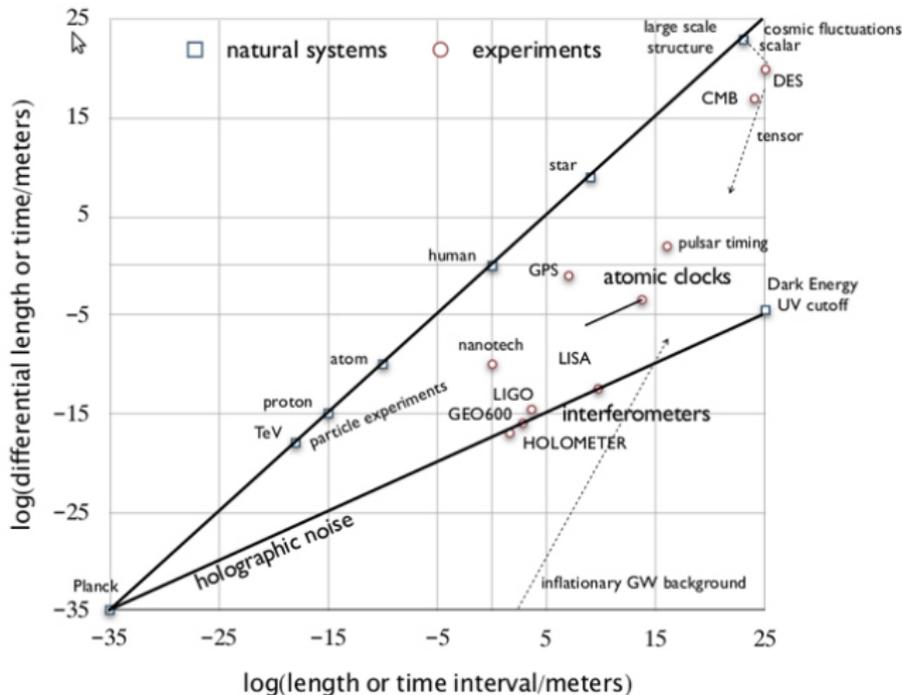


- 1 Background
- 2 Stars
- 3 From Stars to the Big Bang
- 4 Precision Cosmology
- 5 Final Ideas

# Background

# Length Scales



# Wonder

For it is through wondering that human-beings both now and at first began to philosophize, wondering first about the strange things close at hand, and then little by little in this way devotedly exerting themselves and coming to impasses about greater things, such as about the attributes of the moon and things pertaining to the sun and the stars and the coming-into-being of the whole.  
(Aristotle, *Metaphysics*, I.2, 982b11-17)

# We “know” that the Earth is Round

## THEORY

**Aristotle** Every portion of the Earth tends toward the center until by compression and convergence they form a sphere. (De caelo, 297a9-21)

## CIRCUMSTANTIAL EVIDENCE

**Sailors** see tops of mountains; mast disappears last

**Travelers** see Sun and stars at different elevations

**Shadow** of earth on moon during eclipse is round

**Eratosthenes** measured circumference at Summer Solstice in 340 B.C. to  $\sim 10\%$

**Magellan et al.** circumnavigation 1519-1522

## DIRECT OBSERVATION

**NNHS** you could repeat this

**NASA** you'll need more funding for this.

# Naperville North High School



Weather Balloon Launched to 100,000 ft

# Apollo 17



December 7, 1972



# Limits

**Physical Cosmology** Is matter all that matters?

**Observation**  $\leftrightarrow$  **Model** What if 2 models explain all observations?

**Model**  $\leftrightarrow$  **Interpretation** How to pick the “best” interpretation?

**Not Fair Questions** (but great questions for other discussions!)

- Why is there something rather than nothing?
- What is before the beginning?
- What is man and why is he here?

**Godel's Incompleteness** a theory cannot be both consistent and complete

- Wolpert applies this in “Physical limits of inference” (arXiv.0708.1362v2)

# Are you sure about that?

**Perception** your own observation

**Introspection** perception with a special object

**Memory** iirc

**Reason** mathematics, geometry, logic

**Testimony** because I said so!

Different criteria for justifying knowledge.

Which ones are used in scientific methods?

# Scientific Methods

(note the s)

Observation  $\leftrightarrow$  Model  $\leftrightarrow$  Interpretation

- A good mathematical model uses a “few” parameters to describe “many” observations.
- We “know” this model is true.
- Gives us powers of prediction
- Allows us to manipulate matter to do wonderful things. (aka Technology or Engineering.)
- grow more and better food cheaper
- walk on the moon
- see inside your body without cutting you open
- ...
- make us more wealthy

# The story of $F_G = Gm_1m_2/r^2$

- allows stable orbits
- always attracts
- explains the motion of apples, moons, and planets

BUT

- does not explain details of orbits (precession)
- mass bends light
- gravitational redshift

# Olber's Paradox

**Model** Infinite Static Universe, with uniform distribution of stars

**Predict** What would the sky look like?

**Observe** What does the sky look like?

# Stars

# Solar Power

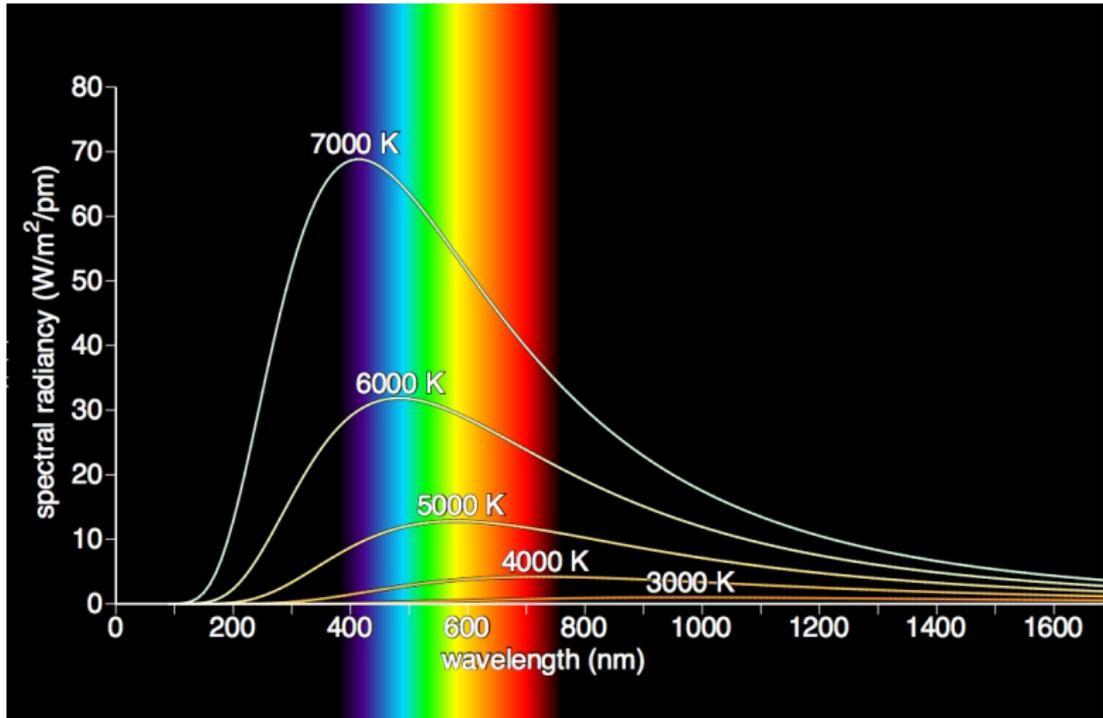
What is the source of the Sun's energy?

