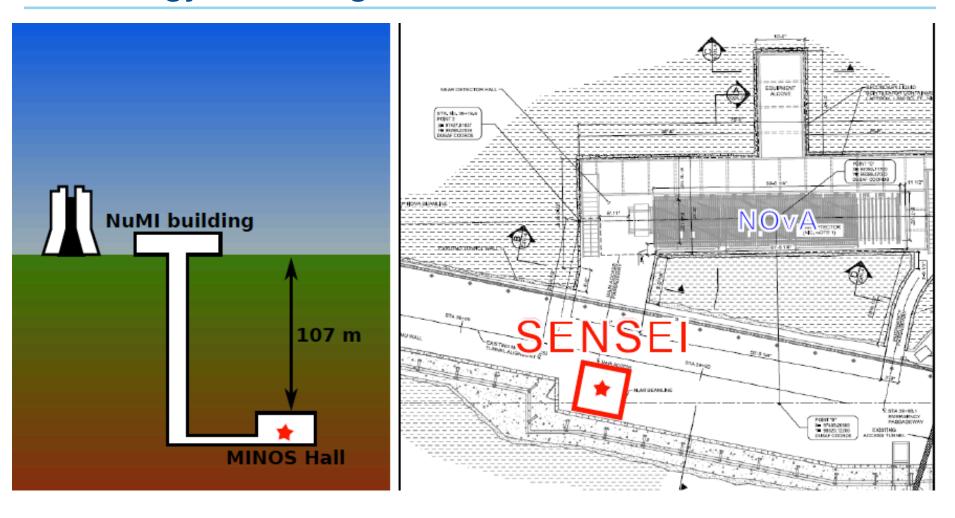








Technology and background demonstrator at FNAL



New low radiation package and clean room area in the MINOS hall



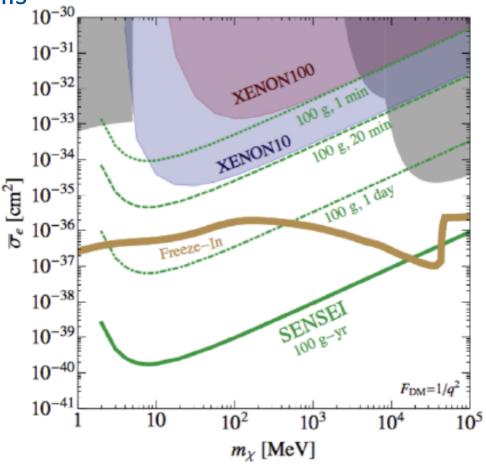
Low mass WIMP Dark Matter last!

New unexplored parameter space for low mass WIMP

dark matter utilizing electron recoils

is enabled by sub-electron noise using CCD with skipper readout.

In addition, skipper CCDs improve the prospects of detecting coherent neutrino interactions at a reactor (CONNIE)



Rouven Essig, Tomer Volansky & Tien-Tien Yu



Conclusions

- Sensors leading to quantum sensors have been at Fermilab
 - Silicon Detectors, Low temperatures devices (Super-CDMS)
 - CMB S3 deployed, MKIDs R&D leadership for cosmology
- New activities in Quantum Sensors benefit FNAL and QIS
 - For high mass axions, need qubits to achieve low bit error rates
 - For SRF cavities, enhance the qubit lifetimes
- QIS activities are enhanced by FNAL core competencies and leverage Chicago-land strength
 - Chicagoland strength in superconducting electronics -- sensors and readout, Fermilab core capability in detector integration, test facilities
 - Facilities can attract outside users (atom interferometry, users of dilution refrigerators with unique capabilities, new experiments)