We need a “model” consistent with what we know about the Sun

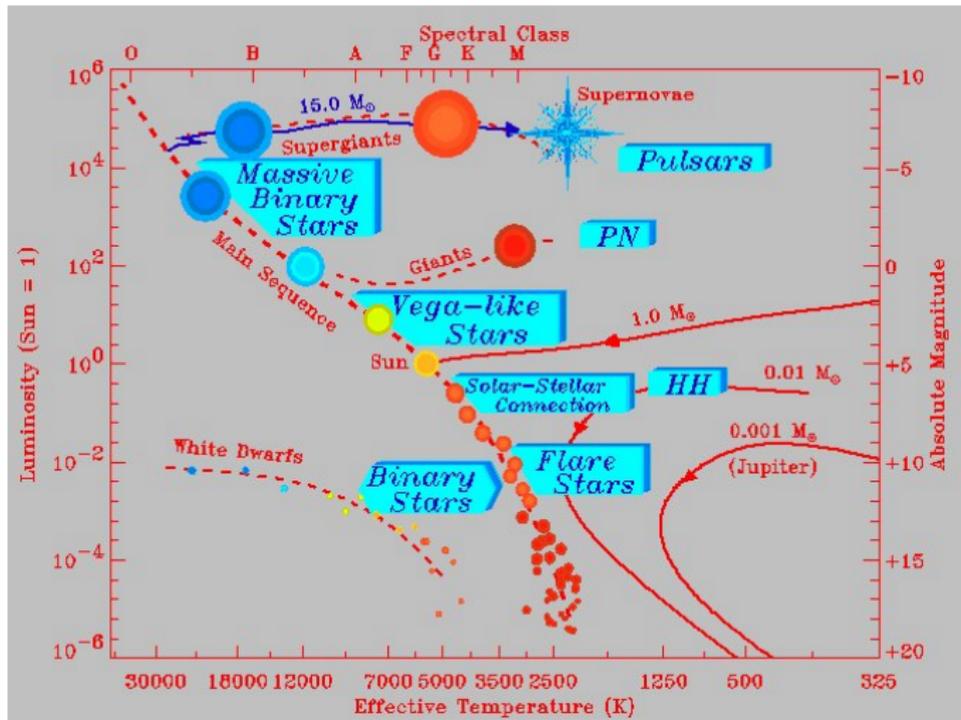
- $M_{Sun} = 2 \times 10^{33} \text{ grams}$
- $T_{Sun} > T_{Earth} > T_{Rocks} \sim 4 \times 10^9 \text{ years}$

Model	Duration (years)
Chemical Burning	$6.3 \times 10^3$
Gravitational Potential Energy	$16 \times 10^6$
Fusion	$100 \times 10^9$ (simple estimate)

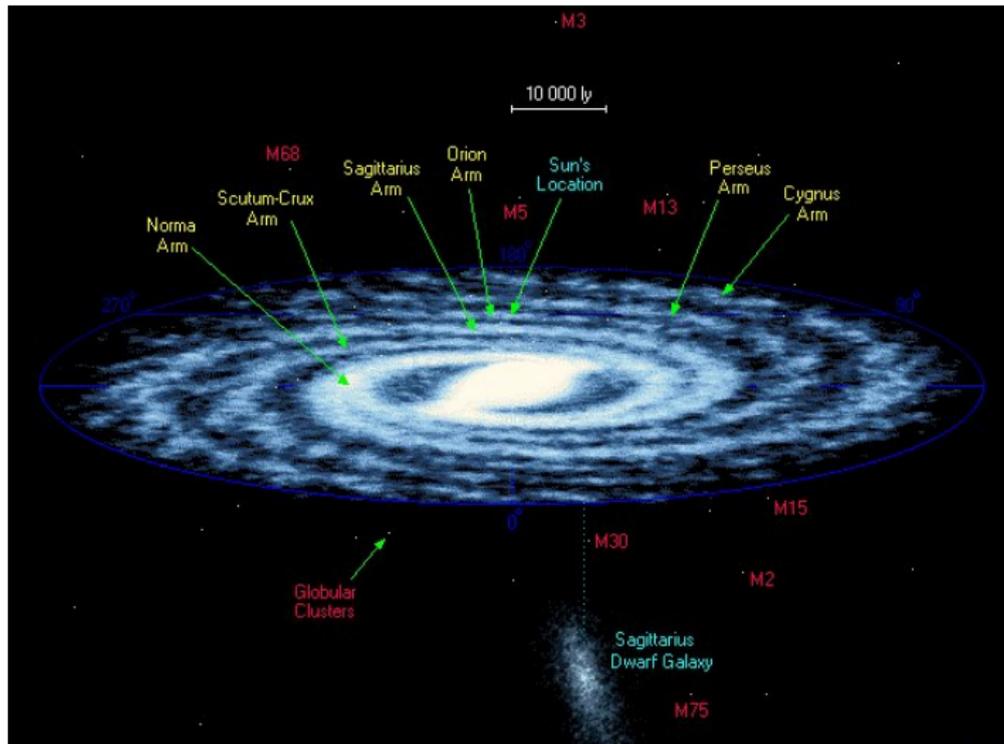
# Stars and Black Body Radiation



# The Lives of Stars



# A Map of the Milky Way



# Looking Out from the Milky Way

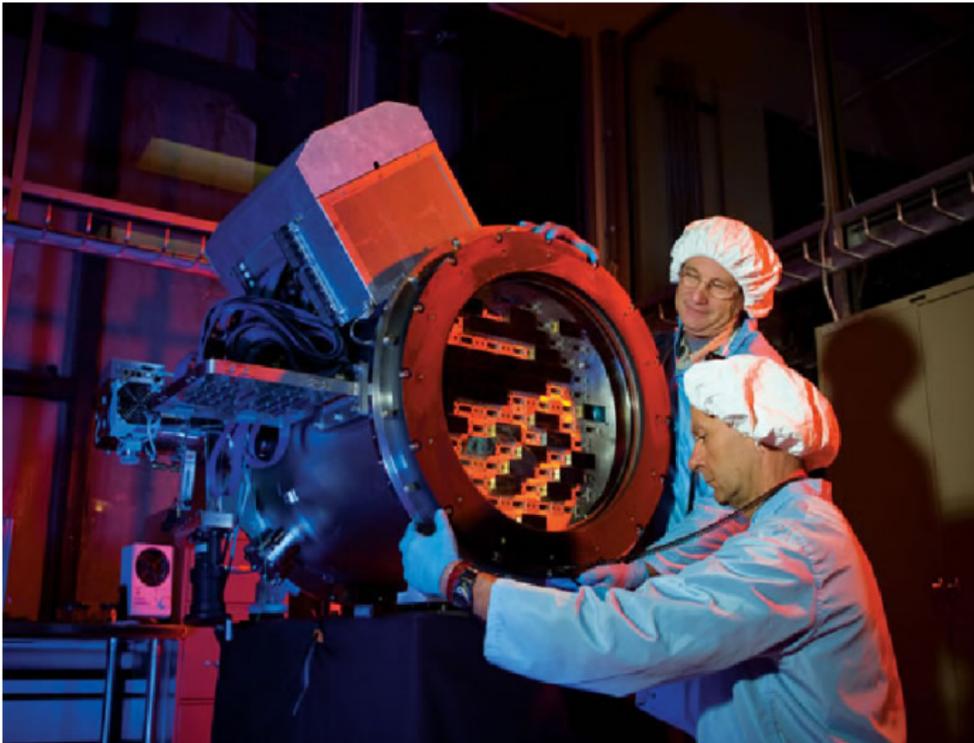


# Stars Essentials for Cosmologists

- one number ( $T$ ) describes overall color and lifetime
- heavy stars are hot and bright and burn out fast
- turn that around: given how bright a star is, we know the mass
- various things can make a star get brighter all of a sudden (nova); if a star is heavy enough, it ends with a bang (Super Nova); a star that is just under this mass can slowly siphon stuff from a neighbor. When it gets to the limit it is SNIa – all blow up with the same mass
- metals from previous generations “contaminate” the atmosphere giving absorption lines

# From Stars to the Big Bang

# The Business End of Modern Cosmology Telescope



# What else is out there?



# Nebulae

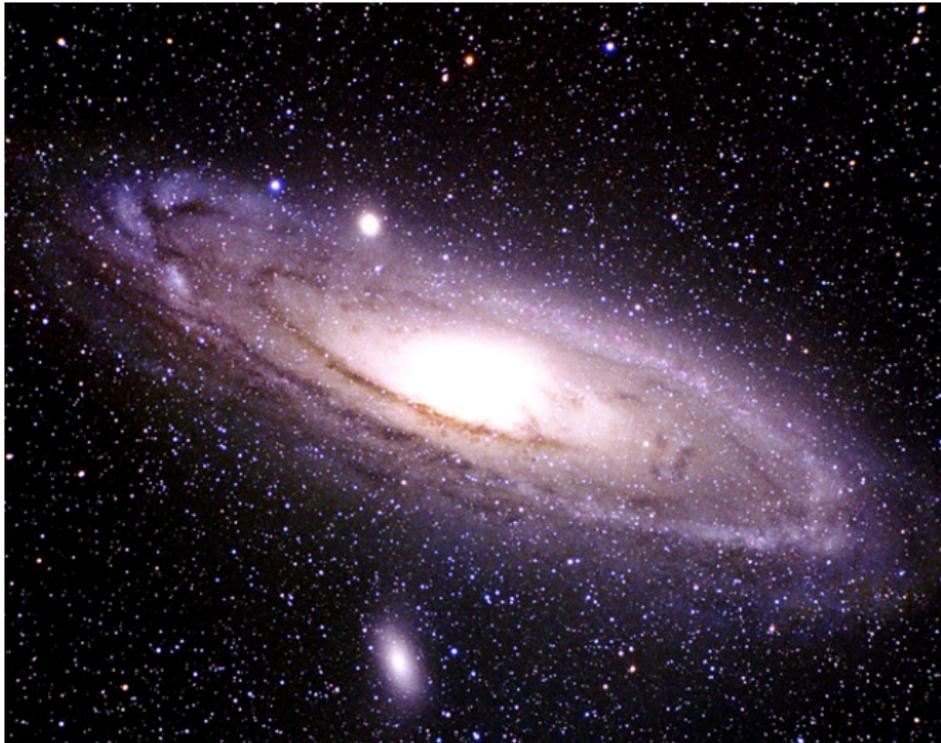
**Thomas Wright** in 1750 – An original theory or new hypothesis of the Universe: Milky Way is a flattened disk; nebulae are “outside”.

**Immanuel Kant** in 1755 –coined the term “Island Universes”

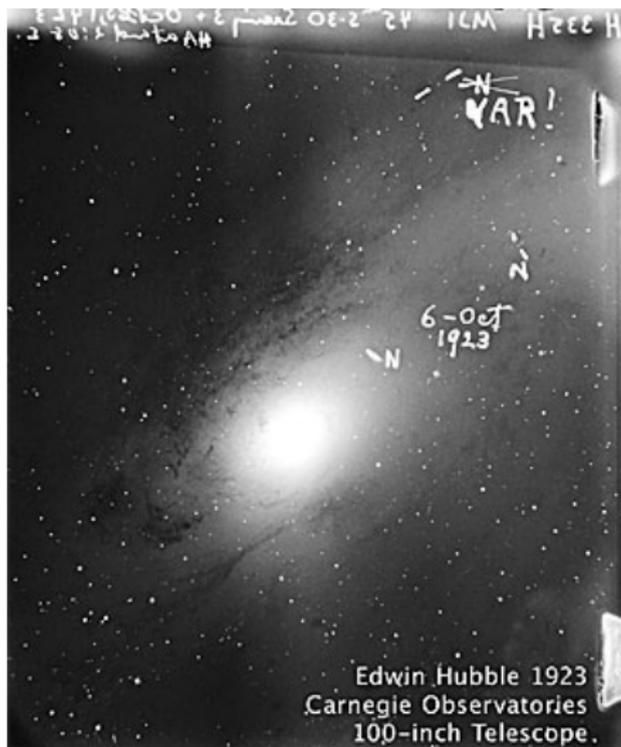
**Vesto Slipher** in 1912 – looking for signs of life in galaxy spectra, saw they were mostly receding, not gravitationally bound to Milky Way.

**Harlow Shapley and Heber Curtis** in 1920 – the “Great Debate”

# Andromeda Nebula – M31

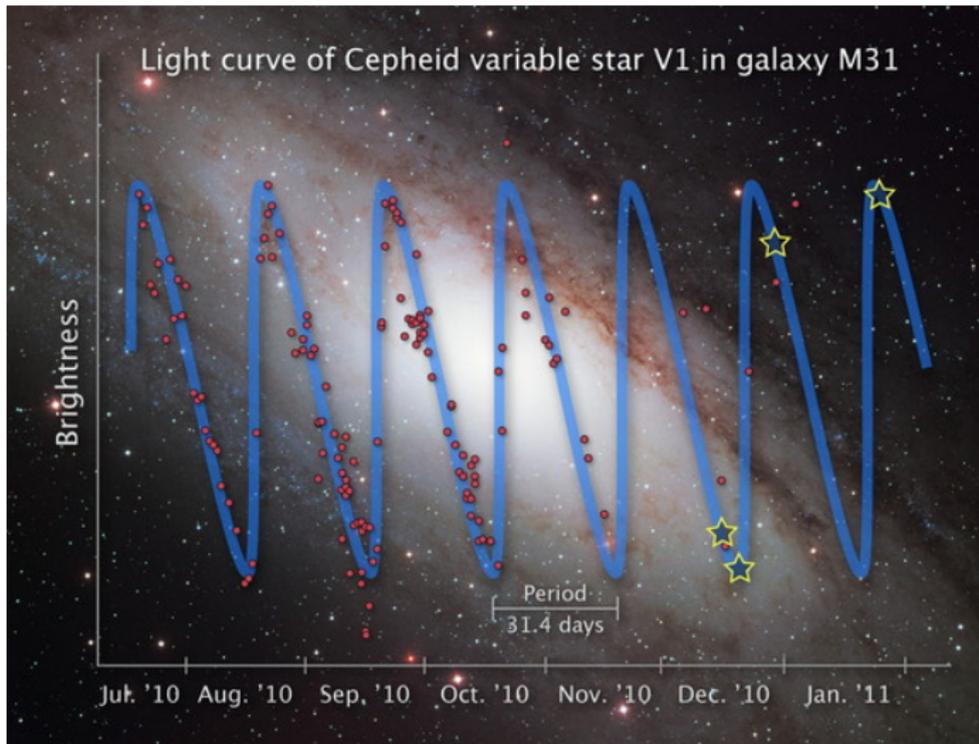


# Andromeda Nebula – Hubble V1



Edwin Hubble 1923  
Carnegie Observatories  
100-inch Telescope.

# Andromeda Nebula – Hubble V1



# Henrietta Leavitt – Cepheid Period/Luminosity

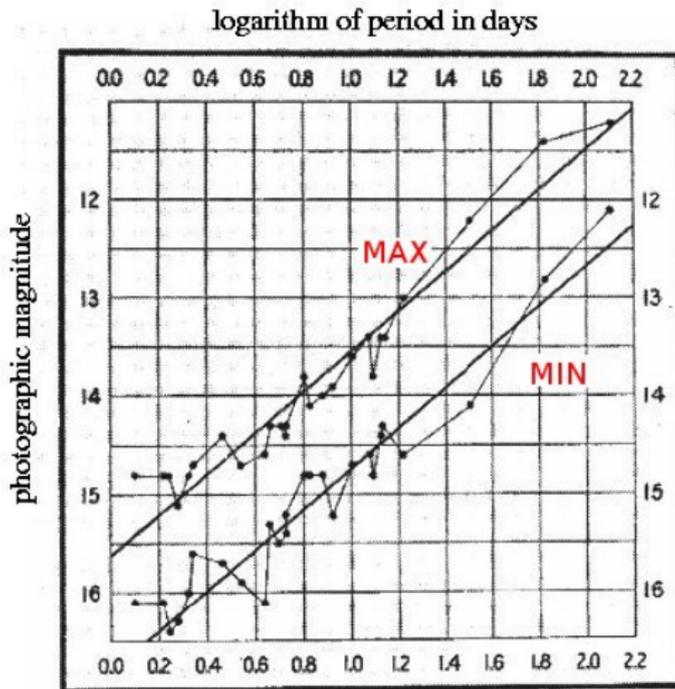
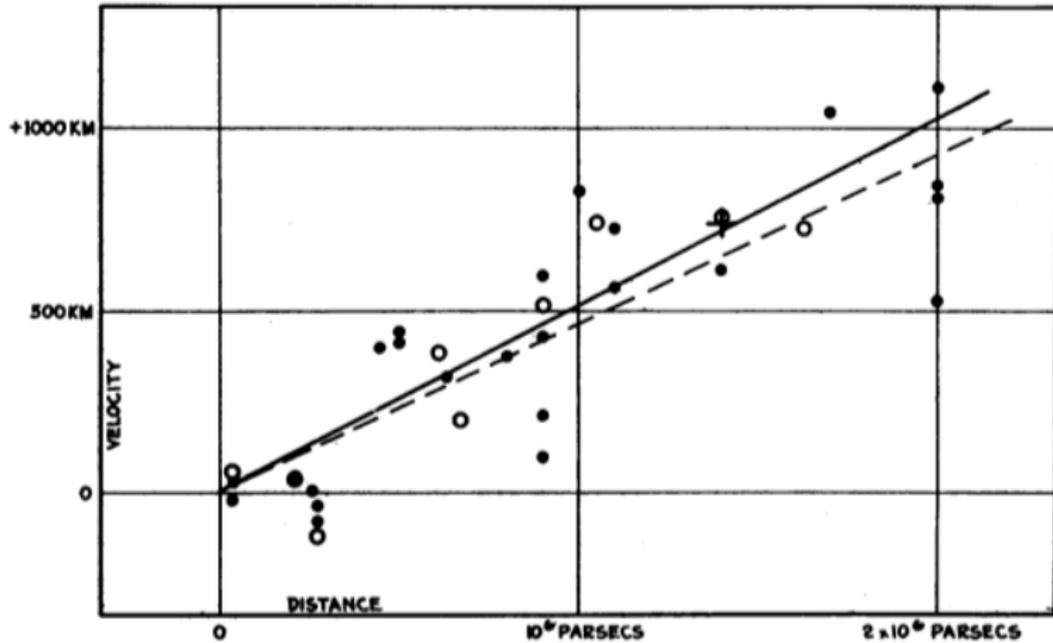


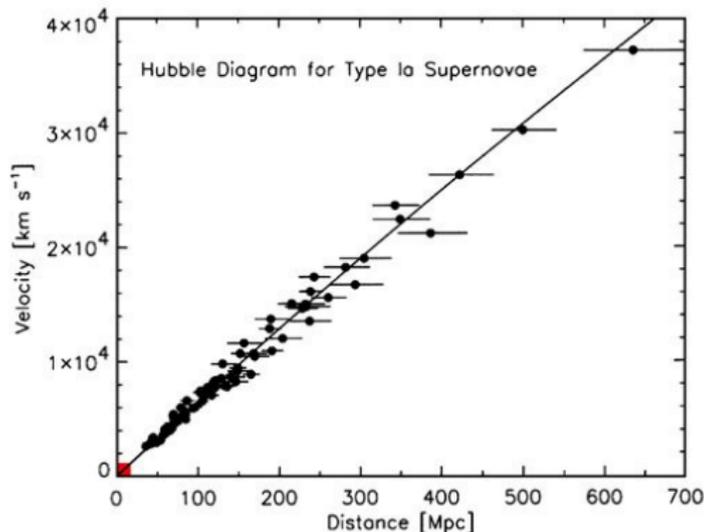
Fig. 2.

# Hubble Diagram – Original



# Hubble Diagram – Modern

Fig. 3. The Hubble diagram for type Ia supernovae



Kirshner, Robert P. (2004) Proc. Natl. Acad. Sci. USA 101, 8-13

## Do we “know” the Universe is expanding?

It did not help adherents of the Big Bang that Gamow was its most vocal supporter. Or that Einstein, now living out his remaining years in Princeton as the world’s most famous scientist, was still philosophically more comfortable with a static universe.

Or, most important, that [Ralph] Alpher and [Robert] Herman’s prediction of the cosmic background radiation, which could not plausibly be accounted for in steady-state cosmology, had been all but forgotten during the 1950s. With problems on both sides, neither was a clear winner. This was how matters stood until the early spring of 1965, cosmology stalemated.

## Do we “know” the Universe is expanding?

Had Arno Penzias and Robert Wilson known in 1964 of the prediction of Alpher and Herman sixteen years earlier, the two Bell scientists would have been spared a year's work trying to uncover the source of the noise in their horn antenna. Had [Robert] Dicke been aware of the prediction, he could have begun work on his own antenna years earlier without having to wait for Jim Peebles to do the theoretical calculations from scratch.

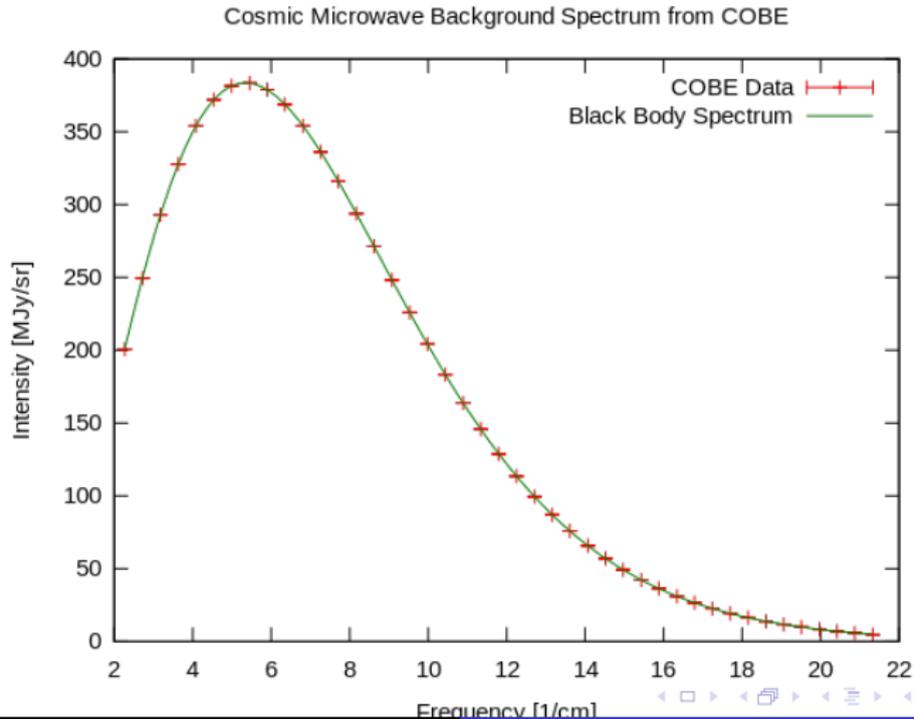
John C. Mather, John Boslough, *The Very First Light*, New York: Basic Books, 1996, pp. 49-50.

# Penzias and Wilson – Cosmic Microwave Background

1964

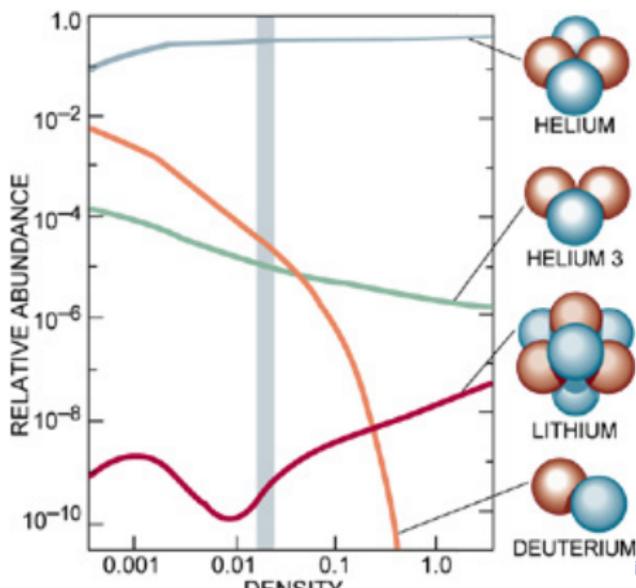


# COBE – measures 3 degree K



# BBN – Big Bang Nucleosynthesis

The fact that helium is nowhere seen to have an abundance below 23% mass is very strong evidence that the Universe went through an early hot phase.



# HALFTIME

- We “know” that the Universe is Expanding
- We’ll talk about HOW it is expanding in the second half
- And what we are doing about it.



# Precision Cosmology

# The Cosmological Principle

**Homogenous** the Universe is the same everywhere

**Isotropic** the Universe looks the same in all directions

- We assume that the Universe is “playing fair” with us
- A stronger condition would be to make it the same at all times
- We can test how “good” these assumptions are, but not completely



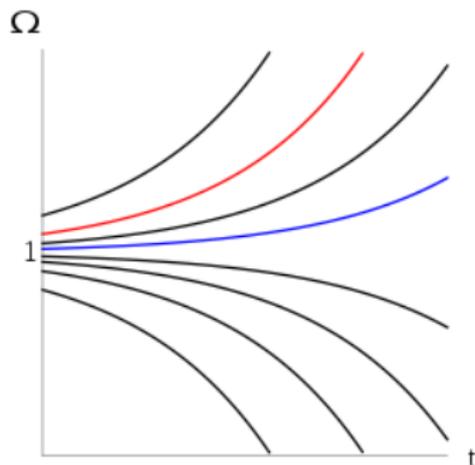
# Problems with the Big Bang

**Flatness** if the density is  $\sim 1$  now, it must have been really close to 1 early on

**Horizon** why is the Cosmic Microwave Background so uniform?  
How does the left side know what the right side is doing?

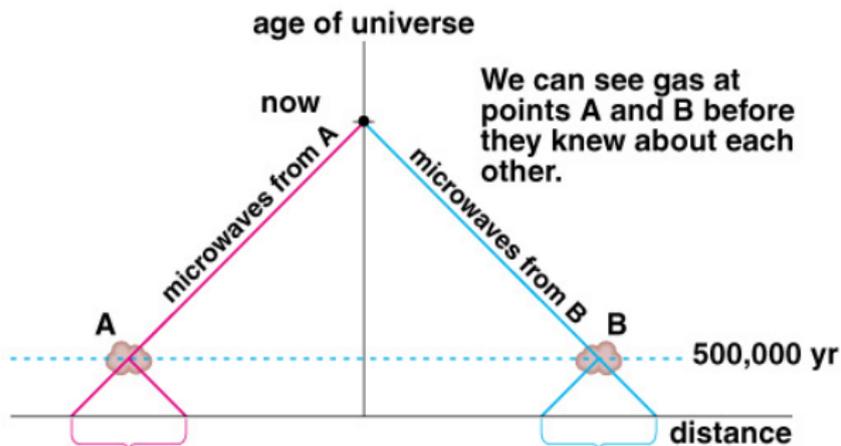
**Magnetic Monopoles** where are they?

# The Flatness Problem



$|\Omega - 1|$  is currently less than 0.01, and therefore must have been less than  $1\text{E-}62$  at the Planck era. This “fine tuning” makes cosmologists nervous.

# The Horizon Problem

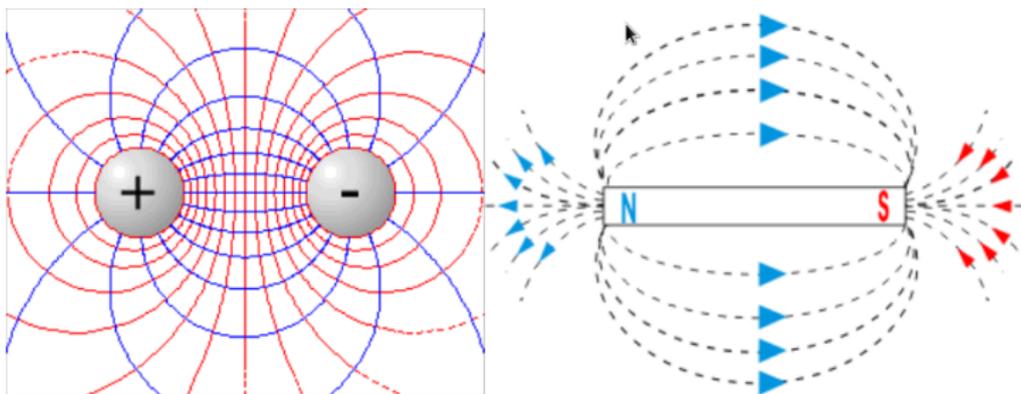


Gas at point A has received signals from this part of the universe.

Gas at point B has received signals from this part of the universe.

Copyright © Addison Wesley.

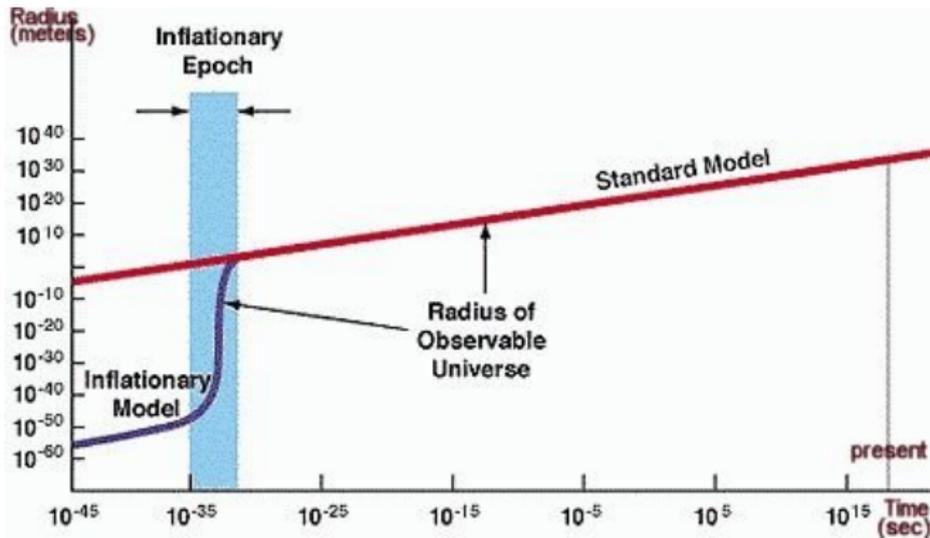
# The Magnetic Monopole Problem



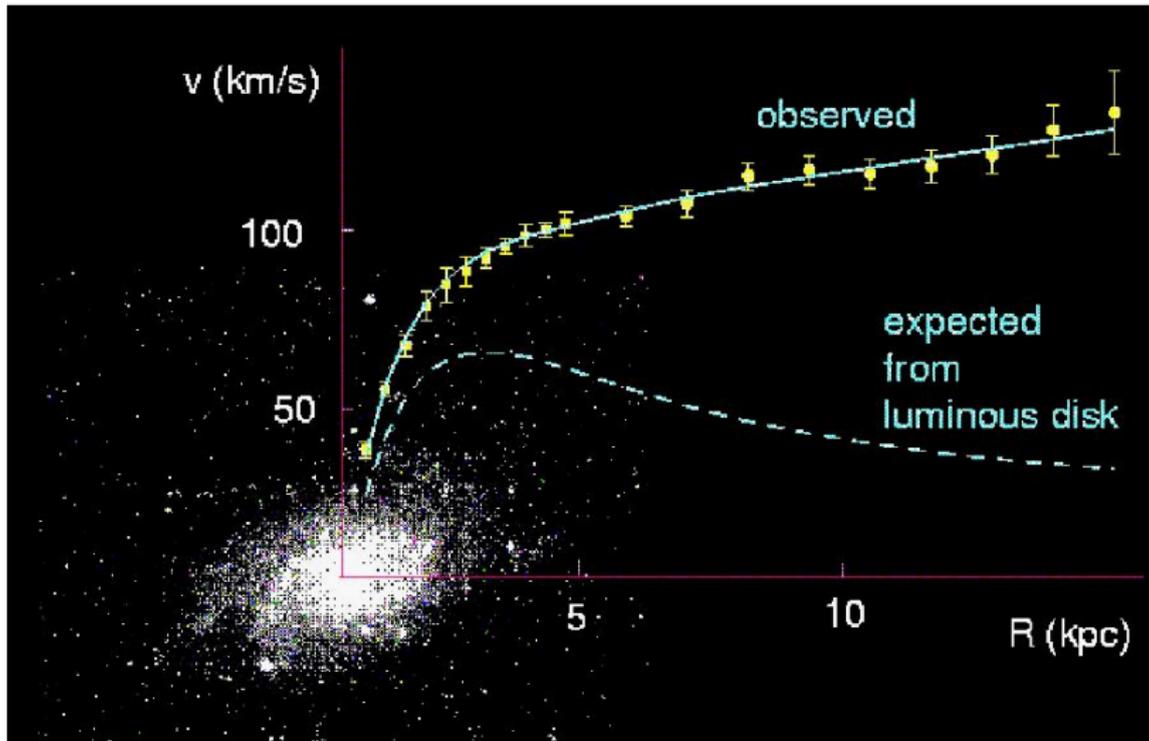
Grand Unified Theories (GUT) predict magnetic monopoles are a dominant part of the universe. BUT – we have not even seen one.

# Cosmic Inflation

The “scalar field” of GUT (invented for particle physics) drives an exponential expansion.



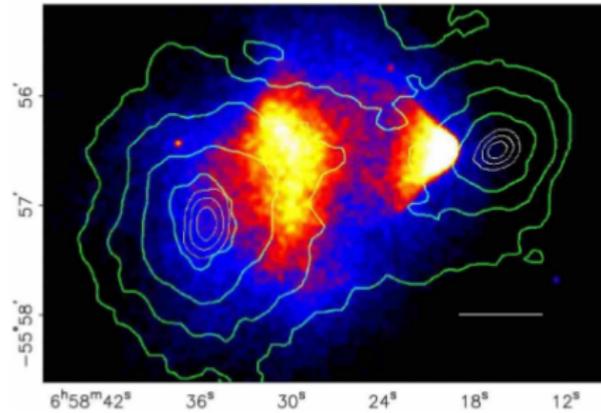
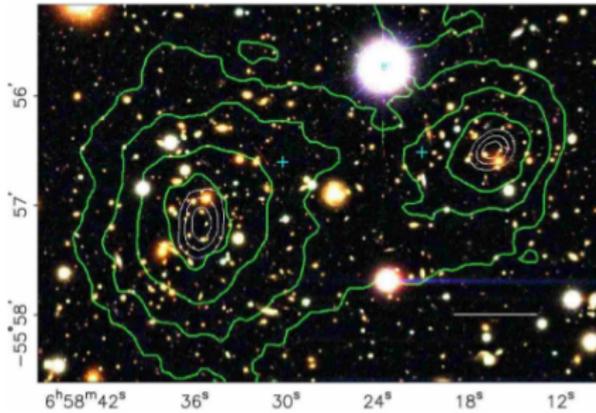
# Dark Matter – galaxy rotation curve



# Dark Matter – strong gravitational lensing



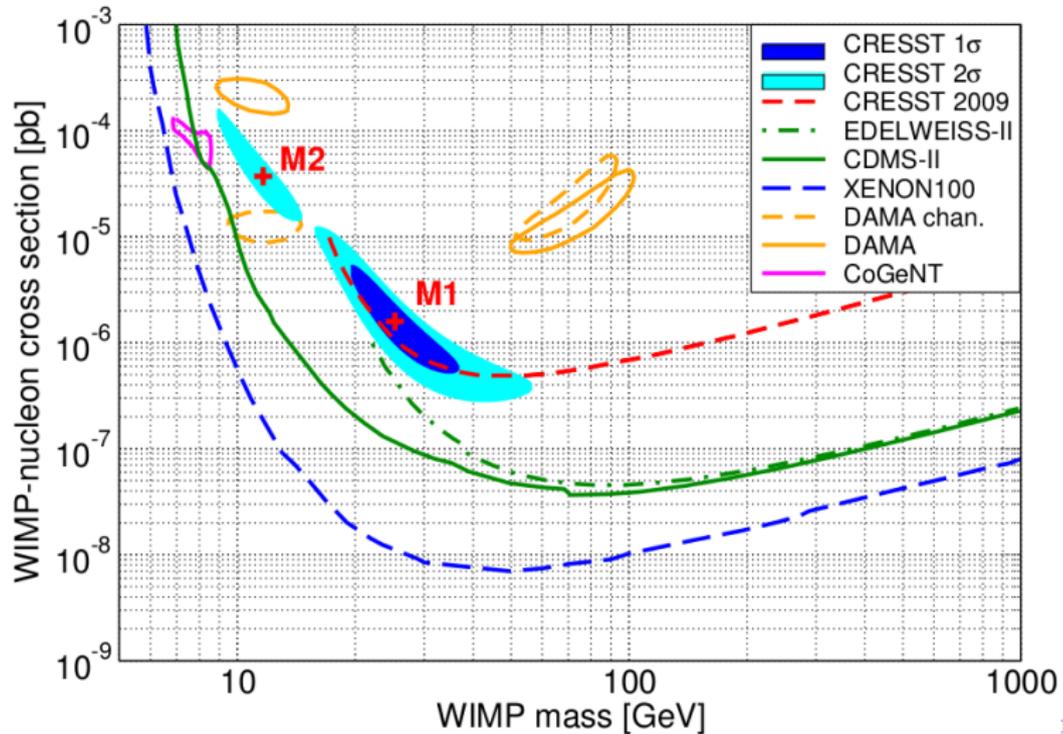
# Dark Matter – weak gravitational lensing



# Dark Matter – direct detection of Particles



# Dark Matter – direct detection



# Dark Matter – indirect detection

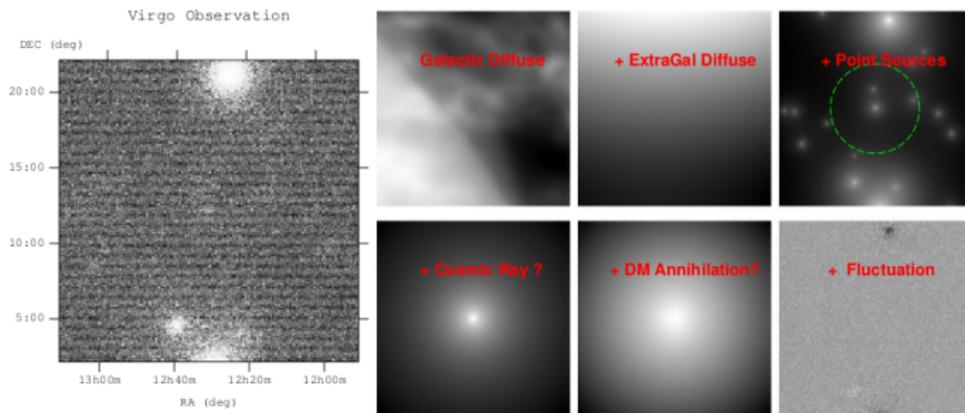
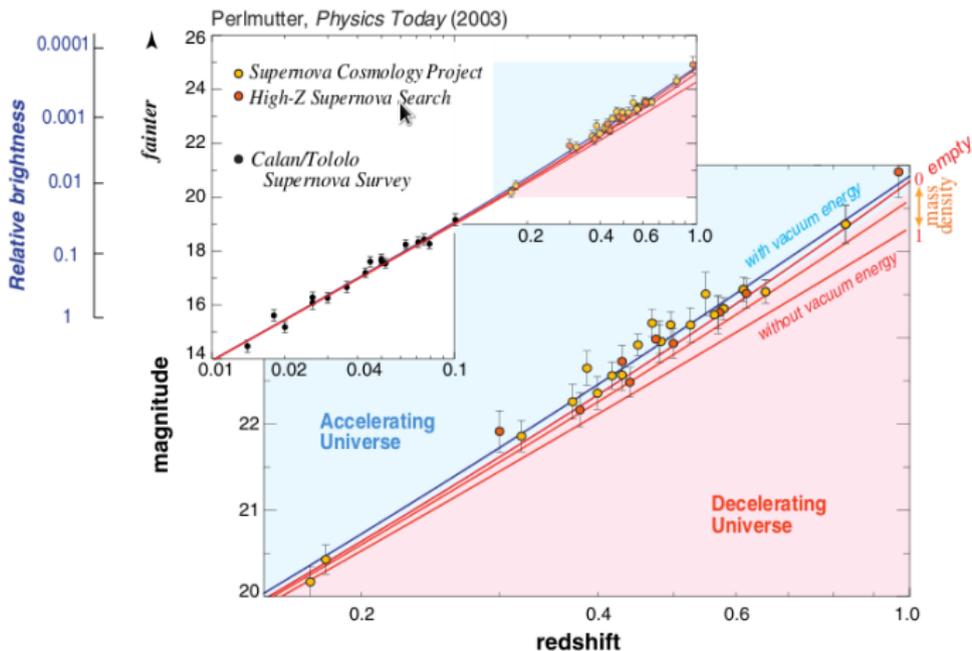


Fig. 1.— Decomposition of the Fermi-LAT image in the region of the Virgo cluster into model components. The observed photon count image from 100MeV to 100 GeV is shown on the left.

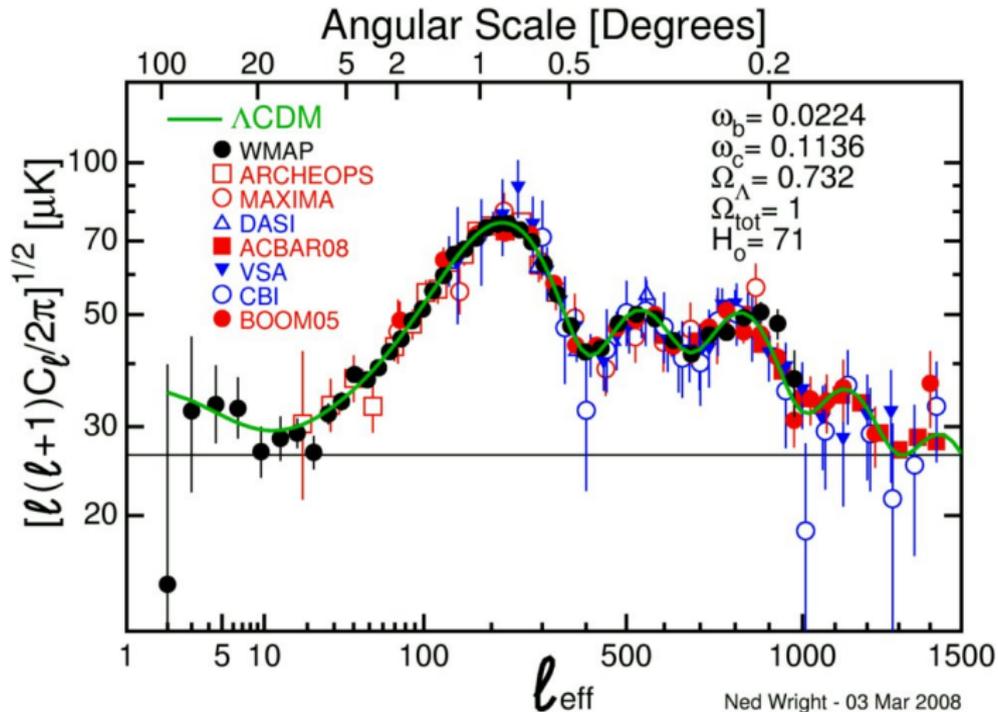
Mass of 20-60 GeV (annihilating into the  $b\bar{b}$  channel)  
consistent with Hooper et al. analysis of Galactic center.

# Dark Energy – first detection

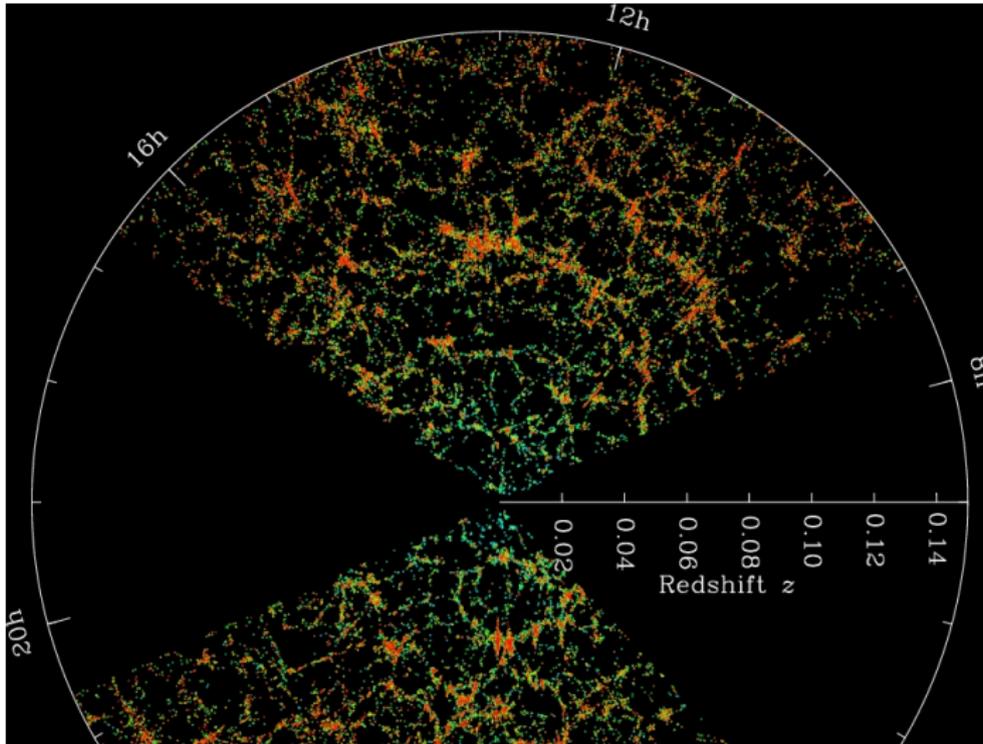
## Type Ia Supernovae



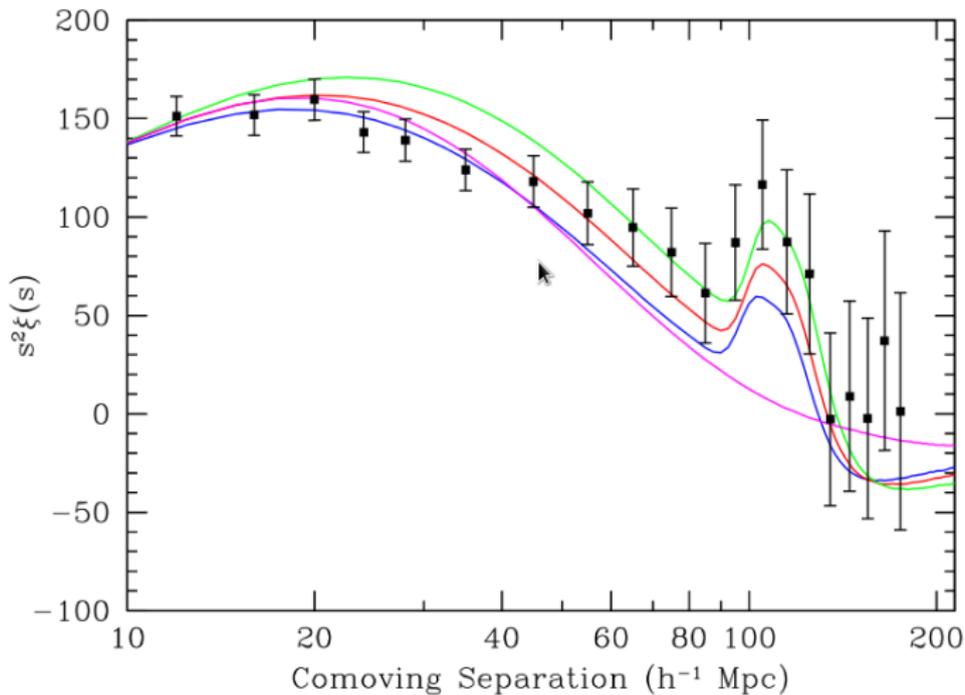
# Dark Energy – CMB Anisotropies



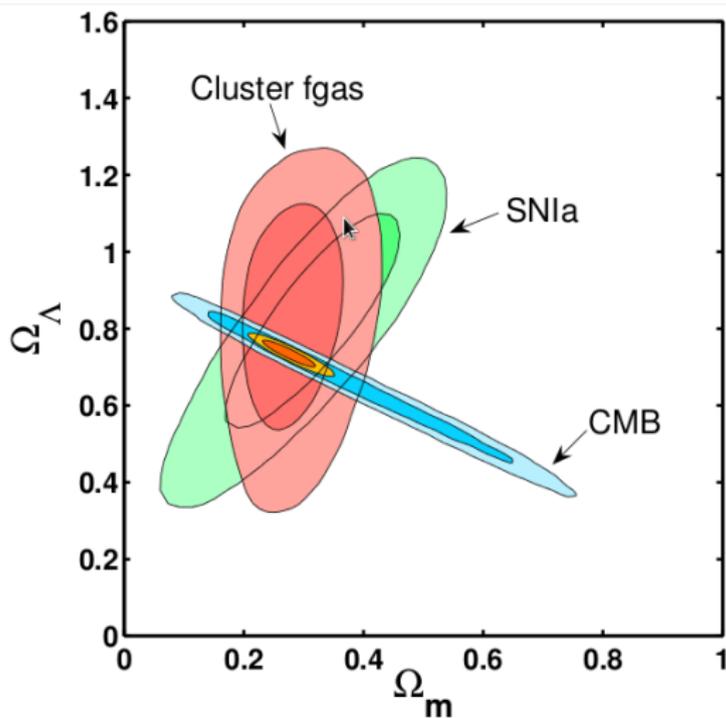
# Dark Energy – Large Scale Structure



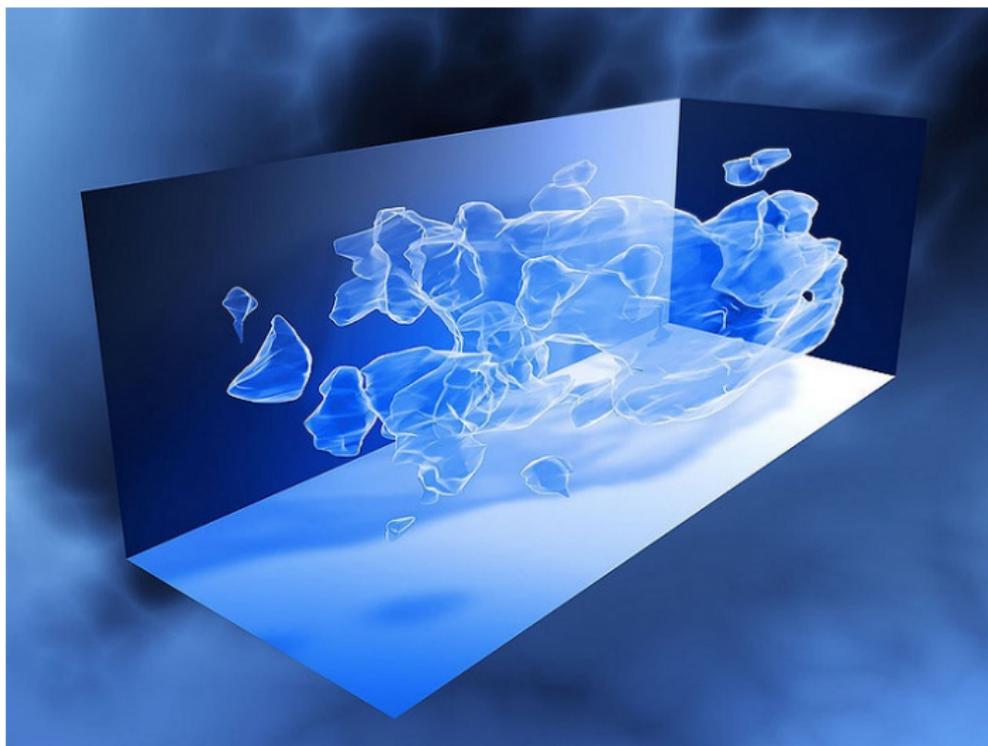
# Dark Energy – Baryon Acoustic Oscillation



# Dark Energy – Galaxy Clusters



## Dark Energy – Weak Lensing



# Lambda Cold Dark Matter Parameters

Parameter	Value	Description
$t_0$	$13.75 \pm 0.11 \times 10^9$ years	Age of the universe
$H_0$	$70.4^{+1.3}_{-1.4}$ km s <sup>-1</sup> Mpc <sup>-1</sup>	Hubble constant
$\Omega_b h^2$	$0.0260 \pm 0.00053$	Physical baryon density
$\Omega_c h^2$	$0.1123 \pm 0.0035$	Physical dark matter density
$\Omega_b$	$0.0456 \pm 0.0016$	Baryon density
$\Omega_c$	$0.227 \pm 0.014$	Dark matter density
$\Omega_\Lambda$	$0.728^{+0.015}_{-0.016}$	Dark energy density
$\Delta_R^2$	$2.441^{+0.088}_{-0.092} \times 10^{-9}$ , $k_0 = 0.002 \text{Mpc}^{-1}$	Curvature fluctuation amplitude
$\sigma_8$	$0.809 \pm 0.024$	Fluctuation amplitude at $8h^{-1} \text{Mpc}$
$n_s$	$0.963 \pm 0.012$	Scalar spectral index
$z_*$	$1090.89^{+0.68}_{-0.69}$	Redshift at decoupling
$t_*$	$377730^{+3205}_{-3200}$ years	Age at decoupling
$\tau$	$0.087 \pm 0.014$	Reionization optical depth
$z_{\text{reion}}$	$10.4 \pm 1.2$	Redshift of reionization

# We “know” Lambda Cold Dark Matter Describes the Universe

- a few parameters describe many different observations
- there are alternate ways to account for some of the observations
- are the parameters constant in time?
- it sure would be nice to have a direct detection of dark matter
- laboratory tests of dark energy? We can't think of any.

## Final Ideas

# Black Holes

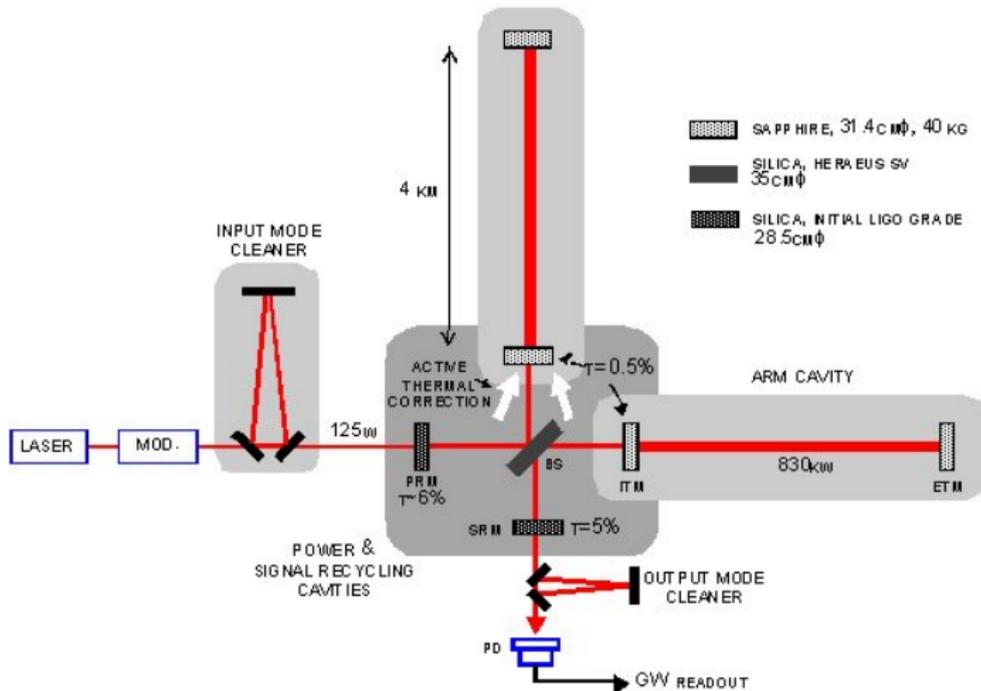
- M is contained inside  $r_{sh} = \frac{2GM}{c^2}$
- light black holes evaporate
- not directly observed, but we see effects
  - rotation of stars near center of Milky Way
  - X-rays from binary systems
  - Quasars/Active Galactic Nuclei/Galaxies
- Entropy  $S = \frac{\pi A k c^3}{2hG}$  where A is surface area. Normally, S is proportional to Volume  $\sim$  information content.
- Leads to “holographic principle” that everything in a volume of spacetime can be written on the surface. (A hologram is a 2d “projection” of 3d space)

# Gravitational Waves

Predicted by General Relativity. Sources include

- black hole/neutron star pair inspiraling
- black hole/neutron star near miss
- black hole mergers
- supernovae
- gamma ray bursts
- spinning neutron stars
- relics from inflation
- the unexpected

# Advanced LIGO – gravitational wave detector



## More Future Projects

The Fermilab Holometer will measure “pixel size” of spacetime

Dark Energy Survey will use the 4 methods to measure dark energy

Large Synoptic Survey Telescope will repeat the 5-year SDSS  
every few weeks!

CMBPOL proposed to measure CMB with higher sensitivity

Square Kilometer Array surveys the radio sky – for example

Hydrogen maps using the 21cm wavelength radiation

TNGI the Next Great Idea....